

Statistical mechanics for the phase separation of interacting self propelled particles

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Discovering the joys of research with Stefano

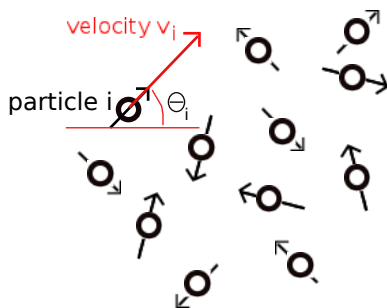


1998 World Cup, Italy vs France. L. Di Biagio misses his penalty kick.

Self propelled particles

- ▶ Particles with an internal source of free energy that they can convert into systematic movement.
- ▶ Used to model flocks of animals (from mammals to insects), bacteria, some artificial systems. . . **This will be a theoretical talk!**
- ▶ **Main question:** understand their collective properties
- ▶ Blooming field; many recent developments (I will not be able to cite all relevant contributions!).

The model system (2D)



- ▶ Point particles with an internal angular variable θ_i
- ▶ Move with speed v_i along direction $\theta_i +$ spatial noise
- ▶ the speed v_i may depend on the local density
- ▶ Particles interact: they tend to align locally

Microscopic equations (2D)

- Spatial variables: transport in direction θ_i (speed may depend on local density) + noise
- Angular variables: interactions promoting local alignment + noise

$$\begin{aligned}\dot{\mathbf{x}}_i &= v(n_i)\mathbf{u}(\theta_i) + \sqrt{2D_x}\sigma_i(t) \\ \dot{\theta}_i &= -\frac{\gamma}{n_i} \sum_{j \text{ neighbor of } i} \sin(\theta_i - \theta_j) + \sqrt{2D_\theta}\eta_i(t)\end{aligned}$$

with

n_i = local density

σ_i, η_i = gaussian white noises, unit covariance

Representative of a class of models with similar large scale properties.

Qualitative behavior

- Strong interactions, or "external field" \rightarrow local orientation order.
Not studied here.
- Weak interactions \rightarrow no local orientation order
Large scale dynamics = diffusive.
- v depends on the local density ρ \rightarrow effective diffusion coefficient depends on ρ
Possibility of "motility induced phase separation" (Cates, Tailleur)

Main questions

- A macroscopic description? Finite N fluctuations? Stationary measure? Probability distribution of the density?
- A very quick review
 - ▶ J. Toner, Y. Tu (1995): phenomenological hydrodynamical equations + noise introduced "by hand"
 - ▶ E. Bertin, M. Droz, G. Grégoire (2006): write a Boltzmann like equation + expansion close to the phase transition threshold → derivation of Toner-Tu like equations, without noise (many developments from there: Chaté et al., Marchetti et al., Ihle...)
 - ▶ Math. literature: P. Degond, S. Motsch (2007); Fokker-Planck like models (locally mean-field); far from the threshold
 - ▶ Keeping finite N fluctuations: J. Tailleur, M. Cates et al. (2008, 2011, 2013): without alignment promoting interactions; Bertin et al. (2013): derive a noise from the microscopic equations for nematics.

Our goals

1. start from microscopic equations
2. **derive** hydrodynamical equations and noise in a controlled way
Noise may have correlations \rightarrow important to have a microscopic derivation
3. exploit these results to study the dynamical fluctuations of the empirical density (cf Macroscopic Fluctuation Theory).
4. obtain large deviation estimate for the stationary spatial density ρ such as

$$\mathbb{P}(\rho \approx u) \asymp e^{NS[u]}$$

S = "entropy", or "quasi-potential".

Simple framework: aligning interactions below threshold for local order; density dependent speed (\rightarrow clustering possible).

Microscopic equations (simplified), adimensionalized

$$\frac{d\tilde{\mathbf{x}}_i}{d\tilde{t}} = \varepsilon \tilde{v}(n_i) \mathbf{u}(\theta_i) + \varepsilon \sqrt{2\tilde{D}_x} \tilde{\sigma}_i(\tilde{t}) \quad (1)$$

$$\frac{d\theta_i}{d\tilde{t}} = -\frac{\tilde{\gamma}}{n_i} \sum_{j \text{ neighbor of } i} \sin(\theta_i - \theta_j) + \sqrt{2} \eta_i(\tilde{t}), \quad (2)$$

with $\varepsilon = v_0/(LD_\theta)$, $\tilde{D}_x = D_x D_\theta / v_0^2$, $\tilde{\gamma} = \gamma / D_\theta$, $\tilde{t} = D_\theta t$.

