

BSM Lecture 4

Semi-classical limit $\hbar = 1$

$$M \rightarrow \infty \quad G_N \rightarrow 0 \quad R = \text{fixed}$$

$$S_{\text{BH}} = \left(\frac{M}{M_p}\right)^2 \rightarrow \infty \quad \text{fixed (rigid geometry)}$$

$(M_p \rightarrow \infty)$ \nearrow

This is also a limit of zero back-reaction

In this limit certain properties are understood exactly.

E.g. thermality of Hawking radiation

Therefore for finite $S_{\text{BH}} \gg 1$

the back-reaction (at least for initial time) must be small.

Correspondingly, the correction to maximised I in semi-classical

(which are exact) are expected to be small, as long as we are in the regime of weakly interacting Einstein gravity.

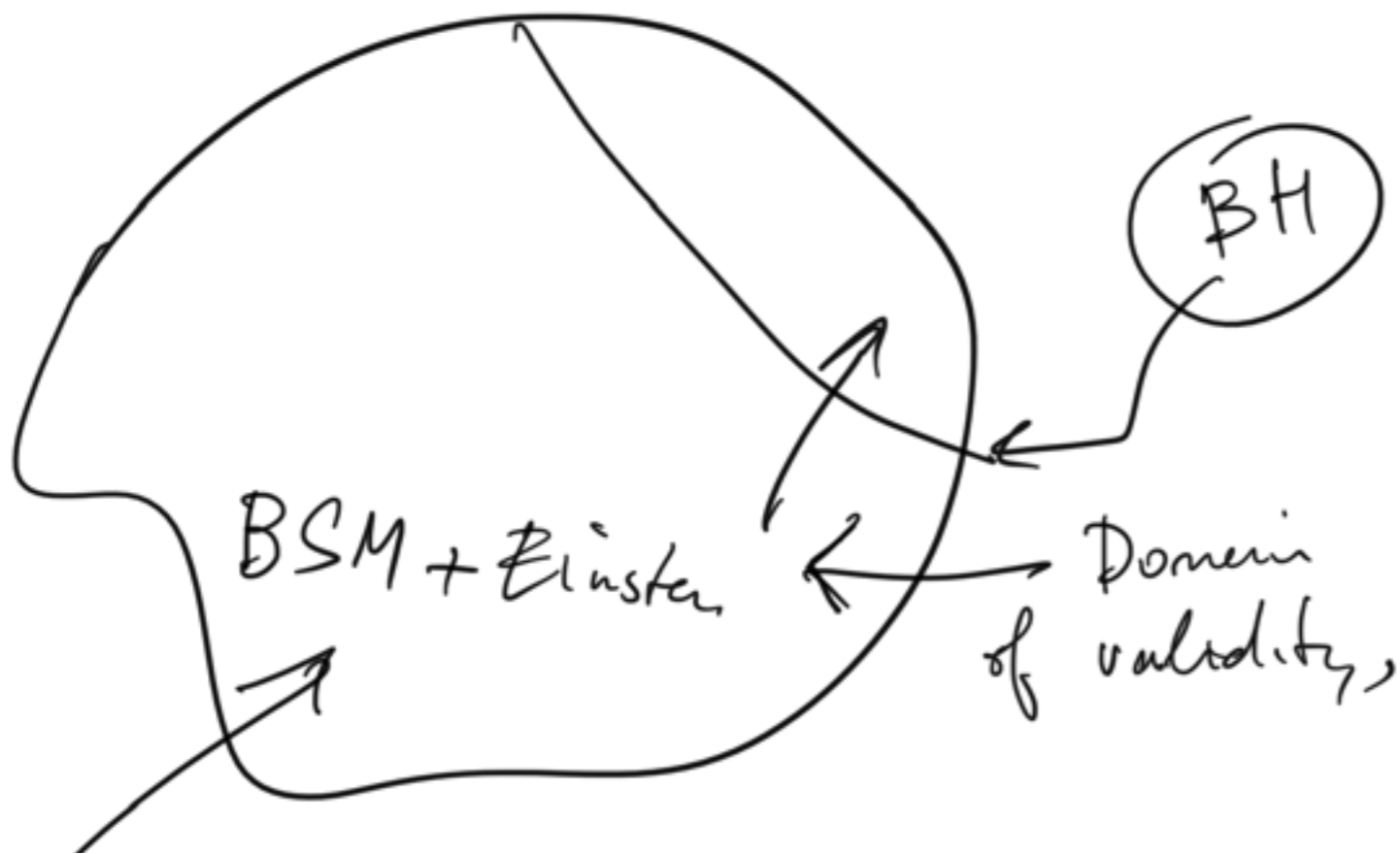
$$R = \text{finite} \quad S_{\text{BH}} = \infty.$$



$$\delta \sim \frac{1}{S_{\text{BH}}}$$

$l \rightarrow R'$

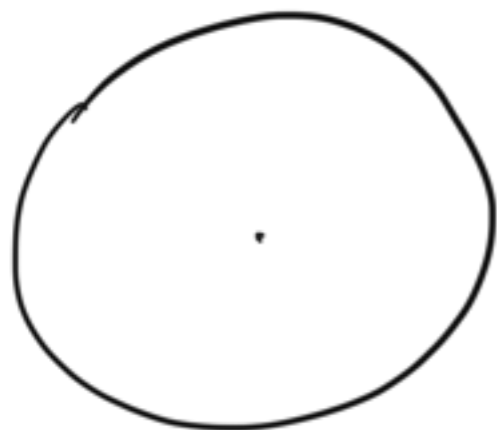
as long as $d_{\text{gr}}\left(\frac{l}{R}\right) \ll 1$



Example: Bound on the number of particle species N_{sp}

L $m \lesssim \frac{1}{L}$

active in processes with momentum-transfer $q \gtrsim \frac{1}{L}$



$R \gg L_p$
 $S_{BH} \gg 1$ $S_{BH} \sim \frac{R^2}{L_p^2}$

$k \sim R$

$T = \frac{1}{R} = \frac{M_p^2}{M}$

$\dot{M} = -T^2 N_{sp}$

democratic in species

$M \sim M_p^2$

$\dot{M} \sim M_p^2 \left(\frac{\dot{T}}{T} \right)$

$$\frac{\dot{T}}{T^2} = N_{sp} \cdot \frac{T^2}{M_p^2}$$

$\frac{\dot{T}}{T^2}$ = parameter of thermality of semi-classicality.

$$\frac{\dot{T}}{T^2} \ll 1$$

we can speak about T reliably but the notion of thermality breaks down for

$$\frac{\dot{T}}{T^2} \sim 1$$

Thus, weak coupling Einstein description breaks down.

$$\frac{\dot{T}_*}{T_*^2} = N_{sp} \frac{T_*^2}{M_p^2} \sim 1$$

$$T_*^2 \quad M_p^2$$

Cutoff temperature.

