

# **GGI lectures on the theory of fundamental interactions**

Firenze 12-18 January 2015

Aspects of Neutrino Physics (F. Feruglio)

## Bibliography

- on-line slides

- Lectures

W. Grimus hep-ph/0307149

P. Hernandez hep-ph/1010.4131

L. Maiani hep-ph/1406.5503

- Book

"Fundamentals of Neutrino Physics and Astrophysics" by C. Giunti and C. Kim

- PDG: Section on "Neutrino Masses, Mixing and Oscillations"

- Further readings:

General: A. Strumia and F. Vissani, hep-ph/0606054

Oscillation formula: L. Stodolsky hep-ph/9802387

GUTS: R. Mohapatra hep-ph/9801235

Leptogenesis: P. Di Bari hep-ph/1206.3168

Lepton Flavor Violation: F. Deppisch hep-ph/1206.5212,

Raidal et al hep-ph/0801.1826

# GGI lectures on the theory of fundamental interactions 2015

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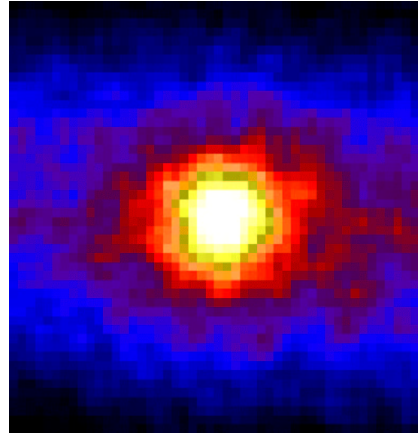
Aspects of neutrino physics (I)  
Neutrino Masses, Mixing and Oscillations:  
the data

Ferruccio Feruglio  
Universita' di Padova

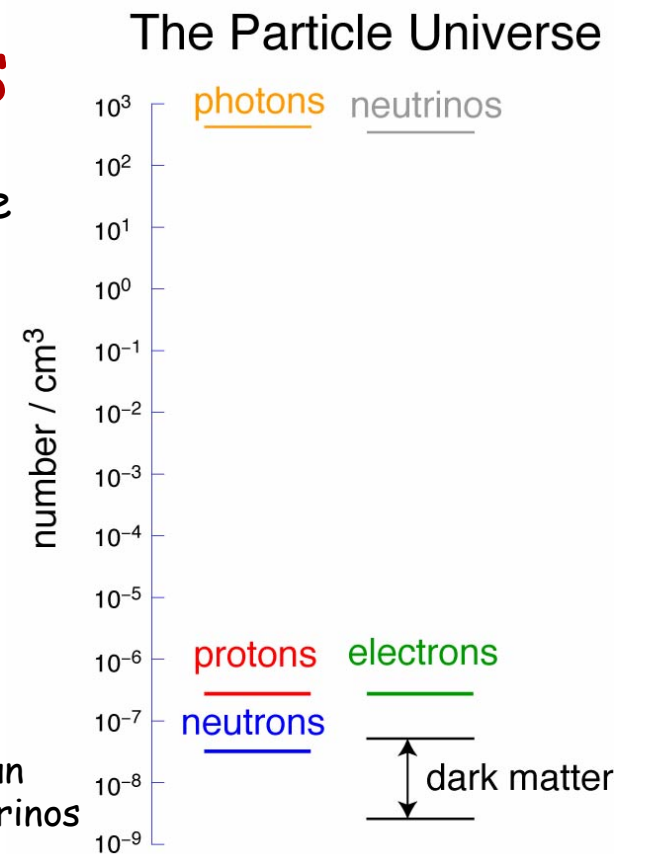
# General remarks on neutrinos

the more abundant particles in the universe after the photons: about 300 neutrinos per  $\text{cm}^3$

produced by stars: **most** of the sun energy emitted in neutrinos. As I speak more than 1 000 000 000 000 solar neutrinos go through your bodies each second.



this is a picture of the sun reconstructed from neutrinos



from Murayama  
talk Aspen 2007

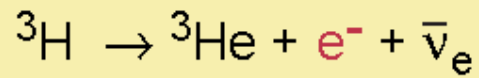
**electrically neutral and extremely light:**

they can carry information about extremely large length scales  
e.g. a probe of supernovae dynamics: neutrino events from a supernova explosion first observed 27 years ago

**in particle physics:**

they have a tiny mass (1 000 000 times smaller than the electron's mass)  
the discovery that they are massive allows us to explore, at least in principle, extremely high energy scales, otherwise inaccessible to present laboratory experiments

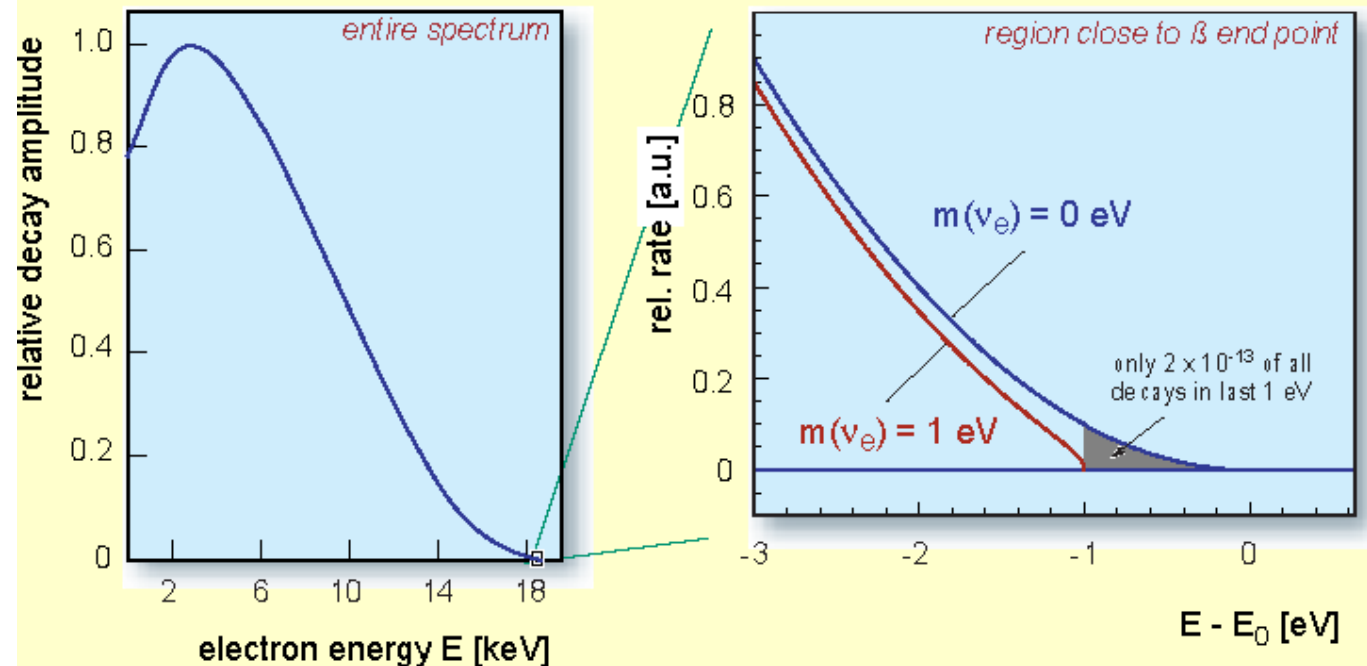
# Upper limit on neutrino mass (laboratory)



superallowed

half life :  $t_{1/2} = 12.32 \text{ a}$

$\beta$  end point energy :  $E_0 = 18.57 \text{ keV}$



$$m_{\nu} < 2.2 \text{ eV} \quad (95\% \text{ CL})$$



# Upper limit on neutrino mass (cosmology)

massive  $\nu$  suppress the formation of small scale structures

$$\sum_i m_i < 0.2 \div 1 \text{ eV}$$

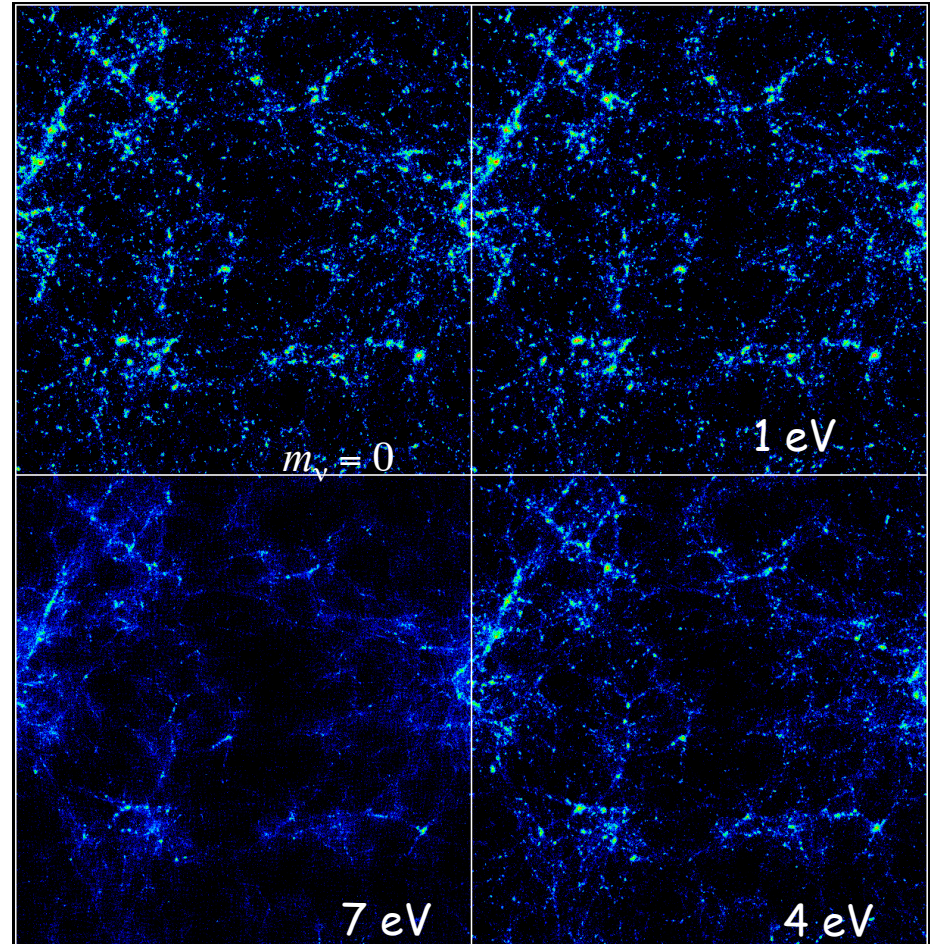
depending on

- assumed cosmological model
- set of data included
- how data are analyzed

$$k_{\text{nr}} \approx 0.026 \left( \frac{m_\nu}{1 \text{ eV}} \right)^{1/2} \Omega_m^{1/2} h \text{ Mpc}^{-1}.$$

The small-scale suppression is given by

$$\left( \frac{\Delta P}{P} \right) \approx -8 \frac{\Omega_\nu}{\Omega_m} \approx -0.8 \left( \frac{m_\nu}{1 \text{ eV}} \right) \left( \frac{0.1 N}{\Omega_m h^2} \right)$$

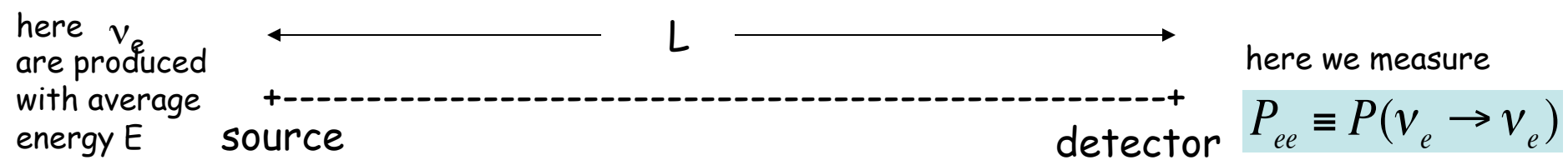


$$\delta(\vec{x}) \equiv \frac{\rho(\vec{x}) - \bar{\rho}}{\bar{\rho}}$$

$$\langle \delta(\vec{x}_1) \delta(\vec{x}_2) \rangle = \int \frac{d^3 k}{(2\pi)^3} e^{i\vec{k} \cdot (\vec{x}_1 - \vec{x}_2)} P(\vec{k})$$

# Two-flavour neutrino oscillations

$(\nu_e, \nu_\mu)$



neutrino interaction eigenstates

$$\begin{pmatrix} \nu_e \\ \nu_\mu \end{pmatrix} = \underbrace{\begin{pmatrix} \cos \vartheta & \sin \vartheta \\ -\sin \vartheta & \cos \vartheta \end{pmatrix}}_U \begin{pmatrix} 1 & 0 \\ 0 & e^{i\alpha} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$-\frac{g}{\sqrt{2}} W_\mu^- \bar{l}_L \gamma^\mu \nu_l$$

$$\gamma/2 = \vartheta$$

as before, but  $t \approx L$

$$E_2 - E_1 = \sqrt{p^2 + m_2^2} - \sqrt{p^2 + m_1^2} \approx \frac{m_2^2}{2E} - \frac{m_1^2}{2E} \equiv \frac{\Delta m_{21}^2}{2E}$$

$$P_{ee} = \left| \langle \nu_e | \psi(L) \rangle \right|^2 = 1 - \underbrace{4|U_{e1}|^2|U_{e2}|^2}_{\sin^2 2\vartheta} \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E} \right)$$

no dependence on the phase  $\alpha$   
more on this later on ....

to see any effect, if  $\Delta m^2$  is tiny, we need both  $\theta$  and  $L$  large

# regimes

$$P_{ee} = \left| \langle \nu_e | \psi(L) \rangle \right|^2 = 1 - \underbrace{4|U_{e1}|^2|U_{e2}|^2}_{\sin^2 2\vartheta} \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E} \right)$$

$$\frac{\Delta m^2 L}{4E} \ll 1$$

$$P_{ee} \approx 1$$

$$\frac{\Delta m^2 L}{4E} \gg 1$$

$$\sin^2 \left( \frac{\Delta m^2 L}{4E} \right) \approx \frac{1}{2}$$

$$P_{ee} \approx 1 - \frac{\sin^2 2\vartheta}{2}$$

by averaging over  $\nu_e$  energy at the source

$$\frac{\Delta m^2 L}{4E} \approx 1$$

$$P_{ee} = P_{ee}(E)$$

useful relation  $\frac{\Delta m^2 L}{4E} \approx 1.27 \left( \frac{\Delta m^2}{1 \text{ eV}^2} \right) \left( \frac{L}{1 \text{ Km}} \right) \left( \frac{E}{1 \text{ GeV}} \right)^{-1}$

source	L(km)	E(GeV)	$\Delta m^2(\text{eV}^2)$
$\nu_e, \nu_\mu$ (atmosphere)	$10^4$ (Earth diameter)	1-10	$10^{-4} - 10^{-3}$
anti- $\nu_e$ (reactor)	1	$10^{-3}$	$10^{-3}$
anti- $\nu_e$ (reactor)	100	$10^{-3}$	$10^{-5}$
$\nu_e$ (sun)	$10^8$	$10^{-3} - 10^{-2}$	$10^{-11} - 10^{-10}$

neglecting  
matter  
effects

# Three-flavour neutrino oscillations

$(\nu_e, \nu_\mu, \nu_\tau)$

survival probability as before, with more terms

$$P_{ff} = P(\nu_f \rightarrow \nu_f) = \left| \langle \nu_f | \psi(L) \rangle \right|^2 = 1 - 4 \sum_{k < j} |U_{fk}|^2 |U_{fj}|^2 \sin^2 \left( \frac{\Delta m_{jk}^2 L}{4E} \right)$$

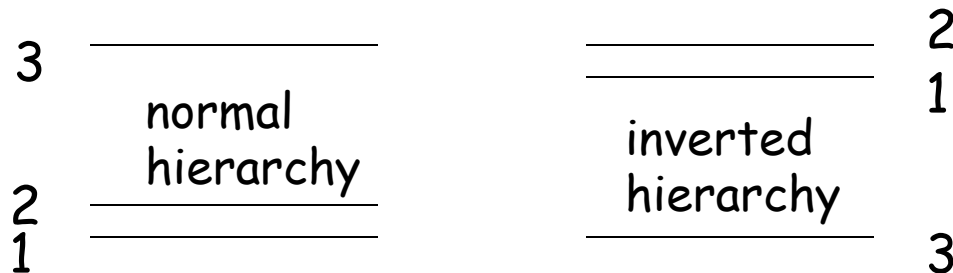
similarly, we can derive the disappearance probabilities  $P_{ff'} = P(\nu_f \rightarrow \nu_{f'})$

conventions:  $[\Delta m_{ij}^2 \equiv m_i^2 - m_j^2]$

$$m_1 < m_2$$

$\Delta m_{21}^2 < |\Delta m_{32}^2|, |\Delta m_{31}^2|$  i.e. 1 and 2 are, by definition, the closest levels

two possibilities:



# Mixing matrix $U=U_{PMNS}$ (Pontecorvo, Maki, Nakagawa, Sakata)

neutrino  
interaction  
eigenstates

$$\nu_f = \sum_{i=1}^3 U_{fi} \nu_i$$

$(f = e, \mu, \tau)$

neutrino mass  
eigenstates

$U$  is a  $3 \times 3$  unitary matrix  
standard parametrization

$$U_{PMNS} = \begin{pmatrix} c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{i\delta} \\ -s_{12} c_{23} - c_{12} s_{13} s_{23} e^{-i\delta} & c_{12} c_{23} - s_{12} s_{13} s_{23} e^{-i\delta} & c_{13} s_{23} \\ -c_{12} s_{13} c_{23} e^{-i\delta} + s_{12} s_{23} & -s_{12} s_{13} c_{23} e^{-i\delta} - c_{12} s_{23} & c_{13} c_{23} \end{pmatrix} \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha} & 0 \\ 0 & 0 & e^{i\beta} \end{pmatrix}$$

$$c_{12} \equiv \cos \vartheta_{12}, \dots$$

three mixing angles

$$\vartheta_{12}, \vartheta_{13}, \vartheta_{23}$$

three phases (in the most general case)

$$\delta$$

$$\underbrace{\alpha, \beta}$$

do not enter  $P_{ff'} = P(\nu_f \rightarrow \nu_{f'})$

oscillations can only test 5 combinations

$$\Delta m_{21}^2, \Delta m_{32}^2, \vartheta_{12}, \vartheta_{13}, \vartheta_{23}, \delta$$

## structure of the mixing matrix

$$\begin{pmatrix}
 c_{12} c_{13} & s_{12} c_{13} & s_{13} e^{i\delta} \\
 -s_{12} c_{23} - c_{12} s_{13} s_{23} e^{-i\delta} & c_{12} c_{23} - s_{12} s_{13} s_{23} e^{-i\delta} & c_{13} s_{23} \\
 -c_{12} s_{13} c_{23} e^{-i\delta} + s_{12} s_{23} & -s_{12} s_{13} c_{23} e^{-i\delta} - c_{12} s_{23} & c_{13} c_{23}
 \end{pmatrix} =$$

$$= \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13} e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

# Analysis of Oscillations Data

we anticipate that there are two small parameters

$$|\alpha| \equiv \left| \frac{\Delta m_{21}^2}{\Delta m_{31}^2} \right| \approx 0.03$$

$$\Delta m_{21}^2 \ll |\Delta m_{32}^2|, |\Delta m_{31}^2|$$

$$|U_{e3}|^2 \approx \sin^2 \vartheta_{13} \approx 0.02$$

we first consider experiments not sensitive to  $\Delta m_{21}^2$  (L not very large, E not very small) and we set  $\Delta m_{21}^2 = 0$

