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School of Physics
and Astronomy



Theoretical implications of large θ_{13}

Steve King
Florence 25th June, 2012



The Galileo Galilei Institute for Theoretical Physics
Arcetri, Florence
What is ν ?, INVISIBLES 12 and Alexei Smirnov Fest



Three Neutrino Mixing



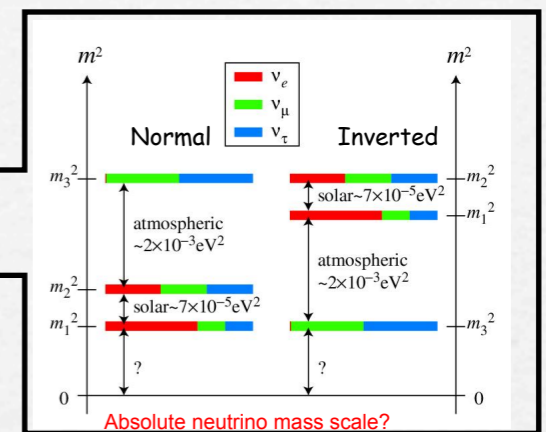
See talks by Brice and de Gouvea

Standard Model states

$$\begin{pmatrix} \nu_e \\ e^- \end{pmatrix}_L, \begin{pmatrix} \nu_\mu \\ \mu^- \end{pmatrix}_L, \begin{pmatrix} \nu_\tau \\ \tau^- \end{pmatrix}_L$$

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Neutrino mass states



Pontecorvo
Maki
Nakagawa
Sakata

$$U_{PMNS} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} e^{i\alpha_1/2} & 0 & 0 \\ 0 & e^{i\alpha_2/2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

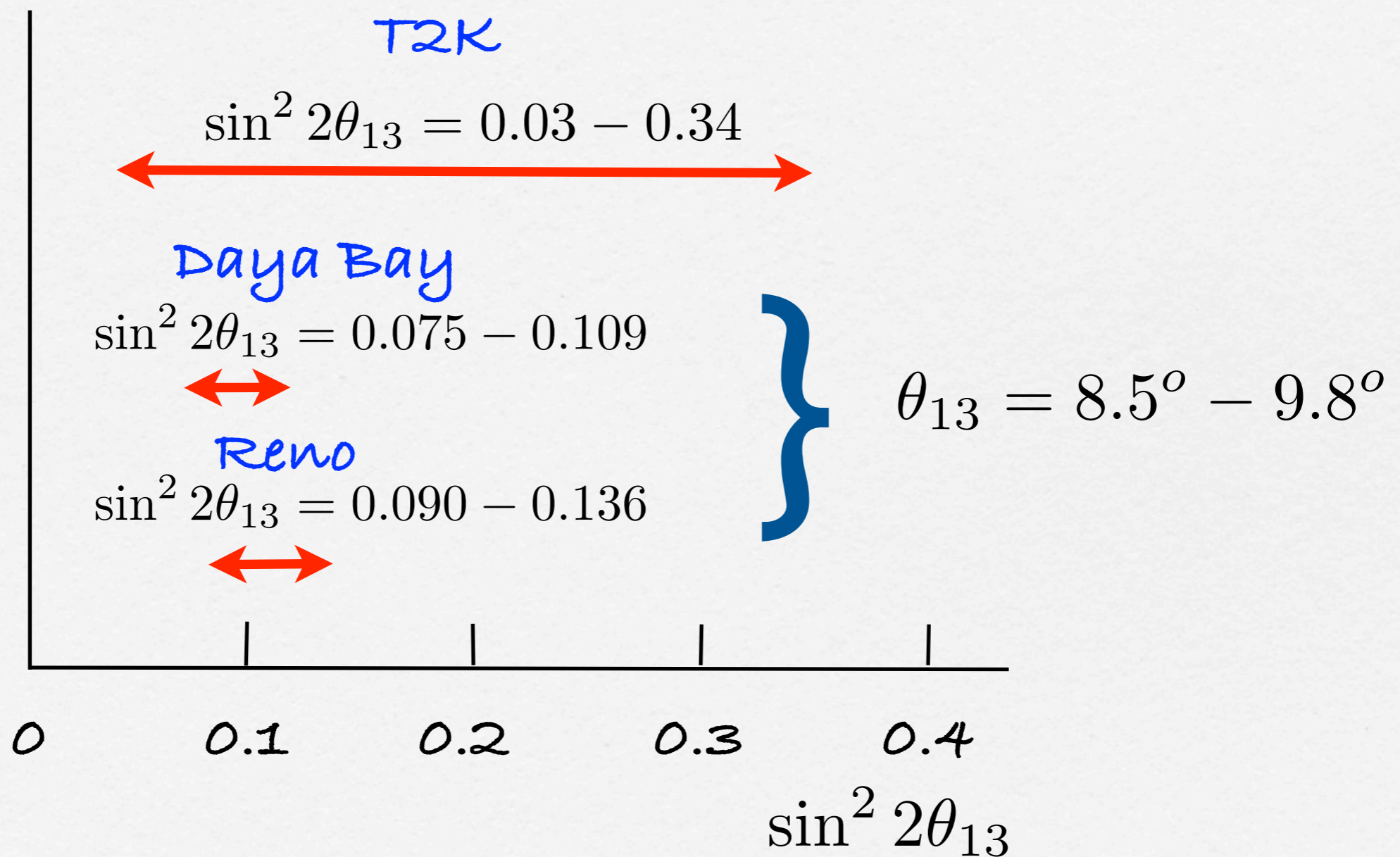
$s_{ij} = \sin\theta_{ij}$
 $c_{ij} = \cos\theta_{ij}$

Atmospheric Reactor Solar Majorana

Oscillation phase δ
Majorana phases α_1, α_2

3 masses + 3 angles + 1 (or 3) phase(s)
= 7 (or 9) new parameters for SM

Theta 13 in 2011/12



Global Fits 2012

Schwetz talk

Forero, Tortola,
valle, Vanegas '12

Fogli, Lisi, Marrone,
Montanino, Palazzo,
Rotunno '12

parameter	best fit $\pm 1\sigma$	best fit $\pm 1\sigma$
$\Delta m_{21}^2 [10^{-5} \text{eV}^2]$	7.62 ± 0.19	$7.54^{+0.26}_{-0.22}$
$\Delta m_{31}^2 [10^{-3} \text{eV}^2]$	$2.53^{+0.08}_{-0.10}$ $-(2.40^{+0.10}_{-0.07})$	$2.43^{+0.07}_{-0.09}$ $-(2.42^{+0.07}_{-0.10})$
$\sin^2 \theta_{12}$	$0.320^{+0.015}_{-0.017}$	$0.307^{+0.018}_{-0.016}$
$\sin^2 \theta_{23}$	$0.49^{+0.08}_{-0.05}$ $0.53^{+0.05}_{-0.07}$	$0.398^{+0.030}_{-0.026}$ $0.408^{+0.035}_{-0.030}$
$\sin^2 \theta_{13}$	$0.026^{+0.003}_{-0.004}$ $0.027^{+0.003}_{-0.004}$	$0.0245^{+0.0034}_{-0.0031}$ $0.0246^{+0.0034}_{-0.0031}$
δ	$(0.83^{+0.54}_{-0.64}) \pi$ $0.07\pi^a$	$(0.89^{+0.29}_{-0.44}) \pi$ $(0.90^{+0.32}_{-0.43}) \pi$

