



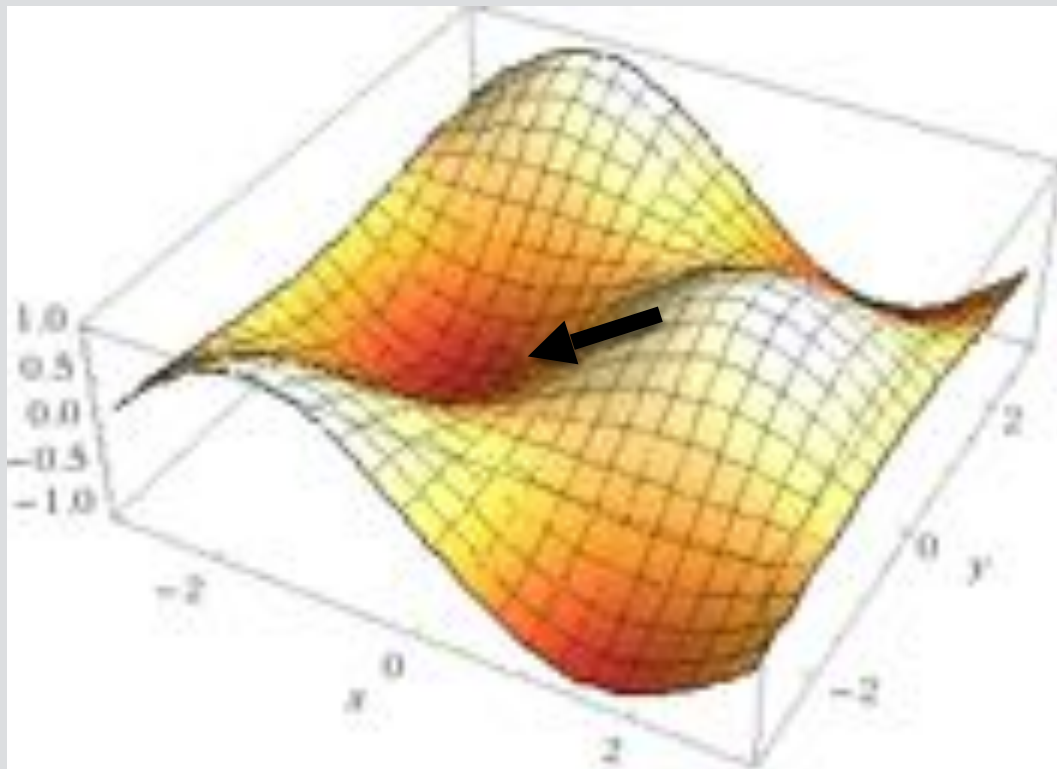
Axion dark matter detection with quantum gyroscopes

William Terrano
Princeton University,
Princeton Axion Project

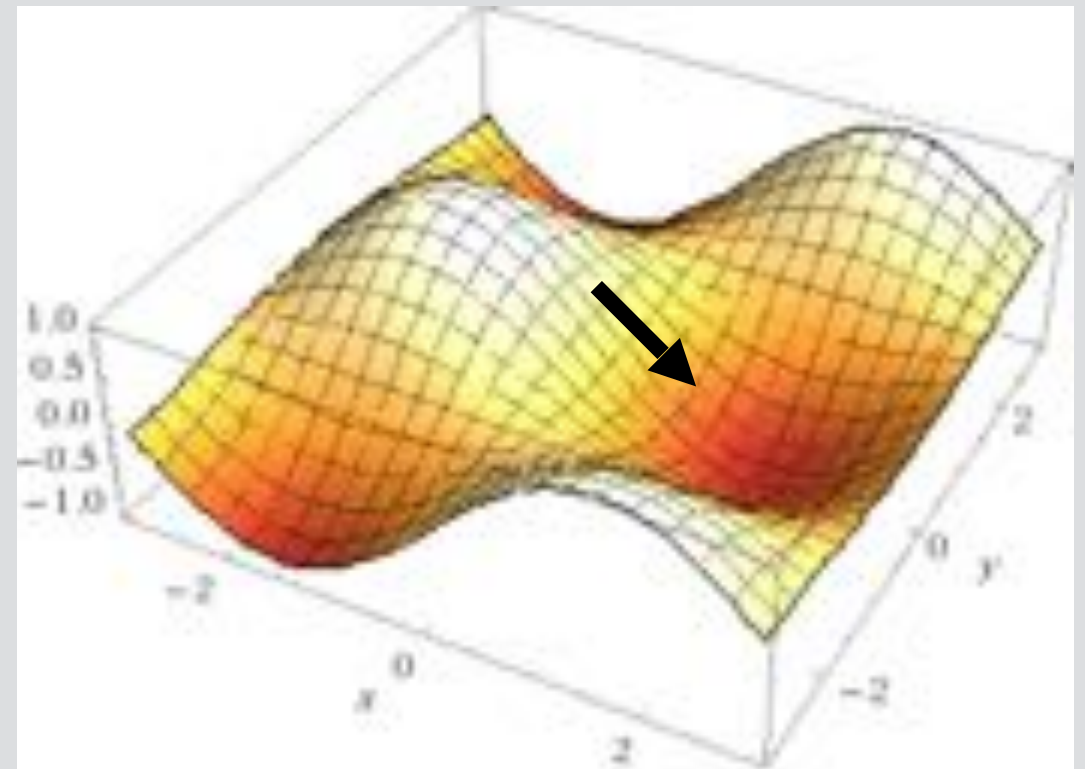
Axion splits Spin Up & Down levels

$$\mathcal{L} = (\partial_\mu a) \bar{\psi} \gamma^\mu \gamma_5 \psi$$

$$a = a_0 \cos \omega_C t$$



a at $t = 0$



a at $t = \pi/\omega_C$

Quantization axis defined by gradient of axion field
amplitude of splitting is oscillating

Axion splits

Spin Up from Down

- magnitude determined by dark matter density, velocity and coupling strength

$$\vec{\nabla} a \sim \vec{p}_a a \sim m_a \vec{v} a_0 \cos m_a t. \quad \rho_{\text{DM}} = \frac{1}{2} m_a^2 a^2$$

$$H_{\text{ax}} \sim 10^{-25} \text{ eV} \left(\frac{g_{a\bar{\psi}\psi}}{10^{-10} \text{ GeV}^{-1}} \right) \left(\frac{v}{10^{-3}} \right) \left(\sqrt{\frac{\rho_{\text{DM}}}{(0.04 \text{ eV})^4}} \right) \cos m_a t$$