## EXERCISE

derive  $P_{ee}, P_{\mu\mu}, P_{\mu e}$  in the limit  $\Delta m_{21}^2 = 0$  (vacuum osc., no matter effects)

$$\Delta_{ij} \equiv \frac{\Delta m_{ij}^2 L}{4E}$$

$$\Delta \equiv \frac{\Delta m_{13}^2 L}{4E} \quad [\Delta m_{21}^2 = 0]$$

$$P_{ee} = 1 - 4|U_{e3}|^2 (1 - |U_{e3}|^2) \sin^2 \Delta$$

$$P_{\mu\mu} = 1 - 4|U_{\mu3}|^2 (1 - |U_{\mu3}|^2) \sin^2 \Delta$$

$$P_{\mu e} = P_{e\mu} = 4|U_{\mu3}|^2 |U_{e3}|^2 \sin^2 \Delta$$

similarly,  $P_{\tau\tau}, P_{\tau\mu}, P_{\mu\tau}, P_{\tau e}, P_{e\tau}$  only depend on  $U_{f3}$  and  $\Delta$  for  $\Delta m_{21}^2 = 0$

we are testing the third column

$$U_{PMNS} = \begin{pmatrix} \cdot & \cdot & U_{e3} \\ \cdot & \cdot & U_{\mu 3} \\ \cdot & \cdot & U_{\tau 3} \end{pmatrix}$$

we also consider the limit  $\vartheta_{13} = 0$

we are left with one frequency and one mixing angle

$$|U_{e3}|^2 \approx \sin^2 \vartheta_{13} \approx 0$$

$$P_{ee} = 1$$

$$P_{\mu\mu} = 1 - \sin^2 2\vartheta_{23} \sin^2 \Delta$$

$$P_{\mu e} = P_{e\mu} = 0$$

$$P_{\tau\tau} = P_{\mu\mu}$$

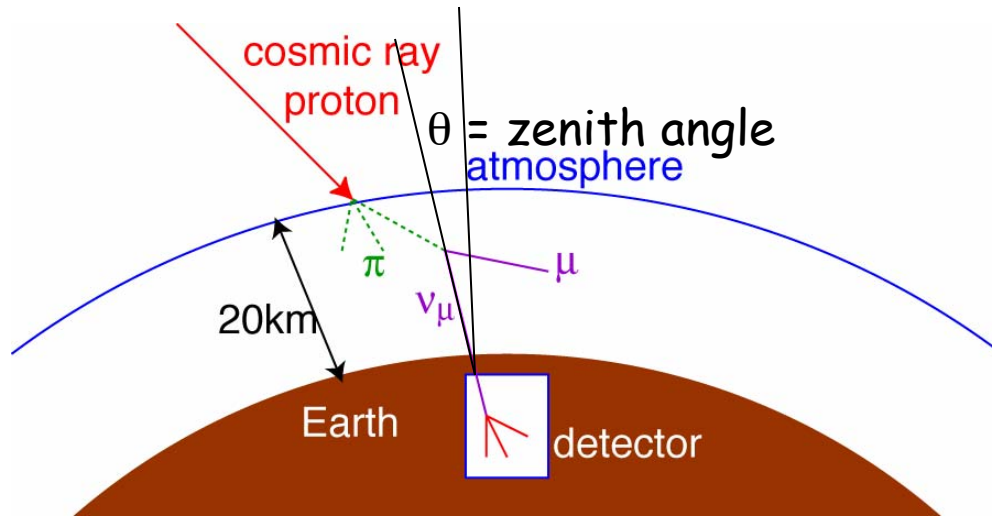
$$P_{\tau\mu} = P_{\mu\tau} = \sin^2 2\vartheta_{23} \sin^2 \Delta$$

$$P_{\tau e} = P_{e\tau} = 0$$

two-flavour oscillations

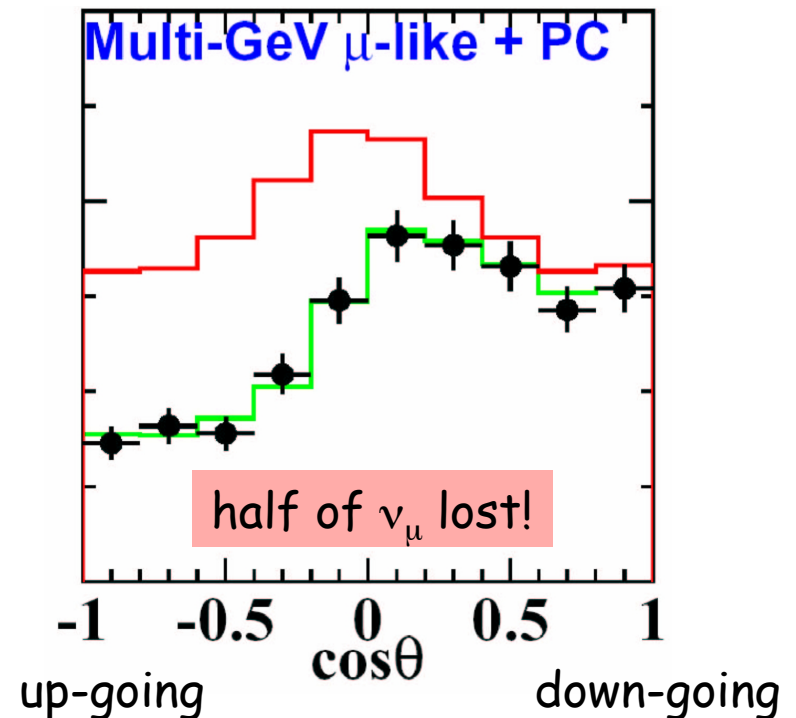
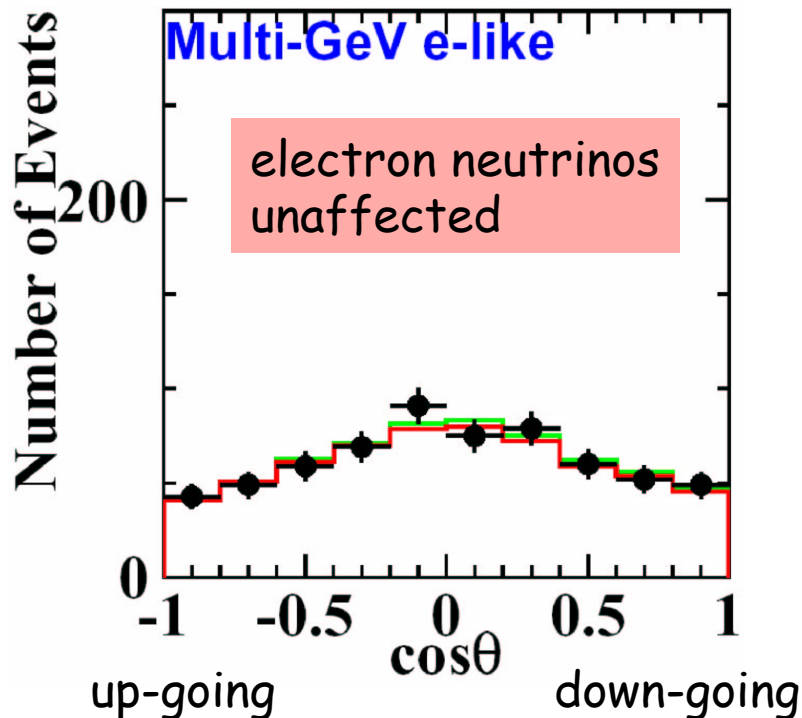


# Atmospheric neutrino oscillations



Electron and muon neutrinos (and antineutrinos) produced by the collision of cosmic ray particles on the atmosphere

Experiment:  
**SuperKamiokande** (Japan)



## electron neutrinos do not oscillate

by working in the approximation  $\Delta m_{21}^2 = 0$

$$P_{ee} = 1 - \underbrace{4|U_{e3}|^2(1-|U_{e3}|^2)}_{\sin^2 2\vartheta_{13}} \sin^2\left(\frac{\Delta m_{31}^2 L}{4E}\right) \approx 1$$

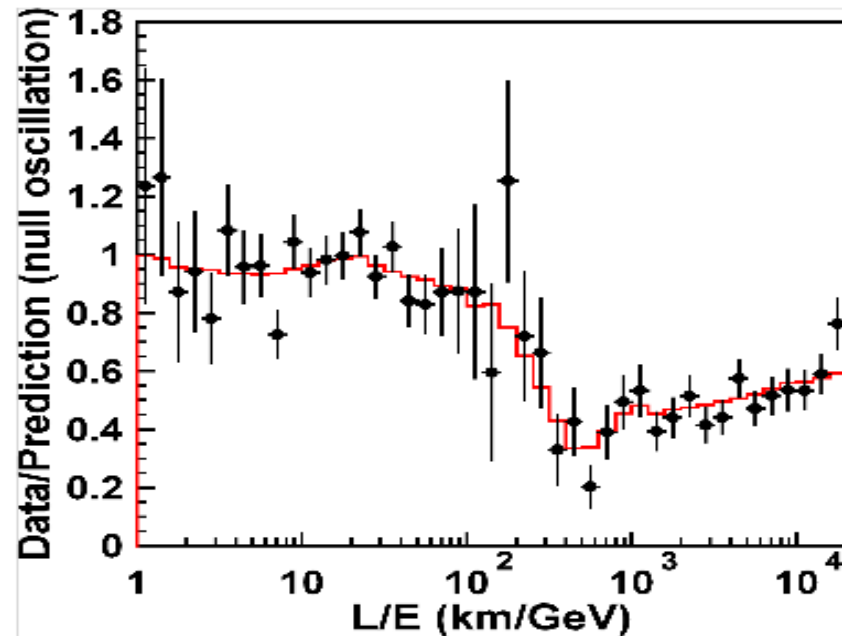
for  $U_{e3} = \sin \vartheta_{13} \approx 0$

## muon neutrinos oscillate

$$P_{\mu\mu} = 1 - \underbrace{4|U_{\mu3}|^2(1-|U_{\mu3}|^2)}_{\sin^2 2\vartheta_{23}} \sin^2\left(\frac{\Delta m_{32}^2 L}{4E}\right)$$

$$|\Delta m_{32}^2| \approx 2 \cdot 10^{-3} \text{ eV}^2$$

$$\sin^2 \vartheta_{23} \approx \frac{1}{2}$$



$$U_{PMNS} = \begin{pmatrix} \cdot & \cdot & 0 \\ \cdot & \cdot & 1 \\ \cdot & \cdot & -\frac{1}{\sqrt{2}} \\ \cdot & \cdot & \frac{1}{\sqrt{2}} \end{pmatrix} + (\text{small corrections})$$

maximal mixing!  
not a replica of the quark  
mixing pattern

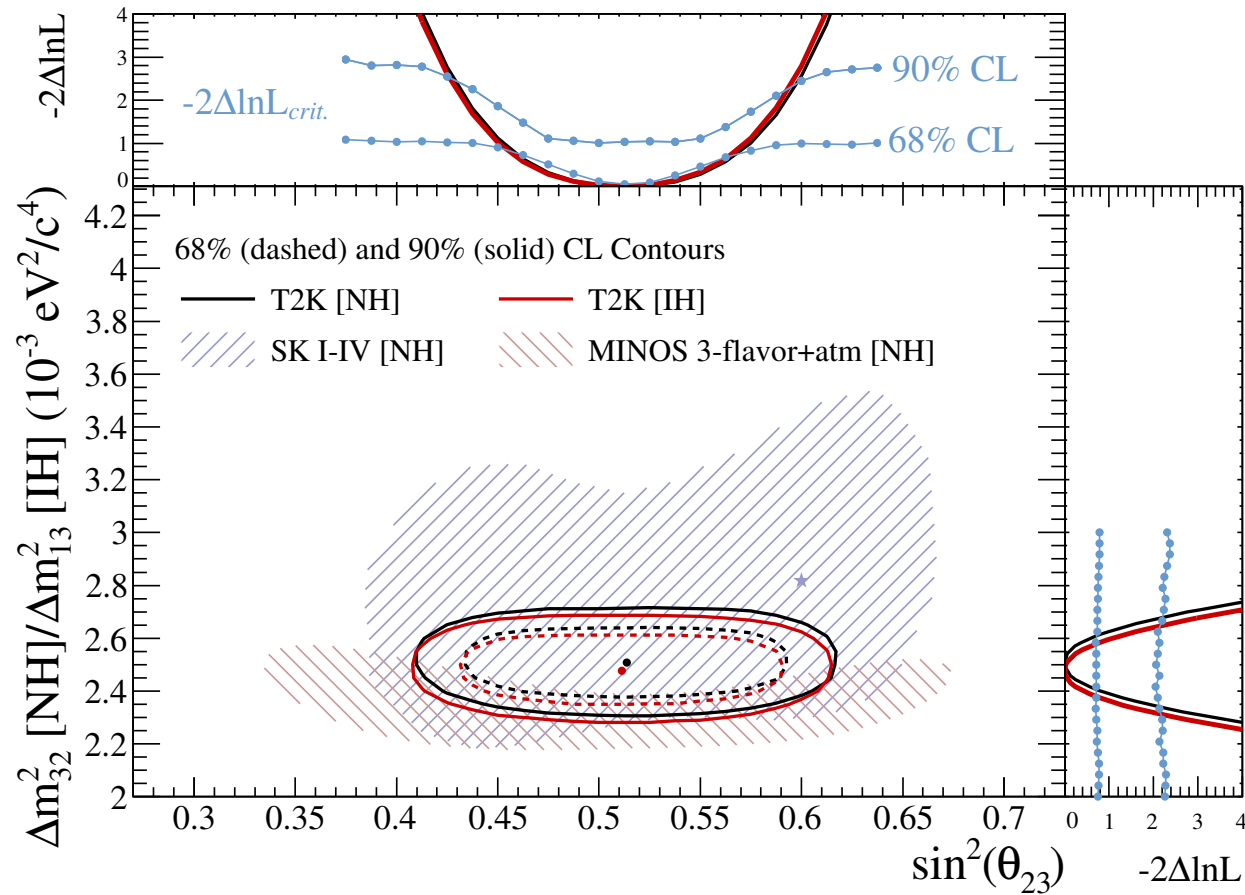
## other terrestrial experiments measuring $P_{\mu\mu}$

- K2K** (Japan, from KEK to Kamioka mine  $L \approx 250$  Km  $E \approx 1.3$  GeV)
- MINOS** (USA, from Fermilab to Soudan mine  $L \approx 735$  Km  $E \approx 3$  GeV)
- T2K** (Japan, from Tokai, J-Park to Kamioka mine  $L \approx 295$  Km  $E \approx 0.6$  GeV)
- OPERA** (CERN-Italy, from CERN to LNGS  $L \approx 732$  Km  $E \approx 17$  GeV)

all sensitive to  $\Delta m_{32}^2$  close to  $10^{-3} \text{ eV}^2$

**OPERA** energy optimized to maximize  $\tau$  production, via  $CC$  events  
by the end of 2014 4  $\tau$  events have been seen

# recent results from T2K [hep-ex/1403.1532]



$$\sin^2 \vartheta_{23} = \begin{cases} 0.514^{+0.055}_{-0.056} & (NO) \\ 0.511 \pm 0.055 & (IO) \end{cases}$$

$$\Delta m_{32}^2 = (2.51 \pm 0.10) \times 10^{-3} \text{ eV}^2 \quad (NO)$$

$$\Delta m_{13}^2 = (2.48 \pm 0.10) \times 10^{-3} \text{ eV}^2 \quad (IO)$$

# KamLAND

previous experiments were sensitive to  $\Delta m^2$  close to  $10^{-3} \text{ eV}^2$   
to explore smaller  $\Delta m^2$  we need larger  $L$  and/or smaller  $E$

KamLAND experiment exploits the low-energy electron anti-neutrinos ( $E \approx 3 \text{ MeV}$ ) produced by Japanese and Korean reactors at an average distance of  $L \approx 180 \text{ Km}$  from the detector and is potentially sensitive to  $\Delta m^2$  down to  $10^{-5} \text{ eV}^2$

by working in the approximation

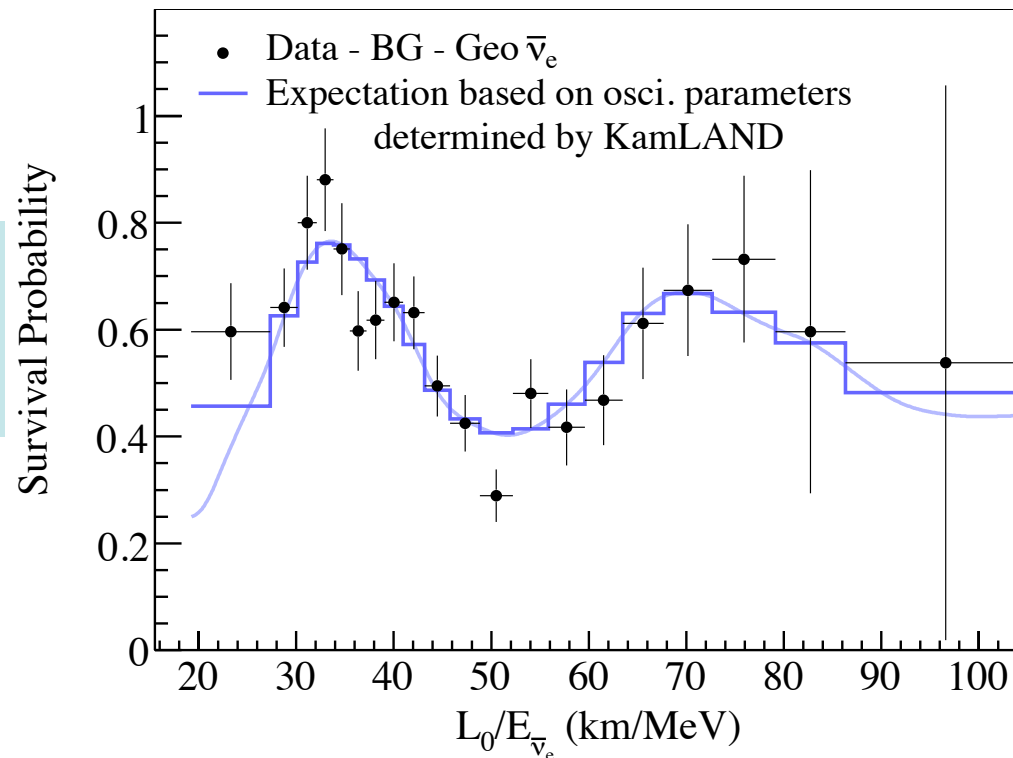
$$U_{e3} = \sin \vartheta_{13} = 0 \quad \text{we get}$$

[Exercise]

$$P_{ee} = 1 - \underbrace{4|U_{e1}|^2|U_{e2}|^2}_{\sin^2 2\vartheta_{12}} \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E} \right)$$

$$\Delta m_{21}^2 \approx 8 \cdot 10^{-5} \text{ eV}^2$$

$$\sin^2 \vartheta_{12} \approx 1/3$$



## EXERCISE

estimate  $\Delta m_{21}^2$  from position of second oscillation dip in previous plot

$$\Delta m_{21}^2 = 6\pi \frac{E}{L} \Big|_{dip} \approx 6\pi \times \frac{1}{50} \text{ MeV} / \text{Km} = 7.5 \times 10^{-5} \text{ eV}^2$$

## EXERCISE

work out  $P_{ee}$  by keeping  $U_{e3}$  non-vanishing

$$P_{ee} \approx |U_{e3}|^4 + (1 - |U_{e3}|^2)^2 \left( 1 - \sin^2 2\vartheta_{12} \sin^2 \Delta_{21} \right)$$

