Particle production in AA collisions in the Color Glass Condensate framework

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- IR & Coll. divergences
- Factorization
- Higher twist
- Goals

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Infrared and collinear divergences

Calculation of some process at LO :

 $\left\{ \begin{array}{c} x_1 = M_{\perp} \ e^{+Y}/\sqrt{s} \\ x_2 = M_{\perp} \ e^{-Y}/\sqrt{s} \\ \end{array} \right\} (M_{\perp}, Y) \qquad \left\{ \begin{array}{c} x_1 = M_{\perp} \ e^{+Y}/\sqrt{s} \\ x_2 = M_{\perp} \ e^{-Y}/\sqrt{s} \end{array} \right\}$

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Infrared and collinear divergences

Calculation of some process at LO :



Radiation of an extra gluon :





ntrod	luction	

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Infrared and collinear divergences

- Large $\log(M_{\perp})$ when M_{\perp} is large
- Large $\log(1/x_1)$ when $x_1 \ll 1$

▷ these logs can compensate the additional α_s , and void the naive application of perturbation theory ▷ resummations are necessary

• Logs of $M_{\perp} \Longrightarrow$ DGLAP. Important when :

- $\bullet \ M_{\perp} \gg \Lambda_{_{QCD}}$
- x_1, x_2 are rather large
- Logs of $1/x \Longrightarrow$ BFKL. Important when :
 - M_{\perp} remains moderate
 - x_1 or x_2 (or both) are small
- Physical interpretation :
 - The physical process can resolve the gluon splitting if $M_{\perp} \gg k_{\perp}$
 - If $x_1 \ll 1$, the gluon that initiates the process is likely to result from bremsstrahlung from another parent gluon



Factorization

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- Logs of M_{\perp} can be resummed by :
 - promoting $f(x_1)$ to $f(x_1, M_{\perp}^2)$
 - letting $f(x_1, M_{\perp}^2)$ evolve with M_{\perp} according to the DGLAP equation

$$\frac{\partial f(x, M^2)}{\partial \ln(M^2)} = \alpha_s(M^2) \int_x^1 \frac{dz}{z} P(x/z) \otimes f(z, M^2)$$

▷ collinear factorization

- Logs of x_1 can be resummed by :
 - promoting $f(x_1)$ to a non integrated distribution $\varphi(x_1, \vec{k}_{\perp})$
 - letting $\varphi(x_1, \vec{k}_{\perp})$ evolve with x_1 according to the BFKL equation

$$\frac{\partial \varphi(x, k_{\perp})}{\partial \ln(1/x)} = \alpha_s \int \frac{d^2 \vec{p}_{\perp}}{(2\pi)^2} K(\vec{k}_{\perp}, \vec{p}_{\perp}) \otimes \varphi(x, \vec{p}_{\perp})$$

 \triangleright k_{\perp} -factorization



Leading twist :

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> 2-point function in the projectile > gluon number



Leading twist :

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> 2-point function in the projectile > gluon number

Higher twist contributions :



> 4-point function in the projectile > higher correlation
 > multiple scattering in the projectile



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Power counting : rescattering corrections are suppressed by inverse powers of the typical mass scale in the process :



- The parameter μ^2 has a factor of α_s , and a factor proportional to the gluon density \triangleright rescatterings are important at high density
- Relative order of magnitude :

$$\frac{{\rm twist}\; {\rm 4}}{{\rm twist}\; {\rm 2}}\sim \frac{Q_s^2}{M_\perp^2} \quad {\rm with} \quad Q_s^2\sim \alpha_s \frac{xG(x,Q_s^2)}{\pi R^2}$$

- When this ratio becomes ~ 1 , all the rescattering corrections become important
- These effects are not accounted for in DGLAP or BFKL



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- 99% of the multiplicity below $p_{\perp} \sim 2 \text{ GeV}$
- Q_s^2 might be as large as 5 GeV² at the LHC ($\sqrt{s} = 5.5$ TeV) \triangleright rescatterings are important, and one should also resum logs of 1/x



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- The Color Glass Condensate framework provides the technology for resumming all the $[Q_s/p_{\perp}]^n$ corrections
- Generalize the concept of "parton distribution"
 - Due to the high density of partons, observables depend on higher correlations (beyond the usual parton distributions, which are 2-point correlation functions)
- If logs of 1/x show up in loop corrections, one should be able to factor them out into the evolution of these distributions
- These distributions should be universal, with non-perturbative information relegated into the initial condition for the evolution
- There may possibly be extra divergences associated with the evolution of the final state



Initial Conditions





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- Basic principles and bookkeeping
- Inclusive gluon spectrum at leading order
- Loop corrections, factorization, instabilities
- Less inclusive quantities
 - FG, Venugopalan, hep-ph/0601209, 0605246
 - Fukushima, FG, McLerran, hep-ph/0610416
 - + work in progress with Lappi, Venugopalan



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Degrees of freedom and their interplay

McLerran, Venugopalan (1994), Iancu, Leonidov, McLerran (2001)

Soft modes have a large occupation number
 b they are described by a classical color field A^µ that obeys Yang-Mills's equation:

$$[D_{\nu}, \boldsymbol{F}^{\boldsymbol{\nu\mu}}]_a = J_a^{\mu}$$

The source term J^μ_a comes from the faster partons. The hard modes, slowed down by time dilation, are described as frozen color sources ρ_a. Hence :

$$J_a^{\mu} = \delta^{\mu +} \delta(x^-) \rho_a(\vec{x}_{\perp}) \qquad (x^- \equiv (t-z)/\sqrt{2})$$

The color sources ρ_a are random, and described by a distribution functional W_Y[ρ], with Y the rapidity that separates "soft" and "hard". Evolution equation (JIMWLK) :

$$\frac{\partial W_{Y}[\rho]}{\partial Y} = \mathcal{H}[\rho] \ W_{Y}[\rho]$$



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Compute the observable O of interest for a configuration of the sources ρ_1 , ρ_2 . Note : the sources are $\sim 1/g$ \triangleright weak coupling but strong interactions

Description of hadronic collisions

At LO, this requires to solve the classical Yang-Mills equations in the presence of the following current :

 $J^{\mu} \equiv \delta^{\mu +} \delta(x^{-}) \,\rho_1(\vec{x}_{\perp}) + \delta^{\mu -} \delta(x^{+}) \,\rho_2(\vec{x}_{\perp})$

(Note: the boundary condition depend on the observable)

• Average over the sources ρ_1 , ρ_2

 $\langle \mathcal{O}_{Y} \rangle = \int \left[D\rho_{1} \right] \left[D\rho_{2} \right] W_{Y_{\text{beam}}-Y}[\rho_{1}] W_{Y+Y_{\text{beam}}}[\rho_{2}] \mathcal{O}[\rho_{1},\rho_{2}]$

Can this procedure – and in particular the above factorization formula – be justified ?



