

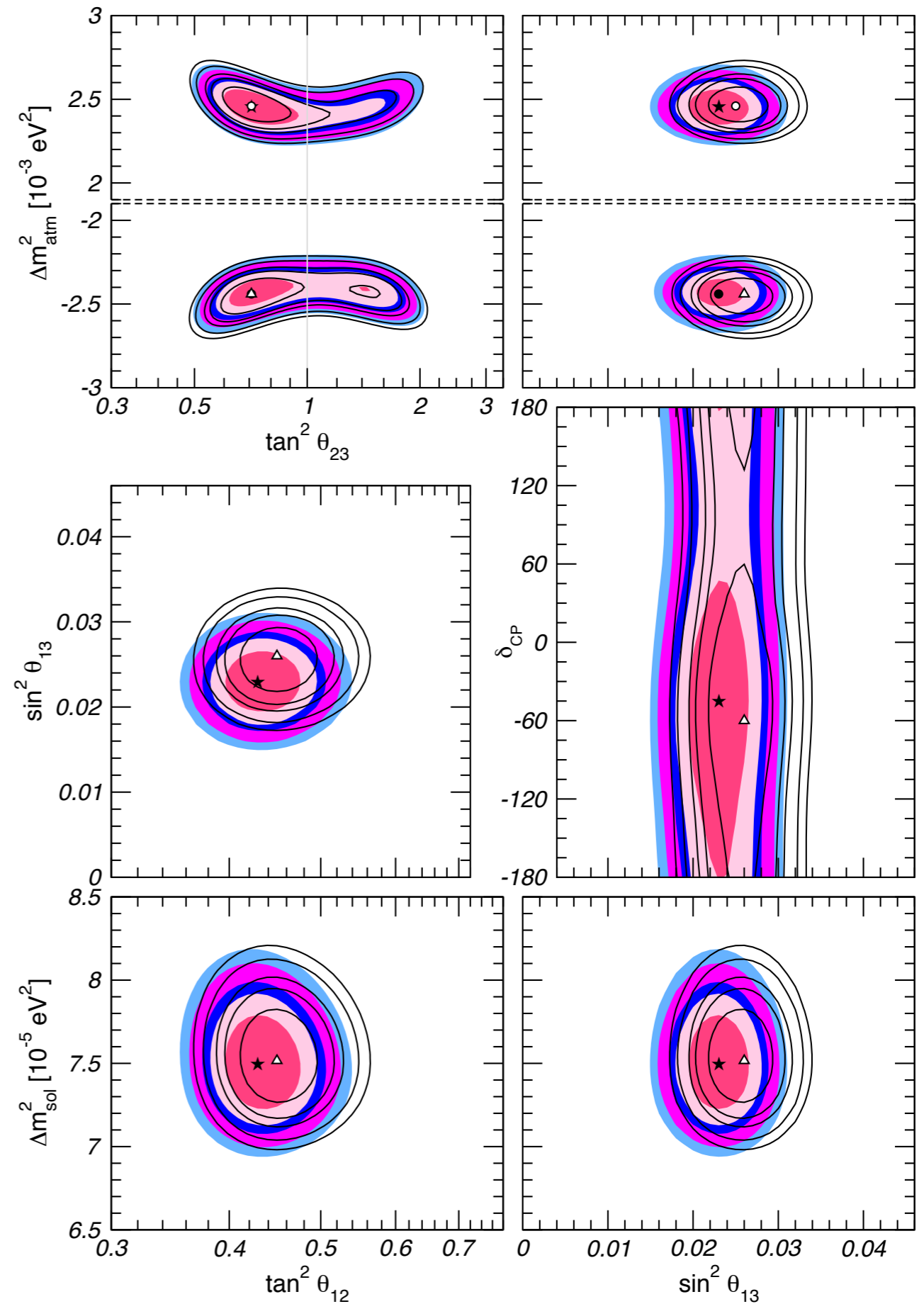
Review on global fits

What's ν - Invisibles12, GGI, Florence, 25 June 2012

Thomas Schwetz



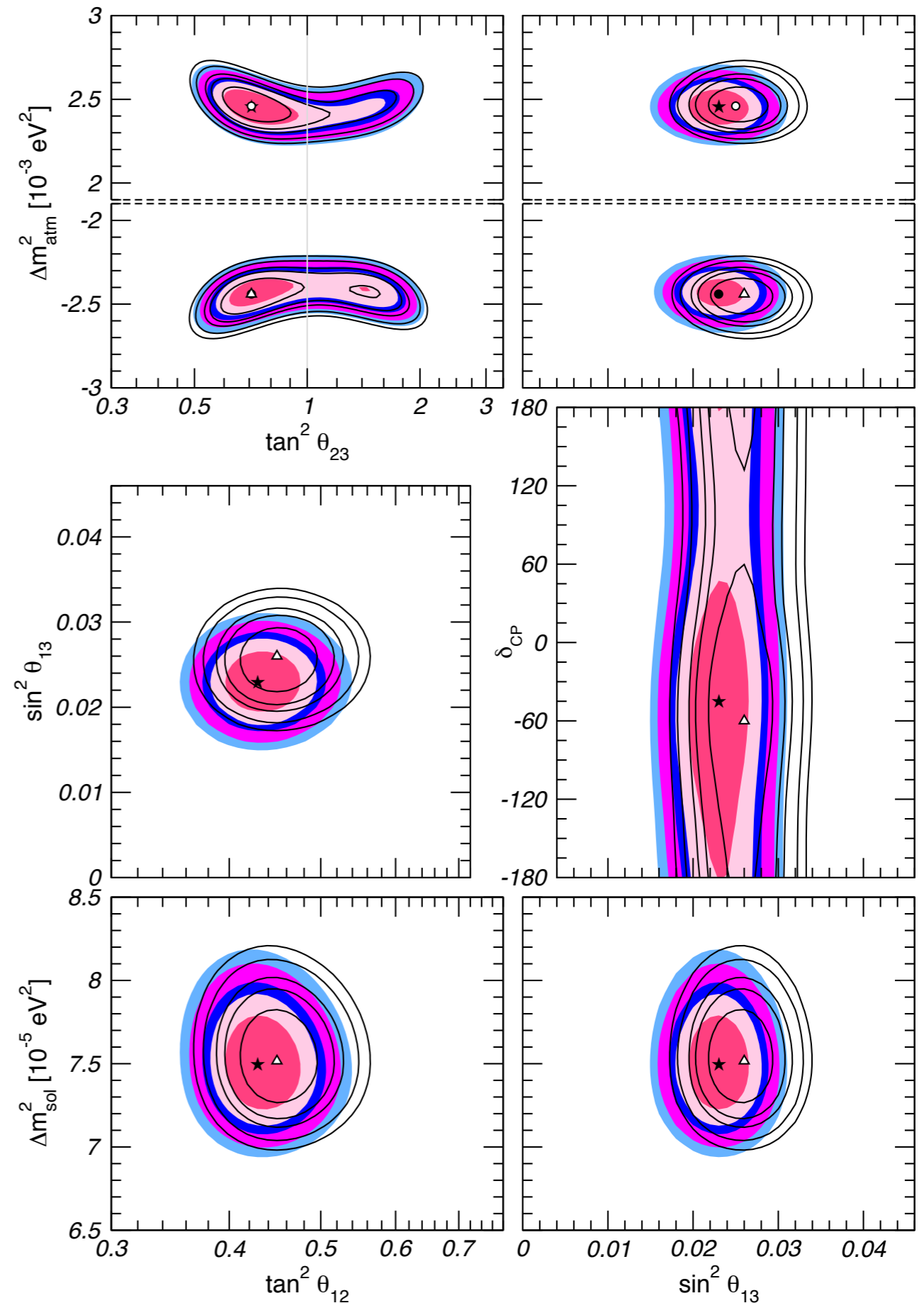
preliminary results from
work in collab. with
C. Gonzalez-Garcia,
M. Maltoni, J. Salvado
including latest data
from Neutrino2012



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Outline:

- *the θ_{13} revolution*
- *non-maximality of θ_{23}*
- *CP phase and hierarchy*



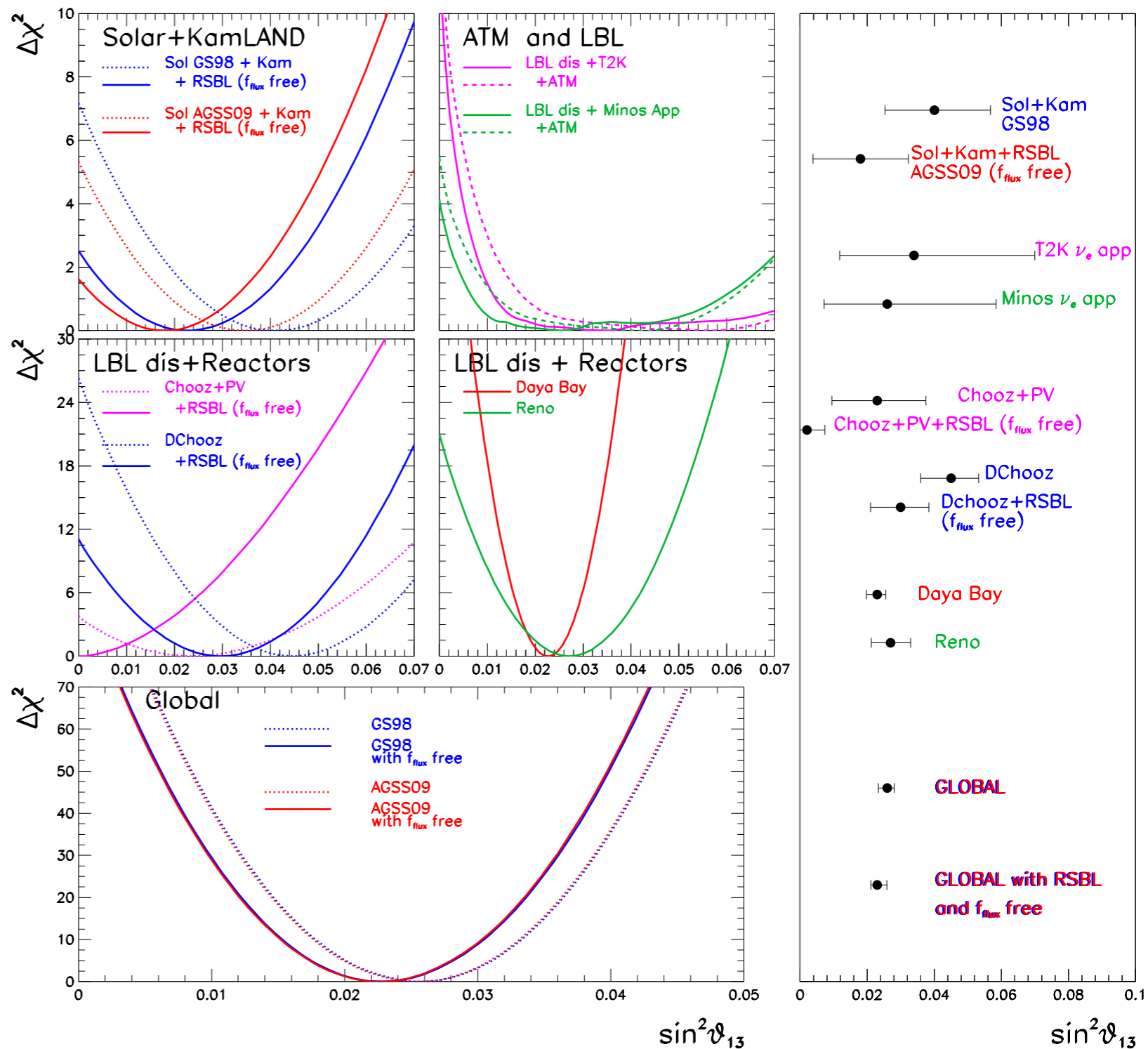
The θ_{13} revolution

- About 1 year ago:
6 events in T2K: 2.5σ
- global fits gave
 $>3\sigma$ for the first
time

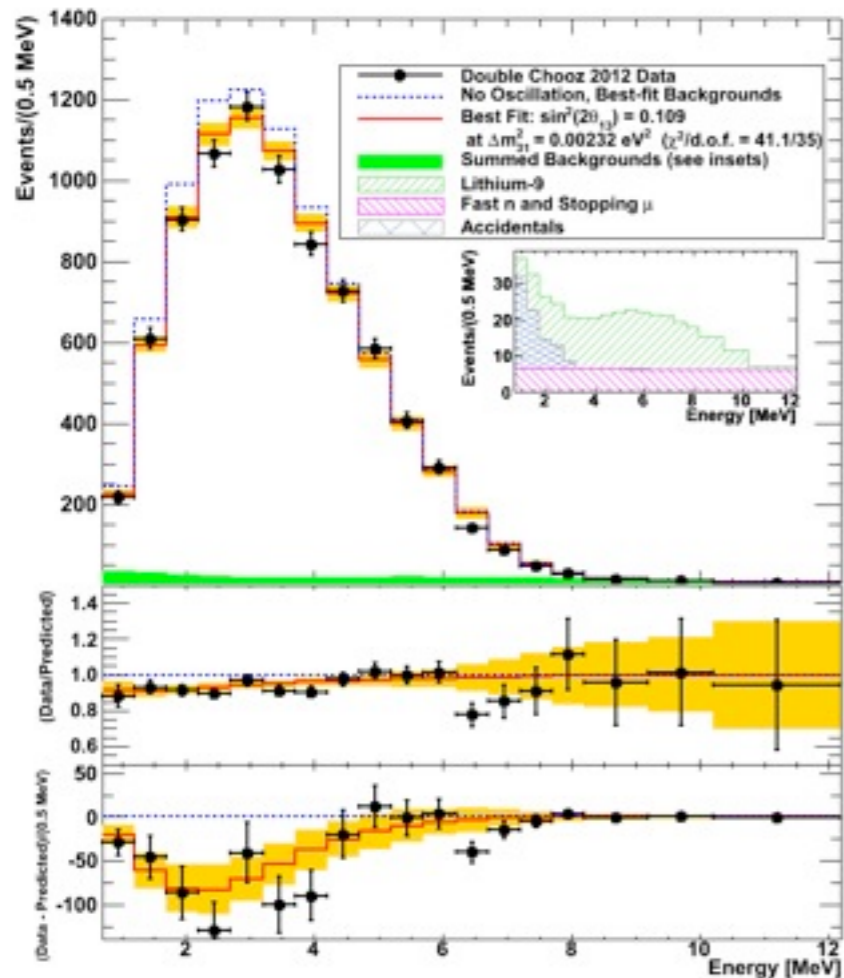
Fogli et al, *I 106.6028*
TS, Tortola, Valle *I 108.1376*

- DoubleChooz,
DayaBay, RENO

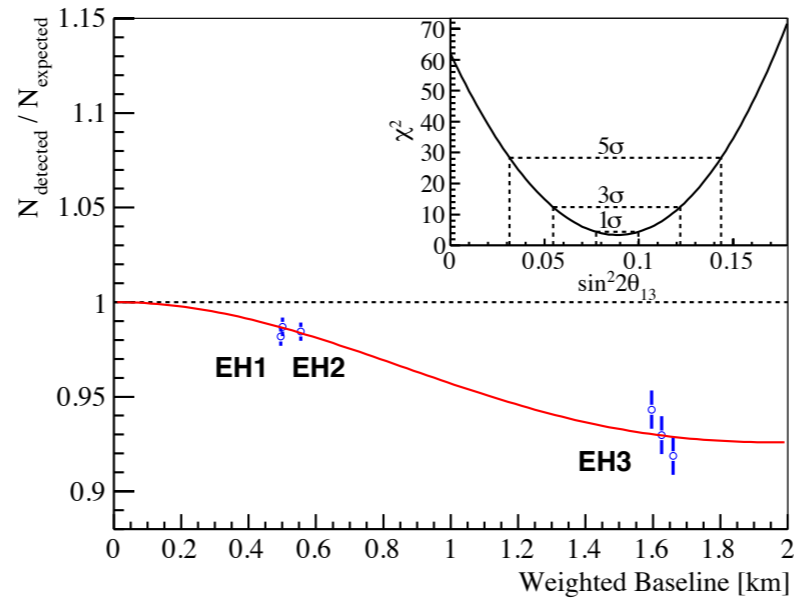
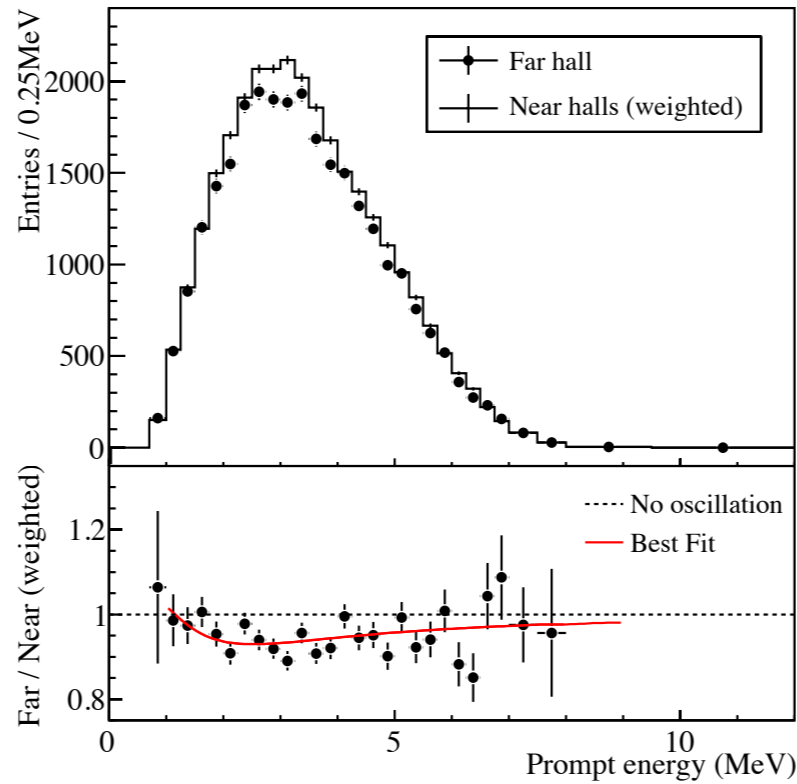
- post-Neutrino 12:
 $\Delta\chi^2 \approx 100$ in the
global fit