$$T_* = \frac{M_p}{\sqrt{N_{sp}}} \equiv M_* \equiv \bar{L}_*$$

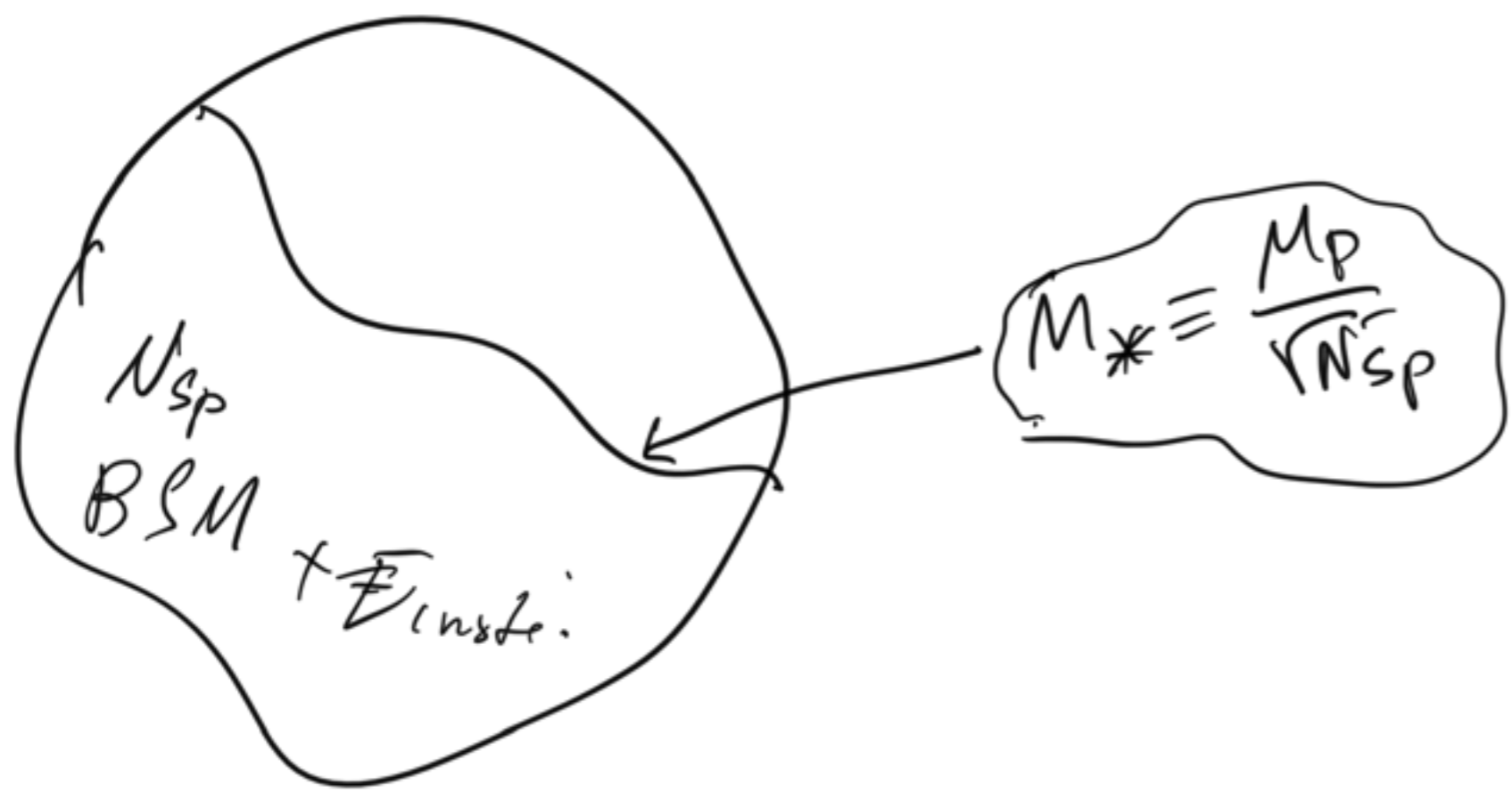
cutoff of weak Einstein gravity



R_*

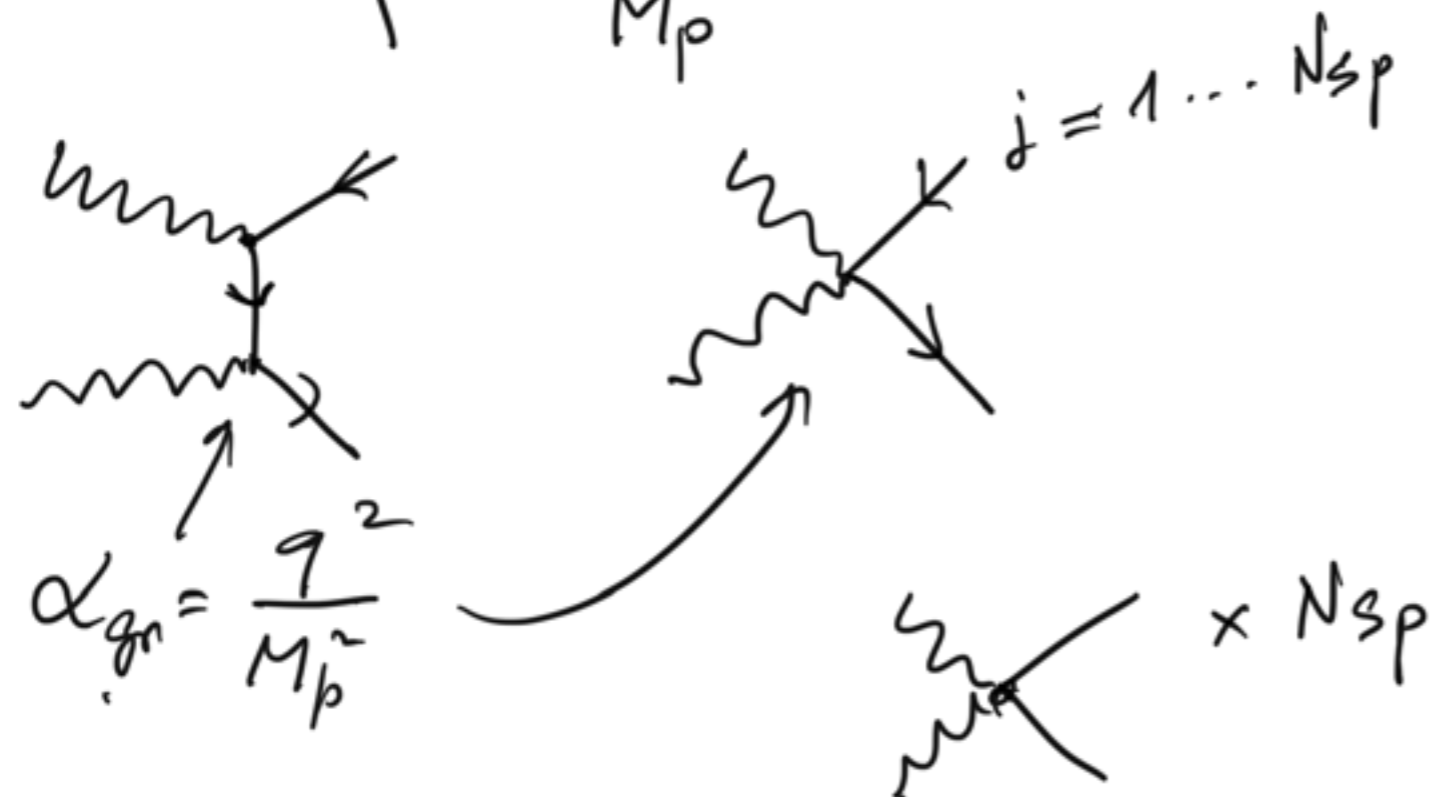
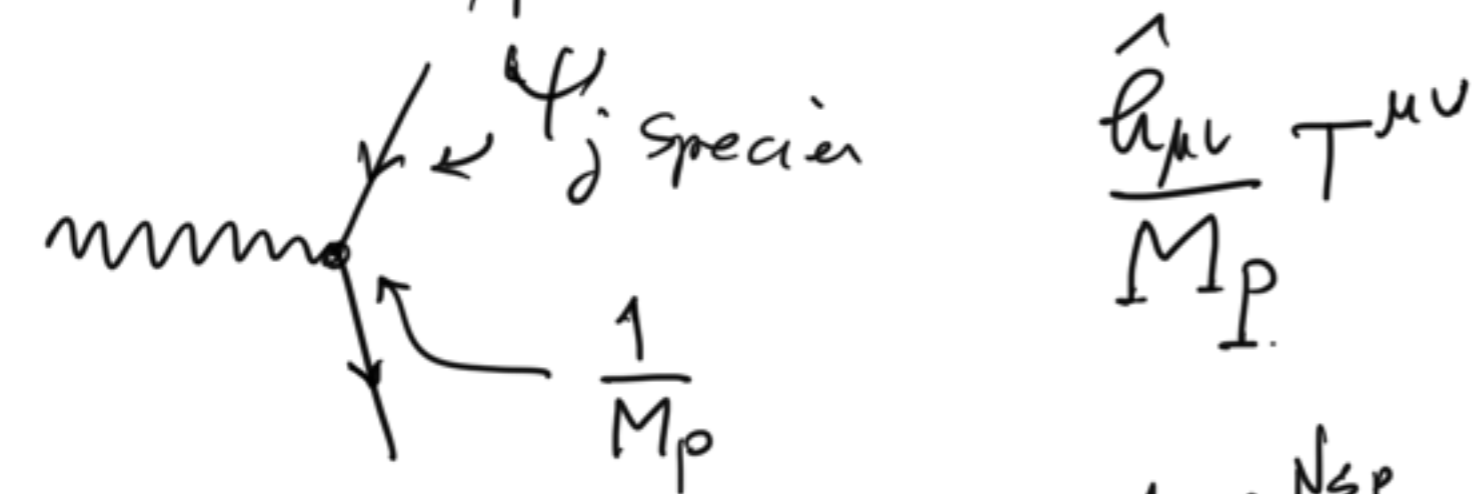
$$R_* = \bar{M}_*^{-1} \gg L_p$$

Collective effect is too strong.



$$M_* \equiv \frac{M_p}{\sqrt{N_{sp}}}$$

Let us explain collective effect.

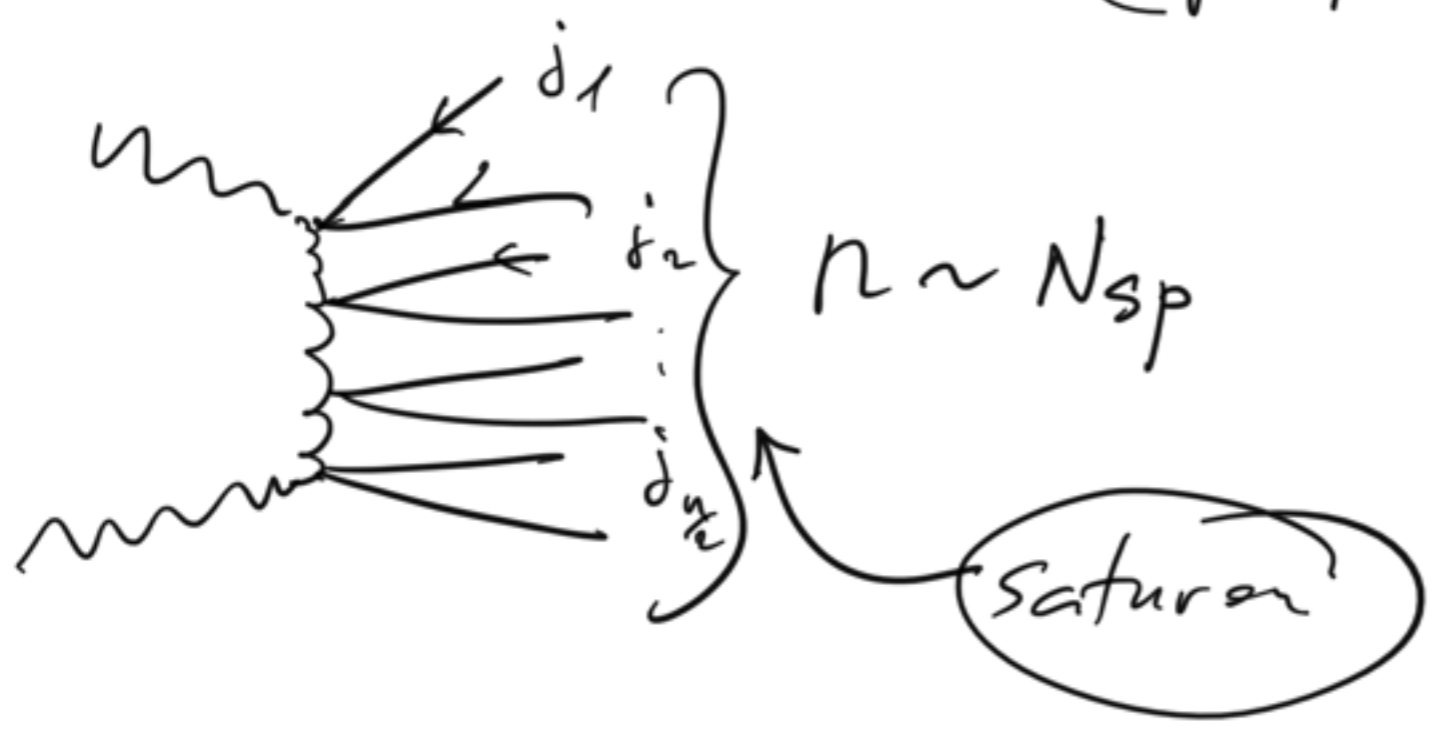


Species coupling λ (gravitational + Higgs coupling)

$$\lambda \equiv \alpha_{gr} N_{sp}$$



$$d_{gr} N_{sp} \quad \lambda^e \equiv (d_{gr} N_{sp})^e$$



$$R_{BH} \sim M_*^{-1} + \boxed{q < M_*}$$

Hierarchical problem puzzle

Naturalness "problem" "puzzle"

$$G_H = \frac{1}{v^2} \gg G_N = \frac{1}{M_P^2} ?$$

UV-sensitivity of the Higgs mass (VEV) to the cutoff of the theory.

