

One-loop quantum gravity and HS from a worldline viewpoint

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F.Bastianelli, R.B., O.Corradini, E.Latini, arXiv:1210.4649

F.Bastianelli, R.B., arXiv:1304.7135

Motivations

- The worldline formalism proved to be useful to compute various one-loop quantities of QFT's, such as n-point functions, propagators, effective actions and so on
- $O(N)$ spinning particles in flat space describe quite efficiently higher spin fields (spin $\frac{N}{2}$ in $D = 4$),
(Berezin and Marinov, Gershun and Tkach, Howe, Penati, Pernici and Townsend,...)
and can be quantized on (A)dS spaces
(Kuzenko, Yarevskaya,...)
- The $O(4)$ model correctly describes massless spin 2 particles in $D = 4$ on (A)dS backgrounds (possible extension to conformally flat), but does not admit a coupling to general backgrounds
- Alternative route to give a worldline description of quantum gravity at one loop

Outline

- The $O(N)$ spinning particle in flat space
- Quantized $O(N)$ spinning particle in (A)dS and effective action
- One-loop Einstein's gravity
- Worldline description of gravity dof
- First quantized approach to the effective action

O(N) spinning particle in flat space

Consider, in D dimensional flat spacetime, the worldline action

$$S = \int dt \left[p_\mu \dot{x}^\mu + \frac{i}{2} \psi_i^\mu \dot{\psi}_{\mu i} - \frac{1}{2} p^2 \right]$$

where $x^\mu(t)$ are target space coordinates, $p_\mu(t)$ conjugate momenta and $\psi_i^\mu(t)$, $i = 1, \dots, N$ are real worldline fermions.

It enjoys O(N)-extended WL supersymmetry, generated by

$$H = \frac{p^2}{2}, \quad Q_i = \psi_i^\mu p_\mu, \quad J_{ij} = \frac{i}{2} [\psi_i^\mu, \psi_{\mu j}]$$

The generators obey the following supersymmetry algebra

$$\{Q_i, Q_j\} = 2 \delta_{ij} H$$

$$[J_{ij}, Q_k] = i \delta_{jk} Q_i - i \delta_{ik} Q_j$$

$$[J_{ij}, J_{kl}] = i \delta_{jk} J_{il} - i \delta_{ik} J_{jl} - i \delta_{jl} J_{ik} + i \delta_{il} J_{jk}$$

- The algebra is first class \rightarrow It can be gauged to make the (super)symmetries local
- One introduces WL gauge fields $e(\tau)$, $\chi_i(\tau)$ and $a_{ij}(\tau)$ to be coupled to the symmetry generators
- By doing so one ends up with the O(N) spinning particle action:

$$S = \int_0^1 dt \left[p_\mu \dot{x}^\mu + \frac{i}{2} \psi_i^\mu \dot{\psi}_{\mu i} - e H - i \chi_i Q_i - \frac{1}{2} a_{ij} J_{ij} \right]$$

- Dirac quantization \rightarrow The constraints J_{ij} , Q_i and H have to annihilate physical states

Dirac constraints J_{ij} , Q_i and $H \Rightarrow$ physical Hilbert space consists of linear, trace-free (or gamma trace-free) HS curvatures

- In $D = 4$ and for $N = 2s$ they give Fronsdal equations for spin s fields with compensators

Francia, Sagnotti; 2003

Bastianelli, Corradini, Latini; 2008

- For general D one gets Fronsdal-Labastida equations for tensors with rectangular $(D/2 - 1) \times s$ Young tableau

Bastianelli, Corradini, Latini; 2008

- For odd N we have Fang-Fronsdal equations for fermionic HS fields

Corradini; 2010

HS on (A)dS backgrounds

- Trying to couple the $O(N)$ spinning particle to a curved space, the SUSY algebra is not first class, obstructed by target space curvature
- On conformally flat spaces the algebra is first class, but with structure functions \rightarrow Cumbersome to quantize (BV...)
- On Maximally symmetric spaces the algebra becomes quadratic in constraints \rightarrow Simpler BRST quantization

Bastianelli, Corradini, Latini; 2008

$$R_{abcd} = b(\eta_{ac}\eta_{bd} - \eta_{ad}\eta_{bc}) \quad \rightarrow$$

$$\{Q_i, Q_j\} = 2\delta_{ij}H - \frac{b}{2} (J_{ik}J_{jk} + J_{jk}J_{ik} - \delta_{ij}J_{kl}J_{kl})$$

