



Neutrino Mass Models- **An SO(10) GUT Perspective**

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“Smirnov-fest” GGI Florence, 2012

.Complete neutrino physicist



Rock star of nu physics

2000-2012



Happy Birth Day Alexei !!



Outline of Talk

- Minimal SO(10) model for neutrinos and θ_{13}
- Proton decay window to GUT models for neutrinos:
- GUT scale baryogenesis revived
- Testing GUT theories by low energy measurements (other than p-decay)



Flavor puzzle in particle physics

- quark flavor puzzle: hierarchy in masses and mixings
- Neutrino mixings: $\theta_{23} \sim 45^\circ$; $\theta_{12} \sim 33^\circ$,
mass hierarchy unknown but $\Delta m_{\odot}^2 / \Delta m_A^2 \sim \theta_C^2$
- +Recent θ_{13} results:
 - T2K: $0.03 \leq \sin^2(2\theta_{13}) \leq 0.28$
 - MINOS: $0.00 \leq \sin^2(2\theta_{13}) \leq 0.12$
 - Double Chooz: $\sin^2(2\theta_{13}) = 0.085 \pm 0.051$
 - Daya Bay: $\sin^2 2\theta_{13} = 0.092 \pm 0.016 \pm 0.005$
 - RENO expt. $\sin^2 2\theta_{13} = 0.103 \pm 0.013 \pm 0.011$
 - Is there a Big Picture unifying quarks and leptons ?



From Data to Physics

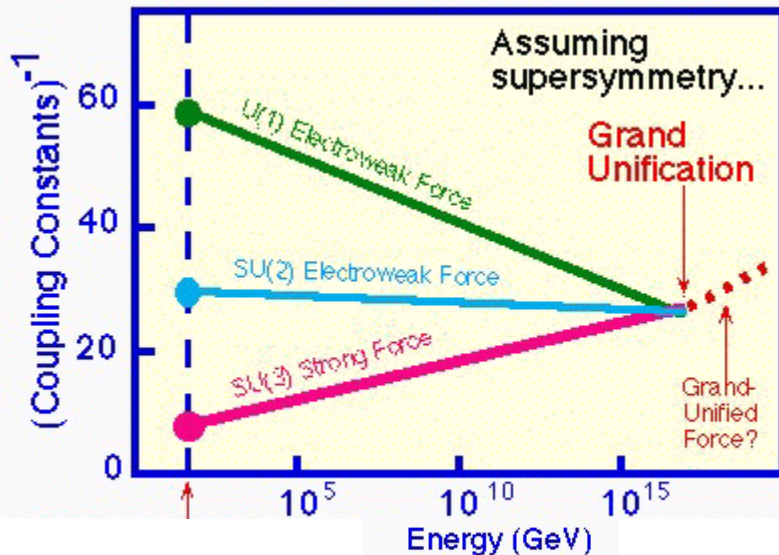
- Important for physics is the mass matrix, M_ν :
- Near Maximal atmospheric angle already gives some idea:

$$M_\nu = \begin{pmatrix} - & - & - \\ - & m - \varepsilon & m \\ - & m & m \end{pmatrix}$$
 large solar $M_\nu = \begin{pmatrix} - & \sim \varepsilon & \sim \varepsilon \\ \sim \varepsilon & m - \varepsilon & m \\ \sim \varepsilon & m & m \end{pmatrix}$
- Quark mass matrix: $M_d \approx m_b \begin{pmatrix} \lambda^4 & \lambda^3 & \lambda^3 \\ \lambda^3 & \lambda^2 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix}$ $\lambda \sim .22$
 $\sim \varepsilon$
- Can a physics big picture ever connect them or quarks and leptons are separate "beasts" ?

GUT approach

Recital of standard virtues of GUTs:

★ $\alpha_i(M_Z) + \text{SUSY at TeV} \rightarrow \text{couplings unify}$



$$M_U \approx 10^{16} \text{ GeV}$$

★ Explains electric charge quantization

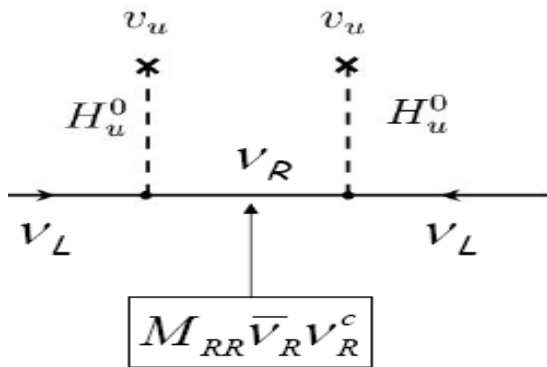
★ Unifies all matter- connecting quark-lepton flavor

Is nu mass a messenger of GUT physics ?

- Coupling unification scale in susy: $M_U \approx 10^{16} GeV$

- Seesaw paradigm

- SM+ right handed neutrinos N_R with Majorana mass: B-L



$$m_\nu \cong - \frac{h_\nu^2 v_{wk}^2}{M_R} \quad (\text{Type I})$$

MR Large \rightarrow neutrino mass tiny

M_R How Large ? GUTs $\rightarrow h_{\nu,33} \approx h_t$; $\Delta m_{atm}^2 \rightarrow M_R \approx 10^{14} GeV$

- Broken B-L sym.- part of BSM (GUT?) physics !!



GUT Scenarios

- **SU(5) with 45+ N_R with Majorana mass near M_U**
 - Why is M_N large?
 - 60 parameters; Need symmetries to predict: (A_4, S_4, T')

(Altarelli, Feruglio; King, Luhn, Antusch, Spinrath; Chen, Mahanthappa; Hagedorn, Meroni, Petcov, Spinrath; Ishimori, Saga, Tanimoto, Shimizu; [Bazzochi, Merlo....](#))

- **Type II alternative: SU(5) +45+15**
 - 45 parameters

(Joaquim, Rossi; Hambye, RNM, Nasri, Yu....)

SO(10) SUSY GUT –dream picture for neutrinos

- SO(10) unifies all fermions/family (including RH nu) in single rep.

$$\begin{pmatrix} u & u & u & \nu \\ d & d & d & e \end{pmatrix}_{L,R}$$

- Scales : $SO(10) \rightarrow MSSM \rightarrow SM$

- Contains all the ingredients for seesaw

- Minimal **renormalizable** models with **126-Higgs** **predictive** for nu masses and mixings in terms of quark masses (18 parameters) (Babu, Mohapatra'93)

- $10+126+120 \rightarrow 15$ parameters (Dutta, Mimura, RNM'05; Bertolini, Frigerio, Malinsky; Aulakh, Garg; Grimus, Kuhbock most recent analysis Altarelli, Blankenburg'10)



Predictive SO(10)