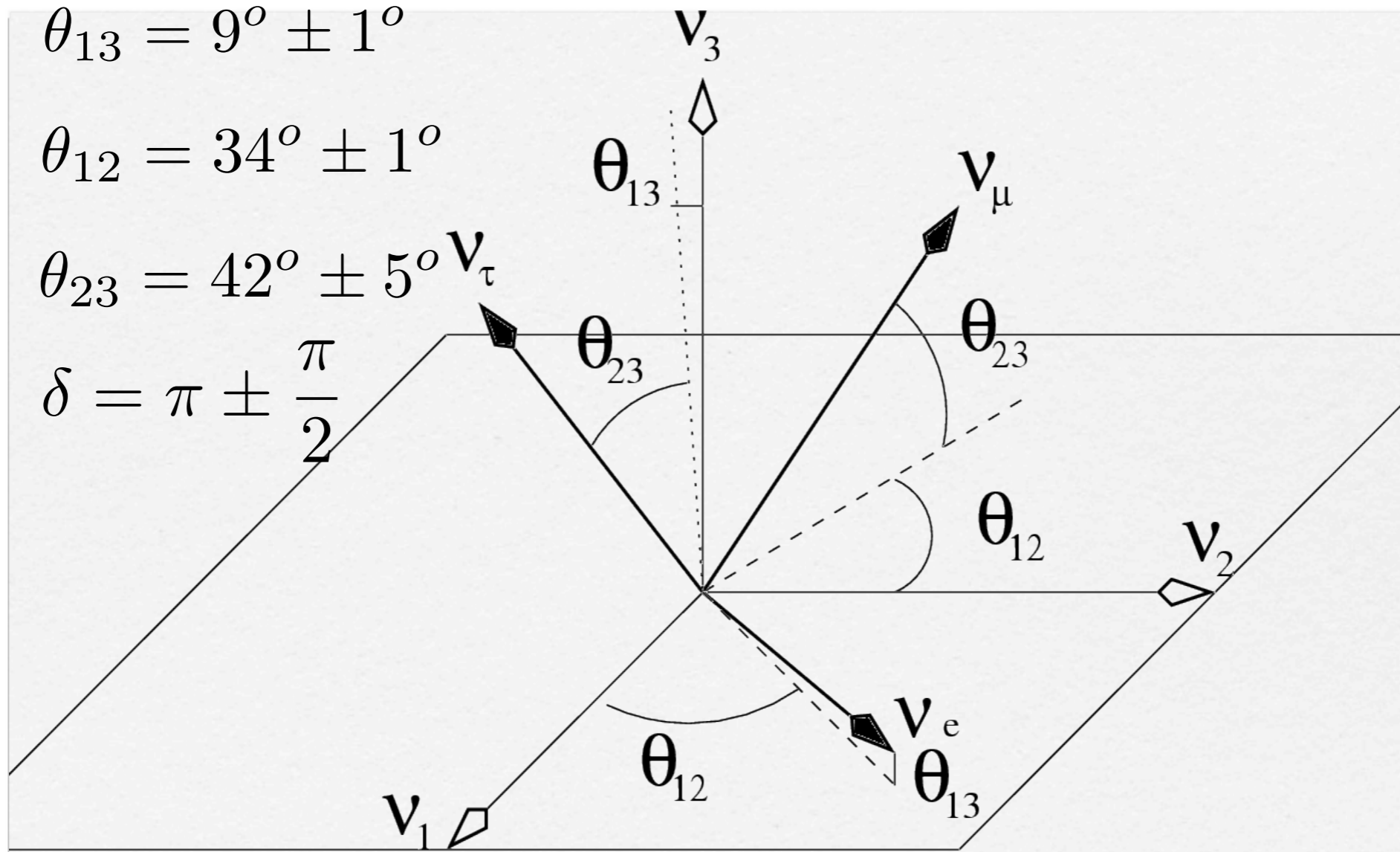
Neutrino Mixing

$$\theta_{13} = 9^\circ \pm 1^\circ$$

$$\theta_{12} = 34^\circ \pm 1^\circ$$

$$\theta_{23} = 42^\circ \pm 5^\circ$$

$$\delta = \pi \pm \frac{\pi}{2}$$



See-saw mechanism

P.Minkowski, PLB67(1977)421 ...

Possible type II contribution

Dirac matrix

$$\begin{pmatrix} \overline{\nu}_L & \overline{\nu}_R^c \end{pmatrix} \begin{pmatrix} 0 & m_{LR} \\ m_{LR}^T & M_{RR} \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix}$$