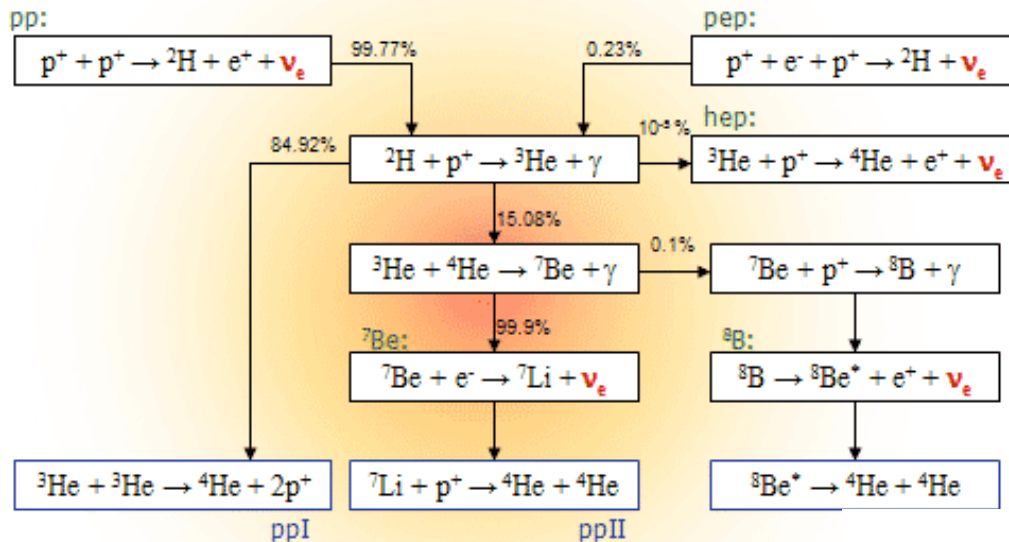
$$U_{PMNS} = \begin{pmatrix} \begin{array}{cc} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} \\ 1 & 1 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} \end{array} & \begin{array}{c} 0 \\ 1 \\ -\frac{1}{\sqrt{2}} \\ 1 \\ \frac{1}{\sqrt{2}} \end{array} \end{pmatrix} + (\text{small corrections})$$

by unitarity

this pattern is called tri-bimaximal  
completely different from the quark  
mixing pattern: two angles are large

historically  $\Delta m_{21}^2$  and  $\sin^2 \theta_{12}$  were first determined by solving the **solar neutrino problem**, i.e. the disappearance of about one third of solar electron neutrino flux, for solar neutrinos above few MeV. The desire of detecting solar neutrinos, to confirm the thermodynamics of the sun, was the driving motivation for the whole field for more than 30 years. Electron solar neutrinos oscillate, but the formalism requires the introduction of matter effects, since the electron density in the sun is not negligible. Experiments: **SuperKamiokande**, **SNO**, **Borexino**

# Solar Neutrinos

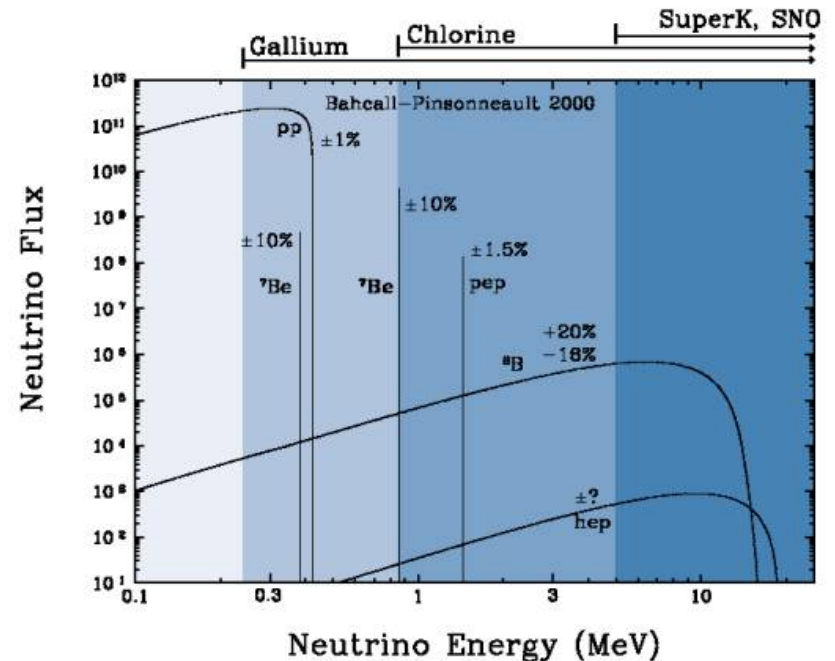


$\nu_e$  produced in the core of the sun through several chains/reactions

with different energy spectrum

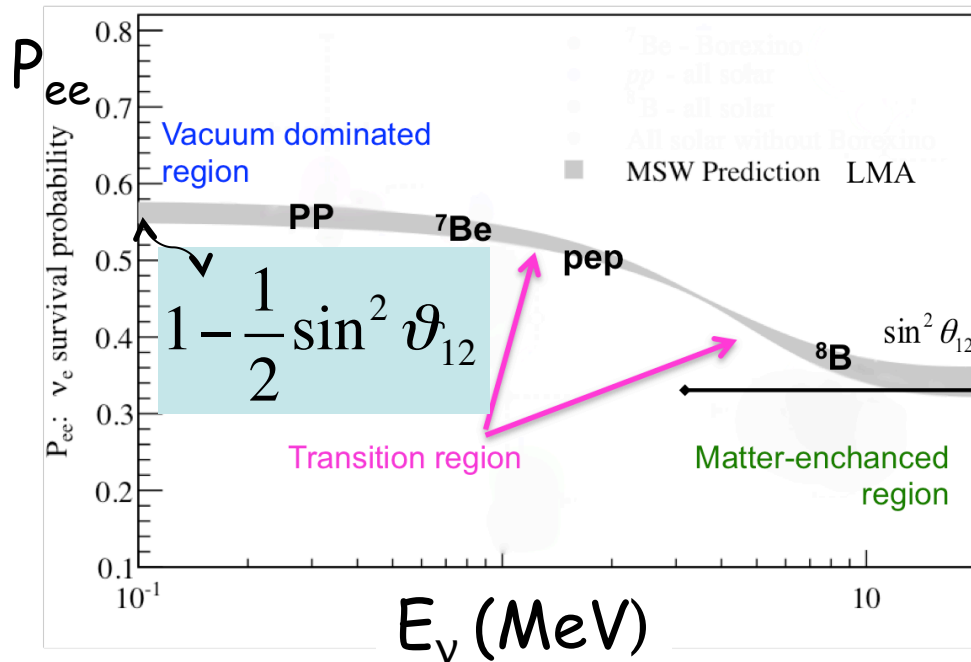
most neutrinos come from pp fusion  $E_{\text{max}} \approx 0.4 \text{ MeV}$

most energetic neutrinos come from  ${}^8\text{B}$  decay  $E_{\text{max}} \approx 15 \text{ MeV}$





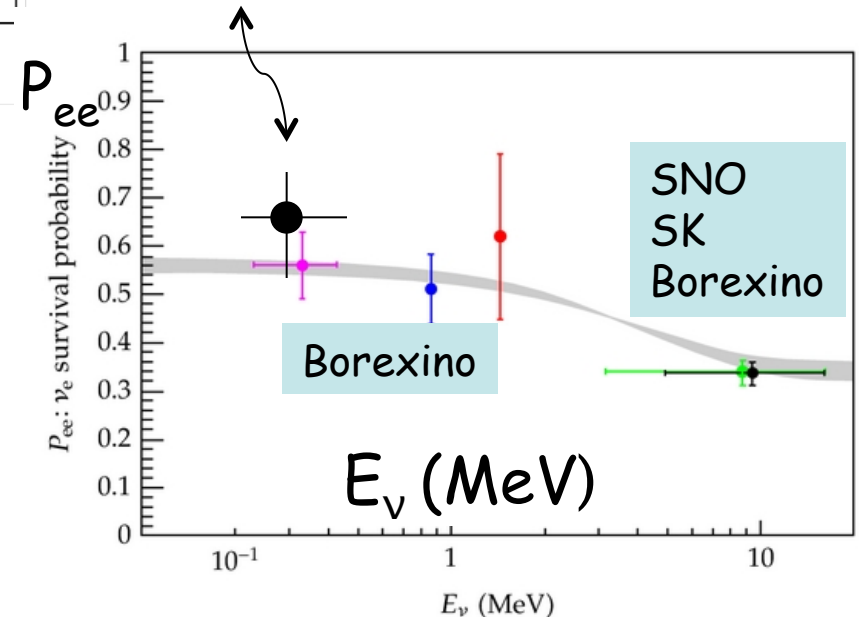
# Theory prediction for $P_{ee}$



$$\sin^2 \vartheta_{12}$$

[Borexino, Nature 512 (2014) 383]

experiments reveal solar neutrinos through different processes and have different energy thresholds



- $\bullet$   $pp$ -all solar
- $\bullet$   ${}^7\text{Be}$ -Borexino
- $\bullet$   $pep$ -Borexino
- $\bullet$   ${}^8\text{B}$ -SNO LETA + borexino
- $\bullet$   ${}^8\text{B}$ -SNO + SK
- $\square$  MSW-LMA prediction

# $\vartheta_{13}$ from disappearance experiments

These experiments have been realized with reactors. Electron anti-neutrinos are produced by a reactor ( $E \approx 3 \text{ MeV}$ ,  $L \approx 1 \text{ Km}$ ) (by CPT the survival probability in vacuum is the same for neutrinos and antineutrinos and matter effects are negligible).

In this range of  $(L, E)$  oscillations driven by  $\Delta m_{21}^2$  are negligible and the survival probability  $P_{ee}$  only depends on  $(|U_{e3}|, \Delta m_{31}^2)$ .

$$P_{ee} = 1 - \underbrace{4|U_{e3}|^2(1 - |U_{e3}|^2)}_{\sin^2 2\vartheta_{13}} \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E} \right)$$

$$E \approx 3 \text{ MeV}$$

$$L \approx 1 \text{ Km}$$

Experiment	Near Detectors	Far Detectors
CHOOZ (France)	–	(1) 1050m
Double CHOOZ	> 2014	(1) 1050m
Reno (Korea)	(1) 290m	(1) 1380m
Daya Bay (China)	(4) (360-530)m	(4) (1600-2000)m

before 2012 there was only an upper bound on  $|U_{e3}|$  by CHOOZ

today (end 2014) the value of  $\vartheta_{13}$  is dominated by the Daya Bay result

$$\sin^2 2\vartheta_{13} = 0.085 \pm 0.005$$

$$|\Delta m_{13}^2| = 2.44_{-0.11}^{+0.10} \times 10^{-3} \text{ eV}^2$$

$$|U_{e3}|^2 = \sin^2 \vartheta_{13} = 0.0215 \pm 0.0013$$

$$\vartheta_{13} = (8.4 \pm 0.3)^\circ$$

# $\theta_{13}$ from appearance experiments

These experiments use a muon-neutrino beam from an accelerator and look for conversion of muon-neutrinos into electron-neutrinos. The (L,E) range is such that they are mainly sensitive to  $\Delta m_{31}^2$

Experiment	E(GeV)	L(Km)
T2K (Japan)	0.6	295
MINOS (USA)	3	735

at the LO (neglecting  $\Delta m_{21}^2$  and matter effects)

$$P_{\mu e} = 4 |U_{\mu 3}|^2 |U_{e 3}|^2 \sin^2 \Delta = \sin^2 \vartheta_{23} \sin^2 2\vartheta_{13} \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E} \right)$$

however in this case corrections from  $\Delta m_{21}^2$  and matter effects are non-negligible

## EXERCISE

by expanding  $P_{\mu e}$  to first order in  $\alpha = \Delta m_{21}^2 / \Delta m_{31}^2$  show that

$$\begin{aligned} P_{\mu e} = & \sin^2 \vartheta_{23} \sin^2 2\vartheta_{13} \sin^2 \Delta_{13} \\ & - 8\alpha J_{CP} \Delta_{13} \sin^2 \Delta_{13} \\ & - 8\alpha J_{CP} \frac{\cos \delta}{\sin \delta} \Delta_{13} \cos \Delta_{13} \sin \Delta_{13} \\ & + O(\alpha^2) + \text{matter effects} \end{aligned}$$

$$\Delta_{13} = \frac{\Delta m_{31}^2 L}{4E}$$

$$\begin{aligned} J_{CP} &= \text{Im}(U_{\mu 3} U_{e 3}^* U_{\mu 2}^* U_{e 2}) \\ &= \frac{1}{8} \sin 2\vartheta_{12} \sin 2\vartheta_{23} \sin 2\vartheta_{13} \cos \vartheta_{13} \sin \delta \end{aligned}$$

T2K works near the first oscillation maximum where  $|\Delta_{13}|=\pi/2$

$$P_{\mu e} = \sin^2 \vartheta_{23} \sin^2 2\vartheta_{13} - 4\pi |\alpha| J_{CP} + O(\alpha^2) + \text{matter effects}$$

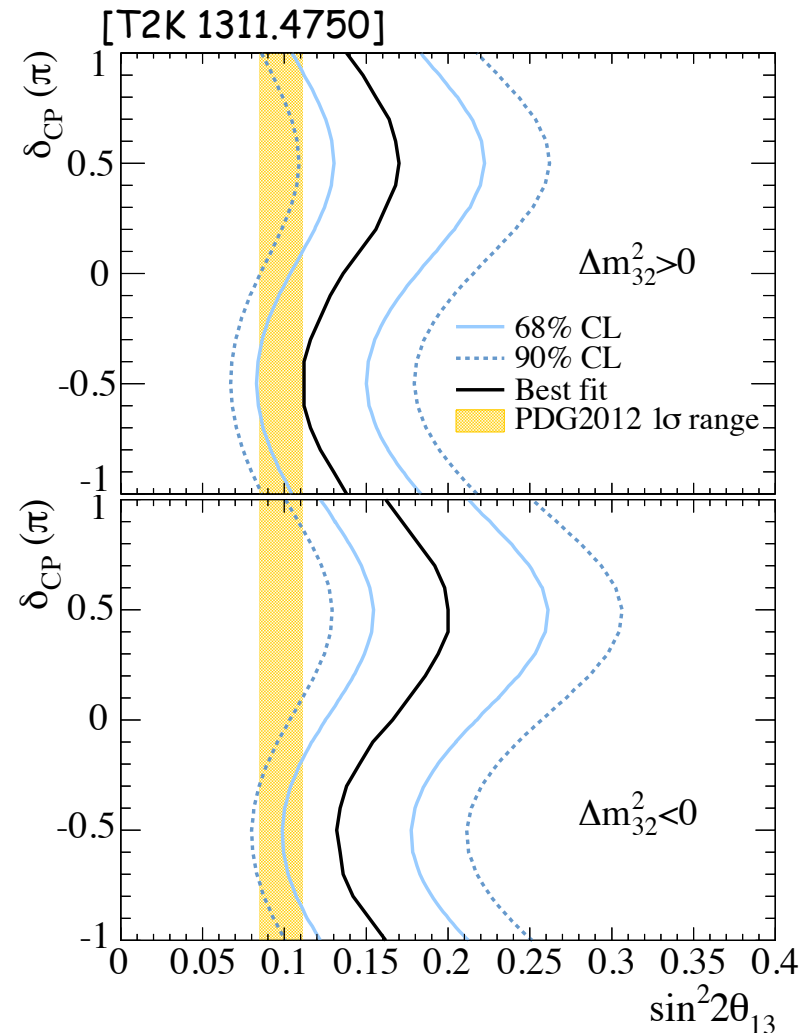
At present (end 2014) agreement with the value of  $\vartheta_{13}$  determined by reactor disappearance experiments requires

$$\sin \delta \approx -1$$

$$\delta \approx \frac{3}{2}\pi$$

i.e. maximal CP violation in the lepton sector

the relative subleading corrections are  $O(20\%)$  and are sensitive to  $\sin \delta$



# main detection processes

Neutrinos	Experiment	Process
atmospheric $\nu$	SK K2K, MINOS, T2K, Opera	$\nu N \rightarrow l X$
solar $\nu$	SK, Borexino SNO	$\nu_X e \rightarrow \nu_X e$ $\nu_X D \rightarrow \nu_X pn, \nu_e D \rightarrow e pp$
reactor $\nu$	KamLand, Chooz, DoubleChooz, Reno, Daya Bay	$\bar{\nu}_e p \rightarrow e^+ n \quad (e^+ D \gamma)$

# Summary of data

$$m_\nu < 2.2 \text{ eV} \quad (95\% \text{ CL}) \quad (\text{lab})$$

$$\sum_i m_i < 0.2 \div 1 \text{ eV} \quad (\text{cosmo})$$

$$\Delta m_{atm}^2 \equiv \begin{cases} \Delta m_{31}^2 = (2.462 \pm 0.033) \times 10^{-3} \text{ eV}^2 & \text{NO} \\ \Delta m_{32}^2 = -(2.453 \pm 0.047) \times 10^{-3} \text{ eV}^2 & \text{IO} \end{cases}$$

$$\Delta m_{sol}^2 \equiv \Delta m_{21}^2 = (7.55^{+0.18}_{-0.17}) \times 10^{-5} \text{ eV}^2$$

$$\sin^2 \vartheta_{13} = 0.0223^{+0.0011}_{-0.0010} \quad \delta_{CP} = (259^{+76}_{-69})$$

$$\sin^2 \vartheta_{23} = [0.451^{+0.026}_{-0.020}] \oplus [0.580^{+0.024}_{-0.039}]$$

$$\sin^2 \vartheta_{12} = 0.311^{+0.013}_{-0.012}$$

[G.-Garcia, Maltoni, Salvado, Schwetz 1209.3023]

violation of individual lepton number  
implied by neutrino oscillations

## Summary of unknowns

absolute neutrino mass  
scale is unknown  
[but well-constrained!]

sign  $[\Delta m_{atm}^2]$  unknown

[complete ordering  
(either normal or inverted  
hierarchy) not known]

$\delta, \alpha, \beta$  unknown

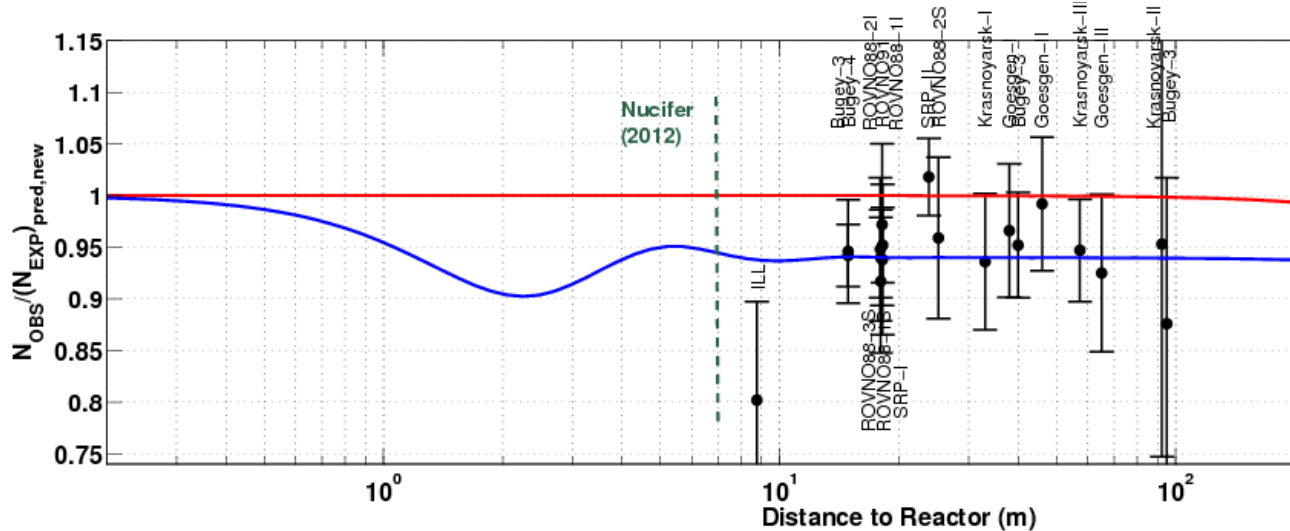
[CP violation in lepton  
sector not yet established]

violation of total lepton number  
not yet established

sterile neutrinos ?

# 1 reactor anomaly (anti- $\nu_e$ disappearance)

re-evaluation of reactor anti- $\nu_e$  flux: new estimate 3.5% higher than old one



$$(\Phi_{\text{exp}} - \Phi_{\text{th}}) / \Phi_{\text{th}} \approx -6\%$$

[th. uncertainty?]