■ Quark lepton masses with 10(H)+126 (Δ) Higgs

$$W = h\psi\psi H + f\psi\psi\Delta$$

$$M_u = hv_{u,10} + fv_{u,126}$$

$$M_d = hv_{d,10} + fv_{d,126}$$

$$M_\ell = hv_{d,10} - 3fv_{d,126}$$

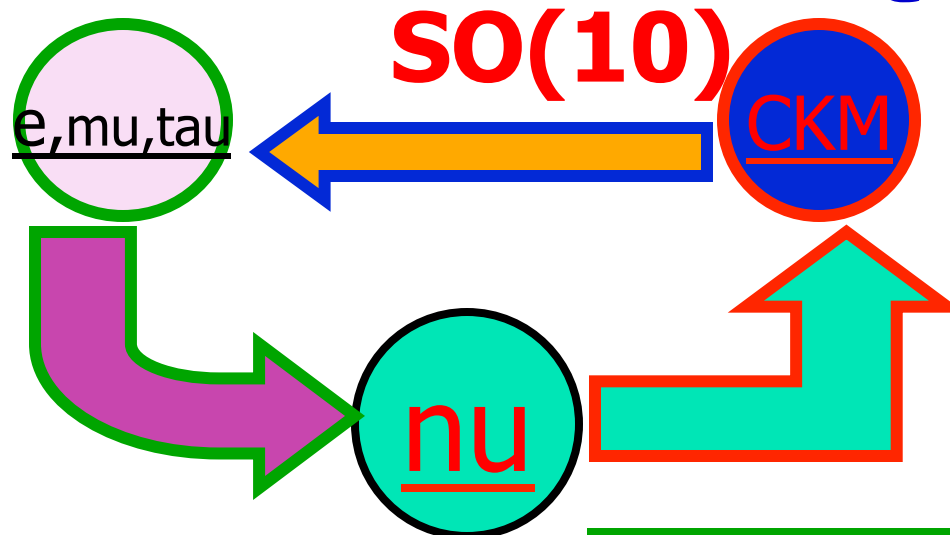
$$M_D = hv_{u,10} - 3fv_{u,126}$$

$$M_\nu = fv_L - M_D(fv_R)^{-1}M_D$$

Predictive Quark-lepton connection in minimal $SO(10)_{126}$

Leptons

Quarks



$$M_\nu$$

$$U_{PMNS} = U_\ell^\dagger U_\nu$$

(Babu, Mohapatra'93; Fukuyama,Okada'02; Bajc, Senjanovic, Vissani'02; Goh, RNM, Ng'03; Aulakh,Bajc,Melfo,Senjanovic, Vissani; Fukuyama,Ilakovic,Meljanac,Kikuchi,Okada; Bertolini, Malisky, Schwetz; Babu, Macesanu; Joshipura,Patel)

SO(10) with Type II seesaw and large mixings

■ II seesaw

$$M_\nu \cong c(M_d - M_l)$$

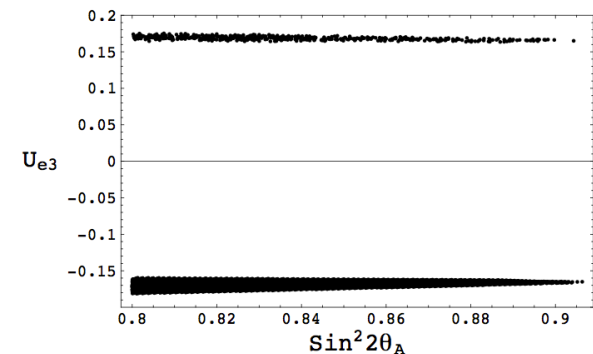
■ GUT relation $m_b \approx m_\tau + \lambda^2 \rightarrow$

$$M_\nu \propto m_b \begin{pmatrix} \lambda^5 & \lambda^3 & \lambda^3 \\ \lambda^3 & \lambda^2 & \lambda^2 \\ \lambda^3 & \lambda^2 & 1 \end{pmatrix} - m_\tau \begin{pmatrix} \lambda^6 & \sim \lambda^3 & \\ \lambda^3 & -3\lambda^2 & \lambda^2 \\ & \lambda^2 & 1 \end{pmatrix} \sim \lambda^2 \begin{pmatrix} \lambda^2 & \lambda^2 & \lambda \\ \lambda^2 & 1 + \lambda & 1 \\ \lambda & 1 & 1 \end{pmatrix}$$

(Bajc, Senjanovic, Vissani' 02)

■ Predicts $\theta_{13} \simeq .17 ; m_{\odot} / m_{\oplus} \sim \lambda$

(Goh, RNM, Ng, Phys.Lett. B570 (2003) 215)



(Daya Bay-RENO value ~ 0.15)



SO(10) with Type I+II

■ Why?

- How well does the GUT relation $m_b \approx m_\tau + \lambda^2$ work ?
- Quite well for susy SO(10) with large $\tan\beta \sim 50$; within 2 sigma for $\tan\beta = 10$
- Large $\tan\beta$ values have problem with p-decay
- Non-Susy no $m_b \approx m_\tau$
- More general analysis with I+II+CPV gives same large θ_{13} (Babu, Macesanu; Bertolini, Malinsky, Schwetz; '05)
- Large θ_{13} generic of 126 models with susy.

SO(10)₁₂₆ without susy

- b-τ does not work.
- Numerical fermion fits work for the minimal model still work with type I

(Joshipura and Patel'11)

$$\sin^2 2\theta_{13} = .093 : \text{Type I}$$

Agrees with Daya Bay

Observables	Type-I		Type-II	
	Fitted value	pull	Fitted value	pull
m_d	0.000810163	-0.687161	0.00101285	-0.264898
m_s	0.0208099	-0.198354	0.0225915	0.0844982
m_b	0.999667	-0.00831657	1.08201	2.05031
m_u	0.000495023	0.0751133	0.000507336	0.13668
m_c	0.237348	0.0670883	0.237096	0.0598882
m_t	73.9427	-0.0154941	74.3006	0.075144
m_e	0.000469652	-	0.000469652	-
m_μ	0.0991466	-	0.0991466	-
m_τ	1.68558	-	1.68558	-
$\left(\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}\right)$	0.030526	0.127968	0.0297114	-0.235285
$\sin \theta_{12}^q$	0.224651	0.0464044	0.224499	-0.0916848
$\sin \theta_{23}^q$	0.0420499	0.0392946	0.0421308	0.103004
$\sin \theta_{13}^q$	0.00349369	-0.0974312	0.00353053	0.0389979
$\sin^2 \theta_{12}^l$	0.323245	0.148134	0.3108	-0.610792
$\sin^2 \theta_{23}^l$	0.435096	-0.369178	0.113306	-7.02461
$\sin^2 \theta_{13}^l$	0.0244287	-	0.0176863	-
$\delta_{CKM} [^\circ]$	69.5262	-0.0314447	69.2051	-0.128759
$\delta_{MNS} [^\circ]$	318.465	-	14.5386	-
$\alpha_1 [^\circ]$	21.5053	-	345.645	-
$\alpha_2 [^\circ]$	215.128	-	141.905	-
$r_{R(L)}$	5.62×10^{-14}	-	2.09×10^{-10}	-



How to test the model ?

- Broken B-L – a key element of SO(10) seesaw
- p-decay tests-(Conventional B-L=0 modes do not test seesaw)
- $\mu \rightarrow e + \gamma$ (does not test nonsusy seesaw)
- Look for p-decay modes violating B-L
- New result: SO(10) GUTs without SUSY \rightarrow
- Observable proton decay modes with ~~B-L~~
- New TeV mass color sextet particles at LHC
- Observable NN-bar oscillation



GUTs and p-decay

- Canonical GUT modes for p-decay:

$$p \rightarrow e^+ + \pi^0; p \rightarrow K^+ \bar{\nu}$$

- Conserve B-L=0

- D=6 operators in SM (Weinberg; Wilczek, Zee'79)

$$\mathcal{O}_1 = (d^c u^c)^* (Q_i L_j) \epsilon_{ij}, \quad \mathcal{O}_2 = (Q_i Q_j) (u^c e^c)^* \epsilon_{ij}, \quad \mathcal{O}_3 = (Q_i Q_j) (Q_k L_l) \epsilon_{ij} \epsilon_{kl}$$

$$\mathcal{O}_4 = (Q_i Q_j) (Q_k L_l) (\vec{\tau} \epsilon)_{ij} \cdot (\vec{\tau} \epsilon)_{kl}, \quad \mathcal{O}_5 = (d^c u^c)^* (u^c e^c)^* .$$

