

BARKANA, NJO, REDIGOLO, VOLANSKY.  
LIU, NJO, REDIGOLO, VOLANSKY.

[ARXIV:1803.03091]  
[ARXIV:1908.06986]

NADAV JOSEPH OUTMEZGUINE

TEL AVIV UNIVERSITY

GGI - SEP 2019

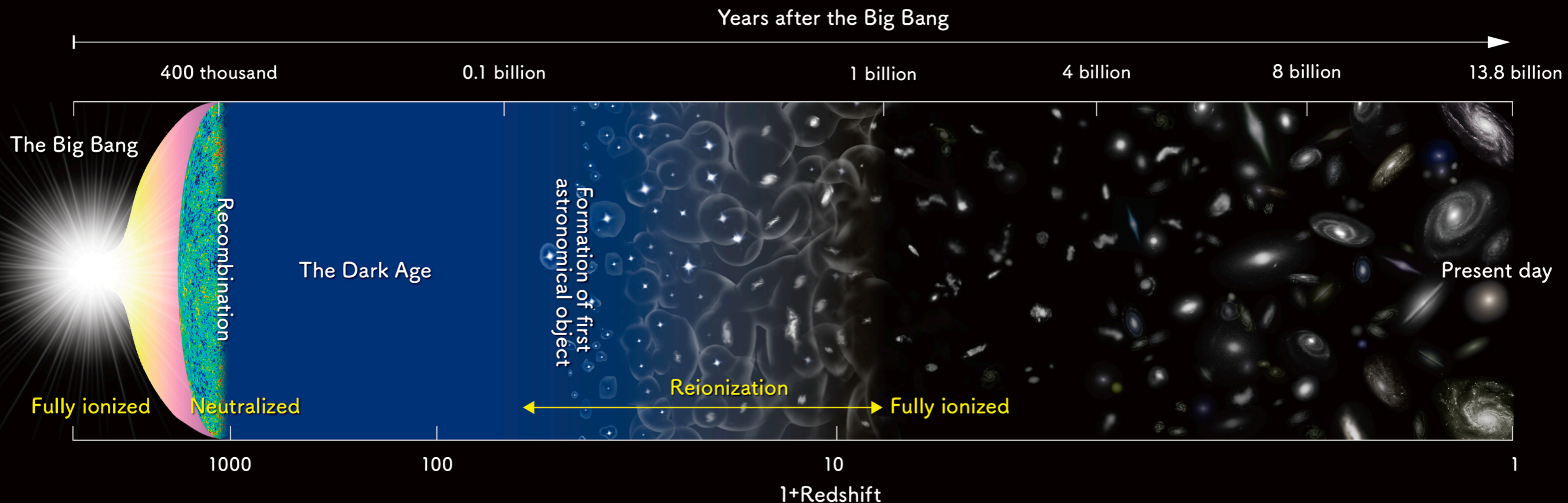
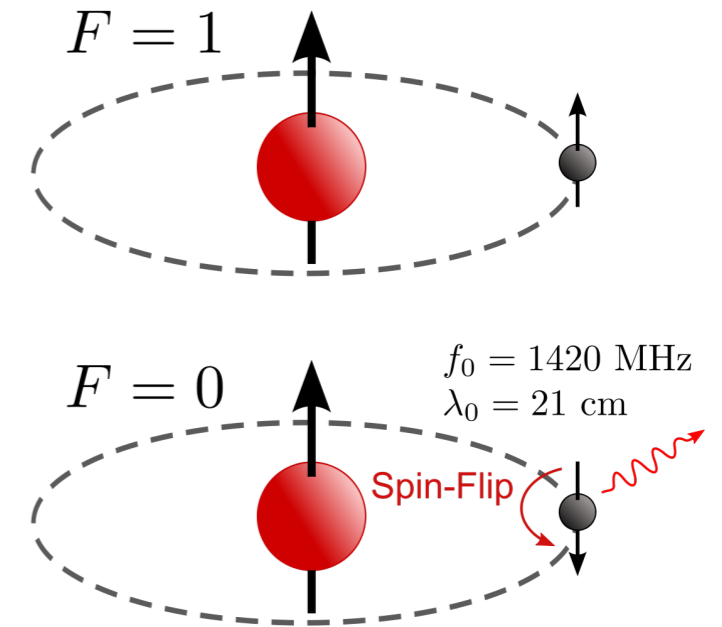
21-CM COSMOLOGY WITH

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**CHARGED DARK MATTER**

# WHY SHOULD I CARE -1. (PHYSICS)

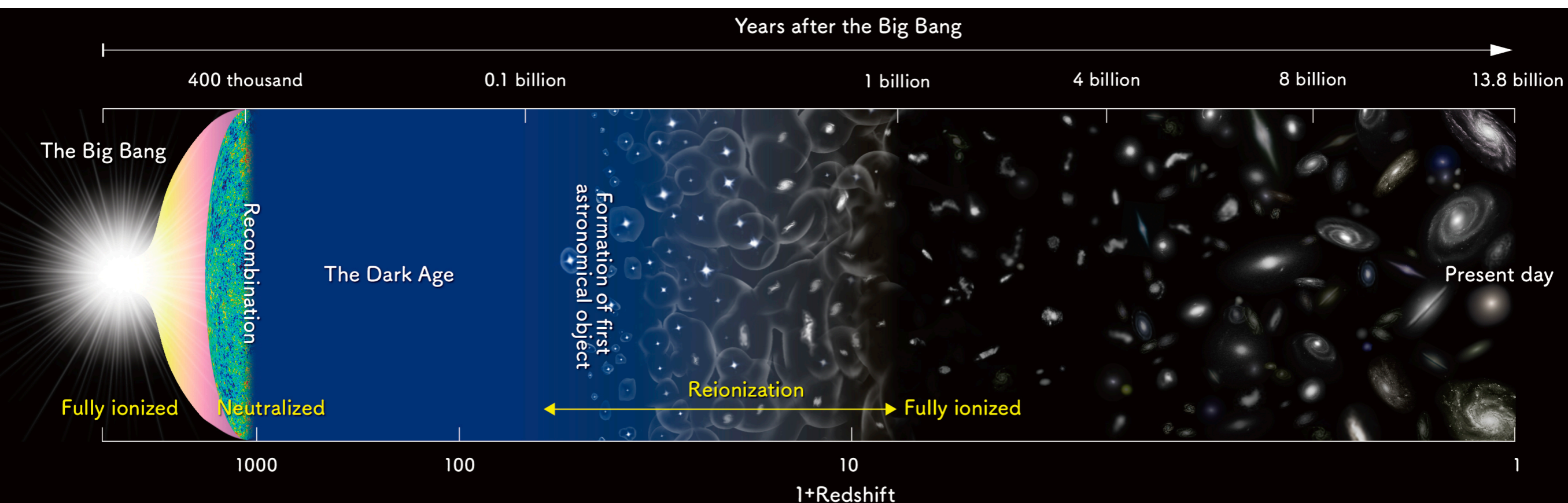
When H atoms are abundant, hyperfine transitions can be used to track the temperature of baryons



## WHY SHOULD I CARE -1. (PHYSICS)

A glance into an epoch yet to be explored:  $(1100 \lesssim z + 1 \lesssim 15)$

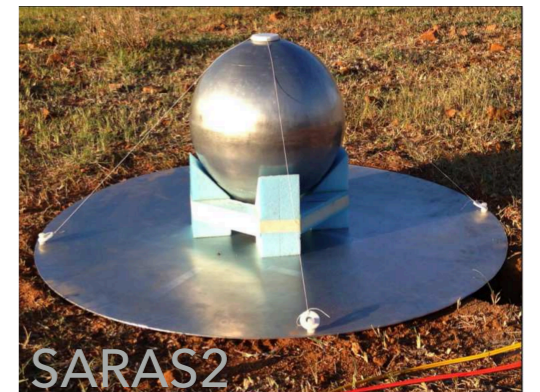
- ▶ Coldest Epoch, post recombination but before reionization
- ▶ Lots of H atoms  $\rightarrow$  Lots of hyperfine transitions
- ▶ Another test of  $\Lambda$ CDM



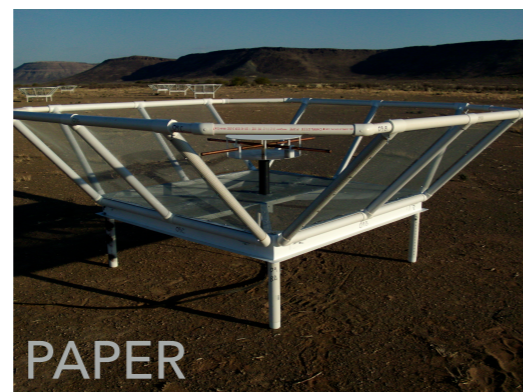
## WHY SHOULD I CARE -2. (EXPERIMENTS)

Many(!) experiments

Global signal



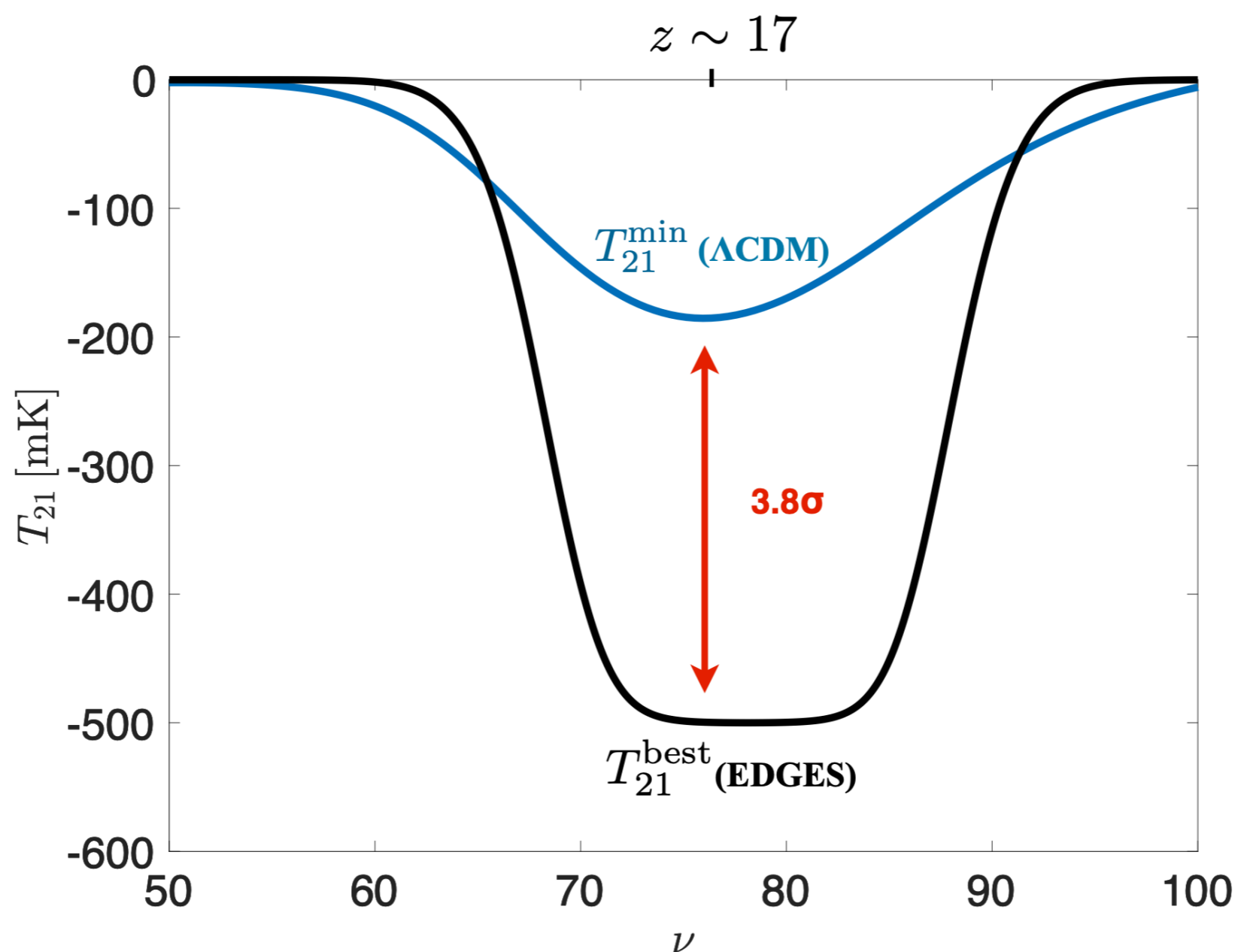
Fluctuations



And more...

## WHY SHOULD I CARE -3. (SOCIOLOGY)

Possibly an anomalous signal - Bowman et al., 2018



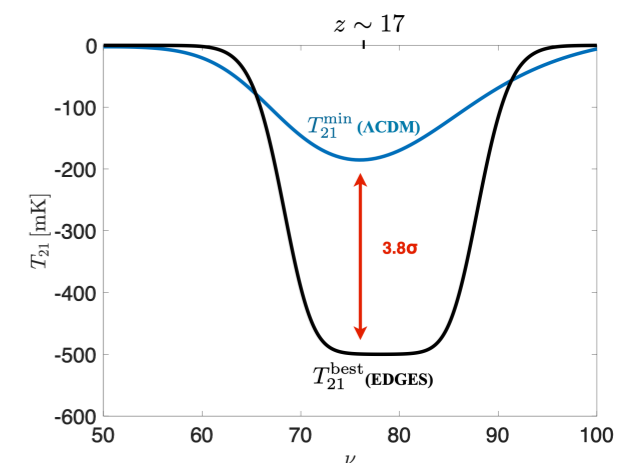
From E.D.Kovetz talk at GGI last year

# BASICS

Spin Temperature: 
$$\frac{n_1}{n_0} \equiv \frac{g_1}{g_0} e^{-E_{21}/T_s} \simeq 3 \left( 1 - \frac{E_{21}}{T_s} \right)$$

Standard Model Evolution:

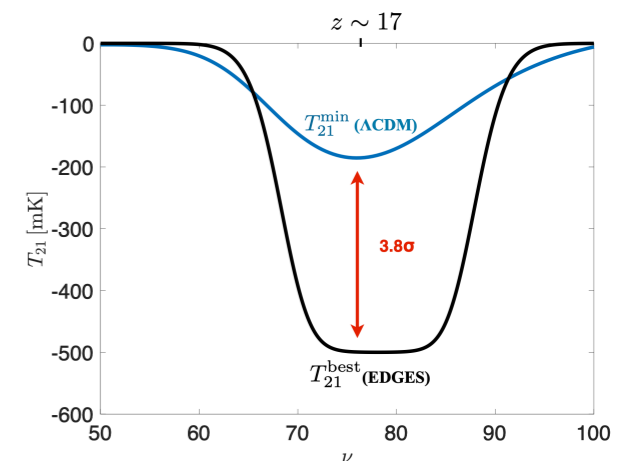
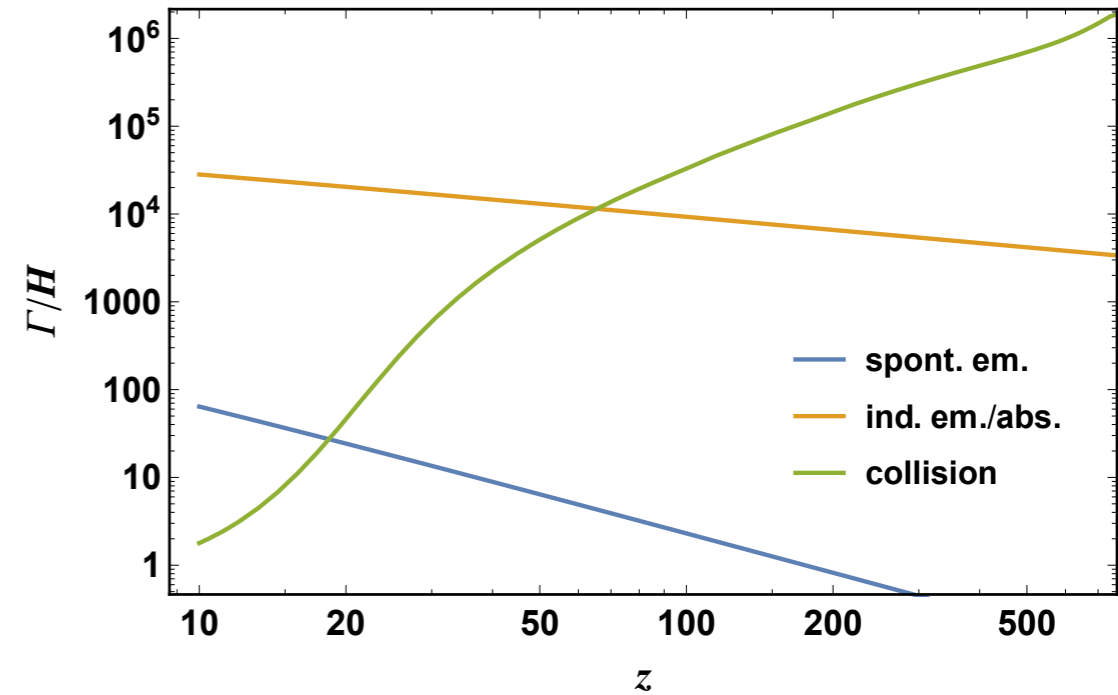
- ▶ H-H and H-e collisions
- ▶ Spontaneous emission and induced emission/absorption
- ▶ Coupling to Ly $\alpha$  radiation