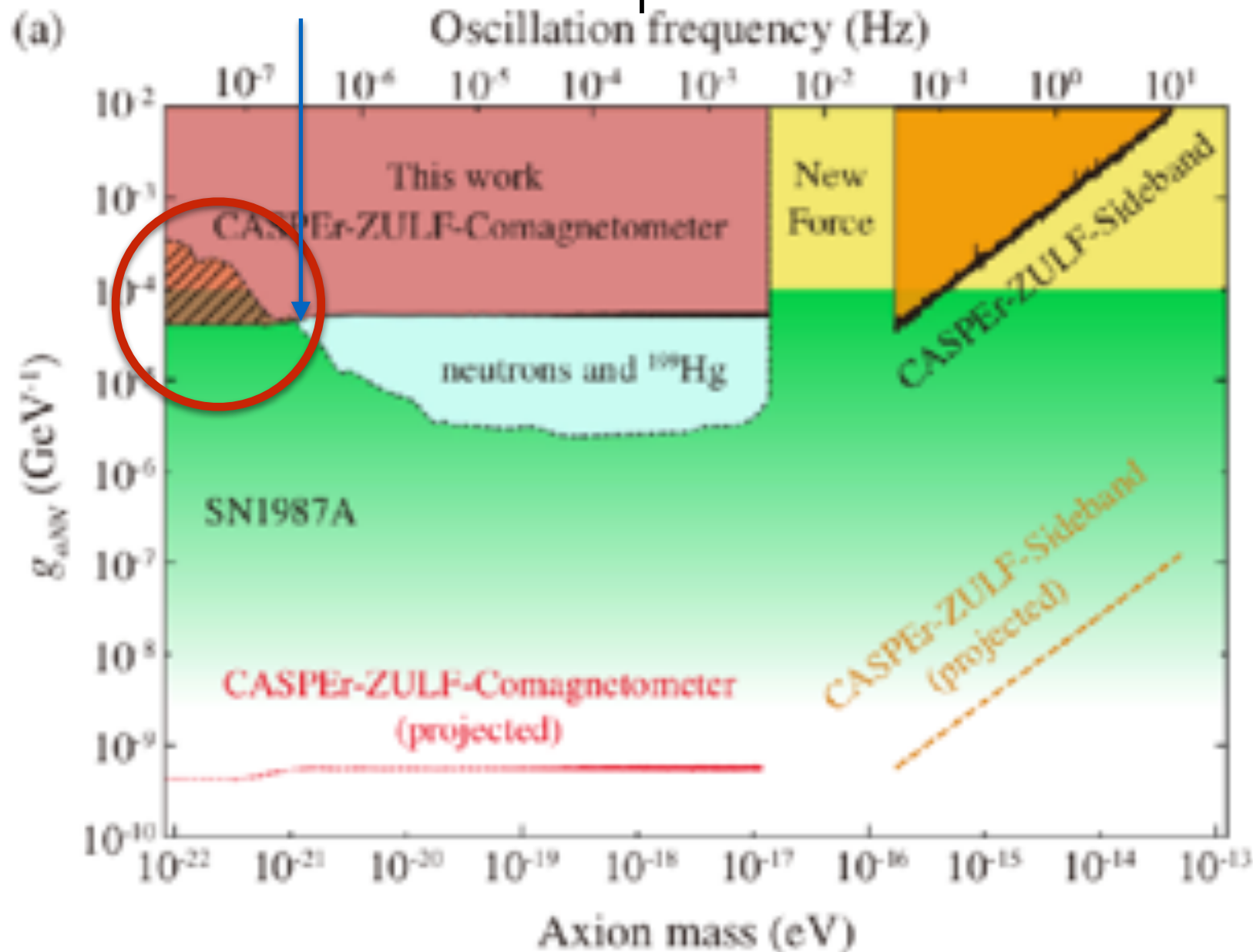
- Intuition: splitting of a magnetic moment in a magnetic field;

$$B \sim g_{a\bar{\psi}\psi} v \sqrt{2\rho_{\text{DM}}} / \gamma_{\psi}$$

- Spin-precession  spin-flip  spin-alignment 
 \Rightarrow Not shielded by magnetic shielding