- Compare with d=5 operator $LHLH$

B-L=2 B-violation

(ii) D=7 : Have **B-L=2** (Weinberg; Weldon, Zee'80)

$$\tilde{\mathcal{O}}_1 = (d^c u^c)^* (d^c L_i)^* H_j^* \epsilon_{ij},$$

$$\tilde{\mathcal{O}}_2 = (d^c d^c)^* (u^c L_i)^* H_j^* \epsilon_{ij},$$

$$\tilde{\mathcal{O}}_3 = (Q_i Q_j) (d^c L_k)^* H_l^* \epsilon_{ij} \epsilon_{kl},$$

$$\tilde{\mathcal{O}}_4 = (Q_i Q_j) (d^c L_k)^* H_l^* (\vec{\tau} \epsilon)_{ij} \cdot (\vec{\tau} \epsilon)_{kl},$$

$$\tilde{\mathcal{O}}_5 = (Q_i e^c) (d^c d^c)^* H_i^*,$$

$$\tilde{\mathcal{O}}_6 = (d^c d^c)^* (d^c L_i)^* H_i,$$

$$\tilde{\mathcal{O}}_7 = (d^c D_\mu d^c)^* (\bar{L}_i \gamma^\mu Q_i),$$

$$\tilde{\mathcal{O}}_8 = (d^c D_\mu L_i)^* (\bar{d}^c \gamma^\mu Q_i),$$

$$\tilde{\mathcal{O}}_9 = (d^c D_\mu d^c)^* (\bar{d}^c \gamma^\mu e^c).$$

$$\rightarrow n \rightarrow e^- \pi^+$$

(iii) D=9: $u^c d^c d^c u^c d^c d^c$ **B-L=2** $\rightarrow n - \bar{n}$
 (RNM, Marshak'80; Glashow'80)

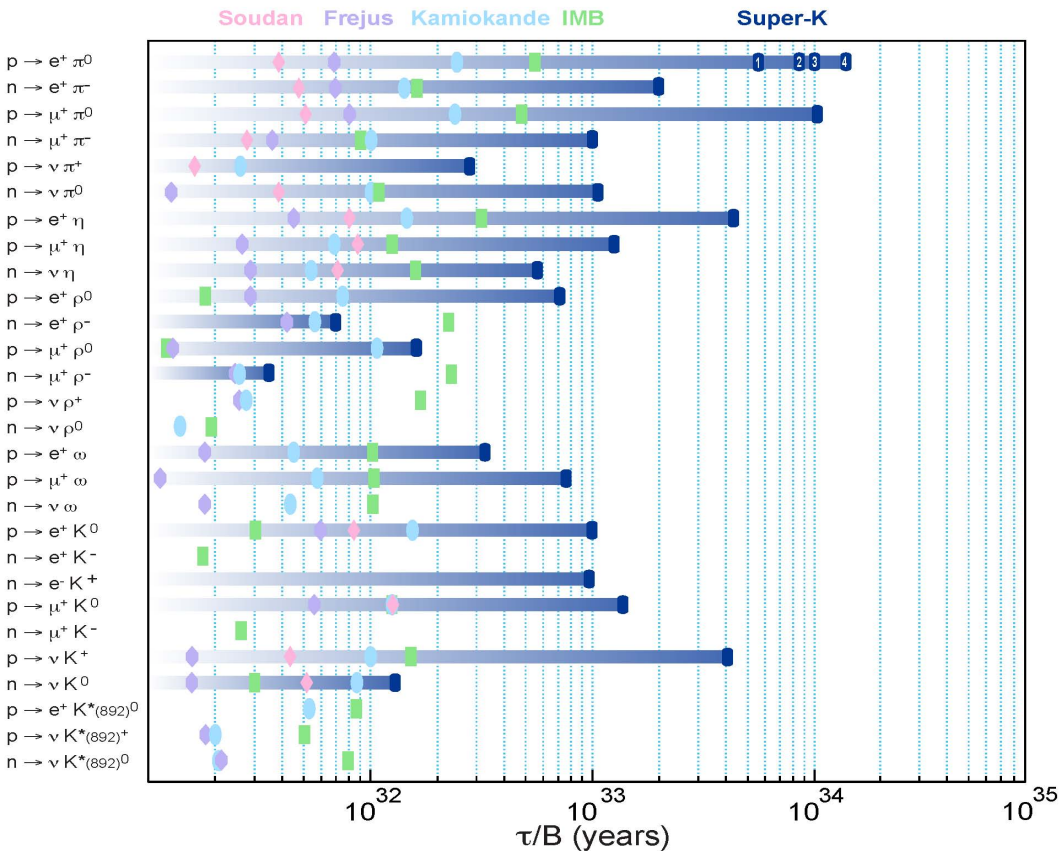
- Same property as seesaw – are they related to neutrino mass? Are they present in GUT theories with seesaw?

~~B-L~~ modes in non-SUSY SO(10) for nu mass

- Key ingredient of minimal renorm. SO(10) models for nu mass is 126-Higgs field that leads to seesaw.
- If some of the $\{126\}$ scalar fields remain at TeV to \sim TeV scale, they can give enhanced B-violation with B-L=2 e.g. $n \rightarrow e^- \pi^+$
 $n - \bar{n}$ oscillation (Babu, Mohapatra; arXiv: 1203.5544; 1205.5701)
- Can be observable in Hyper-K
- Is it consistent with coupling unification ?

Proton decay-expt status

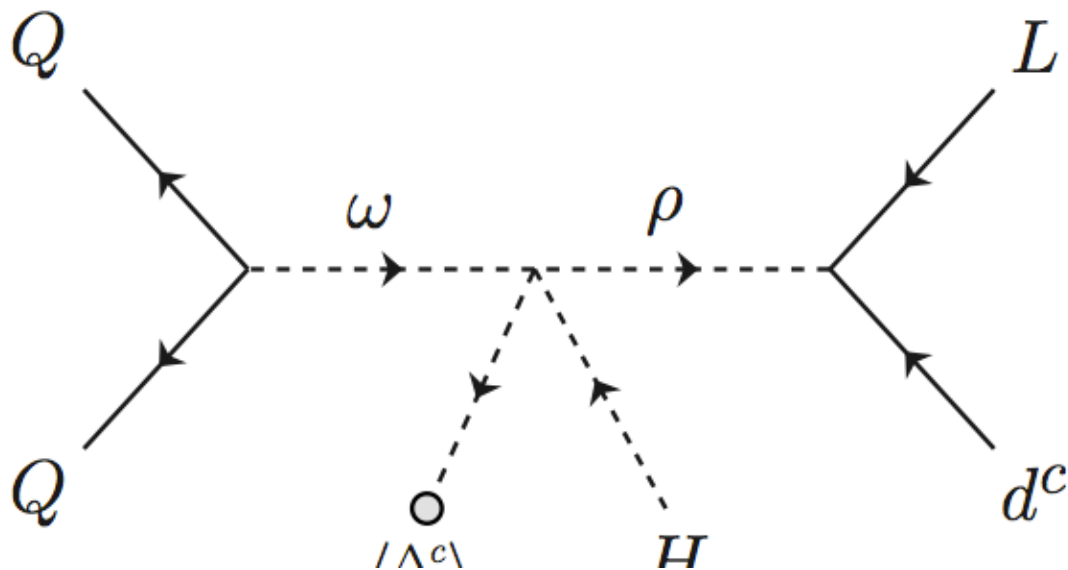
Super-K limits



★ Hyper-K to push limits by factor 10

Intermediate scale Scalars and B-L=2 p-decay

Diagram for neutron-decay: d=7 Operator O_2



$$\tilde{O}_2 = (Q_i Q_j)(d^c L_i)^* H_j^* \epsilon_{ij} \epsilon_{kl}$$

ω in $\{10, 126\}$;