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10 configurations



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100 configurations



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1000 configurations







Dilute regime : one source in each projectile interact







Dilute regime : one source in each projectile interact

Dense regime : non linearities are important







- Dilute regime : one source in each projectile interact
- Dense regime : non linearities are important
- Many gluons can be produced from the same diagram





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- Dilute regime : one source in each projectile interact
- Dense regime : non linearities are important
- Many gluons can be produced from the same diagram
- There can be many simultaneous disconnected diagrams





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- Dilute regime : one source in each projectile interact
- Dense regime : non linearities are important
- Many gluons can be produced from the same diagram
- There can be many simultaneous disconnected diagrams
- Some of them may not produce anything (vacuum diagrams)





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- Dilute regime : one source in each projectile interact
- Dense regime : non linearities are important
- Many gluons can be produced from the same diagram
- There can be many simultaneous disconnected diagrams
- Some of them may not produce anything (vacuum diagrams)
- All these diagrams can have loops (not at LO though)



Power counting





- In the saturated regime, the sources are of order 1/g
- The order of each disconnected diagram is given by :

$$\frac{1}{g^2} g^{\# \text{ produced gluons}} g^{2(\# \text{ loops})}$$

The total order of a graph is the product of the orders of its disconnected subdiagrams > quite messy...



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Consider squared amplitudes (including interference terms) rather than the amplitudes themselves



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- Consider squared amplitudes (including interference terms) rather than the amplitudes themselves
- See them as cuts through vacuum diagrams



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- Consider squared amplitudes (including interference terms) rather than the amplitudes themselves
- See them as cuts through vacuum diagrams
- Consider only the simply connected ones, thanks to :

$$\sum \begin{pmatrix} \text{all the vacuum} \\ \text{diagrams} \end{pmatrix} = \exp \left\{ \sum \begin{pmatrix} \text{simply connected} \\ \text{vacuum diagrams} \end{pmatrix} \right\}$$

Simpler power counting for connected vacuum diagrams :

$$\frac{1}{g^2} g^{2(\# \text{ loops})}$$



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There is an operator D that acts on a pair of vacuum diagrams by removing two sources and attaching a cut propagator instead :




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 $\blacksquare \mathcal{D}$ can also act directly on single diagram, if it is already cut



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 $\blacksquare \mathcal{D}$ can also act directly on single diagram, if it is already cut



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There is an operator D that acts on a pair of vacuum diagrams by removing two sources and attaching a cut propagator instead :



- $\blacksquare \mathcal{D}$ can also act directly on single diagram, if it is already cut
- By repeated action of D, one generates all the diagrams with an arbitrary number of cuts
- Thanks to this operator, one can write :

$$P_{n} = \frac{1}{n!} \mathcal{D}^{n} e^{iV} e^{-iV^{*}} , \quad iV = \sum \begin{pmatrix} \text{connected uncut} \\ \text{vacuum diagrams} \end{pmatrix}$$
$$\sum \begin{pmatrix} \text{all the cut} \\ \text{vacuum diagrams} \end{pmatrix} = e^{\mathcal{D}} e^{iV} e^{-iV^{*}}$$



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$$\overline{N} = \sum_{n} n P_{n} = \mathcal{D} \left\{ e^{\mathcal{D}} e^{iV} e^{-iV^{*}} \right\}$$

- \overline{N} is obtained by the action of \mathcal{D} on the sum of all the cut vacuum diagrams. There are two kind of terms :
 - \mathcal{D} picks two sources in two distinct connected cut diagrams



• \mathcal{D} picks two sources in the same connected cut diagram





At LO, only tree diagrams contribute topologies can be neglected (it starts at 1-loop)

In each blob, we must sum over all the tree diagrams, and over all the possible cuts :



A major simplification comes from the following property :

------ + ----**x**--- retarded propagator

The sum of all the tree diagrams constructed with retarded propagators is the retarded solution of Yang-Mills equations :

$$[D_{\mu}, F^{\mu\nu}] = J^{\nu}$$
 with $A^{\mu}(x_0 = -\infty) = 0$

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$$\frac{dN_{LO}}{dYd^2\vec{p}_{\perp}} = \frac{1}{16\pi^3} \int_{x,y} e^{ip \cdot (x-y)} \Box_x \Box_y \sum_{\lambda} \epsilon^{\mu}_{\lambda} \epsilon^{\nu}_{\lambda} \mathcal{A}_{\mu}(x) \mathcal{A}_{\nu}(y)$$

• $\mathcal{A}^{\mu}(x) =$ retarded solution of Yang-Mills equations





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$$\frac{dN_{LO}}{dYd^2\vec{p}_{\perp}} = \frac{1}{16\pi^3} \int_{x,y} e^{ip \cdot (x-y)} \Box_x \Box_y \sum_{\lambda} \epsilon^{\mu}_{\lambda} \epsilon^{\nu}_{\lambda} \mathcal{A}_{\mu}(x) \mathcal{A}_{\nu}(y)$$

• $\mathcal{A}^{\mu}(x) =$ retarded solution of Yang-Mills equations \triangleright can be cast into an initial value problem on the light-cone