Two important parameters:

\tilde{D}_x : ratio spatial diffusion/"active" diffusion

$\tilde{\gamma}$: strength of the aligning interaction

$\varepsilon =$ spatial time scale/angular time scale: small parameter

Strategy

Main object of interest: the empirical density

$$\rho(\mathbf{x}, \theta, t) = \frac{1}{N} \sum_i \delta(\mathbf{x} - \mathbf{x}_i(t))$$

Phase space empirical density

$$f(\mathbf{x}, \theta, t) = \frac{1}{N} \sum_i \delta(\mathbf{x} - \mathbf{x}_i(t)) \delta(\theta - \theta_i(t))$$

1. Write an equation for f that keeps finite N fluctuations (cf D. Dean 1996): in a sense **exact** in the large N limit
2. Use the time-scale separation to write an equation for ρ that keeps finite N fluctuations: hoped to be exact in a combined $\varepsilon \rightarrow 0$, $N \rightarrow \infty$ limit
3. Write a functional Fokker-Planck equation for $\mu_t[\rho]$, the pdf of ρ .
4. Look for a stationary solution of the form

$$\mu[\rho] \asymp e^{NS[\rho]}$$

A fluctuating non linear Fokker-Planck equation

$$\begin{aligned}
 \frac{\partial f}{\partial t} = & \underbrace{-\varepsilon \nabla \cdot (v(\rho) \mathbf{u}(\theta) f)}_{\text{transport}} + \underbrace{\frac{\gamma}{\rho} \frac{\partial}{\partial \theta} \left(f \int d\theta' \sin(\theta - \theta') f(\theta') \right)}_{\text{interaction}} \\
 & + \underbrace{\sqrt{\frac{2}{N}} \frac{\partial}{\partial \theta} \left(\eta(\mathbf{x}, \theta, t) \sqrt{f} \right) + \varepsilon \frac{\sqrt{2D_x}}{\sqrt{N}} \nabla_x \cdot \left(\vec{\sigma}(\mathbf{x}, \theta, t) \sqrt{f} \right)}_{\text{finite N fluctuations}} \\
 & + \underbrace{\frac{\partial^2 f}{\partial \theta^2} + \varepsilon^2 D_x \nabla_x^2 f}_{\text{angular and spatial diffusions}}
 \end{aligned}$$

Meaning? A dynamical large deviation principle (Dawson 1987).

$$\mathbb{P}(f_t \approx g_t) \asymp \exp(-N J_{[0, T]}[g]); \quad J_{[0, T]}[g] = \frac{1}{4} \int_0^T \|\partial_t g - VFP[g]\|_{-1, g}^2 dt$$

VFP = nonlinear Vlasov-Fokker-Planck operator = red terms

On the computations

- Local equilibrium + small deviation (order ε and $1/\sqrt{N}$ fluctuations)

$$f(\mathbf{x}, \theta, t) = \frac{1}{2\pi} \rho(\mathbf{x}, \varepsilon^\alpha t) + \delta f(\mathbf{x}, \theta, t)$$

- Equation for ρ : slow time scale, depends on δf

$$\frac{\partial \rho}{\partial t} = -\varepsilon \nabla \cdot (v \int u_\theta \delta f) + \varepsilon^2 D_x \nabla^2 \rho + \varepsilon \frac{\sqrt{2D_x}}{\sqrt{N}} \nabla \cdot (\vec{\xi}(x, y, t)) \quad (3)$$

ξ = noise, multiplicative in ρ .

- δf small \rightarrow obtained by solving a linearized equation
- Reintroduce into Eq.(3) \rightarrow the final equation, a fluctuating PDE for ρ .

Dynamical large deviation principle

- Fluctuating PDE for ρ

$$\frac{\partial \rho}{\partial t} = U[\rho](\vec{x}) + \frac{1}{\sqrt{N}} \nu(\vec{x}, t)$$

$$U[\rho](\vec{x}) = \frac{1}{2} \nabla \cdot \left(\frac{v(\rho)}{1 - \frac{\gamma}{2}} \nabla [v(\rho)\rho] \right) + D_x \nabla^2 \rho$$

$$\langle \nu(x, y, t) \nu(x', y', t') \rangle = D[\rho](\vec{x}, \vec{x}') \delta(t - t')$$

- The fluctuating PDE for ρ is a rephrasing of a dynamical large deviation principle "à la Dawson"

$$\mathbb{P}(\rho \approx u) \asymp \exp(-N I_{[0, T]}[u]) \text{ with } I_{[0, T]}[u] = \frac{1}{2} \int_0^T \|\partial_t u - U[u]\|_{-1, D}^2$$

- This kind of dynamical large deviation principle is the starting point for the macroscopic fluctuation theory (in this case, it is actually trivial...)