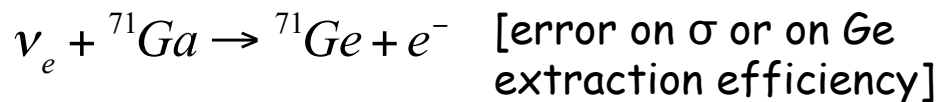
very SBL  $L \leq 100$  m

$$\vartheta_{\text{es}} \approx 0.2$$

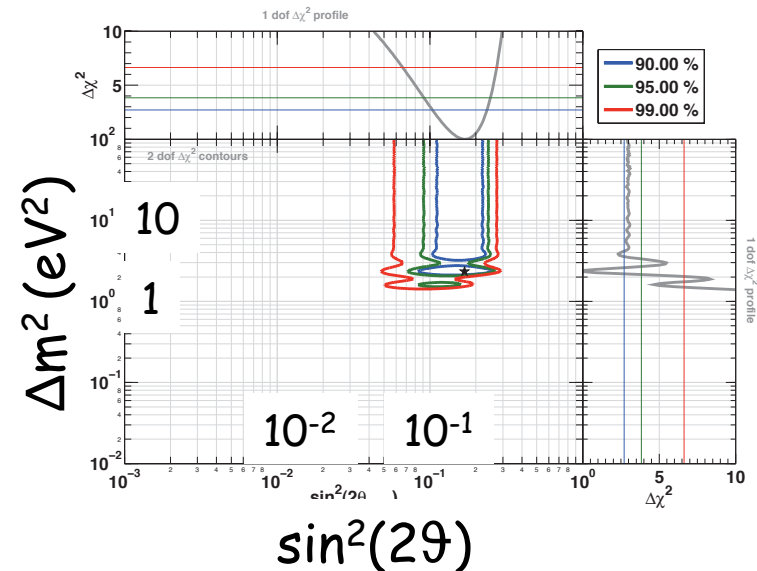
$$\Delta m^2 \approx m_s^2 \geq 1 \text{ eV}^2$$

supported by the **Gallium anomaly**

$\nu_e$  flux measured from high intensity radioactive sources in Gallex, Sage exp



... but disfavoured by cosmological limits





## 2 long-standing claim

evidence for  $\nu_\mu \rightarrow \nu_e$  appearance in accelerator experiments

exp		$E(\text{MeV})$	$L(\text{m})$
<i>LSND</i>	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$10 \div 50$	30
<i>MiniBoone</i>	$\nu_\mu \rightarrow \nu_e$ $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$	$300 \div 3000$	541

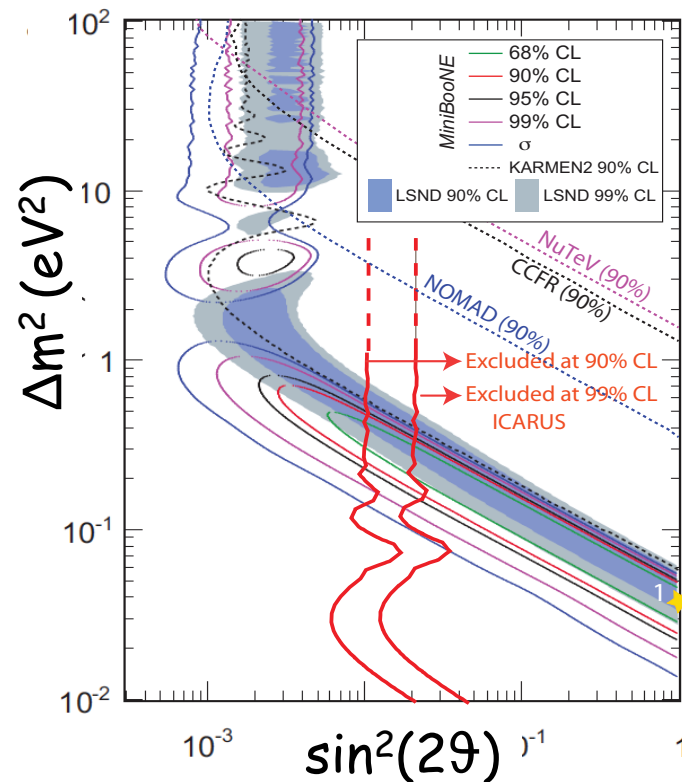
$3.8\sigma$

$3.8\sigma$  [signal from low-energy region]

parameter space limited by  
negative results from Karmen  
and ICARUS

$$\vartheta_{e\mu} \approx 0.035$$

$$\Delta m^2 \approx 0.5 \text{ eV}^2$$

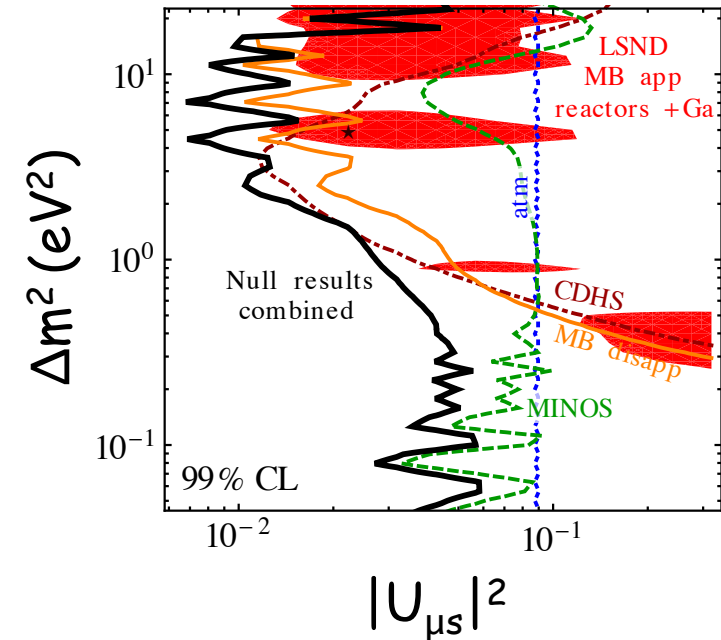


interpretation in 3+1 scheme: **inconsistent**  
(more than 1s disfavored by cosmology)

$$\underbrace{\vartheta_{e\mu}}_{0.035} \approx \underbrace{\vartheta_{es}}_{0.2} \times \vartheta_{\mu s} \quad \Rightarrow \quad \vartheta_{\mu s} \approx 0.2$$

predicted suppression in  $\nu_\mu$  disappearance experiments: **undetected**

by ignoring LSND/Miniboone data the reactor anomaly can be accommodated by  $m_s \geq 1$  eV and  $\vartheta_{es} \approx 0.2$   
[not suitable for Warm DM]



# GGI lectures on the theory of fundamental interactions 2015

Firenze, 12-16 January 2015

Aspects of neutrino physics (II)  
Neutrino Masses, Mixing and Oscillations:  
Implication for Physics BSM

Ferruccio Feruglio  
Universita' di Padova

# Lecture 1

## Neutrino Masses

# Summary of data

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[CP violation in lepton  
sector not yet established]

violation of total lepton number  
not yet established

# Beyond the Standard Model

a non-vanishing neutrino mass is the first evidence of the incompleteness of the Standard Model [SM]

in the SM neutrinos belong to  $SU(2)$  doublets with hypercharge  $Y=-1/2$  they have only two helicities (not four, as the other charged fermions)

$$l = \begin{pmatrix} \nu_e \\ e \end{pmatrix} = (1, 2, -1/2)$$

the requirement of invariance under the gauge group  $G=SU(3)\times SU(2)\times U(1)$  forbids pure fermion mass terms in the lagrangian. Charged fermion masses arise, after electroweak symmetry breaking, through gauge-invariant Yukawa interactions

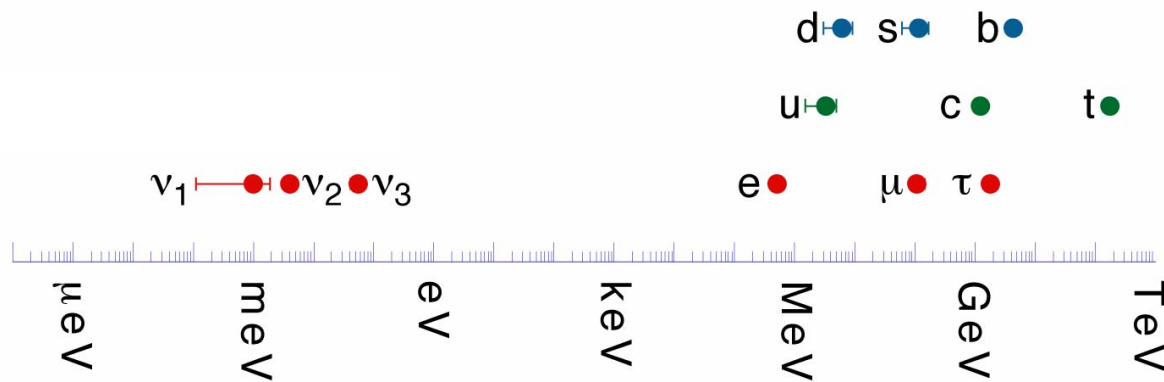
$$\Phi \underbrace{\Psi\Psi'}_{\text{same helicity}}$$

not even this term is allowed for SM neutrinos, by gauge invariance

# Questions

how to extend the SM in order to accommodate neutrino masses?

why neutrino masses are so small, compared with the charged fermion masses?



why lepton mixing angles are so different from those of the quark sector?

$$U_{PMNS} = \begin{pmatrix} \frac{2}{\sqrt{6}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & -\frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} + \text{corrections}$$

$$V_{CKM} \approx \begin{pmatrix} 1 & O(\lambda) & O(\lambda^4 \div \lambda^3) \\ O(\lambda) & 1 & O(\lambda^2) \\ O(\lambda^4 \div \lambda^3) & O(\lambda^2) & 1 \end{pmatrix}$$

$\lambda \approx 0.22$

# How to modify the SM?

the SM, as a consistent QFT, is completely specified by

0. invariance under local transformations of the gauge group  $G=SU(3)\times SU(2)\times U(1)$  [plus Lorentz invariance]
1. particle content      three copies of  $(q, u^c, d^c, l, e^c)$   
                                 one Higgs doublet  $\Phi$
2. renormalizability (i.e. the requirement that all coupling constants  $g_i$  have non-negative dimensions in units of mass:  $d(g_i) \geq 0$ . This allows to eliminate all the divergencies occurring in the computation of physical quantities, by redefining a finite set of parameters.)

(0.+1.+2.) leads to the SM Lagrangian,  $L_{SM}$ , possessing an additional, accidental, global symmetry: (B-L)       $\rightarrow$  EXERCISE

0. We cannot give up gauge invariance! It is mandatory for the consistency of the theory. Without gauge invariance we cannot even define the Hilbert space of the theory [remember: we need gauge invariance to eliminate the photon extra degrees of freedom required by Lorentz invariance]!  
We could extend  $G$ , but, to allow for neutrino masses, we need to modify 1. (and/or 2.) anyway...



## Exercise 1: anomalies of B and L<sub>i</sub>

the anomaly of the baryonic current and the individual leptonic currents are proportional to  $\text{tr}[Q \{T^A, T^B\}]$  and  $\text{tr}[Q \{Y, Y\}]$  where  $Q=(B, L_i)$  and  $(T^A, Y)$  are the generators of the electroweak gauge group  
compute these traces in the SM with 3 fermion generations

$$\frac{1}{2} \text{tr}[B \{T^A, T^B\}] = 3(\text{gen}) \times 3(\text{col}) \times \frac{1}{3}(B) \times \left[ \frac{1}{4}(\text{up}) + \frac{1}{4}(\text{down}) \right] \delta^{AB} = \frac{3}{2} \delta^{AB}$$

$$\frac{1}{2} \text{tr}[L_i \{T^A, T^B\}] = 1(L_i) \times \left[ \frac{1}{4}(\text{nu}) + \frac{1}{4}(e) \right] \delta^{AB} = \frac{1}{2} \delta^{AB}$$

$$\frac{1}{2} \text{tr}[B \{Y, Y\}] = 3(\text{gen}) \times 3(\text{col}) \times \frac{1}{3}(B) \times \left[ \frac{1}{18}(\text{Doubl}) - \frac{10}{18}(\text{Singl}) \right] = -\frac{3}{2}$$

$$\frac{1}{2} \text{tr}[L_i \{Y, Y\}] = 1(L_i) \times \left[ \frac{1}{2}(\text{Doubl}) - 1(\text{Singl}) \right] = -\frac{1}{2}$$

(B+L) is anomalous, (B/3-L<sub>i</sub>) [and (B-L)] are anomaly-free

# First possibility: modify (1), the particle content

there are several possibilities

one of the simplest one is to mimic the charged fermion sector

**Example 1**  $\left\{ \begin{array}{l} \text{add (three copies of)} \\ \text{right-handed neutrinos} \end{array} \right. \quad \nu^c \equiv (1,1,0) \quad \text{full singlet under} \\ G = \text{SU}(3) \times \text{SU}(2) \times \text{U}(1)$

ask for (global) invariance under B-L  
(no more automatically conserved as in the SM)

the neutrino has now four helicities, as the other charged fermions, and we can build gauge invariant Yukawa interactions giving rise, after electroweak symmetry breaking, to neutrino masses

$$L_Y = -d^c y_d (\Phi^+ q) - u^c y_u (\tilde{\Phi}^+ q) - e^c y_e (\Phi^+ l) - \nu^c y_\nu (\tilde{\Phi}^+ l) + h.c.$$

$$m_f = \frac{y_f}{\sqrt{2}} v \quad f = u, d, e, \nu$$

with three generations there is an exact replica of the quark sector and, after diagonalization of the charged lepton and neutrino mass matrices, a mixing matrix  $U$  appears in the charged current interactions

$$-\frac{g}{\sqrt{2}} W_\mu^- \bar{e} \sigma^\mu U_{PMNS} \nu + h.c. \quad U_{PMNS} \text{ has three mixing angles and one phase, like } V_{CKM}$$

## a generic problem of this approach

the particle content can be modified in several different ways in order to account for non-vanishing neutrino masses (additional right-handed neutrinos, new  $SU(2)$  fermion triplets, additional  $SU(2)$  scalar triplet(s), SUSY particles,...). Which is the correct one?

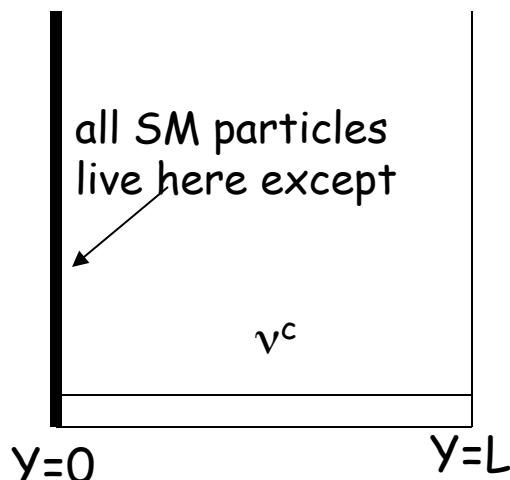
## a problem of the above example

if neutrinos are so similar to the other fermions, why are so light?

$$\frac{y_\nu}{y_{top}} \leq 10^{-12}$$

Quite a speculative answer:

neutrinos are so light, because the right-handed neutrinos have access to an extra (fifth) spatial dimension



neutrino Yukawa coupling

$\nu^c(y=0)(\tilde{\Phi}^+ l) = \text{Fourier expansion}$

$$= \frac{1}{\sqrt{L}} \nu_0^c(\tilde{\Phi}^+ l) + \dots \quad [\text{higher modes}]$$

if  $L \gg 1$  (in units of the fundamental scale)  
then neutrino Yukawa coupling is suppressed

# Second possibility: abandon (2) renormalizability

A disaster?

$$L = L_{d \leq 4}^{SM} + \frac{L_5}{\Lambda} + \frac{L_6}{\Lambda^2} + \dots$$

a new scale  $\Lambda$  enters the theory. The new (gauge invariant!) operators  $L_5, L_6, \dots$  contribute to amplitudes for physical processes with terms of the type

$$\frac{L_5}{\Lambda} \rightarrow \frac{E}{\Lambda} \quad \frac{L_6}{\Lambda^2} \rightarrow \left( \frac{E}{\Lambda} \right)^2 \quad \dots$$

the theory cannot be extrapolated beyond a certain energy scale  $E \approx \Lambda$ .  
[at variance with a renormalizable (asymptotically free) QFT]

If  $E \ll \Lambda$  (for example  $E$  close to the electroweak scale,  $10^2$  GeV, and  $\Lambda \approx 10^{15}$  GeV not far from the so-called Grand Unified scale), the above effects will be tiny and, the theory will *look like* a renormalizable theory!

$$\frac{E}{\Lambda} \approx \frac{10^2 \text{ GeV}}{10^{15} \text{ GeV}} = 10^{-13}$$

an extremely tiny effect, but exactly what needed to suppress  $m_\nu$  compared to  $m_{\text{top}}$ !

**Worth to explore.** The dominant operators (suppressed by a single power of  $1/\Lambda$ ) beyond  $L_{SM}$  are those of dimension 5. Here is a list of all d=5 gauge invariant operators

$$\frac{L_5}{\Lambda} = \frac{(\tilde{\Phi}^+ l)(\tilde{\Phi}^+ l)}{\Lambda} =$$

$$= \frac{v}{2} \left( \frac{v}{\Lambda} \right) \nu \nu + \dots$$

a unique operator!  
[up to flavour combinations]  
it violates (B-L) by two units

it is suppressed by a factor  $(v/\Lambda)$   
with respect to the neutrino mass term  
of Example 1:

$$\nu^c (\tilde{\Phi}^+ l) = \frac{v}{\sqrt{2}} \nu^c \nu + \dots$$

**it provides an explanation for the smallness of  $m_\nu$ :**

the neutrino masses are small because the scale  $\Lambda$ , characterizing (B-L) violations, is very large. How large? Up to about  $10^{15}$  GeV

from this point of view neutrinos offer a unique window on physics at very large scales, inaccessible in present (and probably future) man-made experiments.

since this is the dominant operator in the expansion of  $L$  in powers of  $1/\Lambda$ , we could have expected to find the first effect of physics beyond the SM in neutrinos ... and indeed this was the case!

$L_5$  represents the effective, low-energy description of several extensions of the SM

Example 2:  
see-saw

add (three copies of)  $\nu^c \equiv (1,1,0)$

full singlet under  
 $G = SU(3) \times SU(2) \times U(1)$

this is like Example 1, but without enforcing (B-L) conservation

$$L(\nu^c, l) = -\nu^c y_\nu (\tilde{\Phi}^+ l) - \frac{1}{2} \nu^c M \nu^c + h.c.$$

mass term for right-handed  
neutrinos:  $G$  invariant, violates  
(B-L) by two units.

the new mass parameter  $M$  is independent from the electroweak breaking scale  $v$ . If  $M \gg v$ , we might be interested in an effective description valid for energies much smaller than  $M$ . This is obtained by “integrating out” the field  $\nu^c$

$$L_{eff}(l) = \frac{1}{2} (\tilde{\Phi}^+ l) \left[ y_\nu^T M^{-1} y_\nu \right] (\tilde{\Phi}^+ l) + h.c. + \dots$$

terms suppressed by more  
powers of  $M^{-1}$

this reproduces  $L_5$ , with  $M$  playing the role of  $\Lambda$ . This particular mechanism is called (type I) **see-saw**.