$$M^{\nu} = m_{LR} \cdot \frac{1}{M_{RR}} \cdot m_{LR}^T$$

Light Majorana matrix

Heavy Majorana matrix

□ Neutrinos are light because RH neutrinos are heavy

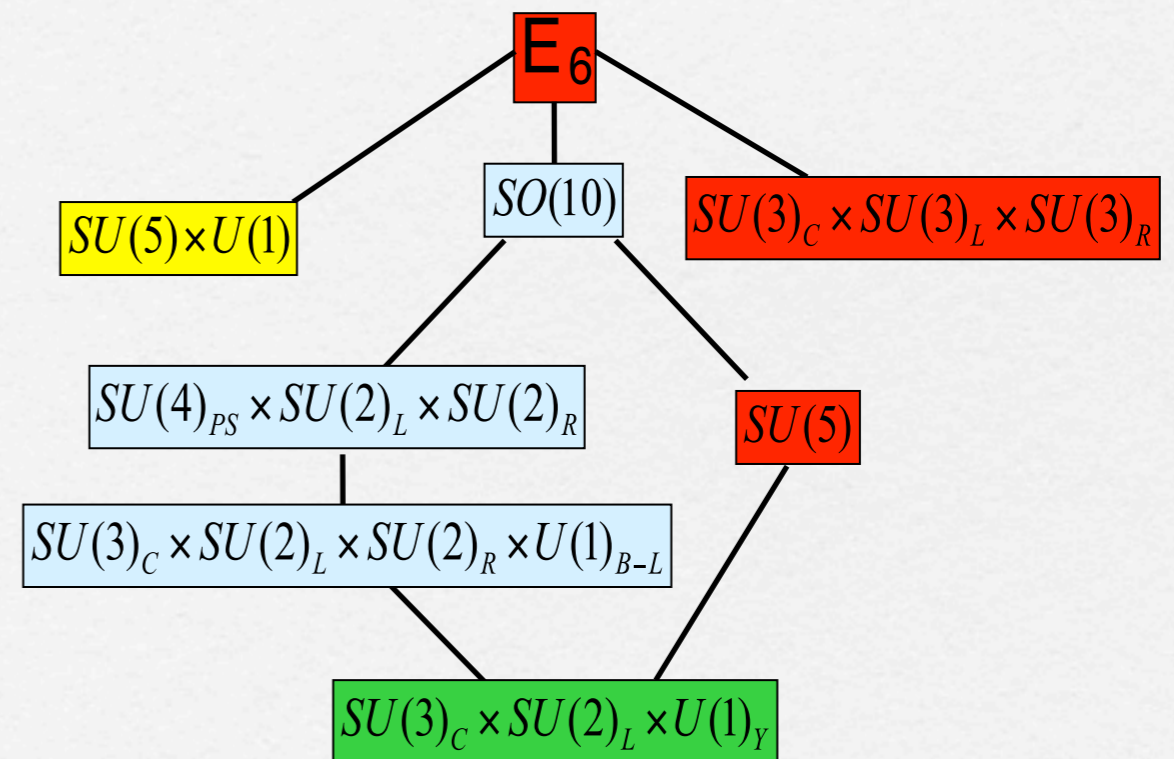
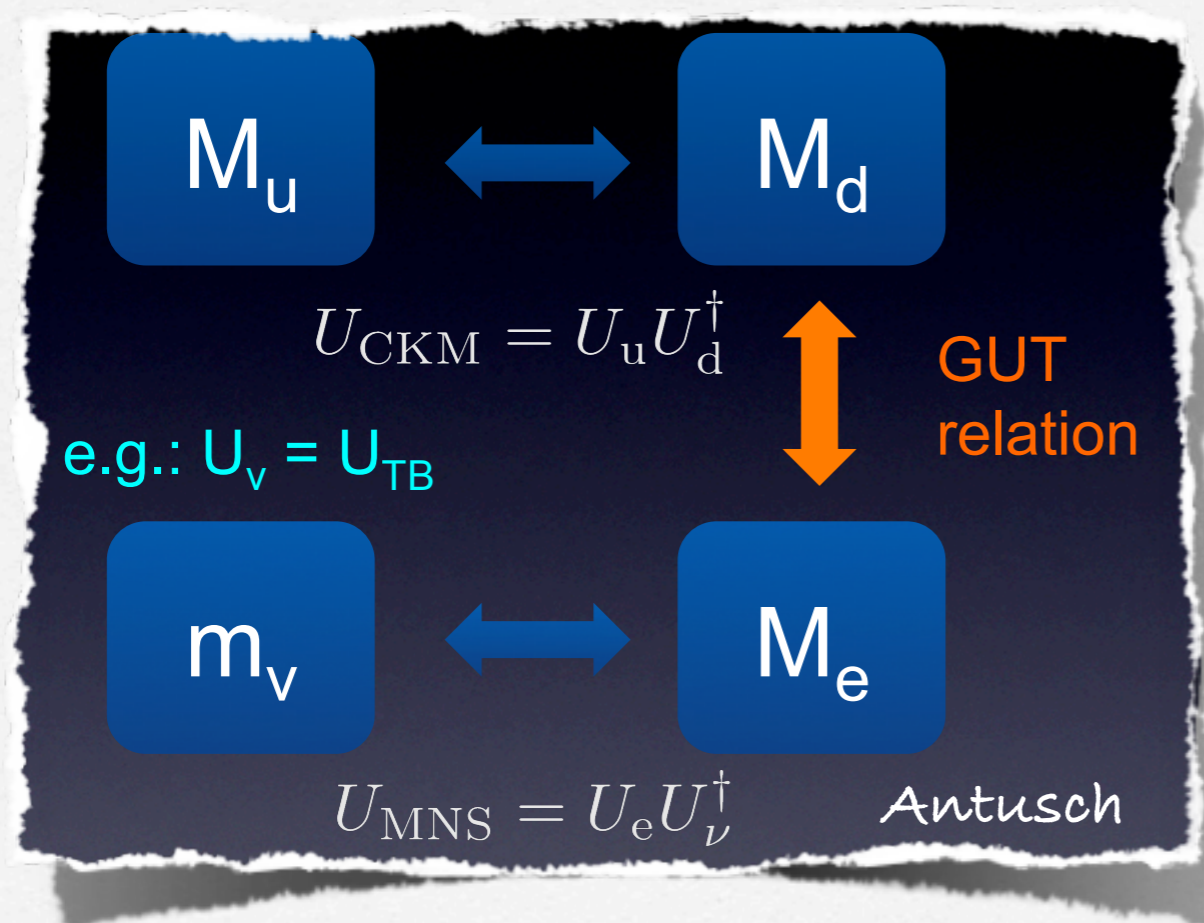
□ Neutrino mass suggests connection with GUT scale physics



