Path to discoveries: Quo Vadis?



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Galileo Galilei Institute, September 04, 2018



Most Serene Prince,

Galileo Galilei most humbly prostrates himself before Your Highness, watching carefully, and with all spirit of willingness, not only to satisfy what concerns the reading of mathematics in the study of Padua, but to write of having decided to present to Your Highness a telescope ("Occhiale") that will be a great help in maritime and land enterprises. I assure you I shall keep this new invention a great secret and show it only to Your Highness. The telescope was made for the most accurate study of distances. This telescope has the advantage of discovering the ships of the enemy two hours before they can be seen with the natural vision and to distinguish the number and quality of the ships and to judge their strength and be ready to chase them, to fight them, or to flee from them; or, in the open country to see all details and to distinguish every movement and preparation.





Beware of fake news!

How to make discoveries?

Following Galileo: New instrumentation and techniques New ideas



The γ-ray source scheme for CERN



A major recent news from CERN!



During a special one-day run, LHC operators injected lead "atoms" containing a single electron into the machine (Image: Maximilien Brice/Julien Ordan/CERN)

Protons might be the Large Hadron Collider's bread and butter, but that doesn't mean it can't crave more exotic tastes from time to time. On Wednesday, 25 July, for the very first time, operators injected not just atomic nuclei but lead "atoms" containing a single electron into the LHC. This was one of the first proof-of-principle tests for a new idea called the Gamma Factory, part of CERN's Physics Beyond Colliders project.

Gamma Factory @ CERN

Partially Stripped Ion beam as a light frequency converter

$$v^{\text{max}} \longrightarrow (4 \gamma_{\text{L}}^2) v_{\text{i}}$$

Tuning of the beam energy, the choice of the ion type, the number of left electrons and of the laser type allows to tune the γ -ray energy, at CERN, in the energy domain of 100 keV – 400 MeV.

Example (maximal energy): LHC, Pb⁸⁰⁺ ion, γ_L = 2887, n=1 \rightarrow 2, λ = 104.4 nm, E_{γ} (max) = 396 MeV

Witek Krasny

Gamma Factory @ CERN

The gamma ray source for Gamma Factory

<u>The idea:</u> replace an electron beam by a beam of highly ionised atoms (Partially Stripped Ions - **PSI**)





K.A. ISPIRIAN, A.T. MARGARIAN, N.G. BASOV, A.N. ORAEVSKI, B.N. CHICHKOV E.G. BESSONOV, K-J. KIIM, M.W. KRASNY...

The expected magnitude of the γ -source intensity leap	
Electrons:	Partially Stripped Ions:
$\sigma_{\rm e} = 8\pi/3 \ {\rm x} \ {\rm r_e}^2$	$\sigma_{\rm res} = \lambda_{\rm res}^2 / 2\pi$
r _e - classical electron radius	λ _{res} - photon wavelength in the ion rest frame
Electrons:	Partially Stripped Ions:
$\sigma_{e} = 6.6 \text{ x} \ 10^{-25} \text{ cm}^{2}$	$\sigma_{\rm res}$ = 5.9 x $10^{-16}{\rm cm}^2$
<u>Numerical example:</u> λ _{laser} = 1540 nm	
~ 9 orders of magnitude difference in the cross-section	

~ 7 orders of magnitude increase of gamma fluxes

Very exciting times for CERN physics!



Really Important Questions... Fundamental Constants?