## BASICS

### Evolution - Boltzmann

- ▶ Dominant rate always larger the Hubble  $\rightarrow$  detailed balance



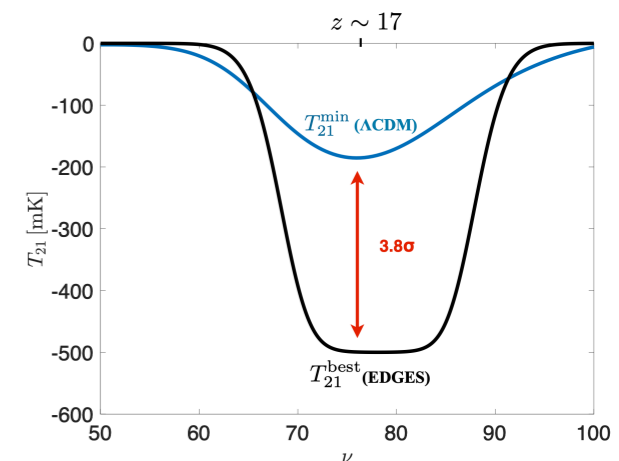
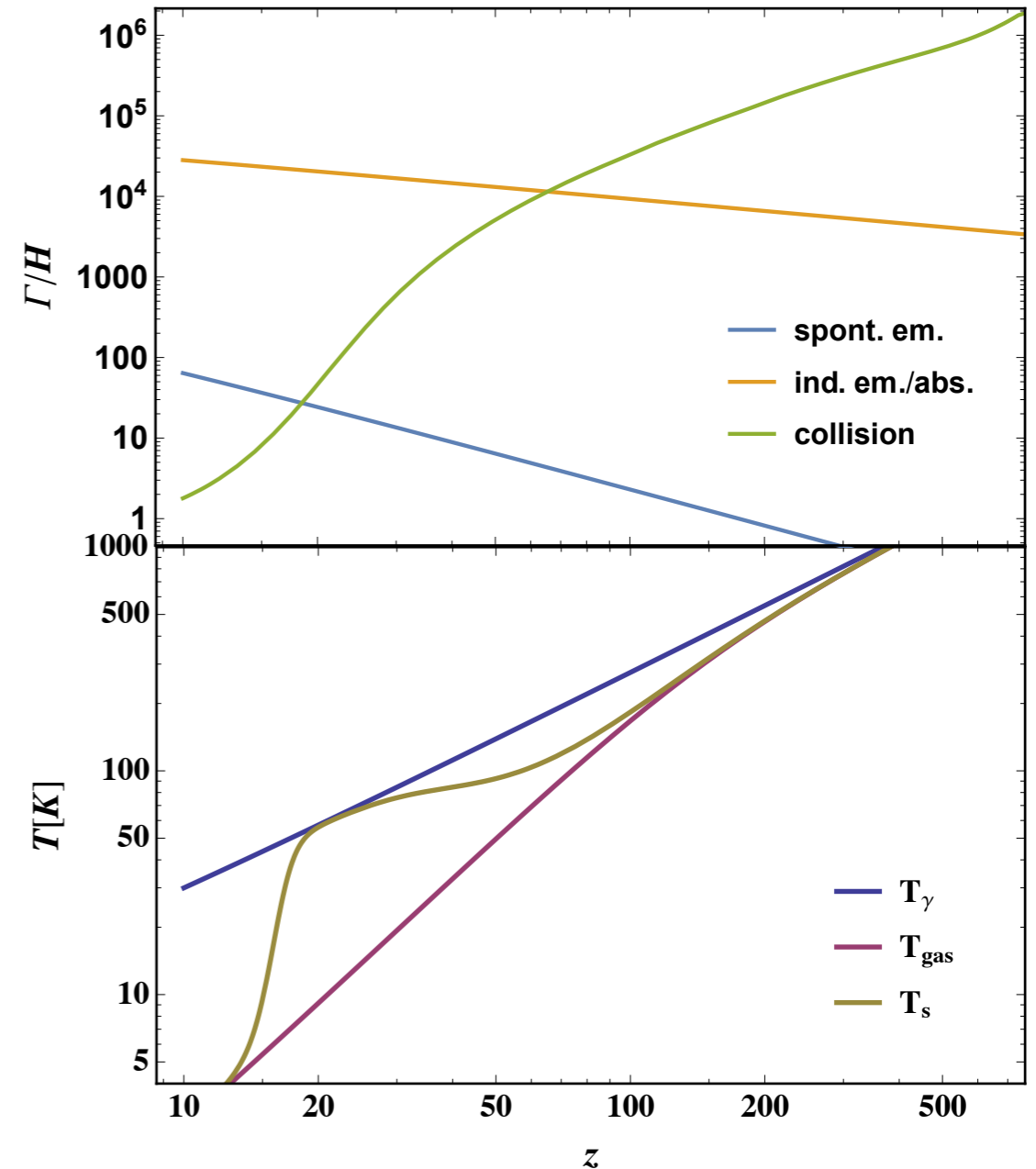
# BASICS

## Evolution - Boltzmann

- ▶ Dominant rate always larger the Hubble  $\rightarrow$  detailed balance
- ▶ Spin temperature follows the dominant process:

$$\Delta T_s = \frac{c_{\text{col}} \Delta T_{\text{gas}} + c_{\text{Ly}\alpha} \Delta T_{\text{Ly}\alpha}}{c_{\text{ind}} + c_{\text{col}} + c_{\text{Ly}\alpha}}$$

$$\Delta T = T - T_{\text{CMB}}$$





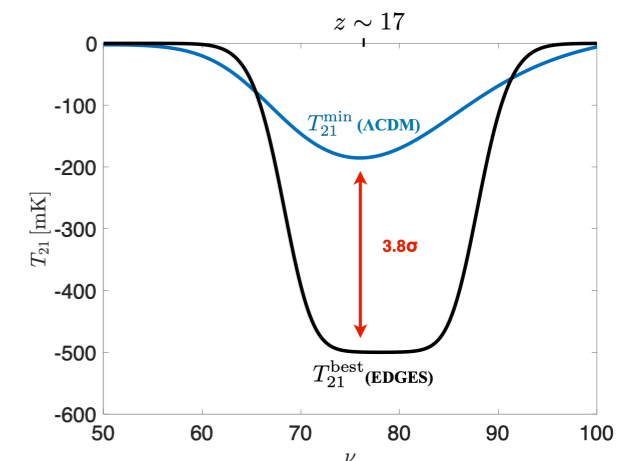
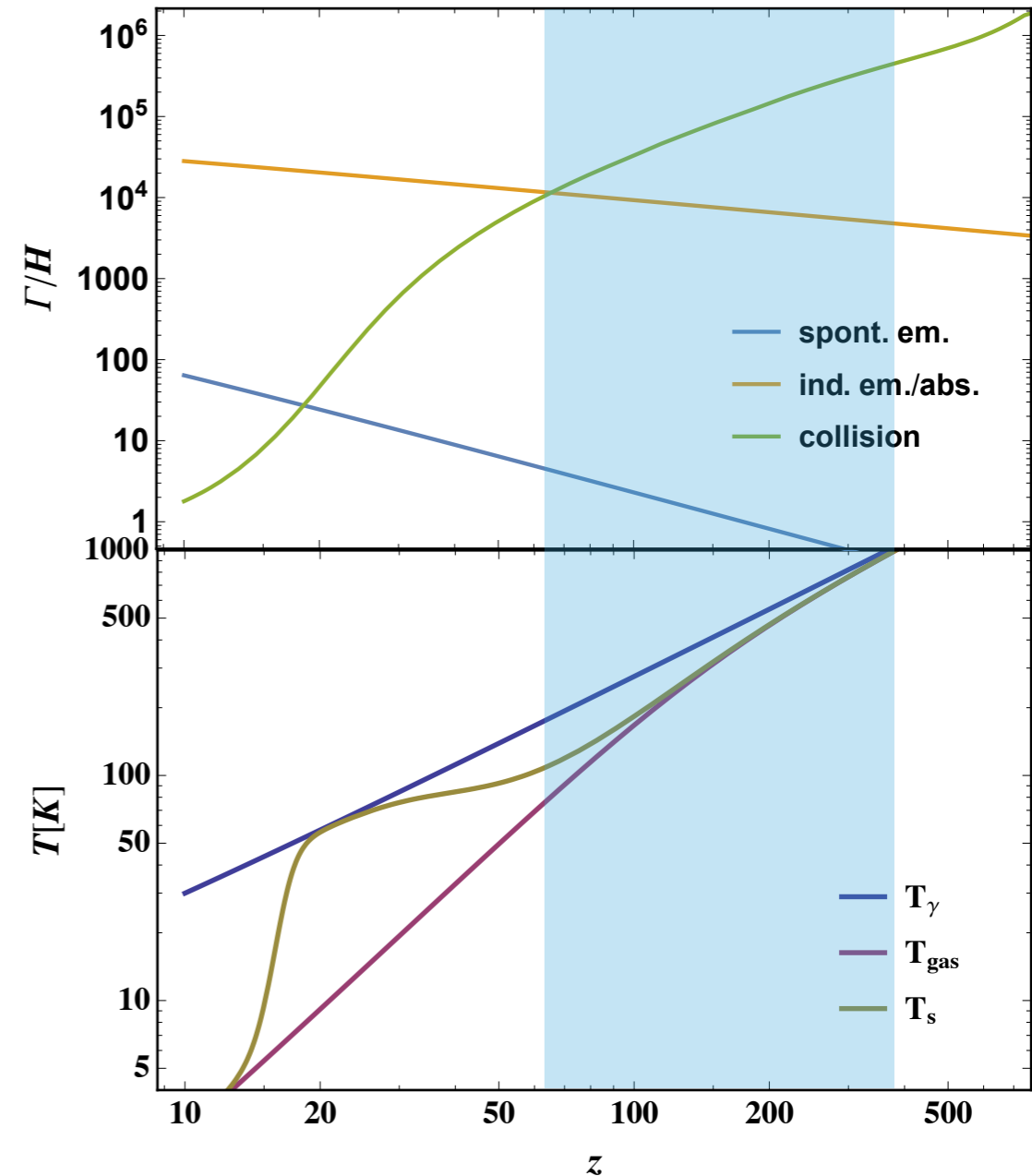
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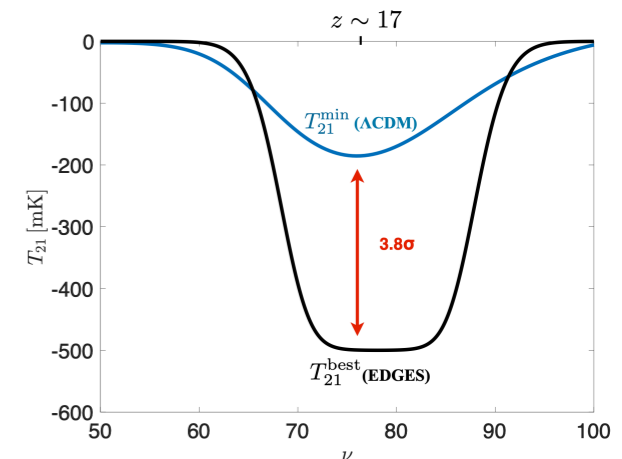
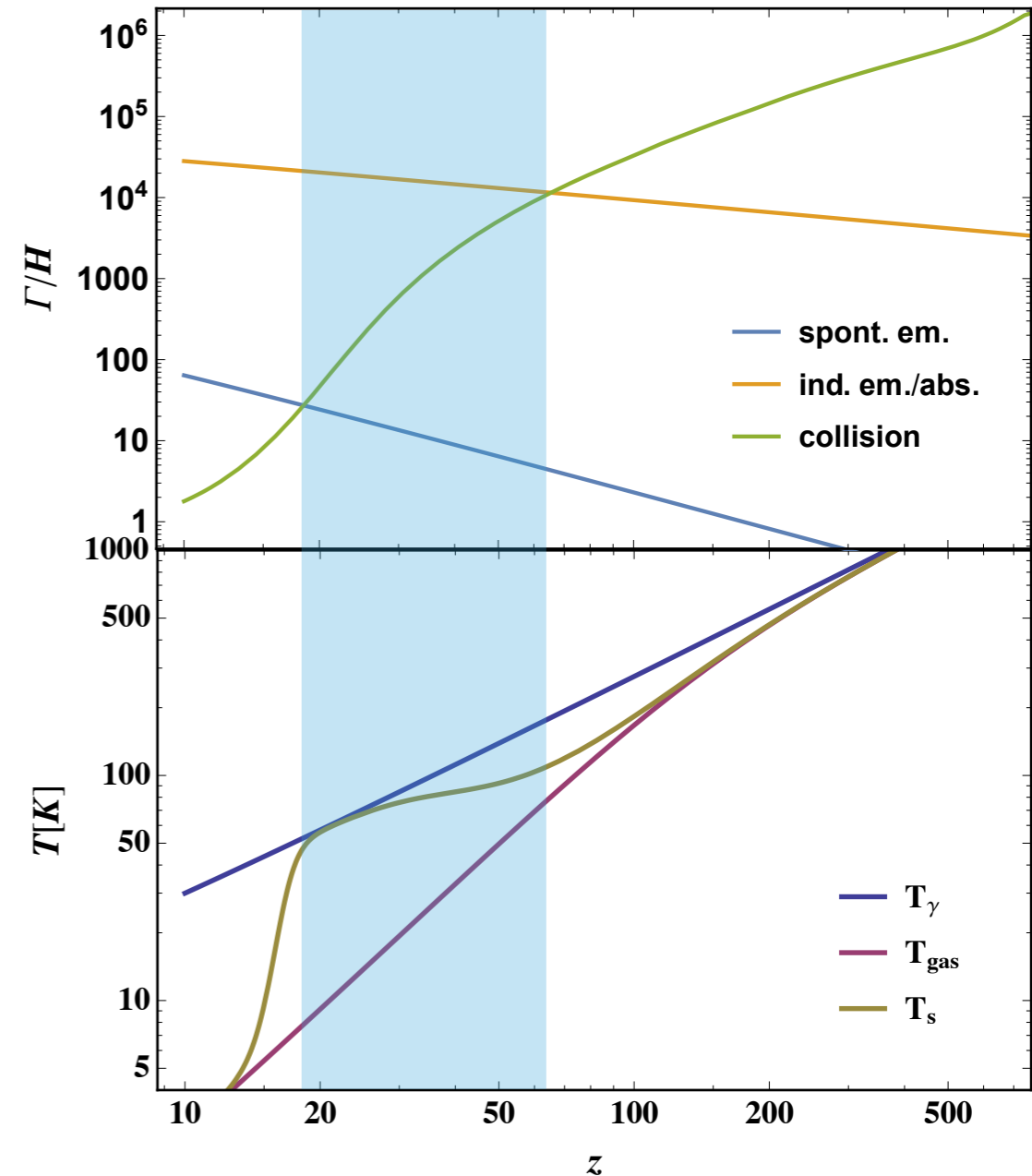
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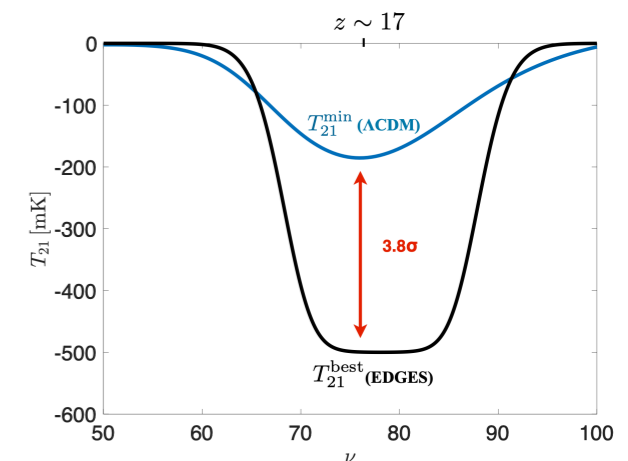
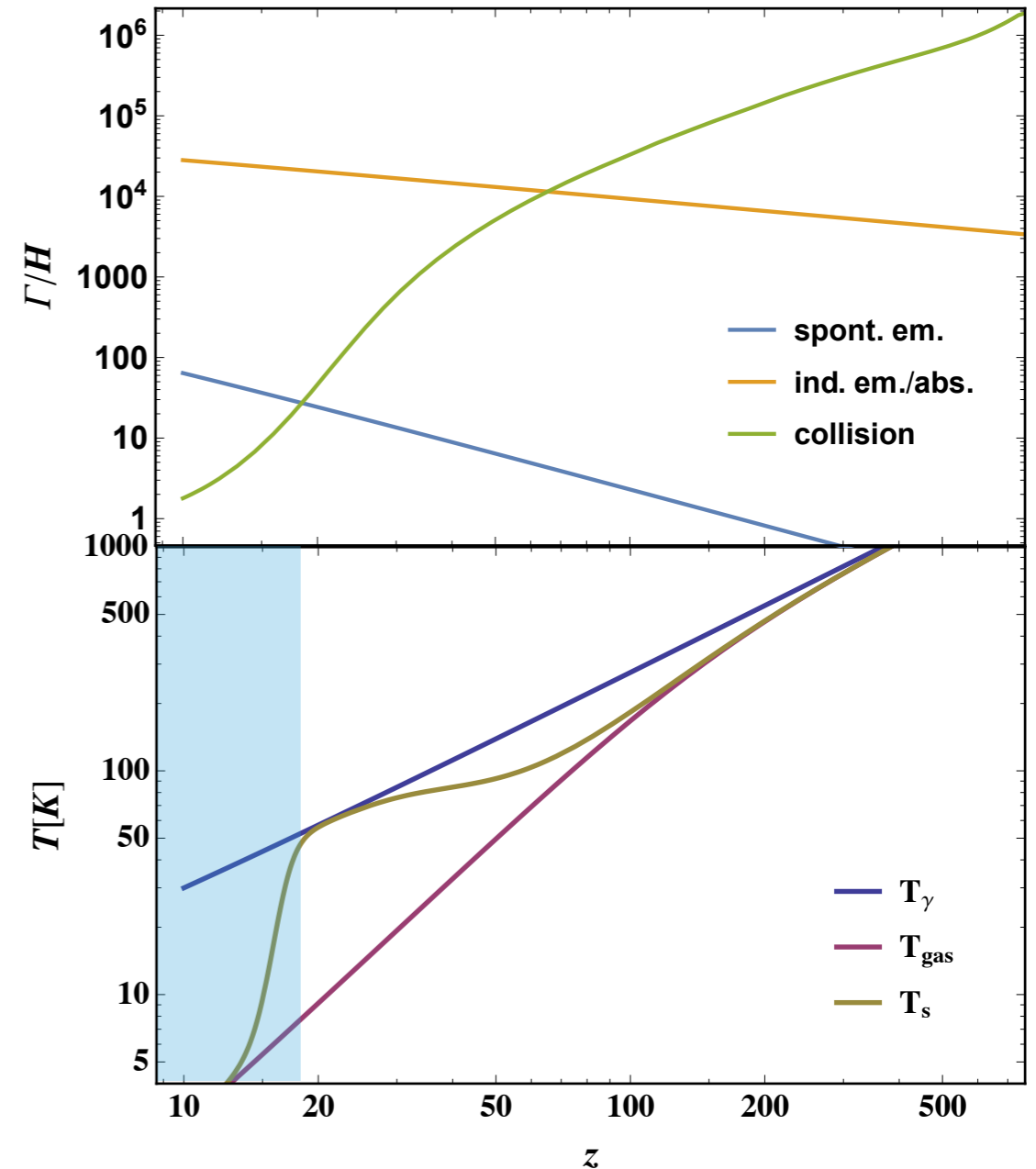
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# BASICS