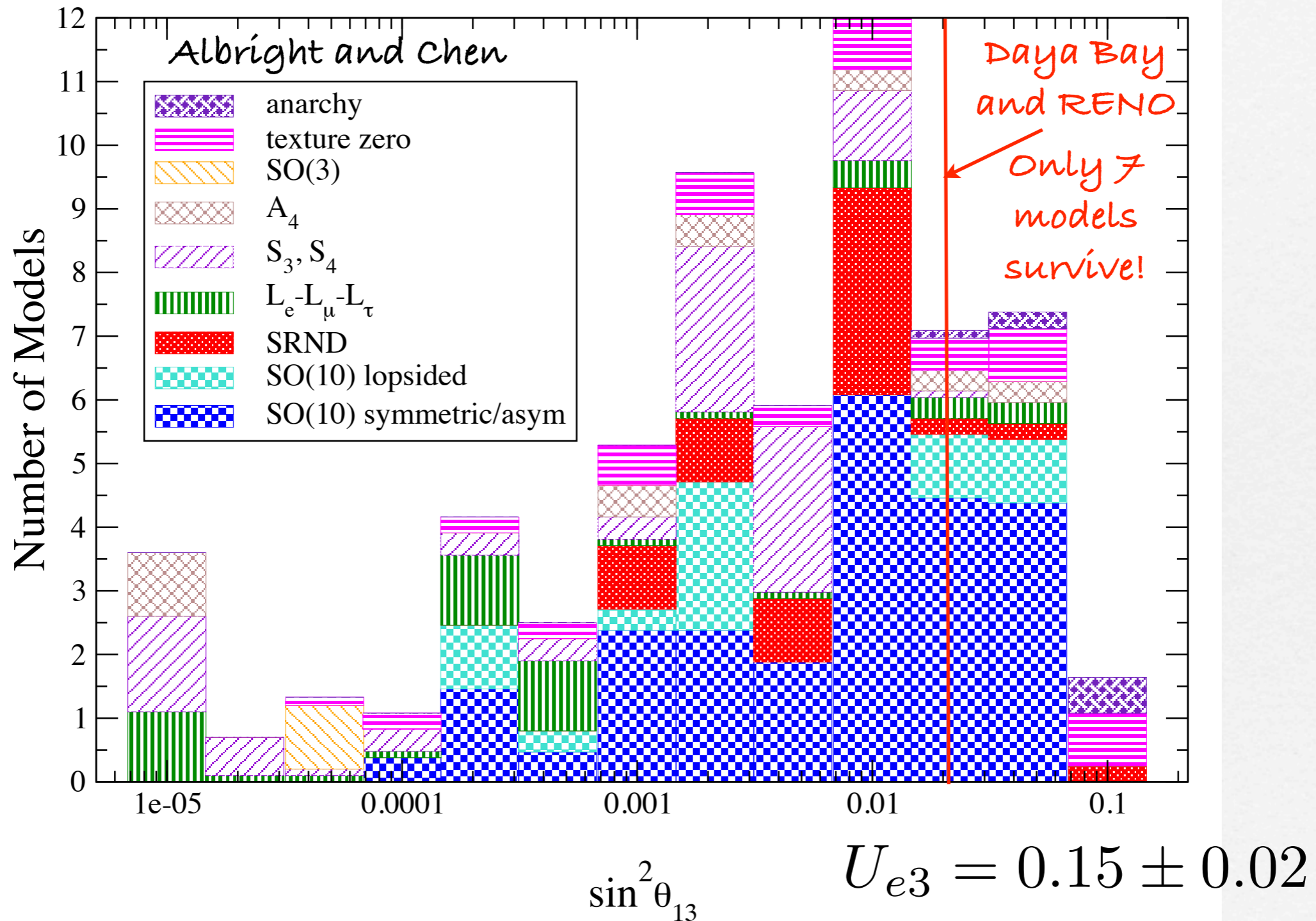
GUTs

Talks by Antusch, Mohapatra

Possible new combinations of Clebsches leading to large Theta13: Antusch, Maurer ('11)
Mazocca, Petcov, Romanino, Spinrath ('11)
Meroni, Petcov, Spinrath ('12)



Models Survey c.2006



Theory Road Map

Daya Bay/RENO



Symmetry

Anarchy

Direct

Indirect

Landscape



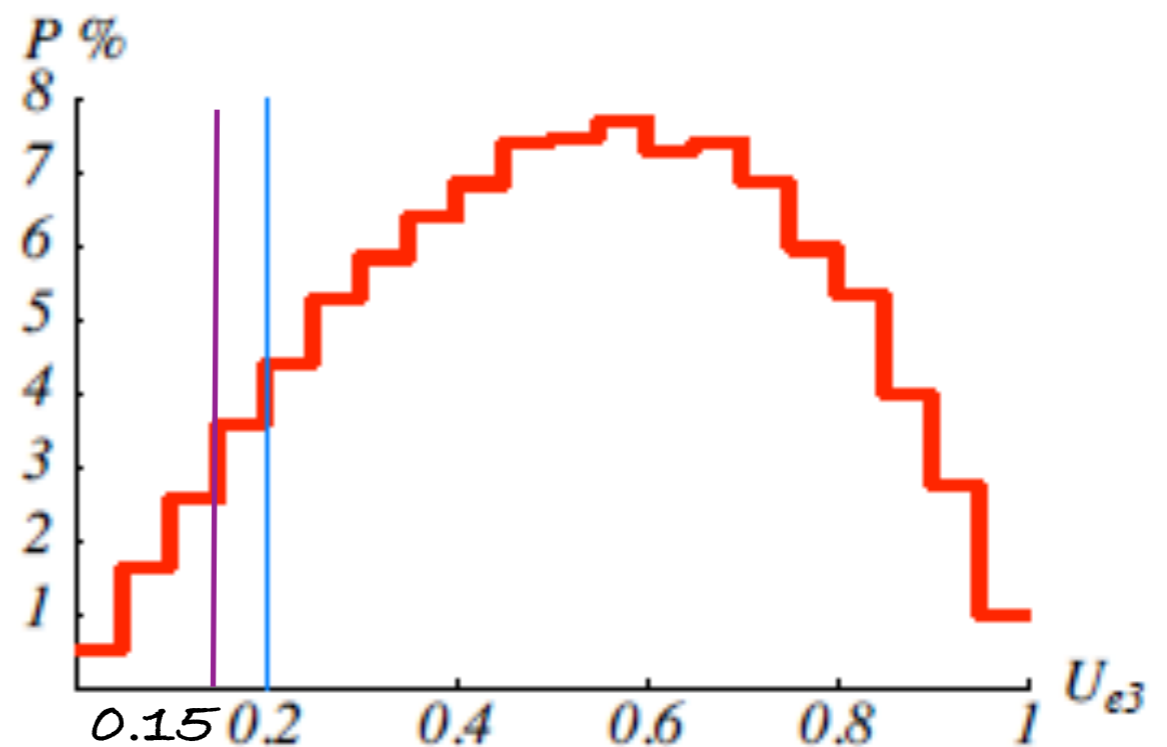
Anarchy

Hall, Murayama,
de Gouvea



- Anarchy: all angles are "large" and unpredictable, so expect $\sin\theta_{13} \sim 0.5$
- Hence larger reactor angle is good news
- Problem is that reactor angle is not that large...
- Also Anarchy not very predictive c.f. landscape