$\rho(3, 2, 1/6) \{126\}$

$\omega \rho^* H$ in $\{126\}^4$

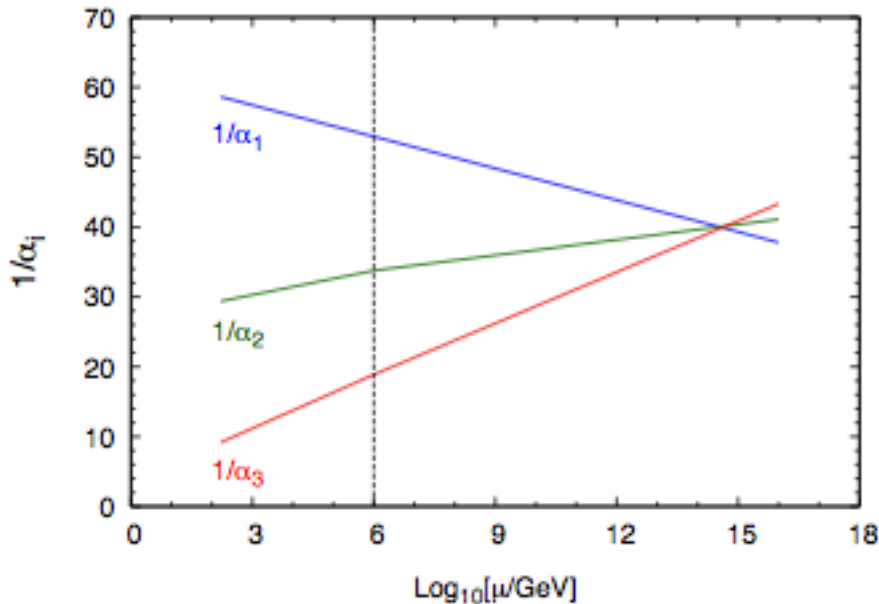
Leads to $n \rightarrow e^- + \pi^+$ decay

Coupling unification and observability

$$A_{n \rightarrow e-\pi^+} \sim \frac{f_{126}^2 v_{BL} v_{wk}}{M_\omega^2 M_\rho^2} \sim \frac{10^{-20}}{M_\rho^2}$$

Observability $\rightarrow M_\rho \sim 10^5 \text{ GeV}$; is it GUT allowed ?

Yes \rightarrow



\rightarrow

$$\tau_p \sim 4 \times 10^{33} \text{ yrs}$$

Alternative SO(10) scenario

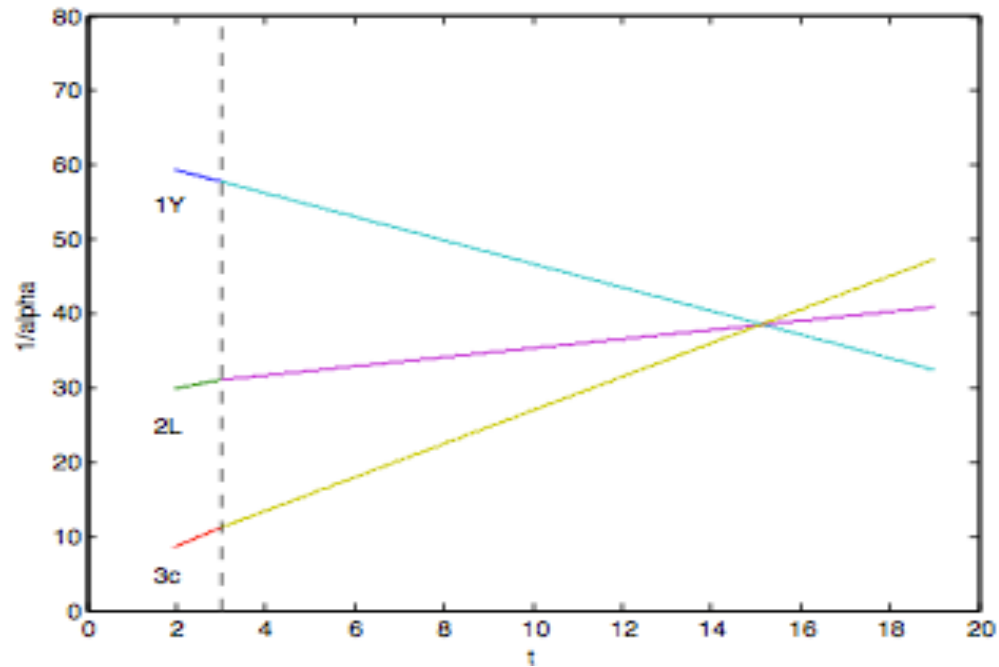
- Non-SUSY SO(10) does not unify without low scale particles, (Babu, Mohapatra, arXiv:1206.5701)

- Coupling unif with sub-TeV $\Delta_{ud}(6, 1, \frac{1}{3})$ + 2 SM triplets;

- Predicts seesaw scale near $M_U \sim 10^{15.7}$ GeV;
 $\tau_{p \rightarrow e^+ + \pi^0} \sim 3 \times 10^{34}$ yrs

- Δ_{ud} mass ~ 2 TeV GeV

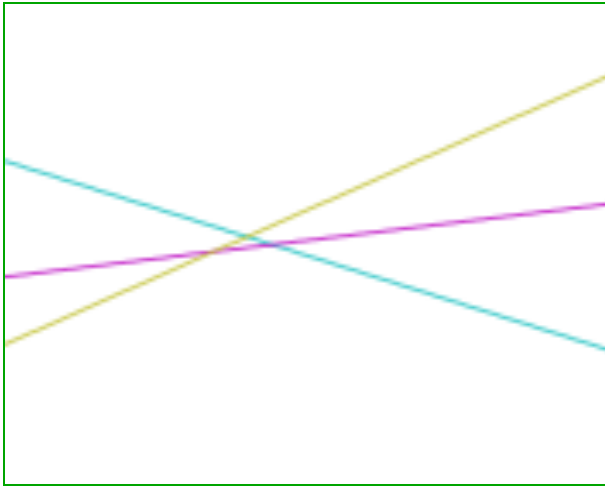
- B-L violation \rightarrow GUT scale coupling $v_{BL} \Delta_{ud} \Delta_{ud} \Delta_{dd}$