## Exercise 2

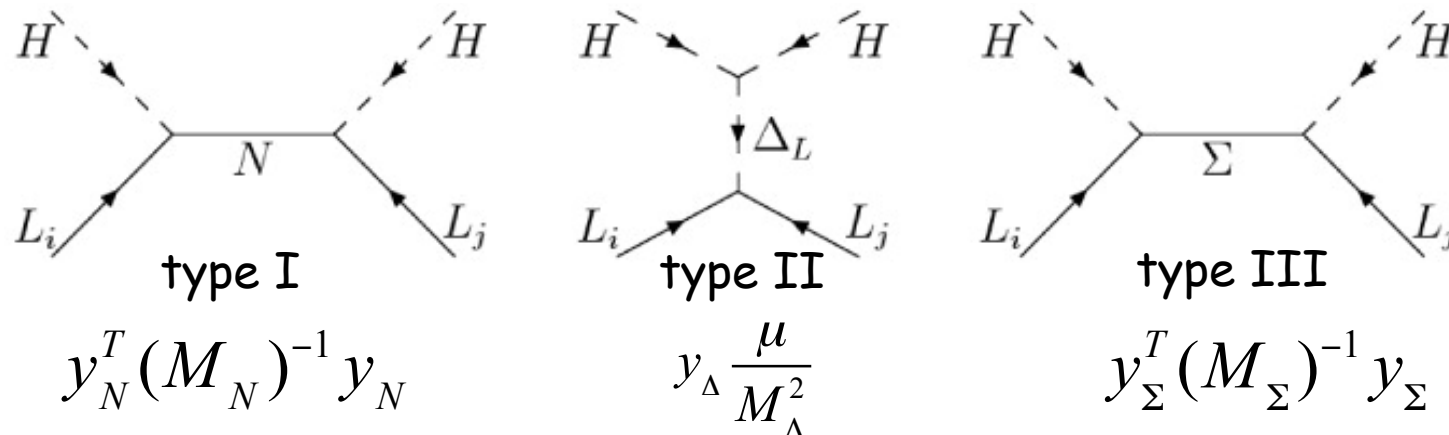
derive the see-saw relation by integrating out the fields  $\nu^c$  through their e.o.m. in the heavy  $M$  limit. Compute the 1<sup>st</sup> order corrections in  $p/M$

equations of motion of  $\nu^c$

$$\begin{pmatrix} \nu^c \\ \bar{\nu}^c \end{pmatrix} = \begin{pmatrix} i\bar{\sigma}^\mu \partial_\mu & -M^+ \\ -M & i\sigma^\mu \partial_\mu \end{pmatrix}^{-1} \begin{pmatrix} y_\nu^* \bar{\omega} \\ y_\nu \omega \end{pmatrix} = \begin{pmatrix} -M^{-1} y_\nu \omega \\ -M^{*-1} y_\nu^* \bar{\omega} \end{pmatrix} + \dots \quad \omega \equiv (\tilde{\Phi}^+ l)$$

$$L_{\text{eff}} = i\bar{l} \bar{\sigma}^\mu \partial_\mu l + \underbrace{\frac{1}{2} \left[ \omega (y_\nu^T M^{-1} y_\nu) \omega + h.c. \right]}_{d=5} + \underbrace{i\bar{\omega} (y_\nu^+ M^{*-1} M^{-1} y_\nu) \bar{\sigma}^\mu \partial_\mu \omega}_{d=6 \text{ renormalizes the KE of } \nu \text{ by } v^2/M^2} + O(M^{-3})$$

there are 3 types of see-saw depending on the particle we integrate out they all give rise to the same  $d=5$  operator



# Theoretical motivations for the see-saw

$\Lambda \approx 10^{15}$  GeV is very close to the so-called unification scale  $M_{\text{GUT}}$ .

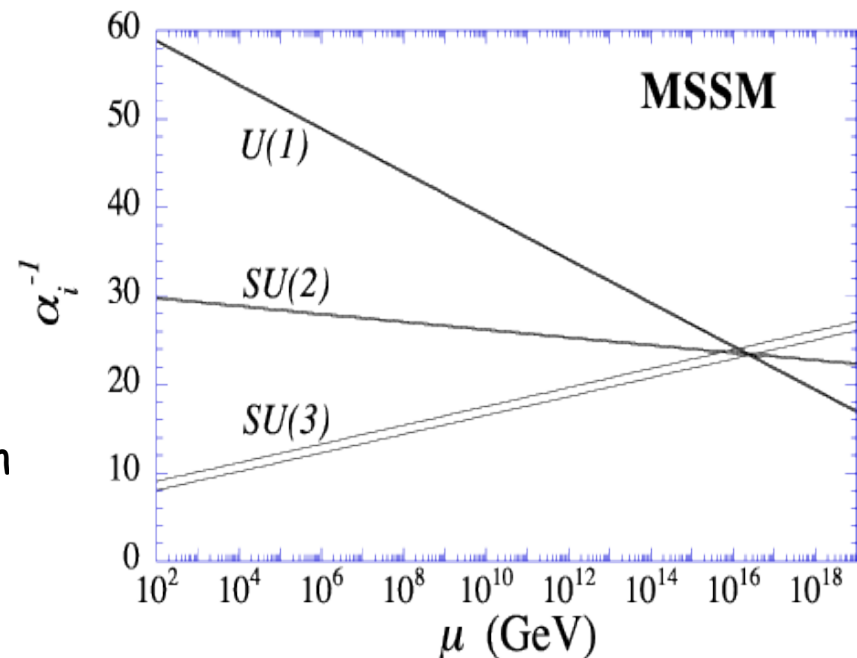
an independent evidence for  $M_{\text{GUT}}$  comes from the **unification of the gauge coupling constants** in (SUSY extensions of) the SM.

such unification is a generic prediction of **Grand Unified Theories** (GUTs): the SM gauge group  $G$  is embedded into a simple group such as  $SU(5)$ ,  $SO(10)$ ,....

**Particle classification:** it is possible to unify all SM fermions (1 generation) into a single irreducible representation of the GUT gauge group. Simplest example:  $G_{\text{GUT}} = SO(10)$

$16 = (q, d^c, u^c, l, e^c, \nu^c)$  a whole family plus a right-handed neutrino!

quite a fascinating possibility. Unfortunately, it still lacks experimental tests. In GUT new, very heavy, particles can convert quarks into leptons and the **proton is no more a stable particle**. Proton decay rates and decay channels are however model dependent. Experimentally we have only lower bounds on the proton lifetime.



Unity of All Elementary-Particle Forces  
Phys. Rev. Lett. 32, (1974) 438  
Howard Georgi and S. L. Glashow

Georgi, H.; Quinn, H.R. and Weinberg, S.  
Hierarchy of interactions in unified gauge theories.  
Phys. Rev. Lett. 33 (1974) 451



## Exercise 3: gauge coupling unification

$O^{\text{th}}$  order approximation

justify this  $\sqrt{\frac{5}{3}}g_Y = g_2 = g_3$

$$\sin^2 \vartheta_W = \frac{g_Y^2}{g_Y^2 + g_2^2} = \frac{3}{8} \approx 0.375$$

include 1-loop running

$$\frac{1}{\alpha_i(Q)} = \frac{1}{\alpha_i(m_Z)} + \frac{b_i}{2\pi} \log \frac{Q}{m_Z}$$

$$\begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix}_{MSSM} = \begin{pmatrix} 33/5 \\ 1 \\ -3 \end{pmatrix} \quad \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix}_{SM} = \begin{pmatrix} 41/10 \\ -19/6 \\ -7 \end{pmatrix}$$

knowledge of b.c.  $M_{GUT}$  and  $\alpha_U = \alpha(M_{GUT})$  would allow to predict  $\alpha_i(m_Z)$   
in practice, we use as inputs

$$\alpha_{em}^{-1}(m_Z) \Big|_{\overline{MS}} = 127.934 \quad \sin^2 \vartheta(m_Z) \Big|_{\overline{MS}} = 0.231$$

to predict  
[MSSM]

$$\alpha_3(m_Z) \Big|_{\overline{MS}} = \frac{7\alpha_{em}(m_Z)}{15\sin^2 \vartheta(m_Z) - 3} \approx 0.118$$

$$\alpha_U = \frac{28\alpha_{em}(m_Z)}{36\sin^2 \vartheta(m_Z) - 3} \approx \frac{1}{25}$$

[corrections from 2-loop RGE,  
threshold corrections at  $M_{SUSY}$ ,  
threshold corrections at  $M_{GUT}$ ]

$$\log \left( \frac{M_{GUT}}{m_Z} \right) = \pi \frac{3 - 8\sin^2 \vartheta(m_Z)}{14\alpha_{em}(m_Z)} \Rightarrow M_{GUT} \approx 2 \times 10^{16} \text{ GeV}$$

## Exercise 4: effective lagrangian for nucleon decay

recognize that, with the SM particle content, the lowest dimensional operators violating B occur at d=6. Make a list of them

$$\frac{1}{\Lambda_B^2} \times \begin{cases} qqu^{c+}e^{c+} & qqql \\ qlu^{c+}d^{c+} & u^cu^cd^ce^c \end{cases} \quad \begin{array}{l} \text{color and SU(2)} \\ \text{indices contracted} \end{array}$$

notice that they respect  $\Delta B = \Delta L$ : nucleon decay into antileptons  
e.g.  $p \rightarrow e^+ \pi^0$ ,  $n \rightarrow e^+ \pi^-$  [  $n \rightarrow e^- \pi^+$  suppressed by further powers of  $\Lambda_B$  ]

naïve estimate

$$\tau_p \approx \frac{\Lambda_B^4}{m_p^5}$$

assuming

$$\tau_p(p \rightarrow e^+ \pi^0) > 1.4 \times 10^{34} \text{ ys} \quad [\text{SK}]$$

we get

$$\Lambda_B > 2.6 \times 10^{16} \text{ GeV}$$

in GUTs  $\Lambda_B$  is related to the scale  $M_{\text{GUT}}$  at which the grand unified symmetry is broken down to SM gauge group

the observed proton stability is guaranteed by the largeness of  $M_{\text{GUT}}$

In SUSY extensions of the SM the lowest dimensional operators violating B occur at d=5: why?

flavor puzzle made simpler in  $SU(5)$  ?

Higgs

$$\bar{5} = (l, d^c) \quad 10 = (q, u^c, e^c) \quad 1 = \nu^c$$

$$\Phi_5 = (\Phi_D, \Phi_T)$$

$$L_Y = -10 y_u 10 \Phi_5 - \bar{5} y_d 10 \Phi_5^+ - 1 y_\nu \bar{5} \Phi_5 - \frac{1}{2} 1 M 1 + h.c.$$

$$y_d = y_e^T$$

$$m_b = m_\tau$$

$$m_s = m_\mu$$

$$m_d = m_e$$

O.K.

wrong, but not by orders of magnitude

can be fixed with additional Higgs

$$m_s \approx m_\mu / 3$$

$$m_d \approx 3 m_e$$

suppose that  $y_u, y_e, y_\nu$  and  $M/\Lambda$  are **anarchical matrices** [ $O(1)$  matrix elements] and that the observed hierarchy is due to the wave function renormalization of matter multiplets (we will see how later on)

$$10 \rightarrow F_{10} 10$$

$$\bar{5} \rightarrow F_{\bar{5}} \bar{5}$$

$$1 \rightarrow F_1 1$$

$$F_X = \begin{pmatrix} \lambda^{Q_{X_1}} & 0 & 0 \\ 0 & \lambda^{Q_{X_2}} & 0 \\ 0 & 0 & \lambda^{Q_{X_3}} \end{pmatrix}$$

$$\lambda \approx 0.22$$

$$Q_{X_1} \geq Q_{X_2} \geq Q_{X_3}$$

$F_1$  dependence  
cancels in  $m_\nu$

$$\mathcal{Y}_u = F_{10} y_u F_{10}$$

$$\mathcal{Y}_d = F_{\bar{5}} y_d F_{10}$$

$$\mathcal{Y}_e = F_{10} y_e^T F_{\bar{5}}$$

$$m_\nu \propto F_{\bar{5}} y_\nu^T M^{-1} y_\nu F_{\bar{5}}$$

large mixing in lepton sector suggests  $F_{\bar{5}} \approx \text{diag}(1, 1, 1)$

hierarchy mostly due to  $F_{10}$   $m_u : m_c : m_t \approx m_d^2 : m_s^2 : m_b^2 \approx m_e^2 : m_\mu^2 : m_\tau^2$

large  $l$  mixing corresponds to a large  $d^c$  mixing: unobservable in weak int. of quarks

how can a wave function renormalization (effectively) arise?

several possibilities

here (Exercise 5 ): bulk fermions in a compact extra dimension  $S^1/Z_2$

$$\mathcal{L} = i\bar{\Psi}_1 \Gamma^M \partial_M \Psi_1 + i\bar{\Psi}_2 \Gamma^M \partial_M \Psi_2 - m_1 \varepsilon(y) \bar{\Psi}_1 \Psi_1 + m_2 \varepsilon(y) \bar{\Psi}_2 \Psi_2 - \left[ \delta(y) \frac{y}{\Lambda} \bar{f}_1 (h + v) f_2 + h.c. \right]$$

$$\Psi_1 = \begin{pmatrix} E_1 \\ \bar{f}_1 \end{pmatrix} \quad \Psi_2 = \begin{pmatrix} f_2 \\ \bar{E}_2 \end{pmatrix}$$

solve the e.o.m. for the fermion zero modes with the b.c.

$$-\gamma_5 \partial_y \Psi_{1,2}^0 \pm m_{1,2} \varepsilon(y) \Psi_{1,2}^0 = 0$$

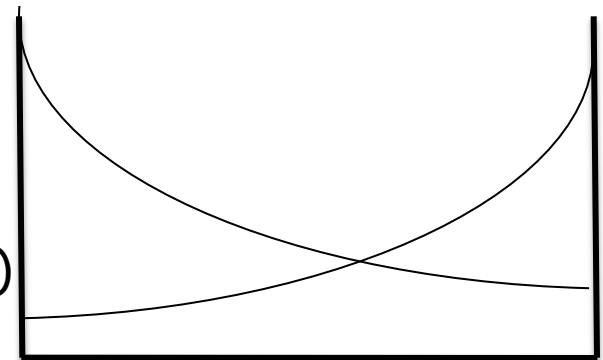
$$\Psi_1(-y) = +\gamma_5 \Psi_1(y)$$

$$\Psi_2(-y) = -\gamma_5 \Psi_2(y)$$

$$f_i^0(y) = \sqrt{\frac{2m_i}{1 - e^{-2m_i \pi R}}} e^{-m_i y}$$

vanishing zero-modes  
for  $(E_1, \bar{E}_2)$

$y \approx O(1)$

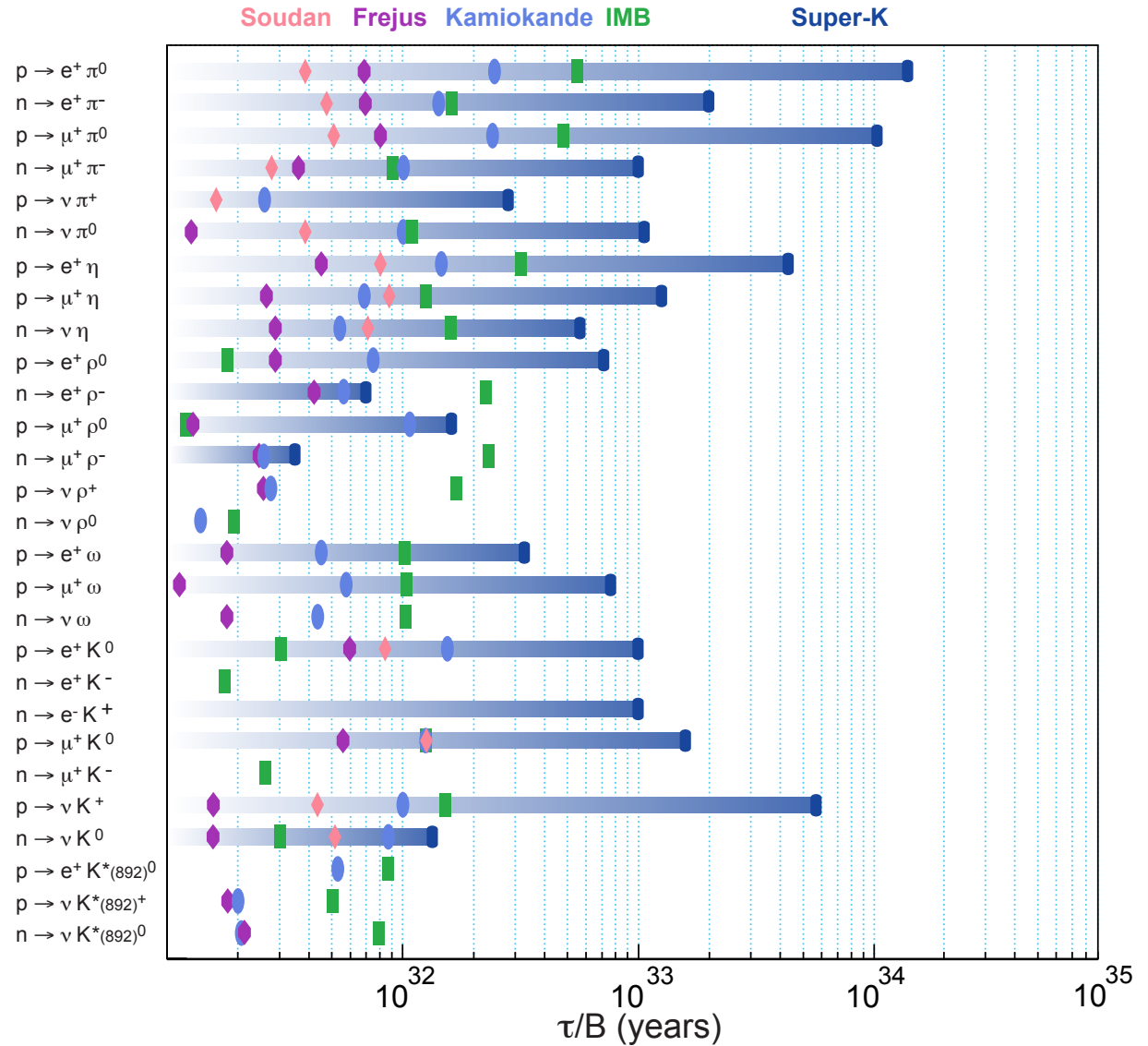


$$\mathcal{L}_Y = -\frac{1}{\Lambda \pi R} \bar{f}_1 (F_1 y F_2) (h + v) f_2$$

$$F_i = \sqrt{\frac{x_i}{1 - e^{-x_i}}} \approx \begin{cases} e^{-x_i/2} & x_i \gg 1 \\ 1 & x_i \approx 0 \\ \sqrt{-x_i} & x_i \ll -1 \end{cases}$$

Back up slides

# Antilepton + meson two-body modes



# Flavor symmetries I (the hierarchy puzzle)

hierarchies in fermion spectrum

quarks	$\frac{m_u}{m_t} \ll \frac{m_c}{m_t} \ll 1$ $\frac{m_d}{m_b} \ll \frac{m_s}{m_b} \ll 1$ $ V_{ub}  \ll  V_{cb}  \ll  V_{us}  \equiv \lambda < 1$
leptons	$\frac{m_e}{m_\tau} \ll \frac{m_\mu}{m_\tau} \ll 1$

$$\frac{\Delta m_{sol}^2}{\Delta m_{atm}^2} = (0.025 \div 0.049) \approx \lambda^2 \ll 1 \quad (2\sigma)$$

$$|U_{e3}| < 0.18 \leq \lambda \quad (2\sigma)$$

call  $\xi_i$  the generic small parameter. A modern approach to understand why  $\xi_i \ll 1$  consists in regarding  $\xi_i$  as small breaking terms of an approximate flavour symmetry. When  $\xi_i = 0$  the theory becomes invariant under a flavour symmetry  $F$

Example: why  $y_e \ll y_{top}$ ? Assume  $F = U(1)_F$

$F(t) = F(t^c) = F(h) = 0$	$y_{top} (h + v) t^c t$	allowed
$F(e^c) = p > 0$ $F(e) = q > 0$	$y_e (h + v) e^c e$	breaks $U(1)_F$ by $(p+q)$ units
if $\xi = \langle \varphi \rangle / \Lambda \ll 1$ breaks $U(1)$ by one negative unit		$y_e \approx O(\xi^{p+q}) \ll y_{top} \approx O(1)$

provides a qualitative picture of the existing hierarchies in the fermion spectrum

# GGI lectures on the theory of fundamental interactions 2015

Firenze, 12-16 January 2015

Aspects of neutrino physics (III)  
Neutrino Masses, Mixing and Oscillations:  
Leptogenesis and Hierarchy problem

Ferruccio Feruglio  
Universita' di Padova



The see-saw (continue)

## 2 additional virtues of the see-saw

The see-saw mechanism can enhance small mixing angles into large ones

$$m_\nu = -[y_\nu^T M^{-1} y_\nu] v^2$$

example

$$y_\nu = \begin{pmatrix} \delta & \delta \\ 0 & 1 \end{pmatrix} \quad \delta \ll 1$$

small mixing

$$M = \begin{pmatrix} M_1 & 0 \\ 0 & M_2 \end{pmatrix} \quad \text{no mixing}$$

$$y_\nu^T M^{-1} y_\nu = \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \frac{\delta^2}{M_1} + \begin{pmatrix} 0 & 0 \\ 0 & 1 \end{pmatrix} \frac{1}{M_2}$$
$$\approx \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \frac{\delta^2}{M_1} \quad \text{for } \frac{M_1}{M_2} \ll \delta^2$$