HS Effective action on (A)dS

- One-loop effective action given by the worldline path integral

$$\Gamma[g] = \int_{S^1} \frac{\mathcal{D}x \mathcal{D}\psi_i \mathcal{D}e \mathcal{D}\chi_i \mathcal{D}a_{ij}}{\text{Vol}(\text{Gauge})} e^{-S} \quad \text{Diagram: A heptagon with wavy lines on each side and vertices labeled 6, 9, 9, 6, 9, 9, 6.$$

- After gauge fixing WL symmetries one has the Heat Kernel expansion

$$\Gamma[g] = \int_0^\infty \frac{dT}{T} \int \frac{d^D x \sqrt{|g|}}{(2\pi T)^{D/2}} a_0 \langle\langle e^{-S_{\text{int}}} \rangle\rangle$$

$$a_0 \langle\langle e^{-S_{\text{int}}} \rangle\rangle = a_0 \left(1 + v_1 R T + v_2 R^2 T^2 + \dots \right)$$

Simplest case: $D = 4$ and even $N = 2s$

The SDW coefficients are defined as

$$\Gamma[g] = \int_0^\infty \frac{dT}{T} \int \frac{d^4x \sqrt{|g|}}{(2\pi T)^2} a_0 \left(1 + v_1 R T + v_2 R^2 T^2 + \dots \right)$$

and we get

$$a_0 = 2 - \delta_{s,0}, \quad v_1 = -\frac{s^2}{6}, \quad v_2 = -\frac{1}{8640} + \frac{s^2}{288} - \frac{s^4}{144}$$

Bastianelli, R.B., Corradini, Latini; arXiv:1210.4649

- Explicit result also for half-integer spins and arbitrary even dimensions

One-loop quantum gravity

- Obstruction to describe quantum spin 2 on generic gravitational background
- Similar problem in coupling the $O(2)$ particle (vector fields in spacetime) to a gauge background was treated with BRST techniques

Dai, Huang, Siegel, 2008

that look not suitable for the gravitational case

- We undertake an alternative route, starting directly from QFT

One-loop Einstein gravity

- Consider four dimensional (euclidian) Einstein gravity with cosmological constant

$$S[G] = -\frac{1}{\kappa^2} \int d^4x \sqrt{G} [R(G) - 2\Lambda]$$

- We employ background field method splitting $G_{\mu\nu} = g_{\mu\nu} + h_{\mu\nu}$. At one-loop order we look for the part quadratic in quantum fluctuations $h_{\mu\nu}$:

$$S_2 = - \int d^4x \sqrt{g} \left\{ \frac{1}{4} h^{\mu\nu} (\nabla^2 + 2\Lambda) h_{\mu\nu} - \frac{1}{8} h (\nabla^2 + 2\Lambda) h + \frac{1}{2} (\nabla^\nu h_{\nu\mu} - \frac{1}{2} \nabla_\mu h)^2 \right. \\ \left. + \frac{1}{2} h^{\mu\lambda} h^{\nu\sigma} R_{\mu\nu\lambda\sigma} + \frac{1}{2} (h^{\mu\lambda} h^\nu_\lambda - h h^{\mu\nu}) R_{\mu\nu} + \frac{1}{8} (h^2 - 2h^{\mu\nu} h_{\mu\nu}) R \right\}$$

One-loop Einstein gravity

We choose covariant gaussian averaged gauge fixing:

$$S_{\text{GF}} = \frac{1}{2} \int d^4x \sqrt{g} (\nabla^\nu h_{\nu\mu} - \frac{1}{2} \nabla_\mu h)^2$$

- The quadratic part of the ghost action then reads

$$S_{bc} = - \int d^4x \sqrt{g} b^\mu \left[\nabla^2 c_\mu + R_{\mu\nu} c^\nu \right]$$

- We split traceless and trace part of metric fluctuations:
 $h_{\mu\nu} = \bar{h}_{\mu\nu} + \frac{1}{4} g_{\mu\nu} h$, in order to avoid worldline non-perturbative vertices

