

Atom Interferometric Tests of Gravity

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Gravitational wave Detection with Atom Interferometry
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Gravity can only be explored through the motion of test particles

Test particles

- Orbits and clocks
- Massive particles and light

What is gravity depends on the structure of the equation of motion

- Existence of inertial systems
- Order of differential equation
- Dependence on particle parameters

Applies to test particles exploring gravitational waves

Gravitational waves may influence physics of test particles

- 1 Questioning Newton's laws
 - Newton's first law: Inertial systems
 - Newton's second law: The law of inertia
 - Newton's third law: Law of reciprocal action

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 - Violation of UFF from space-time fluctuations
 - Universality of Free Fall for charged particles
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Finsler geometry

Motivation

- generic generalization of GR
- leads to deformed light cones and mass shells
- has been discussed within Quantum Gravity (Jacobson, Liberati & Mattingly)
- Very Special Relativity (Cohen & Glashow)

Finsler space

Finsler length function

$$ds^2 = F(x, dx), \quad F(x, \lambda dx) = \lambda^2 F(x, dx)$$

Finsler metric tensor $f_{\mu\nu}(x, dx)$ is defined as

$$ds^2 = g_{\mu\nu}(x, dx) dx^\mu dx^\nu, \quad \text{where} \quad g_{\mu\nu}(x, y) = \frac{1}{2} \frac{\partial^2 F^2(x^k, y^m)}{\partial y^\mu \partial y^\nu}$$

Finsler geometry

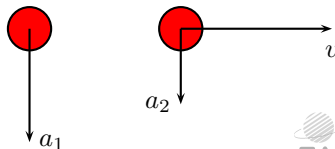
Geodesics

$$\delta \int ds = 0 \quad \Rightarrow \quad 0 = \frac{d^2 x^\mu}{ds^2} + \{ \overset{\mu}{\rho\sigma} \} (x, \dot{x}) \frac{dx^\rho}{ds} \frac{dx^\sigma}{ds}$$

with $\{ \overset{\mu}{\rho\sigma} \} (x, \dot{x}) = g^{\mu\nu} (x, \dot{x}) (\partial_\rho g_{\sigma\nu} (x, \dot{x}) + \partial_\sigma g_{\rho\nu} (x, \dot{x}) - \partial_\nu g_{\rho\sigma} (x, \dot{x}))$

Main characteristics of geodesic motion

- Geodesic equation fulfills Universality of Free Fall
- $\{ \overset{\mu}{\rho\sigma} \} (x, \dot{x})$ cannot be transformed to zero $\forall \dot{x} \Rightarrow$ **gravity cannot be transformed away locally** \Leftrightarrow Einstein's elevator does not hold \Leftrightarrow no inertial system
- Condition to be able to transform away gravity is stronger than pure UFF.
- Acceleration toward the Earth depends on horizontal velocity.
- Speculation: violation of UGR, $G = G(T)$



Parametrizing deviations from Riemann/Minkowski

Special case: “power law” metrics

$$ds^2 = (g_{\mu_1\mu_2\dots\mu_{2n}}(x) dx^{\mu_1} dx^{\mu_2} \dots dx^{\mu_{2n}})^{\frac{1}{r}}$$

Deviation from Riemann/Minkowski

$$ds^{2r} = (g_{\mu_1\mu_2} \dots g_{\mu_{2r-1}\mu_{2r}} + \phi_{\mu_1\dots\mu_{2r}}) dx^{\mu_1} \dots dx^{\mu_{2r}}$$

This gives

$$ds^2 = (g_{\mu\nu} + \phi_{\mu\nu\rho_3\dots\rho_{2r}} n^{\rho_3} \dots n^{\rho_{2r}}) dx^\mu dx^\nu, \quad n^\mu = \frac{dx^\mu}{\sqrt{g_{\rho\sigma} dx^\rho dx^\sigma}}$$

Additional assumption: $\phi_{\mu_1\dots\mu_{2r}}$ possesses spatial indices only (from light propagation)

Quantum mechanics in Finsler space

Finslerian Hamilton operator

$$H = H(p) \quad \text{with} \quad H(\lambda p) = \lambda^2 H(p)$$

“Power-law” ansatz (non-local operator)

$$H = \frac{1}{2m} (g^{i_1 \dots i_{2r}} \partial_{i_1} \dots \partial_{i_{2r}})^{\frac{1}{r}}$$

Simplest case: quartic metric

$$H = \frac{1}{2m} (g^{ijkl} \partial_i \partial_j \partial_k \partial_l)^{\frac{1}{2}}$$

Deviation from standard case

$$\begin{aligned} H &= -\frac{1}{2m} (\Delta^2 + \phi^{ijkl} \partial_i \partial_j \partial_k \partial_l)^{\frac{1}{2}} \\ &= -\frac{1}{2m} \Delta \sqrt{1 + \frac{\phi^{ijkl} \partial_i \partial_j \partial_k \partial_l}{\Delta^2}} \end{aligned}$$