## Evolution - Boltzmann

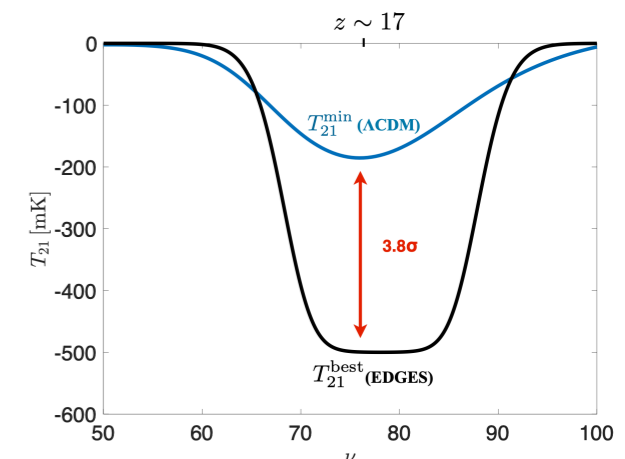
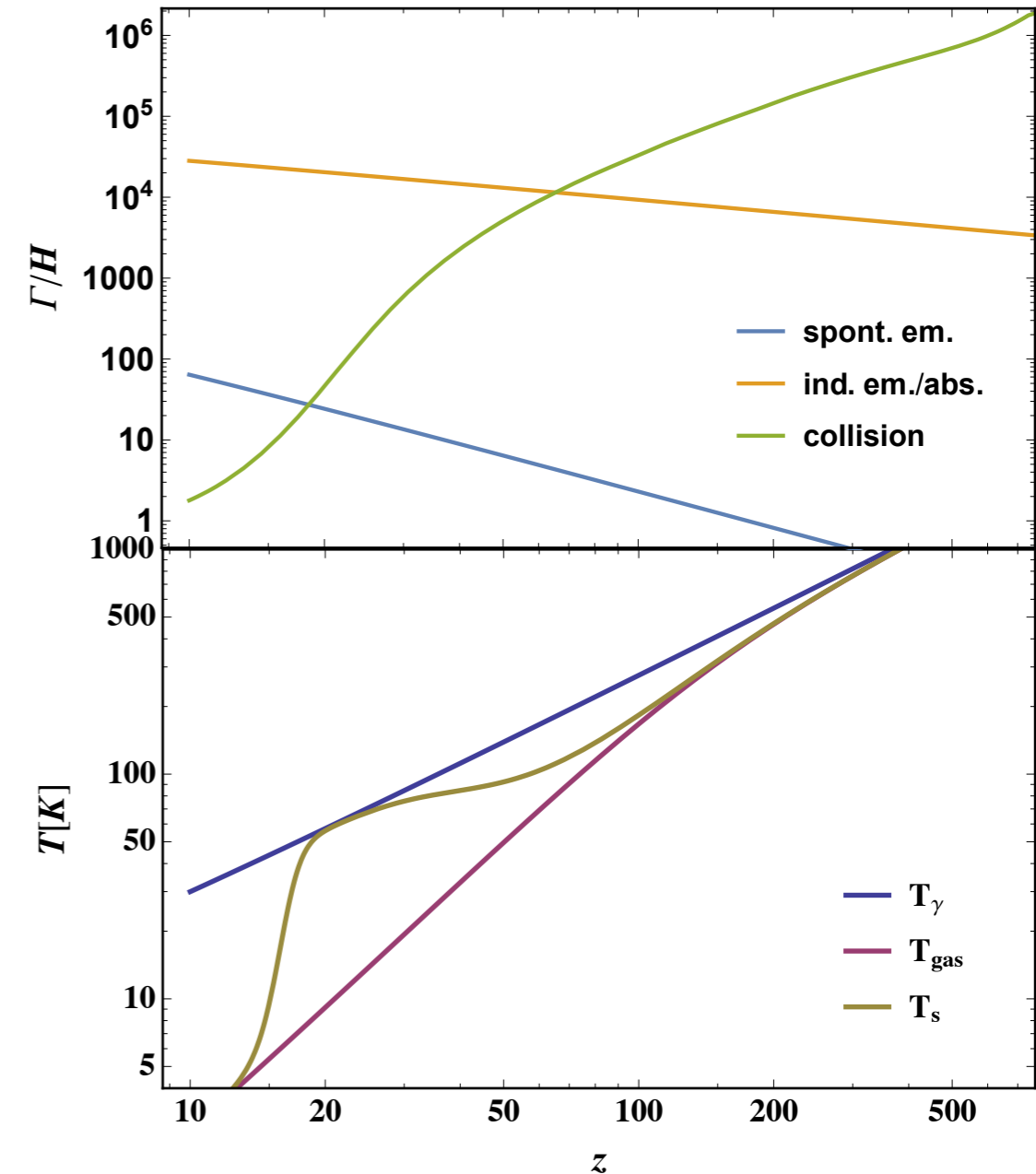
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$$\Delta T = T - T_{\text{CMB}}$$

- ▶ We measure the brightness temperature:

$$T_{21} \simeq (T_s - T_{\text{CMB}}) \frac{1 - e^{-\tau}}{1 + z} \geq (T_{\text{gas}} - T_{\text{CMB}}) \frac{1 - e^{-\tau}}{1 + z} \equiv T_{21}^{\text{min}}$$



# THE EDGES ANOMALY

## Evolution - Boltzmann

▶ Dominant

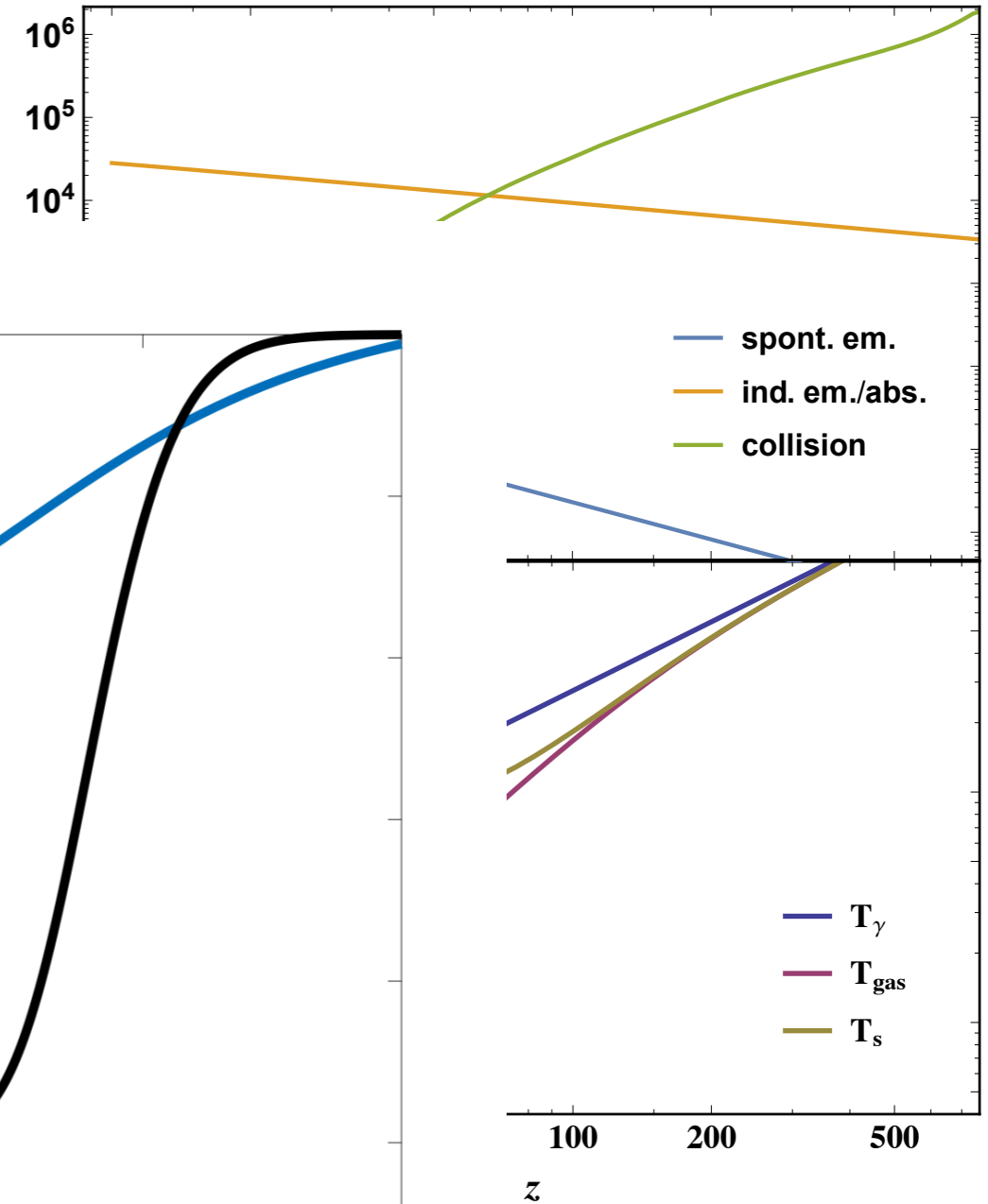
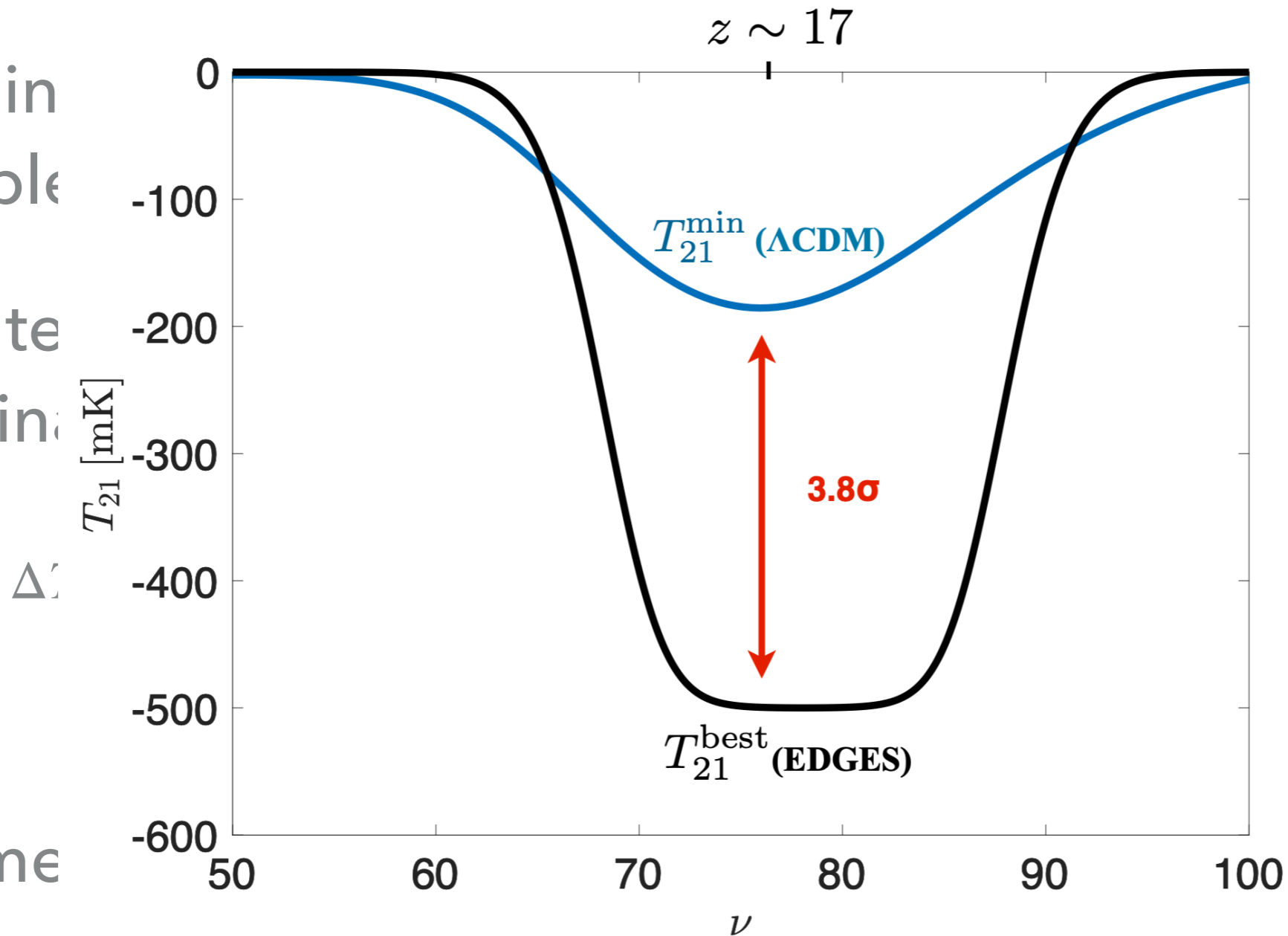
Hubble

▶ Spin temperature

dominant

▶ We measure

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# THE EDGES ANOMALY

Translated to baryons temperature

$$T_b^{\text{EDGES}} \Big|_{z=17} = 3.2^{+1.9}_{-1.55} \text{K}$$

$$\Delta T_b^{\text{EDGES}} \equiv (T_b^{\text{EDGES}} - T_b^{\Lambda\text{CDM}}) \Big|_{z=17} = -3.6^{+1.9}_{-1.55} \text{K}$$

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### ► Heat the CMB

- Radio loud quasars (Feng, Holder, 2018)
- Resonant conversion of hidden photons (Pospelov, Pradler, Ruderman, Urbano, 2018)
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### ▶ Cool the gas

- Earlier photon - baryon decoupling (Falkowski, Petraki)
- DM - baryon scattering (Kadota, Silk, Tashiro, 2014; Ali-Haimoud, Kovetz, Munoz, 2015; Barkana, 2018; Loeb, Munoz, 2018; Berlin, Hooper, Krnjaic, McDermott, 2018; Barkana, **NJO**, Redigolo, Volansky, 2018; Liu, Slatyer 2018; Kovetz, Poulin, Gluscevic, Boddy, Barkana, Kamionkowski, 2018;...)
- ...

▶ ...

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- **DM - BARYON SCATTERING** (Kadota, Silk, Tashiro, 2014; Ali-Haimoud, Kovetz, Munoz, 2015; Barkana, 2018; Loeb, Munoz, 2018; Berlin, Hooper, Krnjaic, McDermott, 2018; Barkana, **NJO**, Redigolo, Volansky, 2018; Liu, Slatyer 2018; Kovetz, Poulin, Gluscevic, Boddy, Barkana, Kamionkowski, 2018;...)
- ...

▶ ...



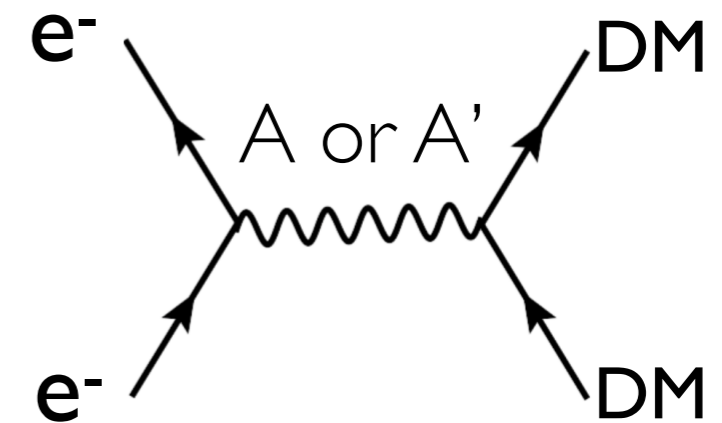
BARKANA, NJO, REDIGOLO,  
VOLANSKY.  
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**MILLICHARGED DARK  
MATTER @ 21-CM**

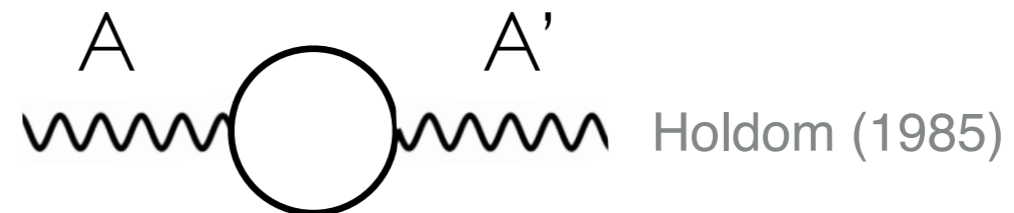
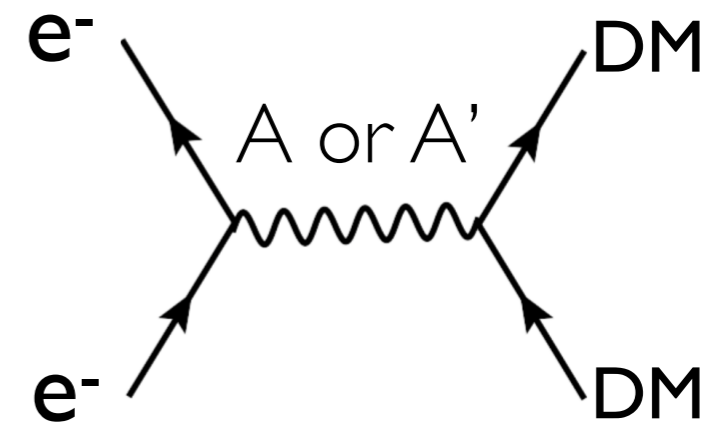
## MILLICHARGED DARK MATTER