Altarelli, Feruglio, Masina



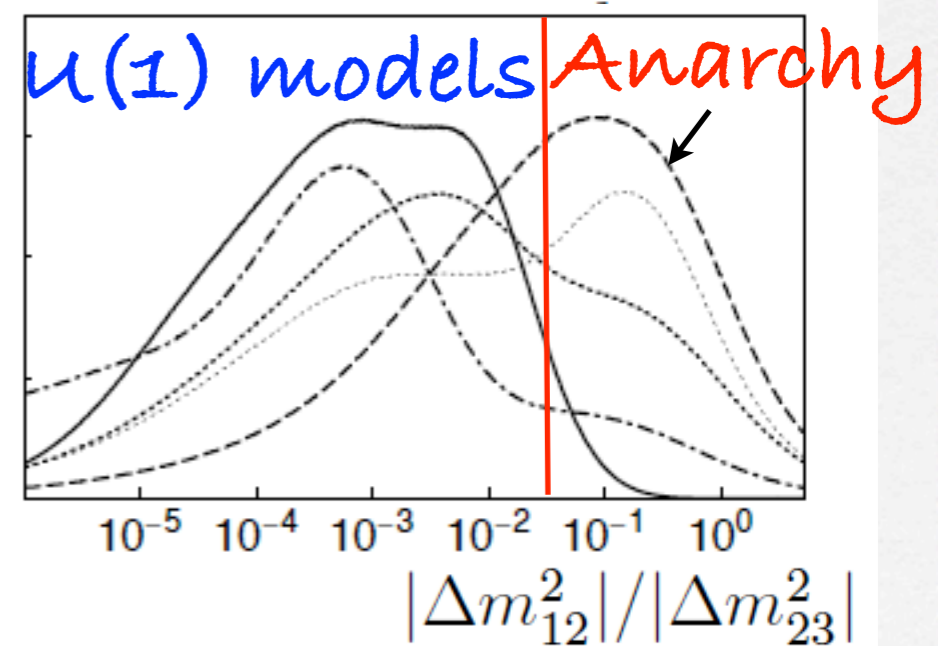
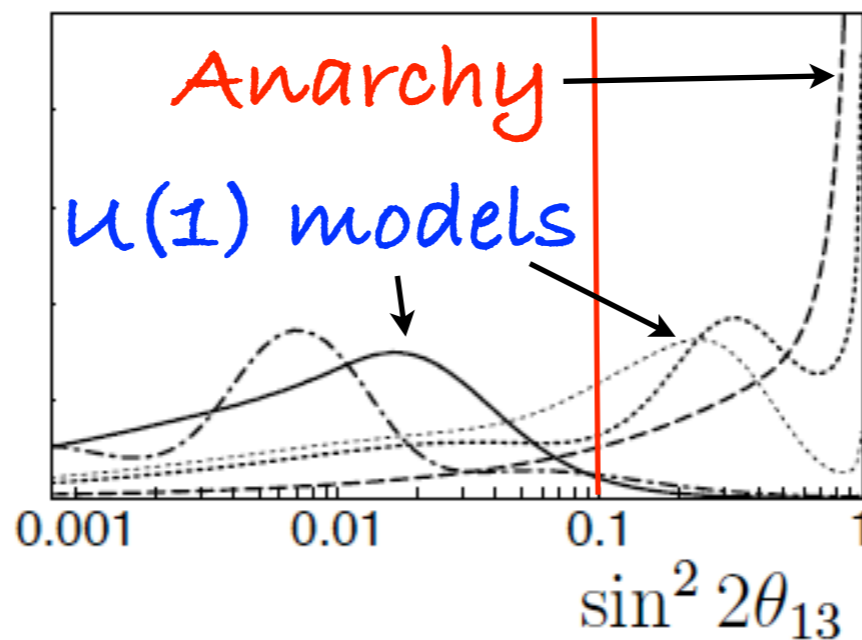
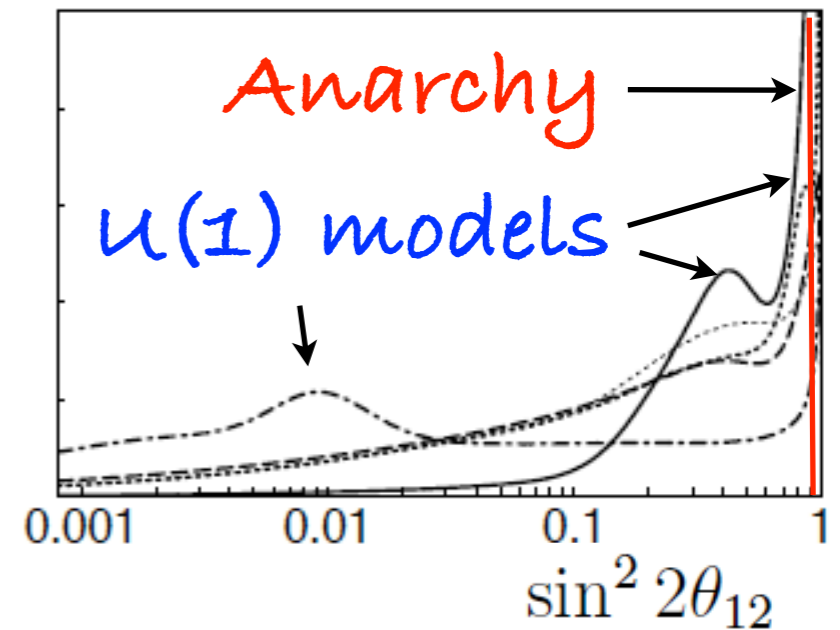
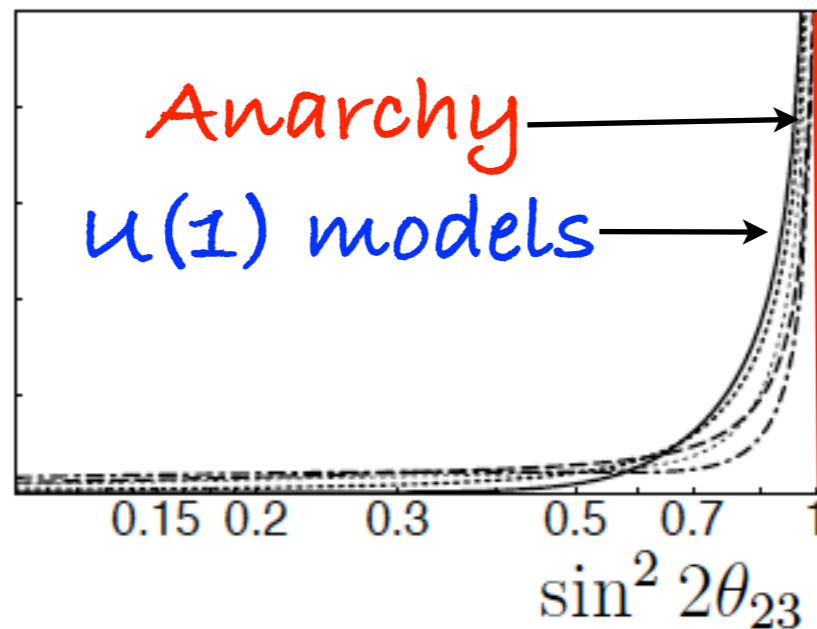
$$U_{e3} = 0.15 \pm 0.02$$

U(1) family symmetry helps...

Hirsch and King ('01)