---

The (out-of equilibrium, CP-violating) decay of heavy right-handed neutrinos in the early universe might generate a net asymmetry between leptons and anti-leptons. Subsequent SM interactions can partially convert it into the observed baryon asymmetry

$$\eta = \frac{(n_B - n_{\bar{B}})}{s} \approx 6 \times 10^{-10}$$

Sakharov conditions met by the see-saw theory

1. (B-L) violation at high-temperature and (B+L) violation by pure SM interactions
2. C and CP violation by additional phases in see-saw Lagrangian (more on this later)
3. out-of-equilibrium condition

restrictions imposed by leptogenesis on neutrinos

active neutrinos should be light

here: thermal leptogenesis  
dominated by lightest  $\nu^c$   
no flavour effects]

out-of-equilibrium controlled  
by rate of RH neutrino decays

$$\frac{M_1}{8\pi} (y_\nu y_\nu^\dagger)_{11} < \frac{T^2}{M_{Pl}} \Big|_{T \approx M_1}$$

$$\frac{(y_\nu y_\nu^\dagger)_{11} v^2}{M_1} \equiv \tilde{m}_1 < 10^{-3} \text{ eV}$$

Exercise 6; compute this

more accurate estimate

$$m_i < 0.15 \text{ eV}$$

RH neutrinos should be heavy

$$\eta_B \approx 10^{-2} \varepsilon_1 \eta$$

[efficiency factor  $\leq 1$   
washout effects]

$$\varepsilon_1 = \frac{\Gamma(\nu_1^c \rightarrow l\Phi) - \Gamma(\nu_1^c \rightarrow \bar{l}\Phi^*)}{\Gamma(\nu_1^c \rightarrow l\Phi) + \Gamma(\nu_1^c \rightarrow \bar{l}\Phi^*)} = -\frac{3}{16\pi} \sum_{j=2,3} \frac{M_1}{M_j} \frac{\text{Im}\{[(yy^\dagger)_{1j}]^2\}}{(yy^\dagger)_{11}} \approx 0.1 \times \frac{M_1 m_i}{v^2}$$

[Yukawas  $y$  in mass eigenstate basis for  $\nu_i^c$ ]

$$M_1 > 6 \times 10^8 \text{ GeV}$$

more refined bound [Davidson and Ibarra 0202239]

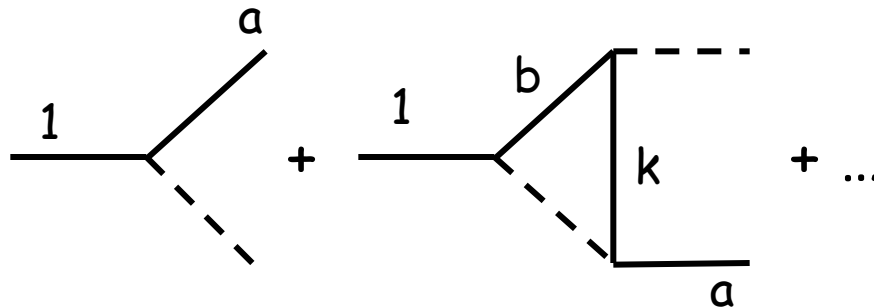
$$|\varepsilon_1^\infty| \leq \varepsilon_1^{DI} = \frac{3}{16\pi} \frac{M_1}{v^2} (m_3 - m_1)$$

$$T_R \approx M_1 > (4 \times 10^8 \div 2 \times 10^9) \text{ GeV}$$

in conflict with the bound on  $T_R$  in SUSY models to avoid overproduction of gravitinos

$$T_R^{SUSY} < 10^{7-9} \text{ GeV}$$

Exercise 7: reconstruct the flavour structure of  $\varepsilon_1$



$$\begin{aligned} \mathcal{A}(v_1^c \rightarrow l_a \Phi) &\propto y_{a1}^+ + W y_{1b} y_{bk}^+ y_{ak}^+ \\ \mathcal{A}(v_1^c \rightarrow \bar{l}_a \Phi^*) &\propto y_{1a} + W y_{b1}^+ y_{kb} y_{ka} \end{aligned}$$

$$\varepsilon_1 \propto \frac{\left| y_{a1}^+ + W y_{1b} y_{bk}^+ y_{ak}^+ \right|^2 - \left| y_{1a} + W y_{b1}^+ y_{kb} y_{ka} \right|^2}{\left| y_{a1}^+ + W y_{1b} y_{bk}^+ y_{ak}^+ \right|^2 + \left| y_{1a} + W y_{b1}^+ y_{kb} y_{ka} \right|^2} \approx \frac{\text{Im}(W) \text{Im}\{[(yy^+)_{1k}]^2\}}{(yy^+)_{11}}$$

[sums understood]

$$\text{Im}(W) \approx \frac{M_1}{M_k}$$

## Exercise 8: count the number of physical parameters in the type I see-saw model distinguish between moduli and phases

$y_e, y_\nu$  and  $M$  depend on  $(18+18+12)=48$  parameters, 24 moduli and 24 phases

we are free to choose any basis leaving the kinetic terms canonical  
(and the gauge interactions unchange)

$$e^c \rightarrow \Omega_{e^c} e^c \quad \nu^c \rightarrow \Omega_{\nu^c} \nu^c \quad l \rightarrow \Omega_l l \quad [U(3)^3]$$

these transformations contain 27 parameters (9 angles and 18 phases)  
and effectively modify  $y_e, y_\nu$  and  $M$

$$y_e \rightarrow \Omega_{e^c}^T y_e \Omega_l \quad y_\nu \rightarrow \Omega_{\nu^c}^T y_\nu \Omega_l \quad M \rightarrow \Omega_{\nu^c}^T M \Omega_{\nu^c}$$

so that we can remove 27 parameters from  $y_e, y_\nu$  and  $M$

we remain with 21 parameters: 15 moduli and 6 phases  
the moduli are 9 physical masses and 6 mixing angles

the same count in the quark sector would give a total of 9 moduli  
(6 masses and 3 mixing angles) and 0 phases <- wrong  
how the above argument should be modified, in general?

# weak point of the see-saw

full high-energy theory is difficult to test

$$L(\nu^c, l) = \nu^c y_\nu (\tilde{\Phi}^+ l) + \frac{1}{2} \nu^c M \nu^c + h.c.$$

depends on many physical parameters:

3 (small) masses + 3 (large) masses

3 (L) mixing angles + 3 (R) mixing angles

6 physical phases = 18 parameters

the double of those

describing  $(L_{SM}) + L_5$ :

3 masses, 3 mixing angles

and 3 phases, as in lecture 1

few observables to pin down the extra parameters:  $\eta, \dots$

[additional possibilities exist under special conditions, e.g. Lepton Flavor Violation at observable rates]

easier to test the low-energy remnant  $L_5$

[which however is “universal” and does not imply the specific see-saw mechanism of Example 2]

look for a process where B-L is violated by 2 units. The best candidate is

$0\nu\beta\beta$  decay:

$$(A, Z) \rightarrow (A, Z+2) + 2e^-$$

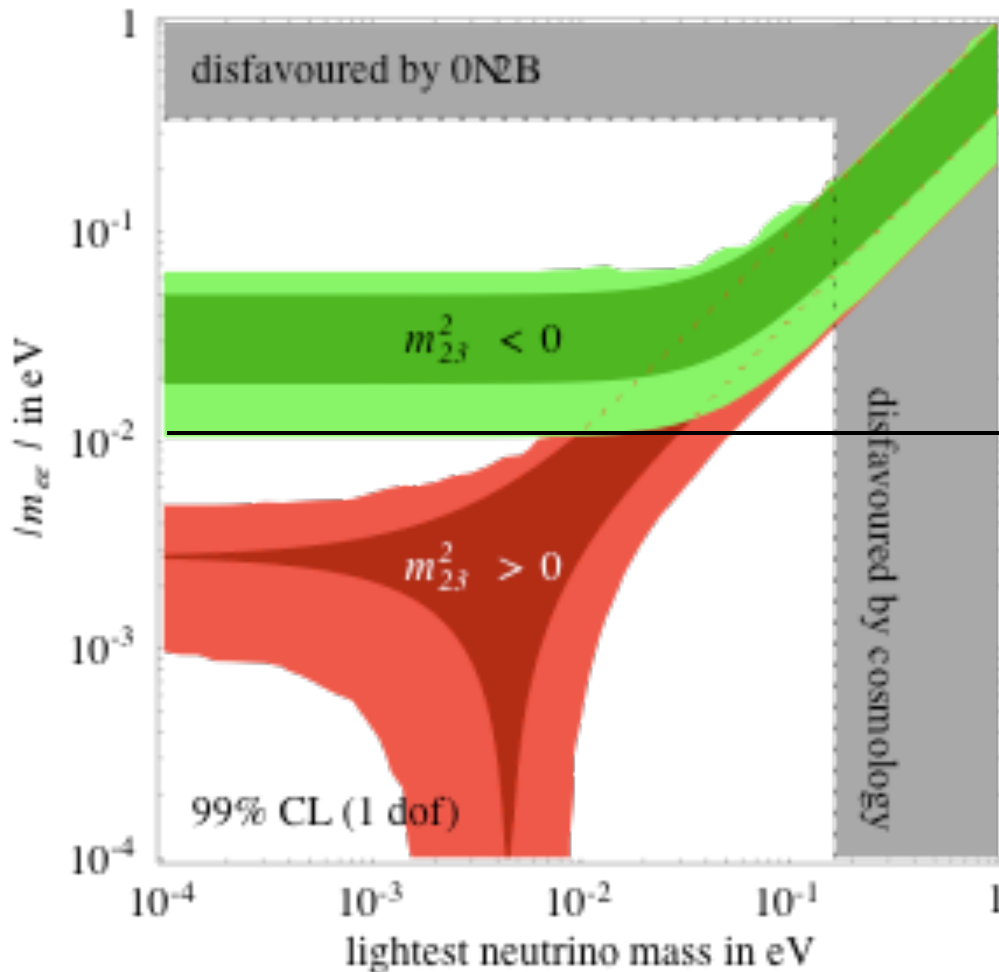
this would discriminate  $L_5$  from other possibilities, such as Example 1.

The decay in  $0\nu\beta\beta$  rates depend on the combination

$$|m_{ee}| = \left| \sum_i U_{ei}^2 m_i \right|$$

$$|m_{ee}| = \left| \cos^2 \vartheta_{13} (\cos^2 \vartheta_{12} m_1 + \sin^2 \vartheta_{12} e^{2i\alpha} m_2) + \sin^2 \vartheta_{13} e^{2i\beta} m_3 \right|$$

[notice the two phases  $\alpha$  and  $\beta$ , not entering neutrino oscillations]



from the current knowledge of  $(\Delta m_{ij}^2, \vartheta_{ij})$  we can estimate the expected range of  $|m_{ee}|$

future expected sensitivity  
on  $|m_{ee}|$

10 meV

a positive signal would test both  $L_5$  and the absolute mass spectrum at the same time!

# Neutrinos and the Higgs boson

1. neutrinos and the hierarchy problem
2. neutrinos and the stability of the electroweak vacuum

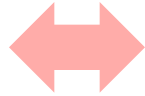


1.

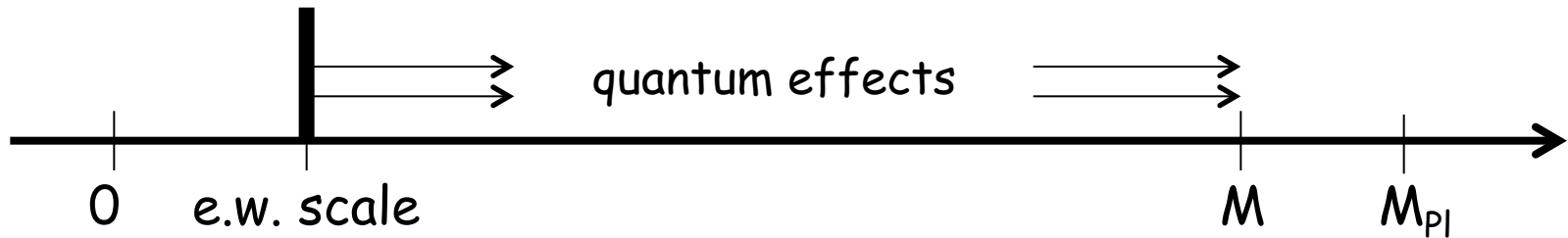
Why

any new particle threshold:  $M_{\text{GUT}, \dots}$

e.w. scale  $\ll \dots, M_{\text{Pl}}$  ?

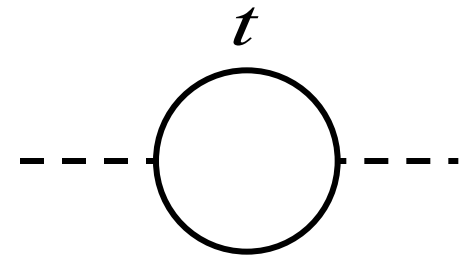


sensitivity of  $m_h$  to UV physics



often discussed in terms of quadratic divergences

$$\delta m_h^2 \propto \frac{y_t^2}{16\pi^2} \Lambda^2$$



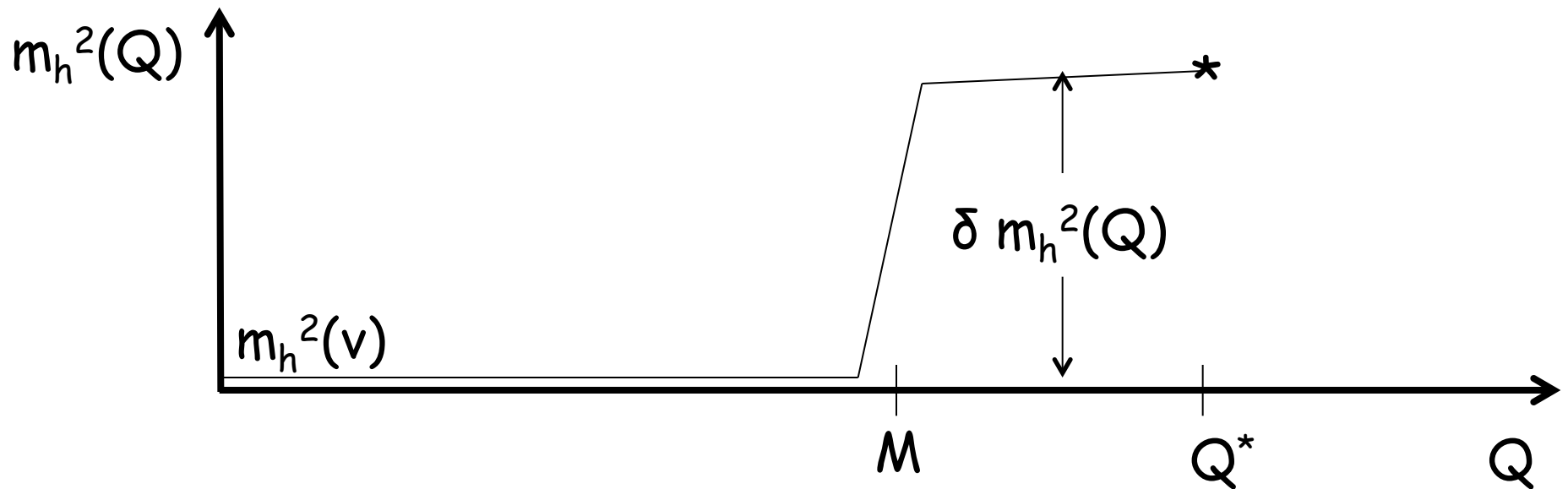
but

- what represents exactly  $\Lambda$  ? Any evidence from experiment?
- can we get rid of  $\Lambda$  in some suitable scheme ?
- technical aspect obscure physics

hierarchy problem can be formulated entirely in terms of renormalized quantities with no reference to regulators

**assumption:** coupling  $y$  of Higgs particle to an heavy state of mass  $M$

running Higgs mass  $\delta m_h^2(Q) \approx \frac{y^2}{16\pi^2} M^2 \log \frac{Q}{M} \quad Q > M$



fine-tune the initial conditions at  $Q^*$  such that

$$m_h^2(v) \approx m_h^2(Q^*) - \frac{y^2}{16\pi^2} M^2 \log \frac{Q^*}{M}$$

consider type I see-saw

heavy state $\nu^c$	mass $M$
Yukawa coupling	$y_\nu$

we will see  
in a moment

$$\delta m_h^2(Q) \approx -\frac{y_\nu^2}{4\pi^2} M^2 \log \frac{Q}{M} \quad Q > M$$

by using  $m_\nu \approx \frac{y_\nu^2 v^2}{M}$  to eliminate the  $y^2$  dependence

$$|\delta m_h^2(Q)| \approx \frac{1}{4\pi^2} \frac{m_\nu M^3}{v^2} \log \frac{Q}{M} < v^2$$

$$M < 1.4 \times 10^7 \text{ GeV}$$

$$\left[ \begin{array}{l} \log \frac{Q}{M} \approx 1 \\ m_\nu \approx 0.05 \text{ eV} \end{array} \right]$$

$$y_\nu \approx \sqrt{\frac{m_\nu M}{v^2}} < 10^{-4}$$

too small for thermal leptogenesis ?