$$\mathcal{L} = -\underline{m_H^2} H^\dagger H - \lambda^2 (H^\dagger H)^2 + g^\psi H \psi \psi + \text{gauge} \dots$$

$$\delta m_H^2 = \text{---} H \text{---} \bigcirc \text{---} H \text{---} +$$

$$+ \text{---} H \text{---} \bigcirc \text{---} H \text{---} + \dots$$

$$\delta M_H^2 \propto M_*^2$$

$$M_P^2$$

$$N_{sp} \sim 1$$

$$\frac{M_H}{M_p} \sim 10^{-34}$$

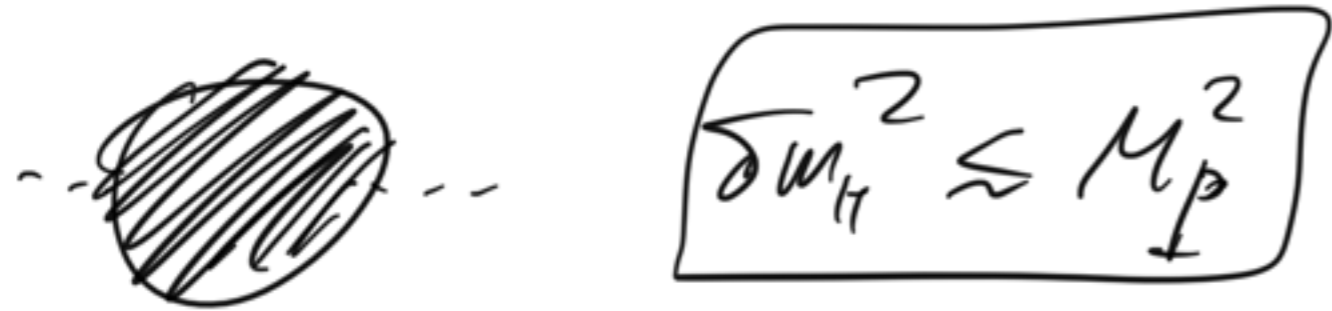
$$m_H^2$$

$$M_{sin}^2$$

$$m_H^2?$$

$$M_P^2$$

$$m_H^2$$

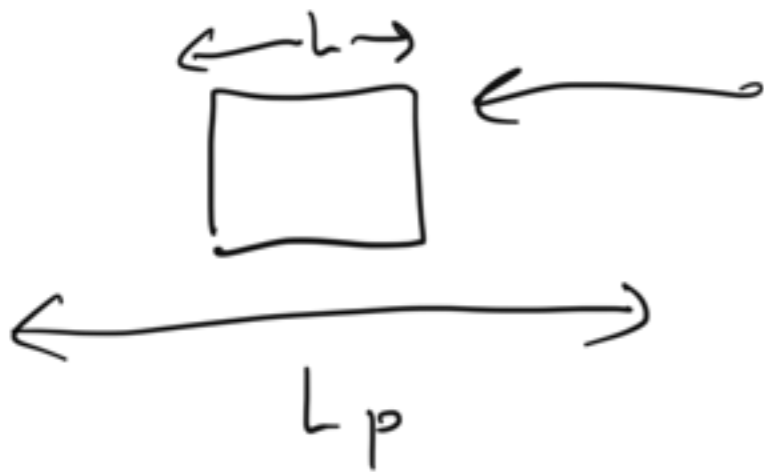


$$q \gg \frac{1}{L}$$

$$E \gg \frac{1}{L}$$

$$L \ll L_p$$

$$E \gg \frac{1}{L}$$



$$\lambda \sim E L_p^2 \gg L_p$$

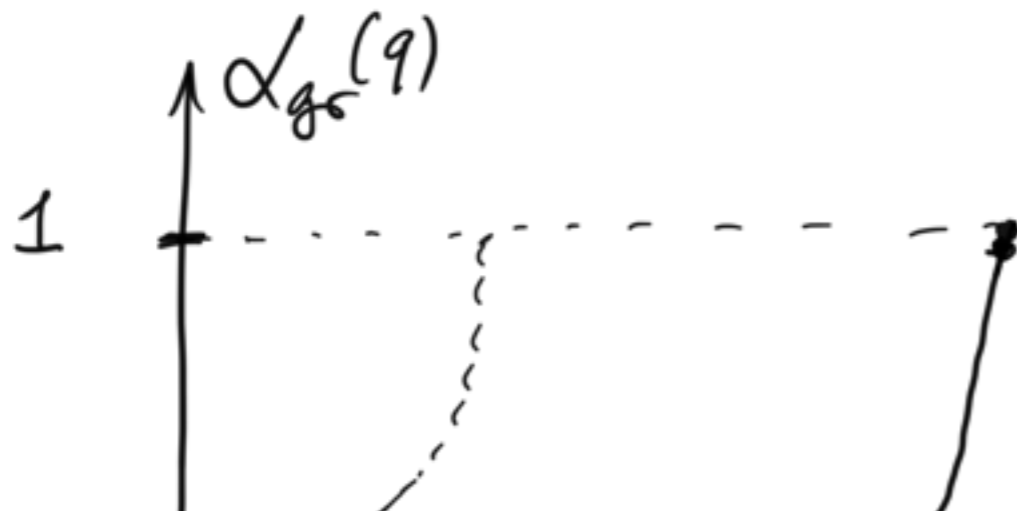
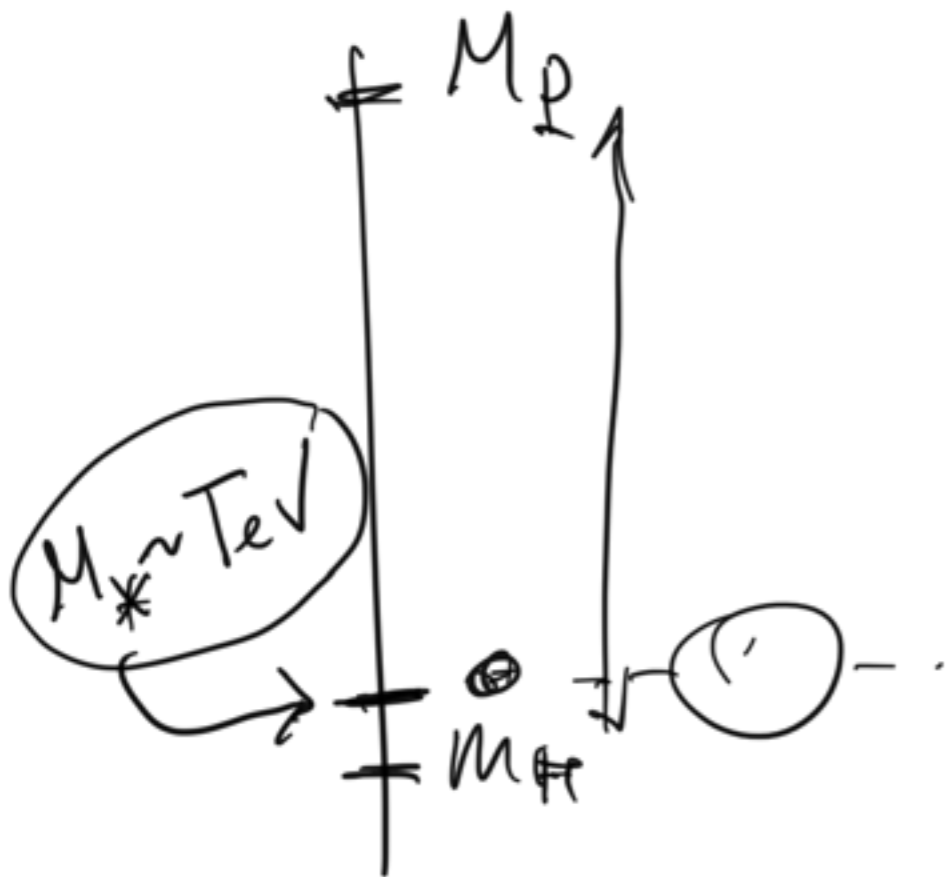
$$m \bar{\Psi}_L \Psi_R$$

$$m \Psi \Psi$$

$$\Psi \rightarrow e^{i\alpha \gamma_5} \Psi$$

$$\delta m \propto m$$

$$M_* = \frac{M_P}{\sqrt{N_{sp}}}$$





$$T^{\mu\nu} \quad \hat{h}_{\mu\nu} \quad T^{\mu\nu}$$

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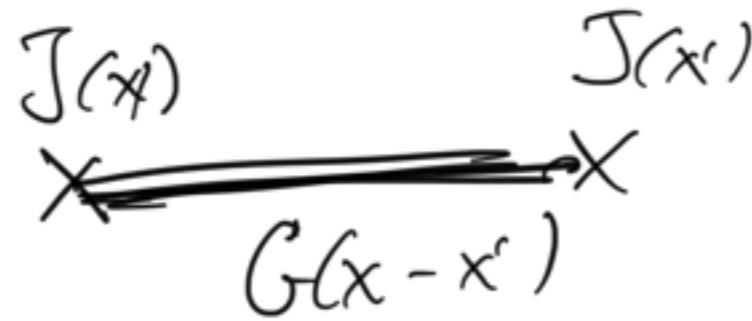
$$S = \int d^4x d^4x' T^{\mu\nu}(x) \underbrace{H_{\mu\nu}(x) H_{\alpha\beta}(x')}_{\text{Köln-Lehman}} T^{\alpha\beta}(x')$$

$$G_{\mu\nu, \alpha\beta}(x-x')$$

Köln-Lehman

$$G(p) = \int_0^{\infty} dm^2 \frac{\rho(m^2)}{p^2 - m^2 + i\epsilon}$$

$$J(x) \overset{\wedge}{\Phi}(x)$$



$$G_{\mu\nu, \alpha\beta}(p) = \int_0^{\infty} dm^2 \frac{\rho(m)_{\mu\nu, \alpha\beta}}{p^2 - m^2 + i\epsilon}$$

$$\mathcal{L}(m=0) = \left( \frac{1}{2} (\eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\mu\beta} \eta_{\nu\alpha}) - \frac{1}{2} \eta_{\mu\nu} \eta_{\alpha\beta} \right)$$

$$\mathcal{L}(m \neq 0) = ?$$

$$\text{Spin} = 2 \quad m \neq 0$$

$$\mathcal{L} = \frac{1}{2} h^{\mu\nu} \epsilon_{\mu\nu} + m^2 (h_{\mu\nu} h^{\mu\nu} - h^2)$$

not free

$$g^{\mu\nu} \rightarrow h^{\mu\nu} + \partial^\mu \xi^\nu + \partial^\nu \xi^\mu$$

Fierz-Pauli

$$\begin{aligned} & \square h_{\mu\nu} - \eta_{\mu\nu} \square h - \partial_\mu \partial_\nu h - \partial_\nu \partial_\mu h \\ & + \partial_\mu \partial_\nu h + \eta_{\mu\nu} \partial^\alpha \partial^\beta h_{\alpha\beta} \\ & + m^2 (h_{\mu\nu} - \eta_{\mu\nu} h) = T_{\mu\nu} \end{aligned}$$

$$\partial^\mu (\dots) = T_{\mu\nu}$$

$10 - 4 = 6$

$\partial^\mu h_{\mu\nu} = \partial_\nu h$

← constraint.

$$\begin{aligned} & \square h_{\mu\nu} - \cancel{\eta_{\mu\nu} \square h} - \cancel{\partial_\mu \partial_\nu h} - \cancel{\partial_\nu \partial_\mu h} \\ & + \cancel{\partial_\mu \partial_\nu h} + \cancel{\eta_{\mu\nu} \square h} + \\ & + m^2 (h_{\mu\nu} - \eta_{\mu\nu} h) = T_{\mu\nu} \end{aligned}$$

$$\square h - \square h - 3m^2 h = T$$

$$h = \frac{-1}{3m^2} T$$

$$6 - 1 = 5$$

$$\square h_{\mu\nu} - \partial_\mu \partial_\nu h + m^2 (h_{\mu\nu} - \eta_{\mu\nu} h) = T_{\mu\nu}$$

$$\underline{(\square + m^2)} h_{\mu\nu} = T_{\mu\nu} - \frac{1}{3} \left( \eta_{\mu\nu} + \frac{\partial_\mu \partial_\nu}{m^2} \right) T$$

$$G_{\mu\nu} = \frac{1}{\Box + m^2} \left( T_{\mu\nu} - \frac{1}{3} \left( \eta_{\mu\nu} + \frac{\partial_\mu \partial_\nu}{m^2} \right) T \right)$$

$$T^{\mu\nu} G_{\mu\nu} = \frac{T^{\mu\nu}}{\Box + m^2} \left( T_{\mu\nu} - \frac{1}{3} \eta_{\mu\nu} T \right)$$

$$\Downarrow$$

$$T^{\mu\nu} G_{\mu\nu, \alpha\beta}^{(m)} T^{\alpha\beta} =$$

$$m \rightarrow 0 \rightarrow \frac{T^{\mu\nu}}{\Box} \left( T_{\mu\nu} - \frac{1}{3} \eta_{\mu\nu} T \right)$$