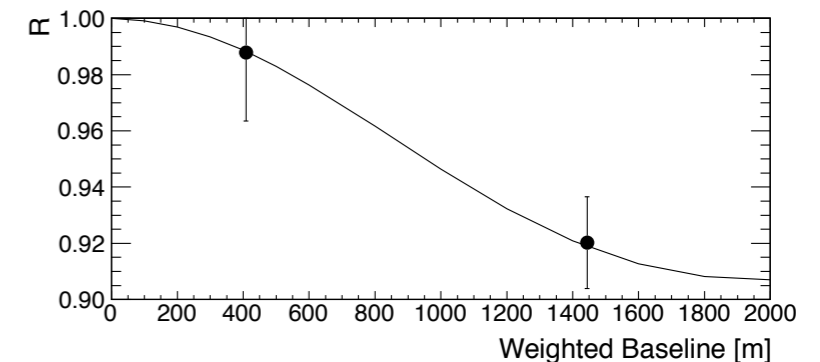
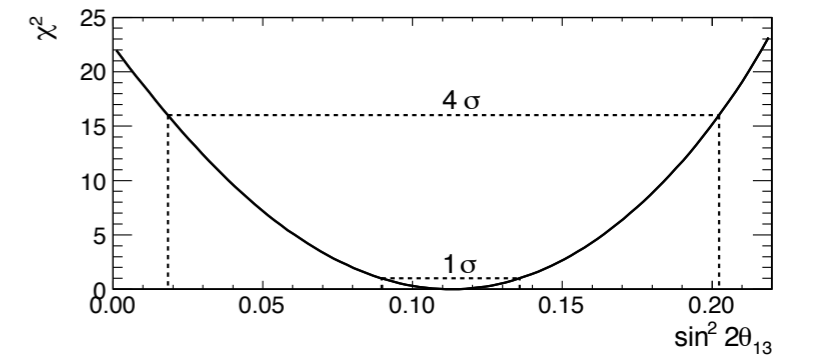
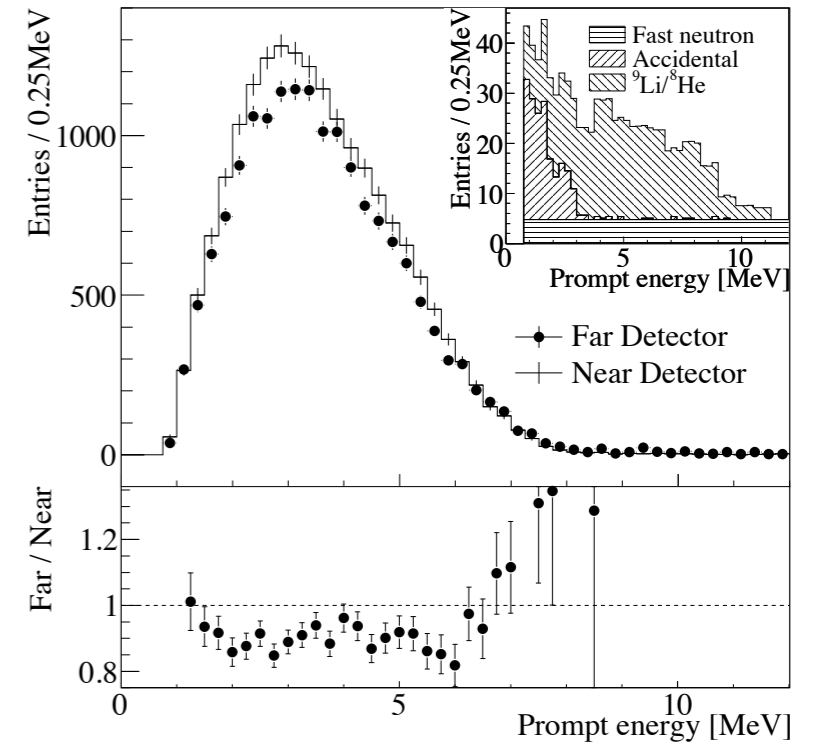
The θ_{13} revolution - the reactors



DoubleChooz @ Neutrino2012



DayaBay @ Neutrino2012



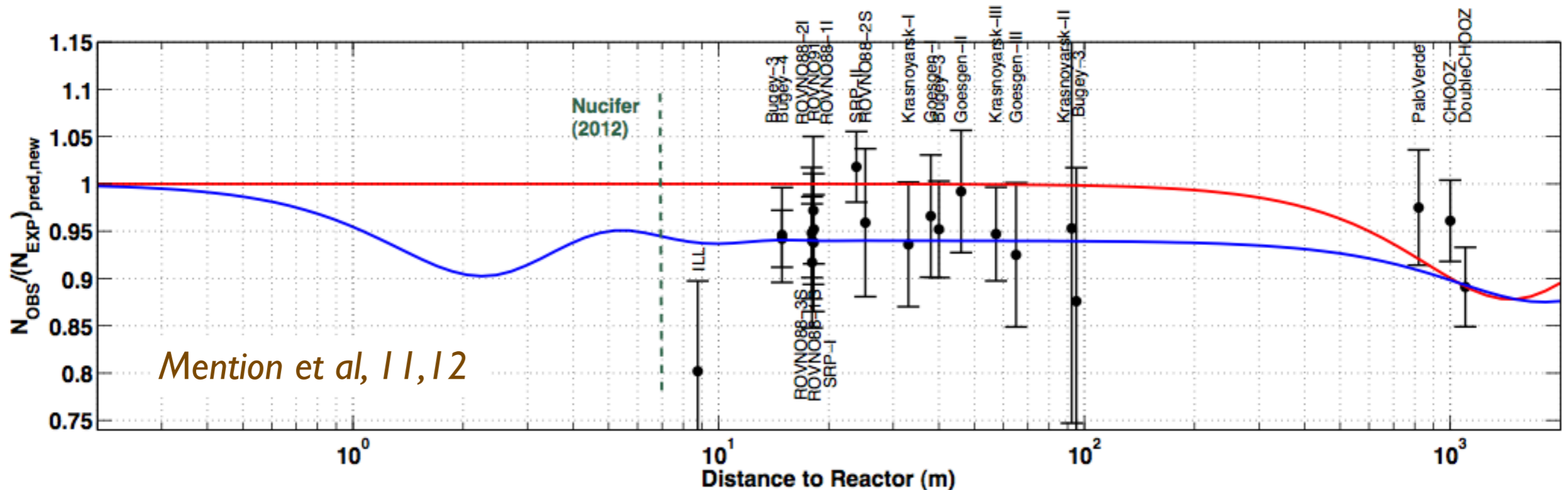
RENO, 1204.0626

The reactor anomaly

- ▶ to predict the $\bar{\nu}_e$ flux from nuclear reactors one has to convert the measured e^- spectra from ^{235}U , ^{239}Pu , ^{241}Pu into neutrino spectra
Schreckenbach et al., 82, 85, 89
- ▶ recent improved calculation *Mueller et al., 1101.2663* $\sim 3\%$ higher fluxes (ab initio calculations + virtual branches for missing part)
- ▶ confirmed by independent calculation *P. Huber, 1106.0687* (virtual branches)
- ▶ increase of predicted number of neutrino-induced events compared to old flux calculations:

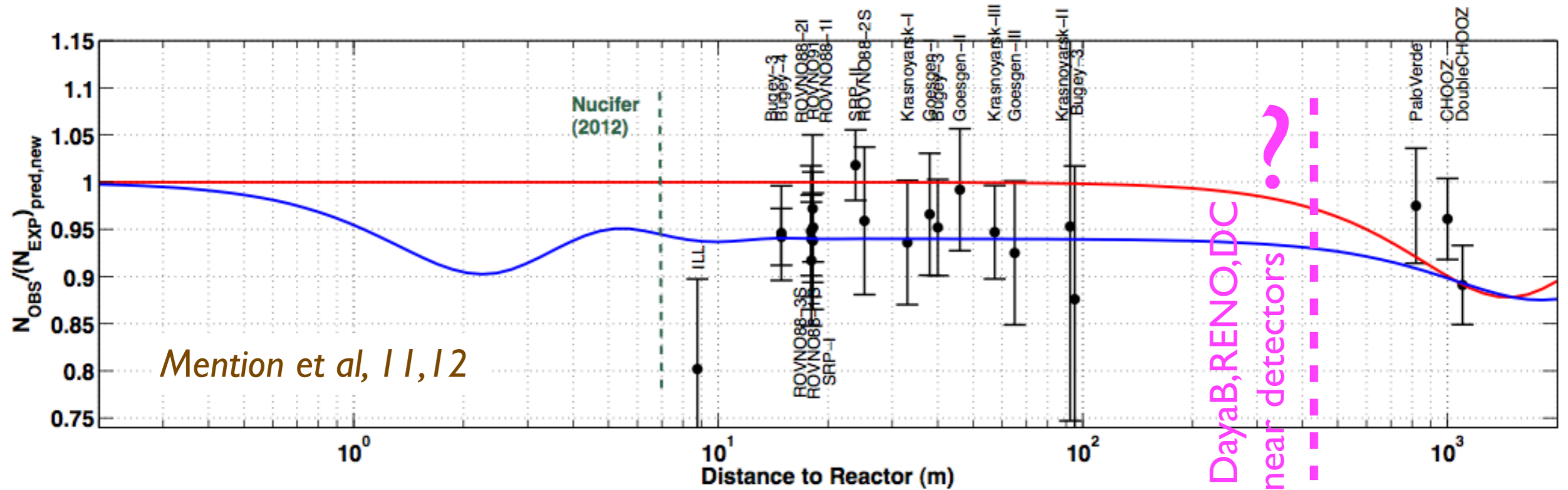
^{235}U	^{239}Pu	^{241}Pu	^{238}U
3.7%	4.2%	4.7%	9.8%

The reactor anomaly



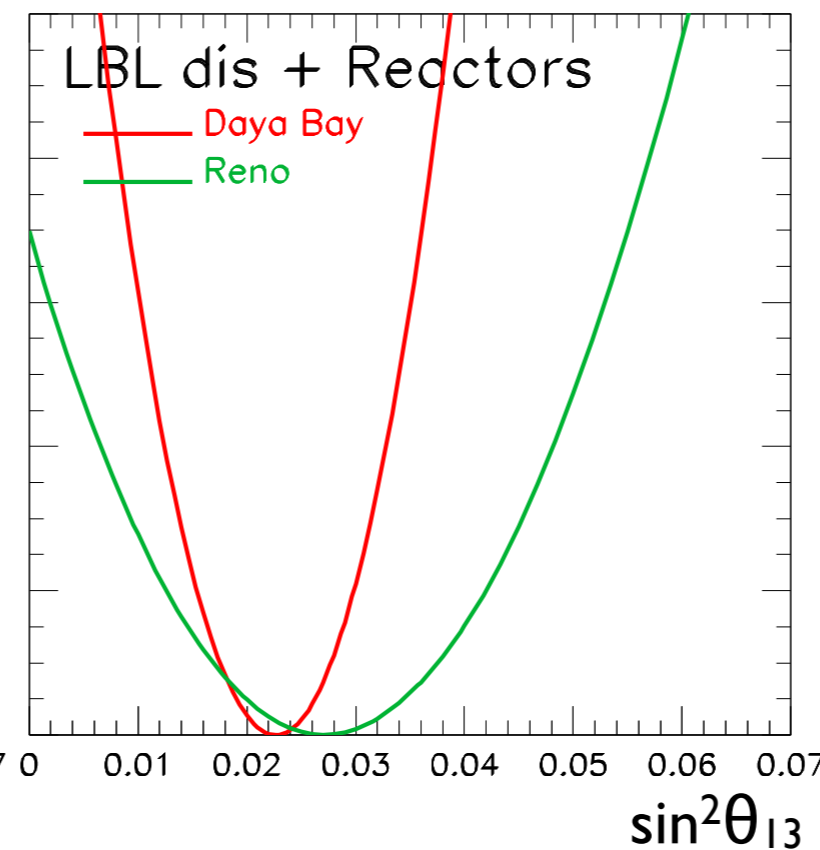
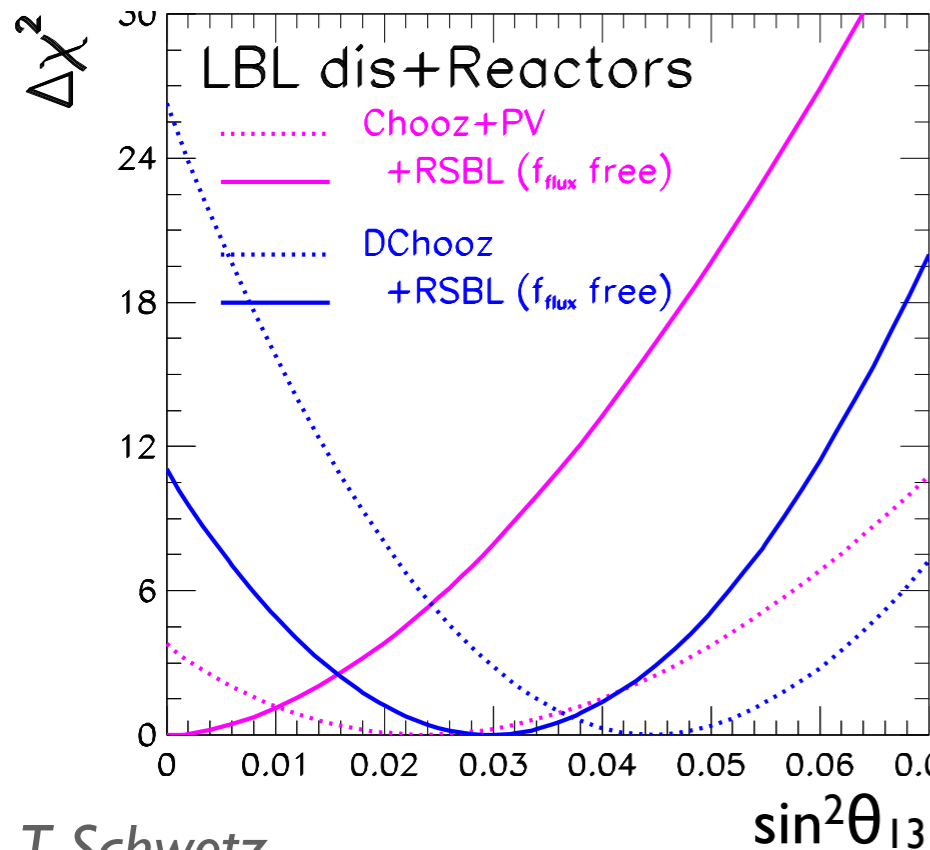
- SBL reactor data ($L < 100\text{m}$) in tension with predicted flux $f = 0.935 \pm 0.024$ (different from 1 @ 2.7σ)
- systematics?
 - ▶ normalization of ILL electron spectra
 - ▶ neutron lifetime (use 2012 PDG value)
- sterile neutrinos at the eV scale? *talk by Maltoni*

The reactor anomaly and the θ_{13} determination



Mention et al, 11,12

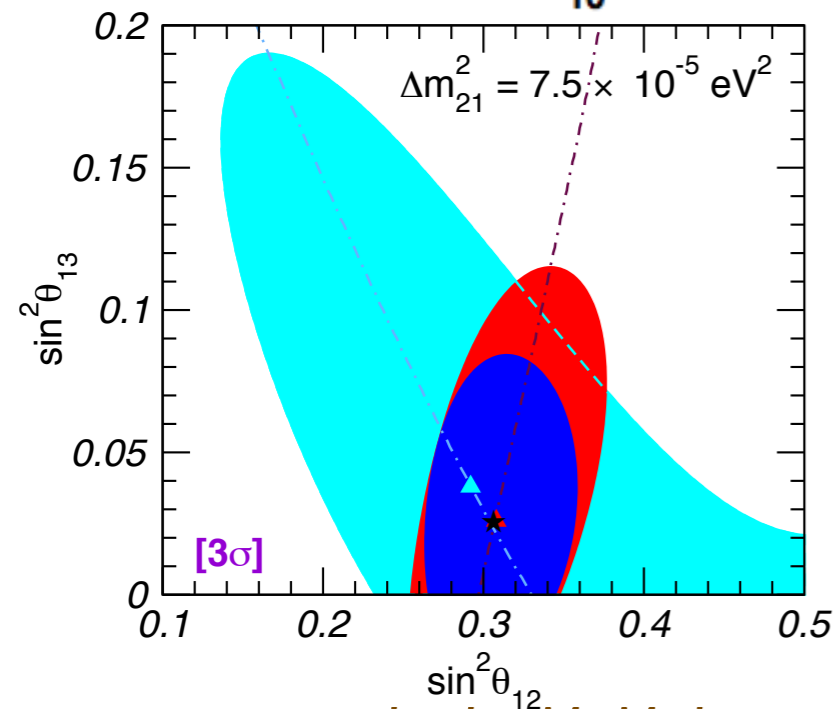
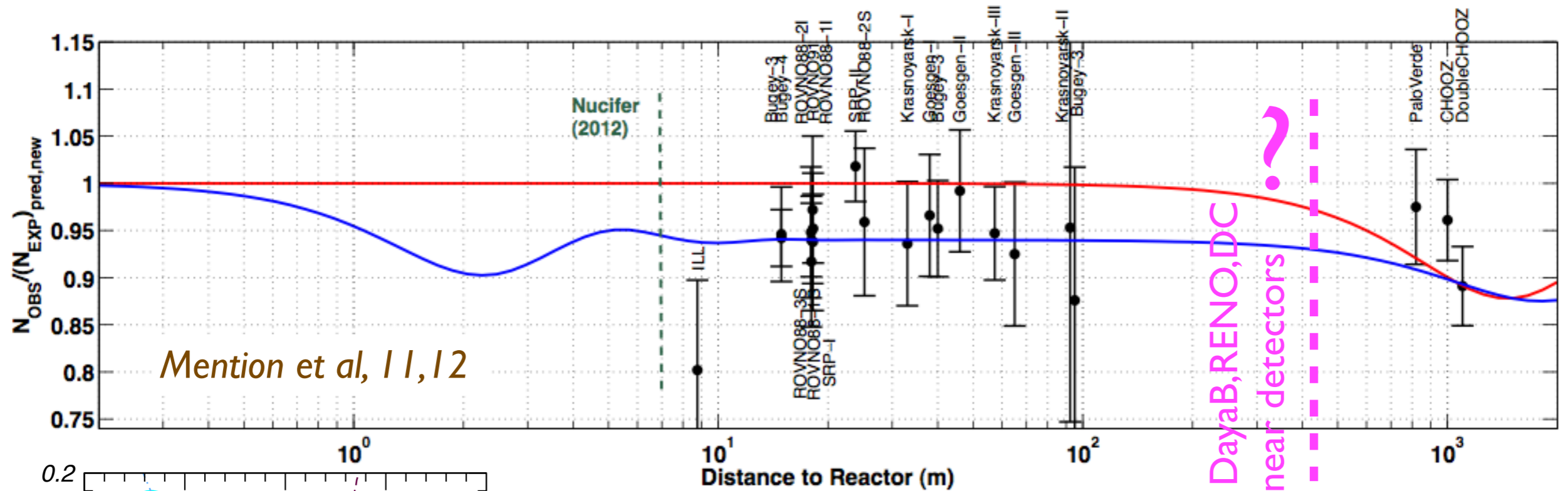
DayaB, RENO, DC
near detectors