Exercise 9: derive the threshold corrections to  $m_\sigma^2(Q)$  in the toy model

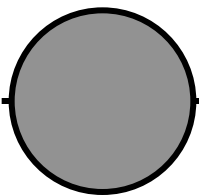
$$L = \frac{1}{2} \partial_\mu \sigma \partial^\mu \sigma + i \bar{\xi} \bar{\sigma}^\mu \partial_\mu \xi - \frac{1}{2} \left[ \xi^T \mathcal{M} \xi + \text{h.c.} \right]$$

assume  $m_\sigma^2(0) = 0$

$$\xi = \begin{pmatrix} \nu \\ \nu^c \end{pmatrix}$$

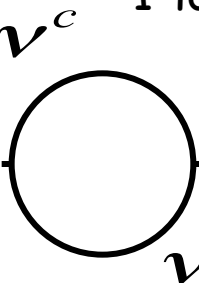
$$\mathcal{M}(\sigma) = \begin{pmatrix} 0 & y(\sigma + \nu) \\ y(\sigma + \nu) & M \end{pmatrix}$$

1. start from the 1-loop renormalized self-energy

$$\sigma \text{ --- } \text{---} \text{---} \sigma = i \left[ Q^2 - \Pi_f(Q^2) \right]$$


$$\Pi_f(Q^2) = \underbrace{\Pi(Q^2)}_{\text{1-loop}} - \underbrace{\Pi(0) + Q^2 \Pi'(0)}_{\text{counterterms}}$$

$$\text{OS scheme} \begin{cases} \Pi_f(0) = 0 \\ \Pi'_f(0) = 0 \end{cases}$$

$$\sigma \text{ --- } \text{---} \text{---} \sigma = -i \Pi(Q^2)$$


2. evaluate 1-loop diagram  $-i\Pi(Q^2)$  in the limit  $0 \approx m_1 \ll m_2 \approx M$

$$m_{1,2} = \frac{1}{2}(M \pm \sqrt{M^2 + 4y^2v^2}) \approx \begin{cases} -y^2v^2 / M \\ M + y^2v^2 / M \end{cases}$$

in dimensional regularization

$$\Pi(Q^2) = \frac{y^2}{2\pi^2} \int_0^1 dx \left[ (D - \log \Omega)(2\Omega - Q^2x(1-x)) + \Omega \right]$$

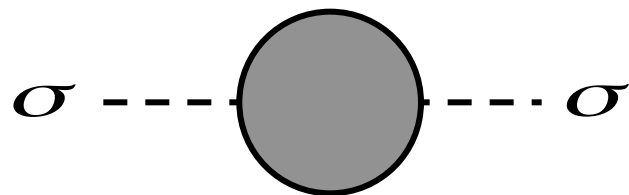
$$D = \frac{2}{\varepsilon} - \gamma + \log 4\pi$$

$$\Omega = -Q^2x(1-x) + M^2x$$

3. compute  $\Pi_f(Q^2)$

$$\Pi_f(Q^2) = \frac{y^2}{2\pi^2} \int_0^1 dx \left[ -2Q^2x(1-x) - (2M^2x - 3Q^2x(1-x)) \log \frac{\Omega}{M^2x} \right] \quad \text{finite}$$

relevant limits  $Q^2 \ll M^2$   $\Pi_f(Q^2) = -\frac{y^2}{12\pi^2} \frac{Q^4}{M^2} + \dots$

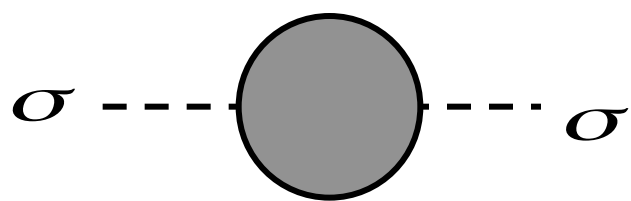


$$= iQ^2 \left[ 1 + \frac{y^2}{12\pi^2} \frac{Q^2}{M^2} + \dots \right]$$

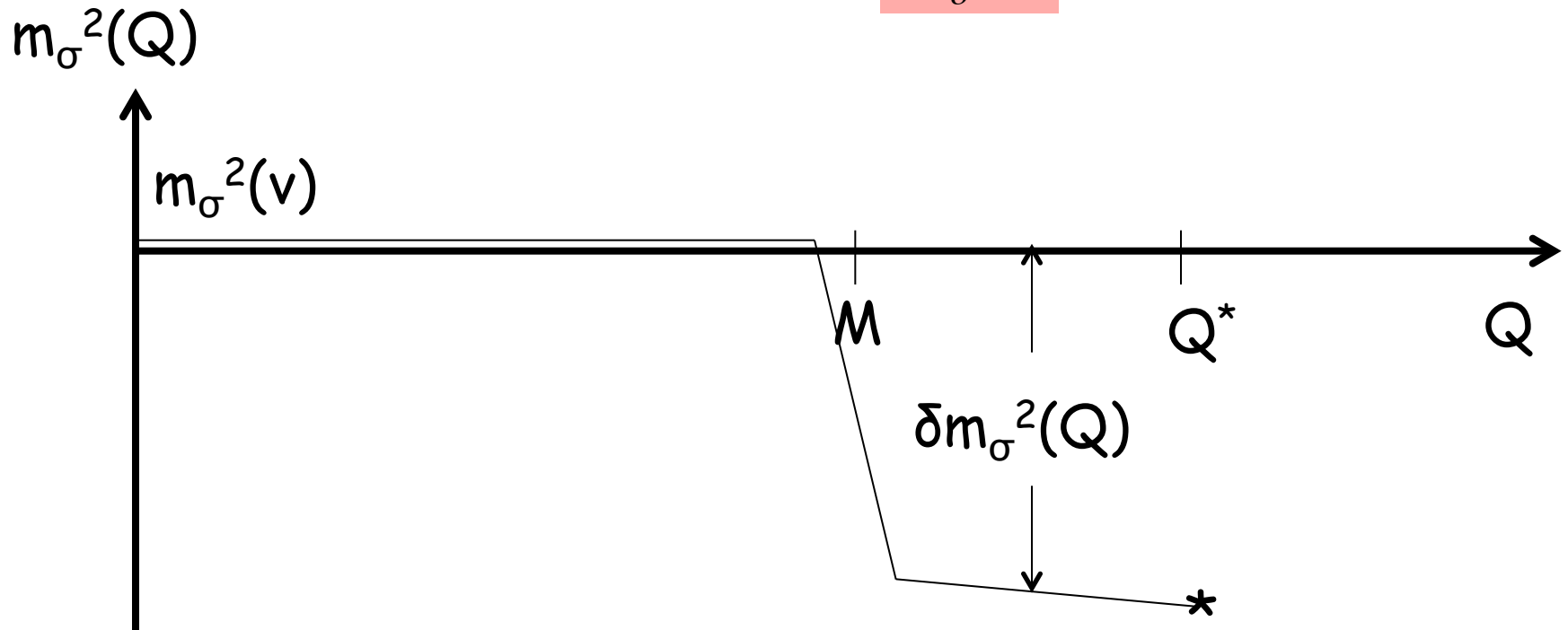
$$m_\sigma^2(Q) = 0$$

decoupling

$$Q^2 \gg M^2 \quad \Pi_f(Q^2) = \frac{y^2}{2\pi^2} \left[ Q^2 \left( -\frac{3}{4} + \frac{1}{4} \log \frac{-Q^2}{M^2} \right) - M^2 \log \frac{-Q^2}{M^2} \right] + \dots$$



$$= i \left( Q^2 + \underbrace{\frac{y^2}{2\pi^2} M^2 \log \frac{-Q^2}{M^2}}_{-m_\sigma^2(Q)} \right) (1 + O(y^2))$$



similar conclusions in type II and type III see-saw where threshold corrections are dominated by 2-loop gauge interactions

$$\text{type III} \quad \delta m_h^2(Q) \approx -\frac{72g^4}{(4\pi)^4} M^2 \log \frac{Q}{M} \quad Q > M \quad M < 940 \text{ GeV}$$

$$\text{type II} \quad M < 200 \text{ GeV}$$

ways out

the initial conditions at the scale  $Q^*$  are fine-tuned to an accuracy of order (e.w. scale)/ $M$

the threshold correction at the scale  $M$  is almost cancelled by an other contribution, as e.g. in supersymmetry with a splitting between neutrinos and sneutrinos of order  $4\pi \times (\text{e.w. scale})$

the Higgs is not an elementary particle and dissolves above a compositeness scale  $\sim \text{TeV}$

gap between the e.w. scale and the compositeness scale if the Higgs is a PGB

## 2. neutrinos and the stability of the electroweak vacuum

for the current values

$$m_h = (125.66 \pm 0.34) \text{ GeV}$$

$$m_t = (173.2 \pm 0.9) \text{ GeV}$$

$$\alpha_s(m_Z) = 0.1184 \pm 0.0007$$

assumption: only SM all the way up to the scale  $\Lambda$

for large values of the field  $h$

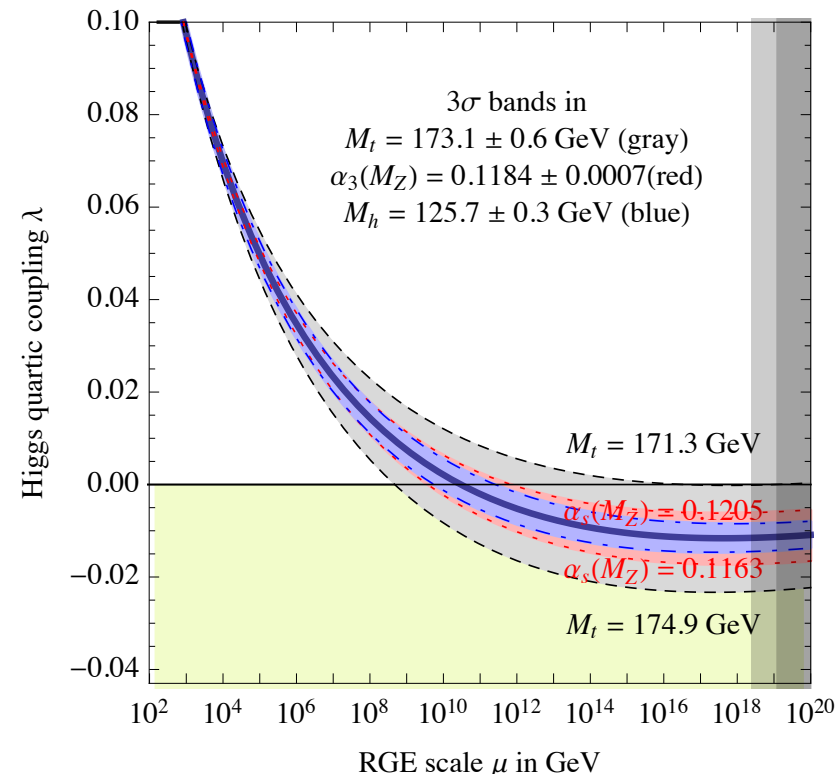
$$V(h) \approx \frac{\lambda}{4} h^4$$

$$(4\pi)^2 \frac{d\lambda}{dt} = -6y_t^4 + \frac{3}{8}[2g^4 + (g^2 + g'^2)] + 12\lambda y_t^2 - 3\lambda(g^2 + 3g'^2) + 24\lambda^2 + \dots$$

$\underbrace{\hspace{10em}}_{O(\lambda)} \qquad \underbrace{\hspace{10em}}_{O(\lambda^2)}$

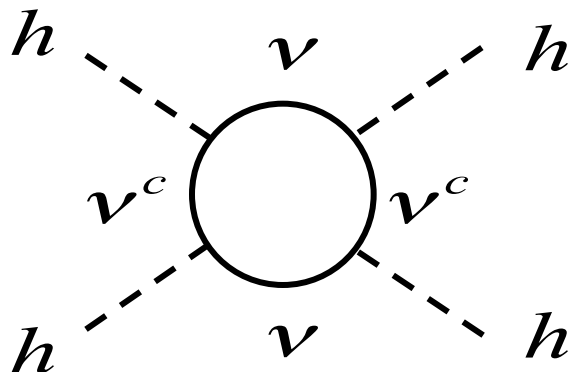
the Higgs potential develops an instability at

$$10^9 \text{ GeV} < \Lambda < 10^{15} \text{ GeV}$$



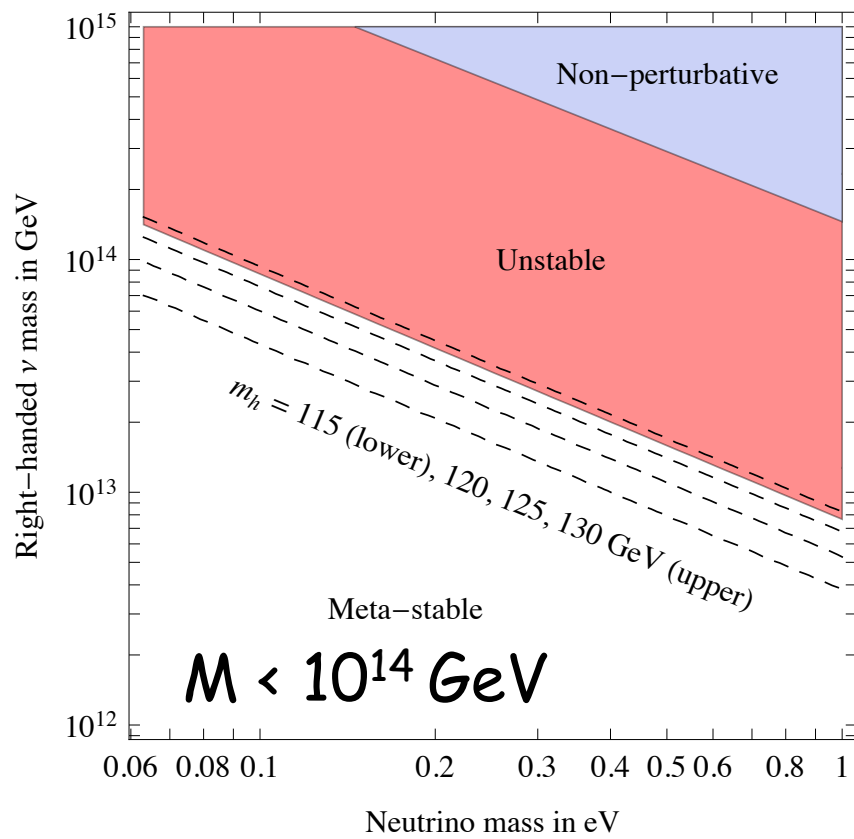


above the scale  $M$  a new contribution to  $\beta_\lambda$  arises from neutrino Yukawa couplings



$$\delta\beta_\lambda = -2\text{tr}(y_\nu y_\nu^+ y_\nu y_\nu^+) < 0$$

contributes to instability above  $M$



the larger  $M$ ,  
the larger the contribution

$$y_\nu \approx \sqrt{\frac{m_\nu M}{v^2}}$$

the bound applies only to the  
portion of SM parameter space  
that guarantees a stable vacuum  
in the limit  $y_\nu=0$   
( $m_t$  on the lower side  
 $\alpha_s$  on the higher side)

Back up slides

# Type-III see-saw at LHC

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## Abstract

Neutrino masses can be generated by fermion triplets with TeV-scale mass, that would manifest at LHC as production of two leptons together with two heavy SM vectors or higgs, giving rise to final states such as  $2\ell + 4j$  (that can violate lepton number and/or lepton flavor) or  $\ell + 4j + \cancel{E}_T$ . We devise cuts to suppress the SM backgrounds to these signatures. Furthermore, for most of the mass range suggested by neutrino data, triplet decays are detectably displaced from the production point, allowing to infer the neutrino mass parameters. We compare with LHC signals of type-I and type-II see-saw.

# GGI lectures on the theory of fundamental interactions 2015

Firenze, 12-16 January 2015

Aspects of neutrino physics (IV)  
Neutrino Masses, Mixing and Oscillations:  
Neutrinos and Lepton Flavor Violation

Ferruccio Feruglio  
Universita' di Padova

# Lecture 3

## Neutrinos and LFV

# LFV expected at some level

neutrino masses  
and  $U_{\text{PMNS}} \neq 1$



$L_i$  violated ( $i=e,\mu,\tau$ )

evidence for lepton flavor conversion

direct

$$\nu_e \rightarrow \nu_\mu, \nu_\tau$$

sol, LBL exp

indirect

$$\nu_\mu \rightarrow \nu_\tau$$

atm

should show up in processes with charged leptons

Process	Relative probability	Present Limit	Experiment	Year	prospects
$\mu \rightarrow e\gamma$	1	$5.7 \times 10^{-13}$	MEG	2012	$6 \times 10^{-14}$
$\mu^- \text{Ti} \rightarrow e^- \text{Ti}$	$Z\alpha/\pi$	$4.3 \times 10^{-12}$	SINDRUM II	2006	} $10^{-15} \div 10^{-16}$
$\mu^- \text{Au} \rightarrow e^- \text{Au}$	$Z\alpha/\pi$	$7 \times 10^{-13}$	SINDRUM II	2006	
$\mu \rightarrow eee$	$\alpha/\pi$	$4.3 \times 10^{-12}$	SINDRUM	1988	
$\tau \rightarrow \mu\gamma$	$(m_\tau/m_\mu)^{2\div 4}$	$3.3 \times 10^{-8}$	B-factories	2011	
$\tau \rightarrow e\gamma$	$(m_\tau/m_\mu)^{2\div 4}$	$4.5 \times 10^{-8}$	B-factories	2011	

Table 1: Relative sensitivities and experimental limits of the main CLFV processes.

here: focus on radiative decays of charged leptons

in the SM, minimally extended to accommodate e.g. Dirac neutrinos

$$BR(\mu \rightarrow e\gamma) \approx \frac{3\alpha}{32\pi} \left| U_{\mu i}^* U_{ei} \frac{m_i^2}{m_W^2} \right|^2 \approx 10^{-53}$$



Exercise 10:  
reproduce this

[solution in  
Cheng and Li]

[unobservable also within type I see-saw]  $m_i \approx 0.05 \text{ eV}$   $U_{fi} \approx O(1)$

depleted by

- weak interactions
- loop factor
- GIM mechanism (mixing angle large, but neutrino masses tiny)

$\leftrightarrow$

GIM suppression  
for quarks:  
small mixing angles  
large top mass

a good place to look for BSM physics

general parametrization of LFV effects BSM

$$L = L_{SM} + \sum_i c_i^5 \frac{O_i^5}{\Lambda} + \sum_i c_i^6 \frac{O_i^6}{\Lambda^2} + \dots$$

$O_i^d$  gauge invariant  
operators dimension  $d$

# low-energy effective Lagrangian in the lepton sector

$$L = L_{SM} + i \frac{e}{\Lambda^2} e^c \left( \sigma^{\mu\nu} F_{\mu\nu} \right) \mathcal{Z} (\Phi^+ l) + \frac{1}{\Lambda^2} [4\text{-fermion}] + h.c. + \dots$$

[relation between the scale  $\Lambda$  and new particle masses  $M'$  can be non-trivial in a weakly interacting theory  $g \Lambda / 4\pi \approx M'$ ]

$\mathcal{Z}_{ij}$  a matrix in flavour space

$$L_Y = -e^c y_e (\Phi^+ l) + h.c. + \dots$$

in the basis where charged leptons are diagonal

$$\text{Im}[\mathcal{Z}]_{ii}$$

$$d_i$$

electric dipole moments

$$\text{Re}[\mathcal{Z}]_{ii}$$

$$a_i = \frac{(g-2)_i}{2}$$

anomalous magnetic moments

$$|\mathcal{Z}_{ij}|^2 \quad (i \neq j)$$

$$R_{ij} = \frac{BR(l_i \rightarrow l_j \gamma)}{BR(l_i \rightarrow l_j \nu_i \bar{\nu}_j)}$$

radiative decays

$$\mu \rightarrow e \gamma \quad \tau \rightarrow \mu \gamma \quad \tau \rightarrow e \gamma$$

[4-fermion operators]

other LFV transitions

$$\mu \rightarrow eee \quad \tau \rightarrow \mu\mu\mu \quad \tau \rightarrow eee \quad \dots$$

$$BR(\mu \rightarrow e \gamma) < 5.7 \times 10^{-13}$$

$$\frac{\mathcal{Z}_{\mu e}}{\Lambda^2} < 2 \times 10^{-9} \text{ TeV}^{-2}$$



either the scale of new physics is very large or flavour violation from New Physics is highly non-generic

$$\Lambda > 2 \times 10^4 \left[ \sqrt{\mathcal{Z}_{\mu e}} \right] \text{ TeV}$$



# not a specific problem of the lepton sector

here: constraints from flavour physics on d=6  $|\Delta F|=2$  operators

FLAVOUR PROBLEM

Operator	Bounds on $\Lambda$ in TeV ( $c_{ij} = 1$ )		Bounds on $c_{ij}$ ( $\Lambda = 1$ TeV)		Observables
	Re	Im	Re	Im	
$(\bar{s}_L \gamma^\mu d_L)^2$	$9.8 \times 10^2$	$1.6 \times 10^4$	$9.0 \times 10^{-7}$	$3.4 \times 10^{-9}$	$\Delta m_K; \epsilon_K$
$(\bar{s}_R d_L)(\bar{s}_L d_R)$	$1.8 \times 10^4$	$3.2 \times 10^5$	$6.9 \times 10^{-9}$	$2.6 \times 10^{-11}$	$\Delta m_K; \epsilon_K$
$(\bar{c}_L \gamma^\mu u_L)^2$	$1.2 \times 10^3$	$2.9 \times 10^3$	$5.6 \times 10^{-7}$	$1.0 \times 10^{-7}$	$\Delta m_D;  q/p , \phi_D$
$(\bar{c}_R u_L)(\bar{c}_L u_R)$	$6.2 \times 10^3$	$1.5 \times 10^4$	$5.7 \times 10^{-8}$	$1.1 \times 10^{-8}$	$\Delta m_D;  q/p , \phi_D$
$(\bar{b}_L \gamma^\mu d_L)^2$	$5.1 \times 10^2$	$9.3 \times 10^2$	$3.3 \times 10^{-6}$	$1.0 \times 10^{-6}$	$\Delta m_{B_d}; S_{\psi K_S}$
$(\bar{b}_R d_L)(\bar{b}_L d_R)$	$1.9 \times 10^3$	$3.6 \times 10^3$	$5.6 \times 10^{-7}$	$1.7 \times 10^{-7}$	$\Delta m_{B_d}; S_{\psi K_S}$
$(\bar{b}_L \gamma^\mu s_L)^2$	$1.1 \times 10^2$		$7.6 \times 10^{-5}$		$\Delta m_{B_s}$
$(\bar{b}_R s_L)(\bar{b}_L s_R)$	$3.7 \times 10^2$		$1.3 \times 10^{-5}$		$\Delta m_{B_s}$

TABLE I: Bounds on representative dimension-six  $\Delta F = 2$  operators. Bounds on  $\Lambda$  are quoted assuming an effective coupling  $1/\Lambda^2$ , or, alternatively, the bounds on the respective  $c_{ij}$ 's assuming  $\Lambda = 1$  TeV. Observables related to CPV are separated from the CP conserving ones with semicolons. In the  $B_s$  system we only quote a bound on the modulo of the NP amplitude derived from  $\Delta m_{B_s}$  (see text). For the definition of the CPV observables in the  $D$  system see Ref. [15].