Existing constraints on this coupling

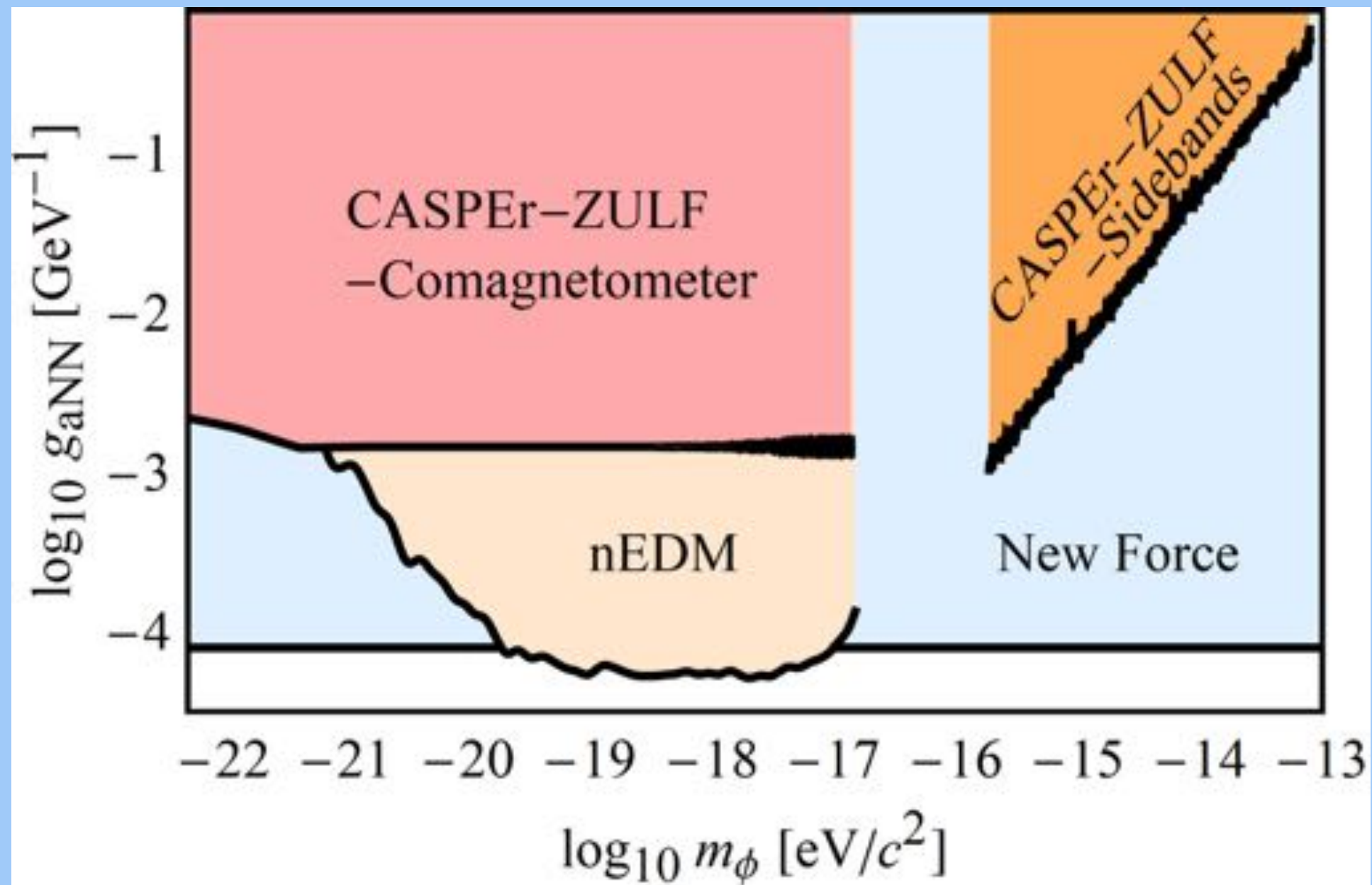
inverse duration of expt



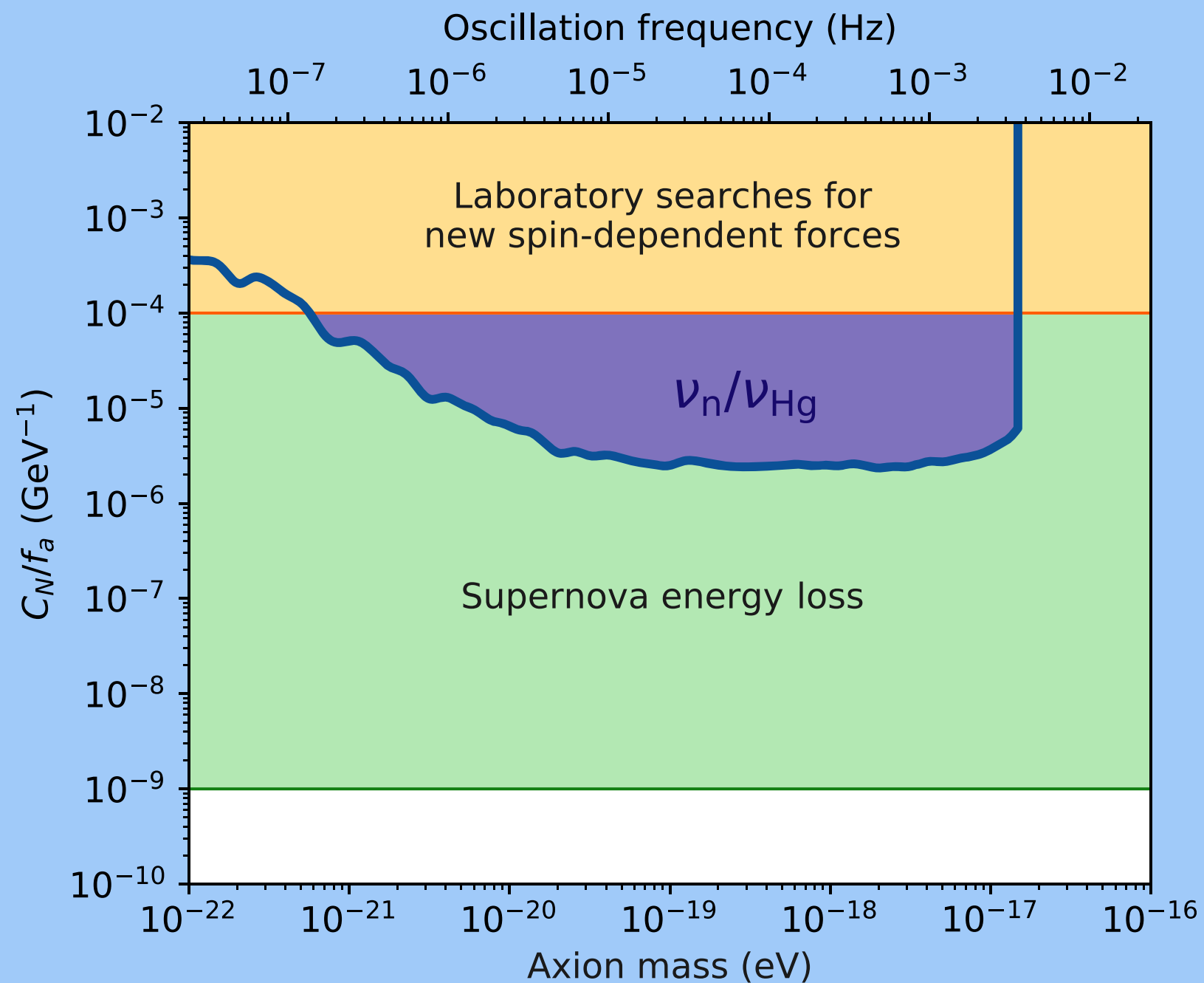
$$\sin(2\pi t/\tau) \approx 2\pi t/\tau$$

$$\cos(2\pi t/\tau) \approx 1 - (2\pi t/\tau)^2/2$$

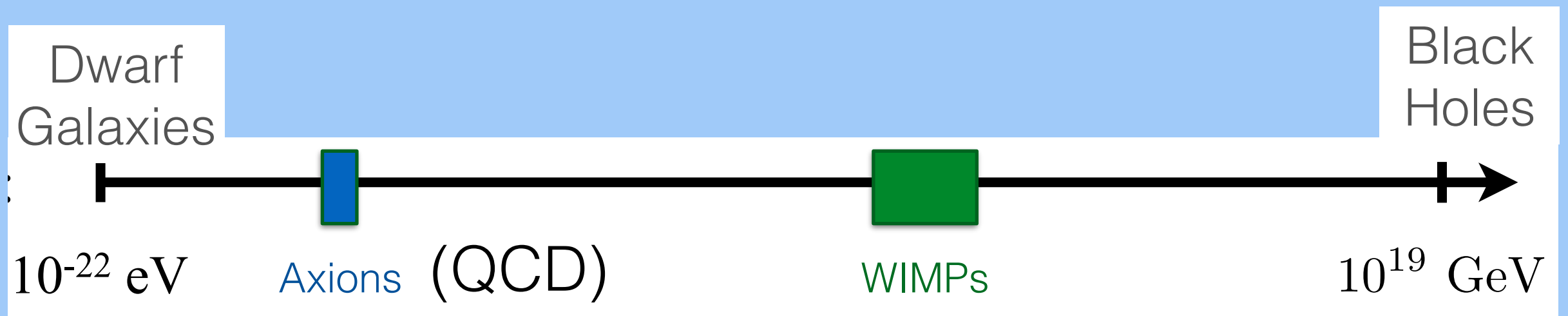
Updated constraints?



Existing constraints



'Fuzzy' Dark Matter



← 1 eV →
Axions (no QCD) = pseudo-Goldstone boson



Fuzzy Dark Matter (Hu, Barkana, Gruzinov, 2001)
(Hui, Ostriker, Tremaine, Witten 2017)

Fuzzy Dark Matter

Light axion dark matter

mass $m \sim 10^{-22}$ eV

Fuzzy dark matter
Hu, Barkana, Gruzinov

- A natural candidate for such a light particle is a pseudo Goldstone boson.
- Concrete realization: an angular field of periodicity $2\pi F$ i.e. an axion-like field with a potential from non-perturbative effects (not QCD axion).

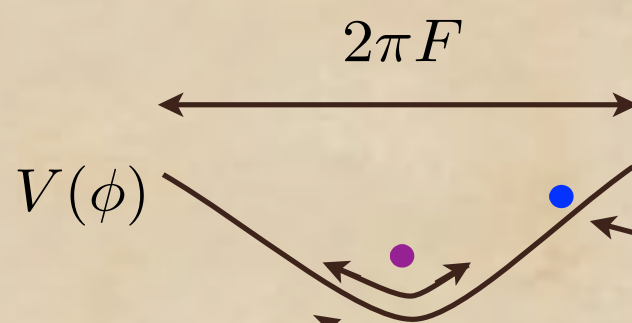
$$\mathcal{L} \sim -\frac{1}{2}(\partial\phi)^2 - \Lambda^4(1 - \cos[\phi/F])$$

$$m \sim \Lambda^2/F$$

(see Monday's talks)

- Relic abundance:

(standard story - see Kolb & Turner;
review by Marsh)



$\phi \sim F$ at early times until $H \sim m$

$$\rho_\phi \sim m^2 F^2, \quad \rho_{\text{rad.}} \sim H^2 M_{\text{pl}}^2 \sim m^2 M_{\text{pl}}^2 \sim T^4$$