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Search for new physics with atoms and molecules

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Marianna Safronova with Erwin L. Hahn and the Budker Group, Berkeley, 2006

Parity Violation in Yb (without T-violation)



Parity died in 1956







ATOMIC PARITY VIOLATION Main Source: Z exchange





Weak interaction (violates parity)

• P-odd, T-even correlation $\vec{\sigma} \cdot \vec{p}$

Yb atomic beam apparatus



First observation of isotopic variation of atomic PV



SM: $Q_W \approx -N + Z(1 - 4\sin^2\theta_W) \rightarrow 1\%$ change per neutron around N=103

Observation: 0.96(15) % change per neutron

Implication for <u>light</u> mediators (V.V. Flambaum)



$$\mathcal{L} = Z'_{\mu} \sum_{f=e,p,n} \overline{f} \gamma^{\mu} (g^{V}_{f} + \gamma_{5} g^{A}_{f}) f$$
PRL 119, 223201 (2017)
PHYSICAL REVIEW LETTERS
Week ending
1 DECEMBER 2017

Probing Low-Mass Vector Bosons with Parity Nonconservation and Nuclear Anapole Moment Measurements in Atoms and Molecules

V. A. Dzuba,1 V. V. Flambaum,1,2 and Y. V. Stadnik2



Reference

D. Antypas, A. Fabricant, J.E. Stalnaker, K. Tsigutkin, V. V. Flambaum, and D. Budker *Isotopic variation of parity violation in atomic ytterbium* <u>arXiv:1804.05747</u>; accepted to Nature Physics

Summary (PV in Yb)

- ➢ Measured PV on a chain of isotopes
- № 0.5% accuracy per isotope
- Next stop: anapole moments of ¹⁷¹Yb& ¹⁷¹Yb











D. Antypas

A. Fabricant

J. Stalnaker

K. Tsigutkin

Axion-like particles (ALPs) search with NMR



So what is DM or what mimics it ?

- A gross misunderstanding of gravity (MOND, ...) ⊗?
- Proca MHD (finite photon mass)
- Black holes, dark planets, interstellar gas, …

• WIMPS

Ultralight bosonic particles

- Axions (pseudoscalar)
- ALPs (pseudoscalar)
- Dilatons (scalar)
- Vector particles
- Tensor particles



?

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"Most Wanted" file on DM What do we know?

- Galactic DM density: ~0.4 GeV/cm³ (10 GeV/cm³ d.g.)
 Has to be nonrelativistic: v/c ~ 10⁻³ (cold DM)
- Has to be bosonic if $m < \sim 20 \text{ eV}$ (1 keV dwarf galaxies)
- "Bosonic Oscillator" with $Q \sim (v/c)^{-2} \sim 10^6$
- \square Cannot be lighter than ~ 10⁻²² eV
- ... (e.g., BEC ?)



Dark Matter search with NMR

Key Ideas:

- Dark Matter could be a "classical" field
- Not screened by shielding
- Oscillating at frequency: *mc*²/*h*
- Relatively narrow line: $\Delta \nu / \nu \sim 10^{-6}$



Cosmic Axion Spin-Precession Experiment(s)
Cosmic Axion Spin-Precession Experiment(s)
CASPEr

Nuclear Magnetic Resonance (NMR)



Resonance: $2\mu B_{\text{ext}} = \omega$

CASPEr



Larmor frequency = axion mass → resonant enhancement SQUID measures resulting transverse magnetization Example materials: liquid ¹²⁹Xe, ferroelectric PbTiO₃

Dark Matter search with ZULF NMR



CASPEr With Zero- to Ultralow-Field NMR



- Search for dark-matter-induced sidebands
- Coherent averaging of arbitrary frequency via post-processing phase cycling
- Upcoming sensitivity improvements with PHIP

 $o(-2\pi i\phi_2)$

 $N = N\omega\tau$

 $\rho(-2\pi i \phi_N)$

Average

•----•

Average

 $\phi_1 = \omega \tau$ $(\bigcirc \phi_2 = 2\omega \tau)$

()

e^(-2πiφ₁)

Signal

FFT

Phase Shift



CASPEr ZULF: preliminary results



Personal opinions...

- BSM vs. BBSM physics
- Merger of particle physics, astronomy, cosmology,... in non-collider physics
- Discoveries are around the corner!
- Adopt new technologies or perish

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Low-energy Tests of Fundamental Physics

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SUMMARY





Number of neutrons

ALPs search with CASPEr