- KNV I KNV II Lappi 5 2 3 4 1 6 k_T/Λ_s
- Lattice artefacts at large momentum (they do not affect much the overall number of gluons)
- Important softening at small k_{\perp} compared to pQCD (saturation)



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$$\begin{array}{rcl} A^{i}(x) & = & \alpha^{i}(\tau,\eta,\vec{x}_{\perp}) \\ A^{\pm}(x) & = & \pm x^{\pm} \ \beta(\tau,\eta,\vec{x}_{\perp}) \end{array}$$



Initial values at $\tau = 0^+$: $\alpha^i(0^+, \eta, \vec{x}_\perp)$ and $\beta(0^+, \eta, \vec{x}_\perp)$ do not depend on the rapidity η

 $ightarrow \alpha^i$ and β remain independent of η at all times (invariance under boosts in the *z* direction)

 \triangleright numerical resolution performed in 1+2 dimensions



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• 1-loop diagrams for \overline{N}



This can be seen as a perturbation of the initial value problem encountered at LO, e.g. :



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This can be seen as a perturbation of the initial value problem encountered at LO, e.g. :



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The 1-loop correction to \overline{N} can be written as a perturbation of the initial value problem encountered at LO :





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The 1-loop correction to \overline{N} can be written as a perturbation of the initial value problem encountered at LO :



- \overline{N}_{LO} is a functional of the initial fields $\mathcal{A}_{in}(\vec{u})$ on the light-cone
- $T_{\vec{u}}$ is the generator of shifts of the initial condition at the point \vec{u} on the light-cone, i.e. : $T_{\vec{u}} \sim \delta/\delta A_{in}(\vec{u})$



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The 1-loop correction to \overline{N} can be written as a perturbation of the initial value problem encountered at LO :



- \overline{N}_{LO} is a functional of the initial fields $\mathcal{A}_{in}(\vec{u})$ on the light-cone
- $T_{\vec{u}}$ is the generator of shifts of the initial condition at the point \vec{u} on the light-cone, i.e. : $T_{\vec{u}} \sim \delta/\delta A_{in}(\vec{u})$
- $\delta A_{in}(\vec{u})$ and $\Sigma(\vec{u}, \vec{v})$ are in principle calculable analytically



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The first two terms involve :

$$\delta \mathcal{A}(x) \equiv \frac{g}{2} \int d^4 z \sum_{\epsilon=\pm} \epsilon \ \mathbf{G}_{+\epsilon}(x,z) \mathbf{G}_{\epsilon\epsilon}(z,z)$$

- The third term involves $G_{+-}(x, y)$
- The propagators $G_{\pm\pm}$ are propagators in the background \mathcal{A} , in the Schwinger-Keldysh formalism. They obey :

$$\begin{cases} \boldsymbol{G}_{+-} = \boldsymbol{G}_{R} G_{R}^{0 - 1} G_{A}^{0 - 1} \boldsymbol{G}_{A} \\ \boldsymbol{G}_{\pm \pm} = \frac{1}{2} \left[\boldsymbol{G}_{R} G_{R}^{0 - 1} (G_{+-}^{0} + G_{-+}^{0}) G_{A}^{0 - 1} \boldsymbol{G}_{A} \pm (\boldsymbol{G}_{R} + \boldsymbol{G}_{A}) \right] \end{cases}$$

 $G_{R,A}$ = retarded/advanced propagators in the background A



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G₊₊ and G₋₋ are only needed with equal endpoints
 b they are both equal to

$$\boldsymbol{G}_{++}(z,z) = \boldsymbol{G}_{--}(z,z) = \frac{1}{2} \left[\boldsymbol{G}_{R} G_{R}^{0 - 1} (G_{+-}^{0} + G_{-+}^{0}) G_{A}^{0 - 1} \boldsymbol{G}_{A} \right] (z,z)$$

 \triangleright thus, δA can be simplified into :

$$\begin{split} \delta \mathcal{A}(x) &= \frac{g}{2} \int d^4 z \left[\mathbf{G}_{++}(x,z) - \mathbf{G}_{+-}(x,z) \right] \mathbf{G}_{++}(z,z) \\ &= \frac{g}{2} \int d^4 z \; \mathbf{G}_R(x,z) \mathbf{G}_{++}(z,z) \end{split}$$

 $\blacksquare G_R G_R^{0 - 1} G_{+-}^0 G_A^{0 - 1} G_A$ can be written as :

$$\left[\boldsymbol{G}_{R} G_{R}^{0 - 1} G_{A - 1}^{0} G_{A}^{0 - 1} \boldsymbol{G}_{A}\right](x, y) = \int \frac{d^{3} \vec{\boldsymbol{p}}}{(2\pi)^{3} 2E_{\boldsymbol{p}}} \, \boldsymbol{\zeta}_{\vec{\boldsymbol{p}}}(x) \boldsymbol{\zeta}_{\vec{\boldsymbol{p}}}^{*}(y) \,,$$

with
$$\left[\Box_x + m^2 + g\mathcal{A}(x)\right]\zeta_{\vec{p}}(x) = 0$$
 and $\lim_{x_0 \to -\infty} \zeta_{\vec{p}}(x) = e^{ip \cdot x}$



Sketch of a proof – III

Green's formulas :