$$m = 0 \rightarrow \frac{T^{\mu\nu}}{\Box} \left( T_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} T \right)$$

extra polarizati

$$G_{\mu\nu, \alpha\beta}^{(p)} = \frac{1}{M_p^2} \frac{\frac{1}{2} (\eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\nu\alpha} \eta_{\mu\beta}) - \frac{1}{2} \eta_{\mu\nu} \eta_{\alpha\beta}}{p^2 + i\epsilon}$$

$$+ \frac{1}{M_p^2} \sum_{m^2 < 0} C_2^{(m)} \frac{\frac{1}{2} (\eta_{\mu\alpha} \eta_{\nu\beta} + \eta_{\nu\alpha} \eta_{\mu\beta}) - \frac{1}{3} \eta_{\mu\nu} \eta_{\alpha\beta}}{p^2 - m^2 + i\epsilon}$$

$$+ \frac{1}{M_p^2} \sum_{m^2 \geq 0} C_0^{(m)} \frac{\eta_{\mu\nu} \eta_{\alpha\beta}}{p^2 - m^2 + i\epsilon}$$

$$T_{\mu\nu} = \delta_{\mu}^0 \delta_{\nu}^0 M \delta^3(\vec{x})$$

$$T_{\mu\nu} = \delta_{\mu}^0 \delta_{\nu}^0 \mu \delta^3(\vec{x} - \vec{r})$$

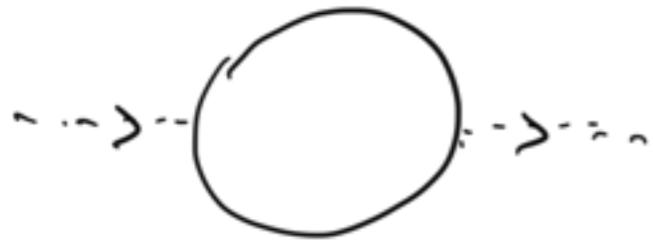
~~$\sum_{\mu} \dots$~~

$$V(r) = -\frac{M M \dot{G}_N}{r} - \sum_{m \neq 0} G_2^{(m)} \frac{e^{-mr}}{r}$$

$$- \frac{M m G_0^{(0)}}{r} - \sum_{m \neq 0} G_0^{(m)} \frac{e^{-mr}}{r}$$

$$m \approx r^{-1}$$





$$H \rightarrow e^{i\alpha} H$$

$$m_H^2 H^\dagger H$$

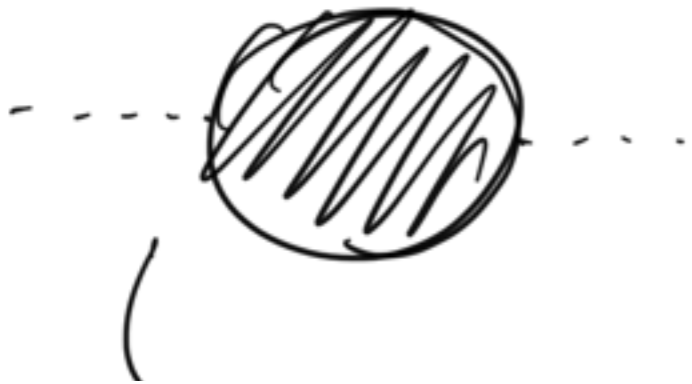
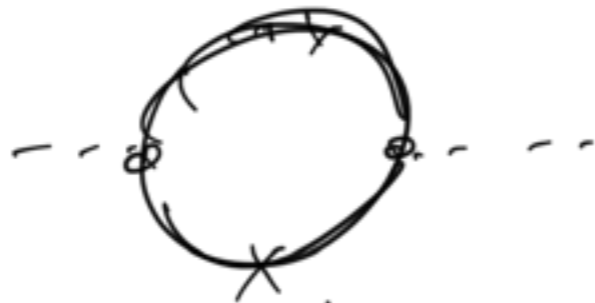
$$\mathcal{L} = \partial_\mu H^\dagger \partial^\mu H - \cancel{m_H^2 H^\dagger H} + \delta m_H^2 H^\dagger H$$

$$H \rightarrow H + \text{const}$$

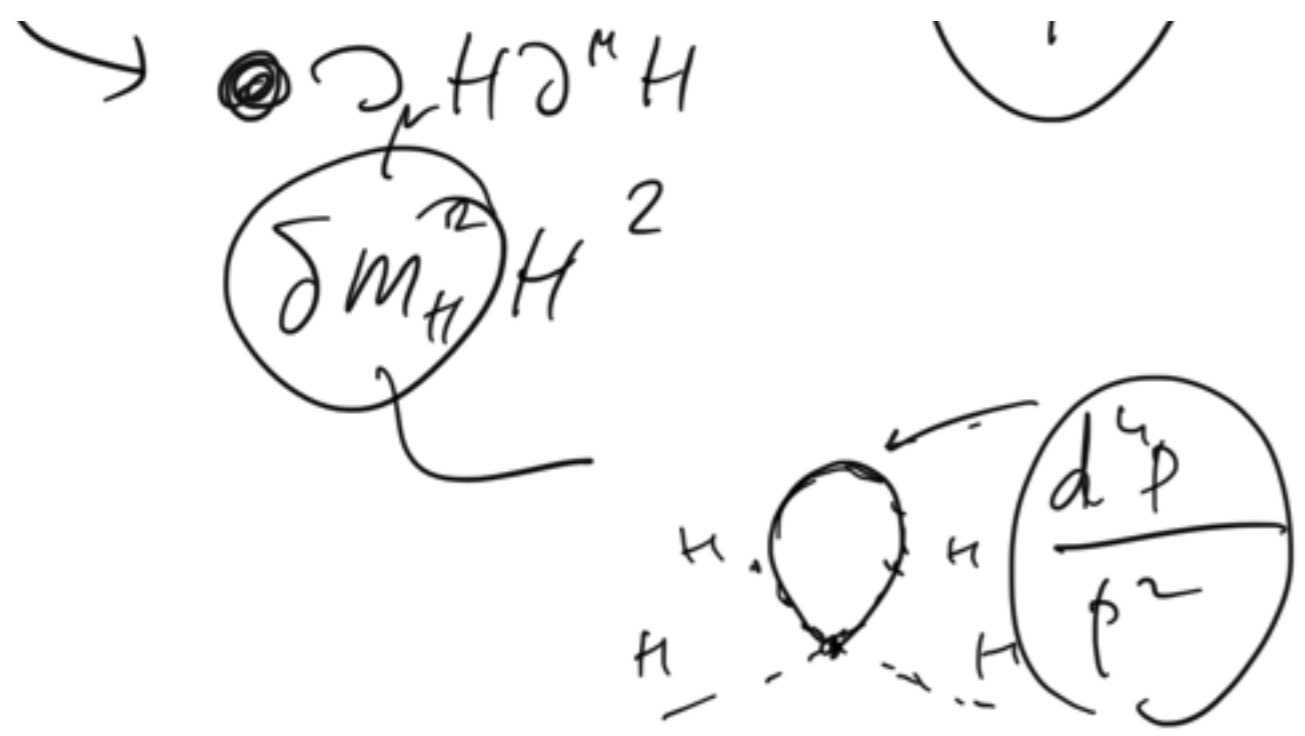
$$-\lambda^2 (H^\dagger H)^2$$

$$\delta m_H^2 \sim M_*^2 \left( \frac{\lambda^2}{\Lambda^2} + \dots \right) + g^4 H^\dagger H$$

$$\mathcal{L} = \partial_\mu H^\dagger \partial^\mu H + \left( \partial_\mu H^\dagger \partial^\mu H \right)^2 + \dots$$

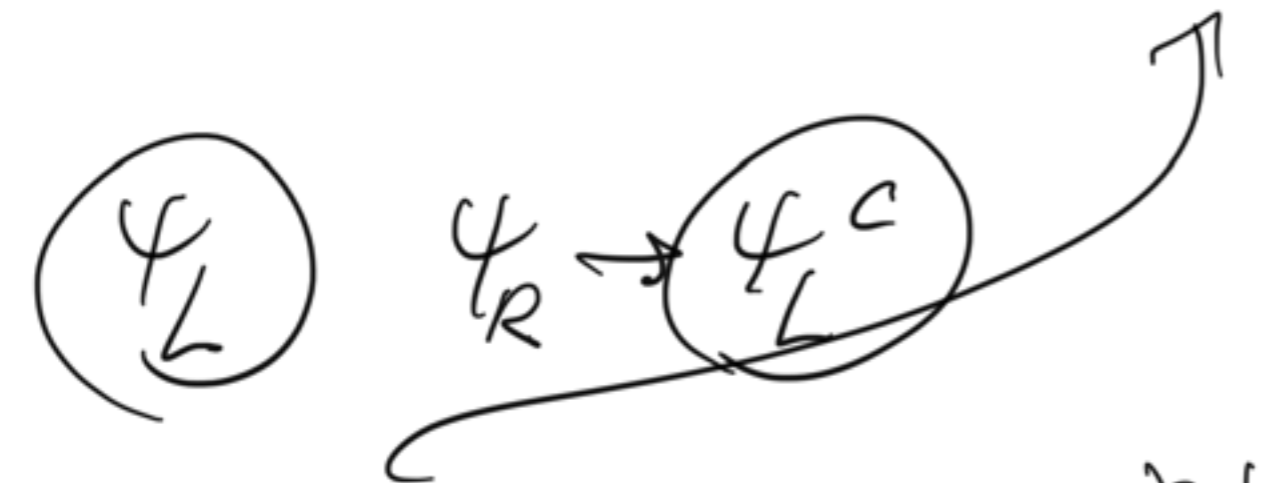


$$\frac{d^4 p}{p^2}$$



$$M_p \bar{\Psi} \Psi \rightarrow M_e \bar{\Psi}_L \Psi_R + h.c$$

$$\mathcal{L} = i \bar{\Psi} \not{\partial} \Psi = i \bar{\Psi}_L \not{\partial} \Psi_L + i \bar{\Psi}_R \not{\partial} \Psi_R$$



chiral symmetry,  $\delta_5 \Psi_L = \Psi_L$   
 $\delta_5 \Psi_R = -\Psi_R$   
 $\Psi \rightarrow e^{i\alpha \gamma_5} \Psi$