Quantum mechanics in Finsler space

$$H = -\frac{1}{2m} \Delta \left(1 + \frac{1}{2} \frac{\phi^{ijkl} \partial_i \partial_j \partial_k \partial_l}{\Delta^2} \right)$$

- Hughes–Drever: $H_{\text{tot}} = H + \boldsymbol{\sigma} \cdot \mathbf{B}$
- Atomic interferometry, atom–photon interaction

$$\delta\phi \sim H(p+k) - H(p) = \frac{k^2}{2m} + \frac{1}{m} \left(\delta^{il} + \frac{\phi^{ijkl} p_j p_k}{p^2} \right) p_i k_l$$

modified Doppler term: gives different Doppler term while rotating the whole apparatus (even in Finsler light still propagates on straight lines, anisotropy – deformed mass shell)

- cf. C.L., Lorek & Dittus 2008 for the photon sector (deformed light cone)

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Order of equation of motion?

Usual framework

$$L = L(t, \mathbf{x}, \dot{\mathbf{x}}) \quad \Rightarrow \quad \frac{d}{dt} (m\dot{\mathbf{x}}) = \mathbf{F}(t, \mathbf{x}, \dot{\mathbf{x}})$$

Most important equation in physics!

More general equations?

- $\mathbf{p} = m\dot{\mathbf{x}}$ is a constitutive law. Can be more general (as is many cases)

$$\mathbf{p} = \mathbf{f}(\dot{\mathbf{x}}, \ddot{\mathbf{x}}, \ddot{\ddot{\mathbf{x}}}, \dots)$$

Then equations of motion of higher order

- Influence of external fluctuations (e.g. space-time fluctuations, gravitational wave background): generalized Langevin equation with extra force term

$$\int_0^t C(t-t')\dot{\mathbf{x}}(t')dt'$$

Order of equation of motion?

Generalized framework

$$L = L(t, \mathbf{x}, \dot{\mathbf{x}}, \ddot{\mathbf{x}}) \quad \Rightarrow \quad \frac{d^2}{dt^2} (\epsilon \ddot{\mathbf{x}}) = \mathbf{F}(t, \mathbf{x}, \dot{\mathbf{x}}, \ddot{\mathbf{x}})$$

Our specific model

$$L(t, \mathbf{x}, \dot{\mathbf{x}}, \ddot{\mathbf{x}}) = L_0(t, \mathbf{x}, \dot{\mathbf{x}}, \ddot{\mathbf{x}}) \quad \underbrace{-q_0 A_a \dot{x}^a}_{\text{1st order gauge fields}} \quad + \quad \underbrace{q_1 A_{ab} \dot{x}^a \dot{x}^b}_{\text{2nd order gauge fields}}$$

with (Pais–Uhlenbeck oscillator)

$$L_0(t, \mathbf{x}, \dot{\mathbf{x}}, \ddot{\mathbf{x}}) = -\frac{\epsilon}{2} \ddot{\mathbf{x}}^2 + \frac{m}{2} \dot{\mathbf{x}}^2$$

C.L. & Rademaker 2009

Equation of motion

simplest case: constant electric field

$$\epsilon \ddot{\mathbf{x}} + m\ddot{\mathbf{x}} = q\mathbf{E}_0$$

solution in 1D with initial conditions $x(0) = 0$, $\dot{x}(0) = 0$, $\ddot{x}(0) = 0$, and $\ddot{\ddot{x}}(0) = 0$

$$x(t) = \frac{q}{m} E_0 \left(\frac{1}{2} t^2 + \frac{\epsilon}{m} (\cos(\omega t) - 1) \right) \quad \text{small deviation}$$

$$\dot{x}(t) = \frac{q}{m} E_0 \left(t - \sqrt{\frac{\epsilon}{m}} \sin(\omega t) \right) \quad \text{small deviation}$$

$$\ddot{x}(t) = \frac{q}{m} E_0 (1 - \cos(\omega t)) \quad \mathcal{O}(1) \text{ deviation}$$

$$\ddot{\ddot{x}}(t) = \frac{q}{m} E_0 \sqrt{\frac{m}{\epsilon}} \sin(\omega t) \quad \omega = \sqrt{\frac{m}{\epsilon}} \quad \text{large deviation}$$

- *zitterbewegung*
- Limit $\epsilon \rightarrow 0$ does not exist

Acceleration variance

(Hadamard) variance

$$\Delta\ddot{x} = \begin{cases} \frac{1}{\sqrt{2}} \frac{q}{m} E_0 & \text{for } \epsilon > 0 \\ 0 & \text{for } \epsilon = 0 \end{cases}$$

Phase shift for this *zitterbewegung* in ion interferometry

Other possibilities

- time of flight measurements
- electronic noise

Applies to mirrors in gw interferometers?

Linearity of law of inertia

Why is the relation between acceleration and force linear?