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$\mathcal{A}(x) = \int_{\Omega} d^4 z \ G^0_R(x,z) \Big[j(z) - \frac{g}{2} \mathcal{A}^2(z) \Big] \\ + \int_{\mathrm{LC}} d^3 \vec{u} \ G^0_R(x,u) \Big[n \cdot \overrightarrow{\partial}_u - n \cdot \overleftarrow{\partial}_u \Big] \mathcal{A}_{\mathrm{in}}(\vec{u})$

$$\begin{split} \delta \mathcal{A}(x) &= \int_{\Omega} d^4 z \; \boldsymbol{G}_R(x,z) \; \frac{g}{2} \boldsymbol{G}_{++}(z,z) \\ &+ \int_{\mathrm{LC}} d^3 \vec{\boldsymbol{u}} \; \boldsymbol{G}_R(x,u) \Big[n \cdot \overrightarrow{\partial}_u \; -n \cdot \overleftarrow{\partial}_u \; \Big] \delta \mathcal{A}_{\mathrm{in}}(\vec{\boldsymbol{u}}) \end{split}$$

$$\zeta_{\vec{p}}(x) = \int_{\mathrm{LC}} d^{3}\vec{u} \, \boldsymbol{G}_{R}(x,u) \Big[n \cdot \overrightarrow{\partial}_{u} - n \cdot \overleftarrow{\partial}_{u} \Big] \zeta_{\vec{p}} \operatorname{in}(\vec{u})$$

$$\boldsymbol{G}_{R}(x,y) = G_{R}^{0}(x,y) + g \int_{\Omega} d^{4}z \ G_{R}^{0}(x,z) \mathcal{A}(z) \boldsymbol{G}_{R}(z,y)$$



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Thanks to the operator

$$a_{
m in}(ec{m{u}})\cdot m{T}_{ec{m{u}}}\equiv a_{
m in}(ec{m{u}})rac{\delta}{\delta \mathcal{A}_{
m in}(ec{m{u}})} + \Big[(n\cdot\partial_u)a_{
m in}(ec{m{u}})\Big]rac{\delta}{\delta(n\cdot\partial_u)\mathcal{A}_{
m in}(ec{m{u}})}~,$$

we can write

$$\begin{split} \zeta_{\vec{p}}(x) &= \int_{\vec{u} \in \mathrm{LC}} \left[\zeta_{\vec{p} \mathrm{in}}(\vec{u}) \cdot \boldsymbol{T}_{\vec{u}} \right] \mathcal{A}(x) \\ \delta \mathcal{A}(x) &= \int_{\Omega} d^4 z \; \boldsymbol{G}_R(x,z) \; \frac{g}{2} \boldsymbol{G}_{++}(z,z) + \int_{\vec{u} \in \mathrm{LC}} \left[\delta \mathcal{A}_{\mathrm{in}}(\vec{u}) \cdot \boldsymbol{T}_{\vec{u}} \right] \mathcal{A}(x) \end{split}$$

 \triangleright from the classical field $\mathcal{A}(x)$, the operator $a_{in}(\vec{u}) \cdot T_{\vec{u}}$ builds the fluctuation a(x) whose initial condition on the light-cone is $a_{in}(\vec{u})$

■ The 3rd diagram can directly be written as :

$$\int \frac{d^3 \vec{p}}{(2\pi)^3 2E_p} \int_{\vec{u},\vec{v} \in \text{LC}} \left[\left[\zeta_{\vec{p} \text{ in}}(\vec{u}) \cdot T_{\vec{u}} \right] \mathcal{A}(x) \right] \left[\left[\zeta_{\vec{p} \text{ in}}^*(\vec{v}) \cdot T_{\vec{v}} \right] \mathcal{A}(y) \right]$$



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Sketch of a proof – V

One can finally prove that

.

$$\int_{\Omega} d^4 z \, \boldsymbol{G}_R(x, z) \, \frac{g}{2} \boldsymbol{G}_{++}(z, z) =$$

$$= \frac{1}{2} \int \frac{d^3 \vec{\boldsymbol{p}}}{(2\pi)^3 2E_{\boldsymbol{p}}} \int_{\vec{\boldsymbol{u}}, \vec{\boldsymbol{v}} \in \mathrm{LC}} \left[\zeta_{\vec{\boldsymbol{p}}\,\mathrm{in}}(\vec{\boldsymbol{u}}) \cdot \boldsymbol{T}_{\vec{\boldsymbol{u}}} \right] \left[\zeta_{\vec{\boldsymbol{p}}\,\mathrm{in}}^*(\vec{\boldsymbol{v}}) \cdot \boldsymbol{T}_{\vec{\boldsymbol{v}}} \right] \, \mathcal{A}(x)$$

$$\triangleright \quad \delta \mathcal{A}(x) = \left[\int_{\vec{u} \in \mathrm{LC}} \left[\delta \mathcal{A}_{\mathrm{in}}(\vec{u}) \cdot T_{\vec{u}} \right] + \frac{1}{2} \int \frac{d^3 \vec{p}}{(2\pi)^3 2E_p} \int_{\vec{u}, \vec{v} \in \mathrm{LC}} \left[\zeta_{\vec{p} \mathrm{in}}(\vec{u}) \cdot T_{\vec{u}} \right] \left[\zeta_{\vec{p} \mathrm{in}}^*(\vec{v}) \cdot T_{\vec{v}} \right] \right] \mathcal{A}(x)$$

• This leads to the announced formula for $\delta \overline{N}$, with

$$\boldsymbol{\Sigma}(\boldsymbol{\vec{u}},\boldsymbol{\vec{v}}) \equiv \int \frac{d^3 \boldsymbol{\vec{p}}}{(2\pi)^3 2E_{\boldsymbol{p}}} \zeta_{\boldsymbol{\vec{p}}\,\mathrm{in}}(\boldsymbol{\vec{u}}) \zeta_{\boldsymbol{\vec{p}}\,\mathrm{in}}^*(\boldsymbol{\vec{v}})$$



Sketch of a proof – VI

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Conjecture : this result can be generalized to any observable that can be written in terms of the gauge field with retarded boundary conditions, $\mathcal{O} \equiv \mathcal{O}[\mathcal{A}]$:

$$\delta \mathcal{O} = \left[\int_{\vec{u} \in \text{ light cone}} \delta \mathcal{A}_{\text{in}}(\vec{u}) T_{\vec{u}} + \int_{\vec{v} \in \text{ light cone}} \frac{1}{2} \Sigma(\vec{u}, \vec{v}) T_{\vec{u}} T_{\vec{v}} \right] \mathcal{O}_{LO}$$

b whatever we conclude for the multiplicity from this formula holds true for any such observable



Divergences

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If taken at face value, this 1-loop correction is plagued by several divergences :