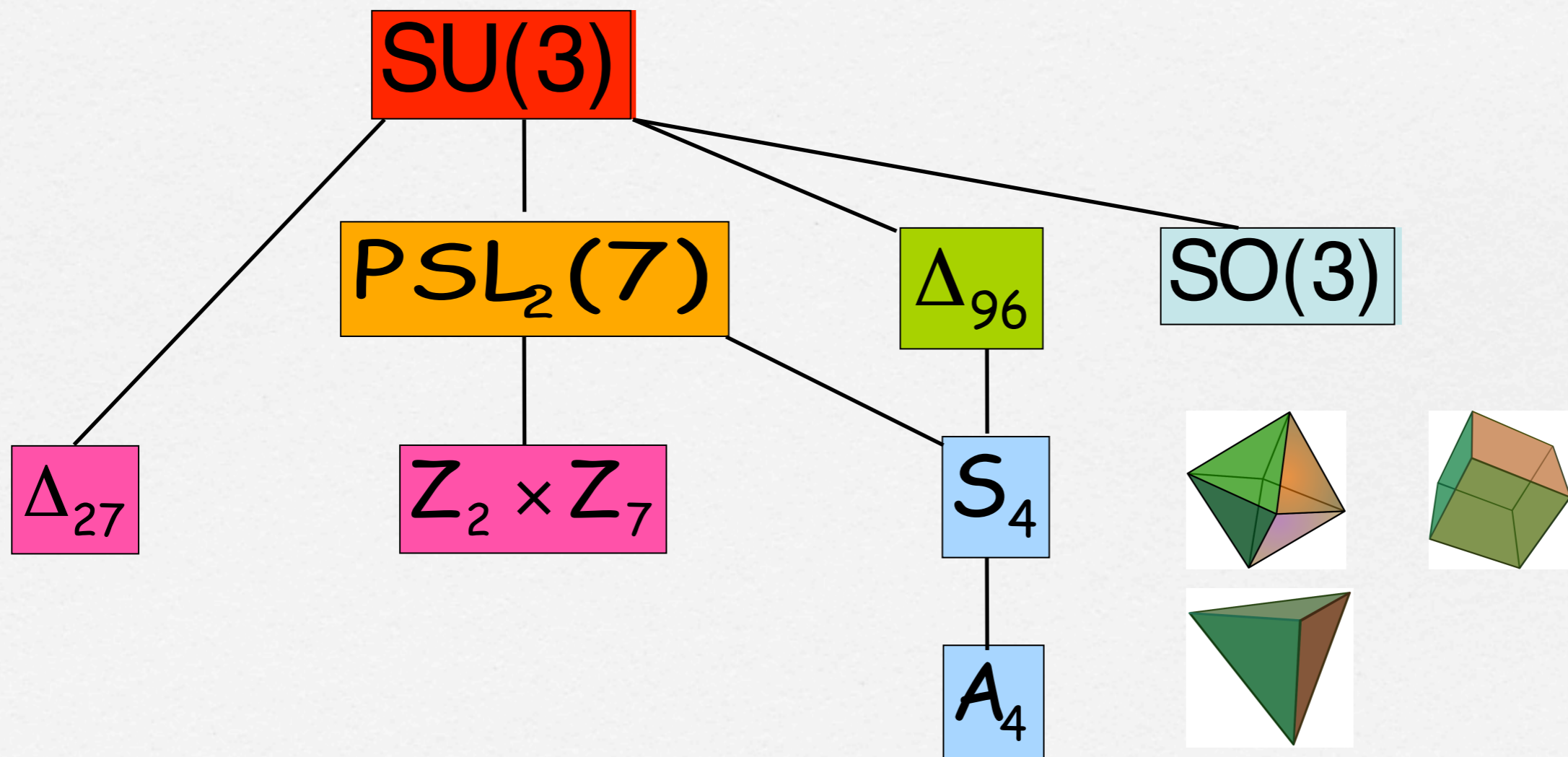
- lepton (quark) generations labelled by U(1) family symmetry

- $\sin^2 2\theta_{13}$ may peak at lower values



Family Symmetry

Family Symmetries G_F
which contain triplet reps
(three families in a triplet)



Partial list of authors who have worked on symmetry approach to large θ_{13}

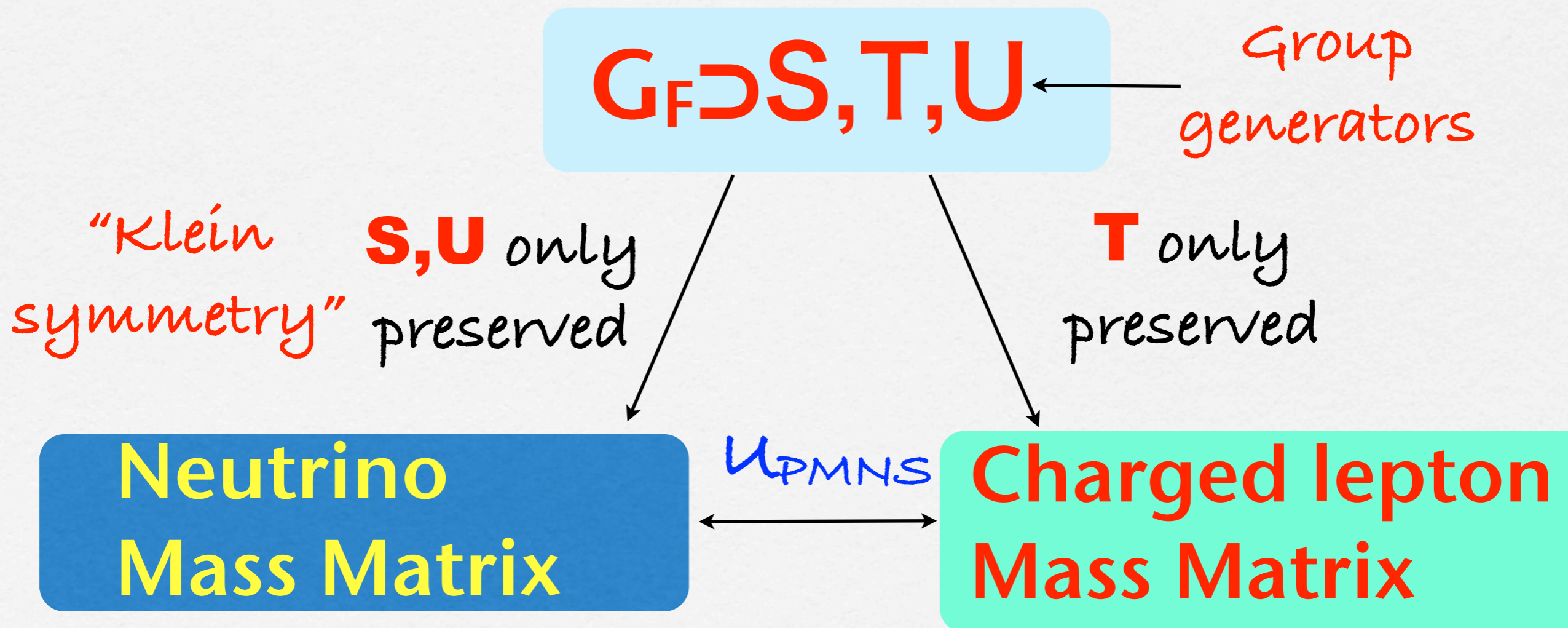
A. Adulpravitchai, Y. H. Ahn, C. H. Albright, G. Altarelli, S. Antusch, A. Aranda, T. Araki, F. Bazzocchi, W. Buchmuller, P. S. Bhupal Dev, G. C. Branco, Q.-H. Cao, H.-Y. Cheng, I. K. Cooper, S. Dev, G. Blankenburg, C. Bonilla, F. Gonzalez Canales, W. Chao, J.-M. Chen, M.-C. Chen, X. Chu, A. Datta, K. N. Deepthi, M. Dhen, D. A. Dicus, G.-J. Ding, P. V. Dong, V. Domcke, L. Dorame, B. Dutta, D. A. Eby, L. Everett, R. P. Feger, F. Feruglio, P. Ferreira, P. H. Frampton, M. Fukugita, R. R. Gautam, S. - F. Ge, D. K. Ghosh, R. Gonzalez Felipe, S. Gollu, S. Gupta, W. Grimus, C. Gross, N. Haba, C. Hagedorn, T. Hambye, J. Kersten, J. E. Kim, Y. Koide, K. Hashimoto, K. Harigaya, H. -J. He, X. -G. He, J. Heek, D. Hernandez, M. Holthausen, R. S. Hundi, M. Ibe, H. Ishimori, F. R. Joaquim, A. S. Joshipura, S. K. Kang, T. W. Kephart, S. Khalil, S. F. King, T. Kobayashi, S. Kumar, L. Lavoura, X.-Q. Li, H. N. Long, P. O. Ludl, C. Luhn, B. Q. Ma, E. Ma, S. K. Majee, K.T. Mahanthappa, D. Marzocca, V. Maurer, D. Meloni, A. Merle, A. Meroni, R. Mohanta, R. N. Mohapatra, E. Molinaro, A. Mondragon, M. Mondragon, S. Morisi, C. H. Nam, H. Nishiura, S. Oh, H. Okada, K. M. Patel, K. M. Parattu, E. Peinado, S. T. Petcov, N. Qin, A. Rashed, W. W. Repko, A. D. Rojas, W. Rodejohann, A. Romanino, G. G. Ross, S. Rigolin, M. A. Schmidt, K. Schmitz, M. Severson, M.-S. Seo, H. Serodio, Y. Shimizu, J. I. Silva-Marcos, L. Singh, K. Siyeon, C. Sluka, A. Yu. Smirnov, M. Spinrath, E. Stamou, A. J. Stuart, R. Takahashi, M. Tanimoto, R. d. A. Toorop, J. W. F. Valle, I. d. M. Varzielas, L. Velasco, V. V. Vien, B. Wang, Q. Wang, A. Watanabe, D. Wegman, A. Wingerter, Yue-Liang Wu, Z. -Z. Xing, T. T. Yanagida, W.-M. Yang, B. Zaldívar, F. -R. Yin, A. Zee, H. Zhang, Y. -j. Zheng, J.-J. Zhong, S. Zhou, R. Zwicky, ...