It is a definition

$\dot{\mathbf{p}} = \mathbf{F}$: exploration of forces through observation of orbits

Meaningful question 1: Test linearity

Taking elements of the field equation into account:

- If $\mathbf{F} = -\nabla U$ with $U = \frac{M}{r}$, then one can ask

$$M \rightarrow \alpha M \quad \stackrel{?}{\implies} \quad \mathbf{F} \rightarrow \alpha \mathbf{F}$$

- Test of field equation/dynamics in the **weak field/small acceleration** domain
- Applies to small relative acceleration of test masses in gw detectors?
- Same for gravity and electromagnetism?
- Experiments
 - Abramovici & Vager, PRD 1986 ($F = q\Delta\phi/L$, down to 10^{-9} m/s^2)
 - Gundlach et al, PRL 2007 (down to 10^{-14} m/s^2)
 - atom interferometry?

Linearity of law of inertia

These questions are motivated by MOND

Meaningful question 2: Free fall experiment

- MOND – dark matter: requires a certain frame of reference (galactic frame)
- MOND–situation possible on Earth once a year for 0.1 s within 1 l volume (Ignatiev, PRL 2007)
- Until now there is no (laboratory) test of MOND — MOND needs space

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Active and passive charges: Dynamics

Bondi RMP 1956; C.L., Macias, Müller, PRA 2007

Dynamics of two electrically bound particles (\mathbf{E} = external electric field)

$$m_{1i}\ddot{\mathbf{x}}_1 = q_{1p}q_{2a}\frac{\mathbf{x}_2 - \mathbf{x}_1}{|\mathbf{x}_2 - \mathbf{x}_1|^3} + q_{1p}\mathbf{E}(\mathbf{x}_1),$$

$$m_{2i}\ddot{\mathbf{x}}_2 = q_{2p}q_{1a}\frac{\mathbf{x}_1 - \mathbf{x}_2}{|\mathbf{x}_1 - \mathbf{x}_2|^3} + q_{2p}\mathbf{E}(\mathbf{x}_2),$$

center-of-mass and relative coordinate

$$\mathbf{X} := \frac{m_{1i}}{M_i}\mathbf{x}_1 + \frac{m_{2i}}{M_i}\mathbf{x}_2, \quad \mathbf{x} := \mathbf{x}_2 - \mathbf{x}_1,$$

$M_i = m_{1i} + m_{2i}$ = total inertial mass. Then

$$\ddot{\mathbf{X}} = \frac{q_{1p}q_{2p}}{M_i}C_{21}\frac{\mathbf{x}}{|\mathbf{x}|^3} + \frac{1}{M_i}(q_{1p} + q_{2p})\mathbf{E} \quad C_{21} := \frac{q_{2a}}{q_{2p}} - \frac{q_{1a}}{q_{1p}}.$$

$C_{21} = 0$: ratio between the active and passive charge is the same for both particles

$C_{21} \neq 0 \Rightarrow$ self-acceleration of center of mass along \mathbf{x}



Active and passive charges: Dynamics

Dynamics of relative coordinate

$$\ddot{\mathbf{x}} = -\frac{1}{m_{\text{red}}} q_{1\text{p}} q_{2\text{p}} D_{21} \frac{\mathbf{x}}{|\mathbf{x}|^3}, \quad (1)$$

where

$$D_{21} = \frac{m_{1i}}{M_i} \frac{q_{1a}}{q_{1\text{p}}} + \frac{m_{2i}}{M_i} \frac{q_{2a}}{q_{2\text{p}}} = \frac{q_{1a}}{q_{1\text{p}}} + \frac{m_{2i}}{M_i} C_{21}$$

In the standard framework, $D_{21} = 1$.

Solutions of equation of motion (1) are ellipses, circles.

The center of mass oscillates at a frequency ω , which is related to the energy of the system.

The acceleration of the center of mass vanishes on average, $\langle \ddot{\mathbf{X}} \rangle = 0$. Thus, not observable for atoms.

Extends to many particle systems, e.g., to atoms having many electrons.

Interpretation

$$\ddot{\mathbf{X}} \neq 0 \quad \Leftrightarrow \quad C_{12} \neq 0 \quad \Leftrightarrow \quad \text{violation of } \textit{actio} = \textit{reactio} \text{ for electromagnetism}$$

Strategy to measure C_{12}

Strategy

- Electromagnetic timescales too short: self-acceleration cannot be observed.
- Electric charges can have different signs. Therefore, we can define

active neutrality $q_{1a} + q_{2a} = 0$

passive neutrality $q_{1p} + q_{2p} = 0$

→ alternative tests of the equality of active and passive charges: An actively neutral system may not be passively neutral and vice versa.

active and passive neutrality $\Leftrightarrow C_{21} = 0$

Experiments

tests of neutrality of atoms and molecules = tests of the equality of active and passive charge

- a **passively neutral** system may still generate an electric field according to

$$\phi(\mathbf{x}) = \frac{q_{1a}}{|\mathbf{x} - \mathbf{x}_1|} + \frac{q_{2a}}{|\mathbf{x} - \mathbf{x}_2|} = \frac{q_{1a} + q_{2a}}{|\mathbf{x}|} + \dots \approx C_{21} \frac{q_{2p}}{|\mathbf{x}|}$$

- an **actively neutral** atom in an external electric field may feel a force

$$M_i \ddot{\mathbf{X}} = (q_{p1} + q_{p2}) \mathbf{E} = \frac{q_{2p}}{q_{2a}} q_{1a} C_{12} \mathbf{E}$$

vanishes if ratios of active and passive charges are the same for all bodies

we can distinguish two types of tests of neutrality:

- Tests of active neutrality, which measure the electric monopole field created by a passively neutral system, and
- tests of passive neutrality, which measure the force imposed by an external field onto an actively neutral system.