- The two coefficients $\delta A_{in}(\vec{x})$ and $\Sigma(\vec{x}, \vec{y})$ are infinite, because of an unbounded integration over a rapidity variable
- At late times, $T_{\vec{x}} \mathcal{A}(\tau, \vec{y})$ diverges exponentially,

$$T_{\vec{x}}\mathcal{A}(\tau,\vec{y}) \underset{\tau \to +\infty}{\sim} e^{\sqrt{\mu\tau}}$$

because of an instability of the classical solution of Yang-Mills equations under rapidity dependent perturbations (Romatschke, Venugopalan (2005))



Initial state factorization

Anatomy of the full calculation :



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Anatomy of the full calculation :



• When the observable $\overline{N}[\mathcal{A}_{in}(\rho_1, \rho_2)]$ is corrected by an extra gluon, one gets divergences of the form $\alpha_s \int dY$ in $\delta \overline{N}$ \triangleright one would like to be able to absorb these divergences into the Y dependence of the source densities $W_{Y}[\rho_{1,2}]$



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- When the observable $\overline{N}[\mathcal{A}_{in}(\rho_1, \rho_2)]$ is corrected by an extra gluon, one gets divergences of the form $\alpha_s \int dY$ in $\delta \overline{N}$ \triangleright one would like to be able to absorb these divergences into the Y dependence of the source densities $W_{Y}[\rho_{1,2}]$
- Equivalently, if one puts some arbitrary frontier Y_0 between the "observable" and the "source distributions", the dependence on Y_0 should cancel between the various factors



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The two kind of divergences don't mix, because the divergent part of the coefficients is boost invariant.

Given their structure, the divergent coefficients seem related to the evolution of the sources in the initial state

In order to prove the factorization of these divergences in the initial state distributions of sources, one needs to establish :

$$\left[\delta \overline{N}\right]_{\text{divergent}\atop\text{coefficients}} = \left[\left(Y_0 - Y\right) \mathcal{H}^{\dagger}[\rho_1] + \left(Y - Y_0^{\prime}\right) \mathcal{H}^{\dagger}[\rho_2] \right] \overline{N}_{LO}$$

where $\mathcal{H}[\rho]$ is the Hamiltonian that governs the rapidity dependence of the source distribution $W_{Y}[\rho]$:

$$\frac{\partial W_{Y}[\rho]}{\partial Y} = \mathcal{H}[\rho] \ W_{Y}[\rho]$$

FG, Lappi, Venugopalan (work in progress)



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- Why is it plausible ?
 - Reminder :

$$\begin{bmatrix} \delta \overline{N} \end{bmatrix}_{\text{divergent}} = \left\{ \int_{\vec{x}} \begin{bmatrix} \delta \mathcal{A}_{\text{in}}(\vec{x}) \end{bmatrix}_{\text{div}} T_{\vec{x}} \\ + \frac{1}{2} \int_{\vec{x}, \vec{y}} \begin{bmatrix} \mathbf{\Sigma}(\vec{x}, \vec{y}) \end{bmatrix}_{\text{div}} T_{\vec{x}} T_{\vec{y}} \right\} \overline{N}_{LO}$$

• Compare with the evolution Hamiltonian :

$$\mathcal{H}[\rho] = \int_{\vec{\boldsymbol{x}}_{\perp}} \sigma(\vec{\boldsymbol{x}}_{\perp}) \frac{\delta}{\delta\rho(\vec{\boldsymbol{x}}_{\perp})} + \frac{1}{2} \int_{\vec{\boldsymbol{x}}_{\perp},\vec{\boldsymbol{y}}_{\perp}} \chi(\vec{\boldsymbol{x}}_{\perp},\vec{\boldsymbol{y}}_{\perp}) \frac{\delta^2}{\delta\rho(\vec{\boldsymbol{x}}_{\perp})\delta\rho(\vec{\boldsymbol{y}}_{\perp})}$$

The coefficients σ and χ in the Hamiltonian are well known. There is a well defined calculation that will tell us if it works...



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Romatschke, Venugopalan (2005)

Rapidity dependent perturbations to the classical fields grow like $\exp(\#\sqrt{\tau})$ until the non-linearities become important :





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• The coefficient $\delta A_{in}(\vec{x})$ is boost invariant.

Hence, the divergences due to the unstable modes all come from the quadratic term in $\delta \overline{N}$:

$$\left[\delta \overline{N}\right]_{\text{unstable}}_{\text{modes}} = \left\{ \frac{1}{2} \int\limits_{\vec{x}, \vec{y}} \boldsymbol{\Sigma}(\vec{x}, \vec{y}) \boldsymbol{T}_{\vec{x}} \boldsymbol{T}_{\vec{y}} \right\} \overline{N}_{LO} [\mathcal{A}_{\text{in}}(\rho_1, \rho_2)]$$

When summed to all orders, this becomes a certain functional Z[T_x]:

$$\left[\delta \overline{N}\right]_{\text{unstable}} = Z[\boldsymbol{T}_{\vec{\boldsymbol{x}}}] \ \overline{N}_{LO}[\mathcal{A}_{\text{in}}(\rho_1, \rho_2)]$$



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This can be arranged in a more intuitive way :

$$\begin{bmatrix} \delta \overline{N} \end{bmatrix}_{\text{unstable}} = \int \begin{bmatrix} Da \end{bmatrix} \widetilde{Z}[a(\vec{x})] e^{i \int_{\vec{x}} a(\vec{x}) T_{\vec{x}}} \overline{N}_{LO}[\mathcal{A}_{\text{in}}(\rho_1, \rho_2)]$$
$$= \int \begin{bmatrix} Da \end{bmatrix} \widetilde{Z}[a(\vec{x})] \overline{N}_{LO}[\mathcal{A}_{\text{in}}(\rho_1, \rho_2) + a]$$

 \triangleright summing these divergences simply requires to add fluctuations to the initial condition for the classical problem \triangleright the fact that $\delta A_{in}(\vec{x})$ does not contribute implies that the distribution of fluctuations is real

