# Sterile Neutrinos in IceCube

### Sandhya Choubey Harish-Chandra Research Institute, India



### What is nu? Invisibles 12 & Alexei Smirnov Fest GGI Firenze, June 25 - 29, 2012

## First there was LSNP...

#### LSND data Beam Excess 17.5 Beam Excess $*3.8\sigma$ excess in $p(\bar{v}_{\mu} \rightarrow \bar{v}_{e'}, e^{+})n$ 15 antineutrinos at p(v̄<sub>e</sub>,e⁺)n 12.5 L/E = 0.4 - 1.2 m/MeVother 10 7.5 5 \* requires presence of 2.5 sterile neutrino states 0 with $\Delta m^2 = 1 eV^2$ 2012 0.4 0.6 0.8 1.2 1 L/E, (meters/MeV)

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1.4

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## and then MiniBooNE...

### \* The idea was to specifically test the LSND signal

### \* Since oscillations depend on L/E.....

### \* The L/E was chosen to be the same as LSND

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## MiniBooNE (2007)

### April 7, 2007 Press Release

"The MiniBooNE results resolve questions raised by observations of the LSND.....

MiniBooNE researchers showed conclusively that the LSNP results could not be due to simple neutrino oscillation,...."

#### arxiv:0704.1500



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# MiniBooNE (2007)

### April 7, 2007 Press Release

"It was very important to verify or refute the surprising LSNP result," said Robin Staffin, POE Associate Director of Science for High Energy Physics. "The MiniBooNE experiment was an important one to do and is to be complimented for a job well done."

\* "We are delighted to see that the work of the MiniBooNE team has led to the resolution of this puzzle," said Marv Goldberg, Program Director for Elementary Particle Physics at the National Science Foundation. "

As in many particle physics experiments, we have a result that answers some questions and raises others," said MiniBooNE co-spokesperson William Louis, Los Alamos National Laboratory, who also worked on the original LSNP experiment. "We live in interesting times."

\*'It is great to get the MiniBooNE results out," said Fermilab Director Pier Oddone. "It clears one mystery but it leaves us with a puzzle that is important to understand."

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# MiniBooNE (2007)

### But wait...LSNP saw oscillations in anti-neutrinos

### MiniBooNE 2007 data was for neutrinos....

## CPT violation?

## No...only CP violation is enough... add 2 sterile neutrinos .....

Sorel, Conrad, Shaevitz **(3+2) mass spectrum** Maltoni, Schwetz, 0705.0107 0305255

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Karagiorgi et al, 0906.1997 June 29, 2012

# Miniboone (2009)

#### From the abstract

antineutrino data sample corresponding to 3.39 x 10<sup>20</sup> protons on target

"No significant excess of events has been observed..... The data are inconclusive with respect to antineutrino oscillations suggested by data from the Liquid Scintillator Neutrino Petector at Los Alamos National Laboratory."

#### arxiv:0907.1958



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# Miniboone (2010)

#### From the abstract

antineutrino data sample corresponding to 5.66 x 10<sup>20</sup> protons on target

"The data are consistent with  $\nabla_{\mu}$  to  $\nabla_{e}$  in the 0.1 to 1.0 eV<sup>2</sup>  $\Delta m^{2}$  range and with the evidence for antineutrino oscillations from the Liquid Scintillator Neutrino Detector at Los Alamos National Laboratory."

#### arxiv:1007.1150



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## Miniboone (2010)

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#### arxiv:1007.1150



### Does that mean that it is inconsistent with their v data?

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## Miniboone (2010)





\*Neutrino and anti-neutrino data now consistent

\* Together they see a  $3.8\sigma$  excess....LSND??

\* If neutrino and anti-neutrino data are consistent then do we need CP violation anymore?

\* What is then the role of 3+2?

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M. Maltoni talk

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#### Talks by M. Maltoni and S. Brice

### \* The reactor anomaly..... $2.9\sigma$

### \* The Gallium anomaly..... $3.3\sigma$

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## Best-fit Parameters



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# Oscillations driven by Sterile Neutrinos

\* oscillation maximum:  $\sin^2\left(\frac{\Delta m^2 L}{4E}\right) = 1$ \*  $E(TeV) = 8.1 \times \left(\frac{\Delta m^2}{1eV^2}\right) \times \left(\frac{L}{10,000km}\right)$ 

\* Atmospheric neutrinos travel 10,000 km to IceCube

\* Atmospheric neutrinos can have energies in 1-10 TeV



## Matter Effects at TeV (-A<sub>NC</sub> is the matter induced NC pot) \* $A_{NC} = \pm \sqrt{2}G_F \rho N_A (1 - Y_e) E$

 $\bigstar (\Delta m_{42}^2)^M = \sqrt{(\Delta m_{42}^2 \cos 2\theta_{24} - A_{NC})^2 + (\Delta m_{42}^2 \sin 2\theta_{24})^2}$ 

 $\star \tan 2\theta_{24}^{M} = \frac{\Delta m_{42}^{2} \sin 2\theta_{24}}{\Delta m_{42}^{2} \cos 2\theta_{24} - A_{NC}} \qquad \underset{0302039}{\text{Nunokawa, Peres, Zukanovich Funchal,}}$   $J_{\text{June 201:}} \star E_{res}^{\nu_{\alpha} \rightarrow \nu_{s}} (TeV) = \frac{\Delta m_{4i}^{2} \cos 2\theta_{4i}}{0.038 \times \rho(gm/cc)}$ 

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## \*For $\Delta m^2_{42}=1 \text{ eV}^2$ , $\rho = 8 \text{ gm/cc}$ , $E_{res}=3 \text{ TeV}$

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## Impact of $\theta_{34}$ on matter effects sc. 0709.1937



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## TeV Oscillations in the (3+2) Scheme SC, 0709.1937



Fluxes and Events

 $\Phi_{\mu} = \Phi^{0}_{\mu}P_{\mu\mu} + \Phi^{0}_{e}P_{e\mu} \approx \Phi^{0}_{\mu}P_{\mu\mu}, \qquad \text{effective area}$ 

$$N_j = 2\pi \int_{\Delta_j \cos \theta_z} d\cos \theta_z \int_{E_{th}} dE \ \Phi^0_\nu(E, \theta_z) A_{\text{eff}}(E, \theta_z) P_{\mu\mu}(E, \theta_z) + \text{antineutrinos.}$$



$$S_j = \frac{N_j}{N_j^0},$$

Smirnov, Razzaque, 1104.1390

$$N_j = N_j^{MC} S_j,$$

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## Limits from IC40



### \* $|U_{\mu4}|^2$ > 0.025 is ruled out at 3 $\sigma$

### \* $|U_{\mu4}|^2 \approx 0.03$ is the global best-fit from the 3+2 fit

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Smirnov, Razzaque,

11041390

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## Future Prospects at IC



## Sterile Neutrinos in Deep Core



### \* Deep Core could probe $|U_{\mu4}|^2$ down to 0.01 or 0.02

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- Many (experimental) reasons to believe that sterile neutrinos with 1 eV mass scale exist
- \* A large variety of experiments being planned to test sterile neutrino osc
- Makes sense to look for their signatures in currently running experiments

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- One should see sterile-driven MSW conversions of TeV range atmospheric neutrinos in IceCube
- \* IceCube could confirm or refute the 3+2 scheme
- \* One could see oscillation effects of sterile nus in the 10-100 GeV range
- \* ICPC could put stringent bounds...

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