- ▶ DM carries as small electric charge
- ▶ Two classes
  - ◉ “Pure” millicharge



# MILLICHARGED DARK MATTER

- ▶ DM carries as small electric charge
- ▶ Two classes
  - ◉ "Pure" millicharge
  - ◉ Hidden photon portal



$$\mathcal{L} = -\frac{1}{4}F^2 - \frac{1}{4}F'^2 - \frac{\epsilon}{2}FF' + m_{A'}^2 A'^2 + J_{\text{SM}}A + J_{\text{DM}}A'$$

$$\rightarrow -\frac{1}{4}F_1^2 - \frac{1}{4}F_2^2 + m_{A'}^2 A_2^2 + J_{\text{SM}}A_1 + (J_{\text{DM}} - \epsilon J_{\text{SM}}) A_2$$

## DM – BARYON SCATTERING

Gas temperature evolution is modified

$$\frac{dT_{\text{gas}}}{d \log a} = -2T_{\text{gas}} + 2\frac{\Gamma_C}{H} (T_{\text{CMB}} - T_{\text{gas}}) + \frac{3}{2} \sum_{I=\{\text{H, He, } p, e\}} \frac{\dot{Q}_{\text{gas}}^I}{H}$$

## DM – BARYON SCATTERING

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$$\frac{dT_{\text{gas}}}{d \log a} = \overbrace{-2T_{\text{gas}} + 2\frac{\Gamma_C}{H} (T_{\text{CMB}} - T_{\text{gas}})}^{\text{SM}} + \frac{3}{2} \sum_{I=\{\text{H, He, p, e}\}} \frac{\dot{Q}_{\text{gas}}^I}{H}$$

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Adiabatic



# DM – BARYON SCATTERING

Gas temperature evolution is modified

$$\frac{dT_{\text{gas}}}{d \log a} = \underbrace{-2T_{\text{gas}}}_{\text{Adiabatic}} + \underbrace{2\frac{\Gamma_C}{H} (T_{\text{CMB}} - T_{\text{gas}})}_{\text{Compton}} + \frac{3}{2} \sum_{I=\{\text{H, He, p, e}\}} \frac{\dot{Q}_{\text{gas}}^I}{H}$$

SM

## DM – BARYON SCATTERING

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$\dot{Q}_{\text{gas}}^I \sim$  Energy (temperature) transfer + momentum transfer (drag)

# DM – BARYON SCATTERING

Gas temperature evolution is modified

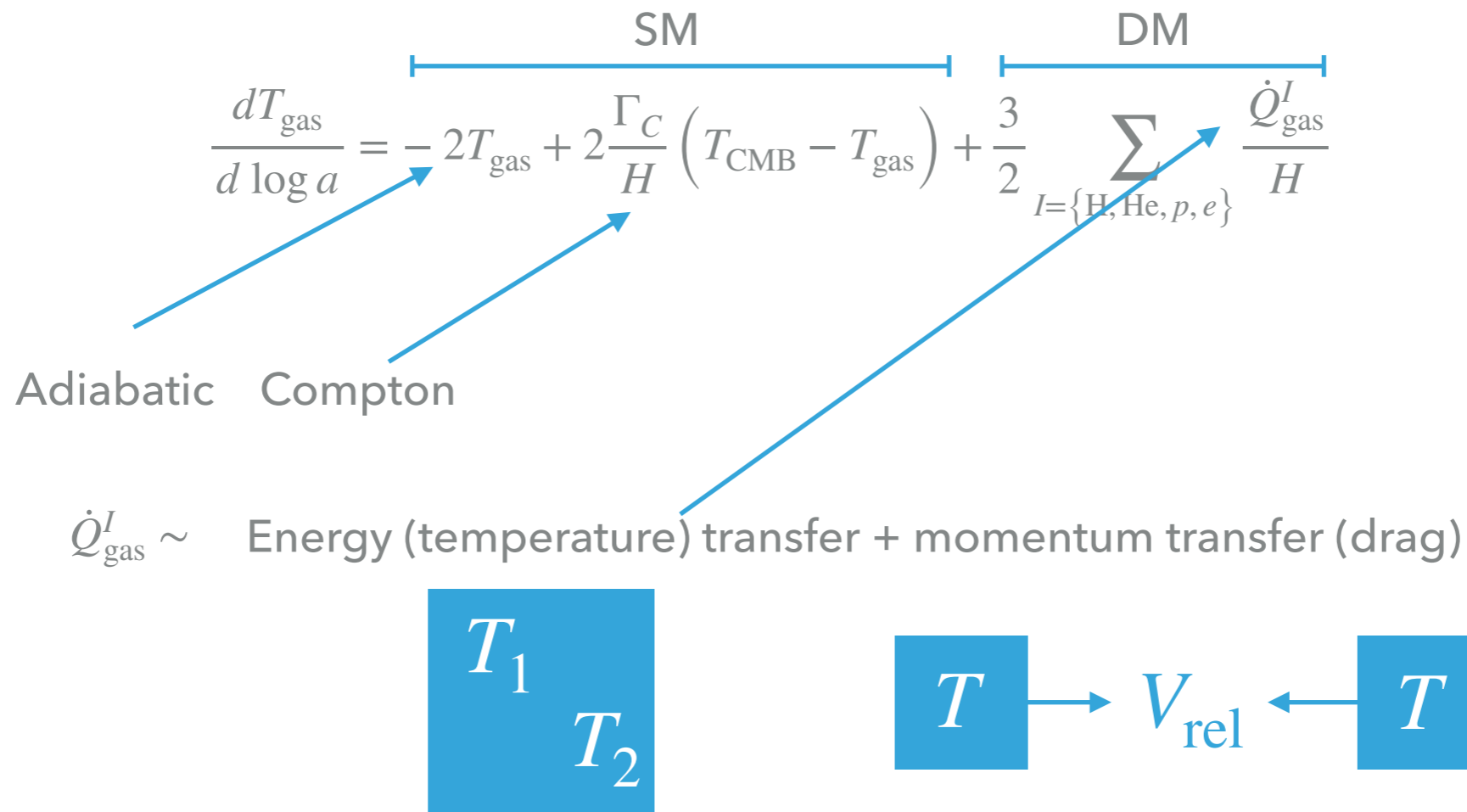
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$T_1$   
 $T_2$

# DM - BARYON SCATTERING

Gas temperature evolution is modified



## DM - BARYON SCATTERING

Gas temperature evolution is modified

$$\frac{dT_{\text{gas}}}{d \log a} = \underbrace{-2T_{\text{gas}}}_{\text{Adiabatic}} + \underbrace{2\frac{\Gamma_C}{H} (T_{\text{CMB}} - T_{\text{gas}})}_{\text{Compton}} + \underbrace{\frac{3}{2} \sum_{I=\{\text{H, He, p, e}\}} \frac{\dot{Q}_{\text{gas}}^I}{H}}_{\text{DM}}$$

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### Requirements:

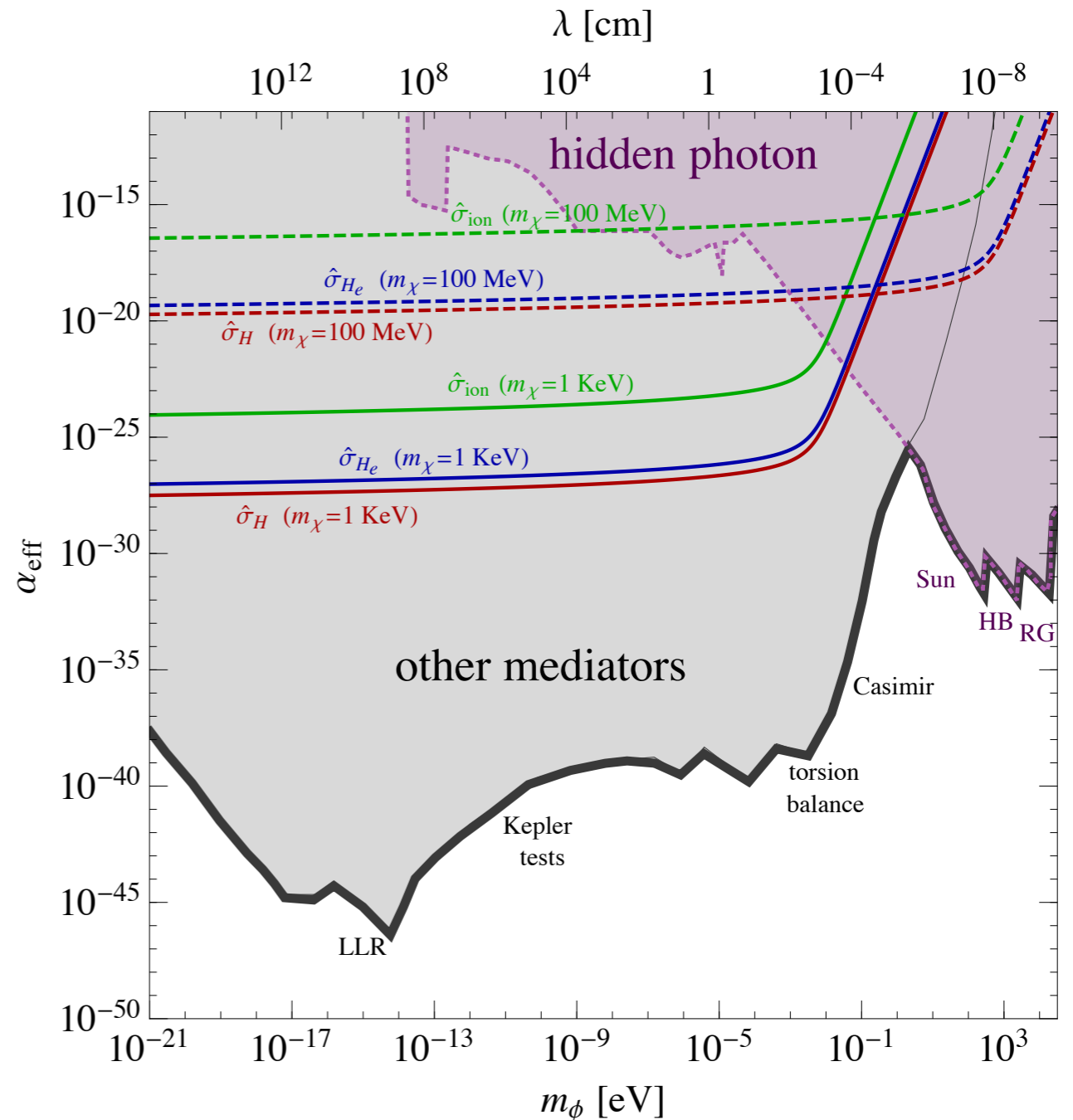
- CMB anisotropies constraints - **small cross section at early times**
- Faster than Hubble - **large cross section at later times**
- IR dominated process - **velocity dependence, light mediator** ( $\sim v_{\text{rel}}^{-4}$ )

# DM - BARYON SCATTERING

Velocity dependence → long range force (light mediator)

## SCREENED OR UNSCREENED?

- ▶ 5th force experiments rules out unscreened force.
- ▶ DM is (effectively?) millicharged
- ▶ Atoms are neutral
- ▶ Cooling only by interacting with electrons and protons



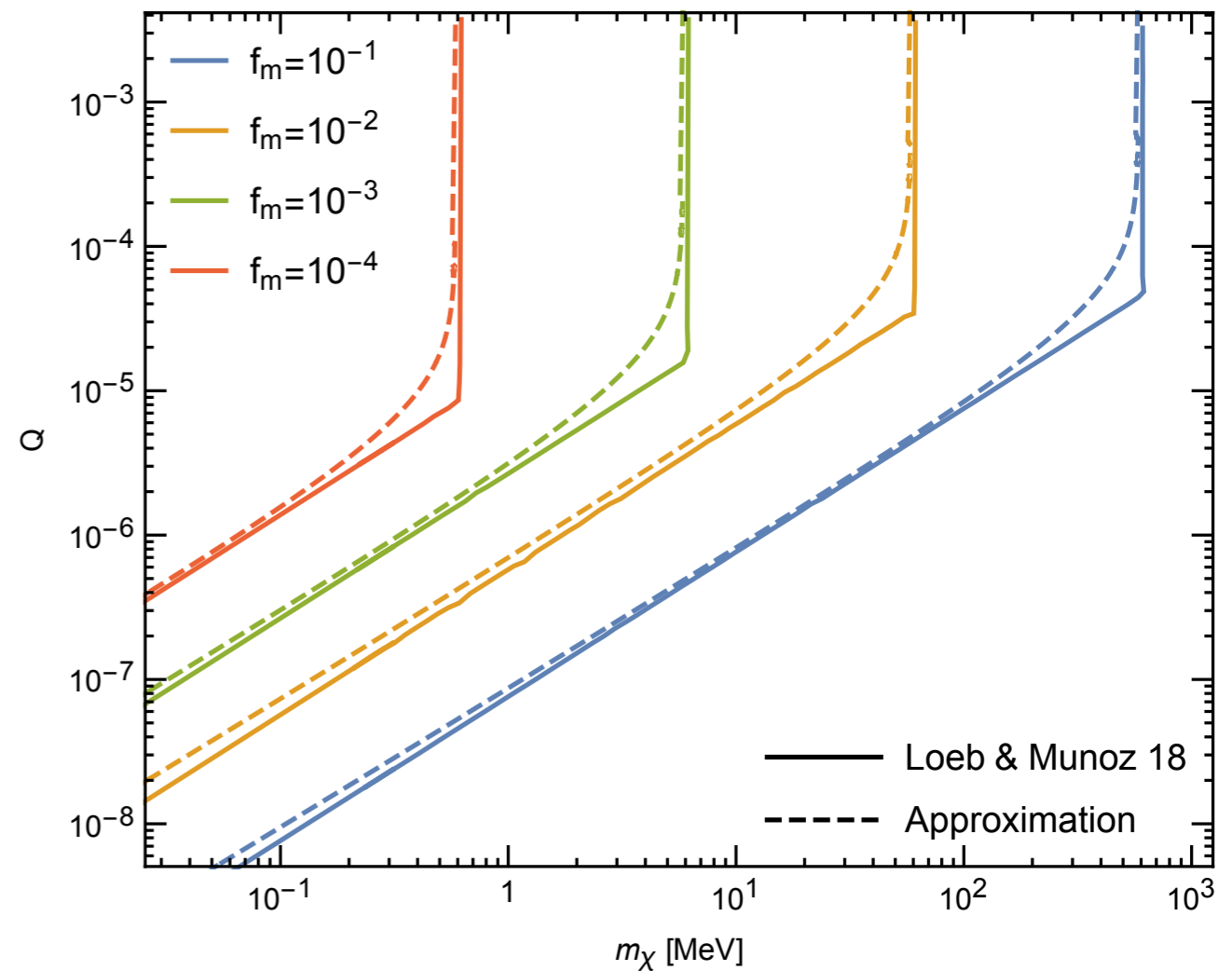
(Barkana, **NJO**, Redigolo, Volansky, 2018)

# ANALYTICS

Very cool estimates (ignoring bulk motion):