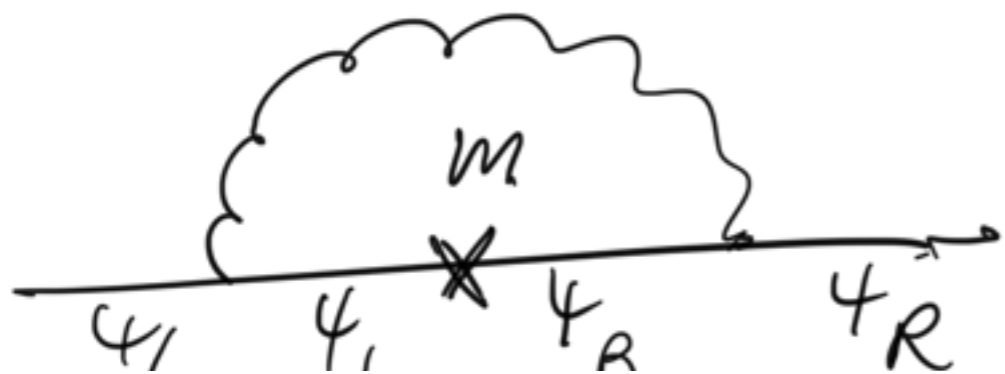
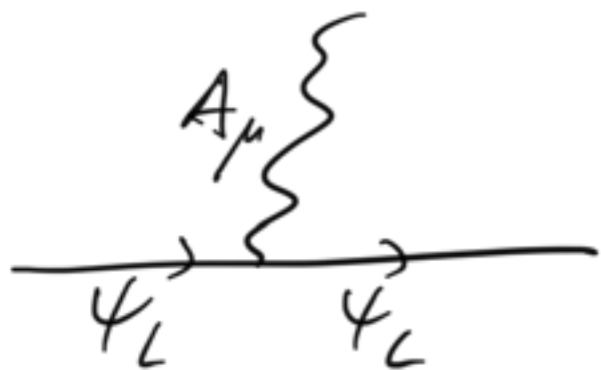
$$\begin{aligned}
 \Psi_L &\rightarrow e^{i\alpha} \Psi_L \\
 \Psi_R &\rightarrow e^{-i\alpha} \Psi_R
 \end{aligned}$$

$$\mathcal{L} = i\bar{\Psi}\not{\partial}\Psi - m\bar{\Psi}\Psi$$

$$m\bar{\Psi}_L\Psi_R + \text{h.c.}$$

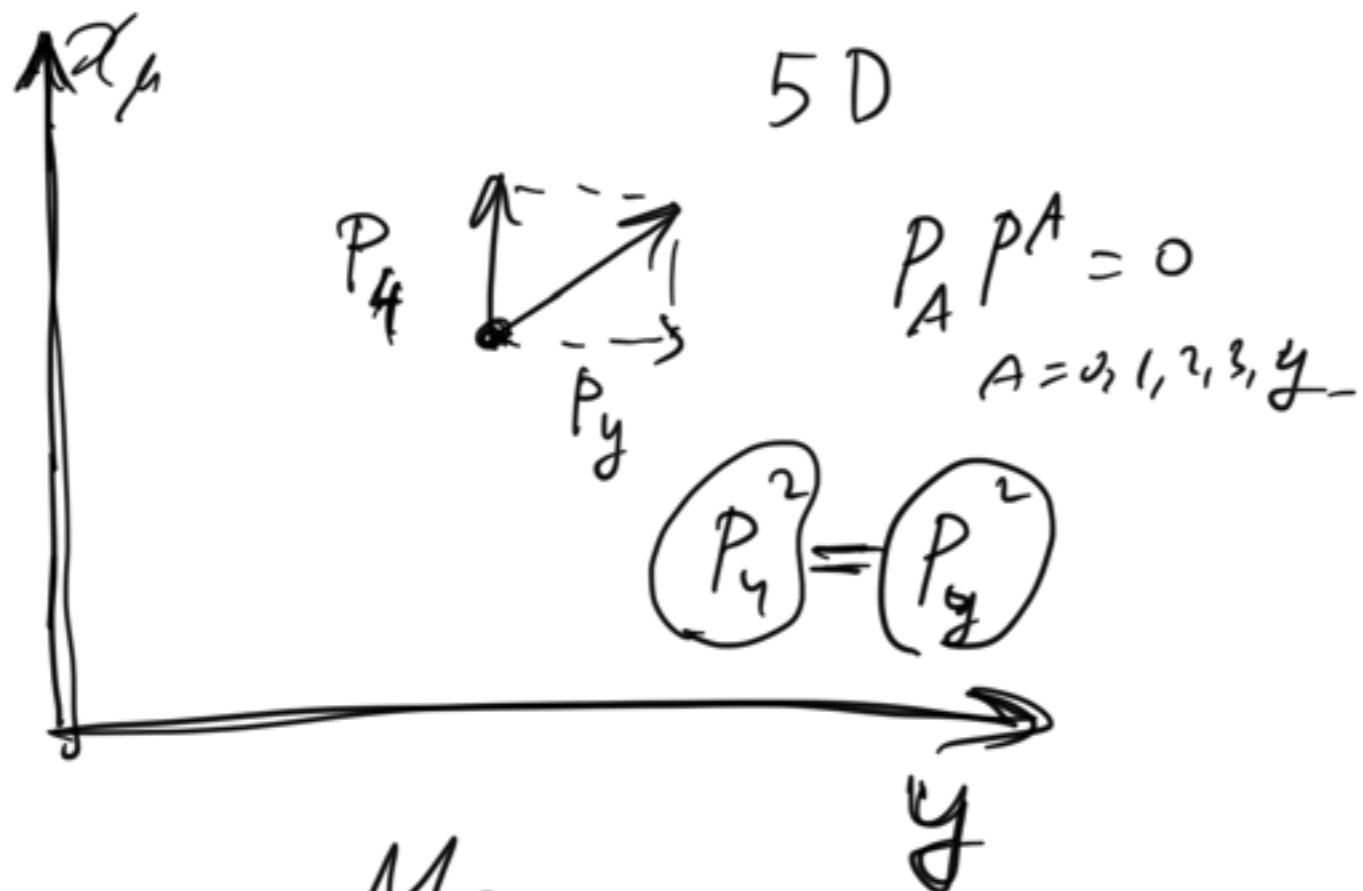
$$i\bar{\Psi}\not{\partial}\Psi \rightarrow$$

$$\bar{\Psi}_L\not{\partial}^M\Psi_L A^M + \bar{\Psi}_R\not{\partial}^M\Psi_R A^M$$





$$\delta m_e \propto m_e$$



$$M_* = \frac{M_p}{\sqrt{N_{sp}}} \leftarrow KK$$

$$m < M_*$$

$$\text{[Symbol]} = \sum$$

$\rightarrow (M_S)$



~~h<sub>0</sub>~~

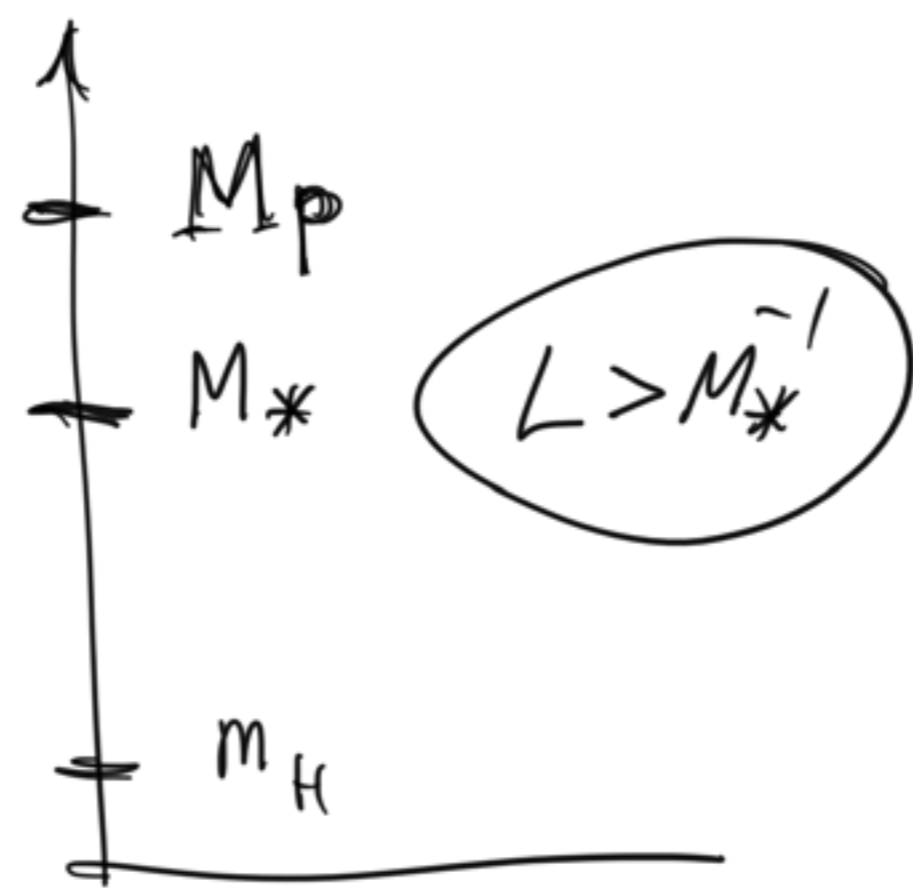
$h_{\mu\nu}$



$$G_{\mu\nu, p} = \int_0^{\infty} \frac{dm^2 \rho(m^2) \gamma_{\mu\nu} d\beta}{p^2 - m^2 + i\epsilon}$$

7/9

$\uparrow$   
 $m^2$



Spin = 2      Spin = 1      2

$$h_{\mu\nu}^{FP} = \underbrace{h_{\mu\nu}^E}_{\text{Spin=2}} + \underbrace{2 \partial_\mu A_\nu + \partial_\nu A_\mu}_{\text{Spin=1}} + \underbrace{\frac{1}{2} \eta_{\mu\nu} \pi + \frac{1}{2} \partial_\mu \partial_\nu \bar{\pi}}_{\text{Spin=0}}$$

The equation above is annotated with circles and arrows. A circle around  $h_{\mu\nu}^E$  has an arrow pointing to "Spin=2". A circle around  $2 \partial_\mu A_\nu + \partial_\nu A_\mu$  has an arrow pointing to "Spin=1". A circle around  $\frac{1}{2} \eta_{\mu\nu} \pi + \frac{1}{2} \partial_\mu \partial_\nu \bar{\pi}$  has an arrow pointing to "Spin=0".

$$5 \quad 2 \quad \left( 6 \quad T^{\mu\nu} \quad 3 \right) \left( \frac{1}{m^2} \right)$$

$$5 = 2 + 2 + 1 \quad M$$

$$\partial_\mu T^{\mu\nu} = 0$$

$$vDVZ$$



$$Gh + m^2 + hT^{\mu\nu}$$

$$m \rightarrow 0$$

## BSM Lecture 5

Köller-Lehman spectral representation  
for generalized graviton propagator.

$$T^{\mu\nu} \quad T^{\mu\nu} \quad T \equiv T_\mu^\mu$$

$$T^{\mu\nu} G_{\mu\nu, \alpha\beta} T'^{\alpha\beta} = \bullet \frac{1}{M_p^2} \left\{ \overbrace{\frac{T_{\mu\nu} T'^{\mu\nu} - \frac{1}{2} T T'}{p^2 + i\epsilon}}^{\text{Einstein}} \right\}$$

$$+ \sum_{m^2 > 0} \left\{ \frac{T_{\mu\nu} T'^{\mu\nu} - \frac{1}{3} T T'}{p^2 - m^2 + i\epsilon} C_2^{(m)} + \frac{T_{\mu\nu} T'^{\mu\nu}}{p^2 - m^2 + i\epsilon} C_0^{(m)} \right\}$$

$$\boxed{C_2^{(m)}, C_0^{(m)} > 0}$$

$$T_{\mu\nu} \longrightarrow M_1 \quad T'_{\mu\nu} \longrightarrow M_2$$

$$V(r) = -\frac{1}{M_p^2} \frac{M_1 M_2}{r} \left\{ 1 + \left(\frac{4}{3}\right) \sum_{m^2 > 0} C_2^{(m)} e^{-mr} \right.$$

$$\left. + 2 \sum_{m=0}^{\infty} C_0^{(m)} e^{-mr} \right\} \leftarrow$$

Equivalence principle.