subsq. oscill. : $\rho_\phi \propto a^{-3}$

self - interaction can be ignored

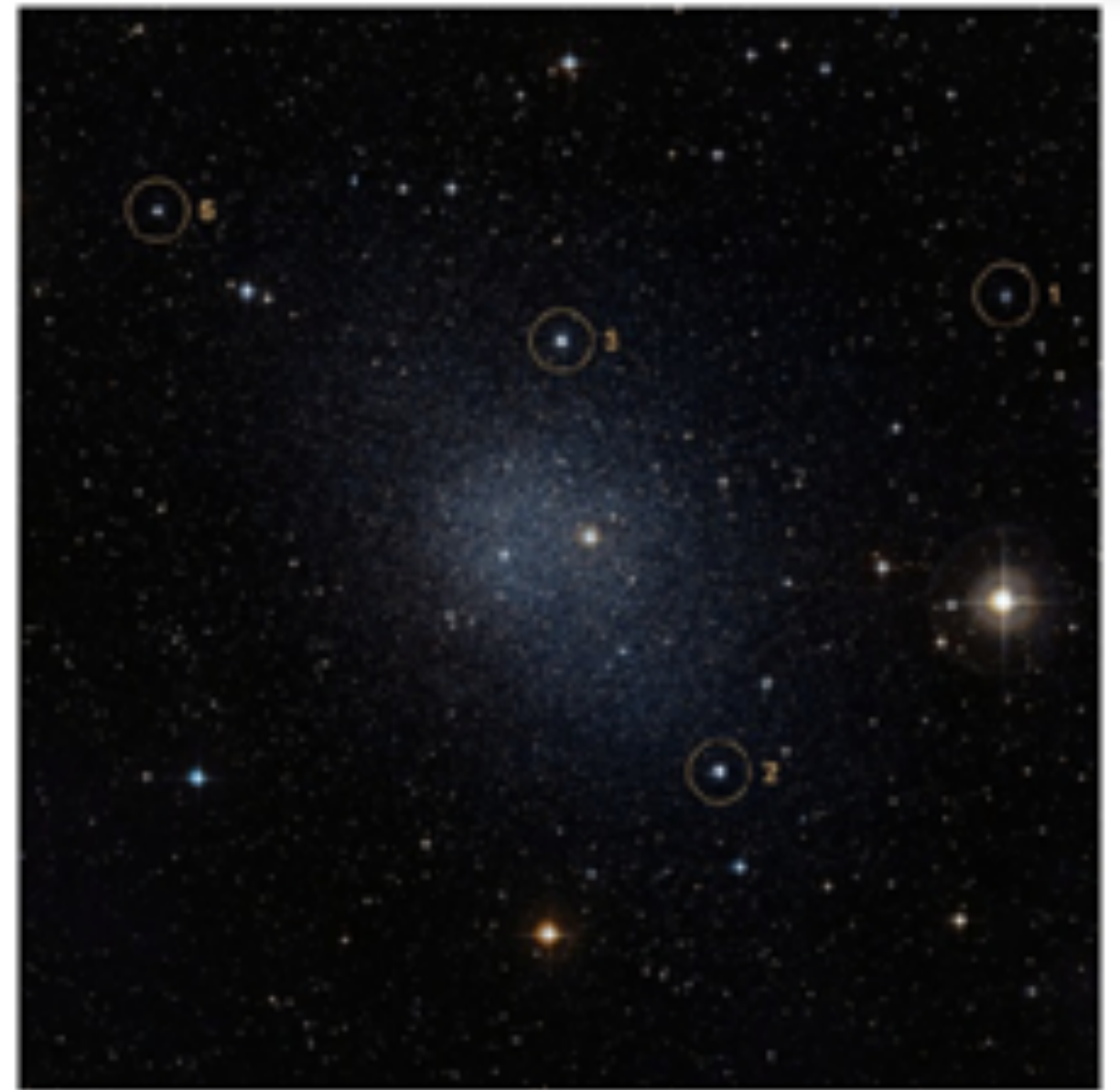
Slide from
L. Hui

$$\Omega_{\text{matter}} \sim \left(\frac{F}{10^{17} \text{ GeV}} \right)^2 \left(\frac{m}{10^{-22} \text{ eV}} \right)^{1/2}$$

(low scale inflation)

Why haven't Globular clusters collapsed?

Fuzzy dark matter reduces dynamical friction by $\times 10$



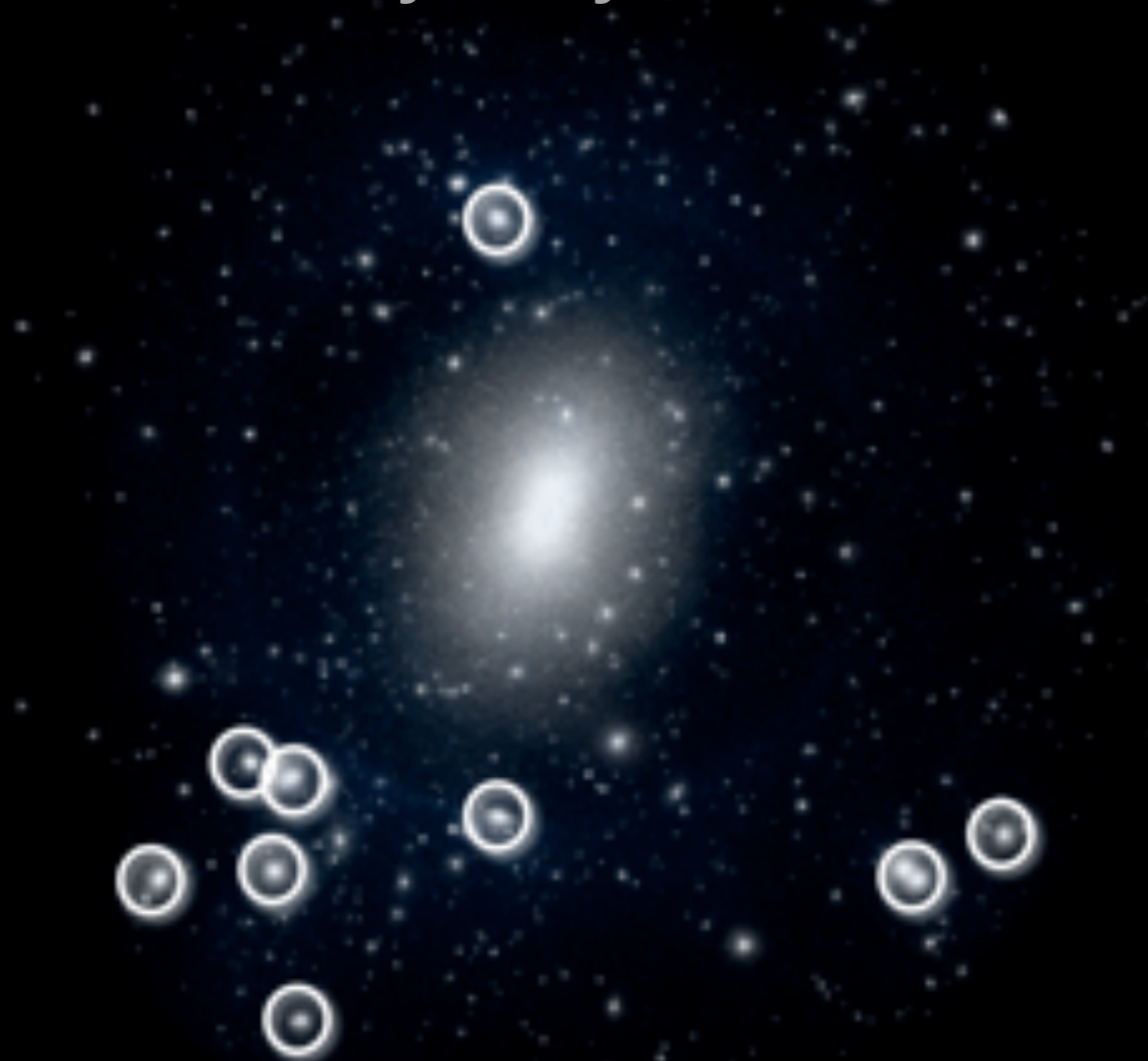
Dynamical friction issue: Tremaine 1976

Why isn't there more small scale structure?