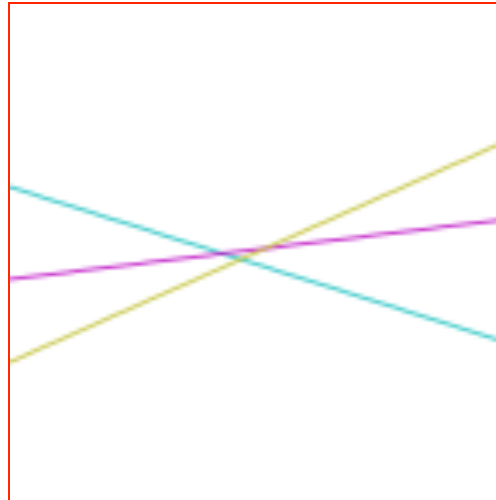
Scale sensitivity-1-loop

- Coupling unification is sensitive to sextet mass

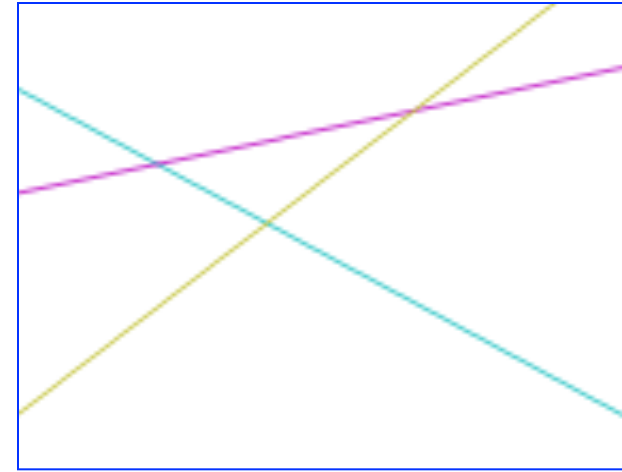
1 TeV



2 TeV



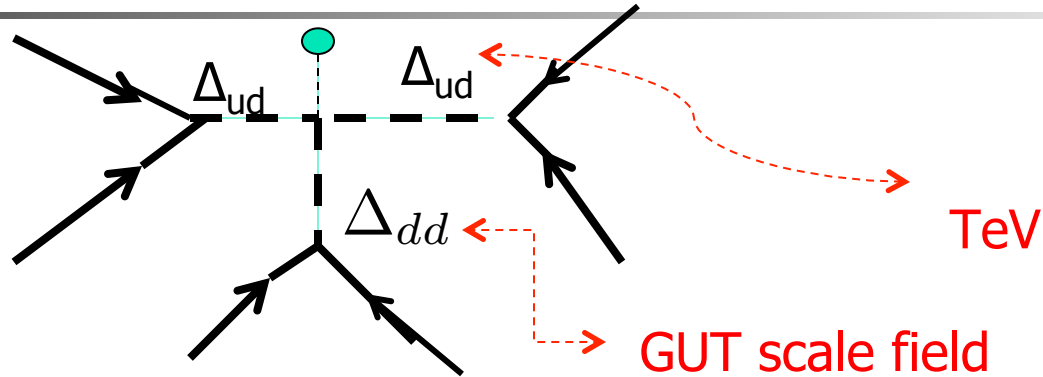
10 TeV



N-N-bar oscillation from

$$\nu_{BL} \Delta_{ud} \Delta_{ud} \Delta_{dd}$$

Diagram:



$$G_{\Delta B=2} \simeq \frac{\lambda f_{11}^3 \eta^3}{\lambda' M_U M_{\Delta_{ud}}^4} \simeq \frac{\lambda}{\lambda'} 10^{-33} \text{GeV}^{-5}$$

Predicts $\tau_{n-\bar{n}} \sim 10^{10} - 10^{13}$ sec.

Observable with current facilities: *Reactors, Project X at Fermilab, etc*

Color sextets Δ_{qq} @LHC

- **TeV scale Color sextets** Can be searched at LHC:

(I) **Single production:** $ud \rightarrow \Delta_{ud} \rightarrow tj$

cross-section calculated in (RNM, Okada, Yu' 07;) resonance peaks above SM background- decay to tj ; $\sigma(tj) > \sigma(\bar{t}\bar{j})$

- **Important LHC signature:**

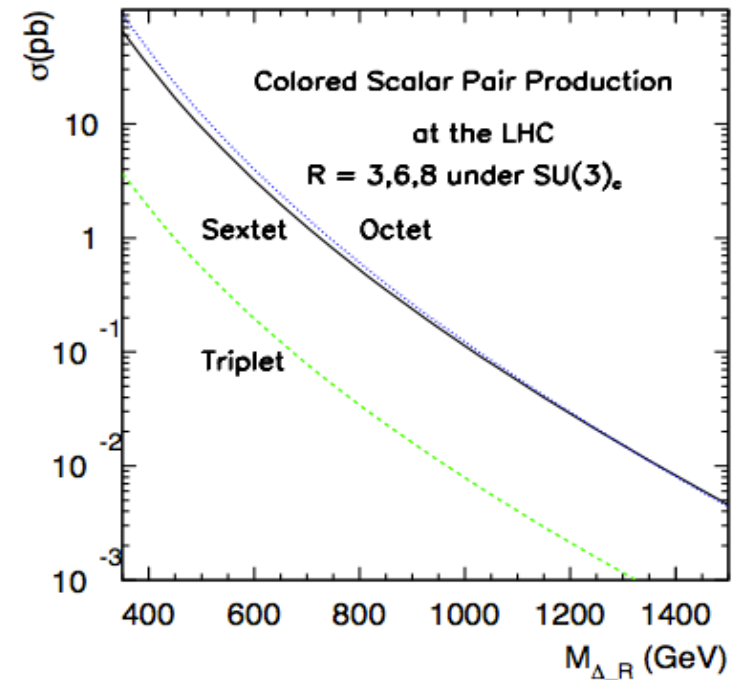
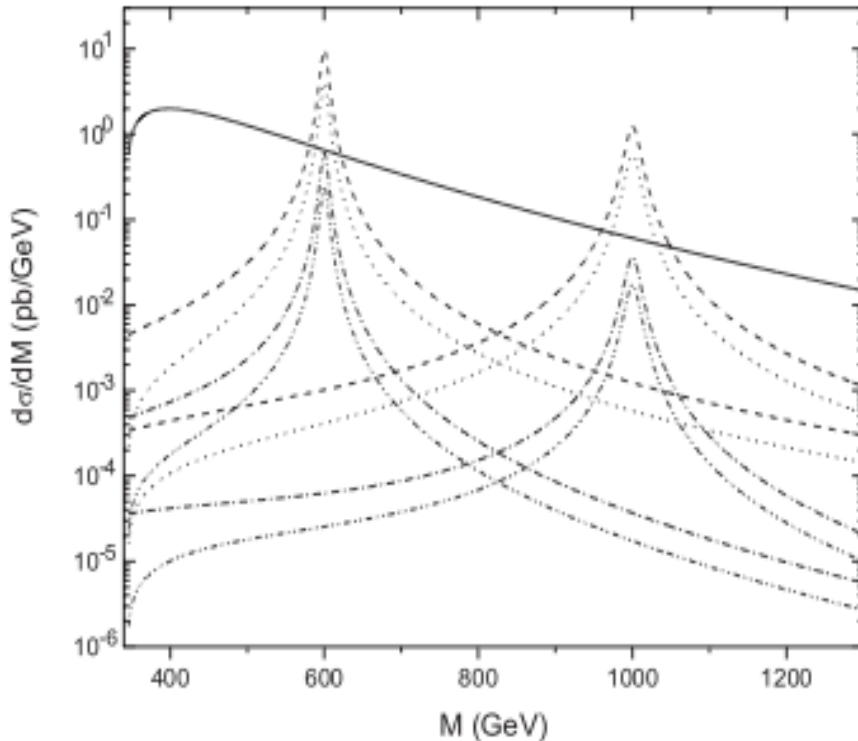
(II) **Drell-Yan pair production**

$$q\bar{q} \rightarrow G \rightarrow \Delta_{ud} \bar{\Delta}_{ud}$$

- Leads to $tjtj$ final states: **LHC reach < TeV**

(Chen, Rentala, Wang; Berger, Cao, Chen, Shaughnessy, Zhang' 10; Han, Lewis' 09)