Heisen gravity

$$H_0 = \langle H_0 \rangle + h_H^{(x)}$$



$$g_\psi H_0 \bar{\Psi}_L \Psi_R \rightarrow \underbrace{g_\psi v}_{m_\psi \equiv g_\psi v} \bar{\Psi}_L \Psi_R + g_\psi h_H \bar{\Psi}_L \Psi_R$$

$$\left(\frac{m_\psi}{v}\right) h_H \bar{\Psi}_L \Psi_R$$

$$\left(\frac{m_u}{v} \quad \frac{m_d}{v}\right)$$

$$m_H = \lambda v$$

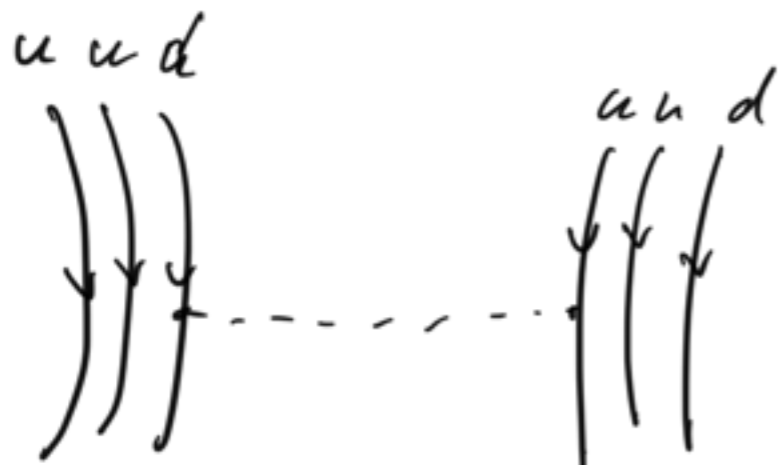
$$\lambda \rightarrow 0$$

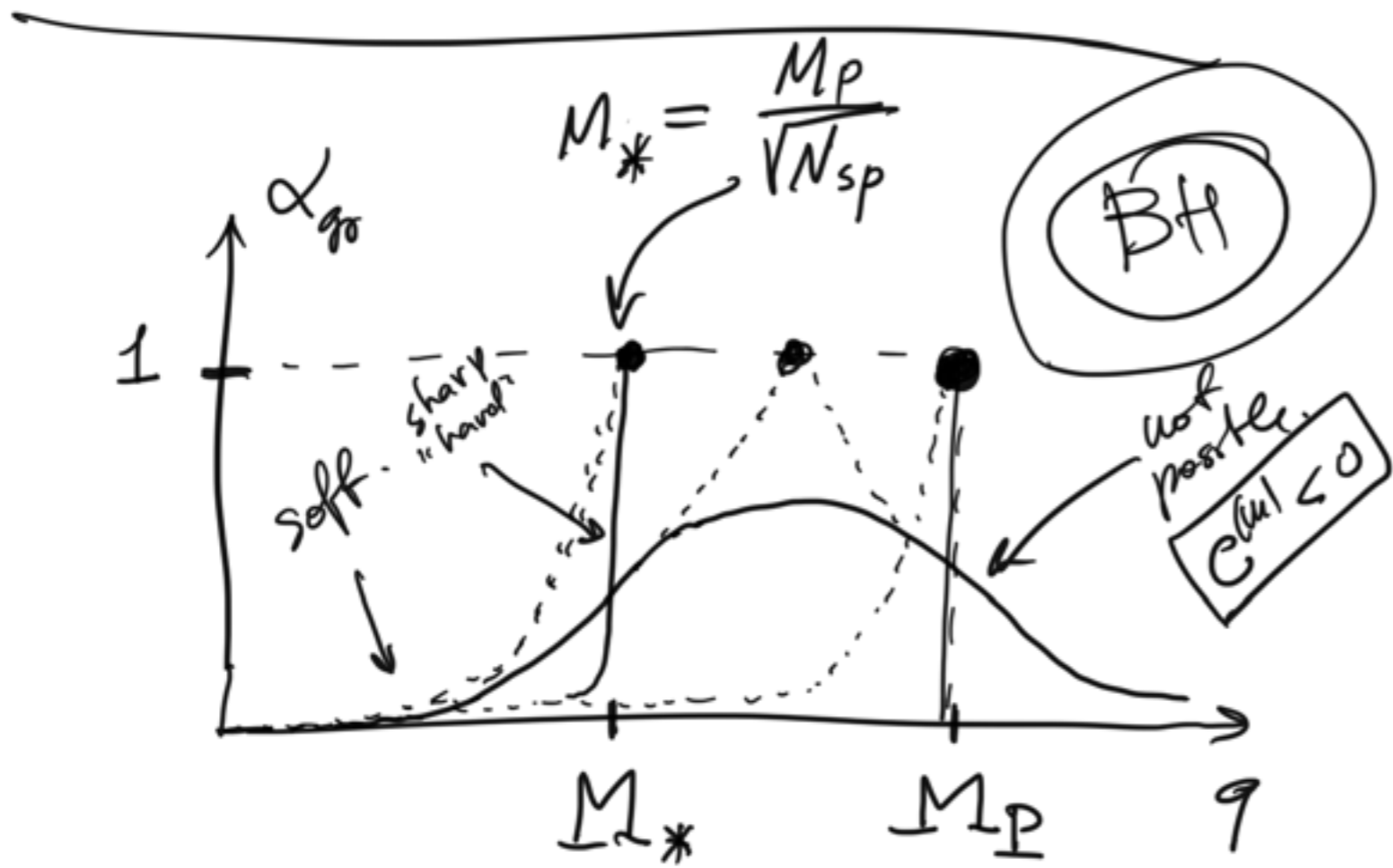
$$V_H(r) = -\frac{m_\psi^2}{v^2} \frac{e^{-m_\psi r}}{r}$$