10

Dwarf galaxies; Rotation curves

CDM predicts 100s
of milky way halos

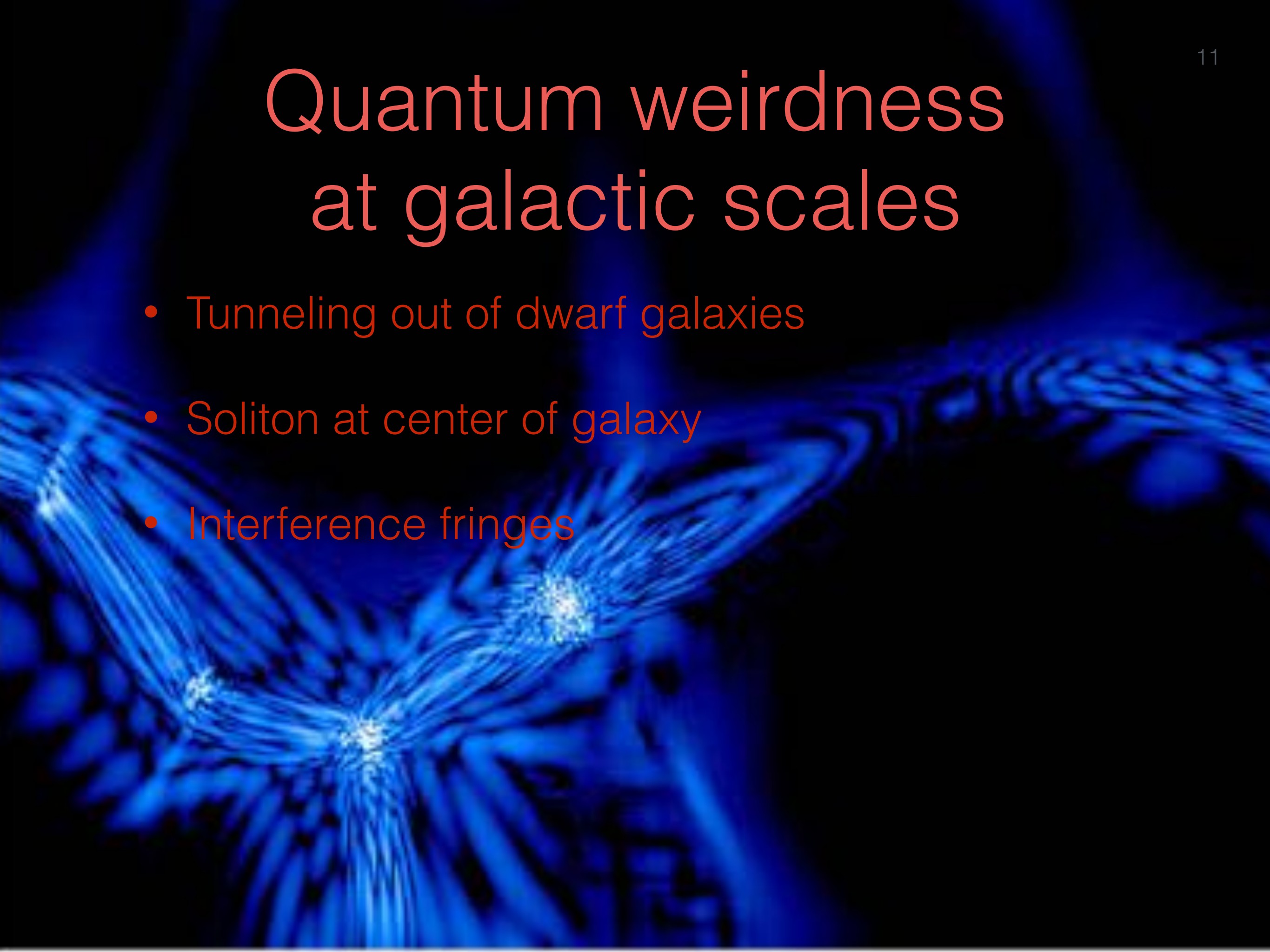


We see 10



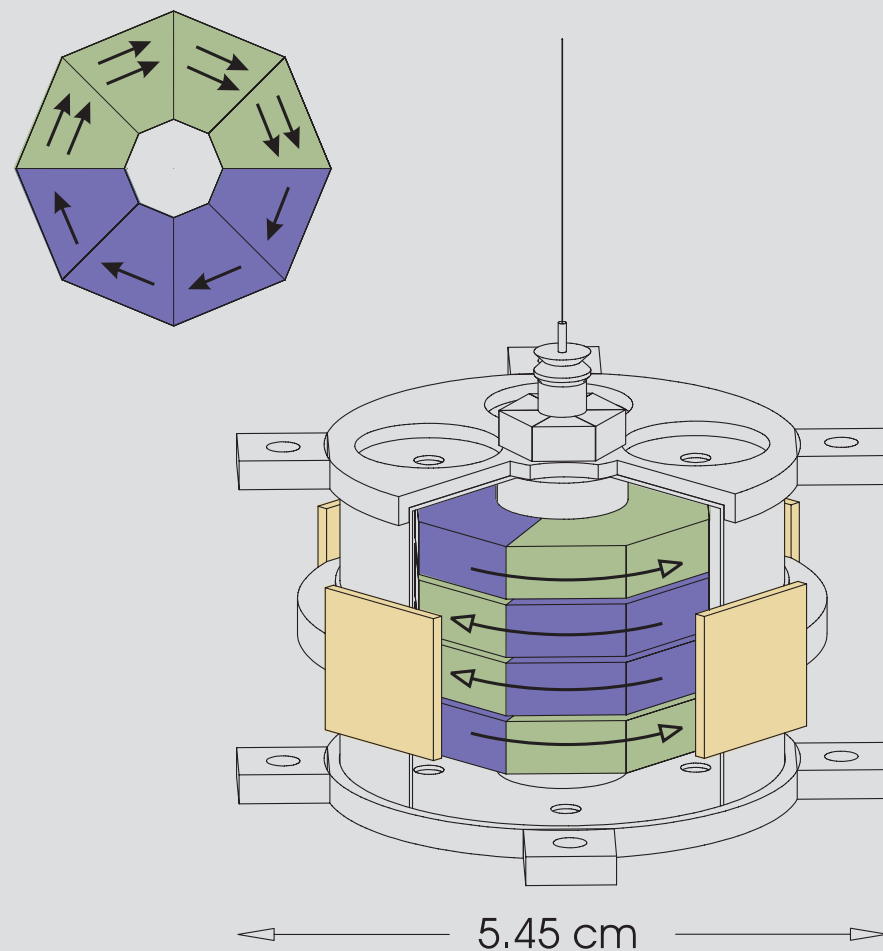
Quantum weirdness at galactic scales

- Tunneling out of dwarf galaxies
- Soliton at center of galaxy
- Interference fringes