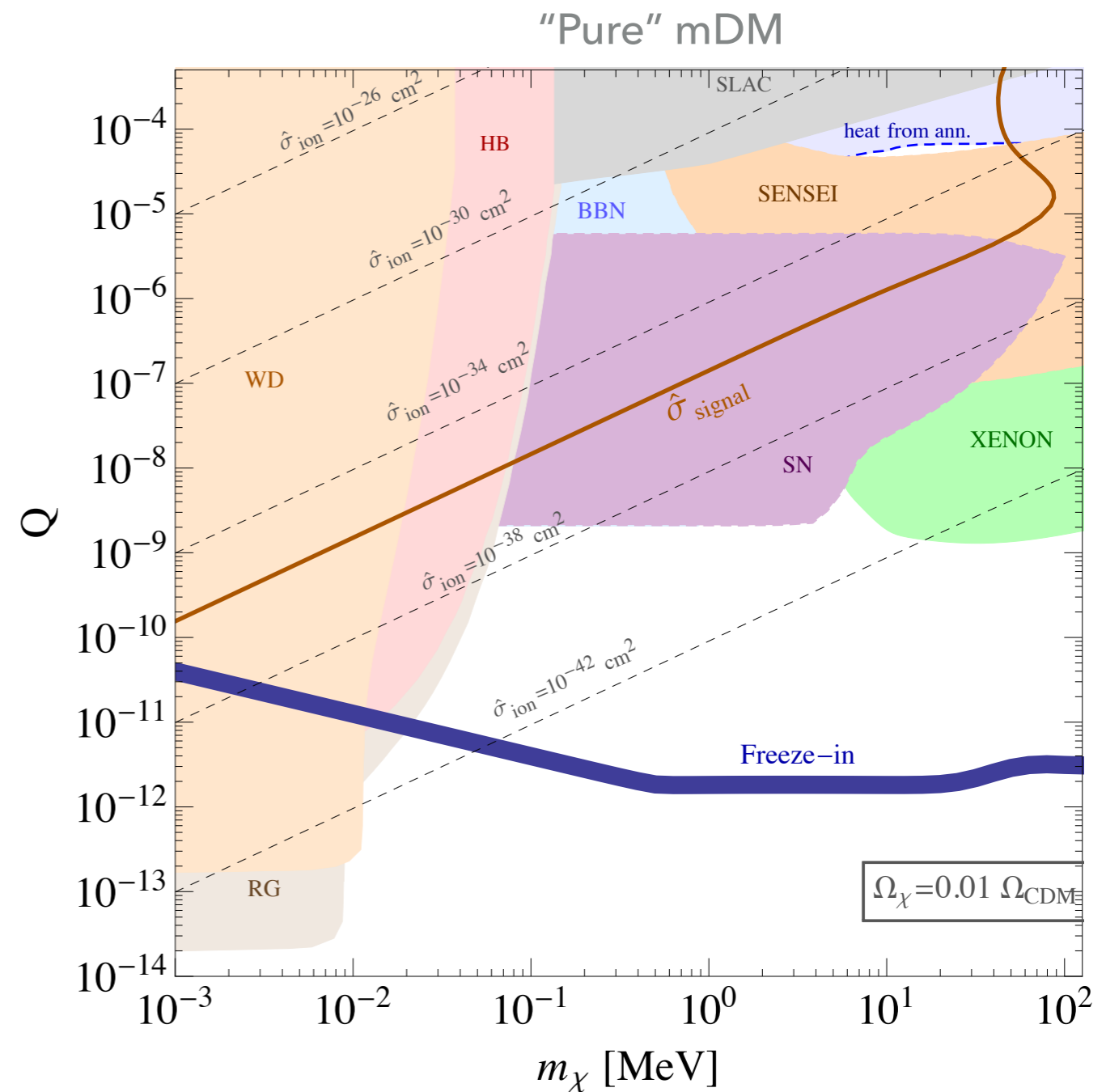
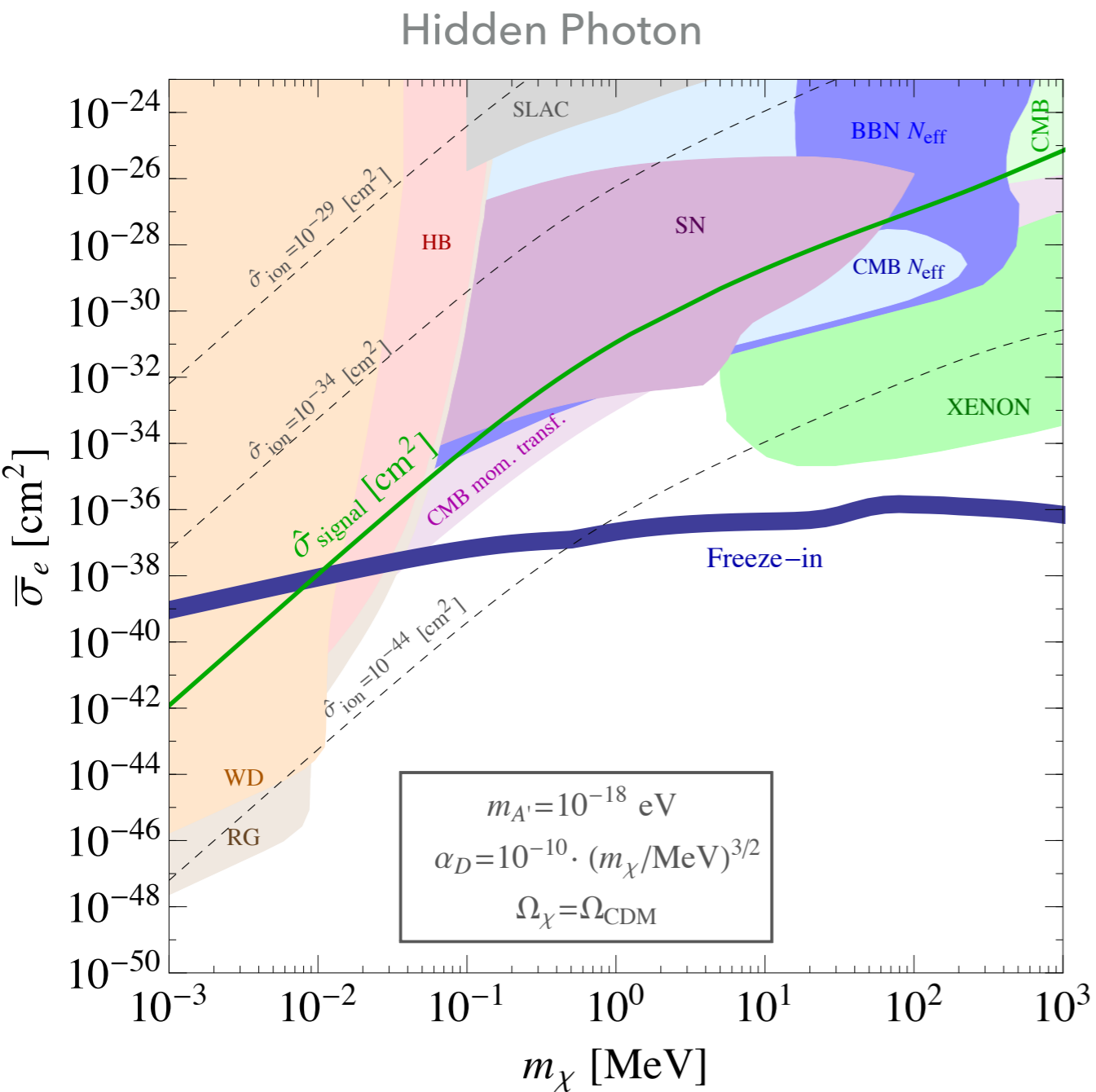
$$m_{\text{mDM}} \lesssim f_{\text{m}} \times 6 \text{ GeV}$$

$$\frac{f_{\text{m}} \rho_{\text{DM}}}{m_{\text{m}}} \frac{x_e \mu_{\text{pm}}}{m_p + m_{\text{m}}} \frac{\langle \sigma_T^{\text{pm}} v_{\text{rel}}^2 \rangle}{\frac{T_{\text{m}}}{m_{\text{m}}} + \frac{T_{\text{b}}}{m_p}} (T_{\text{b}} - T_{\text{m}}) \simeq HT_{\text{b}}$$



# PARAMETER SPACE

Screened force  $\rightarrow$  millicharged DM (mDM)

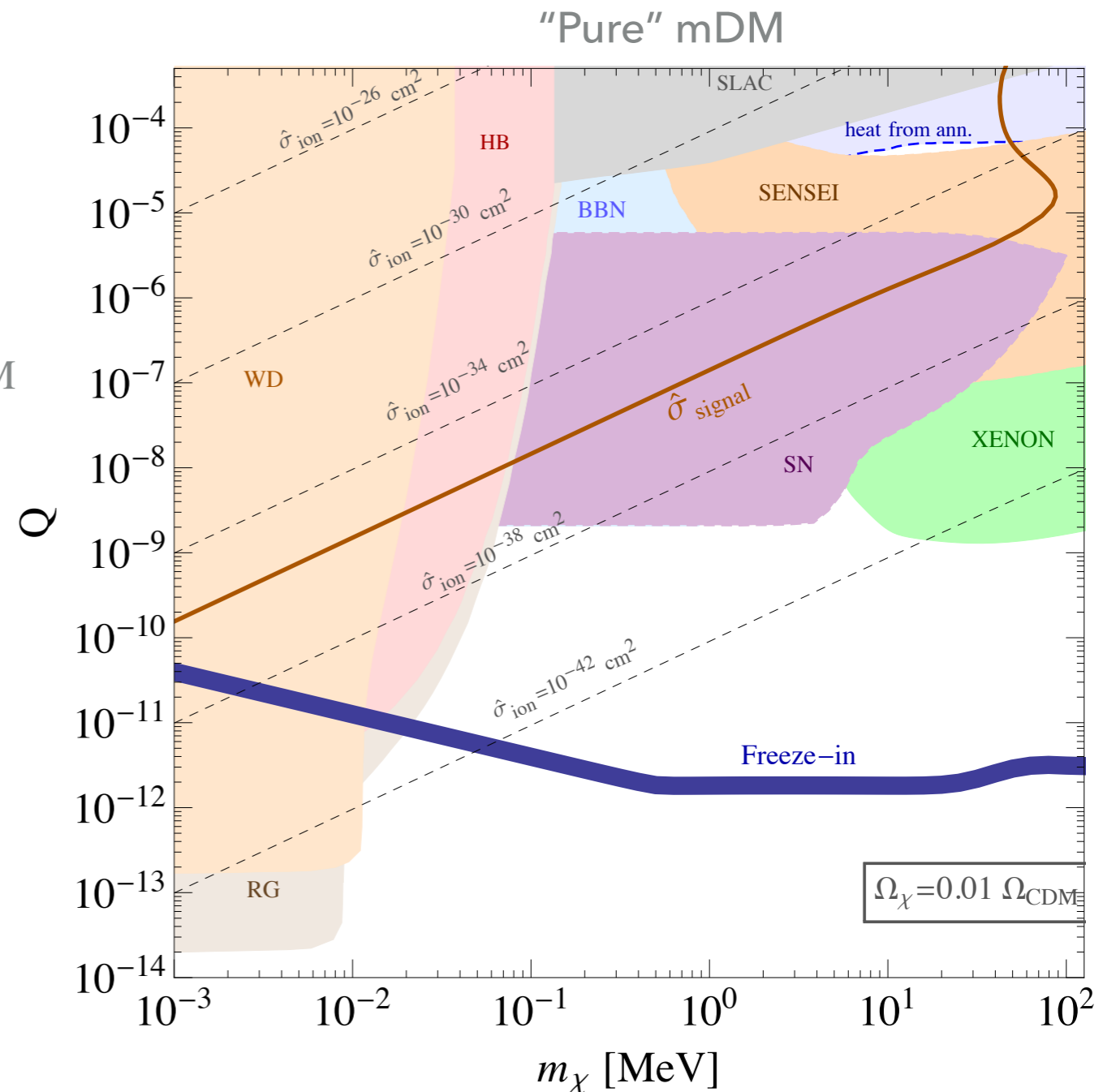
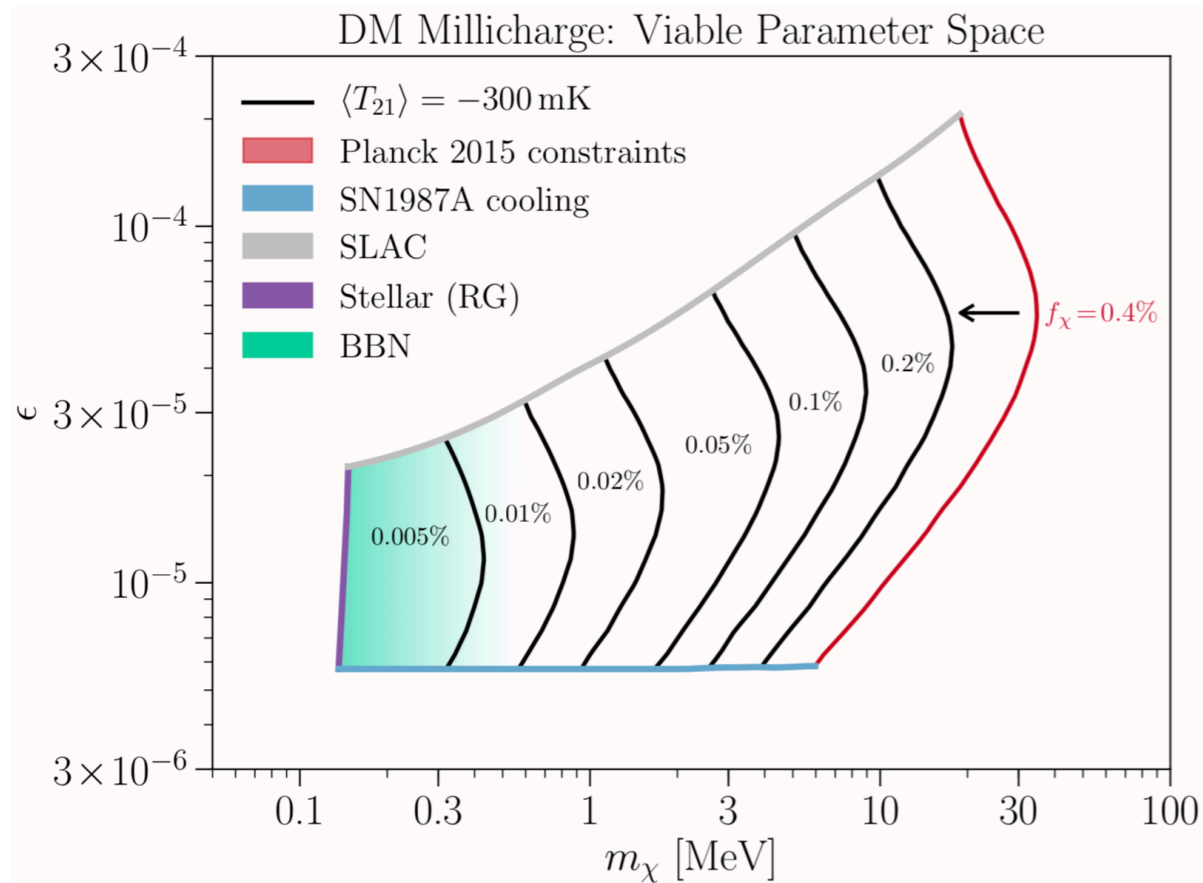




# PARAMETER SPACE

## STATUS

- ▶ E.D.Kovetz et.al (2018) revisited CMB constraints using Planck 2015 concluding  $\Omega_{\text{mDM}} \lesssim 4 \times 10^{-3} \Omega_{\text{DM}}$

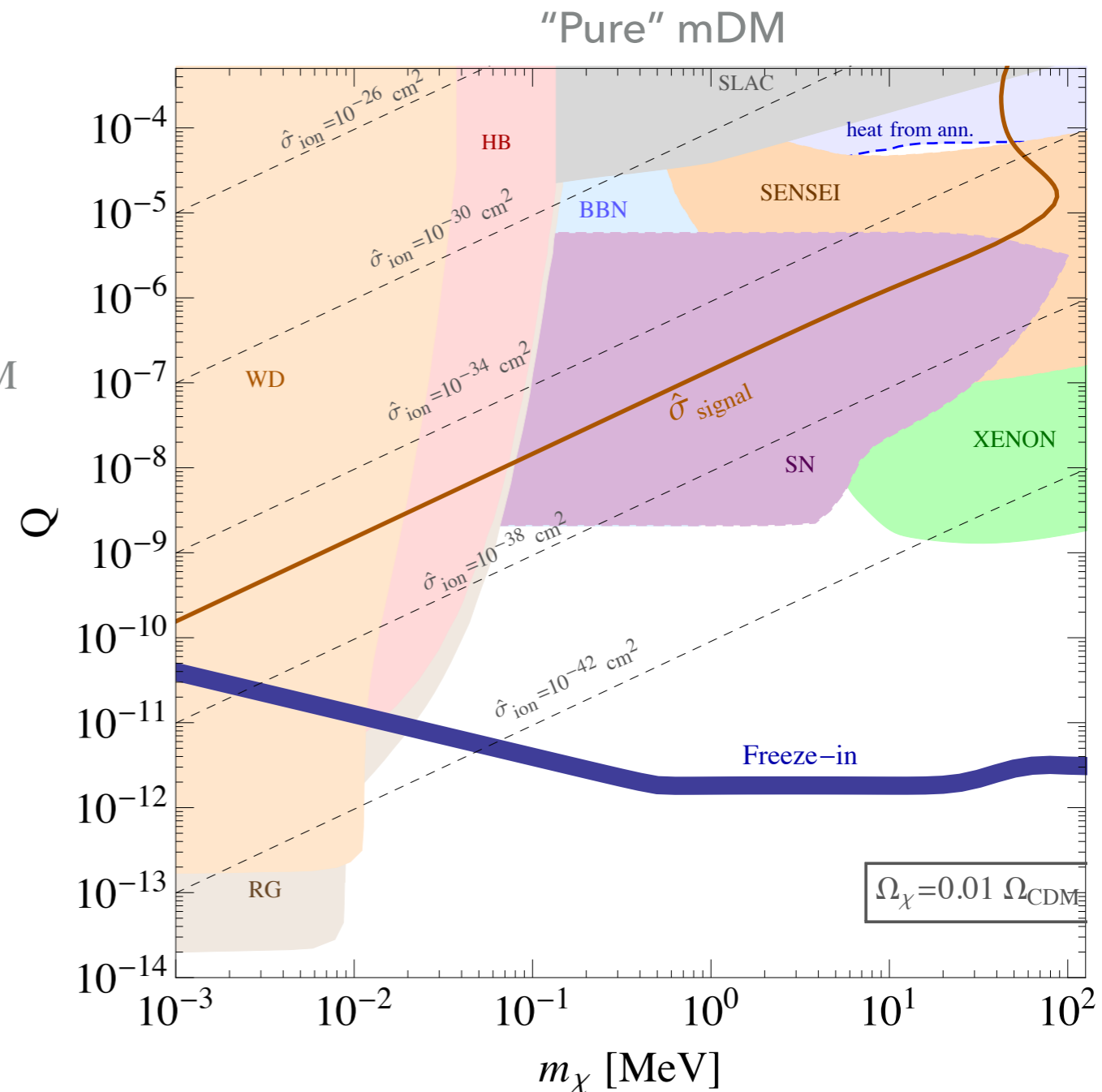


(Barkana, NJO, Redigolo, Volasky, 2018)

# PARAMETER SPACE

## STATUS

- ▶ E.D.Kovetz et.al (2018) revisited CMB constraints using Planck 2015 concluding  $\Omega_{\text{mDM}} \lesssim 4 \times 10^{-3} \Omega_{\text{DM}}$
- ▶ Together with  $N_{\text{eff}}$  constraints, hidden photon explanation is excluded.
- ▶ Probably also without it.
- ▶ The end?