Proton  
neutron

$$M_{\text{proton}} \gg 3M_q$$





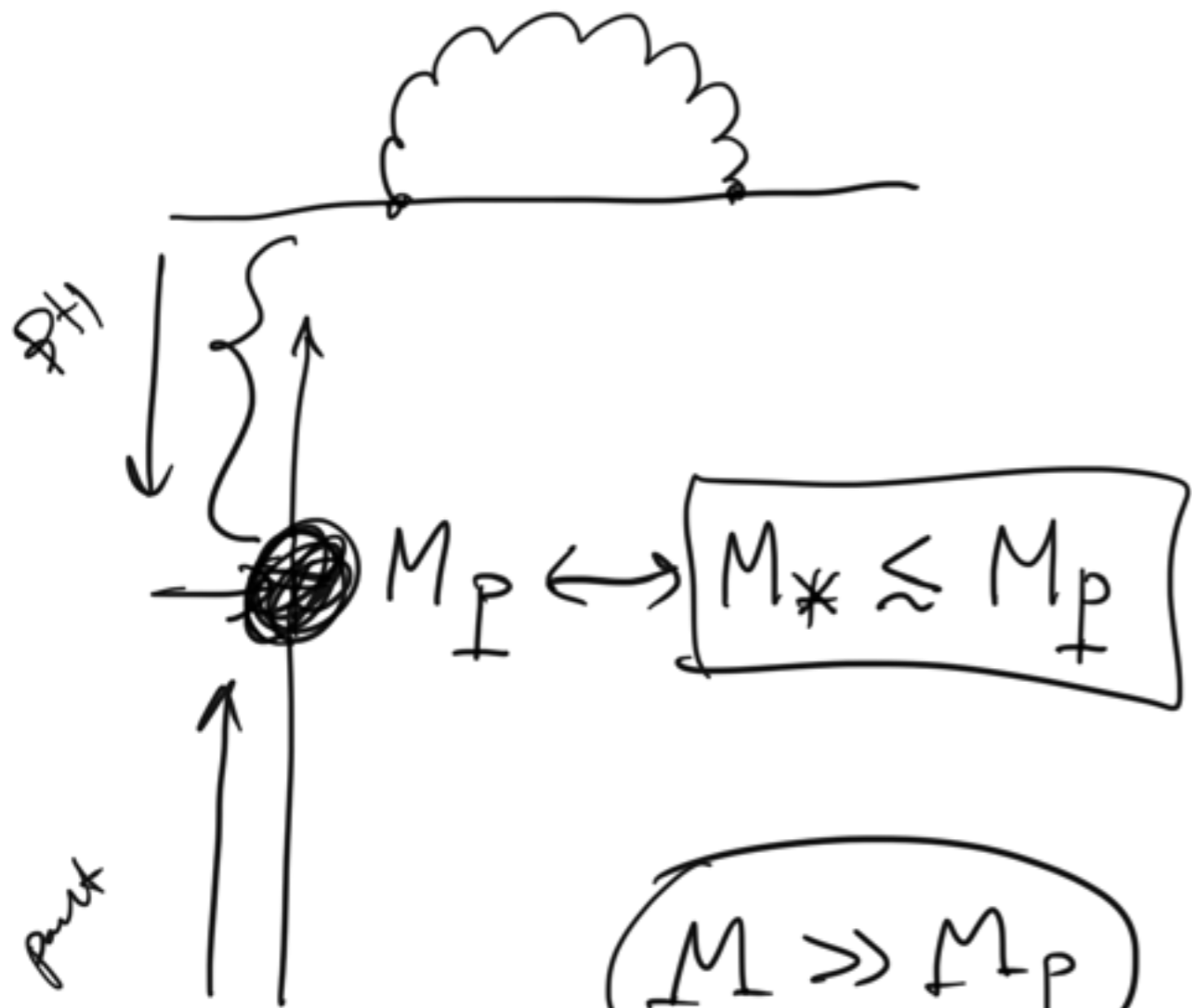
~~ana~~ ←  $L_{gr}$



$q \gg M_P$

toy thin  
 center of an  
 energy.  
 $\sqrt{s} \gg M_P$



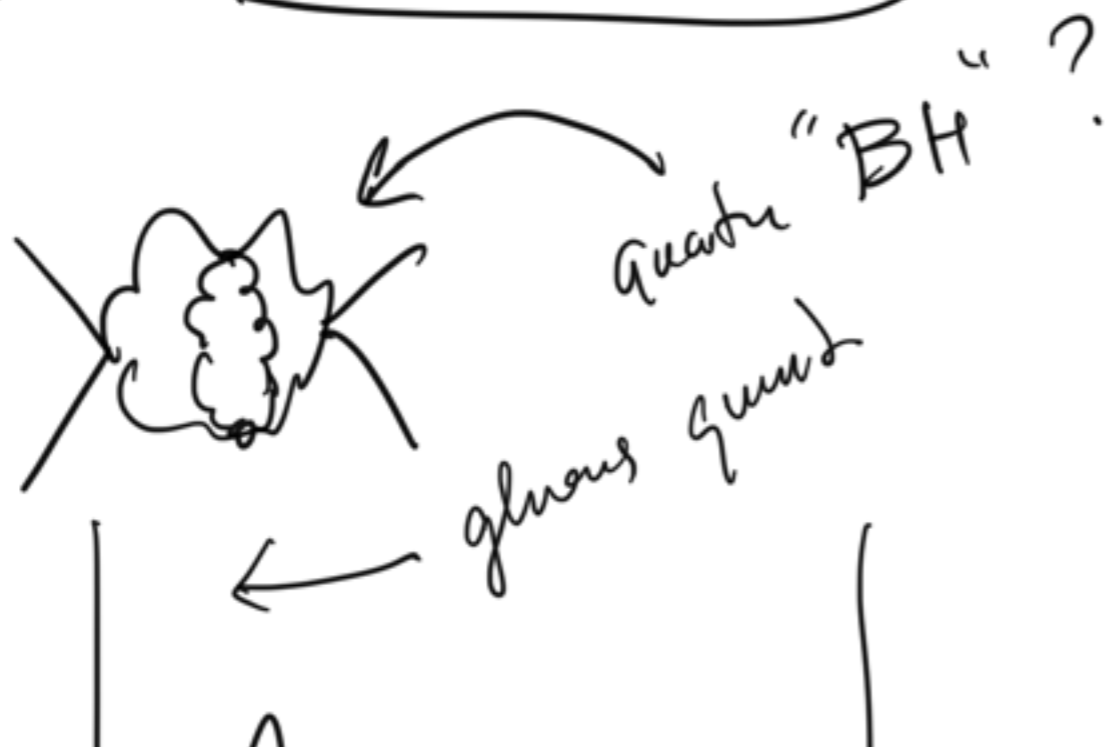


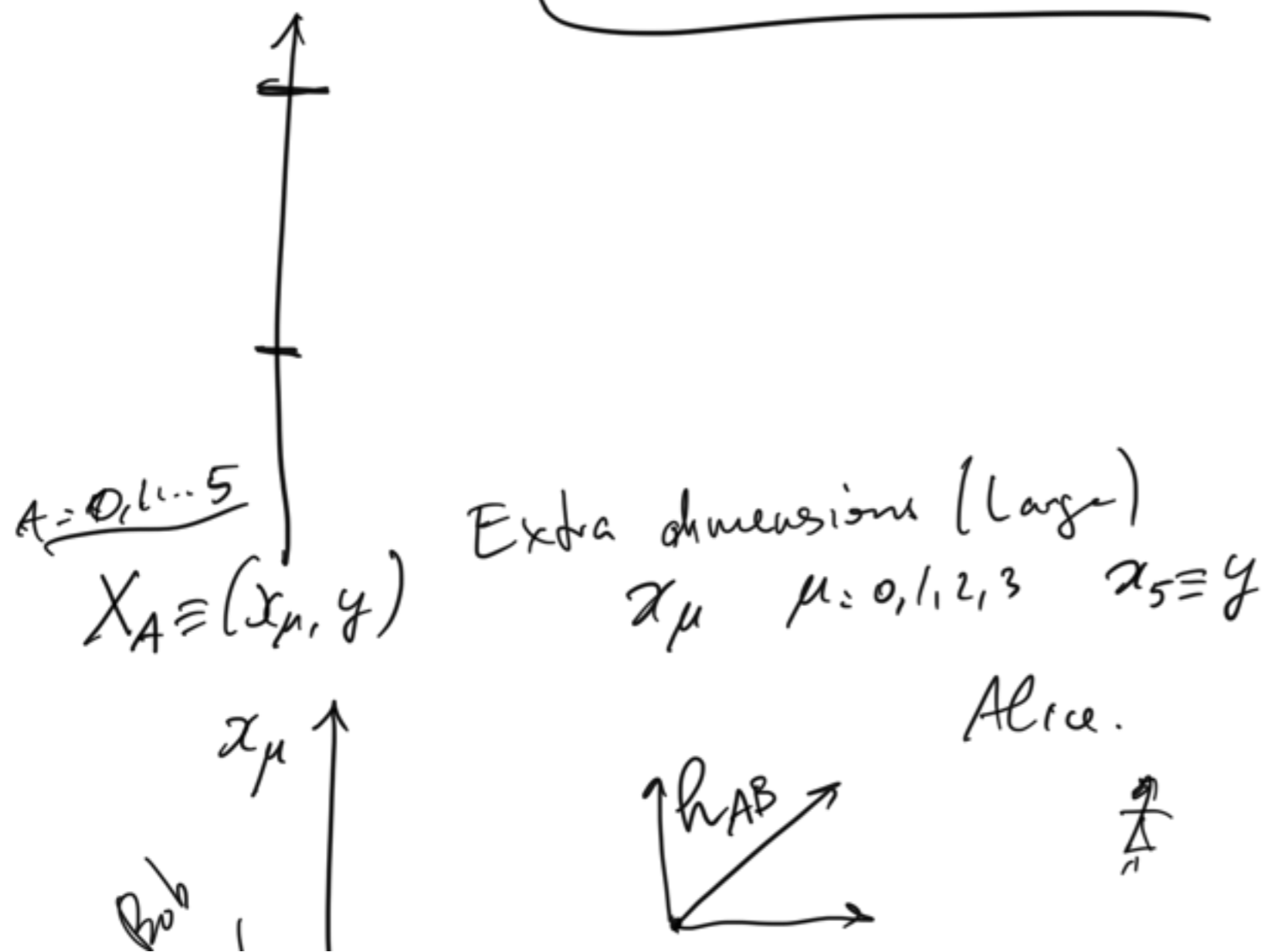
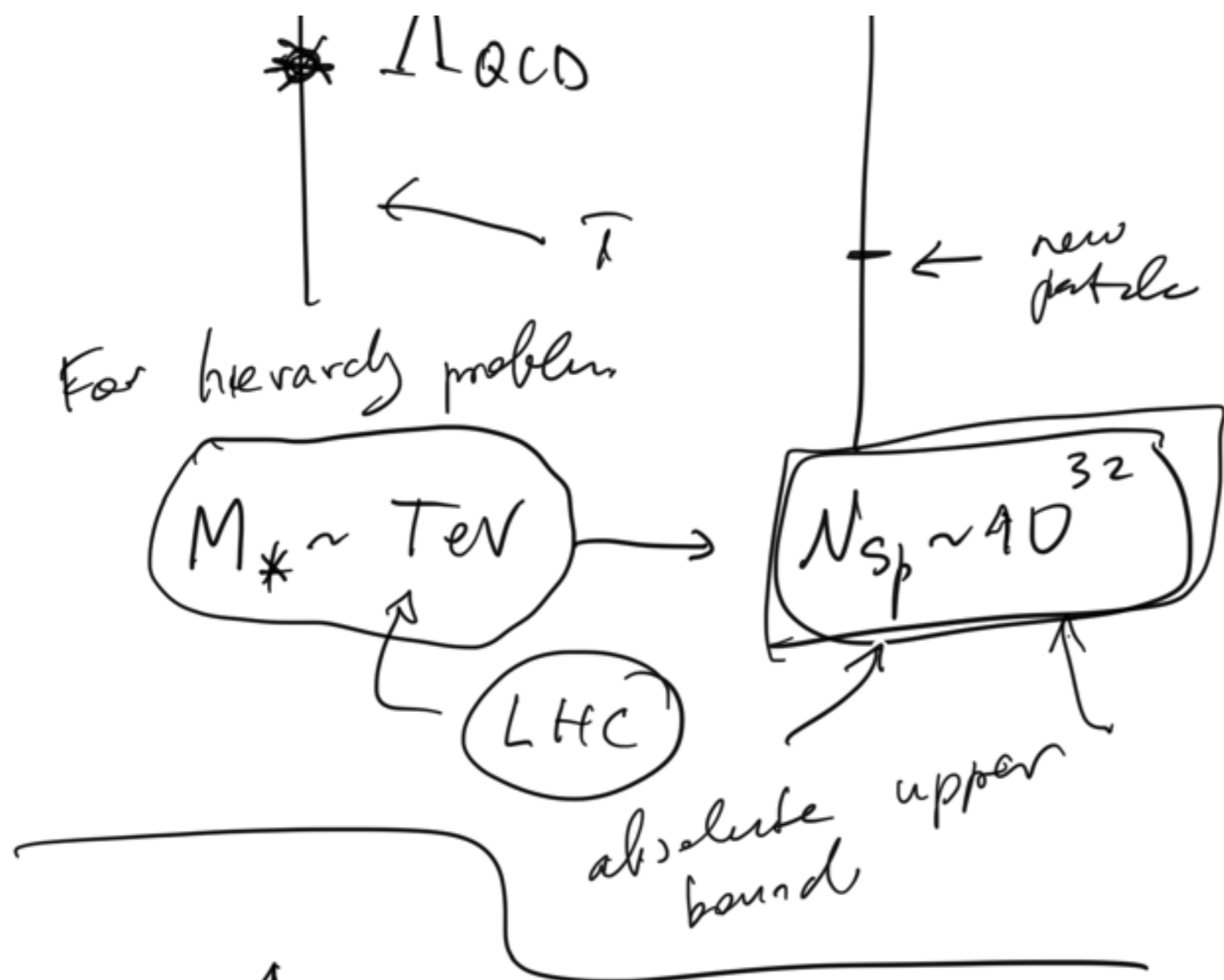
$$M \gg M_P$$

$$\Omega = \left( \frac{M}{M_P} \right)^2 \gg 1$$

$$\frac{1}{\Omega}$$

$$M \rightarrow M_P$$





$\lambda$

$$P_A P^A = 0 \quad m^2$$

$$P_\mu P^\mu - P_4^2 = 0 \rightarrow \boxed{P_\mu P^\mu = m^2}$$

$h_{AB}(x, y)$

$y$

$$\square h_{AB} - \eta_{AB} \square h - \partial_A \partial^C h_{CB} - \partial_B \partial^C h_{CA} + \partial_A \partial_B h + \eta_{AB} \partial^C \partial^D h_{CD} = 0$$

$$h_{AB} \rightarrow h_{AB} + \partial_A \xi_B + \partial_B \xi_A$$

$\begin{pmatrix} \mu\nu \\ \mu 5 \\ 55 \end{pmatrix}$

gauge parameter  $\xi_B(x^\mu, y)$

$$h_{AB} = h_{\mu\nu} + h_{\mu 5} + h_{55}$$

complete basis  $b^{(m)}(y)$   
 $\lambda \equiv \partial y$

symmetric tensor

vector

scalar

$\rho^{(m)}(x)$   $\rho^{(m)}(x)$   $M=0$  Ericksen's eqn.