Experiments: Neutrality of atoms

Table: Various tests of the neutrality of atoms. If no particle is specified, q_p refers to the passive charge of the atoms or molecules used in the experiment, divided by the charge number of that particle (analogous for q_a). See Unnikrishnan & Gillies 2004 for a review.

Method	Limit $/(10^{-20}e)$
Gas efflux (350 g CO ₂) [Piccard & Kessler 1925]	$q_{p,a}/q_{e,a} = 0.1(5)$
Gas efflux (Ar/N) [Hillas & Cranshaw 1960]	$q_{H,a} = 1(3); q_{n,a} = -1(3)$
Gas efflux [King 1960]	$q_{He,a} = -4(2)$
Superfluid He [Classen et al 1998]	$q_{He,a} = -0.22(15)$
Levitorator [Marinelli & Morpurgo 1982]	$ q_p \lesssim 1000$
Acoustic resonator (SF ₆) [Dylla & King 1973]	$ q_p \leq 0.13$
Cs beam [Hughes 1957]	$q_p = 90(20)$
Neutron beam [Baumann et al 1988]	$q_{n,p} = -0.4(1.1)$

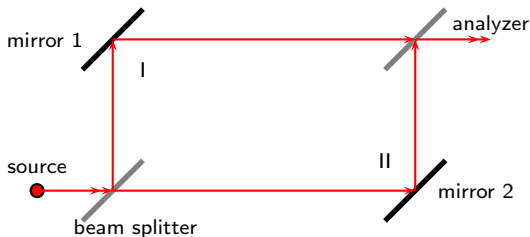
limits go down to $10^{-21} e$ for active and passive charge of various combinations of electrons, protons, and neutrons.

$$\Rightarrow |C_{21}| \leq 10^{-21}.$$

Atom interferometry

Arvanitaki et al 2008

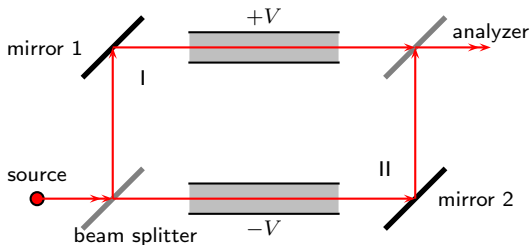
Place atoms in different voltages – requires large beam separation in configuration space



Atom interferometry

Arvanitaki et al 2008

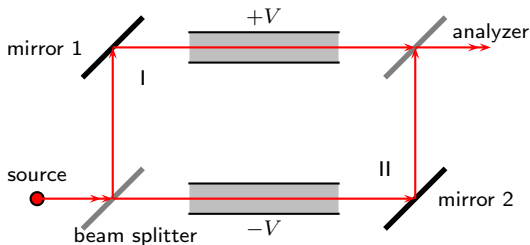
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Phase shift

$$\delta\phi = 2 \text{ charge} \int V dt$$

For Kasevich–setup in small tower: charge $\leq 10^{-30} e$
Improvement by 8 orders of magnitude

Alternative experiment: fine structure constant

- Center-of-mass motion of the two-particle system **cannot be quantized**
- Relative motion quantizable

Hamiltonian

$$H = \frac{\mathbf{p}^2}{2m_{\text{red}}} + D_{21} \frac{q_{1\text{p}}q_{2\text{p}}}{|\mathbf{x}|}.$$

Energy levels are proportional to modified fine structure constant

$$\alpha_{12} = \frac{q_{1\text{p}}q_{2\text{p}}D_{12}}{\hbar c} = \frac{1}{\hbar c} q_{1\text{p}}q_{2\text{p}} \left(\frac{q_{1\text{a}}}{q_{1\text{p}}} + \frac{m_{2\text{i}}}{M_{\text{i}}} C_{21} \right).$$

Spacing between energy levels depends nonlinearly on active charges \Rightarrow comparison of energy levels in different atoms yields test of C_{21} .