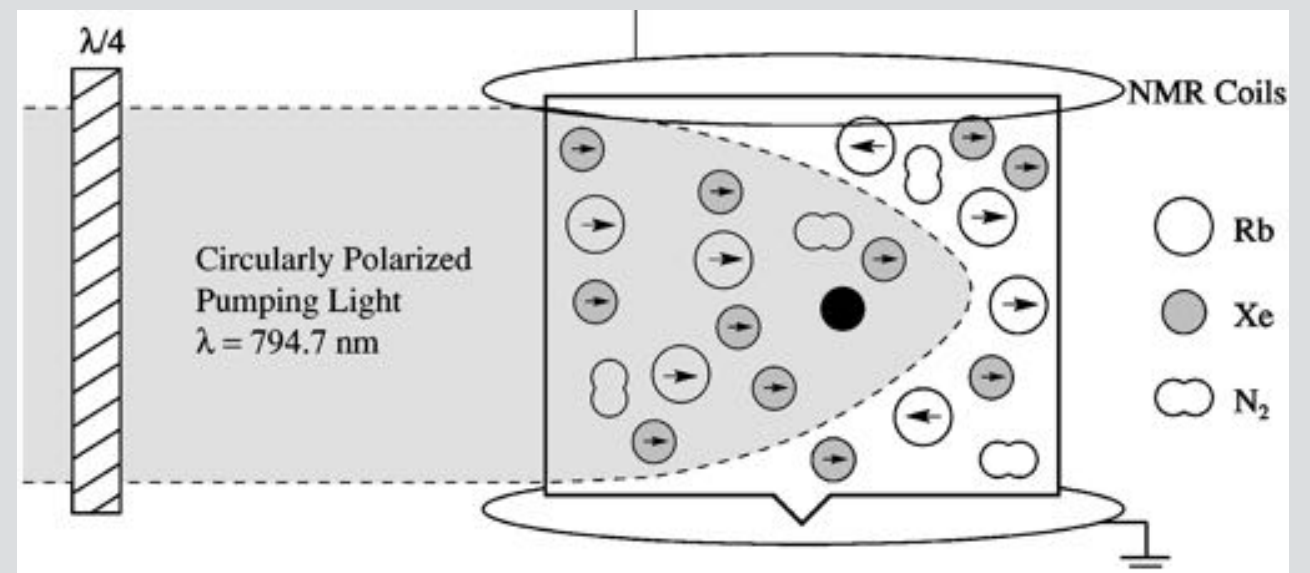


Quantum-mechanical (spin) Gyroscopes

Electronic Spins



Nucleonic Spins

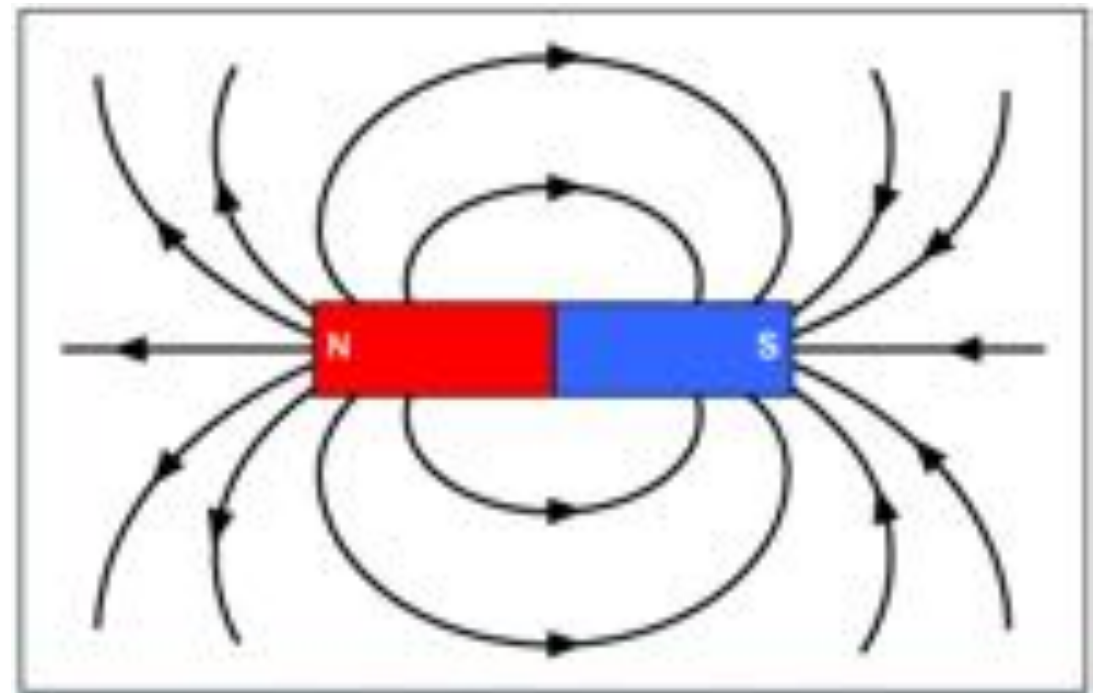
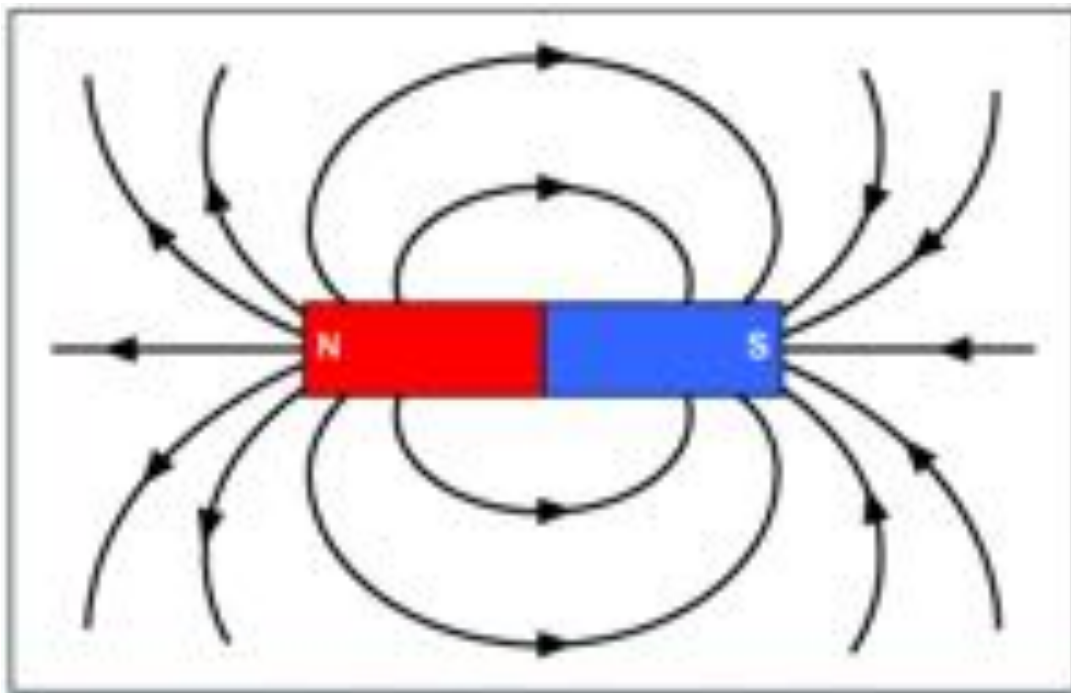


Making a Spin Polarized Pendulum

- Alternating Alnico and SmCo5 magnets
 - At 1 Tesla, Alnico has a spin density of $\sim 7.8 \cdot 10^{22}$ spins/cc
 - At 1 Tesla, SmCo5 has a spin density of $\sim 4.2 \cdot 10^{22}$ spins/cc