Cross sections for single and pair productions



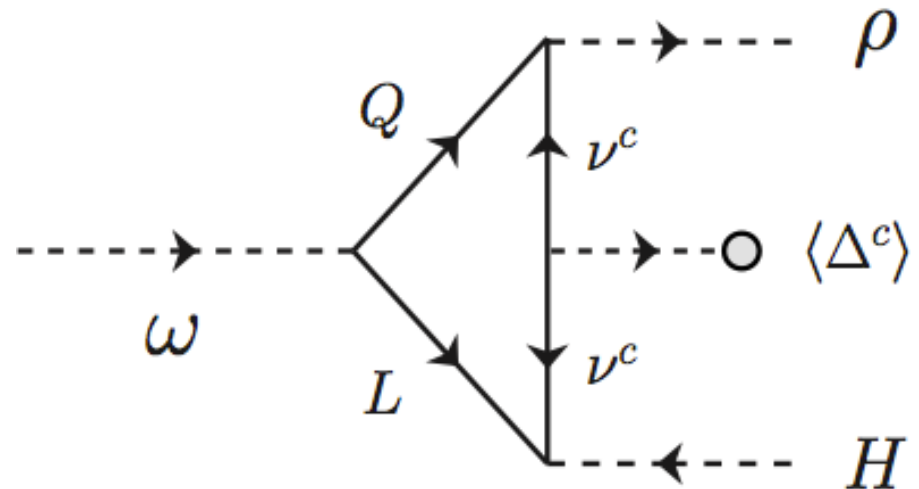
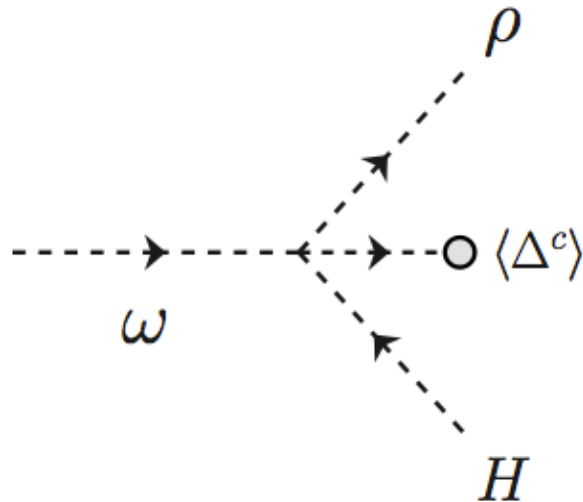
14 TeV LHC

Seesaw in SU(5) and ~~B-L~~ nucleon decays

- SU(5) with type I and III seesaw:
- Required exotic Higgs couplings for BLV decay:
$$\omega \rho^* h; \eta^* \rho h; \rho^* \Phi h; \chi^* \eta h$$
- But ρ and η absent in SU(5) nu models with type I or III seesaw. \rightarrow hence no ~~B-L~~ p-decays
- **Way to distinguish SU(5) vs SO(10)**
- SU(5) with type II seesaw \rightarrow {15}-field
- Leads to observable BLV n-decay but not nn-bar

Proton decay and baryogenesis connection

- Baryogenesis without leptogenesis
- Decays of $\omega \rightarrow \rho H$ can produce baryon asym.



- Graphs responsible; right order.
- Sphalerons do not wash away- it has $B-L=2$

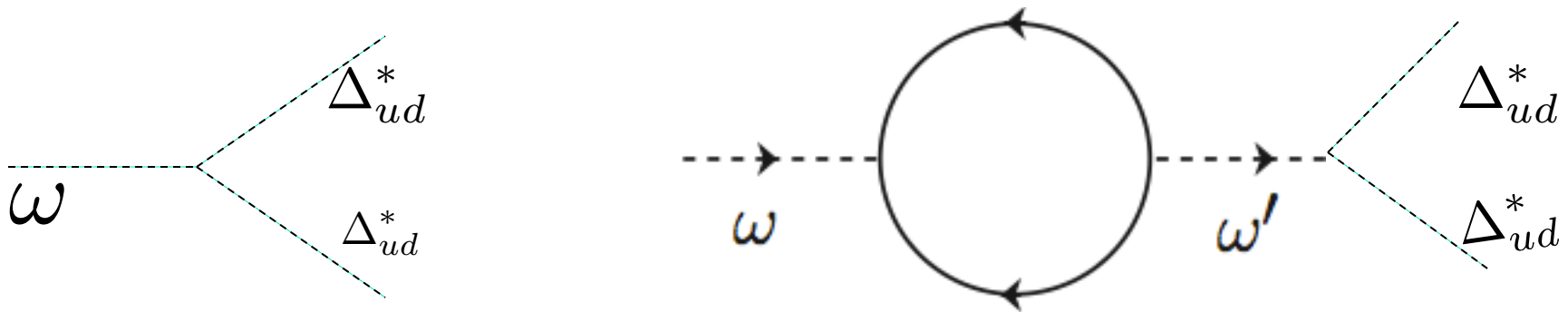
Asymmetry related to neutrino mass

- Asymmetry related to ω -couplings that are couplings of $\{10\}$ and $\{126\}$ and related to neutrino masses:

$$\epsilon_{B-L}^{(c)} = \frac{1}{\pi} \text{Im} \left[\frac{\text{Tr}\{Y_{d^c\nu^c\omega} Y_{d^cL\rho}^\dagger Y_{\nu^cLH} M_{\nu^c} F_2(M_{\nu^c})\} \lambda v_R}{|\lambda v_R|^2} \right] \text{Br},$$

The color sextet model

- N-N-bar interactions go out of eq. around 10^{15} GeV;
- Two sources of matter-anti-matter asymmetry:
 - (a) Leptogenesis
 - (b) B-L violating GUT scale by $\Delta_{dd}(\omega)$ decay





Summary

- SO(10) GUTs with Type I or II seesaw promising
 - Is a natural home for neutrino masses and a unified description of quark and lepton flavor.
- Dayabay value of θ_{13} remarkably close to minimal SO(10)₁₂₆ predictions
- True test of GUTs-proton decay:
- Proton decay modes bearing seesaw signature (i.e. B-L=2) e.g. $n \rightarrow e^- \pi^+$, $n - \bar{n}$ observable !!
- Distinguishes SO(10) from SU(5) models:
- Revival of GUT scale baryogenesis

Understanding Flavor

- Where does the Yukawa texture come from ?
- **Usual Lore:** Zero fermion masses \rightarrow SM + RH nu \rightarrow flavor symmetry group:

$$U(3)_Q \times U(3)_u \times U(3)_d \times U(3)_\ell \times U(3)_e \times U(3)_N$$

- Hope is that observed flavor structure is a consequence of breaking this symmetry-
- Implement this strategy using flavons whose vevs are **Yukawa couplings !** (Altarelli, Feruglio; Hagedorn King, Luhn, Medeiros Varzielas, Ross; Chen, Mahanthappa,..)



Flavor in $SO(10)_{126}$ GUTs

- Note that dominant type II seesaw models demand \rightarrow

$$f_{126} = \lambda^2 \begin{pmatrix} \lambda^2 & \lambda^2 & \lambda \\ \lambda^2 & 1 + \lambda & 1 \\ \lambda & 1 & 1 \end{pmatrix}$$

- We need to understand this pattern that leads to the 10+126 model below the sym. Breaking scale !!

- An attempt with discrete family sym S_4 above GUT scale with triplet flavons lead to this (Dutta, Mimura and RNM, 2010):



S_4 Theory predictions

- Realistic $10+126+10'$ model with triplet flavons:

$$\phi_1 = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}, \quad \phi_2 = \begin{pmatrix} 0 \\ -1 \\ 1 \end{pmatrix}, \quad \phi_3 = \begin{pmatrix} 1 \\ 1 \\ 1 \end{pmatrix}$$

- Yukawa texture:

$$h \propto \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}, \quad f \propto \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & -1 \\ 0 & -1 & 1 \end{pmatrix} + \lambda \begin{pmatrix} 0 & 1 & 1 \\ 1 & 0 & 1 \\ 1 & 1 & 0 \end{pmatrix}, \quad h' \propto \begin{pmatrix} 0 & 1 & -1 \\ 1 & 0 & 0 \\ -1 & 0 & 0 \end{pmatrix}$$