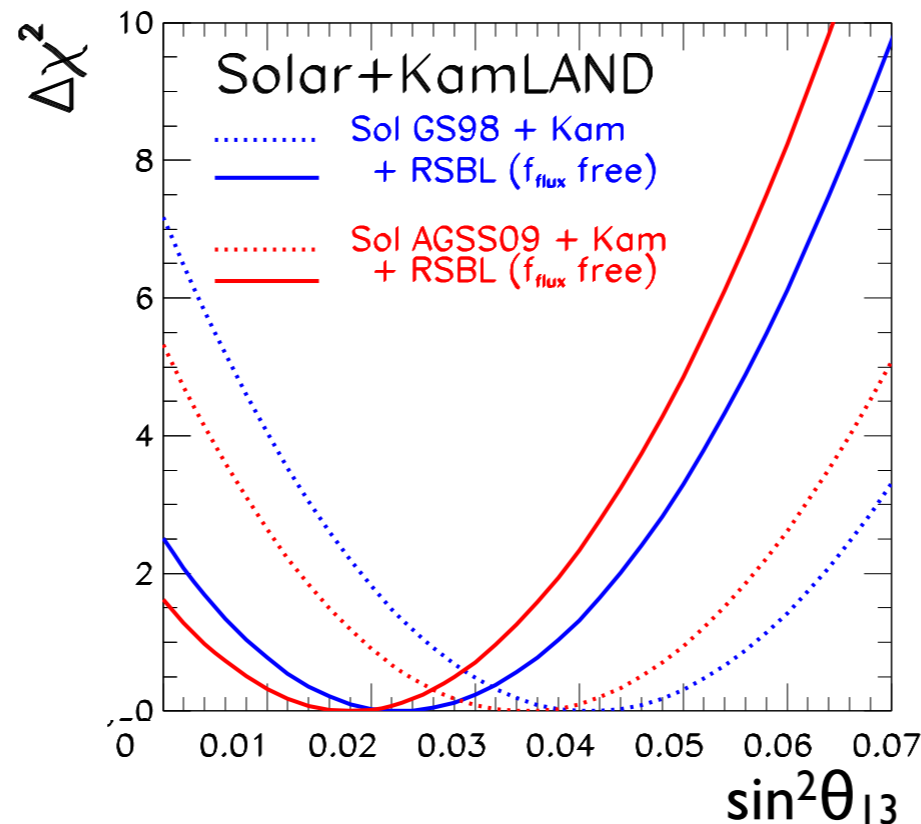
two extreme assumptions:

- use fluxes from *Huber, 1106.0687* without SBL reactor data
- leave reactor flux free and include SBL data in fit

The reactor anomaly and the θ_{13} determination



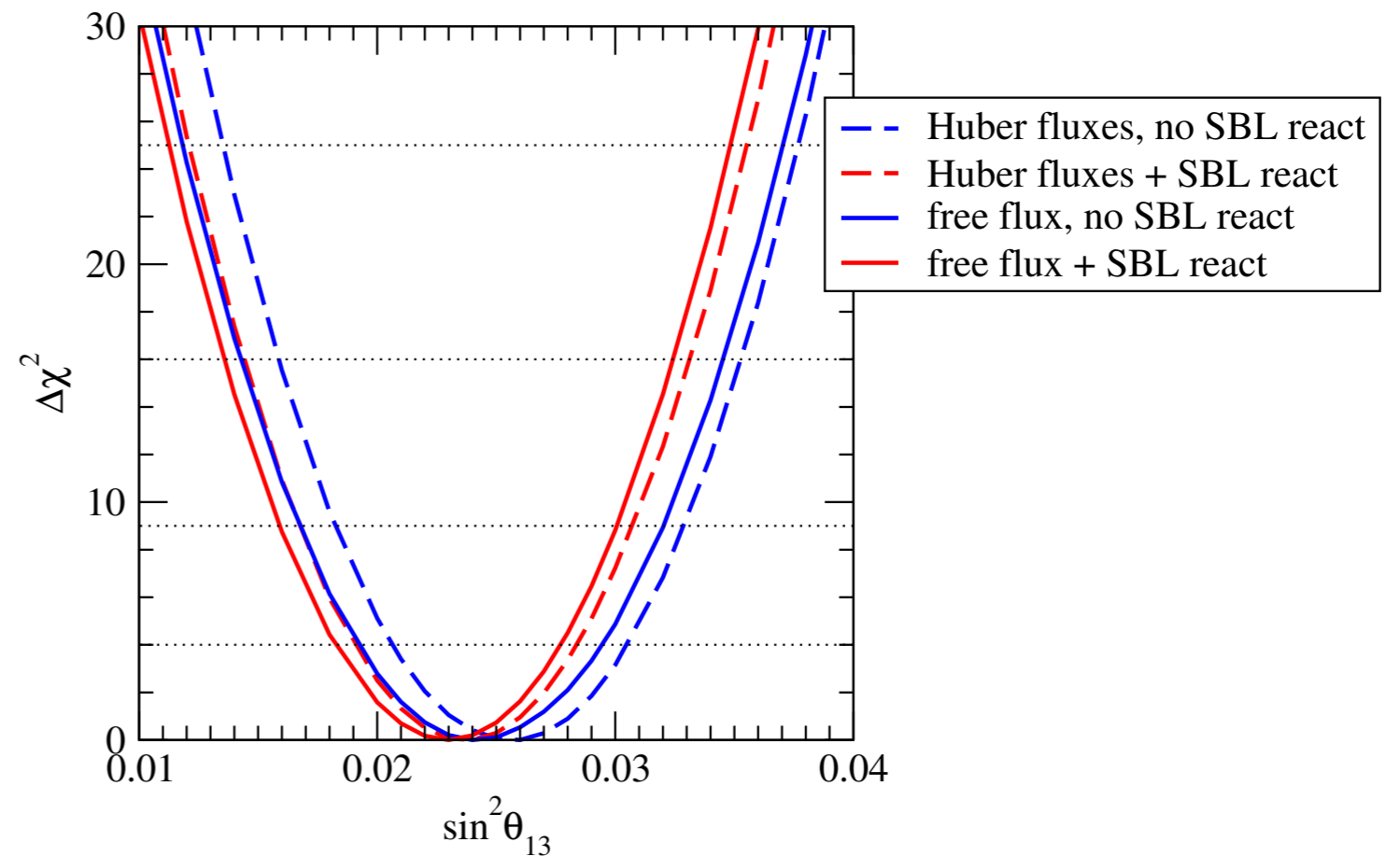
plot by M. Maltoni;
Fogli et al, 08;
TS, Tortola, Valle, 08



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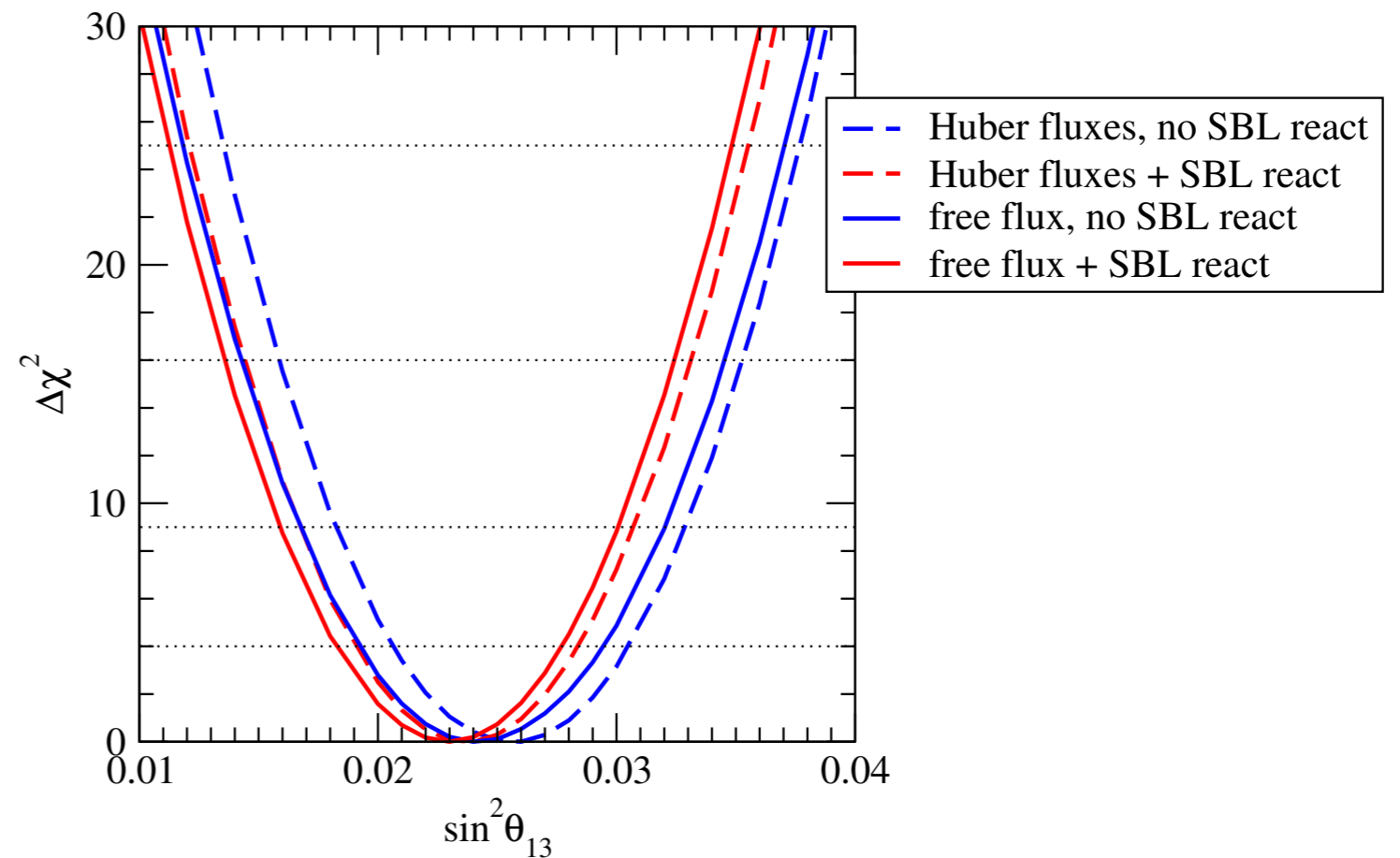
$$\sin^2 \theta_{13} = 0.0257 \pm 0.0025, \quad \theta_{13} = (9.2 \pm 0.46)^\circ, \quad \sin^2 2\theta_{13} = 0.100 \pm 0.0095$$

- *leave reactor flux free and include SBL data in fit*

$$\sin^2 \theta_{13} = 0.0230 \pm 0.0023, \quad \theta_{13} = (8.7 \pm 0.44)^\circ, \quad \sin^2 2\theta_{13} = 0.090 \pm 0.0090$$

The reactor anomaly and the θ_{13} determination

- result depends on data which in principle is not sensitive to θ_{13}
- shift at the level of 1σ



two extreme assumptions:

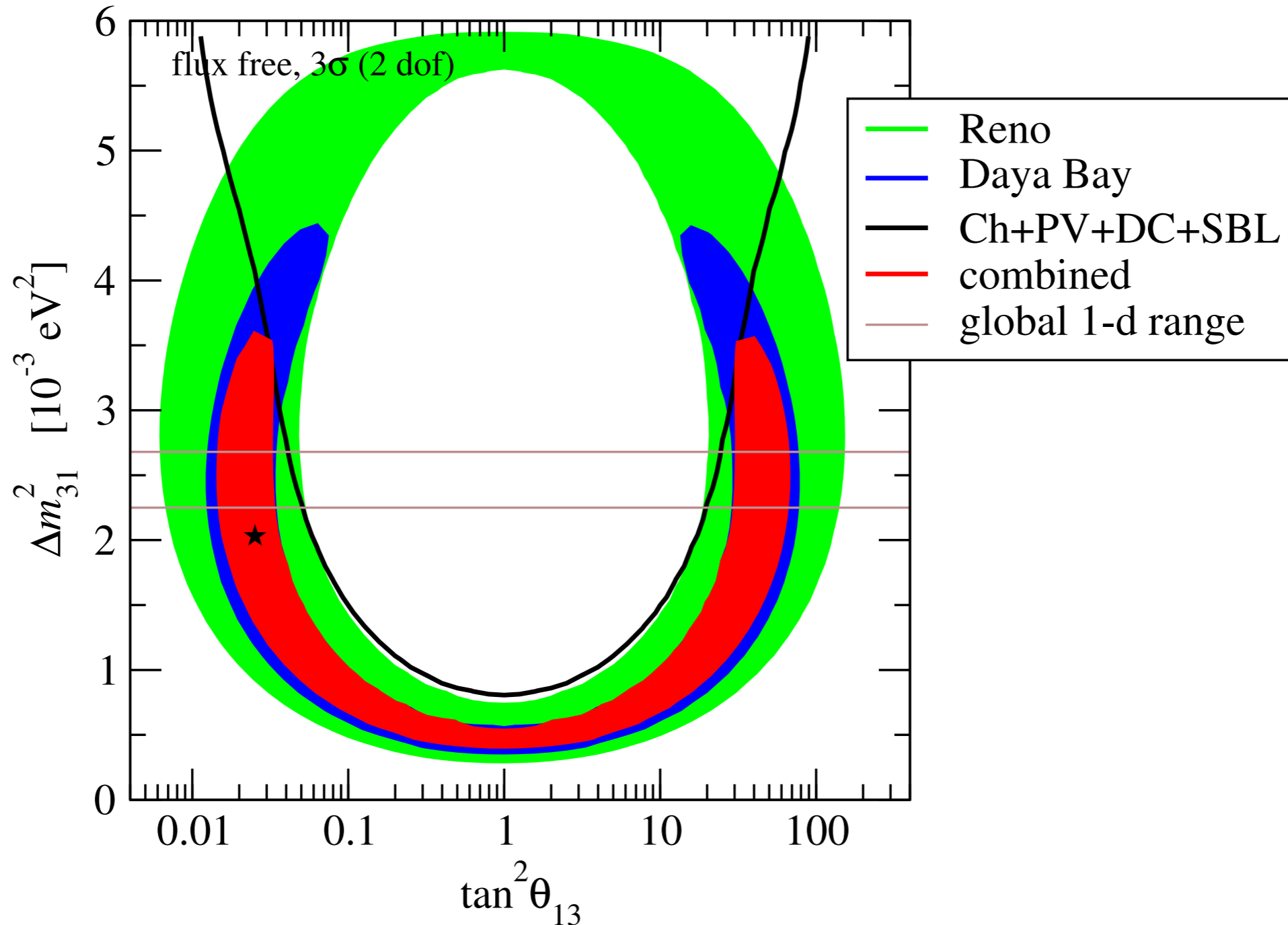
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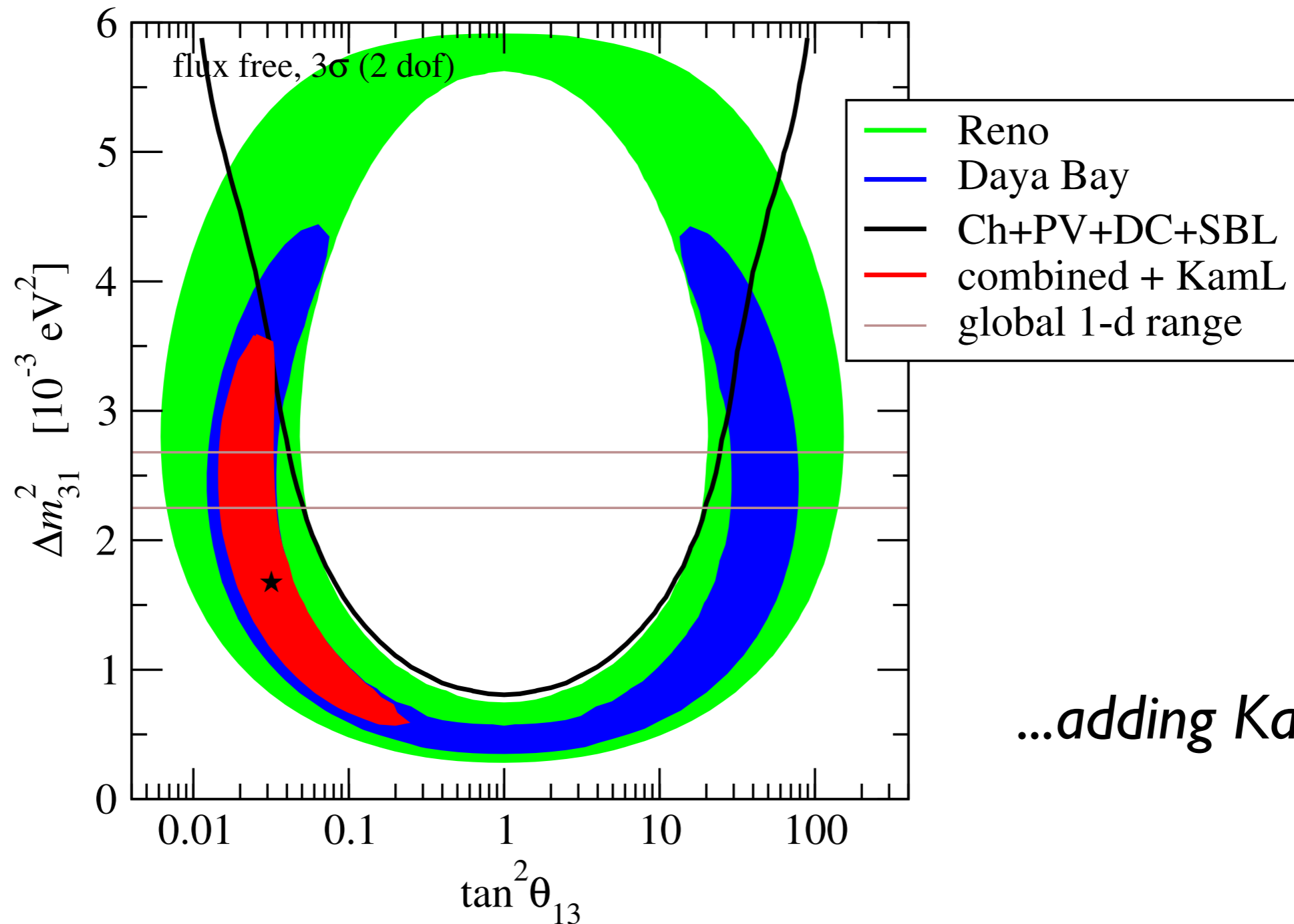
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Measuring Δm^2_{31} with reactors