(Barkana, **NJO**, Redigolo, Volasky, 2018)



LIU, NJO, REDIGOLO,  
VOLANSKY,

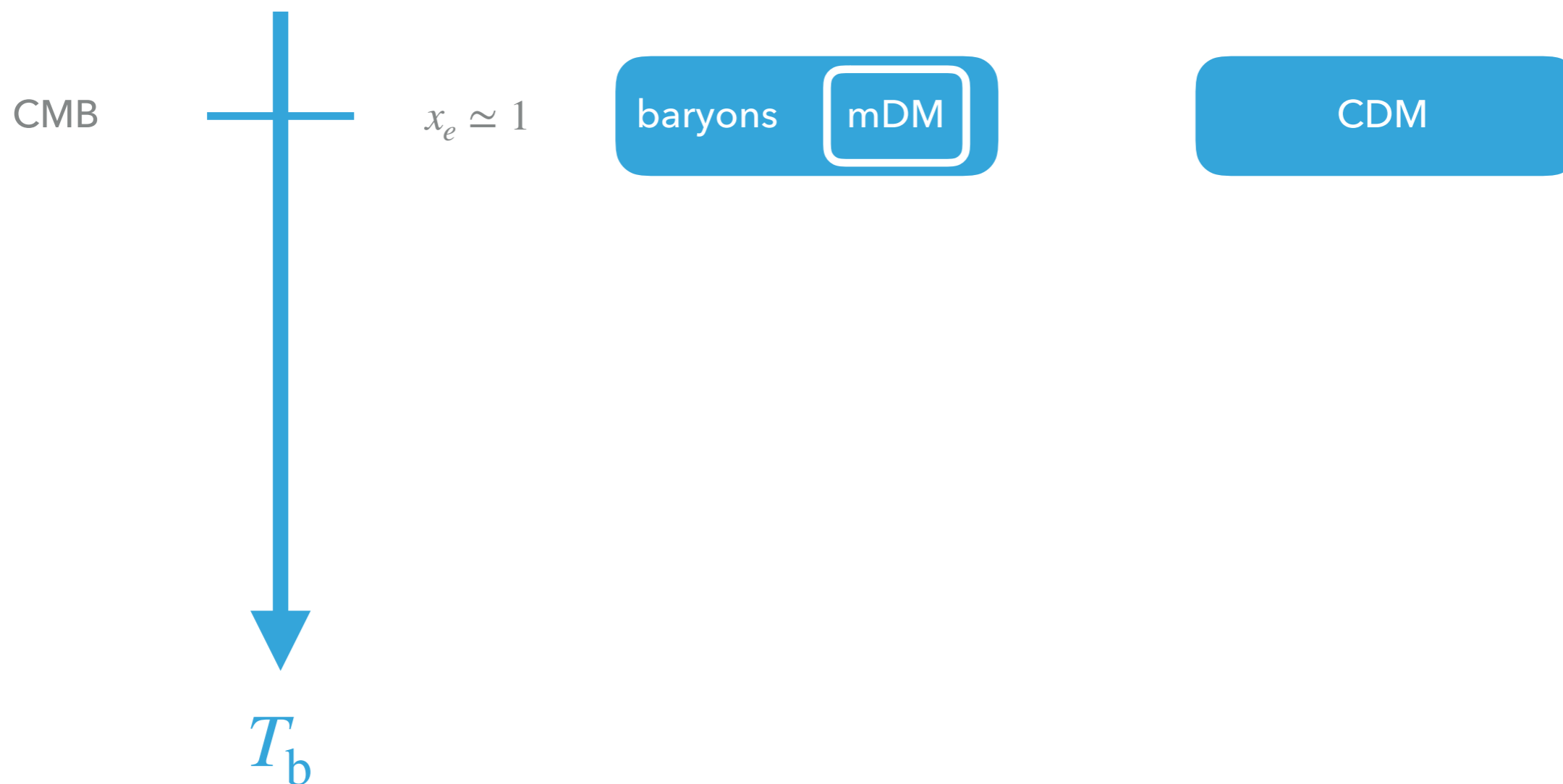
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**REVIVING  
MILLICHARGED DARK  
MATTER**

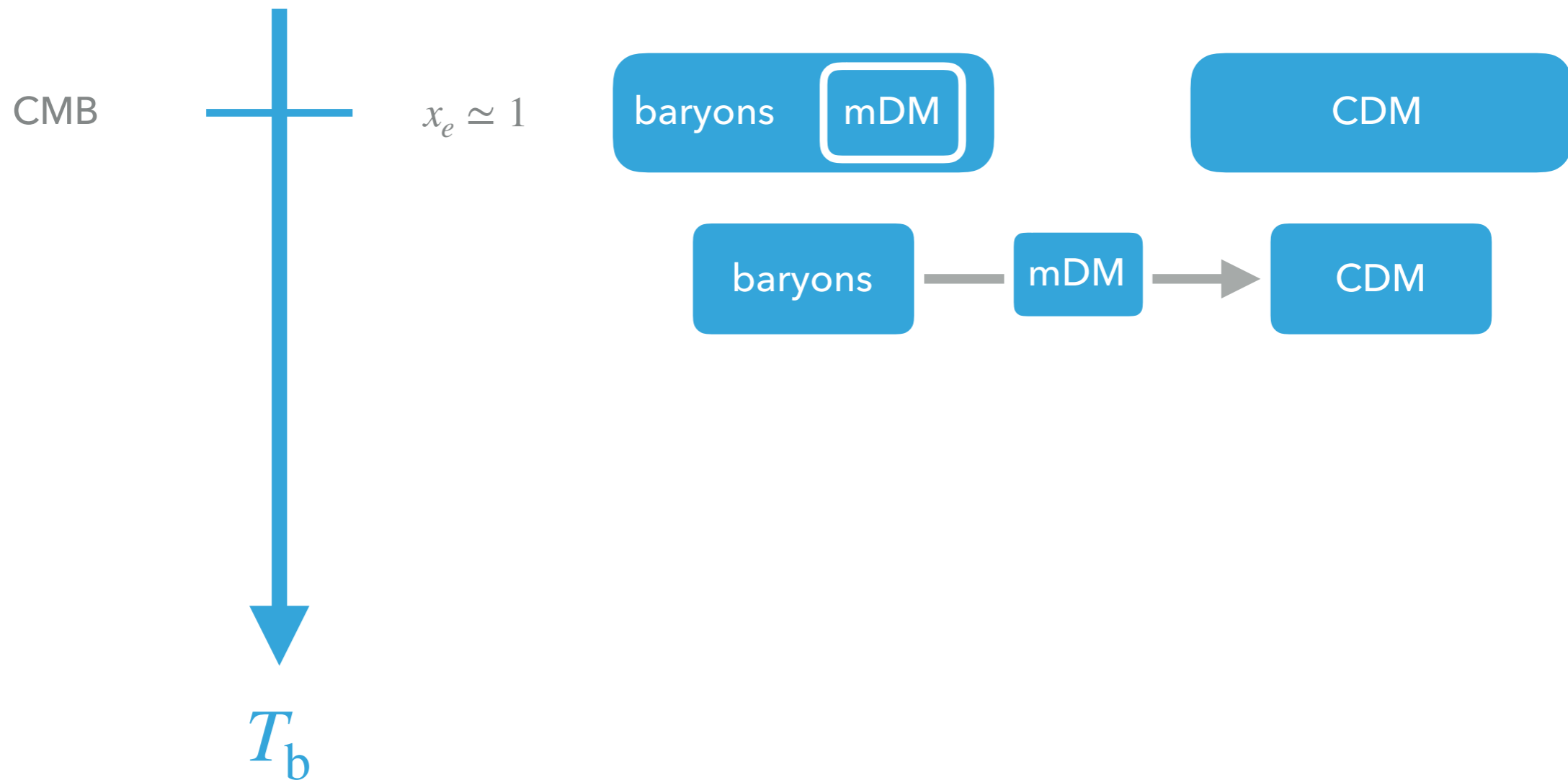
## FRAMEWORK

General idea: use the CDM bath as a reservoir.



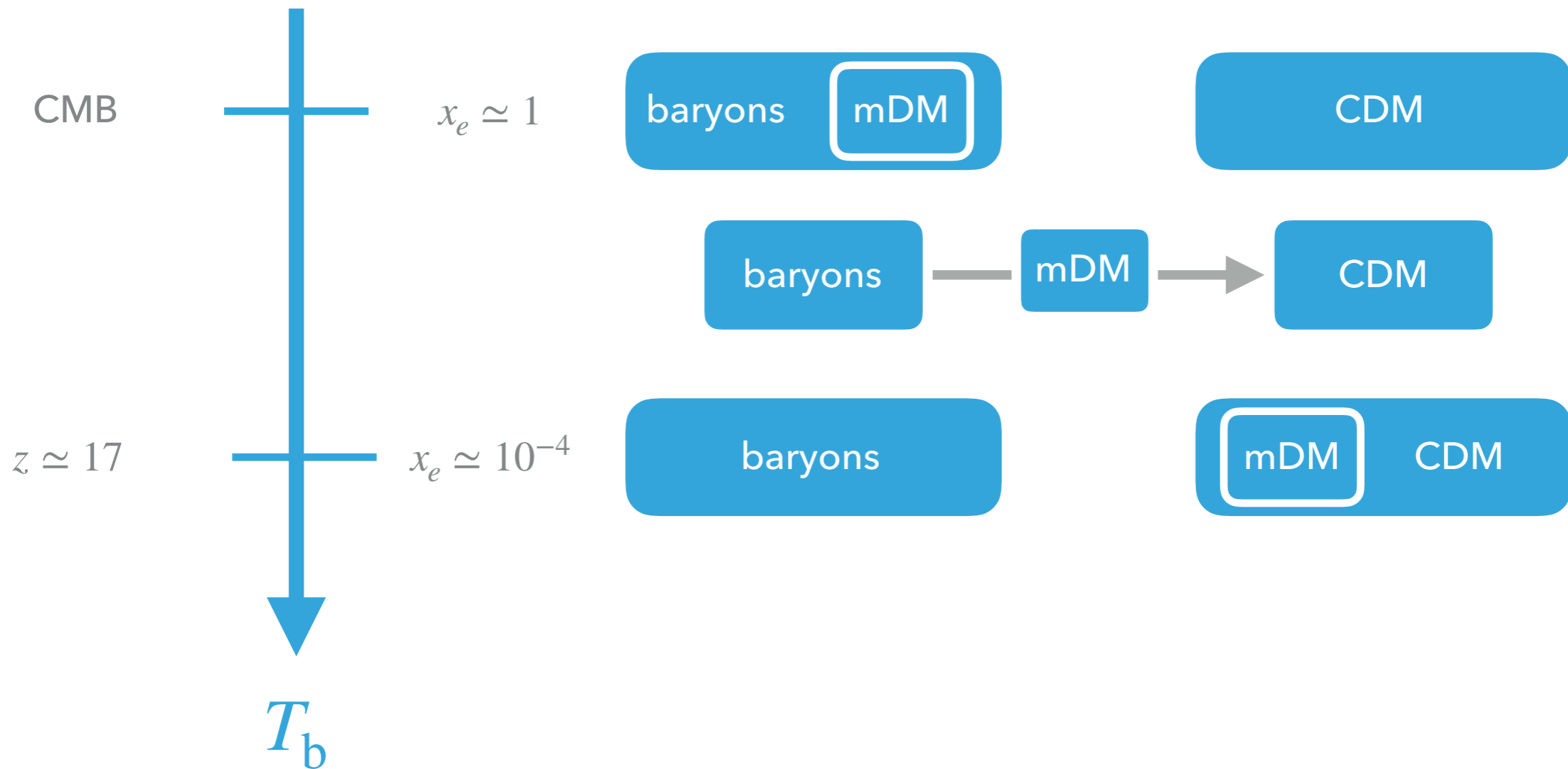
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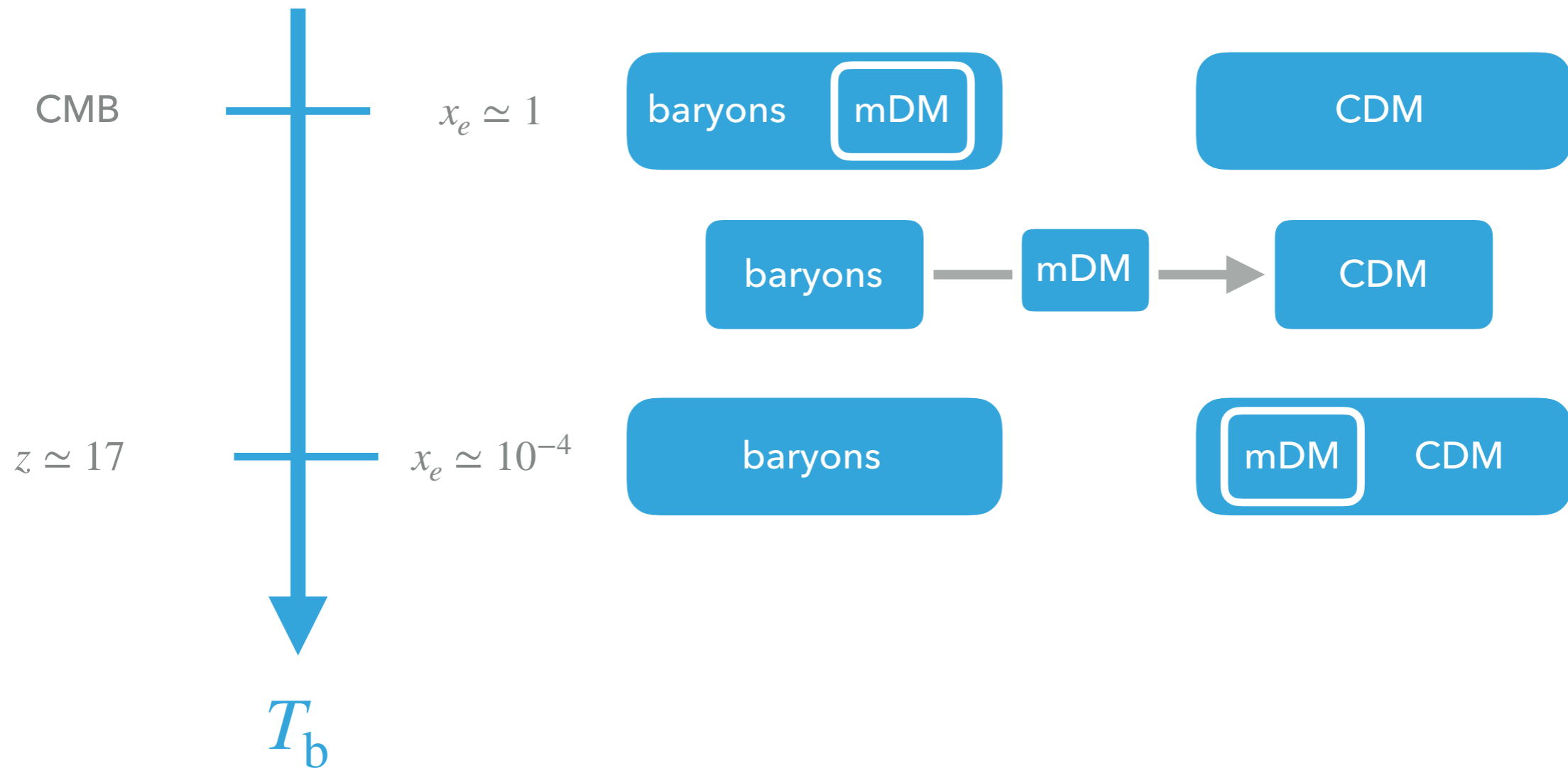
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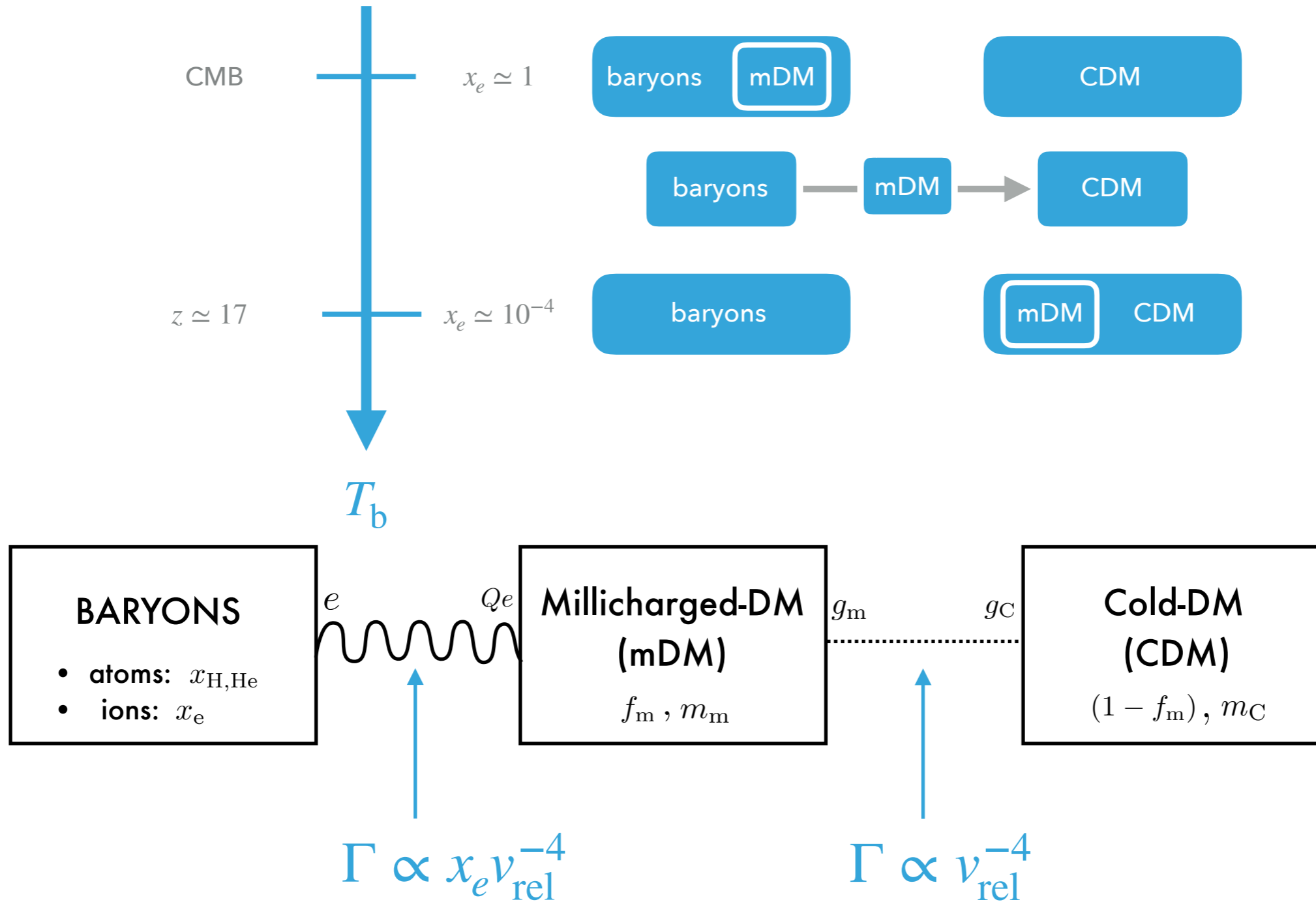
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General idea: use the CDM bath as a reservoir.