$$h_{\mu\nu}(x,y) = \sum_m b(y) h_{\mu\nu}^{(m)}(x)$$

$$h_{\mu 5}(x,y) = \sum b^{(m)}(y) A_\mu^{(m)}(x)$$

$$h_{55}(x,y) = \sum b^{(m)}(y) \phi^{(m)}(x)$$

$m \neq 0$  Fierz  
 $m=0$  Maxwell  
 $m \neq 0$   
 $m=0$  scalar  
 $m \neq 0$

$$b^{(m)}(y) = -\omega^2 b(y)$$

$$b(y) = e^{imy}$$

$$\bar{A}_\mu^{(m)} \equiv A_\mu^{(m)} + \frac{1}{2} \partial_\mu \phi^{(m)}$$

$$(\mathcal{E} \hat{h}^{(m)})_{\mu\nu} - m^2 \left( \hat{h}_{\mu\nu}^{(m)} - \eta_{\mu\nu} \hat{h}^{(m)} + \partial_\mu \bar{A}_\nu^{(m)} + \partial_\nu \bar{A}_\mu^{(m)} - \eta_{\mu\nu} 2 \partial^\alpha \bar{A}_\alpha^{(m)} \right) = 0$$

$$\hat{h}_{\mu\nu}^{(m)} \rightarrow \hat{h}_{\mu\nu}^{(m)} + \partial_\mu \xi_\nu^{(m)} + \partial_\nu \xi_\mu^{(m)}$$

$$\bar{A}_\nu^{(m)} \rightarrow A_\nu^{(m)} - \xi_\nu^{(m)}$$

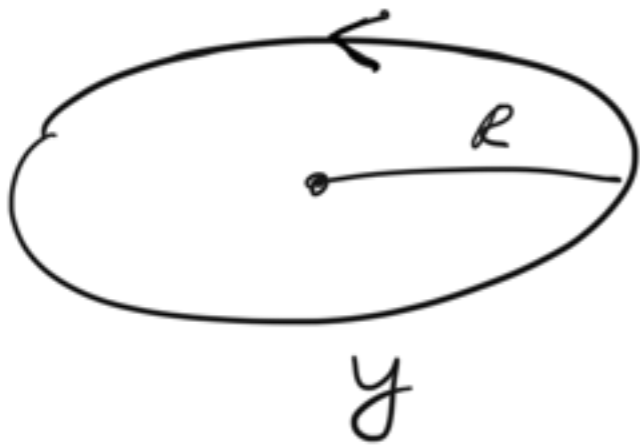
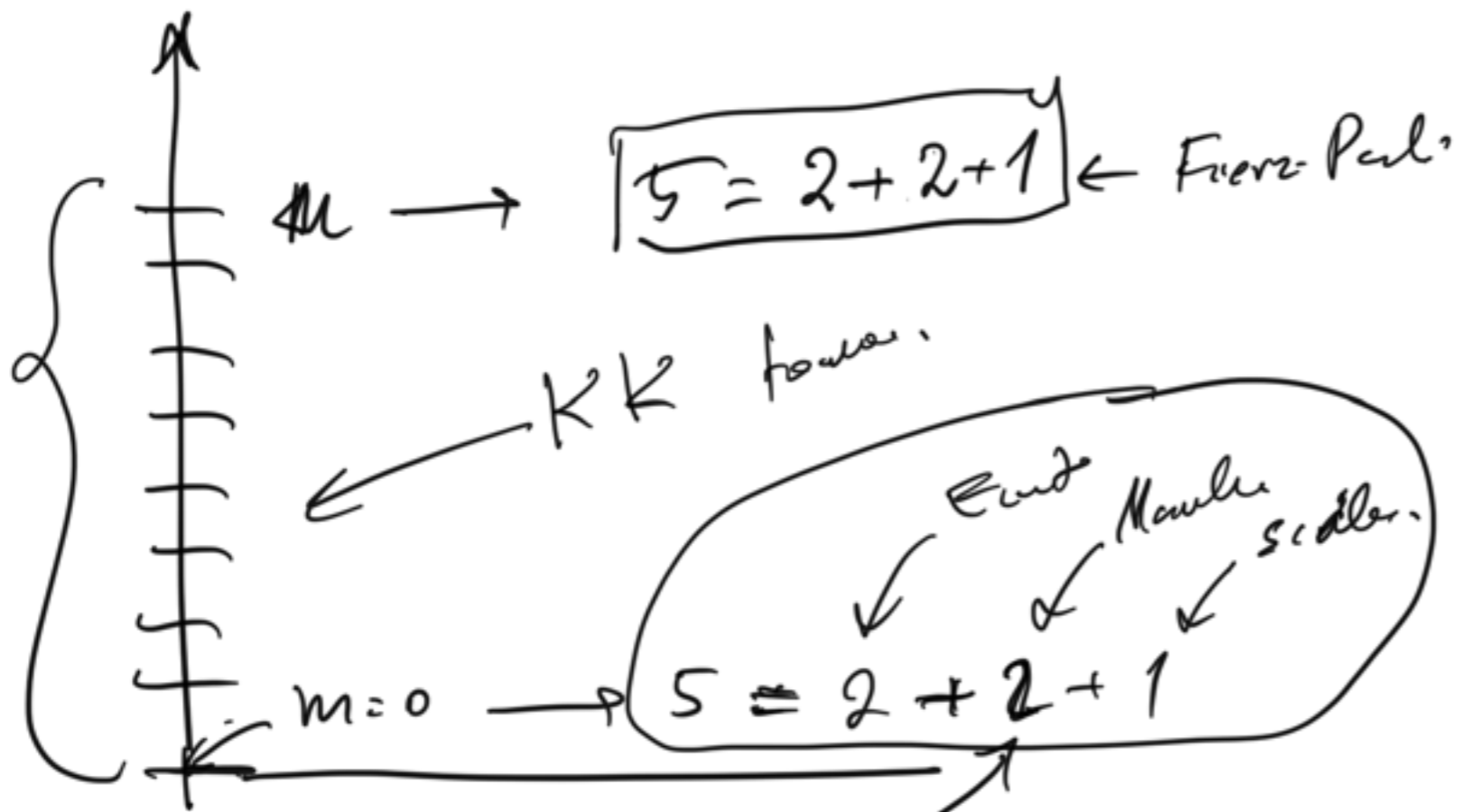
Fierz-Pauli

Stückelberg

- in energy field.

$$\partial_{\mu} F_{\mu\nu}^{(m)} + \partial^{\mu} (\underbrace{h_{\mu\nu}^{(m)} - \eta_{\mu\nu} h^{(m)}}_{\text{trace}}) = 0$$

Maxwell-like with a sum.



$$m = \frac{n}{R}$$

Goldstone.

Poincare 5  $\rightarrow$  Poincare 4  $\otimes$  (U(1))

$$y \rightarrow y + \text{const.}$$

$$S = \int dy d^3x \underline{M_*^3} \left( h^{AB} \epsilon_{AB} \right)$$

$$\rightarrow \int d^4x \underline{M_{\text{pl}}^2} \left( h^{\mu\nu} \epsilon_{\mu\nu} \right) + \dots$$

$$\underline{M_{\text{pl}}^2} = \int dy \underline{M_*^3} = M_*^2 \left( \underline{M_* R} \right)$$

$$\underline{M_{\text{pl}}^2} = \underline{M_*^2} N_{KK}$$

$$m = \frac{\hbar}{R}$$





$L \pm$

$$M_{\pm p}^2 = M_*^2 (M_*^d V^{-(d)})$$

For addressing hierarchy puzzle.

$$N_{sp} \sim 10^{32}$$

$$M_w \sim e^{-\frac{1}{g^2}} M_{\pm p}$$

Tower of "BH"s



$$M_* \sim \text{TeV}$$

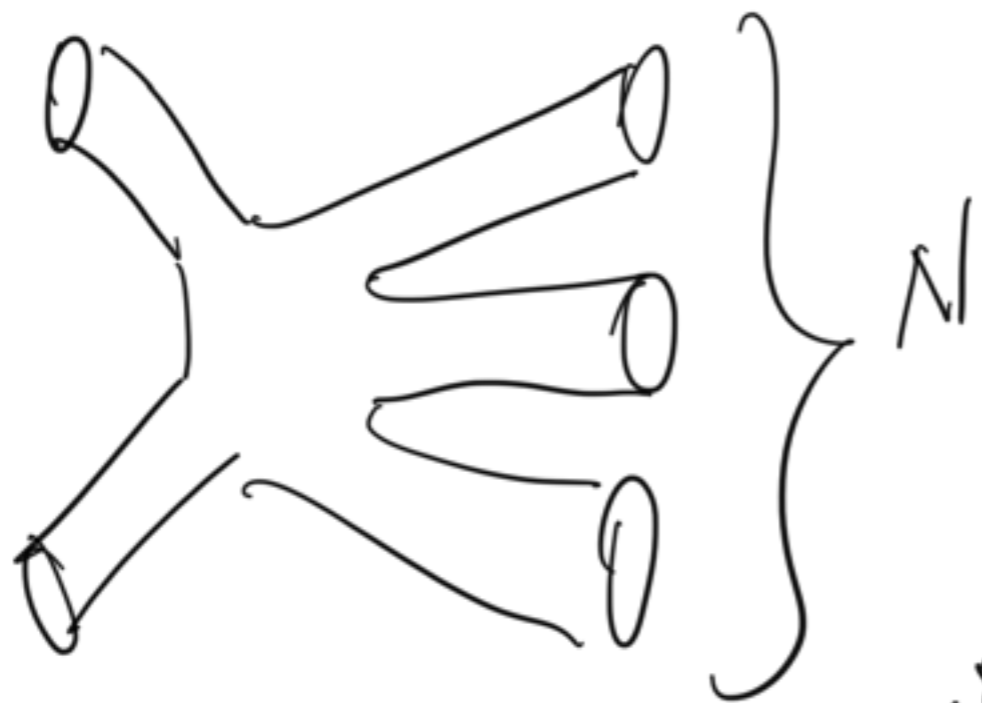
predictions are resonance.  
"quantum BH"

heavier resonance

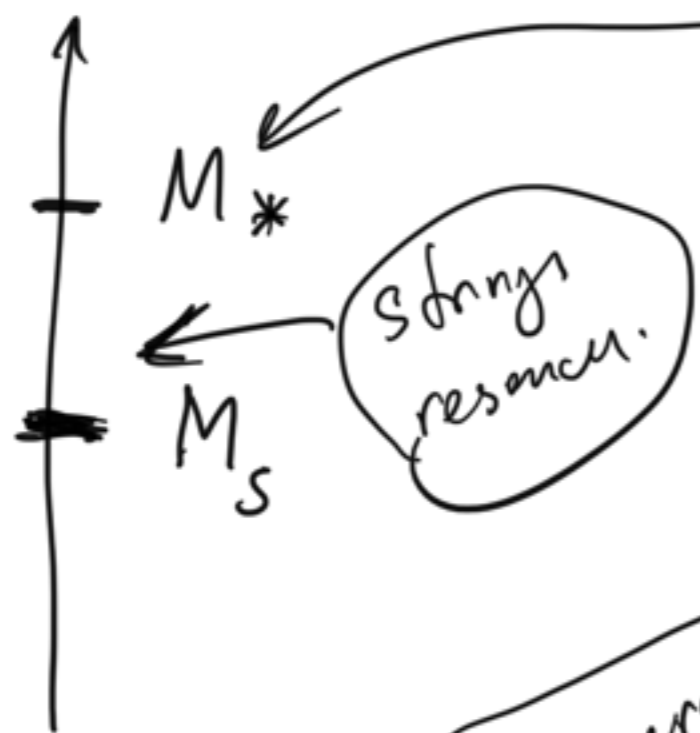
$$M \gg M_*$$

must be lower lived

and decay into more number  
of softer quanta.



$$M_{(10)}^8 = \frac{1}{g_s^2} M_s^8$$



$$g_s \ll 1$$

requires a  
microscopic theory



~ ~ ~ ~ ~

$S \sim 1$

$\mu$   
 $\mu$

$e^{-S}$

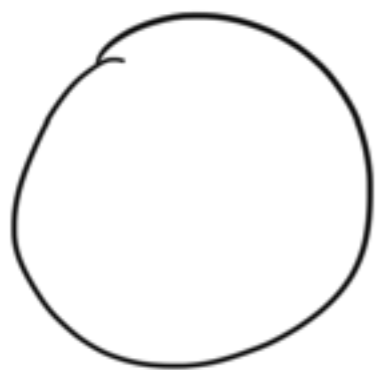
$N \sim S$

$\frac{1}{S}$



$N \gg 1$

$\frac{1}{N}$



BH



BH'



$T^2$