E.g., Hydrogen (one proton $q_1 = q_p$ and one electron $q_2 = q_e$) and ionized Helium He^+ ($q_1 = 2q_p$ and $q_2 = q_e$). Then

$$\alpha_{12}(\text{He}^+) - 2\alpha_{12}(\text{H}) \approx -\frac{q_{pp}q_{ep}}{\hbar c} \frac{m_{ei}}{m_{pi}} C_{21}$$

can deduce limit

$$|C_{21}| \leq \frac{\delta\alpha}{\alpha} \frac{m_{ei}}{m_{pi}} \approx 7 \times 10^{-13} \frac{m_{ei}}{m_{pi}} \approx 4 \times 10^{-16}$$

Summary Newton's axioms

- Fundamental postulates have to be tested as good as possible
- Finsler: Hints from quantum gravity
- Order of equation of motion: influence from space–time fluctuations
- Small accelerations: Hints from unexplained observations (dark matter)
- No model until now for active \neq passive charge;
it is a symmetry of physics which unfortunately is not yet well analyzed
- Systematic experimental study of Newton's axioms seems worthwhile
- Atom interferometry may be of help

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The Universality of Free Fall with quantum matter

Overview

Violations of UFF from

- Scalar tensor theories (Damour, Polyakov, Piazza, Veneziano).
Also influences Universality of Gravitational Redshift,
PPN parameters γ, β, \dots
estimate from low energy string theory $\eta \leq 10^{-13}$, $|\gamma - 1| \leq 10^{-5}$, ...
- Scalar tensor theory coupled to quintessence (Wetterich)
Prediction $\eta \approx 10^{-14}$.
- String theory: scattering of particles at branes = gravity. Estimate $\eta \leq 10^{-18}$
- Varying e model (Bekenstein)
Estimate $\eta \leq 10^{-13}$
- Particle moving through space-time fluctuations (Göklü & C.L.)
According to model and strength $\eta \leq 10^{-9}$,

Space-time fluctuations

The model

- General belief for Quantum Gravity: space-time fluctuates
- Simplest model of space-time fluctuations $g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$
Amelino-Camelia PRD 2000, Schiller et al 2004
- Simplest matter system: Klein-Gordon equation in this fluctuating space-time metric $g^{\mu\nu} D_\mu D_\nu \psi - m^2 \psi = 0$

- Relativistic approximation (of metric and of quantum field) + time-dependent transformation $\psi \rightarrow \psi' = A\psi$: hermitian Hamiltonian

$$\begin{aligned}
 H'\psi' &= -({}^{(3)}g)^{1/4} \frac{\hbar^2}{2m} \Delta_{\text{cov}} \left(({}^{(3)}g)^{-1/4} \psi' \right) + \frac{m}{2} \left(\tilde{h}_{(0)}^{00} - h_{(0)}^{00} \right) \psi' \\
 &\quad - \frac{1}{2} \left\{ i\hbar \partial_i, h_{(1)}^{i0} - \tilde{h}_{(1)}^{i0} \right\} \psi'
 \end{aligned}$$

- **Spatial average** over **Compton length of particle under consideration**

$$i\hbar \partial_t \psi = -\frac{\hbar^2}{2m} (\delta^{ij} + \alpha^{ij}(t)) \partial_i \partial_j \psi - mU\psi$$

Space-time fluctuations

$$\alpha^{ij}(t) = \tilde{\alpha}^{ij} + \gamma^{ij}(t), \quad \langle \gamma^{ij}(t) \rangle_t = 0$$

- $\tilde{\alpha}^{ij} \leftrightarrow$ spectral noise density of fluctuations
- Particular model: $m_i = \frac{m}{1 + \tilde{\alpha}}$, $\tilde{\alpha} \sim \left(\frac{l_{\text{Planck}}}{l_{\text{Compton}}} \right)^\beta$

Result

\Rightarrow anomalous inertial mass \rightarrow **apparent** violation of UFF

- **Alternative** route for violation of UFF and LLI.
- $\beta = \frac{1}{2}$ holographic noise (Ng 2001, Hogan 2008, GEO600)
- applies also to stochastic gravitational waves

Example

For Cesium and Hydrogen: $\eta_{\beta=1} = 10^{-17}$, $\eta_{\beta=2/3} = 10^{-12}$, $\eta_{\beta=1/2} = 10^{-9}$
 $\beta = \frac{1}{2}$ already ruled out (?)

Göklü & C.L. CQG 2008

Space-time fluctuations and Dirac equation

Quantum particle with spin: Dirac equation

Fluctuating metric couples to Dirac equation

$$i\gamma^\mu D_\mu \psi - m\psi = 0, \quad \gamma^{(\mu} \gamma^{\nu)} = g^{\mu\nu}$$

Nonrelativistic limit \Rightarrow expected terms

$$i\hbar\partial_t\psi = \frac{1}{2m} (\delta^{ij} + \alpha^{ij} + \beta^{ij}{}_k \sigma^k) \partial_i \partial_j \psi + m(1 + \gamma_i \sigma^i) U\psi + \delta_i \sigma^i \psi$$

Question: Why to keep Clifford algebra? \leftrightarrow fluctuations may be adapted to the structure of the field equation under consideration (fluctuation of coefficients – space-time is what particles explore).