Interpretation :

Despite the fact that the fields are coupled to strong sources, the classical approximation alone is not good enough, because the classical solution has unstable modes that can be triggered by the quantum fluctuations



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Fukushima, FG, McLerran (2006)

By a different method, one obtains Gaussian fluctuations characterized by :

$$\begin{split} \left\langle a_i(\eta, \vec{x}_{\perp}) \, a_j(\eta', \vec{x}'_{\perp}) \right\rangle &= \\ &= \frac{1}{\tau \sqrt{-(\partial_\eta/\tau)^2 - \partial_{\perp}^2}} \left[\delta_{ij} + \frac{\partial_i \partial_j}{(\partial_\eta/\tau)^2} \right] \delta(\eta - \eta') \, \delta(\vec{x}_{\perp} - \vec{x}'_{\perp}) \\ \left\langle e^i(\eta, \vec{x}_{\perp}) \, e^j(\eta', \vec{x}'_{\perp}) \right\rangle &= \\ &= \tau \sqrt{-(\partial_\eta/\tau)^2 - \partial_{\perp}^2} \left[\delta_{ij} - \frac{\partial_i \partial_j}{(\partial_\eta/\tau)^2 + \partial_{\perp}^2} \right] \delta(\eta - \eta') \, \delta(\vec{x}_{\perp} - \vec{x}'_{\perp}) \end{split}$$



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Classical solution in 2+1 dimensions


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Combining everything, one should write

$$\frac{d\overline{N}}{dYd^{2}\vec{p}_{\perp}} = \int \left[D\rho_{1}\right] \left[D\rho_{2}\right] W_{Y_{\text{beam}}-Y}\left[\rho_{1}\right] W_{Y_{\text{beam}}+Y}\left[\rho_{2}\right]$$
$$\times \int \left[Da\right] \quad \widetilde{Z}\left[a\right] \quad \frac{d\overline{N}\left[\mathcal{A}_{\text{in}}\left(\rho_{1},\rho_{2}\right)+a\right]}{dYd^{2}\vec{p}_{\perp}}$$

▷ This formula resums (all?) the divergences that occur at one loop



Unstable modes – Interpretation

Tree level :

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Unstable modes – Interpretation



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Unstable modes – Interpretation

After summing the fluctuations, things may look like this :



 \triangleright these splittings may help to fight against the expansion ? Note : the timescale for this process is $\tau \sim Q_s^{-1} \ln^2(1/\alpha_s)$



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Definition

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• One can encode the information about all the probabilities P_n in a generating function defined as :

$$F(z)\equiv\sum_{n=0}^{\infty}\ P_n\ z^n$$

From the expression of P_n in terms of the operator \mathcal{D} , we can write :

$$F(z) = e^{z\mathcal{D}} e^{iV} e^{-iV^*}$$

- Reminder :
 - $e^{\mathcal{D}} e^{iV} e^{-iV^*}$ is the sum of all the cut vacuum diagrams
 - The cuts are produced by the action of \mathcal{D}
- Therefore, F(z) is the sum of all the cut vacuum diagrams in which each cut line is weighted by a factor z



What would it be good for ?

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• Let us pretend that we know the generating function F(z). We could get the probability distribution as follows :

$$P_n = \frac{1}{2\pi} \int_0^{2\pi} d\theta \ e^{-in\theta} \ F(e^{i\theta})$$

Note : this is trivial to evaluate numerically by a FFT :





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F(z) at Leading Order

We have :
$$F'(z) = \mathcal{D} \left\{ e^{z\mathcal{D}} e^{iV} e^{-iV^*} \right\}$$

• By the same arguments as in the case of \overline{N} , we get :



- The major difference is that the cut graphs that must be evaluated have a factor z attached to each cut line
- At tree level (LO), we can write F'(z)/F(z) in terms of solutions of the classical Yang-Mills equations, but these solutions are not retarded anymore, because :

 $\cdots + z \cdots \neq retarded propagator$



F(z) at Leading Order

The derivative F'/F has an expression which is formally identical to that of \overline{N} ,

$$rac{F'(z)}{F(z)} = \int rac{d^3 ec p}{(2\pi)^3 2 E_{oldsymbol{p}}} \int_{x,y} \, e^{i p \cdot (x-y)} \, \Box_x \Box_y \, \sum_\lambda \epsilon^\mu_\lambda \epsilon^
u_\lambda \, A^{(+)}_\mu(x) A^{(-)}_
u(y) \; ,$$

with $A_{\mu}^{(\pm)}(x)$ two solutions of the Yang-Mills equations If one decomposes these fields into plane-waves,

$$\boldsymbol{A}_{\mu}^{(\varepsilon)}(\boldsymbol{x}) \equiv \int \frac{d^{3}\boldsymbol{\vec{p}}}{(2\pi)^{3}2E_{\boldsymbol{p}}} \left\{ f_{+}^{(\varepsilon)}(\boldsymbol{x}^{0},\boldsymbol{\vec{p}})e^{-i\boldsymbol{p}\cdot\boldsymbol{x}} + f_{-}^{(\varepsilon)}(\boldsymbol{x}^{0},\boldsymbol{\vec{p}})e^{i\boldsymbol{p}\cdot\boldsymbol{x}} \right\}$$

the boundary conditions are :

$$\begin{split} f_{+}^{(+)}(-\infty,\vec{p}) &= f_{-}^{(-)}(-\infty,\vec{p}) = 0 \\ f_{+}^{(-)}(+\infty,\vec{p}) &= z \, f_{+}^{(+)}(+\infty,\vec{p}) \quad , \quad f_{-}^{(+)}(+\infty,\vec{p}) = z \, f_{-}^{(-)}(+\infty,\vec{p}) \end{split}$$

There are boundary conditions both at $x_0 = -\infty$ and $x_0 = +\infty \triangleright$ not an initial value problem \triangleright hard...