will improve with spectral data from DayaBay / RENO

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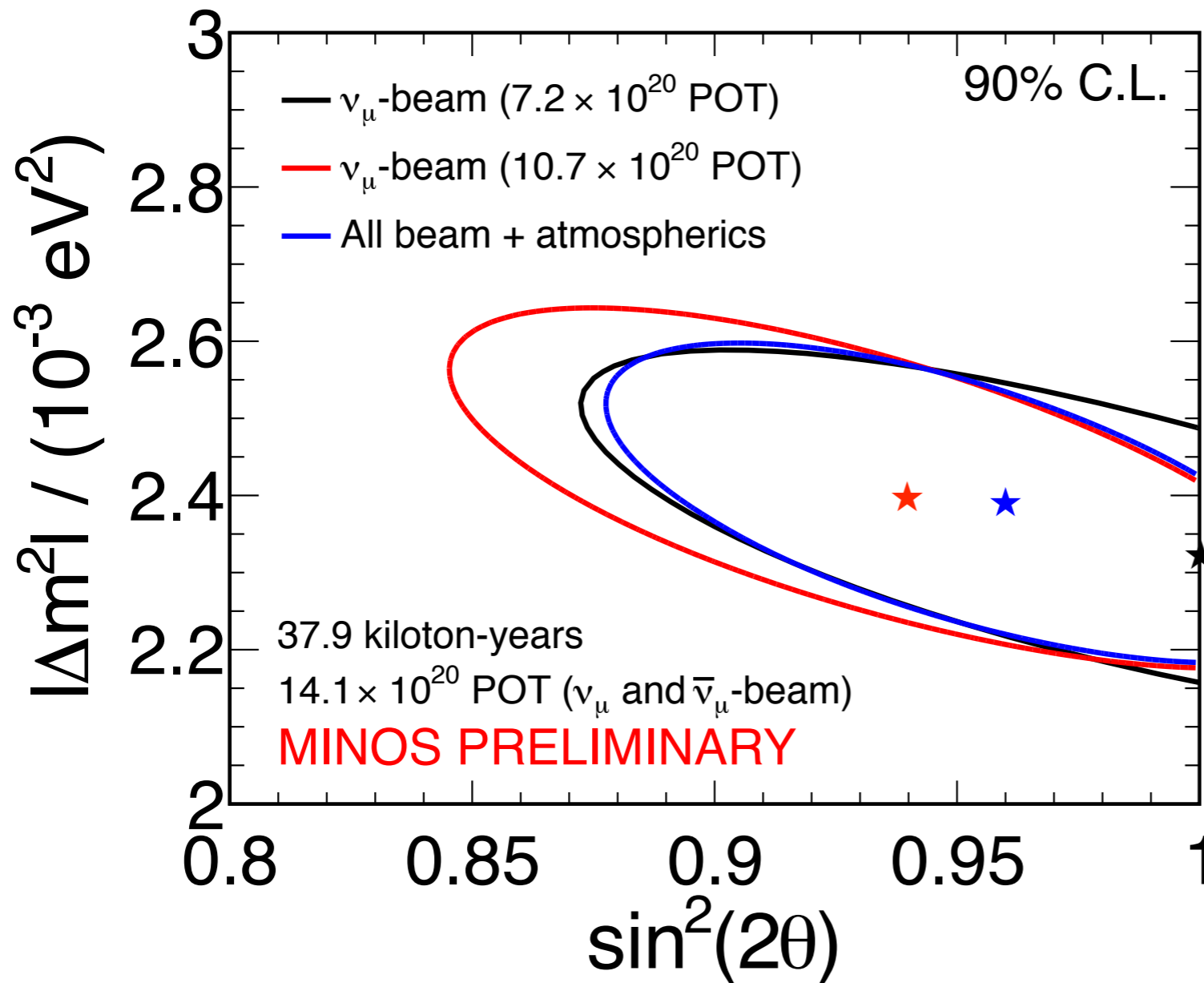


...adding KamLAND

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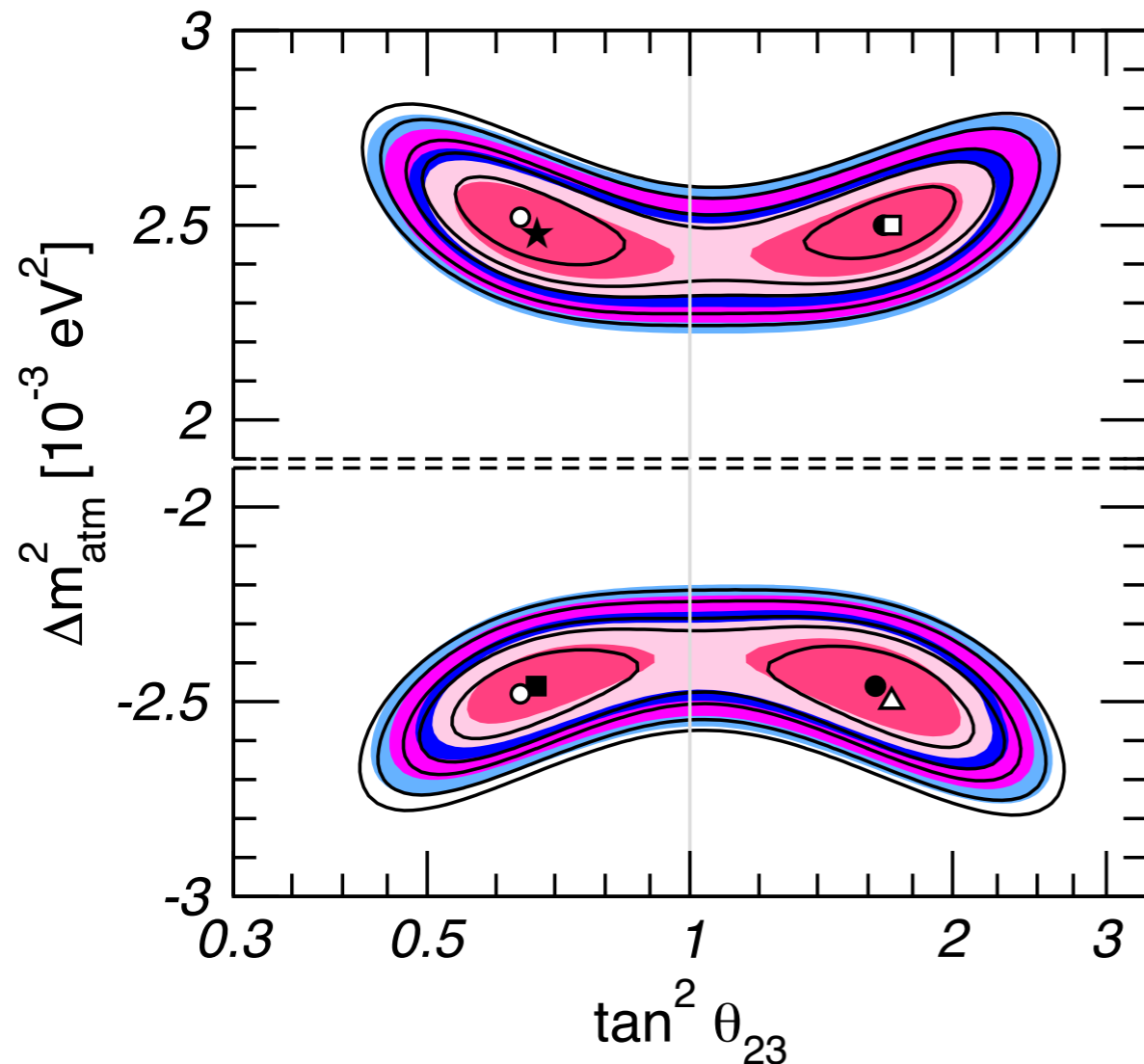
On non-maximal 23 mixing

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Nichol (MINOS), talk
at Neutrino2012

On non-maximal 23 mixing



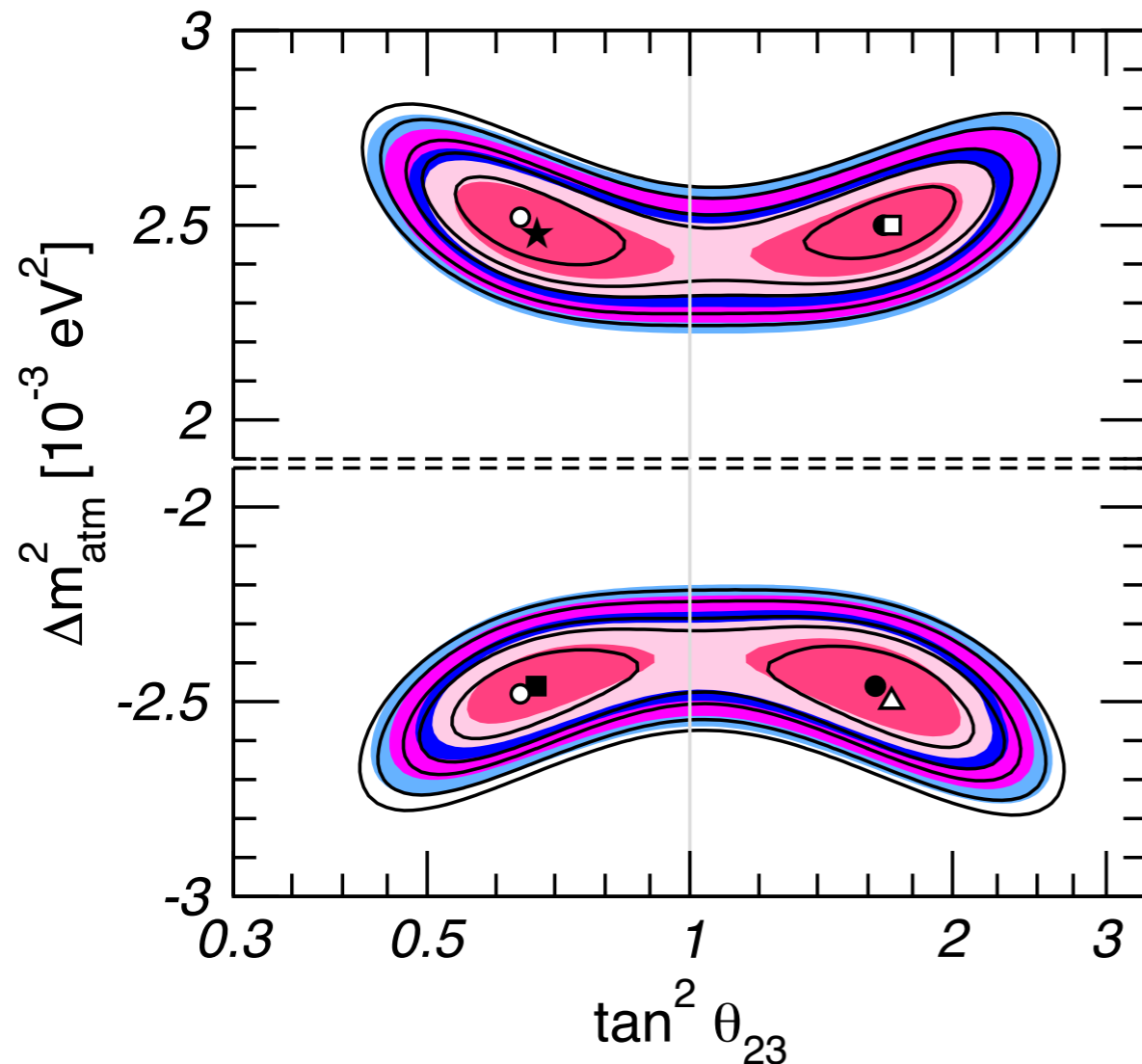
global data without
atmospheric
(MINOS and T2K
disappearance most
important)

$$\sin^2 \theta_{23} \approx 0.40$$

$$\sin^2 \theta_{23} \approx 0.62$$

degeneracy between the two θ_{23} octants

On non-maximal 23 mixing



global data without
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$$\sin^2 \theta_{23} \approx 0.62$$

degeneracy between the two θ_{23} octants

neglecting Δm^2_{21} : $P_{\mu\mu} \approx 1 - 4|U_{\mu 3}|^2(1 - |U_{\mu 3}|^2) \sin^2 \frac{\Delta m^2_{\text{atm}} L}{4E} \Rightarrow \sin^2 \theta_{23} = \frac{|U_{\mu 3}|^2}{\cos^2 \theta_{13}}$

slight shift to larger values of $\sin^2 \theta_{23}$

Octant degeneracy and LBL appearance

Fogli, Lisi, hep-ph/9604415

$$P_{\mu e} \simeq \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2(1-A)\Delta}{(1-A)^2} \\ + \sin 2\theta_{13} \hat{\alpha} \sin 2\theta_{23} \frac{\sin(1-A)\Delta}{1-A} \frac{\sin A\Delta}{A} \cos(\Delta + \delta_{\text{CP}}) \\ + \hat{\alpha}^2 \cos^2 \theta_{23} \frac{\sin^2 A\Delta}{A^2}$$

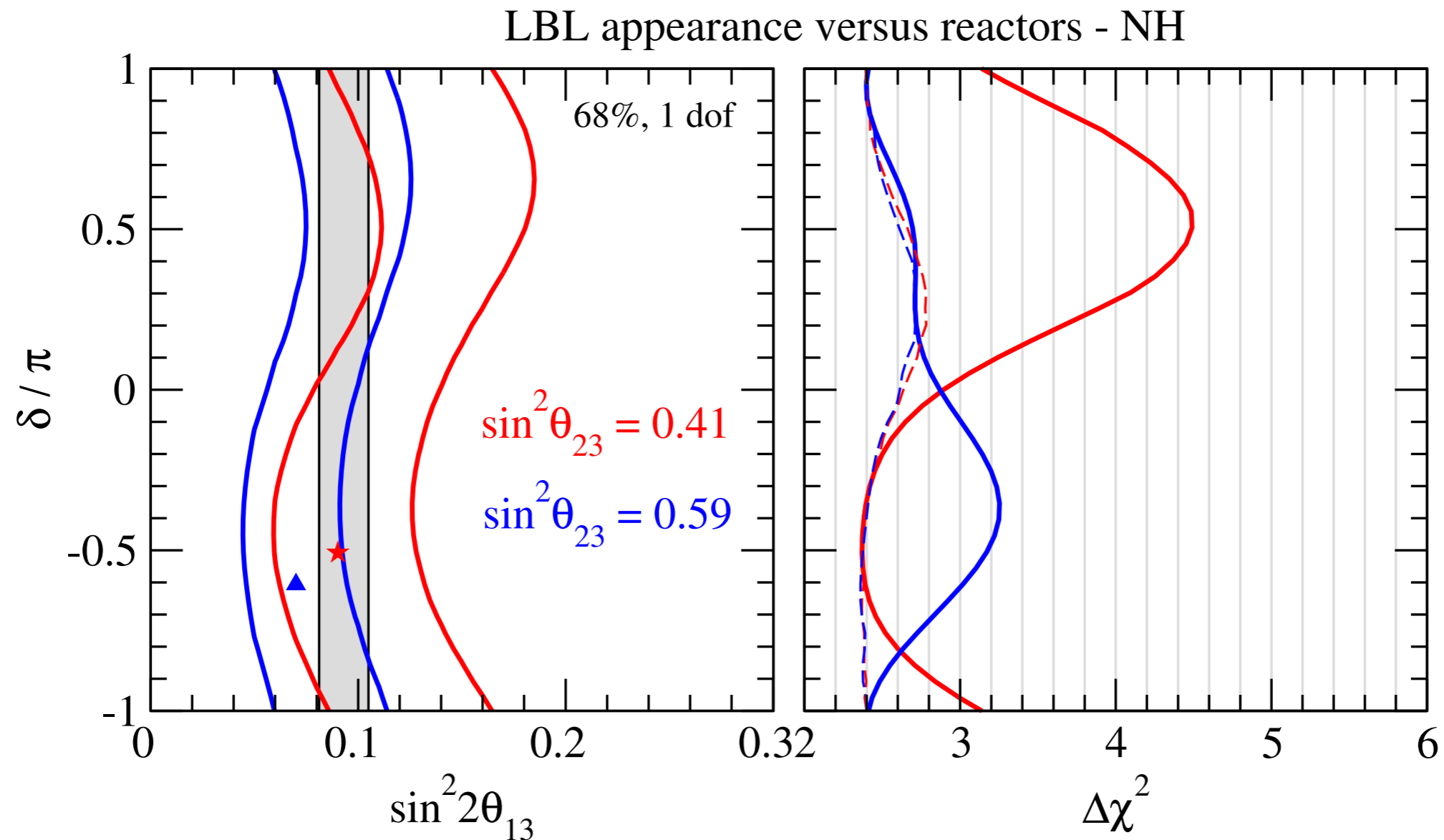
with

$$\Delta \equiv \frac{\Delta m_{31}^2 L}{4E_\nu}, \quad \hat{\alpha} \equiv \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \sin 2\theta_{12}, \quad A \equiv \frac{2E_\nu V}{\Delta m_{31}^2}$$

- for large θ_{13} the leading term depends on octant
- beam+reactor combination may be sensitive to octant

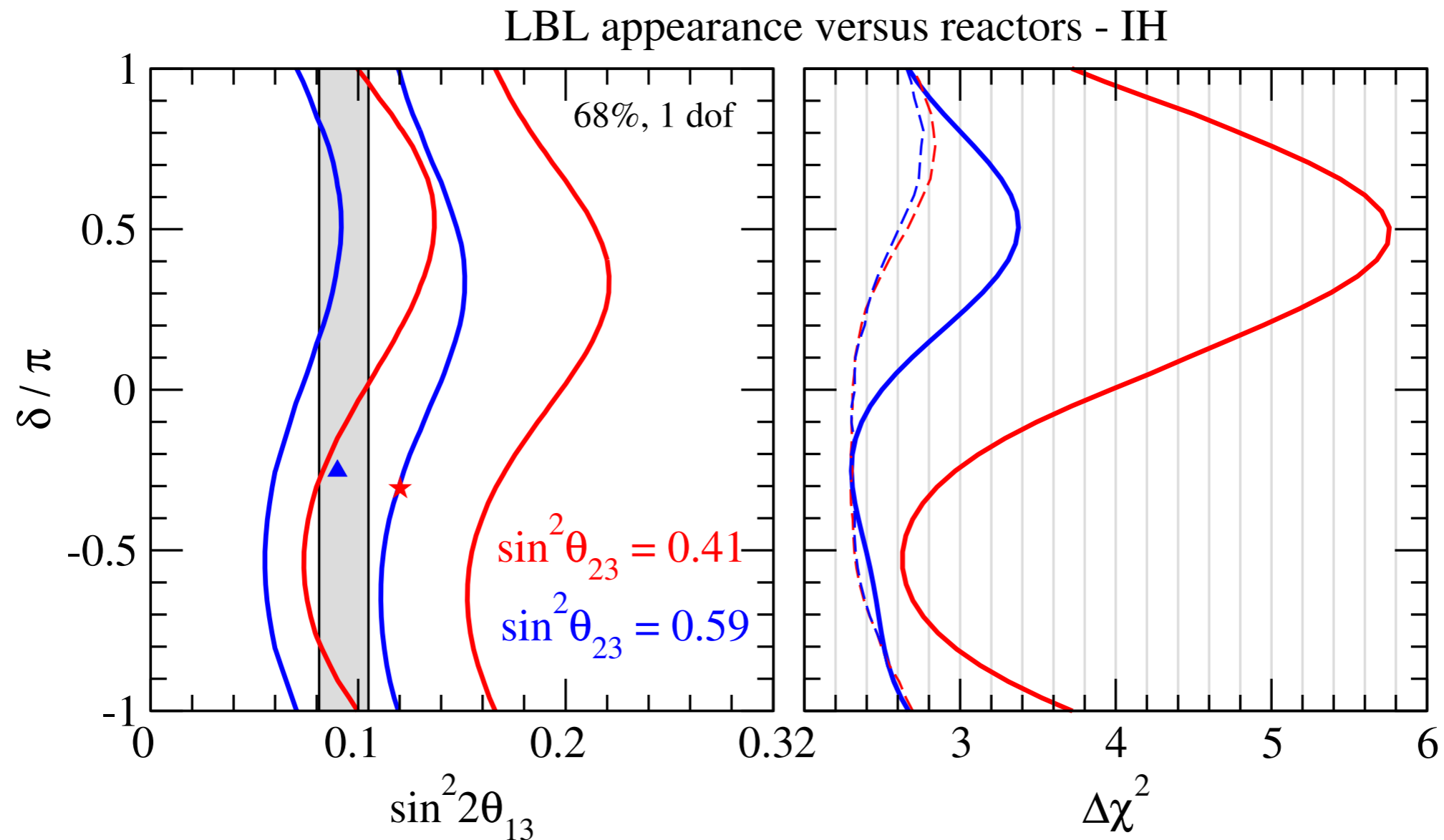
Minakata et al. hep-ph/0211111; McConnel, Shaevitz, hep-ex/0409028