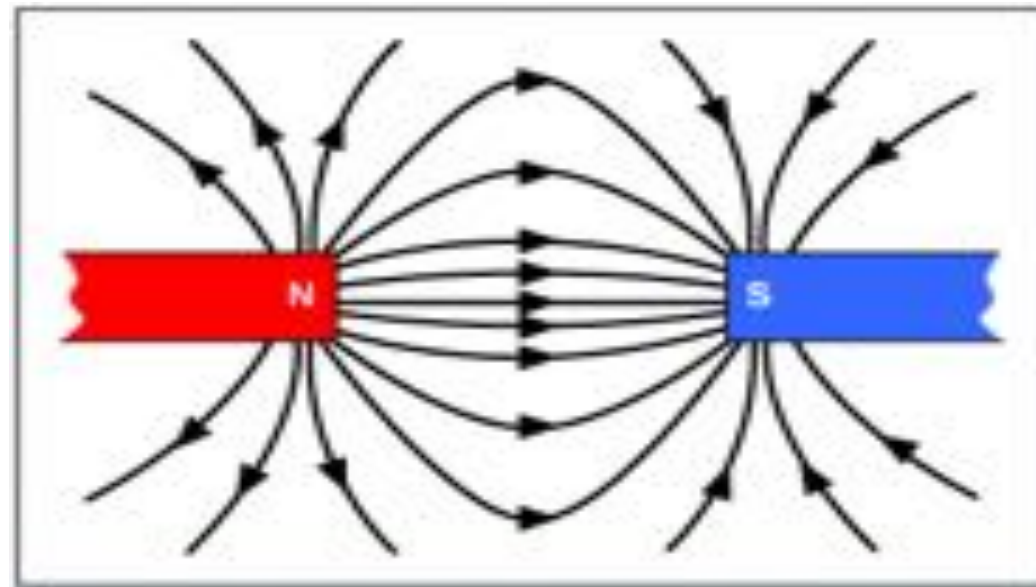
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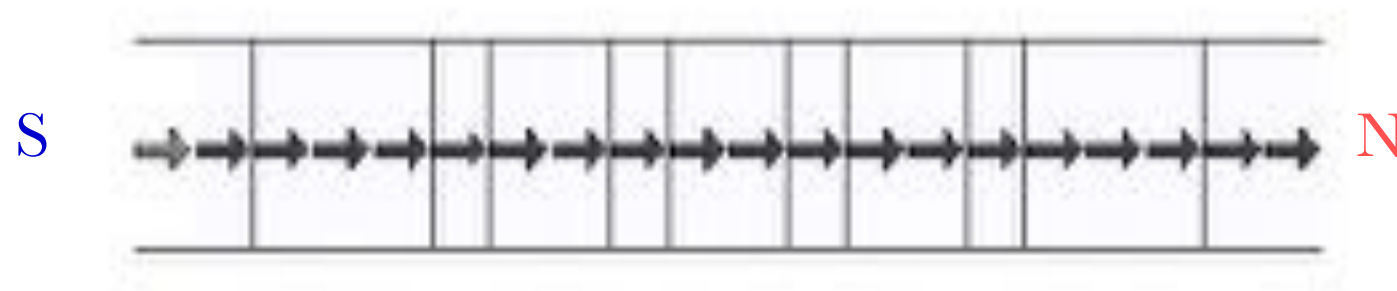
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Making a Spin Polarized Pendulum

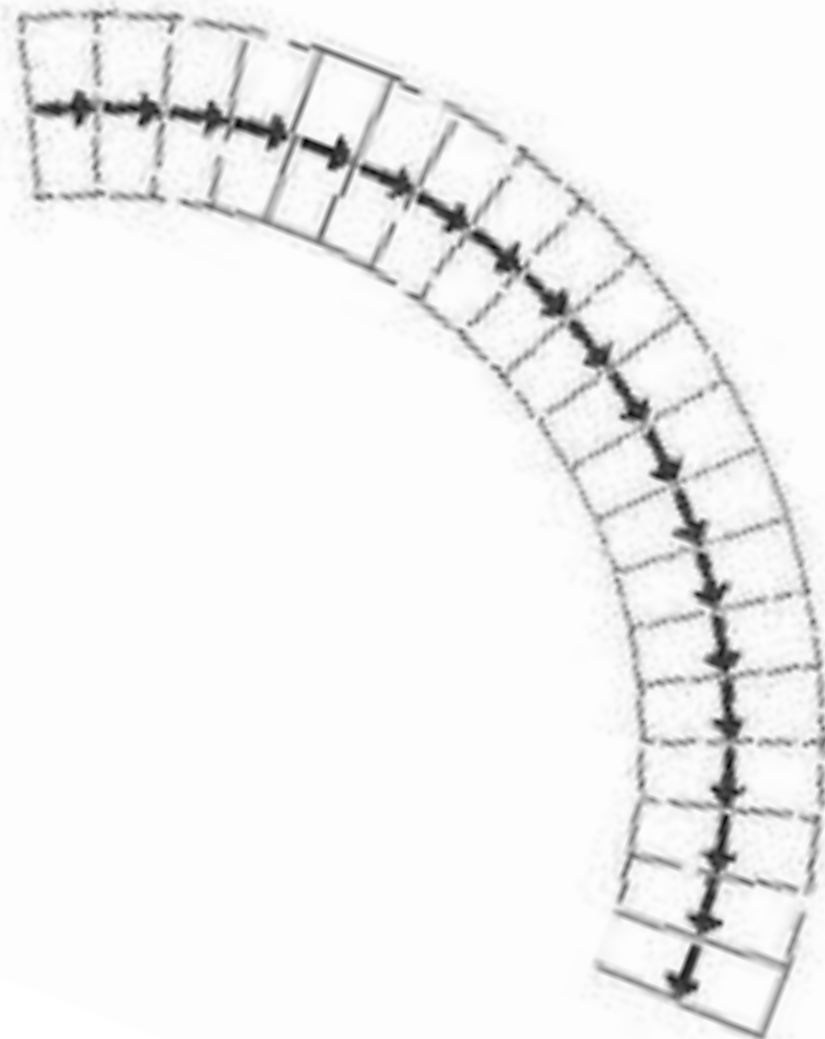
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Making a Spin Polarized Pendulum

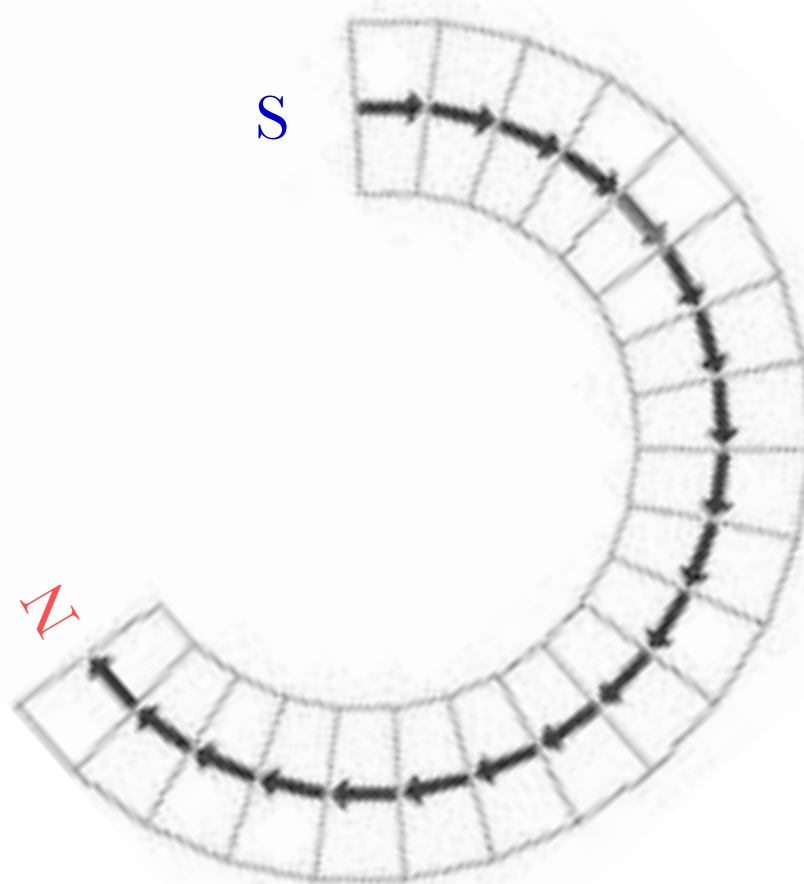
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S