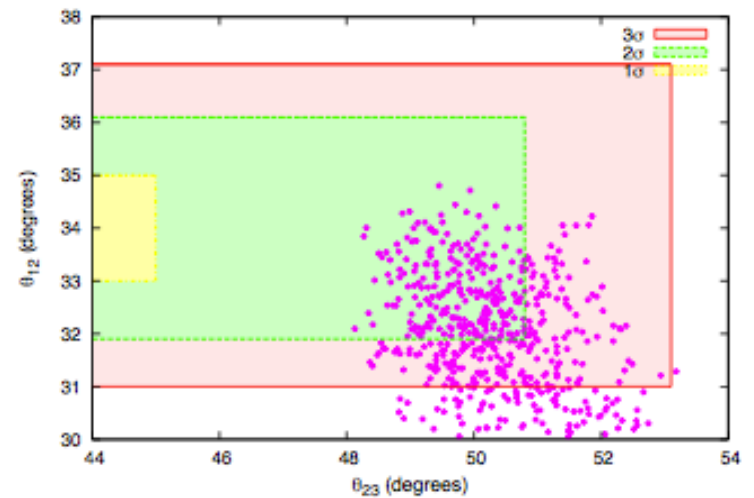
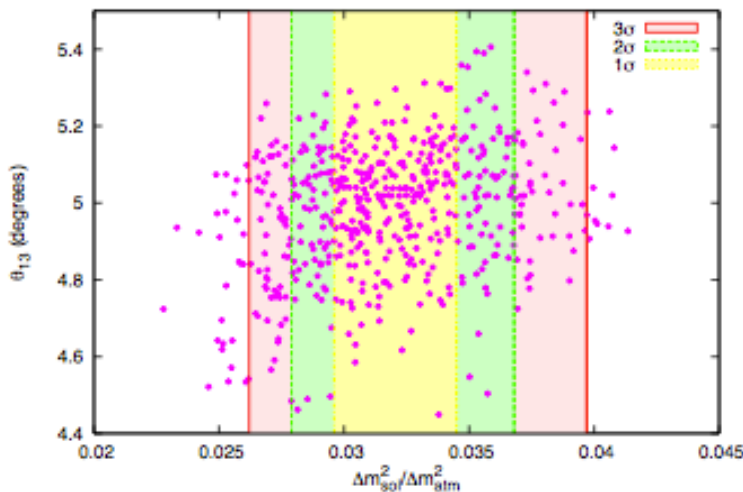
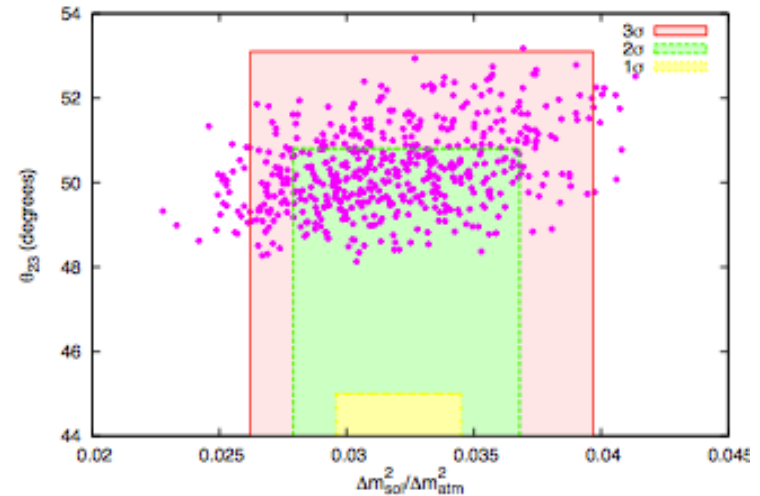
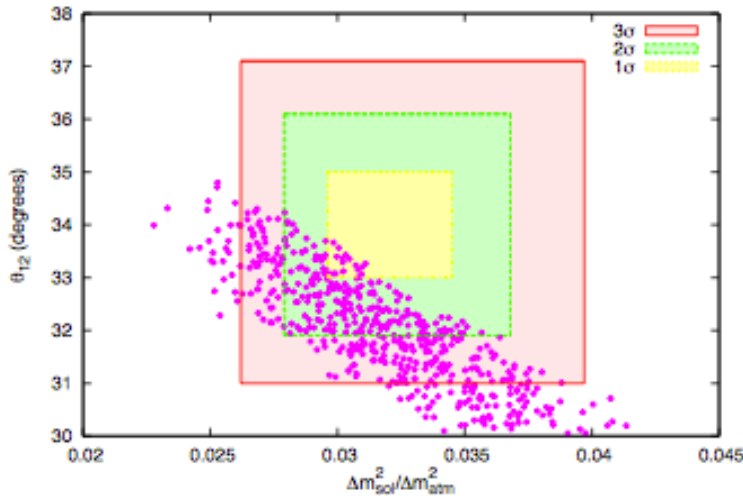
Prediction of

$$\theta_{13}, \theta_{23}, \theta_{12}$$

Model

correl.

(Dev, Dutta,
RNM,
Severson'12)

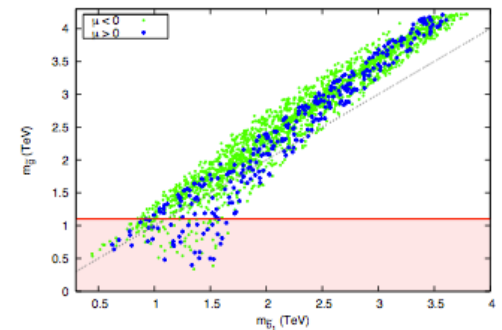
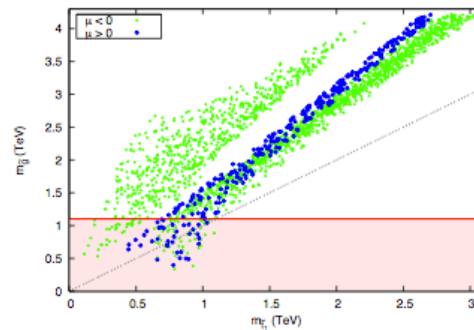
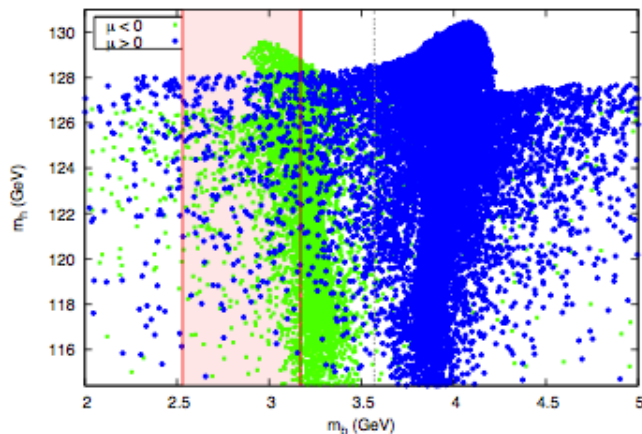


Implications for Higgs mass

- To fit m_b large threshold corrections needed—that requires large A-term in susy →

- Has implications for susy spectrum:

stop-gluino masses



- $120 \text{ GeV} < M_H < 128 \text{ GeV}$.