Klein-Gordon equation: fluctuating $g^{\mu\nu}$

Dirac equation: fluctuating γ^μ : $\rightarrow \gamma^{(\mu} \gamma^{\nu)} = g^{\mu\nu} + \delta g^{\mu\nu} + X^{\mu\nu}$

Maxwell equation: fluctuating $g^{\mu[\rho} g^{\sigma]\nu}$ $\rightarrow k^{\mu\nu\rho\sigma} = g^{\mu[\rho} g^{\sigma]\nu} + \kappa^{\mu\nu\rho\sigma}$



Outline

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 - Newton's first law: Inertial systems
 - Newton's second law: The law of inertia
 - Newton's third law: Law of reciprocal action
- 2 **The Universality of Free Fall with quantum matter**
 - Violation of UFF from space-time fluctuations
 - **Universality of Free Fall for charged particles**
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UFF and charge

Standard theory

- In standard theory from ordinary coupling (deWitt & Brehme 1968)
 $a^\mu = \alpha \lambda_C c R^\mu{}_\nu v^\nu \Rightarrow$ violation of UFF $\sim 10^{-35} \text{ m/s}^2$

Anomalous coupling

- Anomalous coupling (Dittus, C.L., Selig, GRG 2004)

$$H = \frac{\mathbf{p}^2}{2m} + mU(\mathbf{x}) + \kappa e U(\mathbf{x}) = \frac{\mathbf{p}^2}{2m} + m \left(1 + \kappa \frac{e}{m} \right) U(\mathbf{x}).$$

- Charge dependent anomalous gravitational mass tensor
 - Also charge dependent anomalous inertial mass tensor (e.g. Rohrlich 2000)
- \Rightarrow Charge dependent Eötvös factor
- It is possible to choose κ 's such that for neutral composite matter UFF is fulfilled while for **isolated charges** UFF is violated \Rightarrow no constraints on κ from present UFF experiments

No underlying fundamental theory known

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UFF and spin

Standard theory

- In standard theory from ordinary coupling: $a^\mu = \lambda_C R^\mu{}_{\nu\rho\sigma} v^\nu S^{\rho\sigma} \Rightarrow$ violation of UFF at the order 10^{-20} m/s^2 , beyond experiment

Anomalous coupling

- Speculations: violation P , C , and T symmetry in gravitational fields (Leitner & Okubo 1964, Moody & Wilczek 1974) suggest

$$V(r) = U(r) [1 + A_1(\boldsymbol{\sigma}_1 \pm \boldsymbol{\sigma}_2) \cdot \hat{\mathbf{r}} + A_2(\boldsymbol{\sigma}_1 \times \boldsymbol{\sigma}_2) \cdot \hat{\mathbf{r}}] ,$$

- One body (e.g., the Earth) is unpolarized \rightarrow

$$V(r) = U(r) (1 + A\boldsymbol{\sigma} \cdot \hat{\mathbf{r}}) .$$

Hyperfine splittings of H ground state: $A_p \leq 10^{-11}$, $A_e \leq 10^{-7}$

- Hari Dass 1976, 1977, includes velocity of the particles

$$V(r) = U_0(r) \left[1 + A_1 \boldsymbol{\sigma} \cdot \hat{\mathbf{r}} + A_2 \boldsymbol{\sigma} \cdot \frac{\mathbf{v}}{c} + A_3 \hat{\mathbf{r}} \cdot \left(\boldsymbol{\sigma} \times \frac{\mathbf{v}}{c} \right) \right]$$

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Decoherence

- Fluctuations

$$\alpha^{ij}(t) = \underbrace{\tilde{\alpha}^{ij}}_{\text{renormalization of } m_i} + \underbrace{\gamma^{ij}(t)}_{\text{decoherence}}, \quad \langle \gamma^{ij}(t) \rangle_t = 0$$

- Quantum master equation
- White noise $\gamma^{ij} = \sigma \delta^{ij} \xi(t)$, $\tau_c = \sigma^2$: Markovian master equation
 - preserves trace and positivity of the density matrix
 - generates complete positive dynamical map
 - quantum dynamical semigroup

- Increase of entropy
- Exponential decay of coherences in energy representation
- Coherence time

$$\tau_D = 2 \left(\frac{\hbar}{\tau_c \Delta E} \right) \tau_c, \quad \tau_c = T_{\text{PI}}$$

applies also to stochastic gravitational waves

(Breuer, Göklü & C.L. 2008 – also Wang, Bingham & Mendoca 2006, 2008)

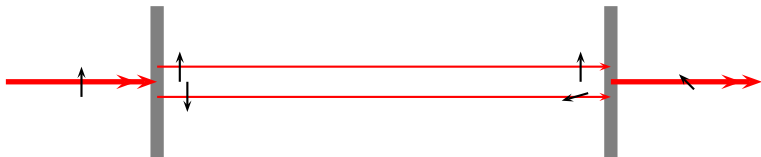


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Interference with spin

Interference in spin space



phase shift (Audretsch, Bleyer & CL, PRD 1993; CL, CQG 1998):

$$\begin{aligned} \delta\phi &= (H(\mathbf{x}, \mathbf{p}, \mathbf{S}) - H(\mathbf{x}, \mathbf{p}, -\mathbf{S})) \delta t \\ &= \left(\frac{\delta m_{ik}^{ij}}{m} p_i \delta_{jk} l^k - 2(\lambda_j^i + m\Lambda_j^i) \delta_{ik} l^k + 2mB_j \delta t + 2C_j mU \delta t + T_j \delta t \right) S^j \end{aligned}$$