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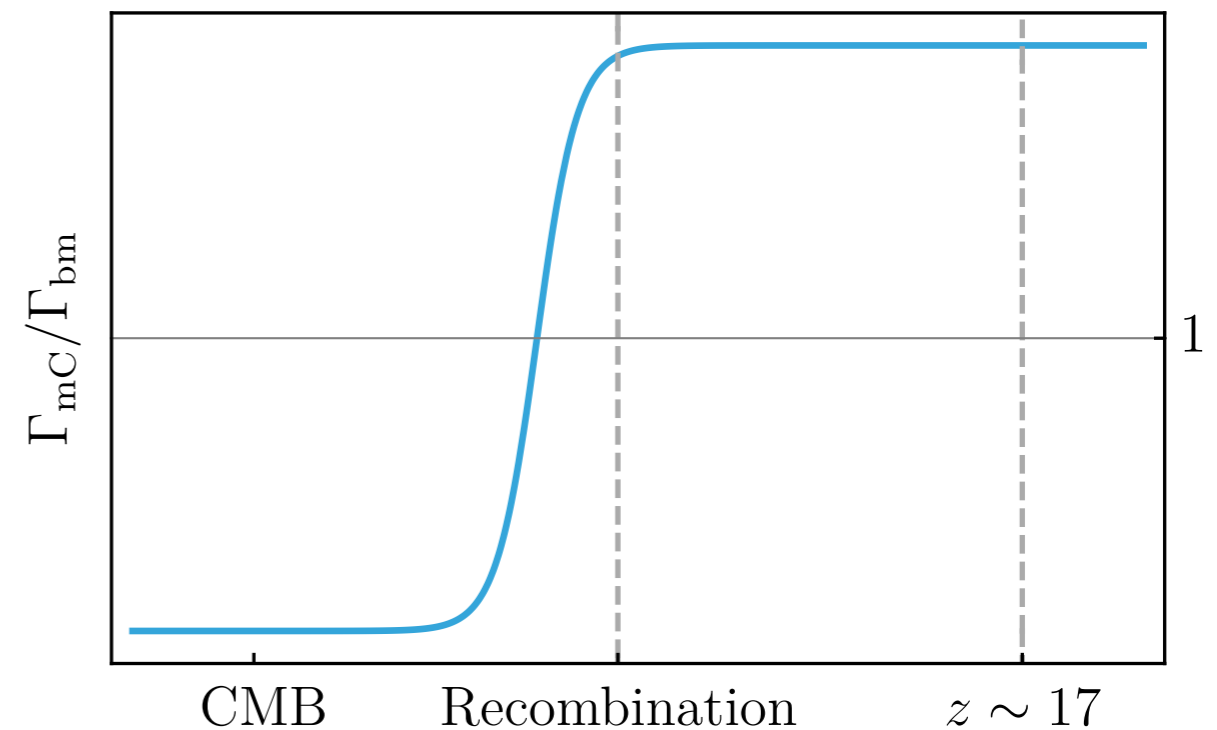
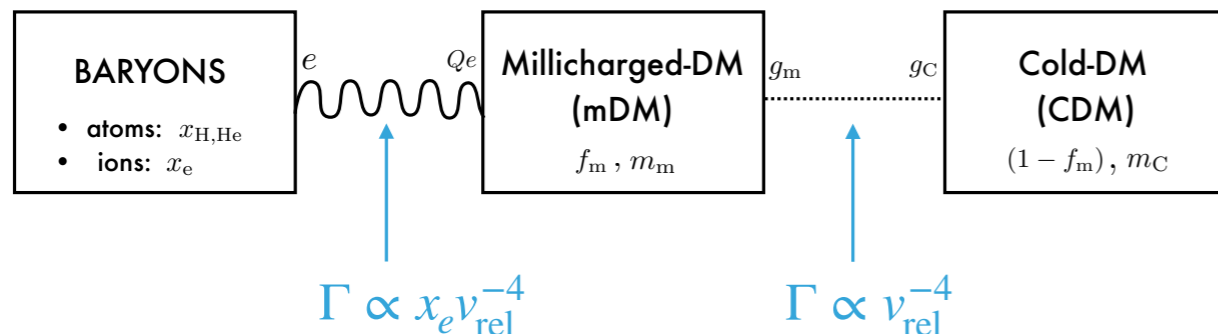
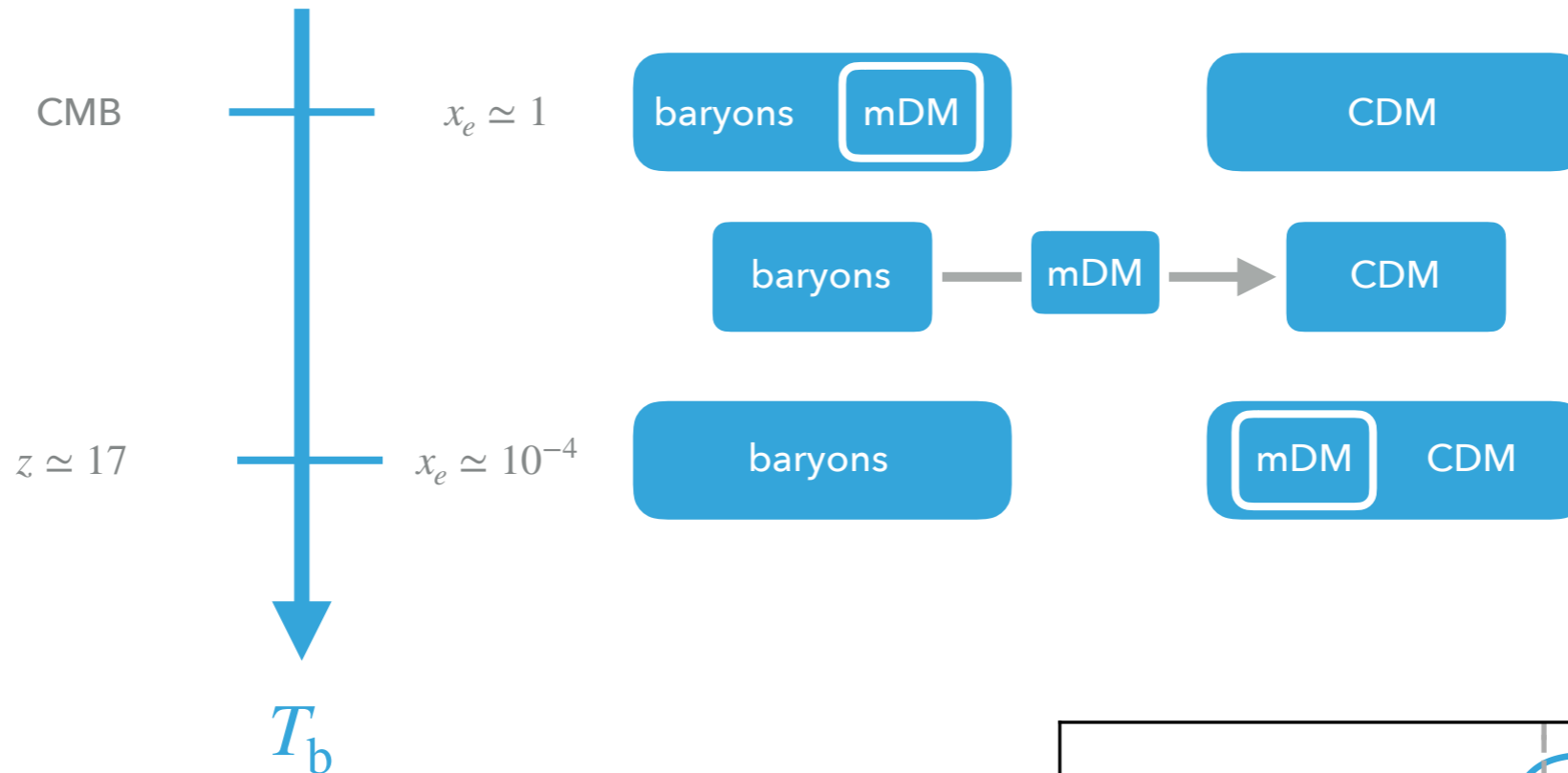
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# FRAMEWORK

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## FRAMEWORK

### Implications

- CMB anisotropies constraints
  - ▶ **hide mDM in the baryonic fluid,  $f_m \lesssim 10^{-4}$**
  - ▶ **large mDM-SM cross section at early times**
  - ▶ **small mDM-CDM cross section at early times**
- EDGES at  $z \simeq 17$  - **large mDM-CDM cross section at later times**
- IR dominated processes - **velocity dependence, light mediator** ( $\sim v_{\text{rel}}^{-4}$ )
- Relative interactions strength is set by  $x_e$

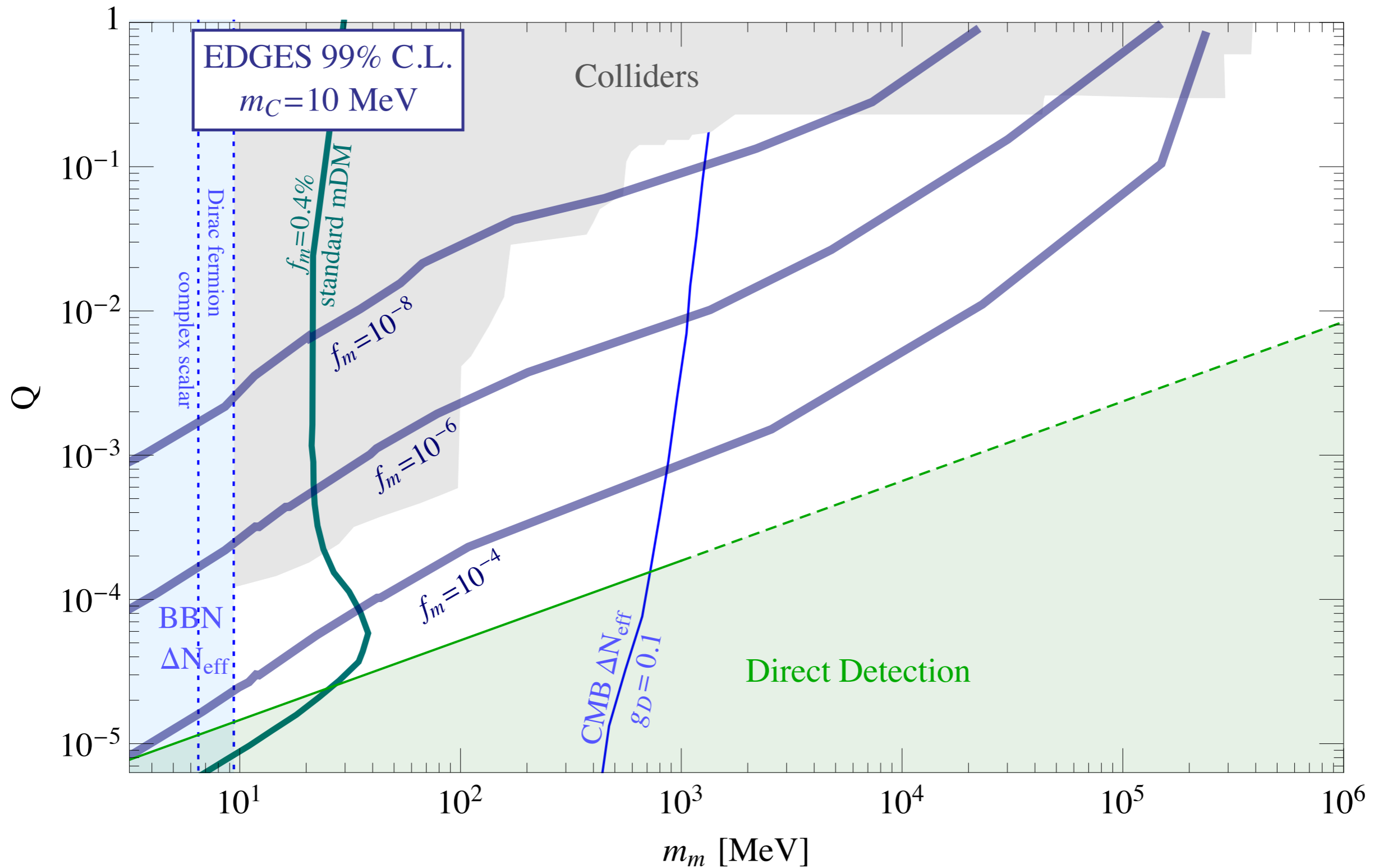
## FRAMEWORK

### What did we gain?

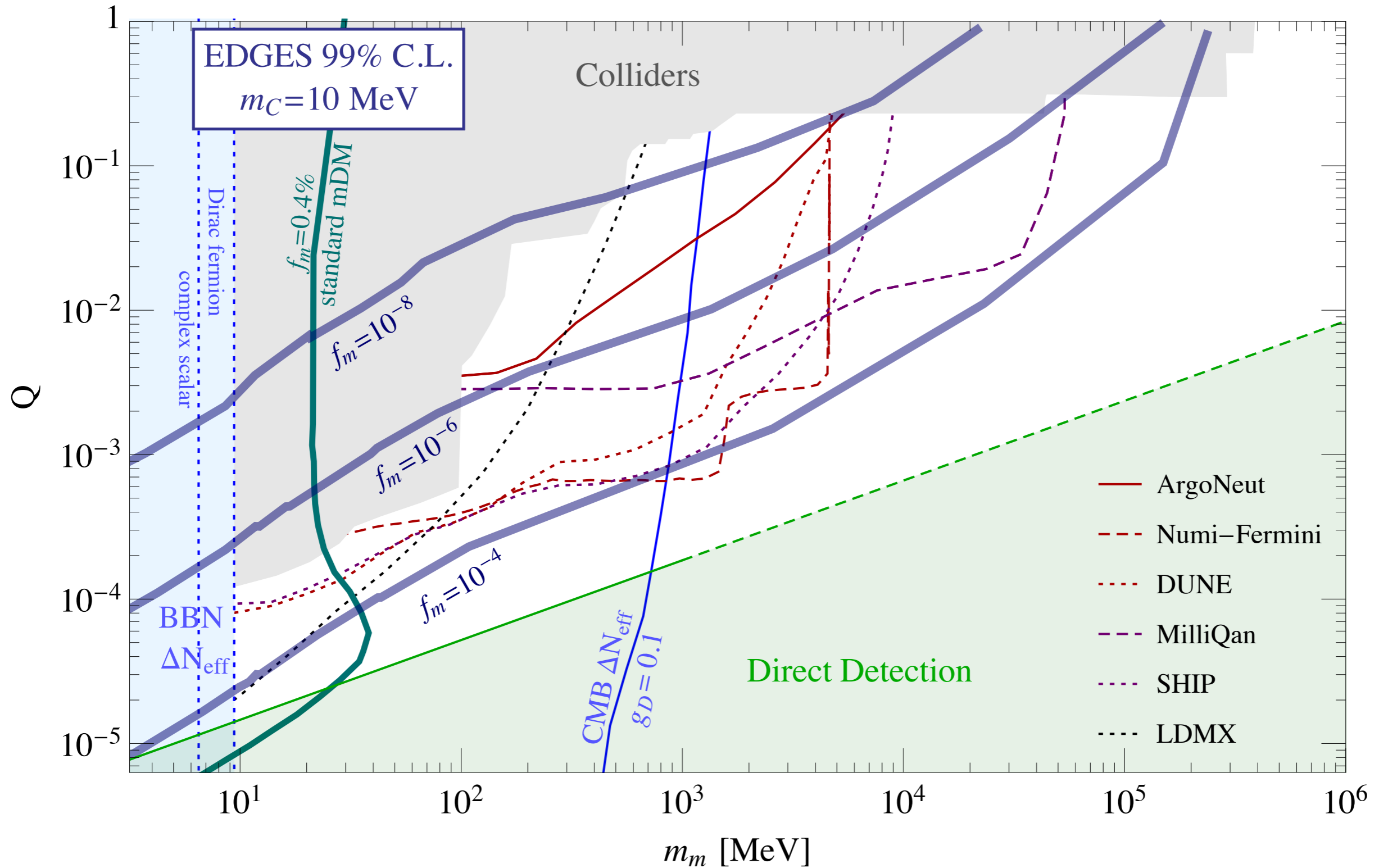
- Upper bound on CDM mass  $m_C \lesssim 6 \text{ GeV}$
- Upper bound on mDM mass\*  $m_m \lesssim 200 \text{ GeV}$
- Lower bound on the mDM fraction is significantly smaller  $f_m \gtrsim 10^{-8}$
- Model lives below colliders reach, but above direct detection reach
- Many prediction for near future experiments