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Remarks on factorization

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• Exclusive processes

- As we have seen, the fact that the calculation of the first moment N can be formulated as an initial value problem seems quite helpful for proving factorization
- If the retarded nature of the fields is crucial, then factorization does not hold for the generating function F(z), or equivalently for the individual probabilities P_n
- Note : by differentiating the result for F(z) with respect to z, and then setting z = 1, we can obtain formulas for higher moments of the distribution



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Summary

So far, we have considered only inclusive quantities – i.e. the *P_n* are defined as probabilities of producing particles anywhere in phase-space

What about events where a part of the phase-space remains unoccupied ? e.g. rapidity gaps





Main issues

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1. How do we calculate the probabilities P_n^{excl} with an excluded region in the phase-space ? Can one calculate the total gap probability $P_{\text{gap}} = \sum_n P_n^{\text{excl}}$?

- 2. What is the appropriate distribution of sources $W_Y^{\text{excl}}[\rho]$ to describe a projectile that has not broken up ?
- 3. How does it evolve with rapidity ?
 - See : Hentschinski, Weigert, Schafer (2005)
- 4. Are there some factorization results, and for which quantities do they hold ?



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Summarv

The probabilities $P_n^{\text{excl}}[\Omega]$, for producing *n* particles – only in the region Ω – can also be constructed from the vacuum diagrams, as follows :

$$\mathcal{D}_n^{ ext{excl}}[\mathbf{\Omega}] \;=\; rac{1}{n!}\; \mathcal{D}_{\mathbf{\Omega}}^n \;\; e^{iV}\; e^{-iV^*}$$

where \mathcal{D}_{Ω} is an operator that removes two sources and links the corresponding points by a cut (on-shell) line, for which the integration is performed only in the region Ω

One can define a generating function,

Exclusive probabilities

$$F_{_{oldsymbol{\Omega}}}(z)\equiv\sum_n \, P^{\mathrm{excl}}_n[oldsymbol{\Omega}] \; z^n$$
 ,

whose derivative is given by the same diagram topologies as the derivative of the generating function for inclusive probabilities



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Differences with the inclusive case :

- In the diagrams that contribute to $F'_{\Omega}(z)/F_{\Omega}(z)$, the cut propagators are restricted to the region Ω of the phase-space
 - \triangleright at leading order, this only affects the boundary conditions for the classical fields in terms of which one can write $F'_{\Omega}(z)/F_{\Omega}(z)$

▷ not more difficult than the inclusive case

• Contrary to the inclusive case – where we know that F(1) = 1 – the integration constant needed to go from $F'_{\Omega}(z)/F_{\Omega}(z)$ to $F_{\Omega}(z)$ is non-trivial. This is due to the fact that the sum of all the exclusive probabilities is smaller than unity

 $ightarrow F_{\Omega}(1)$ is in fact the probability of not having particles in the complement of Ω – i.e. the gap probability



Survival probability

■ We can write :

$$F_{\Omega}(z) = F_{\Omega}(1) \exp\left\{\int_{1}^{z} d\tau \frac{F_{\Omega}'(\tau)}{F_{\Omega}(\tau)}\right\}$$

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 \vartriangleright the prefactor $F_{\Omega}(1)$ will appear in all the exclusive probabilities

This prefactor is nothing but the famous "survival probability" for a rapidity gap

One can in principle calculate it by the general techniques developed for calculating inclusive probabilities :

```
F_{\boldsymbol{\Omega}}(1) = F_{1-\boldsymbol{\Omega}}^{\text{incl}}(0)
```

Note : it is incorrect to say that a certain process with a gap can be calculated by multiplying the probability of this process without the gap by the survival probability



Factorization ?

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- In order to discuss factorization for exclusive quantities, one must calculate their 1-loop corrections, and study the structure of the divergences...
- Except for the case of Deep Inelastic Scattering, nothing is known regarding factorization for exclusive processes in a high density environment
- For the overall framework to be consistent, one should have factorization between the gap probability, $F_{\Omega}(1)$, and the source density studied in Hentschinski, Weigert, Schafer (2005) (and the ordinary $W_{\gamma}[\rho]$ on the other side)
- The total gap probability is the "most inclusive" among the exclusive quantities one may think of. For what quantities if any does factorization work ?



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Introduction

When the parton densities in the projectiles are large, the study of particle production becomes rather involved

> non-perturbative techniques that resum all-twist contributions are needed

- At Leading Order, the inclusive gluon spectrum can be calculated from the classical solution with retarded boundary conditions on the light-cone
- At Next-to-Leading Order, the gluonic corrections can be seen as a perturbation of the initial value problem encountered at LO
- Resummation of the leading divergences to all orders :
 - \triangleright Evolution with Y of the distribution of sources
 - > Quantum fluctuations on top of initial condition for the classical solution in the forward light-cone



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- Diagrammatic interpretation
- Quark production
- Longitudinal expansion
- AGK identities

Extra bits



Parton evolution

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 \triangleright assume that the projectile is big, e.g. a nucleus, and has many valence quarks (only two are represented)

 \triangleright on the contrary, consider a small probe, with few partons

 \triangleright at low energy, only valence quarks are present in the hadron wave function



Parton evolution



- Quark production
- Longitudinal expansion
- AGK identities





> when energy increases, new partons are emitted

▷ the emission probability is $\alpha_s \int \frac{dx}{x} \sim \alpha_s \ln(\frac{1}{x})$, with x the longitudinal momentum fraction of the gluon ▷ at small-x (i.e. high energy), these logs need to be resummed



Parton evolution







▷ as long as the density of constituents remains small, the evolution is linear: the number of partons produced at a given step is proportional to the number of partons at the previous step (BFKL)



Parton evolution



- Longitudinal expansion
- AGK identities





▷ eventually, the partons start overlapping in phase-space

⊳ parton recombination becomes favorable

▷ after this point, the evolution is non-linear: the number of partons created at a given step depends non-linearly

on the number of partons present previously



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Gribov, Levin, Ryskin (1983)