Octant degeneracy and LBL appearance



present data from LBL appearance versus reactor cannot discriminate between the octants

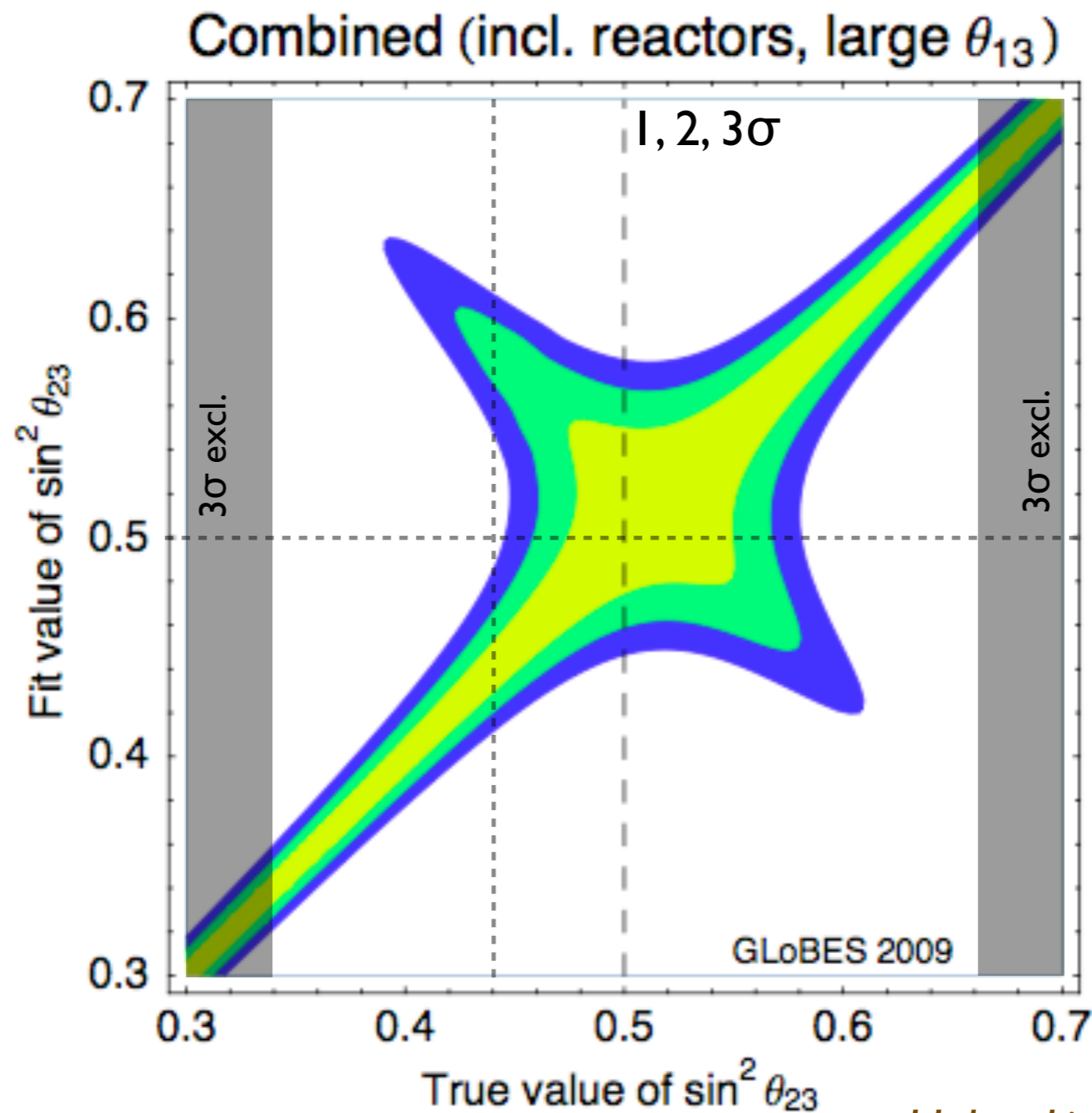
Octant degeneracy and LBL appearance



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Global fit ~ 2020 - θ_{23} octant

final exposure of T2K, NOvA, DayaBay combined



$$\sin^2 2\theta_{13} = 0.1$$
$$\delta = 0$$

Huber, Lindner, TS, Winter, 0907.1896

3-flavor effects in atmospheric neutrinos

excess in electron-like events:

$$\begin{aligned} \frac{N_e}{N_e^0} - 1 &\simeq (r s_{23}^2 - 1) P_{2\nu}(\Delta m_{31}^2, \theta_{13}) && \theta_{13}\text{-effects} \\ &+ (r c_{23}^2 - 1) P_{2\nu}(\Delta m_{21}^2, \theta_{12}) && \Delta m_{21}^2\text{-effects} \\ &- 2s_{13}s_{23}c_{23} r \operatorname{Re}(A_{ee}^* A_{\mu e}) && \text{interference: } \delta_{\text{CP}} \end{aligned}$$

$$r = r(E_\nu) \equiv \frac{F_\mu^0(E_\nu)}{F_e^0(E_\nu)} \quad \begin{array}{l} r \approx 2 \quad (\text{sub-GeV}) \\ r \approx 2.6 - 4.5 \quad (\text{multi-GeV}) \end{array}$$

3-flavor effects in atmospheric neutrinos

Peres, Smirnov, 99;
Gonzalez-Garcia, Maltoni, Smirnov, 04

excess in electron-like events:

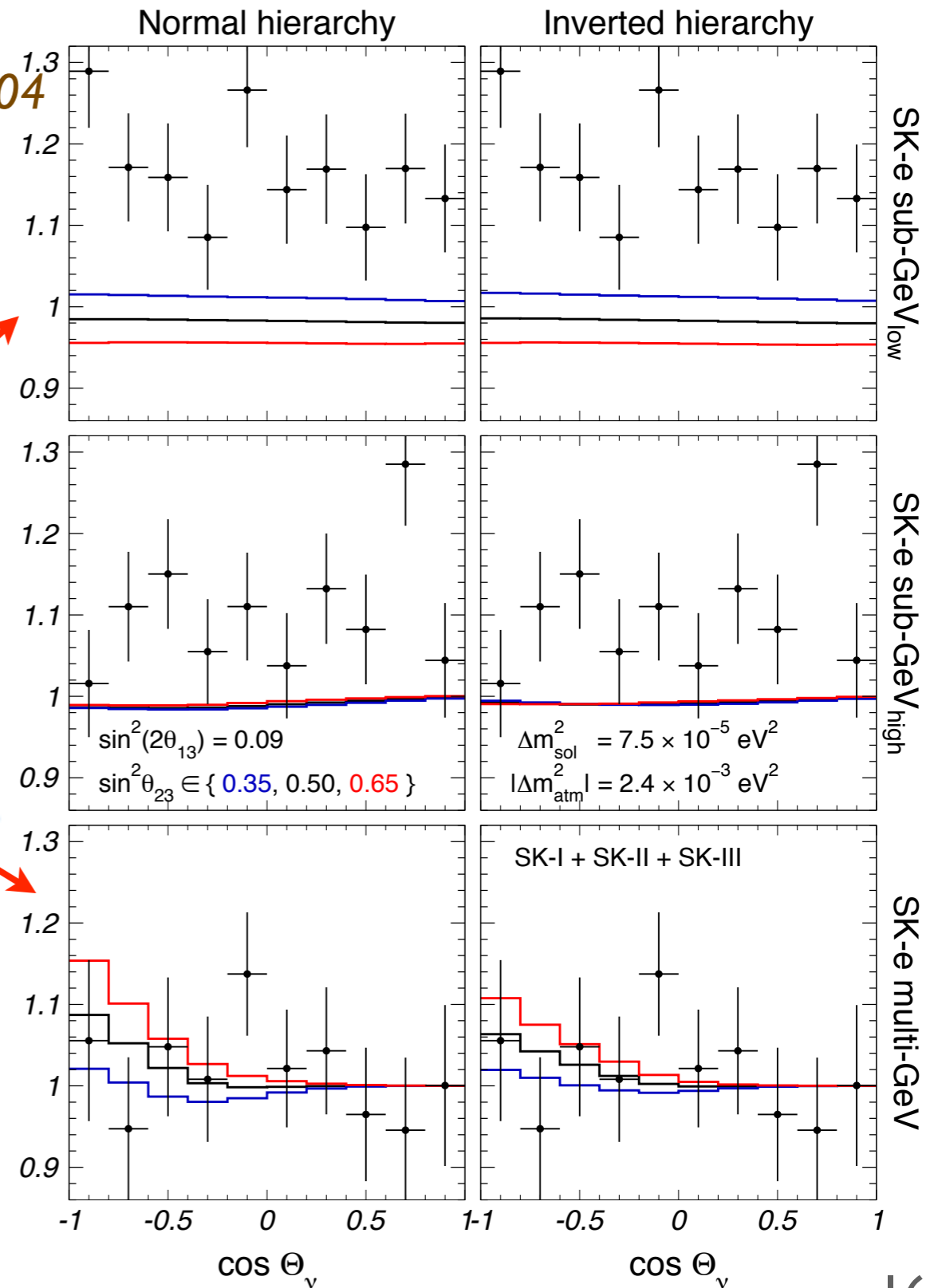
$$\frac{N_e}{N_e^0} - 1 \simeq (r s_{23}^2 - 1) P_{2\nu}(\Delta m_{31}^2, \theta_{13}) \quad \theta_{13}\text{-effects}$$

$$+ (r c_{23}^2 - 1) P_{2\nu}(\Delta m_{21}^2, \theta_{12}) \quad \Delta m_{21}^2\text{-effects}$$

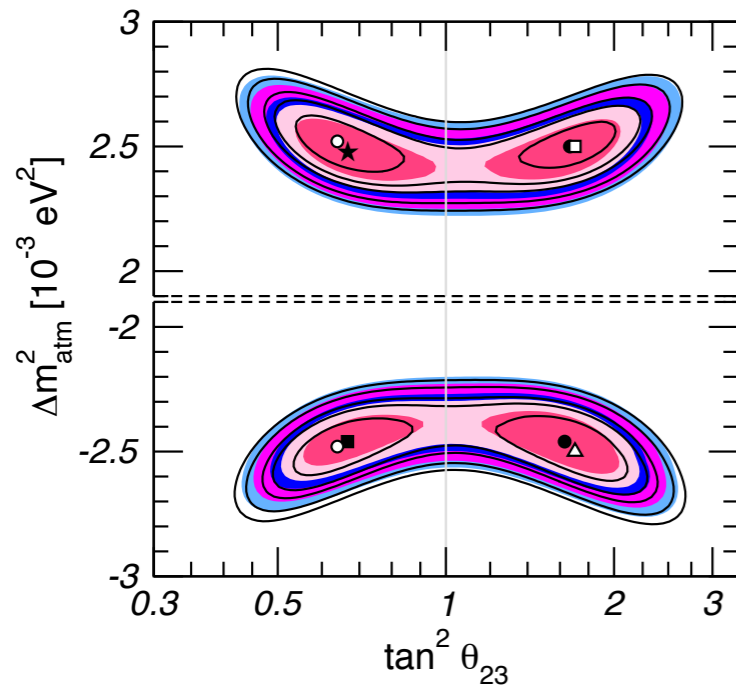
$$- 2s_{13}s_{23}c_{23} r \operatorname{Re}(A_{ee}^* A_{\mu e}) \quad \text{interference: } \delta_{CP}$$

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The octant and atmospheric neutrino data

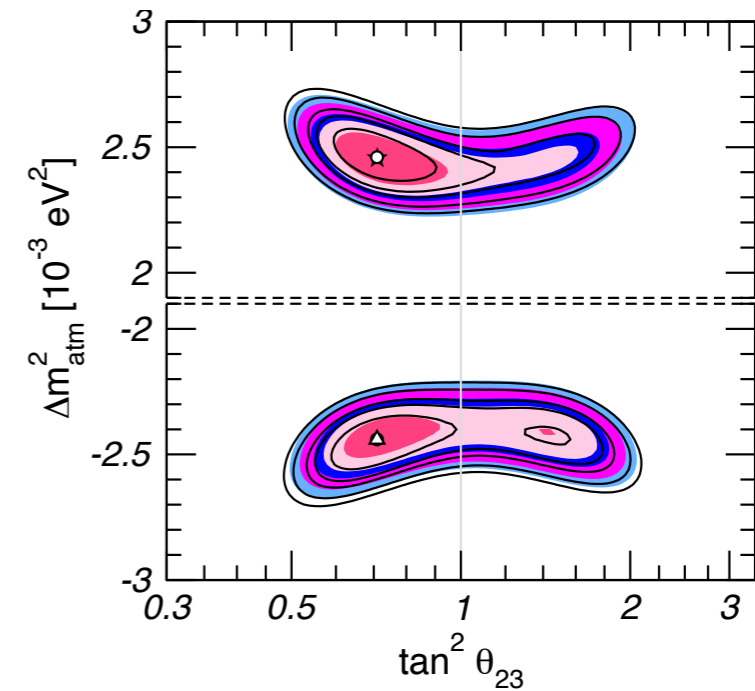


adding atmospheric

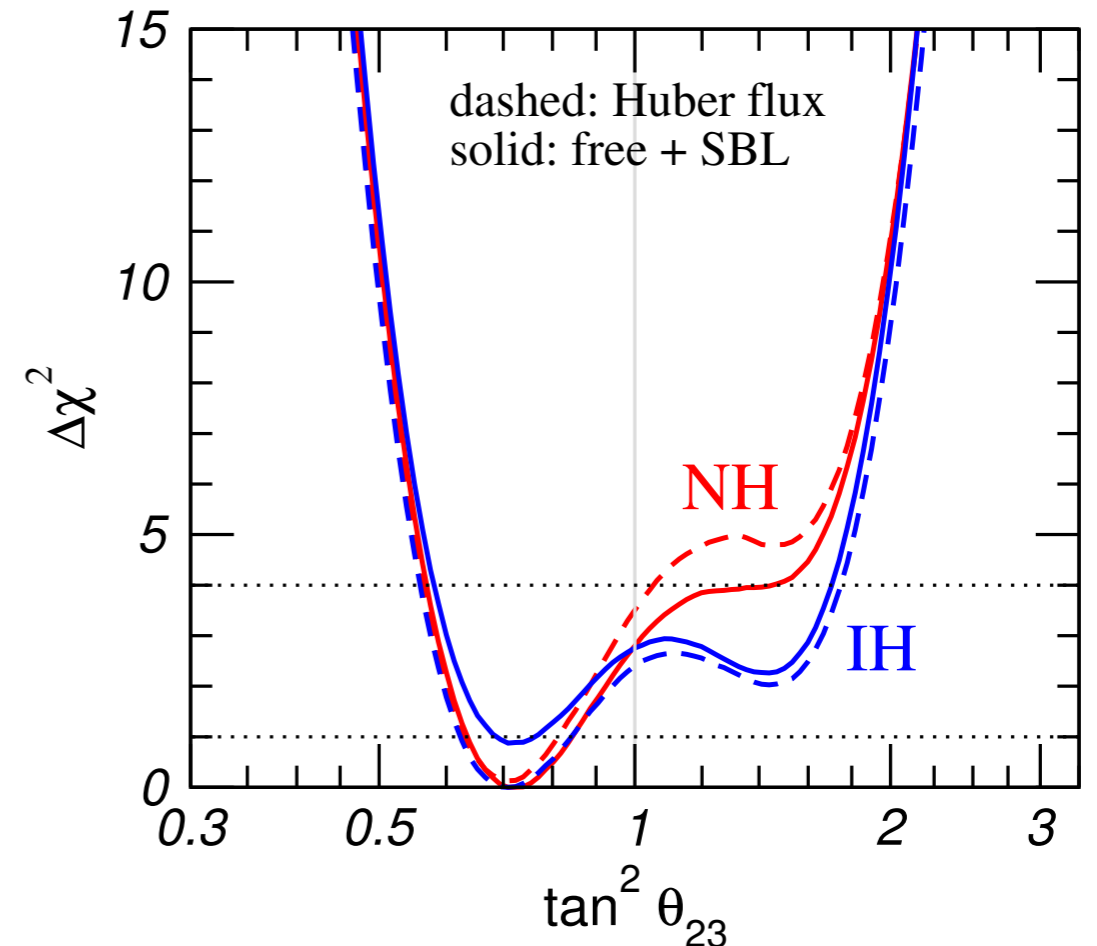
$$\sin^2 \theta_{23} = 0.42^{+0.037}_{-0.031}$$