One-loop Einstein gravity

The gauge fixed action reads

$$\begin{aligned} S_q = \int d^4x \sqrt{g} \left\{ -\frac{1}{4} \bar{h}^{\mu\nu} (\nabla^2 + 2\Lambda) \bar{h}_{\mu\nu} \right. \\ - \frac{1}{2} \bar{h}^{\mu\lambda} \bar{h}^{\nu\sigma} R_{\mu\nu\lambda\sigma} - \frac{1}{2} \bar{h}^{\mu\lambda} \bar{h}^{\nu}_{\lambda} R_{\mu\nu} - \frac{1}{4} \bar{h}^{\mu\nu} \bar{h}_{\mu\nu} R \\ \left. + \frac{1}{16} h(\nabla^2 + 2\Lambda)h - b^\mu \left[\nabla^2 c_\mu + R_{\mu\nu} c^\nu \right] \right\} \end{aligned}$$

- As it is well known, the scalar piece has the wrong sign. We choose the prescription given by Gibbons and Hawking to Wick rotate its path integral
- The one-loop effective action splits in three contributions:

$$\Gamma[g] = \frac{1}{2} \left\{ \text{Tr}_{TT} \ln [K_{\mu\nu,\lambda\sigma} + M_{\mu\nu,\lambda\sigma}] + \text{Tr}_S \ln \left[-\frac{1}{2} (\nabla^2 + 2\Lambda) \right] - 2 \text{Tr}_V \ln [G_\nu^\mu] \right\}$$

Worldline description

- Use Schwinger's proper time integral to exponentiate logs:

$$\Gamma[g] = -\frac{1}{2} \int_0^\infty \frac{dT}{T} \left\{ \text{Tr}_{\mathcal{T}\mathcal{T}} \left[e^{-T(\hat{K} + \hat{M})} \right] + \text{Tr}_{\mathcal{S}} \left[e^{\frac{T}{2}(\nabla^2 + 2\Lambda)} \right] - 2\text{Tr}_{\mathcal{V}} \left[e^{-T\hat{G}} \right] \right. \\ \left. - \frac{1}{2} \{ \Gamma_2 + \Gamma_0 - 2\Gamma_1 \} \right.$$

- Γ_0 is represented by an ordinary bosonic QM path integral
- For Γ_2 and Γ_1 we need two worldline actions containing traceless symmetric tensors and vectors, respectively, in their Hilbert spaces
- Their quantum hamiltonians have to be $\hat{K} + \hat{M}$ and \hat{G} , respectively

Tensor model

Graded phase space: bosonic coordinates and momenta $x^\mu(t)$, $p_\mu(t)$ and WL fermions that are **symmetric traceless tensors** in target space $\psi^{ab}(t)$, $\bar{\psi}^{ab}(t)$

- Hilbert space:

$$|\bar{h}\rangle = h(x) + \bar{h}_{ab}(x)\psi^{ab} + \frac{1}{2}\phi_{(ab)(cd)}(x)\psi^{ab}\psi^{cd} + \dots$$

we need to project away the unwanted fields \rightarrow gauge the $U(1)$ fermion number

- Covariant momenta:

$$\pi_\mu = p_\mu - i\omega_{\mu ab}\psi^{ac}\bar{\psi}^b$$

- Hamiltonian:

$$\hat{H} = \frac{1}{2}g^{-1/4}\pi_\mu g^{\mu\nu}g^{1/2}\pi_\nu g^{-1/4} - \Lambda - \frac{1}{2}R_{abcd}\psi^{ac}\bar{\psi}^{bd} - \frac{1}{2}R_{ab}\psi^a\bar{\psi}^b + \frac{1}{2}R$$

it is indeed $\hat{K} + \hat{M}$

Tensor model

Worldline action

$$S_{\text{TT}} = \int_0^1 d\tau \left[\frac{1}{2T} g_{\mu\nu} \dot{x}^\mu \dot{x}^\nu + \frac{1}{2T} \bar{\psi}_{ab} (D_\tau + iA) \psi^{ab} - \frac{1}{2} R_{abcd} \psi^{ac} \bar{\psi}^{bd} \right. \\ \left. - \frac{1}{2} R_{ab} \psi^a \cdot \bar{\psi}^b - T \left(\frac{3}{8} R + \Lambda \right) \right] - \frac{7i}{2} \int_0^1 d\tau A$$