\* a complicated combination of CMB constraints + the required rate for EDGES

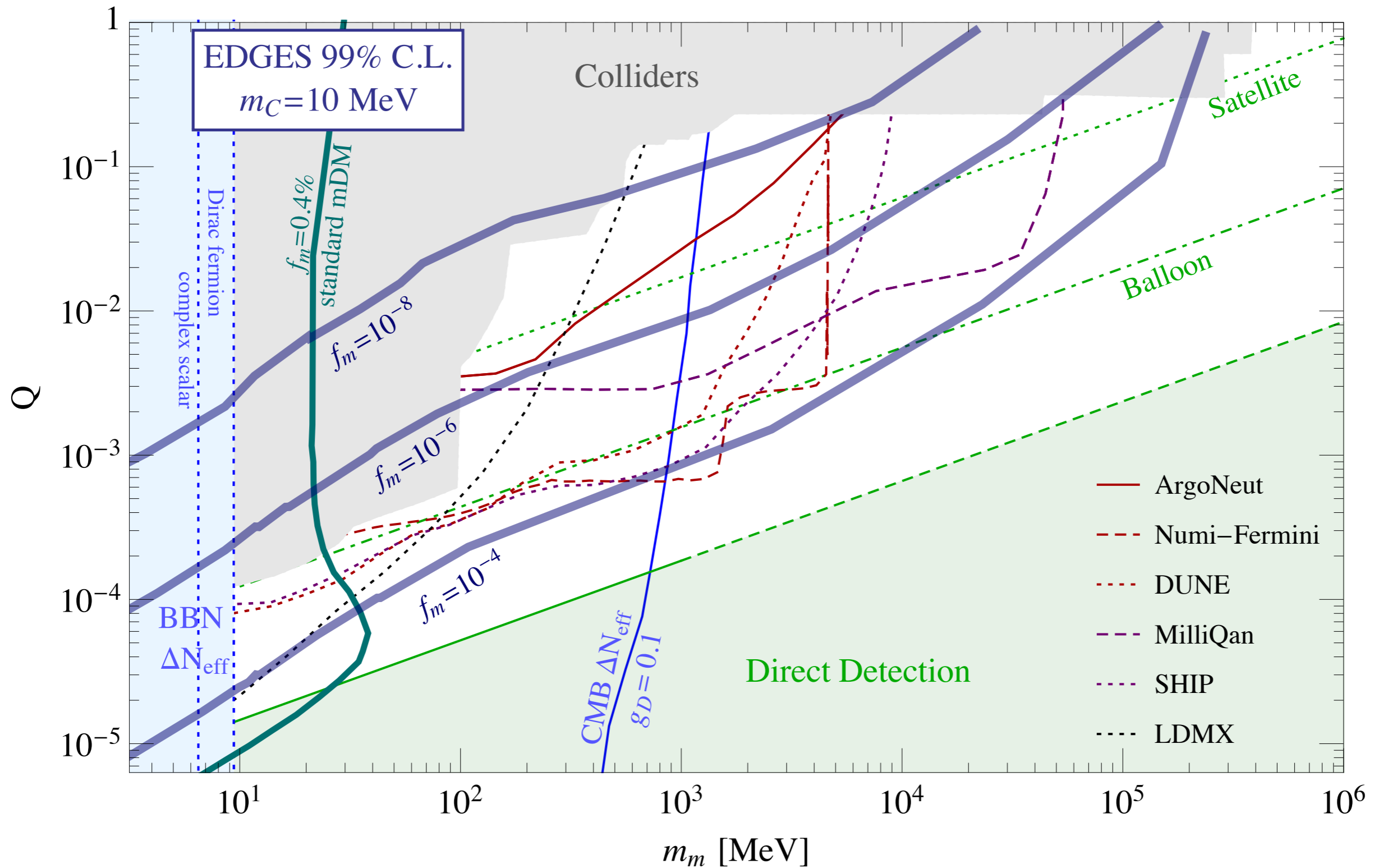
# RESULTS - MDM PARAMETERS



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## RESULTS

### Other general statements

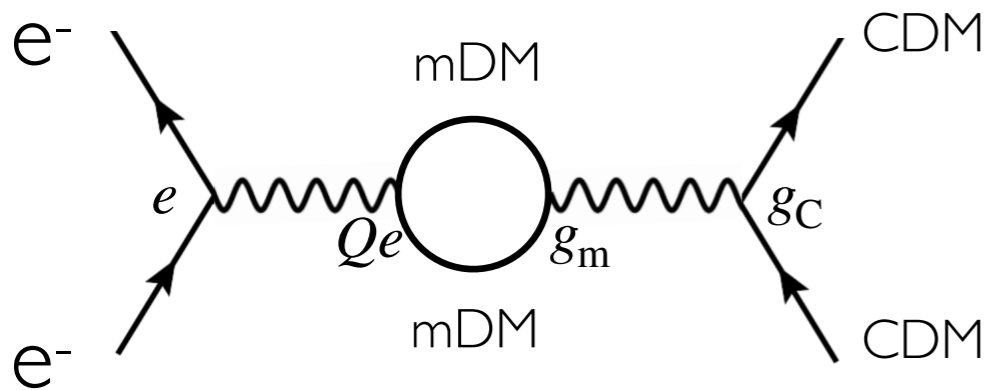
- Lower bound on CDM and mDM mass from  $N_{\text{eff}}$ :  $m_{\text{C,m}} \gtrsim 10 \text{ MeV}$
- Mass ordering\*:  $m_{\text{m}} \gtrsim m_{\text{C}}/10$
- For  $m_{\text{m}} \gtrsim 100 \text{ MeV}$  the inner structure of H and He atoms is sufficiently resolved, cooling is dominated by mDM-atoms interactions

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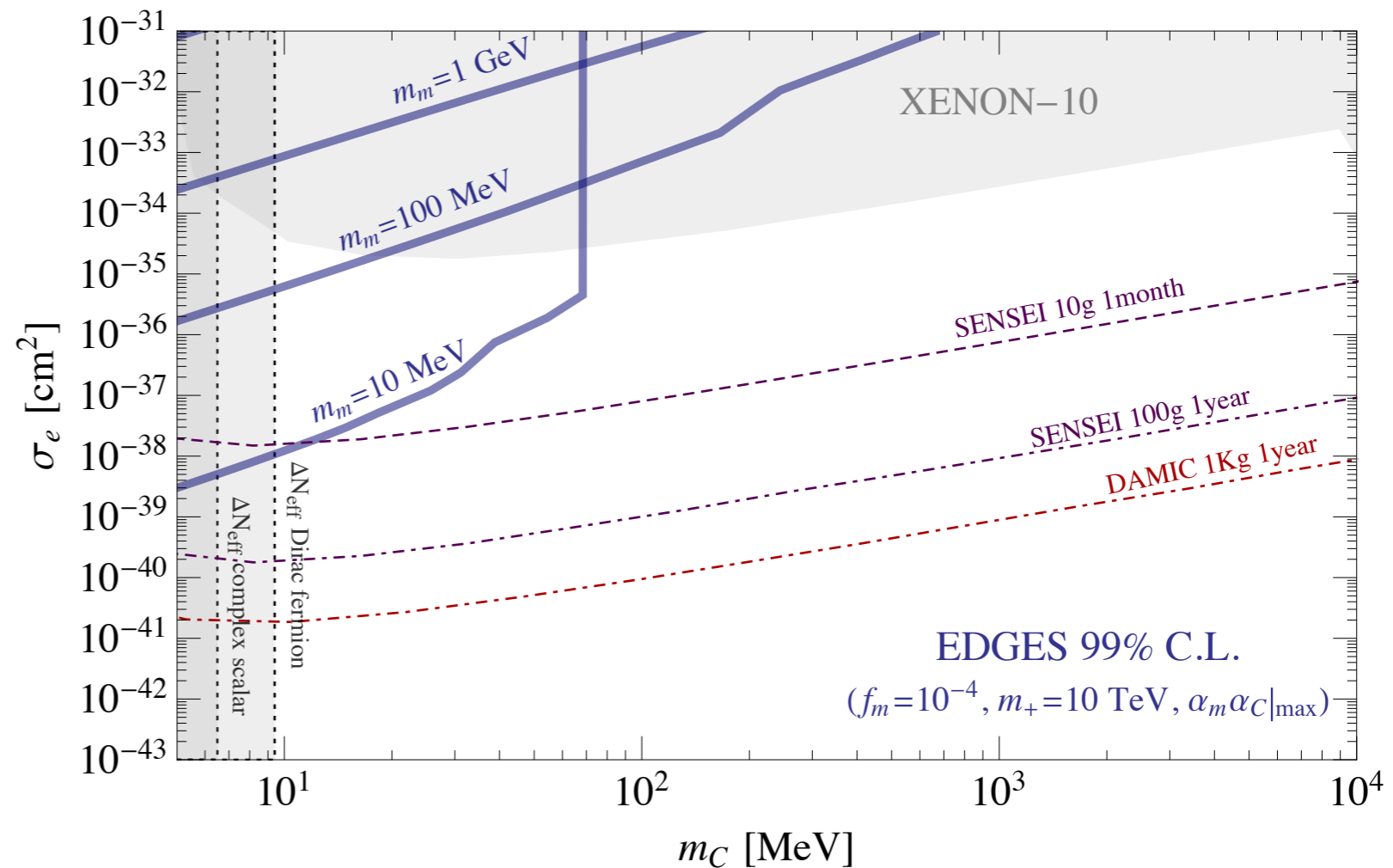
# RESULTS – CDM PARAMETERS

## Constraints

- Conservative self interactions constraints:  $\frac{\alpha_C^2}{m_C^3} \lesssim 10^{-11} \text{GeV}^{-3}$
- Loop induced CDM-SM coupling



$$\Rightarrow \bar{\sigma}_e \approx \frac{8Q^2\alpha_C\alpha_m}{\alpha_{EM}^2\mu_{em}^2} \left( \log \frac{\Lambda}{m_m} \right)^2$$

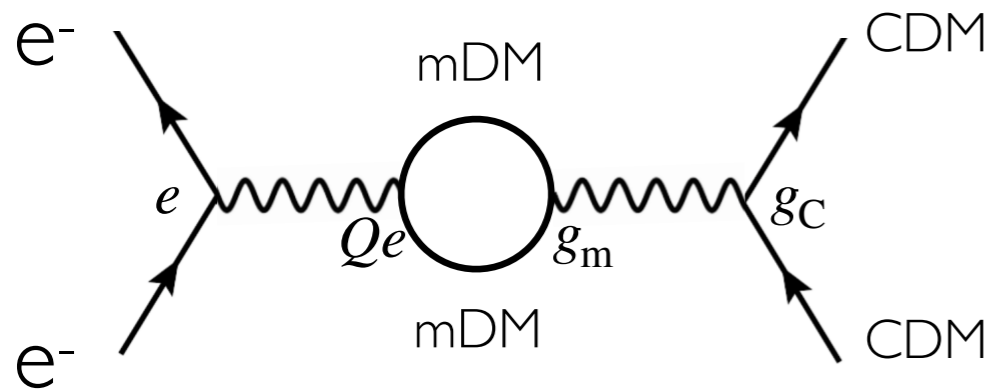




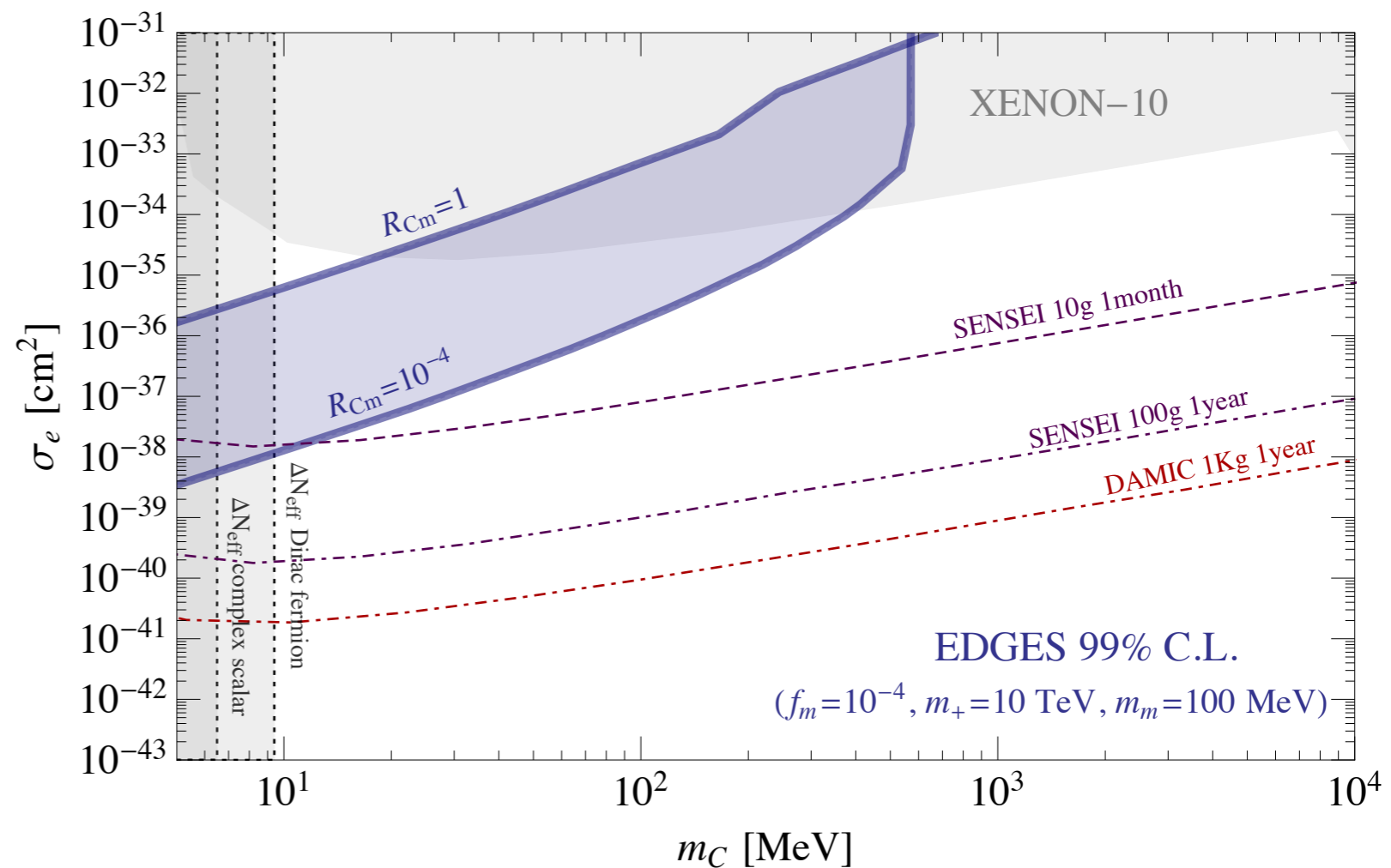
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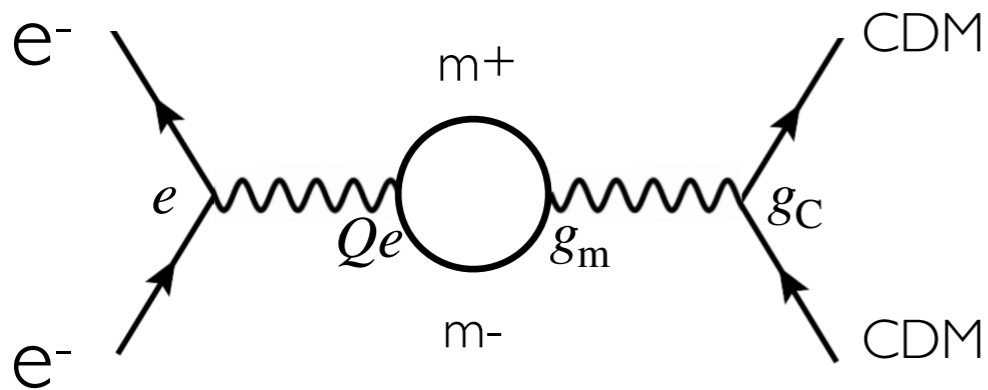
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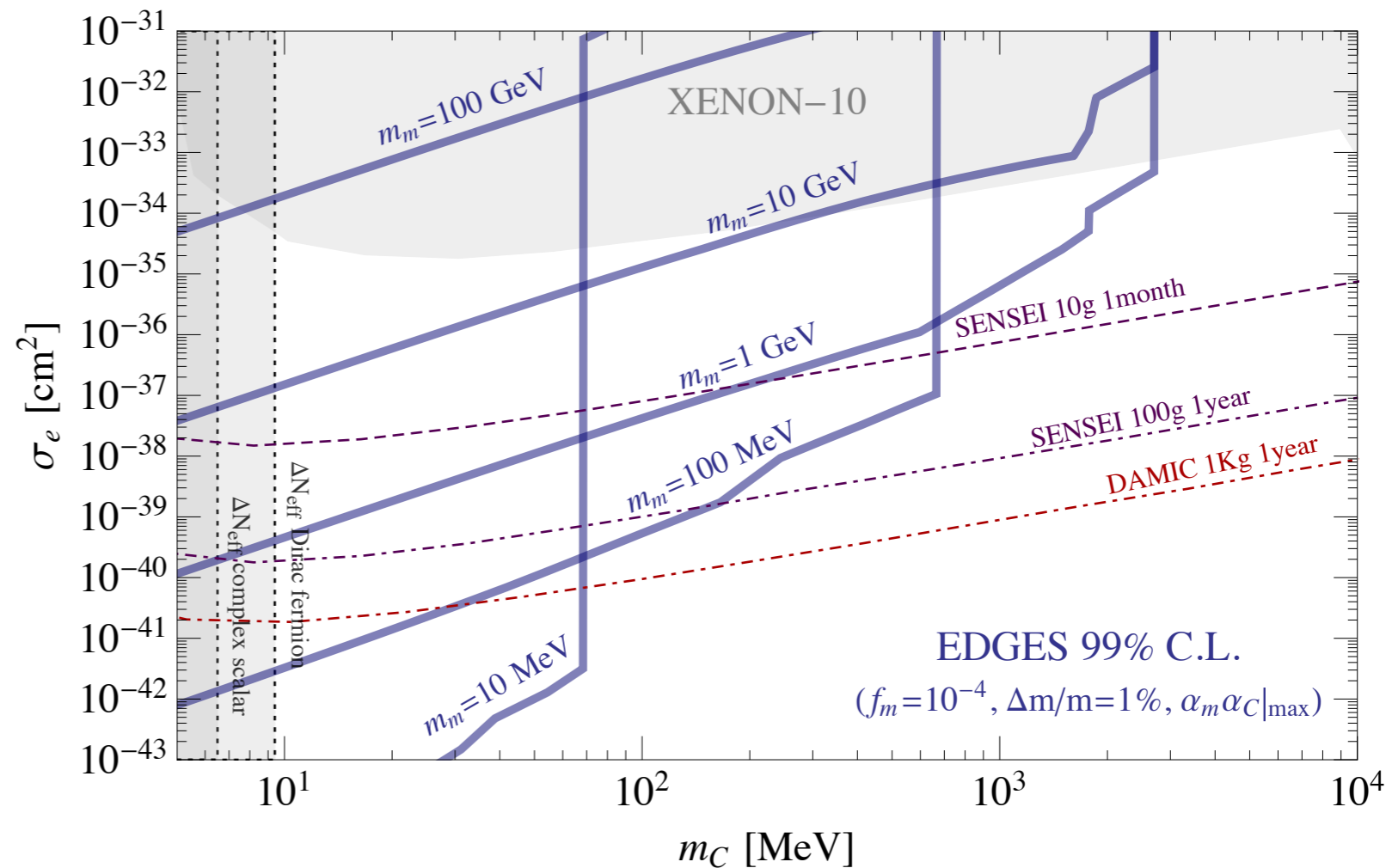
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- Galactic dynamics and the implications on direct detection?
- 21 cm spectrum and fluctuations?

**THANKS.**

**EXTRAS?**