Current best fit values

■ Schwetz, Tortola, Valle :1108

Curious feature!

	best fit $\pm 1\sigma$	3σ range	prec@ 3σ	
$\frac{\Delta m_{21}^2}{10^{-5} \text{eV}^2}$	$7.59^{+0.20}_{-0.18}$	7.09–8.19	7%	KamLAND
$\frac{\Delta m_{31}^2}{10^{-3} \text{eV}^2}$	$2.50^{+0.09}_{-0.16}$ $-(2.40^{+0.08}_{-0.09})$	2.14 – 2.76 $-(2.13 - 2.67)$	12%	MINOS
$\sin^2 \theta_{12}$	$0.312^{+0.017}_{-0.015}$	0.27–0.36	14%	SNO
$\sin^2 \theta_{23}$	$0.52^{+0.06}_{-0.07}$ 0.52 ± 0.06	0.39–0.64	24%	SuperK
$\sin^2 \theta_{13}$	$0.013^{+0.007}_{-0.005}$ $0.016^{+0.008}_{-0.006}$	0.001–0.035 0.001–0.039	120%	T2K + global data
δ	$(-0.61^{+0.75}_{-0.65}) \pi$ $(-0.41^{+0.65}_{-0.70}) \pi$	0 – 2π	—	

upper: normal hierarchy, lower: inverted hierarchy

$$\frac{\Delta m_{\odot}^2}{\Delta m_{\oplus}^2} \sim \theta_{Cabibbo}^2$$

General SM multiplets responsible for B-L decays

- Multiplets: $h(1, 2, +1/2)$, $\bar{h}(1, 2, -1/2)$, $\omega(3, 1, -1/3)$, $\omega^c(\bar{3}, 1, 1/3)$,
 $\rho(3, 2, 1/6)$, $\bar{\rho}(\bar{3}, 2, -1/6)$, $\eta(3, 1, 2/3)$, $\bar{\eta}(\bar{3}, 1, -2/3)$,
 $\Phi(3, 3, -1/3)$, $\bar{\Phi}(\bar{3}, 3, 1/3)$, $\chi(3, 2, 7/6)$, $\bar{\chi}(\bar{3}, 2, -7/6)$,
 $\delta(3, 1, -4/3)$, $\bar{\delta}(\bar{3}, 1, 4/3)$.

- Couplings: $\omega \rho^* h$; $\eta^* \rho h$; $\rho^* \Phi h$; $\chi^* \eta h$
 Present in $(126)^4$ coupling in SO(10)

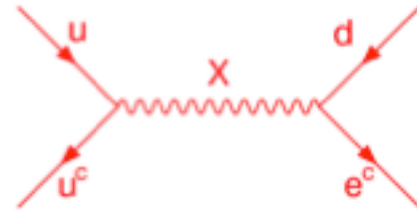
$$\mathcal{L}(16_i 16_j 10_H) = h_{ij} \left[(u_i^c Q_j + \nu_i^c L_j) h - (d_i^c Q_j + e_i^c L_j) \bar{h} + \left(\frac{\epsilon}{2} Q_i Q_j + u_i^c e_j^c - d_i^c \nu_j^c \right) \omega + (\epsilon u_i^c d_j^c + Q_i L_j) \omega^c \right], \quad (6)$$

$$\mathcal{L}(16_i 16_j \bar{126}_H) = f_{ij} \left[(u_i^c Q_j - 3\nu_i^c L_j) h - (d_i^c Q_j - 3e_i^c L_j) \bar{h} + \sqrt{3}i \left(\frac{\epsilon}{2} Q_i Q_j - u_i^c e_j^c + \nu_i^c d_j^c \right) \omega_1 + \sqrt{3}i (Q_i L_j - \epsilon u_i^c d_j^c) \omega_1^c + \sqrt{6} (d_i^c \nu_j^c + u_i^c e_j^c) \omega_2 + 2\sqrt{3}i d_i^c L_j \rho - 2\sqrt{3}i \nu_i^c Q_j \bar{\rho} + 2\sqrt{3} u_i^c \nu_j^c \eta - 2\sqrt{3}i u_i^c L_j \chi + 2\sqrt{3}i e_i^c Q_j \bar{\chi} - 2\sqrt{3} d_i^c e_j^c \delta + \sqrt{6}i Q_i L_j \bar{\Phi} + \dots \right], \quad (7)$$

B-L conserving modes -- consistency check on model

Gauge exchange

$$p \rightarrow e^+ \pi^0$$



$$\Gamma^{-1}(p \rightarrow e^+ \pi^0) = (2.0 \times 10^{35} \text{ yr})$$

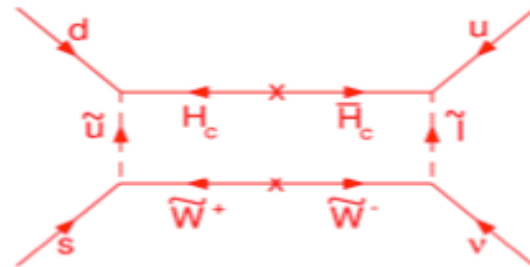
$$\times \left(\frac{\alpha_H}{0.01 \text{ GeV}^3} \right)^{-2} \left(\frac{\alpha_G}{1/25} \right)^{-2} \left(\frac{A_R}{2.5} \right)^{-2} \left(\frac{M_X}{10^{16} \text{ GeV}} \right)^4$$

Higgsino exchange

Model dependent –

Potentially “fatal”

(e.g. minimal SU(5))



$$p \rightarrow \bar{\nu} K^+$$

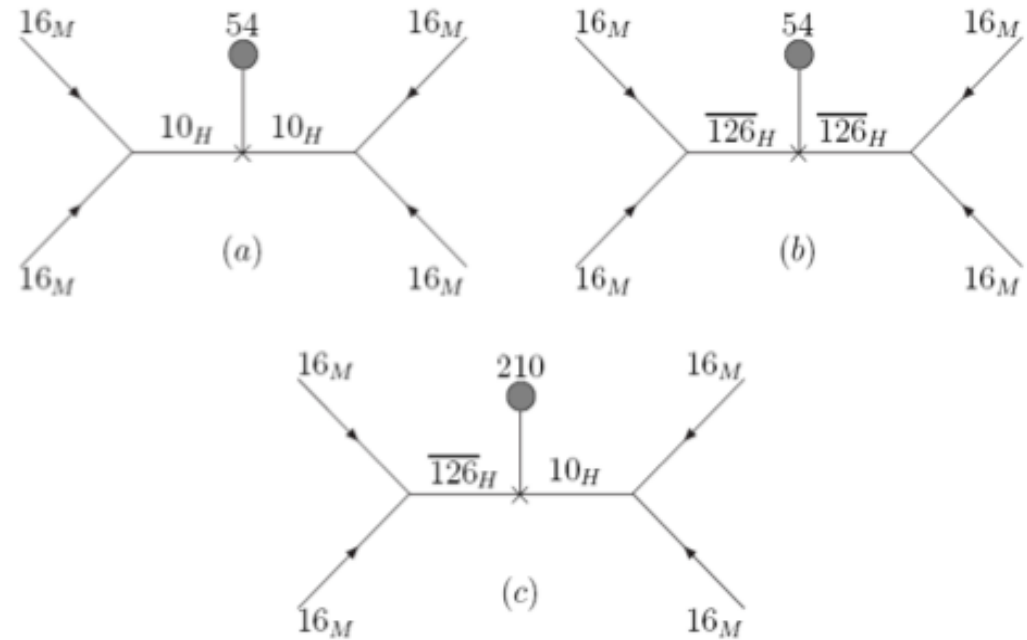
$$\tau_p^{-1} \approx \left[\frac{f^2}{M_{H_c} M_{SUSY}} \right]^2 \left(\frac{\alpha}{4\pi} \right)^2 m_p^5 \approx [10^{28} - 10^{32} \text{ yr}]^{-1}$$

Proton decay - SUSY SO(10)

- Minimal 10+126
- cancellation possible for low tan β

- A prediction

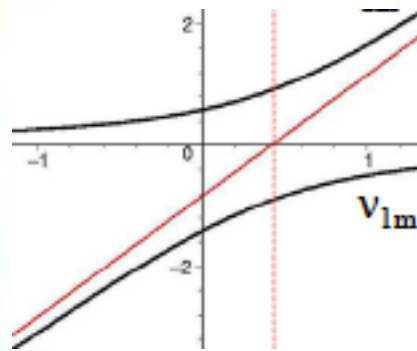
$$\tau_{p \rightarrow \pi + \nu} < 6 \times 10^{32} \text{ yrs}$$



(Goh, RNM, Nasri, Ng'05)

. Alexei enhancing

Alexei in 10 yrs



Rock star of nu Phys

Wise Guru of nu physics