(sincere apologies for incompleteness)

Altarelli, Feruglio, Ma, Hagedorn, Merlo, Luhn, ...

The Direct Approach

Family Symmetry G_F broken in special way
subgroups preserved in neutrino/charged lepton sectors



Altarelli, Feruglio,
Merlo, Hagedorn,
Luhn, King...

Direct Models

de Adelhart Toorop,
Feruglio, Hagedorn ('11)
Ding ('12),
King, Luhn, Stuart ('12)

Smaller groups
A4, S4, A5...

Simple LO Mixing
Patterns $\theta_{13} = 0$

S, U preserved in
Neutrino sector,
T preserved in
Charged Lepton

Larger groups
 $\Delta(96), \dots$

Richer LO Mixing
Patterns $\theta_{13} \neq 0$

T broken

Charged Lep
corrects

Solar Sum
Rules

U broken

Special
HO corrects

e.g. Tri-maximal

Atmospheric
Sum Rules

S, U broken

General
HO corrects

Unpredictive

Corrections not required
(but may be present)

Plus RG,
Canonical
Normalisation, ...

Simple LO Mixing Patterns

$$\theta_{13} = 0 \quad \theta_{23} = 45^\circ$$

□ Bimaximal

V. Barger, S. Pakvasa, T. Weiler and K. Whisnant

$$U_{BM} = \begin{pmatrix} \frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} & 0 \\ -\frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \\ \frac{1}{2} & -\frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix} P \quad \theta_{12} = 45^\circ$$

□ Tri-bimaximal

Harrison, Perkins and Scott

$$U_{TB} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} P \quad \theta_{12} = 35.26^\circ$$

□ Golden ratio

Datta, Ling, Ramond; Kajirama, Raidal, Strumia; Everett, Stuart, Ding; Feruglio, Paris

$$\phi = \frac{1 + \sqrt{5}}{2}$$

$$U_{GR} = \begin{pmatrix} c_{12} & s_{12} & 0 \\ -\frac{s_{12}}{\sqrt{2}} & -\frac{c_{12}}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{s_{12}}{\sqrt{2}} & -\frac{c_{12}}{\sqrt{2}} & \frac{1}{\sqrt{2}} \end{pmatrix} P$$
$$\tan \theta_{12} = \frac{1}{\phi} \quad \theta_{12} = 31.7^\circ$$

Large Charged Lepton Corrections to the rescue

$$\theta_{12}^e \approx \theta_C$$



- BM, TBM, GR might only apply to neutrino mixing and $U_{PMNS} = U_e U_\nu^\dagger$ implies $\theta_{13} \approx \frac{\theta_{12}^e}{\sqrt{2}}$

Solar Sum Rules

- Sum Rule: King ('05); Masina ('05); Antusch, King ('05)

Charged Lepton Corrections: King ('02), Frampton, Petcov, Rodejohann ('04), Altarelli, Feruglio, Masina ('04), Antusch, King ('04), Ferrandis, Pakvasa ('04), Feruglio ('05), Datta, Everett, Ramond ('05), Mohapatra, Rodejohann ('05) Antusch, Maurer ('11) Mazocca, Petcov, Romanino, Spinrath ('11)

- Bimaximal $\theta_{12} = 45^\circ + \theta_{13} \cos \delta \rightarrow \delta \approx \pi$

- Tri-bimaximal $\theta_{12} = 35^\circ + \theta_{13} \cos \delta \rightarrow \delta \approx \pm \frac{\pi}{2}$

- Golden ratio $\theta_{12} = 32^\circ + \theta_{13} \cos \delta$

Experiment $\theta_{12} = 34^\circ \pm 1^\circ$

$\theta_{13} = 9^\circ \pm 1^\circ$

Tri-Bimaximal Parametrisation

$$U_{\text{PMNS}} \approx \begin{pmatrix} \frac{2}{\sqrt{6}}(1 - \frac{1}{2}s) & \frac{1}{\sqrt{3}}(1 + s) & \frac{1}{\sqrt{2}}re^{-i\delta} \\ -\frac{1}{\sqrt{6}}(1 + s - a + re^{i\delta}) & \frac{1}{\sqrt{3}}(1 - \frac{1}{2}s - a - \frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}}(1 + a) \\ \frac{1}{\sqrt{6}}(1 + s + a - re^{i\delta}) & -\frac{1}{\sqrt{3}}(1 - \frac{1}{2}s + a + \frac{1}{2}re^{i\delta}) & \frac{1}{\sqrt{2}}(1 - a) \end{pmatrix} P$$

$$\sin \theta_{12} = \frac{1}{\sqrt{3}}(1 + s), \quad \sin \theta_{23} = \frac{1}{\sqrt{2}}(1 + a), \quad \sin \theta_{13} = \frac{r}{\sqrt{2}}$$

$$s = -0.03 \pm 0.03 \quad a = -0.02 \pm 0.10 \quad r = 0.22 \pm 0.02$$

s = solar **a = atmospheric** **r = reactor**

Allows for deviations from TB mixing

E.g. TB solar sum rule recast as $s = r \cdot \cos \delta$

Tri-bimaximal Hydras



□ Tri-bimaximal
($s=a=r=0$)

$$U_{TB} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} P$$

Harrison, Perkins, Scott

□ Tri-bimaximal-reactor
($s=a=0$)

$$U_{TBR} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} re^{-i\delta} \\ -\frac{1}{\sqrt{6}}(1 + re^{i\delta}) & \frac{1}{\sqrt{3}}(1 - \frac{1}{2} re^{i\delta}) & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}}(1 - re^{i\delta}) & -\frac{1}{\sqrt{3}}(1 + \frac{1}{2} re^{i\delta}) & \frac{1}{\sqrt{2}} \end{pmatrix} P$$