Yet another formulation: functional Fokker-Planck equation

- Ordinary stochastic differential equation for $x \in \mathbb{R}^d \rightarrow$ PDE (Fokker-Planck) for the pdf of x .
- Stochastic PDE for a field $\rho \rightarrow$ **functional** equation for $\mu_t[\rho]$, "pdf" of ρ .

$$\begin{aligned} \frac{\partial \mu_t}{\partial t} = & \underbrace{- \int d\vec{x} \frac{\delta}{\delta \rho(\vec{x})} (U[\rho](\vec{x}) \mu_t)}_{\text{drift part}} \\ & + \underbrace{\frac{1}{2N} \int d\vec{x} \frac{\delta}{\delta \rho(\vec{x})} \left\{ \int d\vec{x}' D[\rho](\vec{x}, \vec{x}') \frac{\delta}{\delta \rho(\vec{x}')} \mu_t \right\}}_{\text{diffusion part}} \end{aligned}$$

Results and discussion

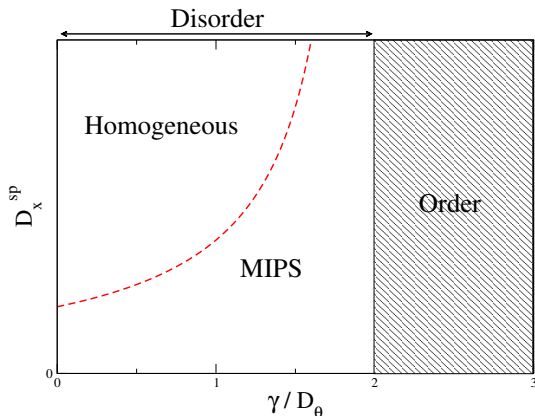
- When $\gamma < \gamma_c$, system effectively at equilibrium (case without aligning interactions: Cates, Tailleur et al. 2007, 2011, 2013)
→ computing S is possible, $S[\rho] = \int s(\rho(\mathbf{x}))d\mathbf{x}$

$$s''(\rho) = - \left(\frac{v^2(\rho) + \rho v(\rho)v'(\rho)}{(1 - \frac{\gamma}{2}) b[\rho]} + \frac{2D_x}{b[\rho]} \right)$$

- → compute $S[\rho]$

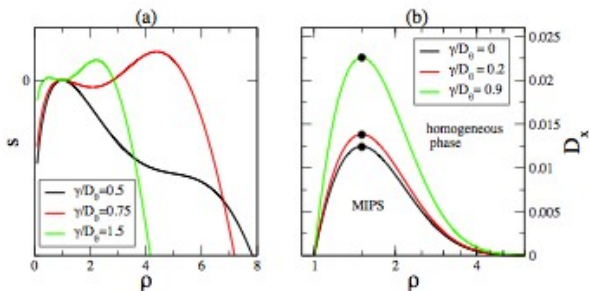
Results and discussion

- Reasonable to assume $v(\rho)$ decreasing \rightarrow possible phase separation (MIPS = Motility Induced Phase Separation). Role of the interactions?



Sketch of the $D_x - \tilde{\gamma}$ phase diagram.

Results and discussion



Left: entropy $s(\rho)$. Right: $D_x - \rho$ phase diagram.

- ▶ Spinodal line very sensitive to the interaction strength (observed in simulations)
- ▶ Density fluctuations increase when approaching the ordered phase
- ▶ A strong enough spatial diffusion always prevent phase separation

Conclusion

- ▶ Nice example where the limiting procedures seem well controlled + a general strategy
- ▶ Some physical insight in the "Motility Induced Phase Separation" with aligning interactions
- ▶ Next step: with a local orientation order
 - hyperbolic hydrodynamic limit
 - one cannot expect an effective equilibrium in the same sense
- ▶ Mathematical theory much less developed in this case... (recent works by Mariani, Bertini et al.) Work in progress...