$$0.34 - 0.66 @ 3\sigma$$

$$(\theta_{23} \approx 40^\circ)$$

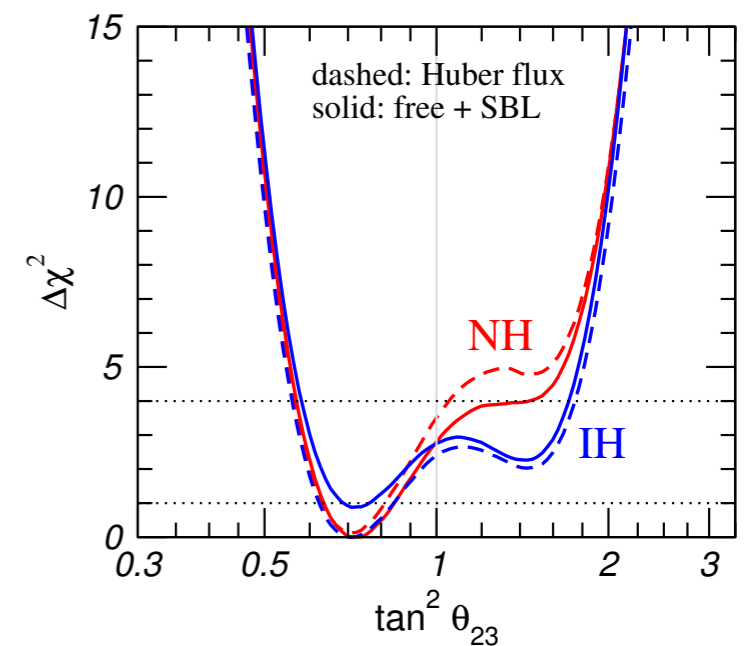
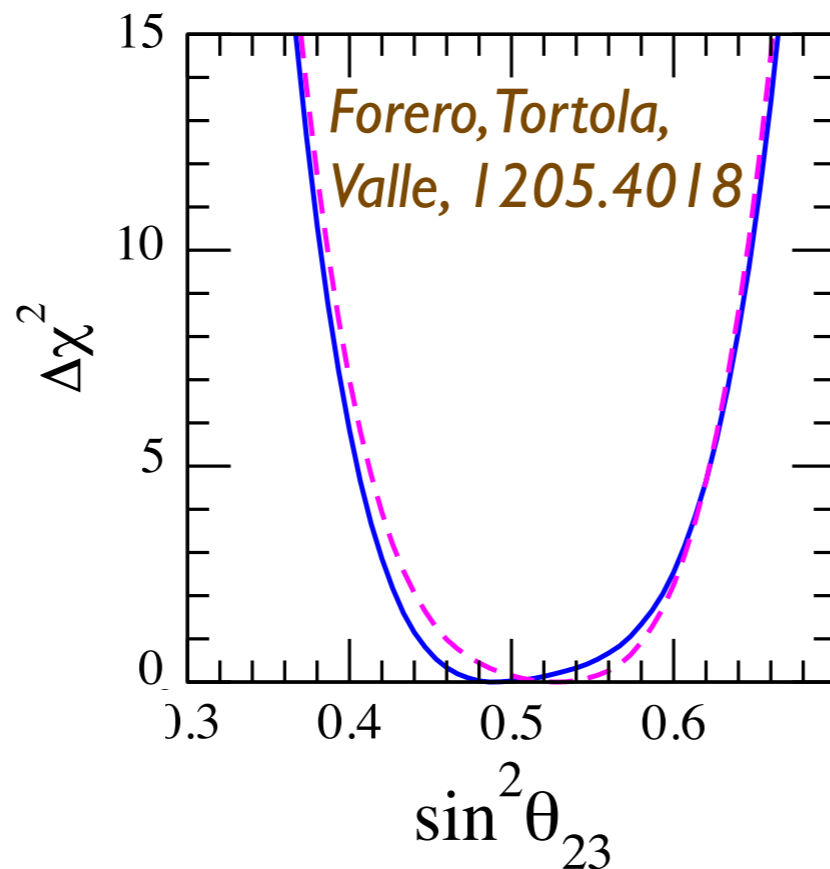
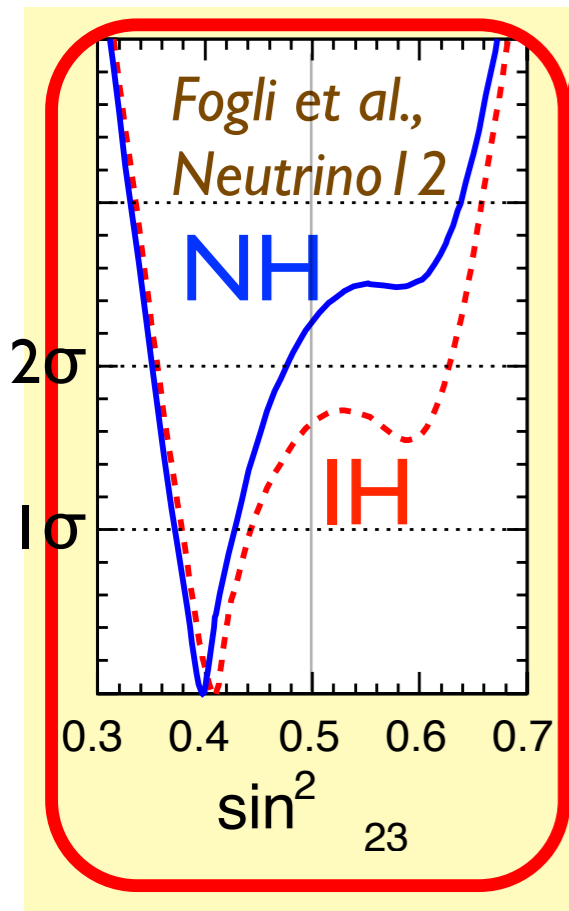
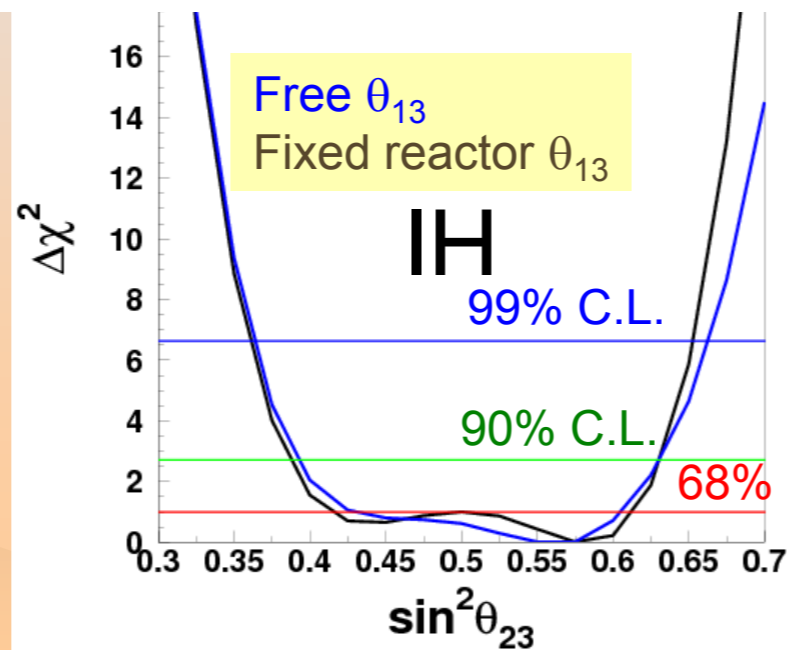
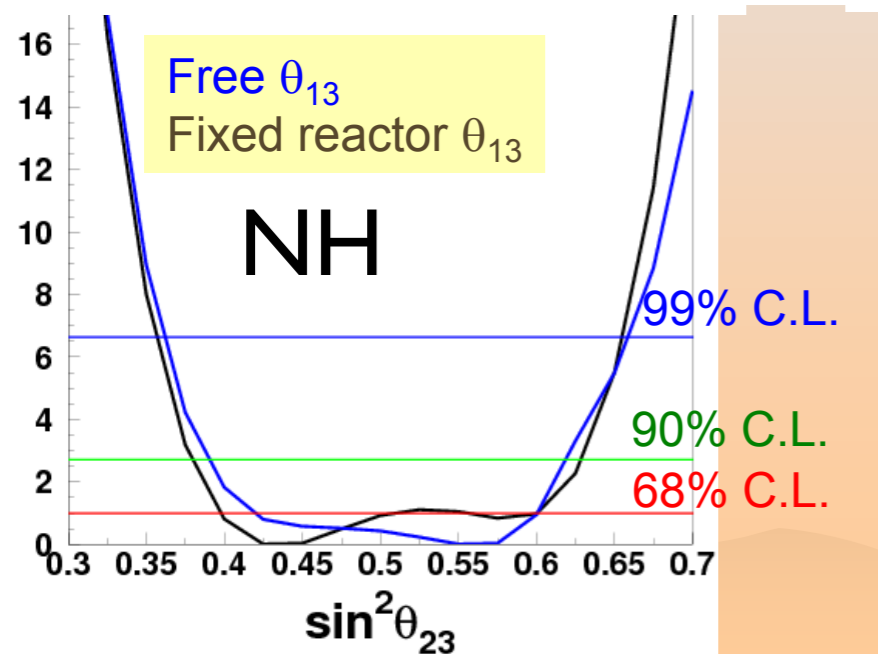


- slight preference for 1st octant by atmospheric data
- more pronounced for NH ($\sim 2\sigma$)
- depends on θ_{13}



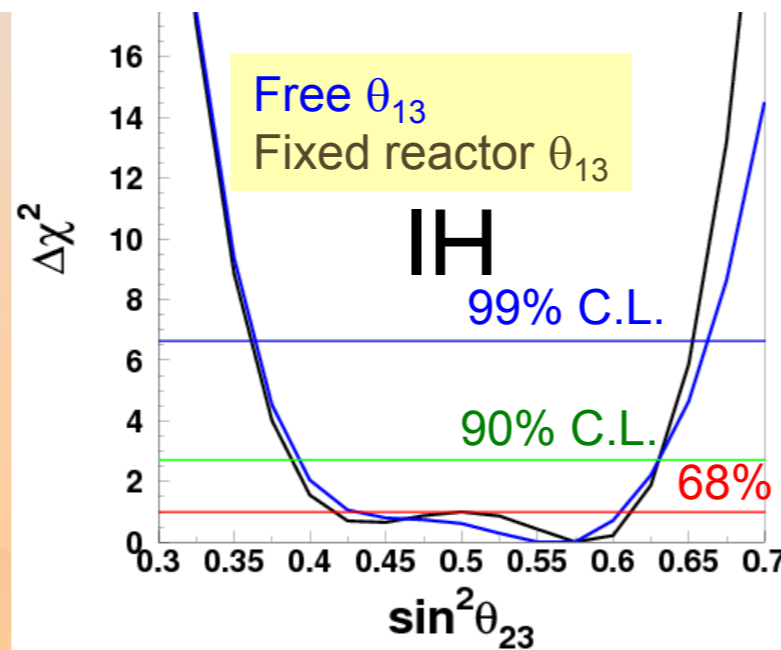
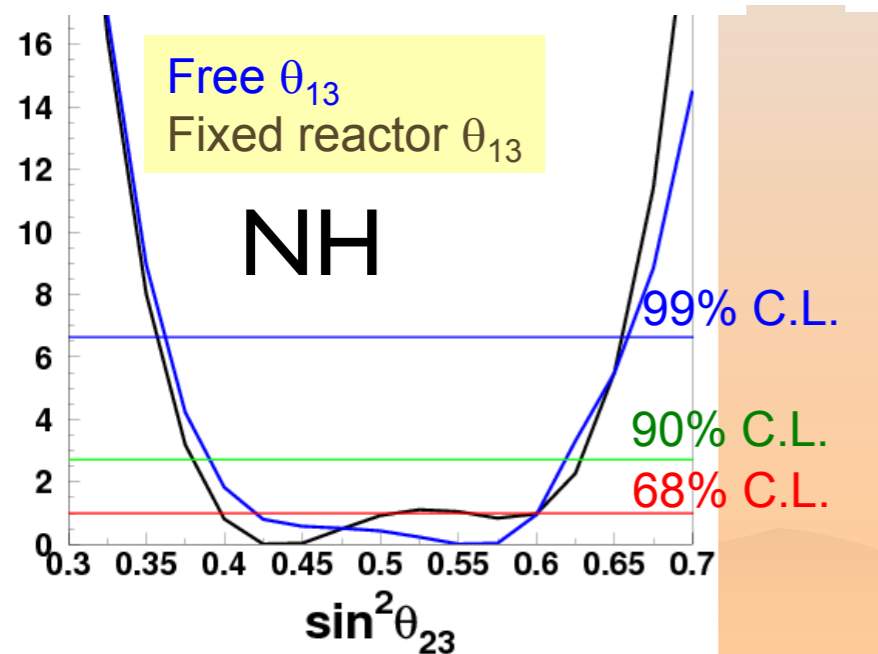
A word of warning

Itow (SuperK), talk at Neutrino2012

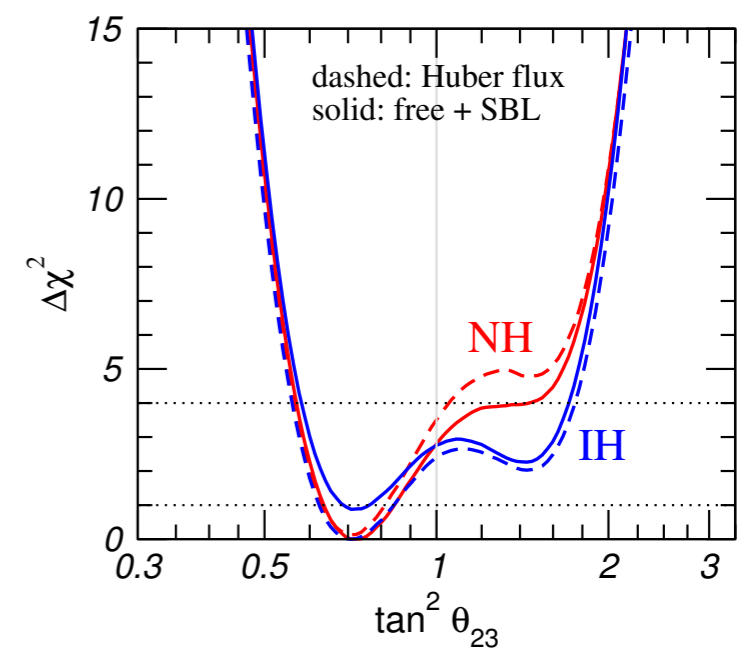
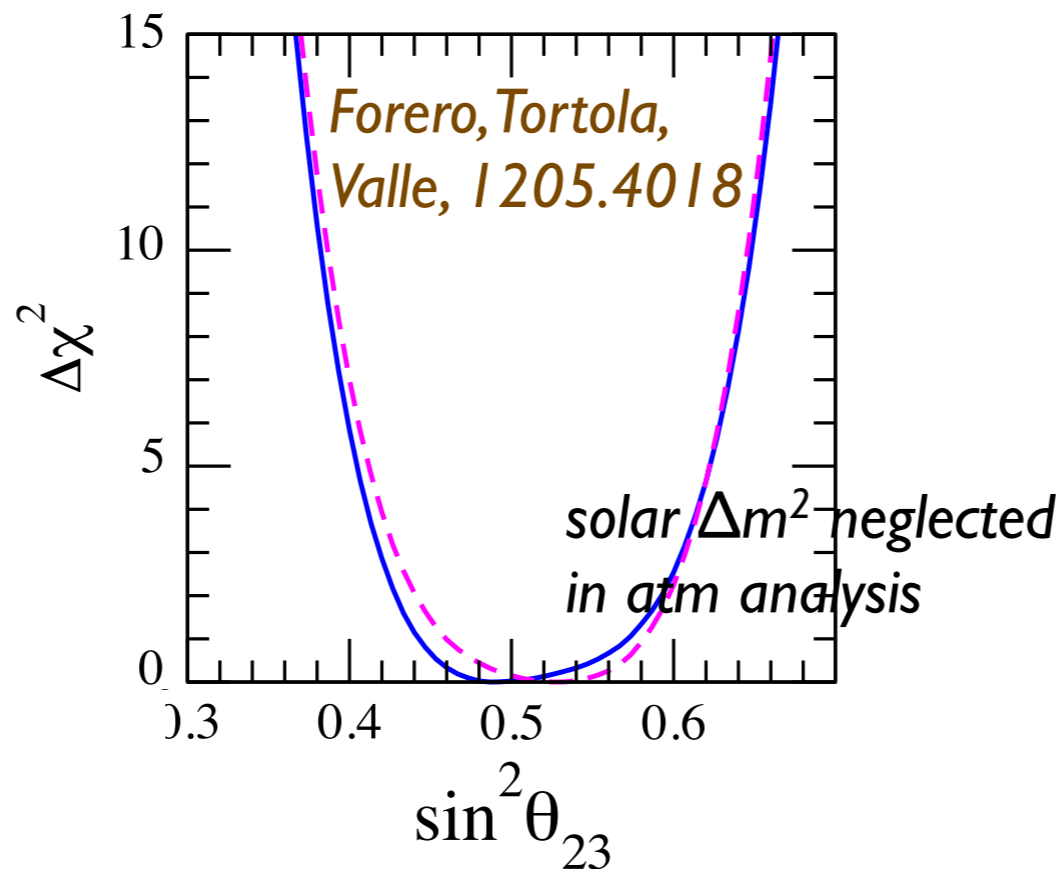
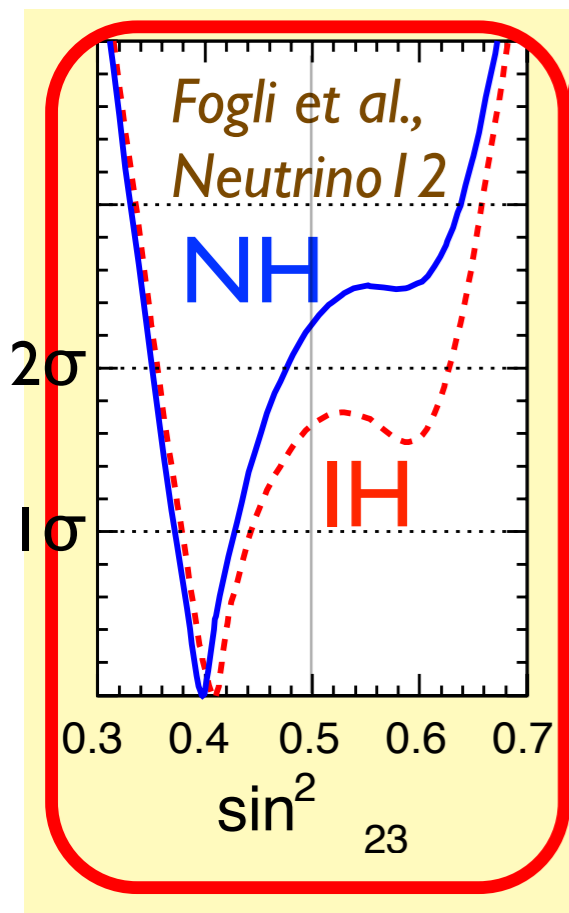