gives estimates on anomalous Lorentz-violating spin-coupling terms

$$\begin{aligned} \left| \frac{\delta m_k^{ij}}{m} \right| &\leq 5 \cdot 10^{-15}, & |\Lambda_j^i| &\leq 8 \cdot 10^{-24}, & |\lambda_j^i| &\leq 5 \cdot 10^{-6} \text{m}^{-1}, & |B_k| &\leq 3 \cdot 10^{-30} \\ |T_k| &\leq 2 \cdot 10^{-9} \text{m}^{-1}, & |C_k| &\leq 3 \cdot 10^{-23} \end{aligned}$$

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Space-time fluctuations and spin

Classical particle with spin I

Equation of motion for pole-dipole particle (v is auxiliary quantity)

$$\begin{aligned} v &= \frac{dx}{ds} \\ D_v p &= R(\cdot, v, S) \\ D_v S &= v \wedge p \\ p(S) &= 0 \end{aligned}$$

Initial conditions $x(t_0), p(t_0), S(t_0)$

Classical article with spin II

From quasiclassical limit of Dirac

$$\begin{aligned} D_v v &= \lambda_C R(\cdot, v, S) \\ D_v S &= 0 \\ p(S) &= 0 \quad \text{automatically} \end{aligned}$$

Since for Dirac $p \sim v$ no auxiliary velocity needed

- Spin particle sees curvature directly!
- For fluctuations: $D_v v = R(\cdot, v, S) + \delta R(\cdot, v, S)$
- fluctuations will give additional acceleration term which may be arger than the standard term \Rightarrow anomalous spin coupling \Rightarrow violation of UFF for spin
- $\delta R(\cdot, v, S) \sim (\partial h)^2 v S$ – enhancement for short wavelength fluctuations

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Superposition principle

General non-linear Schrödinger equation

$$i\frac{\partial\psi}{\partial t} = -\frac{1}{2m}\Delta\psi + F(\psi^*\psi)\psi$$

Separability of quantum systems: non-linear Schrödinger equation of Bialnicky–Birula PRL 1977; Shimony, PRA 1978

$$i\frac{\partial\psi}{\partial t} = -\frac{1}{2m}\Delta\psi + a[\ln(b\psi^*\psi)]\psi$$



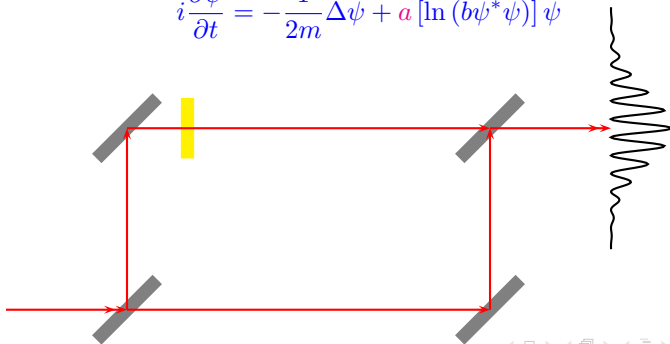
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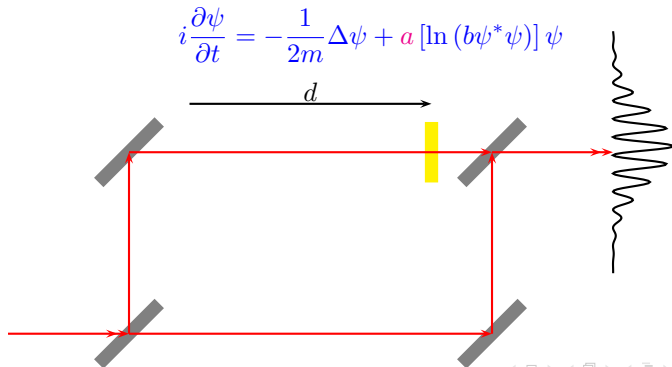


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Superposition principle

phase shift:

$$\delta\phi = \frac{d}{\hbar} \sqrt{\frac{m}{2E}} (F(|\varphi|^2) - F(\alpha^2|\varphi|^2))$$

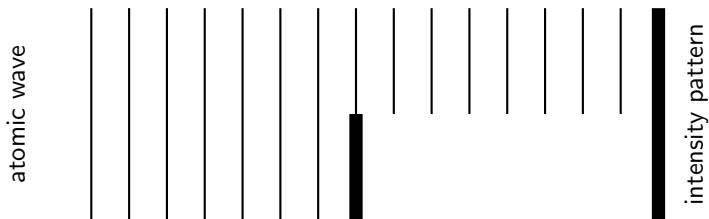
α = intensity attenuation.