[Isidori, Nir, Perez, 2010]

# Minimal Flavour Violation (quarks)

[Chivukula, Georgi 1987  
D' Ambrosio, Giudice, Isidori, Strumia 2002]

useful benchmark: a framework where the only source of flavour violation beyond the SM are the Yukawa coupling. Well-defined in the quark sector.

in the limit  $y_u = y_d = 0$ , the SM lagrangian is invariant under a  $U(3)^3$  flavour symmetry

$$G_q = SU(3)_{u^c} \times SU(3)_{d^c} \times SU(3)_q \times \dots$$

$$q = (1, 1, 3) \quad u^c = (\bar{3}, 1, 1) \quad d^c = (1, \bar{3}, 1)$$

if the Yukawa couplings  $y_u$  and  $y_d$  are promoted to non-dynamical fields (spurions) transforming conveniently, the SM lagrangian remains formally invariant under the flavour group  $G_q$

$$L_{SM} = \dots - d^c y_d (\Phi^+ q) - u^c y_u (\tilde{\Phi}^+ q) + h.c.$$

$$y_u = (3, 1, \bar{3})$$

$$y_d = (1, 3, \bar{3})$$

MFV assumes that new operators coming from New Physics do not involve any additional field/spurions and that they are still invariant under  $G_q$   
[additional assumption: no additional sources of CPV other than those in  $y_{u,d}$ ]

# Exercise 11: build the leading operator contributing to $b \rightarrow s \gamma$ in MFV

a convenient basis:

$$y_d = \hat{y}_d \quad y_u = \hat{y}_u V_{CKM}$$

$$\hat{y}_{u,d} \text{ diagonal}$$

leading order MFV invariant

$$i \frac{e}{\Lambda^2} d^c \left( \sigma^{\mu\nu} F_{\mu\nu} \right) Z^d (\Phi^+ q) + h.c.$$

$$\begin{aligned} Z^d &= y_d y_u^+ y_u \\ &= \frac{2\sqrt{2}}{v^3} \left( \hat{m}_d V_{CKM}^+ \hat{m}_u^2 V_{CKM} \right) \\ \hat{m}_u &\approx \text{diag}(0, 0, m_t) \end{aligned}$$

$$b \rightarrow s \gamma \quad \Leftrightarrow \quad \left( Z^d \right)_{32}^*, \quad \left( Z^d \right)_{23}$$

$$\left( Z^d \right)_{32}^* = \frac{2\sqrt{2}}{v^3} m_b \left( m_t^2 V_{tb} V_{ts}^* \right)$$

$$\left( Z^d \right)_{23} = \frac{2\sqrt{2}}{v^3} m_s \left( m_t^2 V_{tb} V_{ts}^* \right)$$

MFV is nothing but the  
GIM mechanism extended  
to BSM contributions

$$\left[ b^c \left( \sigma F \right) s \right]^+ \text{ dominates over } s^c \left( \sigma F \right) b \text{ by } (m_t/m_b)$$

$$BR(B \rightarrow X_s \gamma) = (3.55 \pm 0.24 \pm 0.09) \times 10^{-4}$$



$$\Lambda > 6.1 \text{ TeV}$$

## Exercise 12: build the leading operator with $\Delta F=2$ in MFV

same basis as before:

$$y_d = \hat{y}_d \quad y_u = \hat{y}_u V_{CKM} \quad \hat{y}_{u,d} \text{ diagonal}$$

leading MFV invariant

$$\bar{q}_{Li} \gamma^\mu (y_u^\dagger y_u)_{ij} q_{Lj} \bar{q}_{Lk} \gamma_\mu (y_u^\dagger y_u)_{kl} q_{Ll}$$

looking at the down quark sector and selecting  $i=k=d,s$  and  $j=l=b$   
we get the MFV operator contributing to  $\Delta B=2$

$$O_{MFV}(|\Delta B|=2) = \frac{c}{\Lambda_{NP}^2} y_t^4 (V_{tb} V_{tq}^*)^2 \bar{q}_L \gamma^\mu b_L \bar{q}_L \gamma_\mu b_L \quad (q = d,s) \quad \text{where we used } \hat{m}_u \approx \text{diag}(0,0,m_t)$$

again same CKM suppression as in the SM. Now the bound on the scale of New Physics reads

$$\Lambda_{NP} > 5.9 \text{ TeV}$$

define 2 New Physics parameters

$$\Delta_q \equiv \frac{M_{12}^q}{M_{12}^{q,SM}} \quad (q=d,s)$$

[ $O_{MFV}$  modify  $M_{12}$  for  $B_d$  and  $B_s$  in the same way:  
i.e  $\Delta_d$  and  $\Delta_s$  are identical and real in MFV]

# bound on the scale of New Physics in MFV

Operator	Bound on $\Lambda$	Observables
$H^\dagger (\overline{D}_R Y^{d\dagger} Y^u Y^{u\dagger} \sigma_{\mu\nu} Q_L) (e F_{\mu\nu})$	6.1 TeV	$B \rightarrow X_s \gamma, B \rightarrow X_s \ell^+ \ell^-$
$\frac{1}{2} (\overline{Q}_L Y^u Y^{u\dagger} \gamma_\mu Q_L)^2$	5.9 TeV	$\epsilon_K, \Delta m_{B_d}, \Delta m_{B_s}$
$H_D^\dagger (\overline{D}_R Y^{d\dagger} Y^u Y^{u\dagger} \sigma_{\mu\nu} T^a Q_L) (g_s G_{\mu\nu}^a)$	3.4 TeV	$B \rightarrow X_s \gamma, B \rightarrow X_s \ell^+ \ell^-$
$(\overline{Q}_L Y^u Y^{u\dagger} \gamma_\mu Q_L) (\overline{E}_R \gamma_\mu E_R)$	2.7 TeV	$B \rightarrow X_s \ell^+ \ell^-, B_s \rightarrow \mu^+ \mu^-$
$i (\overline{Q}_L Y^u Y^{u\dagger} \gamma_\mu Q_L) H_U^\dagger D_\mu H_U$	2.3 TeV	$B \rightarrow X_s \ell^+ \ell^-, B_s \rightarrow \mu^+ \mu^-$
$(\overline{Q}_L Y^u Y^{u\dagger} \gamma_\mu Q_L) (\overline{L}_L \gamma_\mu L_L)$	1.7 TeV	$B \rightarrow X_s \ell^+ \ell^-, B_s \rightarrow \mu^+ \mu^-$
$(\overline{Q}_L Y^u Y^{u\dagger} \gamma_\mu Q_L) (e D_\mu F_{\mu\nu})$	1.5 TeV	$B \rightarrow X_s \ell^+ \ell^-$

TABLE II: Bounds on the scale of new physics (at 95% C.L.) for some representative  $\Delta F = 1$  [27] and  $\Delta F = 2$  [12] MFV operators (assuming effective coupling  $\pm 1/\Lambda^2$ ), and corresponding observables used to set the bounds.

[Isidori, Nir, Perez, 2010]

# Minimal Flavour Violation (leptons)

extension of MFV to leptons is ambiguous:

we can describe neutrino masses in several ways

- 1 B-L conserved, pure Dirac neutrino masses  
just copy the quark sector

$$G_l = SU(3)_{\nu^c} \times SU(3)_{e^c} \times SU(3)_l \times \dots$$

$$l = (1, 1, 3) \quad \nu^c = (\bar{3}, 1, 1) \quad e^c = (1, \bar{3}, 1)$$

$$y_\nu = (3, 1, \bar{3})$$

$$y_e = (1, 3, \bar{3})$$

$$i \frac{e}{\Lambda^2} e^c \left( \sigma^{\mu\nu} F_{\mu\nu} \right) \mathcal{Z} (\Phi^\dagger l) + h.c.$$

choose as basis:

$$y_e = \hat{y}_e \quad y_\nu = \hat{y}_\nu U_{PMNS}^+$$

$$\mathcal{Z} = y_e y_\nu^\dagger y_\nu$$

$$= \frac{2\sqrt{2}}{v^3} \left( \hat{m}_e U_{PMNS} \hat{m}_\nu^2 U_{PMNS}^+ \right)$$

dominant contribution to  $\mu \rightarrow e \gamma$

$$\begin{aligned} \left( \mathcal{Z} \right)_{21}^* &= \frac{2\sqrt{2}}{v^3} m_\mu \left( U_{\mu i}^* U_{ei} m_i^2 \right) \\ &\approx 10^{-28} \end{aligned}$$

$\mu \rightarrow e \gamma$  unobservable  
even for  $\Lambda \approx 1 \text{ TeV}$

## 2 B-L violated, neutrino masses from d=5 operator

[Cirigliano, Grinstein, Isidori, Wise 2005]

$$L = \dots + e^c y_e (\Phi^+ l) + \frac{1}{2\Lambda_L} (\tilde{\Phi}^+ l)_w (\tilde{\Phi}^+ l) + h.c.$$

an important assumption:  $\Lambda_L \neq \Lambda$

$$G_l = SU(3)_{e^c} \times SU(3)_l \times \dots$$

$$l = (1, 3) \quad e^c = (\bar{3}, 1)$$

$$y_e = (3, \bar{3})$$

$$w = (1, \bar{6})$$

the only sources of  $G_l$  breaking

$$y_e = \sqrt{2} \frac{m_e^{diag}}{v}$$

$$w = \frac{2\Lambda_L}{v^2} U^* m_v^{diag} U^+$$

spurions expressed in terms of known quantities and  $\Lambda_L$

$$\mathcal{Z} = y_e w^+ w$$

$$= \frac{4\sqrt{2}}{v^3} \frac{\Lambda_L^2}{v^2} \left( \hat{m}_e U_{PMNS} \hat{m}_v^2 U_{PMNS}^+ \right)$$

$\mu \rightarrow e \gamma$  dominated by

$$\left( \mathcal{Z} \right)_{21}^* = \frac{4\sqrt{2}}{v^3} \frac{\Lambda_L^2}{v^2} m_\mu \left( U_{\mu i}^* U_{ei} m_i^2 \right)$$

enhancement factor  
can be huge

$$\frac{\Lambda_L^2}{v^2}$$

experimental bound satisfied  
by  $(\Lambda_L/\Lambda) < 10^9$

$\mu \rightarrow e \gamma$  observable if  $\Lambda_L \gg \Lambda$

[qualitatively similar conclusion when MFV extended to the type I see-saw case]

# Exercise 13: show that

$$Z_{ij} = \frac{4\sqrt{2}}{v^3} \frac{\Lambda_L^2}{v^4} \left[ \Delta m_{sol}^2 U_{i2} U_{j2}^* \pm \Delta m_{atm}^2 U_{i3} U_{j3}^* \right]$$

+ for normal hierarchy  
- for inverted hierarchy

and estimate

$$\frac{R_{\mu e}}{R_{\tau\mu}} = \frac{BR(\mu \rightarrow e\gamma)}{BR(\tau \rightarrow \mu\gamma)} \times \frac{BR(\tau \rightarrow \mu\nu_\tau \bar{\nu}_\mu)}{BR(\mu \rightarrow e\nu_\mu \bar{\nu}_e)}$$

solution

$$\frac{R_{\mu e}}{R_{\tau\mu}} \approx \left| \frac{2}{3} r \pm \sqrt{2} \sin \vartheta_{13} e^{i\delta} \right|^2 \approx (0.035 \div 0.055)$$

$$r \equiv \frac{\Delta m_{sol}^2}{\Delta m_{atm}^2}$$

from present bound  
on  $\mu \rightarrow e \gamma$

$$R_{\tau\mu} < (1.0 \div 1.6) \times 10^{-11}$$

hints:

-- use unitarity relation for  $U_{PMNS}$

-- use approximate values

$$U_{\mu 3} \approx U_{\tau 3} \approx 1/\sqrt{2}$$

$$U_{e2} \approx U_{\mu 2} \approx -U_{\tau 2} \approx 1/\sqrt{3}$$



# LFV in the limit of vanishing neutrino masses

MFV extended to the lepton sector reproduces the GIM suppression in particular LF is conserved when  $m_i=0$

GIM suppression can be evaded in several models of fermion masses e.g. in partial compositeness where elementary fermions acquire a mass through their mixing with a composite sector

a toy model

$$L_Y = -e^c \Delta_E E - L^c \Delta_L l$$

$$- E^c M E - L^c M L$$

$$- E^c Y (\Phi^+ L) - (L^c \tilde{\Phi}^+) \tilde{Y} E + h.c.$$

$\Leftrightarrow$  elementary-composite mixing

$\Leftrightarrow$  Dirac masses for composite fermions

$\Leftrightarrow$  Yukawa coupling of composite fermions

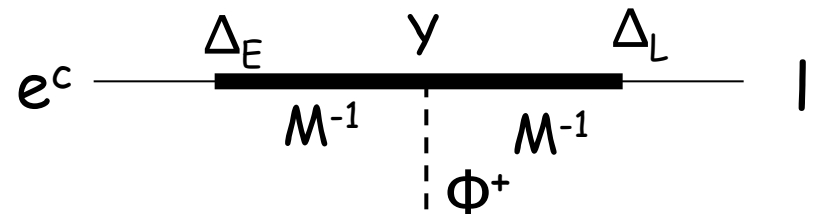
by integrating out the composite sector

[Exercise 14]

$$L_Y = -e^c y_e (\Phi^+ l) + h.c.$$

$$y_e = (\Delta_E M^{-1}) Y (M^{-1} \Delta_L) + \dots$$

higher-orders in  $(\Phi/M)$



## Exercise 15

compute the corrections to previous LO relations by using the equation of motion for the composite sector. Start with 1 generation and then discuss the 3 generation case.

write  $L_Y$  in matrix notation

$$L_Y = - \begin{pmatrix} e^c & E^c & L^c \end{pmatrix} \begin{pmatrix} 0 & \Delta_E & 0 \\ 0 & M & Y\Phi^+ \\ \Delta_L & \tilde{\Phi}^+ \tilde{Y} & M \end{pmatrix} \begin{pmatrix} l \\ E \\ L \end{pmatrix} + h.c.$$

write the e.o.m. for the composite fields  $(E^c, L^c)$  and  $(E, L)$  in the limit of negligible kinetic term and substitute them back into  $L_Y$

$$L_Y = e^c \begin{pmatrix} \Delta_E & 0 \end{pmatrix} \begin{pmatrix} M & Y\Phi^+ \\ \tilde{\Phi}^+ \tilde{Y} & M \end{pmatrix}^{-1} \begin{pmatrix} 0 \\ \Delta_L \end{pmatrix} l + h.c.$$

expand this expression in powers of the Higgs field

At the LO

$$y_e = F_{E^c} Y F_L$$

$$F_{E^c} = \Delta_E M^{-1}$$

$$F_L = M^{-1} \Delta_L$$

an intriguing possibility (anarchic scenario):

-- Yukawa coupling  $Y$  in the composite sector are  $O(1)$

-- fermion mass hierarchy entirely due to the amount of mixing  $F$

it arises in many SM extensions

## split fermions in an Extra Dimension

$$F_{X_i} = \sqrt{\frac{2\mu_i}{1 - e^{-2\mu_i r}}}$$

ED	$\mu_i$	$r$
Flat $[0, \pi R]$	$M_i / \Lambda$	$\Lambda \pi R$
Warped $[R, R']$	$1/2 - M_i R$	$\log R' / R$

no symmetry:  
hierarchy produced by geometry

$M_i$  = bulk mass of fermion  $X_i$

$Y_{u,d} = O(1)$  Yukawa couplings between bulk fermions and a Higgs localized at one brane

## fermion masses from abelian flavour symmetries $Q(X_i) \geq 0$

$$F_{X_i} = \text{diag}(\lambda^{Q(X_1)}, \lambda^{Q(X_2)}, \lambda^{Q(X_3)}) \quad \lambda = \frac{\langle \varphi \rangle}{\Lambda}$$

chiral multiplets  $X_i$  of the MSSM coupled to a superconformal sector

[Nelson-Strassler 0006251]

$$F_{X_i} = \left( \frac{\Lambda_c}{\Lambda} \right)^{\frac{\gamma_i}{2}} < 1$$

$\gamma_i$  anomalous dimension of  $X_i$



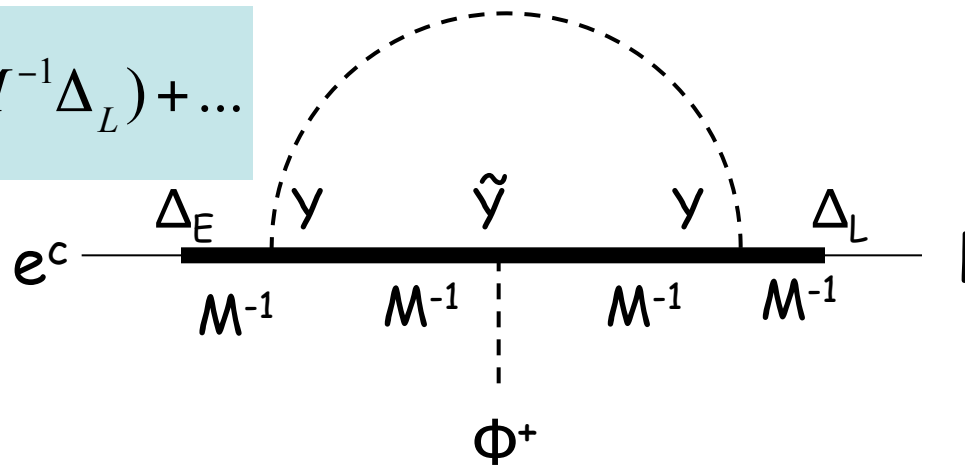
so far **neutrino are massless**  
do we expect LFV in our toy model?

one-loop contribution to lepton dipole operator from Higgs exchange  
(assuming  $M$  proportional to identity)

$$\frac{Z}{\Lambda^2} \approx \frac{1}{16\pi^2 M^2} (\Delta_E M^{-1}) Y \tilde{Y} Y (M^{-1} \Delta_L) + \dots$$

in general these combinations  
not diagonal in the same basis

$$y_e = (\Delta_E M^{-1}) Y (M^{-1} \Delta_L) + \dots$$



LFV not suppressed by neutrino masses and unrelated to (B-L) breaking scale

rough estimate

$$\frac{Z_{\mu e}}{\Lambda^2} < 2 \times 10^{-9} \text{ TeV}^{-2}$$



$$M > 10 \text{ TeV}$$

$$\begin{aligned} \Delta_E &\approx \Delta_L \\ \frac{\Delta_f}{M} &\approx \sqrt{\frac{m_f}{v}} \\ Y &\approx \tilde{Y} \approx O(1) \end{aligned}$$

## Exercise 16: reproduce flavour pattern of $Z$ from a spurion analysis

$$\frac{Z}{\Lambda^2} \approx \frac{1}{16\pi^2 M^2} (\Delta_E M^{-1}) Y \tilde{Y} Y (M^{-1} \Delta_L) + \dots$$

- identify the maximal flavour symmetry  $G$  of our toy model
- identify the transformation properties of the spurions  $\Delta_L, \Delta_E, Y, \tilde{Y}$ , that guarantee the invariance of  $L_Y$
- using previous tools, build the relevant dipole operator invariant under  $G$

## summary

LFV expected in charged leptons = CLFV

CLFV probes physics **beyond** the  $\nu$ SM [=SM minimally extended to accommodate  $\nu$  masses]

observable rates for CLFV require **new physics** at a scale well below the GUT or the L-violation scales  
[ $\Lambda \ll \Lambda_L$  in our example of MFV]

GIM suppression in CLFV is a special feature of MFV:  
it can be violated in models of fermion masses  
and relation to neutrino masses and mixing angles can be more indirect