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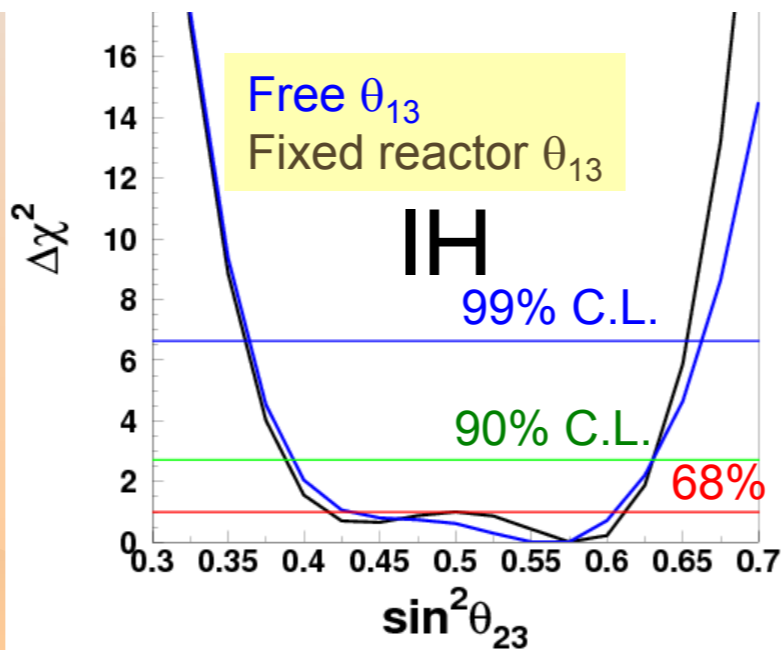
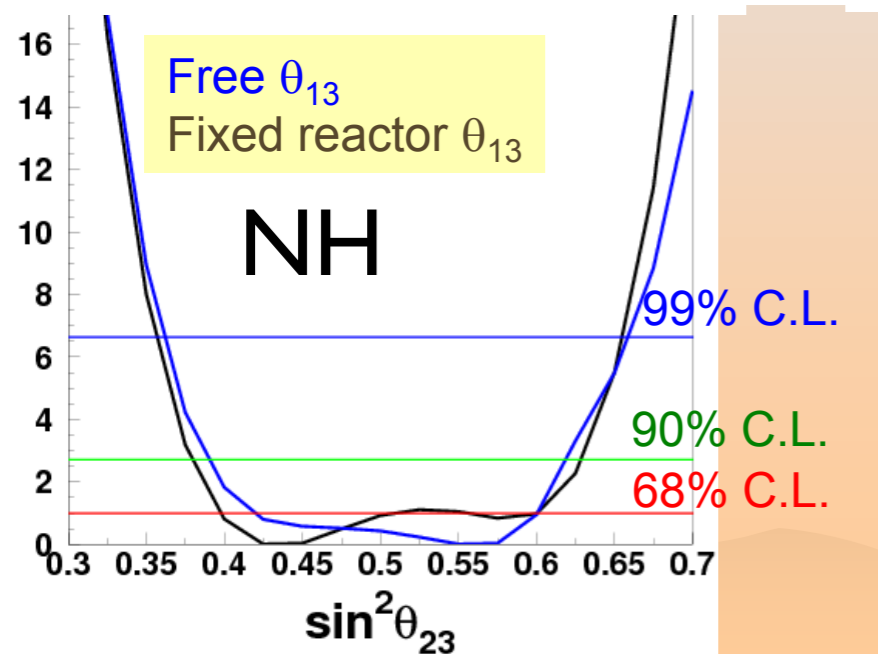
Itow (SuperK), talk at Neutrino2012



MINOS data from Neutrino2012 not included yet



A word of warning



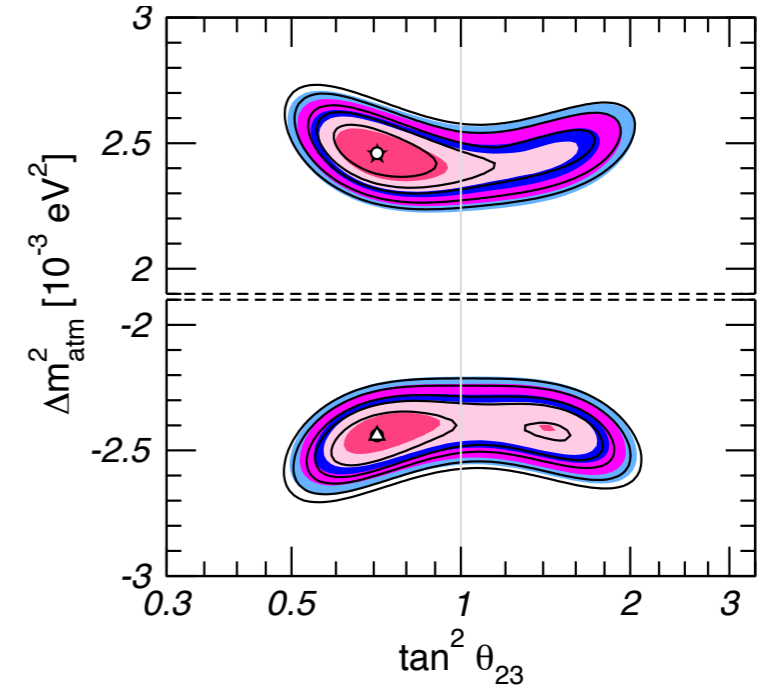
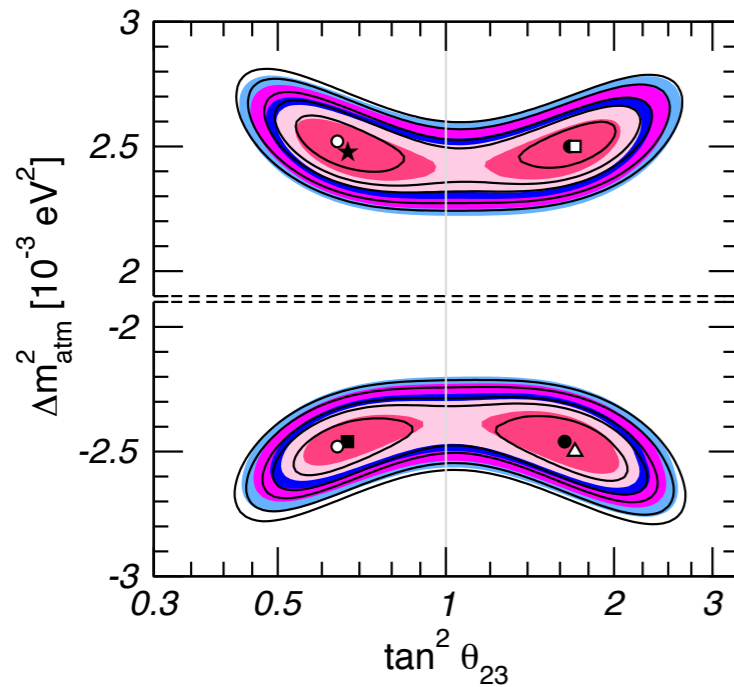
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personal interpretation:

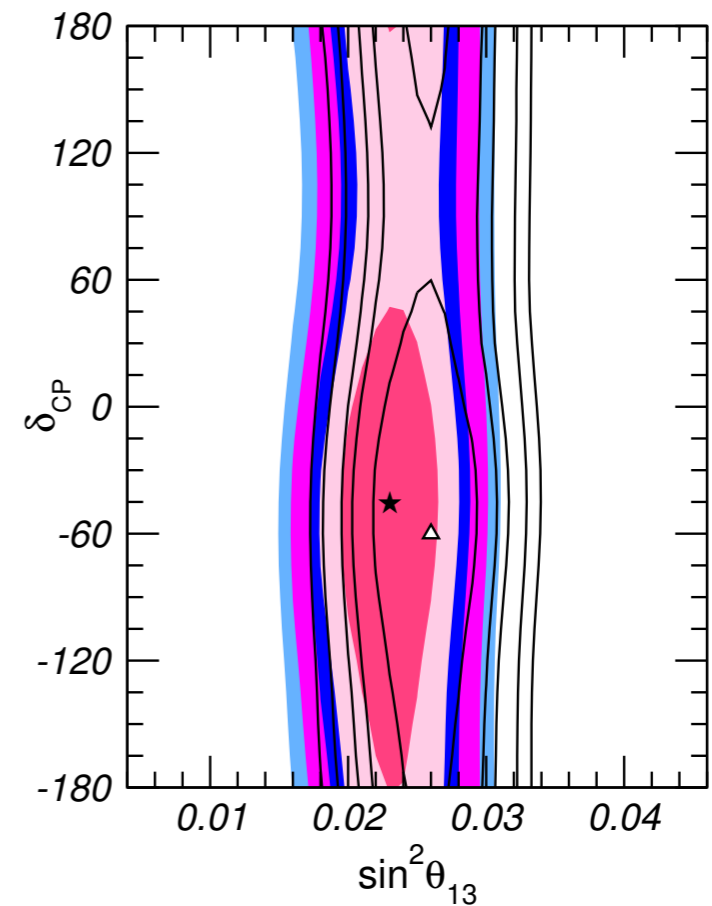
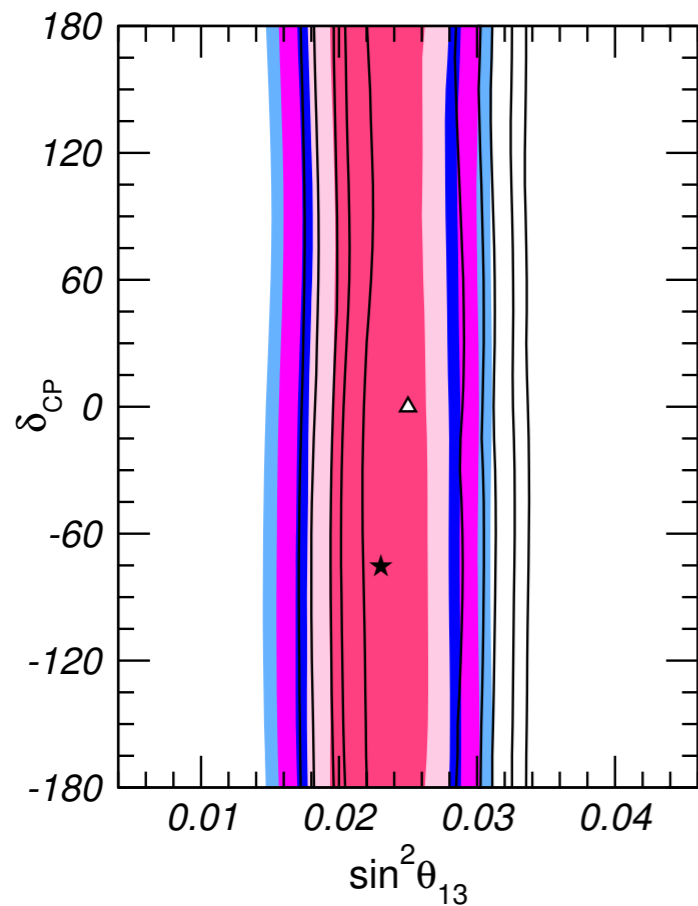
- there seem to be subtle effects in SK atm data pulling into different directions
- sub-GeV data vs multi-GeV (?) data (θ_{13} / hierarchy dependent)
- which effect wins depends on fine details of the analysis / system. treatment

before drawing definite conclusions it is mandatory to identify the physics and understand the origin of the different results from SK and phenomenologists

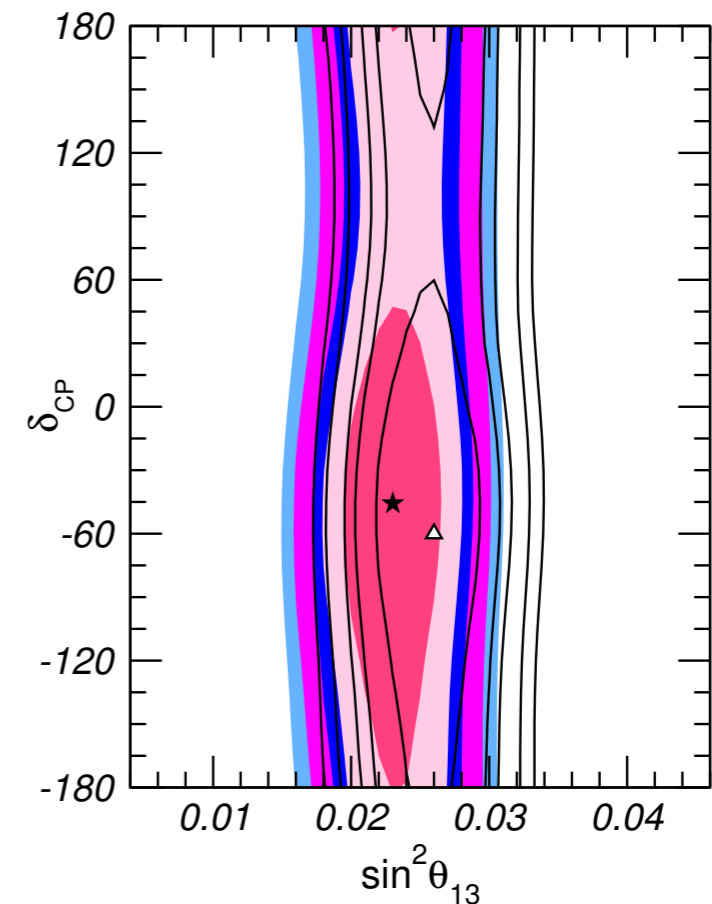
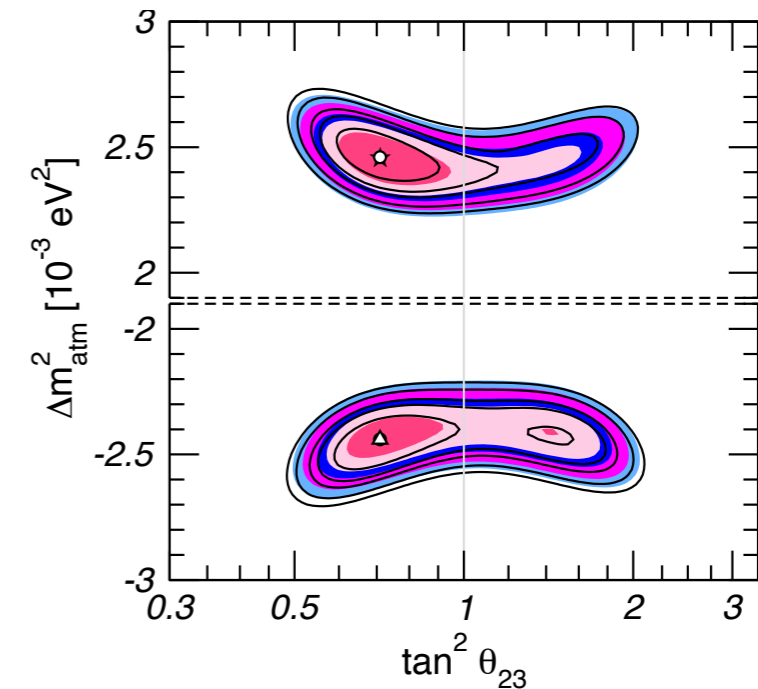
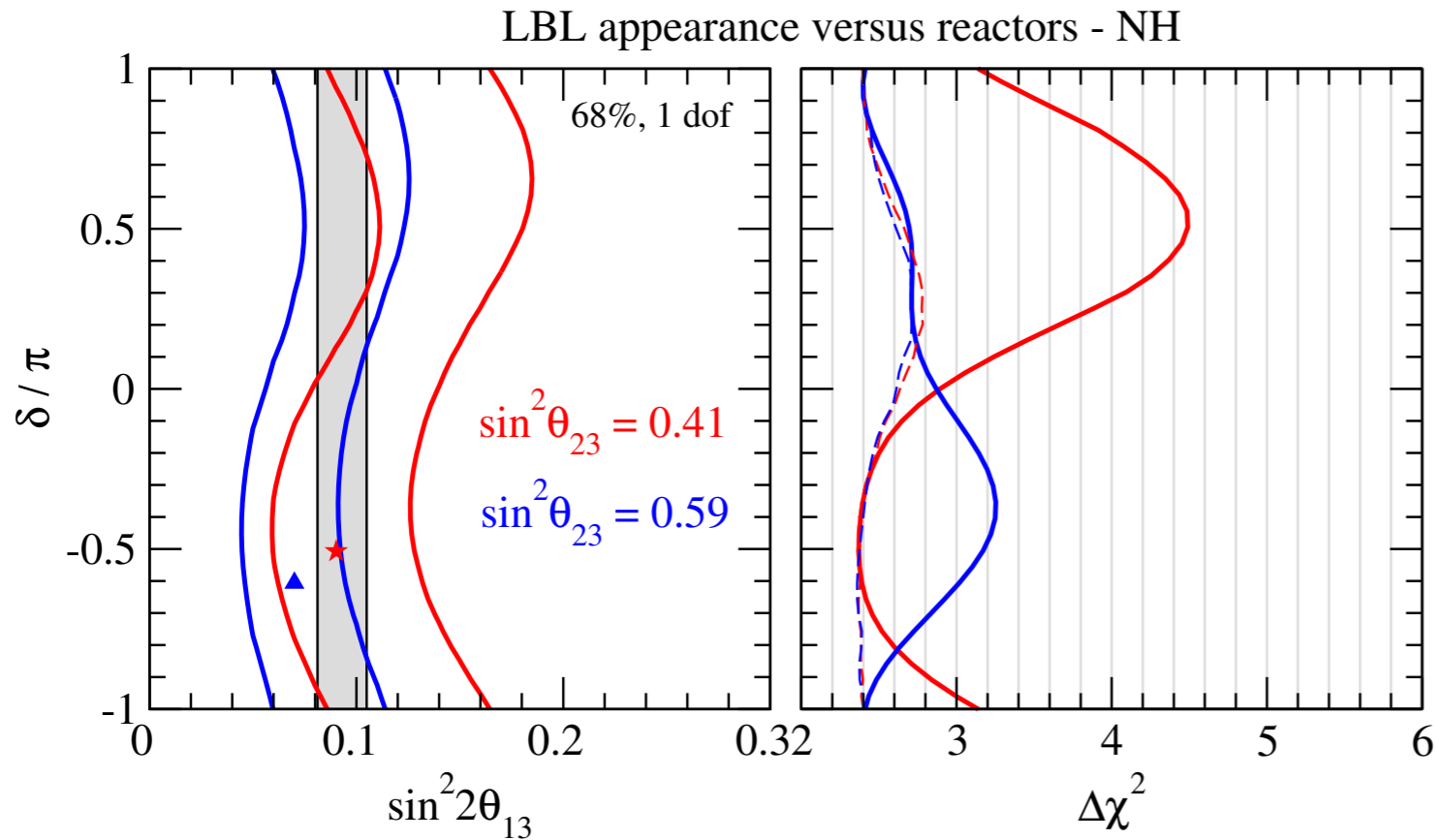
The CP phase and atmospheric neutrino data