King; Antusch, Boudjemaa, King; Morisi, Patel, Peinado; Luhn, King

□ Tri-maximal 1
($s=0, a=r \cdot \cos\delta$)

$$U_{TM_1} = P' \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} re^{-i\delta} \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}}(1 - \frac{3}{2} re^{i\delta}) & \frac{1}{\sqrt{2}}(1 + re^{-i\delta}) \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}}(1 + \frac{3}{2} re^{i\delta}) & -\frac{1}{\sqrt{2}}(1 - re^{-i\delta}) \end{pmatrix} P$$

Lam; Albright, Rodejohann; Antusch, King, Luhn, Spinrath

□ Tri-maximal 2
($s=0, a=-r/2 \cdot \cos\delta$)

Haba, Watanabe, Yoshioka; He, Zee; Grimus, Lavoura; Albright, Rodejohann; King, Luhn

$$U_{TM_2} = P' \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} re^{-i\delta} \\ -\frac{1}{\sqrt{6}}(1 + \frac{3}{2} re^{i\delta}) & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}}(1 - \frac{1}{2} re^{-i\delta}) \\ -\frac{1}{\sqrt{6}}(1 - \frac{3}{2} re^{i\delta}) & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}}(1 + \frac{1}{2} re^{-i\delta}) \end{pmatrix} P$$

N.B. Atmospheric sum rules: $a=r \cdot \cos\delta$, $a=-r/2 \cdot \cos\delta$

Indirect Models

King, Ross, de Medeiros Varzielas, Antusch, Malinsky,...



Starting point is type I see-saw

$$m_{LR} = \begin{pmatrix} A_1 & B_1 & C_1 \\ A_2 & B_2 & C_2 \\ A_3 & B_3 & C_3 \end{pmatrix} \quad M_{RR} = \begin{pmatrix} M_1 & 0 & 0 \\ 0 & M_2 & 0 \\ 0 & 0 & M_3 \end{pmatrix}$$



$$m^v = \frac{AA^T}{M_1} + \frac{BB^T}{M_2} + \frac{CC^T}{M_3}$$

$$A^T = (A_1, A_2, A_3) \quad B^T = (B_1, B_2, B_3)$$

Promote the columns (A,B,C) to dynamical fields
 G_F yields special vacuum alignments, for example:

- (A,B,C) proportional to columns of PMNS called **Form Dominance (FD)**
Chen, King('09)
- $AA^T/M_1 \rightarrow 0$ gives hierarchy ($m_1 \rightarrow 0$) called **Sequential Dominance (SD)**
King('98,'02)
- SD with $B \sim (1,1,-1)$ and $C \sim (0,1,1)$ called **Constrained SD** gives TB Mixing
King('05)
- SD with $B \sim (1,1,-1)$ and $C \sim (r,1,1)$ called **Partially CSD** gives TBR mixing
King('09), King,Luhn('11)
- SD with $B \sim (1,2,0)$ and $C \sim (0,1,1)$ called **CSD2** gives TM1 mixing
Antusch, King, Luhn, Spinrath ('11)

Tri-bimaximal-Cabibbo Mixing

See also: Antusch, Gross, Maurer, Sluka 1205.1051; Relation also appears in "Quark Lepton Complementarity" ... : Minakata, Smirnov ('04); and "Cabibbo Haze" Everett, Ramond ('05)

Combine TB mixing with $\theta_{13} \approx \frac{\theta_C}{\sqrt{2}} \approx 9.2^\circ$

$$s_{13} = \frac{\lambda}{\sqrt{2}}, \quad s_{12} = \frac{1}{\sqrt{3}}, \quad s_{23} = \frac{1}{\sqrt{2}} \quad \lambda = 0.2253 \pm 0.0007$$

$$U_{TBC} \approx \begin{pmatrix} \sqrt{\frac{2}{3}}(1 - \frac{1}{4}\lambda^2) & \frac{1}{\sqrt{3}}(1 - \frac{1}{4}\lambda^2) & \frac{1}{\sqrt{2}}\lambda e^{-i\delta} \\ -\frac{1}{\sqrt{6}}(1 + \lambda e^{i\delta}) & \frac{1}{\sqrt{3}}(1 - \frac{1}{2}\lambda e^{i\delta}) & \frac{1}{\sqrt{2}}(1 - \frac{1}{4}\lambda^2) \\ \frac{1}{\sqrt{6}}(1 - \lambda e^{i\delta}) & -\frac{1}{\sqrt{3}}(1 + \frac{1}{2}\lambda e^{i\delta}) & \frac{1}{\sqrt{2}}(1 - \frac{1}{4}\lambda^2) \end{pmatrix}$$

Describes all current data!



Obtained from PCSD with $B \sim (1, 1, -1)$ and $C \sim (\lambda, 1, 1)$

Conclusions

- With large θ_{13} , still two theory approaches: Symmetry or Anarchy
- Family Symmetry may be implemented Directly or Indirectly
- Simplest Direct models A_4, S_4, A_5 with S, U and T conservation predict Bimaximal, Tri-bimaximal, Golden Ratio at LO
- However T broken in GUT models due to Charged Lepton corrections, (Cabibbo-like charged lepton angle required) imply solar sum rules
- U breaking at HO leads to TM_2 mixing, atmospheric sum rules
- Larger Finite Groups such as Delta (96) predict e.g. $\theta_{13} \sim 12^\circ$ at LO
- Indirect family symmetry models can lead to TM_1 or TBC mixing
- vital to measure the mixing angles and CP phase delta to good precision to test the sum rules, hence discriminate between models, decide if the universe is based on Symmetry or if Anarchy Rules

Summary of Sum Rule Predictions

- Quark-Lepton Complementarity $\theta_{12} + \theta_C = 45^\circ$
- Solar sum rules
 - Bimaximal $\theta_{12} = 45^\circ + \theta_{13} \cos \delta$
 - Tri-bimaximal $\theta_{12} = 35^\circ + \theta_{13} \cos \delta$
 - Golden Ratio $\theta_{12} = 32^\circ + \theta_{13} \cos \delta$
- Atm. sum rules
 - Tri-bimaximal-Cabibbo $\theta_{12} = 35^\circ \quad \theta_{23} = 45^\circ$
 $\theta_{13} = \theta_C / \sqrt{2} = 9.2^\circ$
 - Trimaximal₁ $\theta_{23} = 45^\circ + \sqrt{2} \theta_{13} \cos \delta$
 - Trimaximal₂ $\theta_{23} = 45^\circ - \frac{\theta_{13}}{\sqrt{2}} \cos \delta$

Plus HO corrections...

Plus charged Lepton Corrections...

Now that θ_{13} is measured these predict $\cos \delta$