Test with neutron interferometry (Shull et al, PRL 1980)

- result: $a \leq 3.4 \cdot 10^{-13}$ eV
- atomic interferometry should lead to orders of magnitude improvement
- from energy levels: $a \leq 4 \cdot 10^{-10}$ eV

Superposition principle

Alternative measurement: Scattering at edges



- yields best estimates for neutrons: $\alpha \leq 3 \cdot 10^{-15} \text{ eV}$
- depends on velocity of particles \rightarrow should be better by many orders of magnitude for atoms
- Van der Waals, Casimir forces etc. should be included in calculation, and parameters determined by independent experiments, or by scattering at edges made of different materials

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Anomalous dispersion – non-local field equations

The model

Ansatz in 3+1-form

$$i\partial_t\psi = -i\alpha^i\partial_i\psi + \alpha^{ij}\partial_i\partial_j\psi + \beta m\psi$$

$\alpha^{ij} \neq 0 \implies$ violation of LLI (in a certain sense)

Non-relativistic limit

$$i\partial_t\psi = -\frac{1}{2m}\Delta\psi + \frac{1}{m^2}\left(A^{ijk} + B_l^{ijk}\sigma^l\right)\partial_i\partial_j\partial_k\psi$$

- interference experiments: Spin-flip-experiment:

$$\Delta\phi = \frac{1}{\hbar}(H(S) - H(-S))\Delta t = \frac{2}{\hbar c}B_m^{jki}S^m m v_i v_j l_k$$

spin- $\frac{1}{2}$ -field, $m = 2 \times 10^{-23}$ g, $l = 10$ m, $v = 1000$ m/sec \rightarrow

$$B_m^{(ijk)} \leq \frac{\hbar c}{2} \frac{1}{S m v^2 l} \approx 10^{-10}$$

Anomalous dispersion

Is a generic effect of all QG approaches (string theory, LQG, NCG)

Non-locality \leftrightarrow higher order derivatives

$$i\hbar \frac{\partial}{\partial t} \psi = H(p) \psi = \sum_{i=0}^n \alpha_i p^i$$

may also come from field equations with higher order time derivative (C.L. & Bordé 2001)

Gives dispersion relation

$$E = H(p)$$

Atom-photon interaction yields phase shift

$$\delta\varphi \sim H(p + \hbar k) - H(p)$$

Under consideration with G. Amelino-Camelia, G. Tino, ...

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Summary

- Worth to do better tests of Newton's axioms
 - Existence of inertial systems
 - Order of equation of motion
 - Active vs. passive charge/mass/spin
- Many influences of space-time fluctuations (stochastic gravitational waves)
 - Violation of UFF and LLI
 - Decoherence
 - Order of equations of motion

Summary

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Thank you!

Thanks to

- DLR
- DFG
- Center of Excellence QUEST

Proceedings

- Publication as special issue in the journal *General Relativity and Gravitation*
- Deadline May 29, 2009
- Papers will be refereed
- All formats are welcome (LaTeX, Word)

Space-time fluctuations and spin

- Geodesic equation for static spherically symmetric metric

$$\left(\frac{dr}{ds}\right)^2 = \frac{1}{g_{tt}g_{rr}} \left(E^2 - g_{tt} \left(\epsilon + m \frac{L^2}{r^2} \right) \right)$$

- Assuming $g_{rr} = 1/g_{tt}$ and $g_{tt} = 1 + h \cos(kr)$ (everywhere C^∞)
Then for radial motion ($L = 0$)

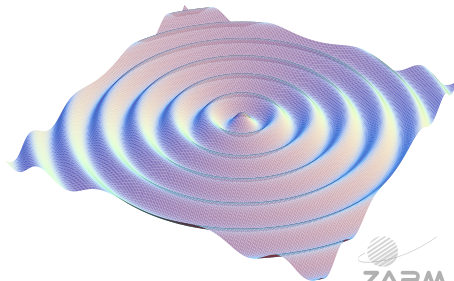
$$\left(\frac{dr}{ds}\right)^2 = E^2 - 1 - h \cos(kr)$$

Can be solved by elliptic function

- for $h \ll 1$

$$r = \sqrt{E^2 - 1} s + h \frac{\sin(\sqrt{E^2 - 1} ks)}{2(E^2 - 1)k}$$

like zitterbewegung



Space-time fluctuations and spin

Kretschmann scalar (indicator of space-time singularities)

$$K = R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma} \sim \frac{g''_{tt}}{g_{tt}} \sim hk^2 \frac{\cos(kr)}{1 + h \cos(kr)}$$

\Rightarrow force term $R^\mu{}_{\nu\rho\sigma}u^\nu S^{\rho\sigma}$ may become large \Rightarrow violation of UFF

Consequence

- For $k \rightarrow \infty$: solution of geodesic equation approaches straight line
- Kretschmann scalar $\rightarrow \infty$
- \Rightarrow point particles do not see fluctuating curvature
 particles with spin should be sensitive to curvature
 - \rightarrow may be of importance in atomic interferometry, spectroscopy, ... if fluctuations are regarded as being due to quantum gravity
 - \rightarrow estimates from experiments
- needs to be compared with analysis of Dirac equation in fluctuating space-time metric

Göklü & C.L. in preparation