Number of gluons per unit area:

$$\rho \sim \frac{x G_A(x, Q^2)}{\pi R_A^2}$$

Recombination cross-section:

$$\sigma_{gg \to g} \sim \frac{\alpha_s}{Q^2}$$

Recombination happens if $\rho\sigma_{gg\rightarrow g}\gtrsim 1$, i.e. $Q^2 \lesssim Q_s^2$, with:

$$Q_s^2 \sim \frac{\alpha_s x G_A(x, Q_s^2)}{\pi R_A^2} \sim A^{1/3} \frac{1}{x^{0.3}}$$

At saturation, the phase-space density is:

$$\frac{dN_g}{d^2 \vec{\boldsymbol{x}}_\perp d^2 \vec{\boldsymbol{p}}_\perp} \sim \frac{\rho}{Q^2} \sim \frac{1}{\alpha_s}$$



Saturation domain



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Diagrammatic interpretation

One loop :



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Diagrammatic interpretation

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Diagrammatic interpretation

Diagrammatic interpretation

One loop :



 \triangleright The sum of tree diagrams for fluctuations on top of the classical field with initial condition A_{in} gives the classical field with a shifted initial condition $A_{in} + a$

 \triangleright If we keep only the fastest growing terms, we need only the leading two-point correlation of the initial fluctuation *a*



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• Diagrammatic interpretation

Quark production

FG, Kajantie, Lappi (2004, 2005)

$$E_{\boldsymbol{p}} \frac{d \langle n_{\text{quarks}} \rangle}{d^{3} \boldsymbol{\vec{p}}} = \frac{1}{16\pi^{3}} \int_{x,y} e^{i \boldsymbol{p} \cdot (\boldsymbol{x}-\boldsymbol{y})} \, \boldsymbol{\partial}_{x} \boldsymbol{\partial}_{y} \, \left\langle \overline{\boldsymbol{\psi}}(\boldsymbol{x}) \boldsymbol{\psi}(\boldsymbol{y}) \right\rangle$$

Dirac equation in the classical color field :





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Quark production

FG, Kajantie, Lappi (2004, 2005)

$$E_{\boldsymbol{p}} \frac{d \langle n_{\text{quarks}} \rangle}{d^{3} \boldsymbol{\vec{p}}} = \frac{1}{16\pi^{3}} \int_{x,y} e^{i \boldsymbol{p} \cdot (\boldsymbol{x} - \boldsymbol{y})} \, \boldsymbol{\partial}_{x} \boldsymbol{\partial}_{y} \, \left\langle \overline{\boldsymbol{\psi}}(\boldsymbol{x}) \boldsymbol{\psi}(\boldsymbol{y}) \right\rangle$$

Dirac equation in the classical color field :





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For a system finite in the η direction, the gluons will have a longitudinal velocity tied to their space-time rapidity



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For a system finite in the η direction, the gluons will have a longitudinal velocity tied to their space-time rapidity





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For a system finite in the η direction, the gluons will have a longitudinal velocity tied to their space-time rapidity



▷ at late times : if particles fly freely, only one longitudinal velocity can exist at a given η : $v_z = \tanh(\eta)$

▷ the expansion of the system is the main obstacle to local isotropy


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Let P_n be the probability of producing n particles
Define the generating function :

$$F(z) \equiv \sum_{n=0}^{\infty} P_n \, z^n$$

From unitarity, $F(1) = \sum_{n=0}^{\infty} P_n = 1$. Thus, we can write

$$\ln(F(z)) \equiv \sum_{r=1}^{\infty} b_r \left(z^r - 1\right)$$

• At the moment, we need to know only very little about the b_r :

- F(z) is a sum of diagrams that may or may not be connected
- $\ln(F(z))$ involves only connected diagrams. Hence, the b_r 's are given by certain sums of connected diagrams
- Every diagram in b_r produces r particles



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Example : typical term in the coefficient of z^{11} , with contributions from b_5 and b_6 :





Distribution of connected subdiagrams

From this form of the generating function, one gets :

$$P_n = \sum_{p=0}^n e^{-\sum_r b_r} \frac{1}{p!} \sum_{\alpha_1 + \dots + \alpha_p = n} \frac{b_{\alpha_1} \cdots b_{\alpha_n}}{\sum_{\alpha_1 + \dots + \alpha_p = n} b_{\alpha_1} \cdots b_{\alpha_n}}$$

probability of producing n particles in p cut subdiagrams

Summing on n, we get the probability of p cut subdiagrams :

$$\boldsymbol{R_p} = \frac{1}{p!} \left[\sum_{r=1}^{\infty} \boldsymbol{b_r} \right]^p e^{-\sum_r \boldsymbol{b_r}}$$

Note : Poisson distribution of average $\langle N_{\rm subdiagrams} \rangle = \sum_r b_r$

By expanding the exponential, we get the probability of having p cut subdiagrams out of a total of m :

$$\boldsymbol{R_{p,m}} = \frac{(-1)^{m-p}}{(m-p)! \, p!} \left[\sum_{r=1}^{\infty} \boldsymbol{b_r} \right]^m$$

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The quantities $R_{p,m}$ obey the following relations :

$$egin{aligned} &orall m\geq 2\;, \quad \sum_{p=1}^m p\,R_{p,m}=0\;, \ &orall m\geq 3\;, \quad \sum_{p=1}^m p(p-1)\,R_{p,m}=0\;, \cdots \end{aligned}$$

- Interpretation : contributions with more than 1 subdiagram cancel in the average number of cut subdiagrams, etc...
- Correspondence with the original relations by Abramovsky-Gribov-Kancheli :
 - The original derivation is formulated in the framework of reggeon effective theories
 - ◆ Dictionary: reggeon → subdiagram
 - These identities are more general than "reggeons", and are valid for any kind of subdiagrams



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The AGK relations, obtained by "integrating out" the number of produced particles, describe the combinatorics of connected diagrams

⊳ by doing that, a lot of information has been discarded

For instance, to compute the average number of produced particles, one would write :

$$\langle n \rangle = \langle N_{\text{subdiagrams}} \rangle \times \langle \# \text{ of particles per diagram} \rangle$$

requires a more detailed description