- U(1) gauge field $A(\tau)$ to constrain fermion number J , CS term undoes normal ordering constant in J
- DR \rightarrow we added the corresponding counterterm $V = -\frac{7}{8} R$
- Path integral on the circle gives Γ_2

$$\Gamma_2 = \int_0^\infty \frac{dT}{T} \int_0^{2\pi} \frac{d\phi}{2\pi} \int_P \mathcal{D}X \int_A D\bar{\psi} D\psi e^{-S_{\text{TT}}} \\ = \int_0^\infty \frac{dT}{T} e^{T\Lambda} \int \frac{d^4x \sqrt{g}}{(2\pi T)^2} \langle\langle e^{-S_{\text{int}}} \rangle\rangle$$

Vector model

Graded phase space: coordinates and momenta $x^\mu(t)$, $p_\mu(t)$ and vector WL fermions $\lambda^a(t)$, $\bar{\lambda}^a(t)$

- Hilbert space:

$$|v\rangle = v(x) + v_a(x)\lambda^a + \frac{1}{2} v_{ab}(x)\lambda^a\lambda^b + \dots$$

project away the unwanted fields \rightarrow gauge the $U(1)$ fermion number

- Covariant momenta:

$$\pi_\mu = p_\mu - i\omega_{\mu ab}\lambda^a\bar{\lambda}^b$$

- Hamiltonian:

$$\hat{H} = \hat{G} = \frac{1}{2} g^{-1/4} \pi_\mu g^{\mu\nu} g^{1/2} \pi_\nu g^{-1/4} - \frac{1}{2} R_{ab} \lambda^a \bar{\lambda}^b$$

Vector model

Worldline action

$$S_V = \int_0^1 d\tau \left[\frac{1}{2T} g_{\mu\nu} \dot{x}^\mu \dot{x}^\nu + \frac{1}{T} \bar{\lambda}_a (D_\tau + ia) \lambda^a - \frac{1}{2} R_{ab} \lambda^a \bar{\lambda}^b - \frac{3T}{8} R \right] \\ - i \int_0^1 d\tau a$$

- U(1) gauge field $a(\tau)$ constrains fermion number to one, CS undoes normal ordering constant in it
- DR \rightarrow counterterm $V = -\frac{3}{8} R$
- Path integral on the circle gives Γ_1

$$\Gamma_1 = \int_0^\infty \frac{dT}{T} \int_0^{2\pi} \frac{d\theta}{2\pi} \int_P \mathcal{D}x \int_A D\bar{\lambda} D\lambda e^{-S_V} \\ = \int_0^\infty \frac{dT}{T} \int \frac{d^4x \sqrt{g}}{(2\pi T)^2} \langle\langle e^{-S_{\text{int}}} \rangle\rangle$$

Gravity effective action

Proper time perturbative expansion, assembling the three contribution we get

$$\Gamma[g] \propto \int_0^\infty \frac{dT}{T} \int \frac{d^4x \sqrt{g}}{(2\pi T)^2} \left\{ 2 + T \left(-\frac{23}{6} R + 10\Lambda \right) \right. \\ \left. T^2 \left[\frac{53}{180} R_{\mu\nu\lambda\sigma}^2 - \frac{361}{360} R_{\mu\nu}^2 + \frac{43}{144} R^2 - \frac{19}{60} \nabla^2 R + 5\Lambda^2 - \frac{13}{6} R\Lambda \right] + \mathcal{O}(T^3) \right\}$$

It reproduces well known one-loop log divergencies of pure gravity:

$$\mathcal{L}_{div} \propto \frac{1}{8\pi^2} \frac{1}{\epsilon} \left\{ \frac{53}{90} E + \frac{7}{20} R_{\mu\nu}^2 + \frac{1}{120} R^2 + 10\Lambda^2 - \frac{13}{3} R\Lambda \right\}$$

Bastianelli, R.B.; arXiv:1304.7135

Conclusions and outlook

- $O(N)$ spinning particle \rightarrow spin $N/2$ field in spacetime
- Results for first SDW coefficients of quantum HS in AdS
- Possibly useful in HS/CFT duality?
- Extension to HS on conformally flat backgrounds?
- The $O(4)$ particle describes massless spin 2 fields but does not admit coupling to arbitrary background
- The alternative description does reproduce known SDW coefficients in $D=4$
- Desirable a gauge invariant description, not directly related to QFT

THANKS FOR YOU ATTENTION!