adding atmospheric

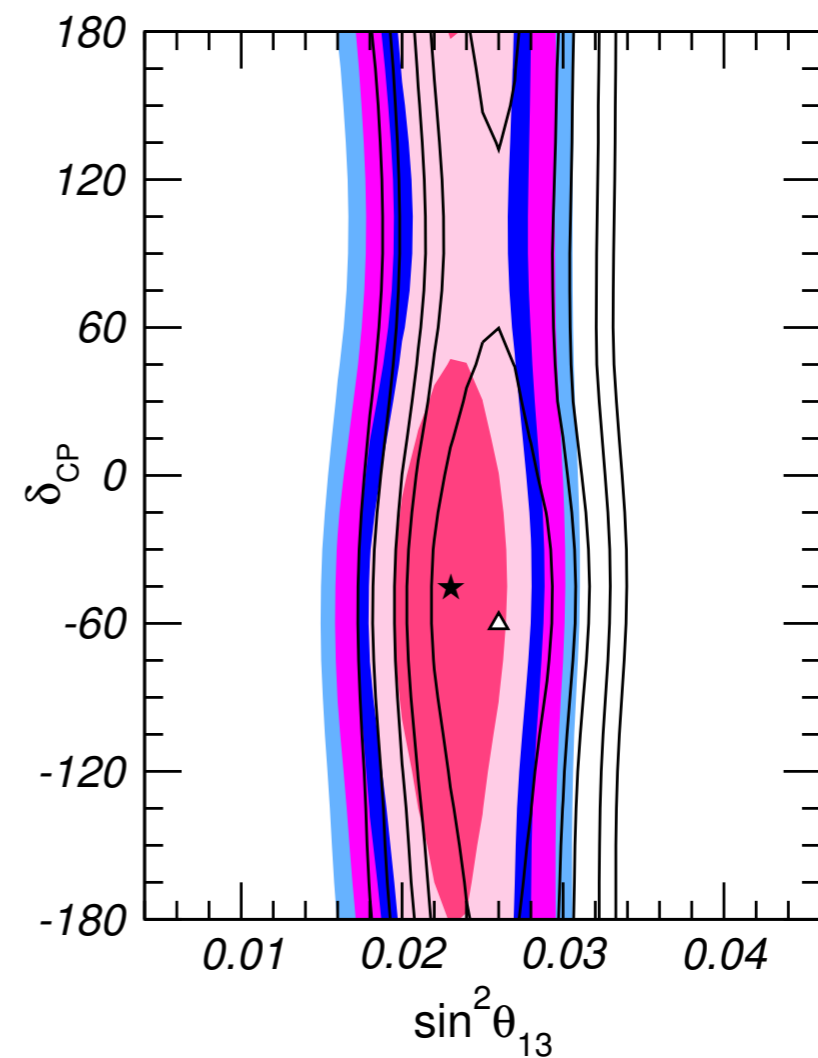
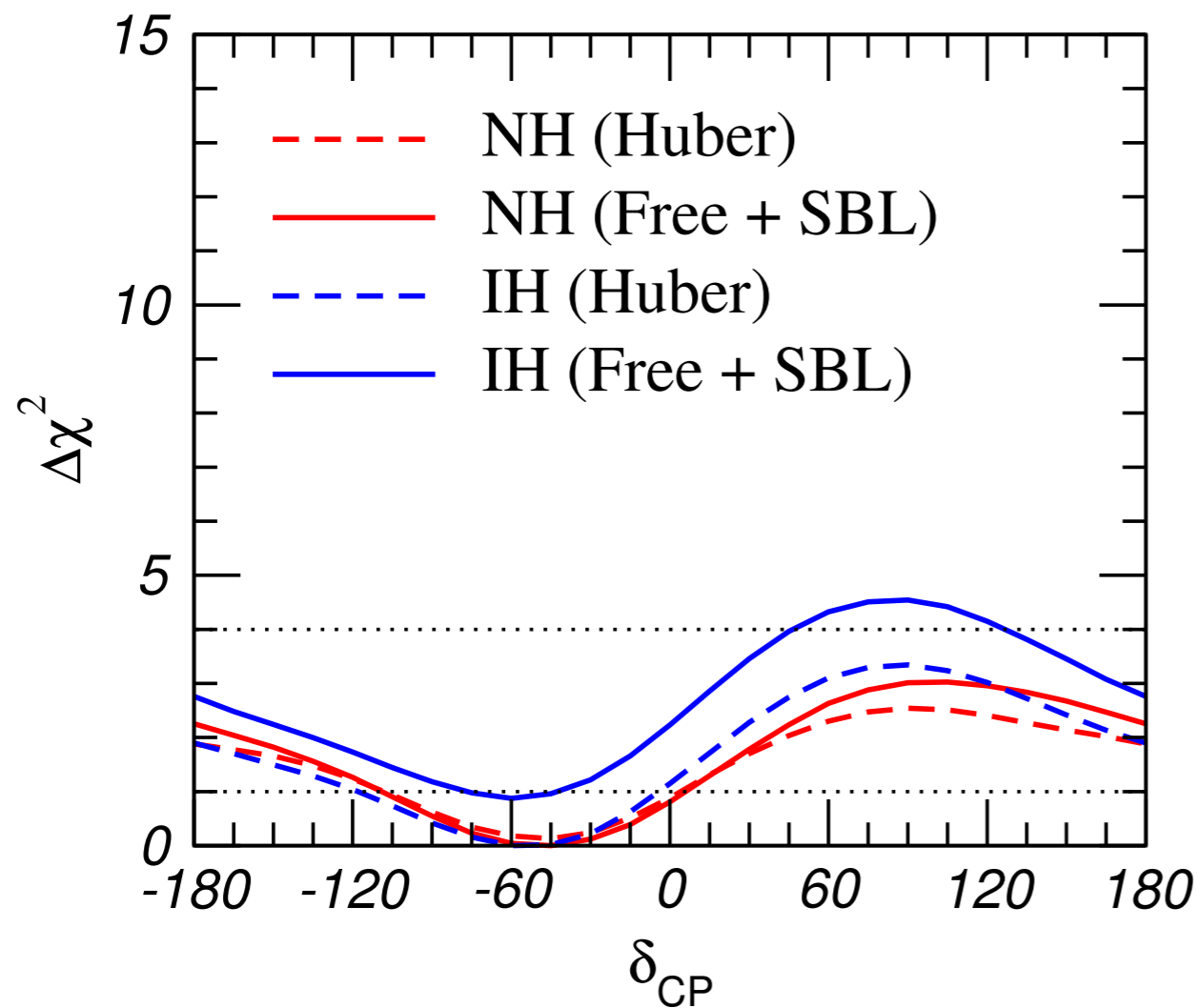


The CP phase and atmospheric neutrino data



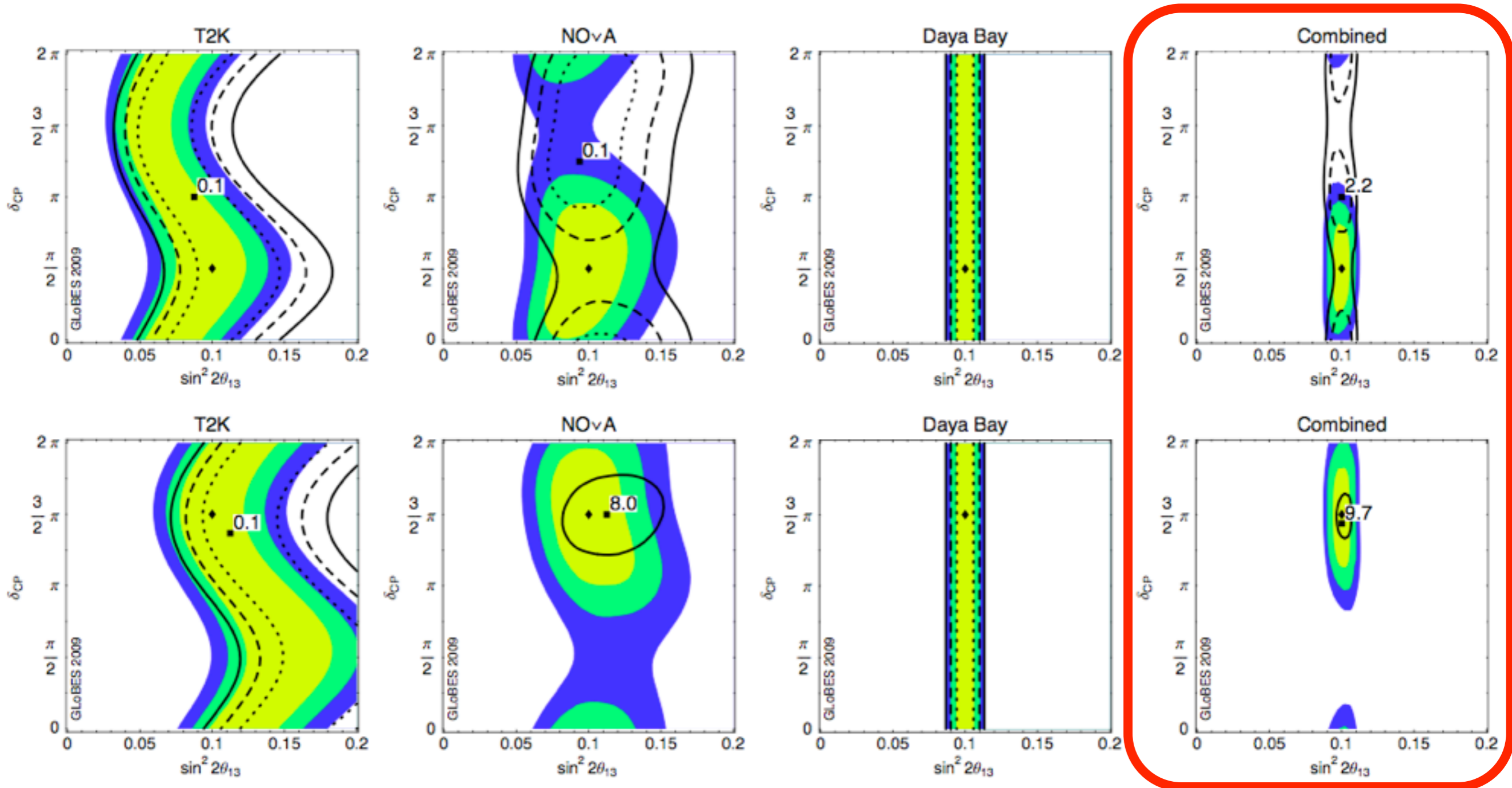
- ➔ non-maximality from LBL disapp
- ➔ preference for 1st octant from atm
- ➔ sensitivity to δ from LBL appearance + reactors

The CP phase



- “preferred” regions for $\delta \sim -60^\circ$ at 1σ (everything allowed at 2σ)
- difference between NH and IH of $\Delta\chi^2 \approx 1$

Global fit ~ 2020 Huber, Lindner, TS, Winter, 0907.1896

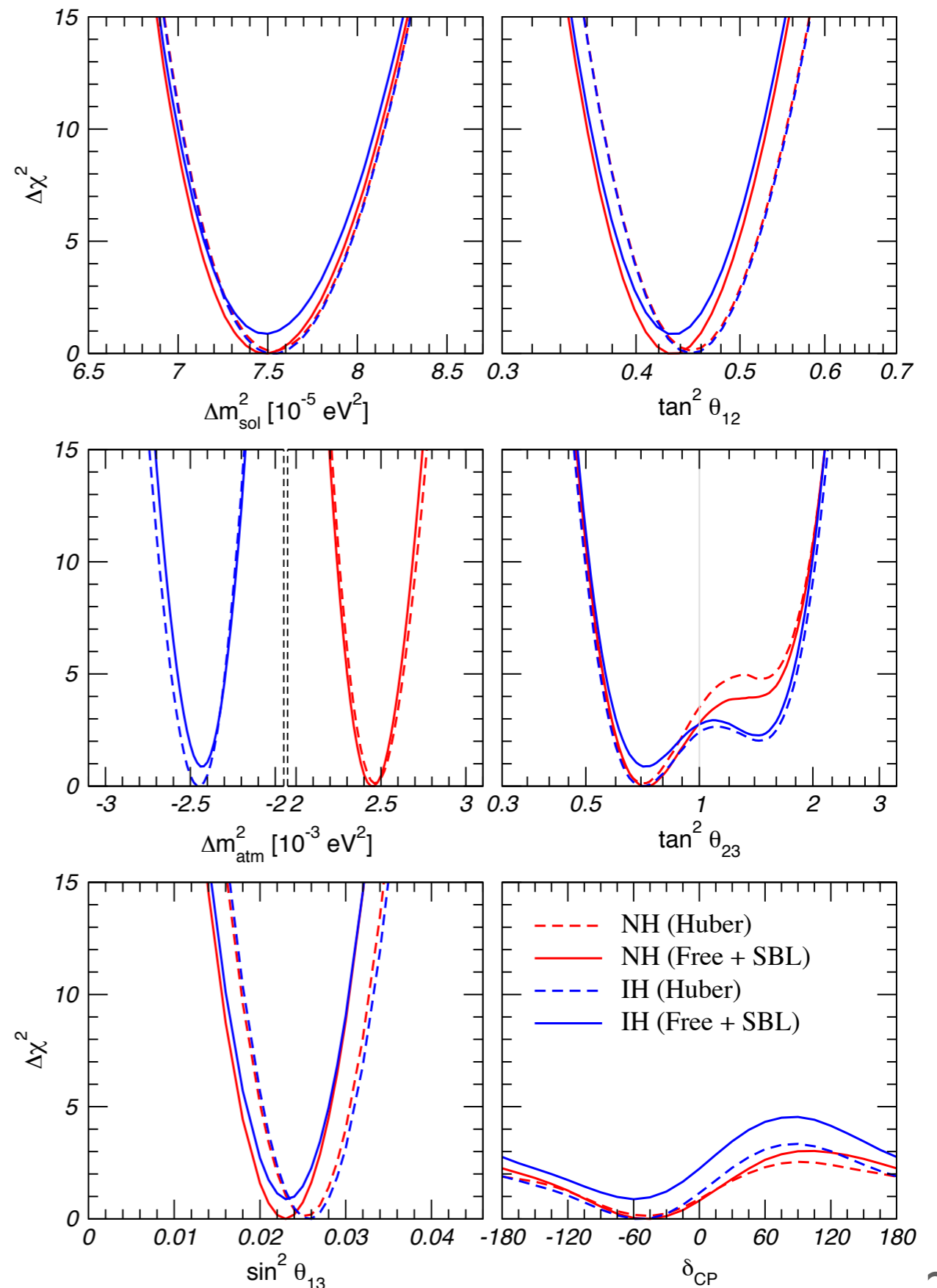


some regions of δ will appear but CP violation and hierarchy will be very hard with this generation of experiments

Summary

	bfp $\pm 1\sigma$	3σ range
$\sin^2 \theta_{12}$	$0.30^{+0.13}_{-0.12}$	0.27 – 0.34
$\sin^2 \theta_{23}$	$0.42^{+0.037}_{-0.031}$	0.34 – 0.66
$\sin^2 \theta_{13}$	0.023 ± 0.0023	0.016 – 0.030
$\delta/^\circ$ (NH)	-48^{+53}_{-59}	—
$\delta/^\circ$ (IH)	-59^{+49}_{-60}	—
$\Delta m_{21}^2 / 10^{-5} \text{eV}^2$	7.50 ± 0.185	7.00 – 8.09
$\Delta m_{31}^2 / 10^{-3} \text{eV}^2$ (NH)	$2.45^{+0.067}_{-0.071}$	2.25 – 2.67
$ \Delta m_{32}^2 / 10^{-3} \text{eV}^2$ (IH)	2.43 ± 0.068	2.23 – 2.65

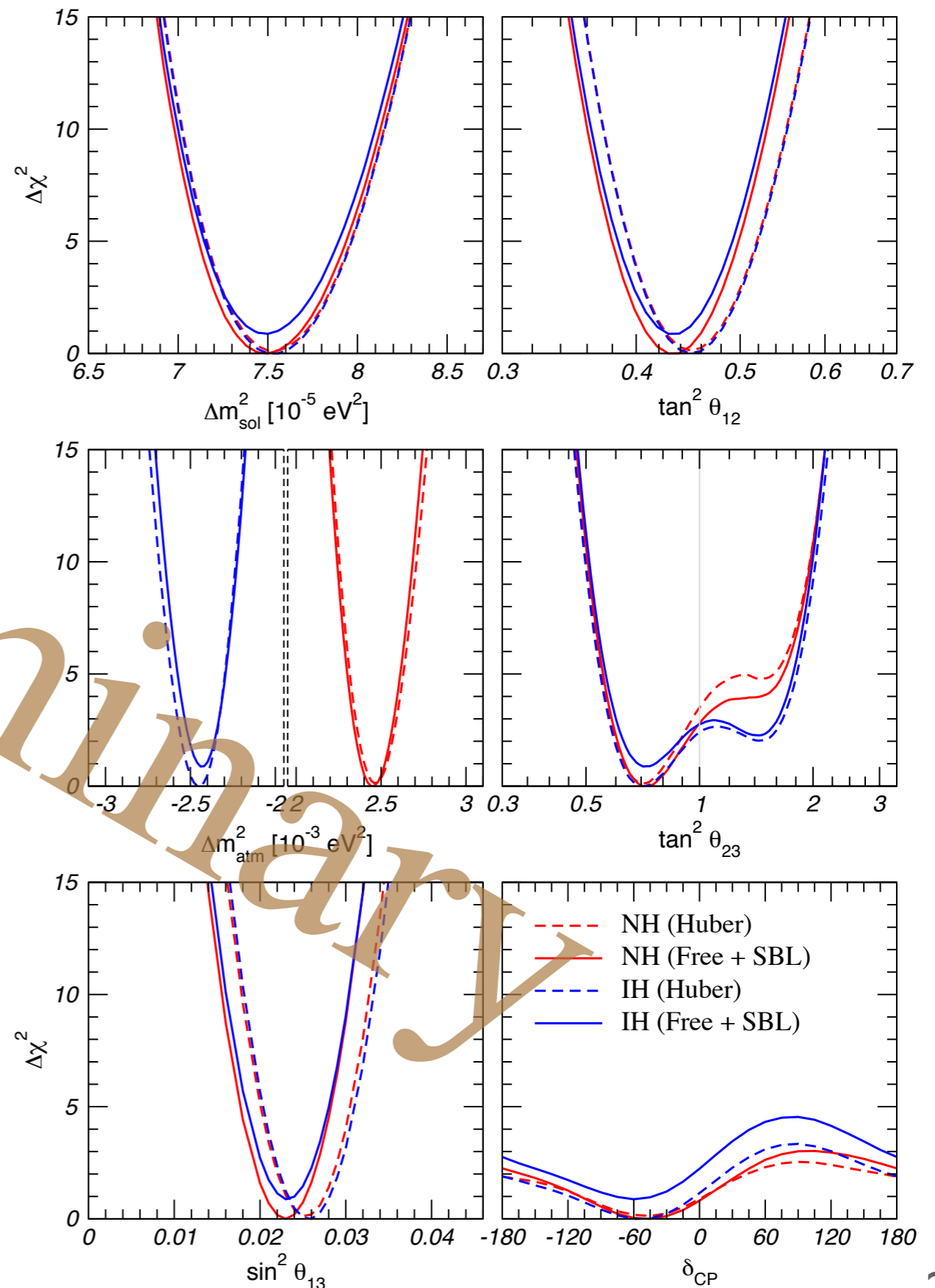
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- *this work launches an Invisibles initiative:*

ν -fit collaboration

C. Gonzalez-Garcia (UB), M. Maltoni (UAM), TS (MPIK)

- *involve and train young scientists in global analysis of neutrino data*

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Thank you for your attention!