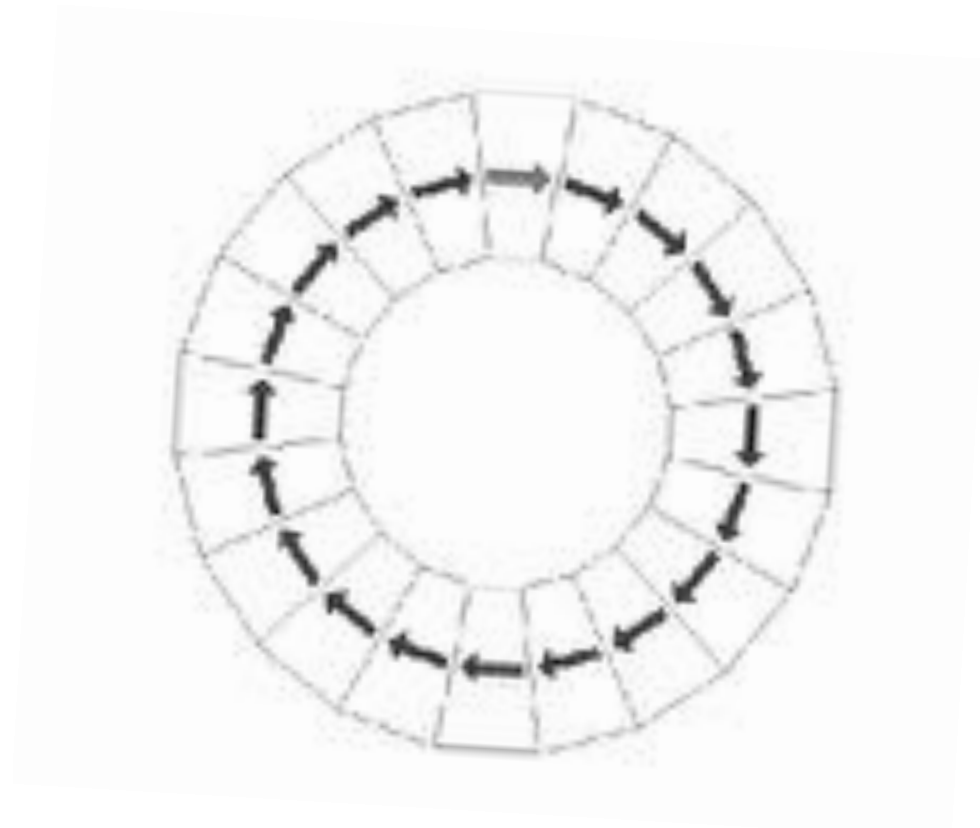
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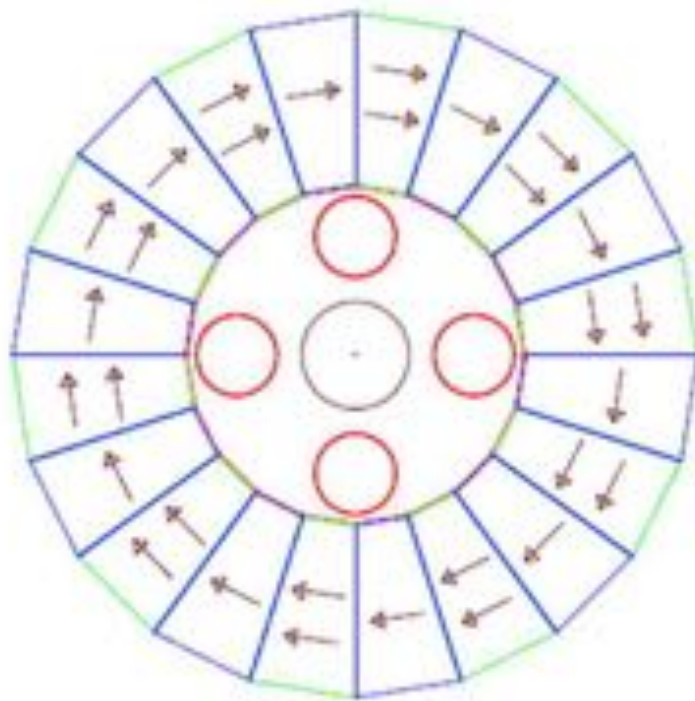
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- Place them North pole to South Pole to contain the magnetic flux:

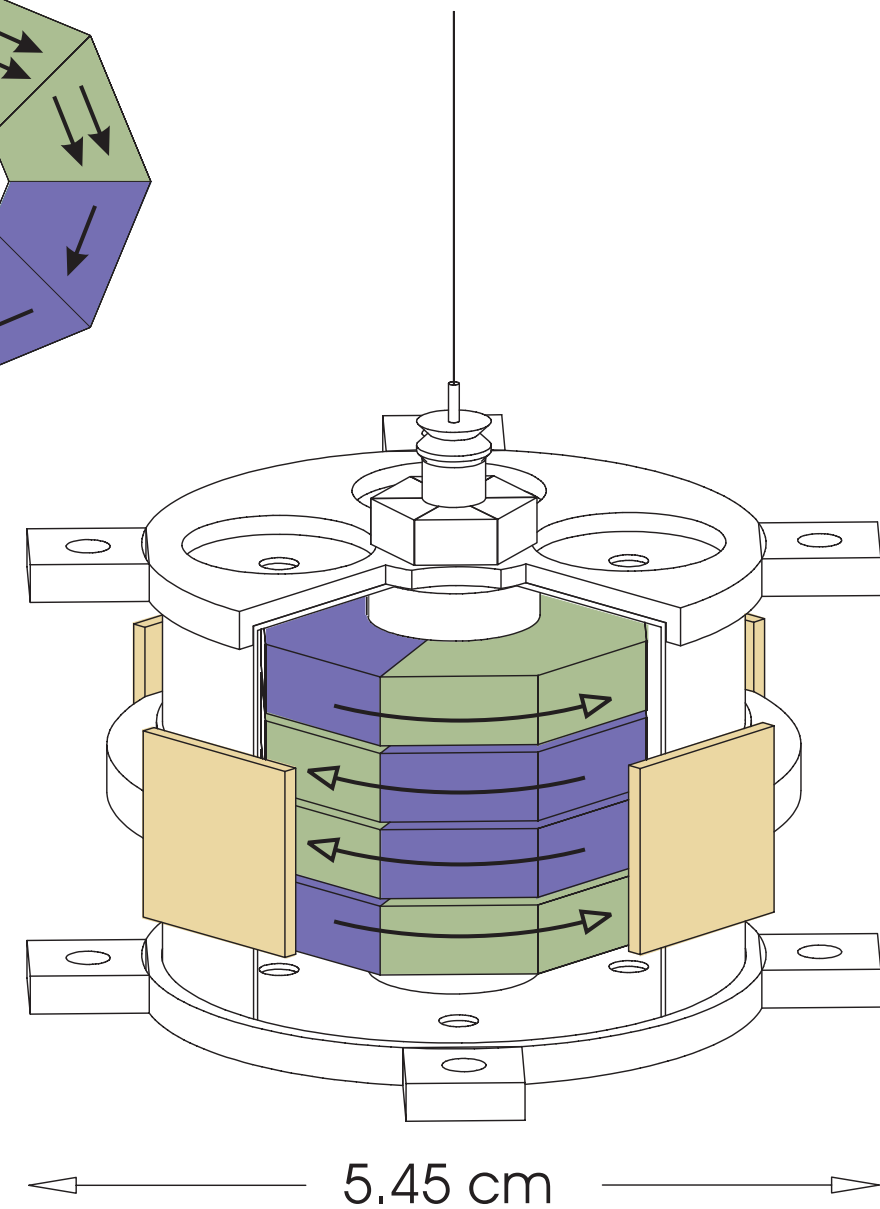
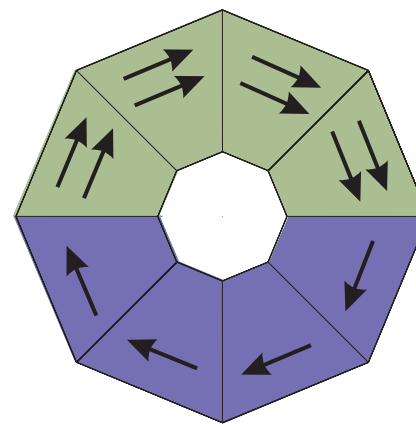
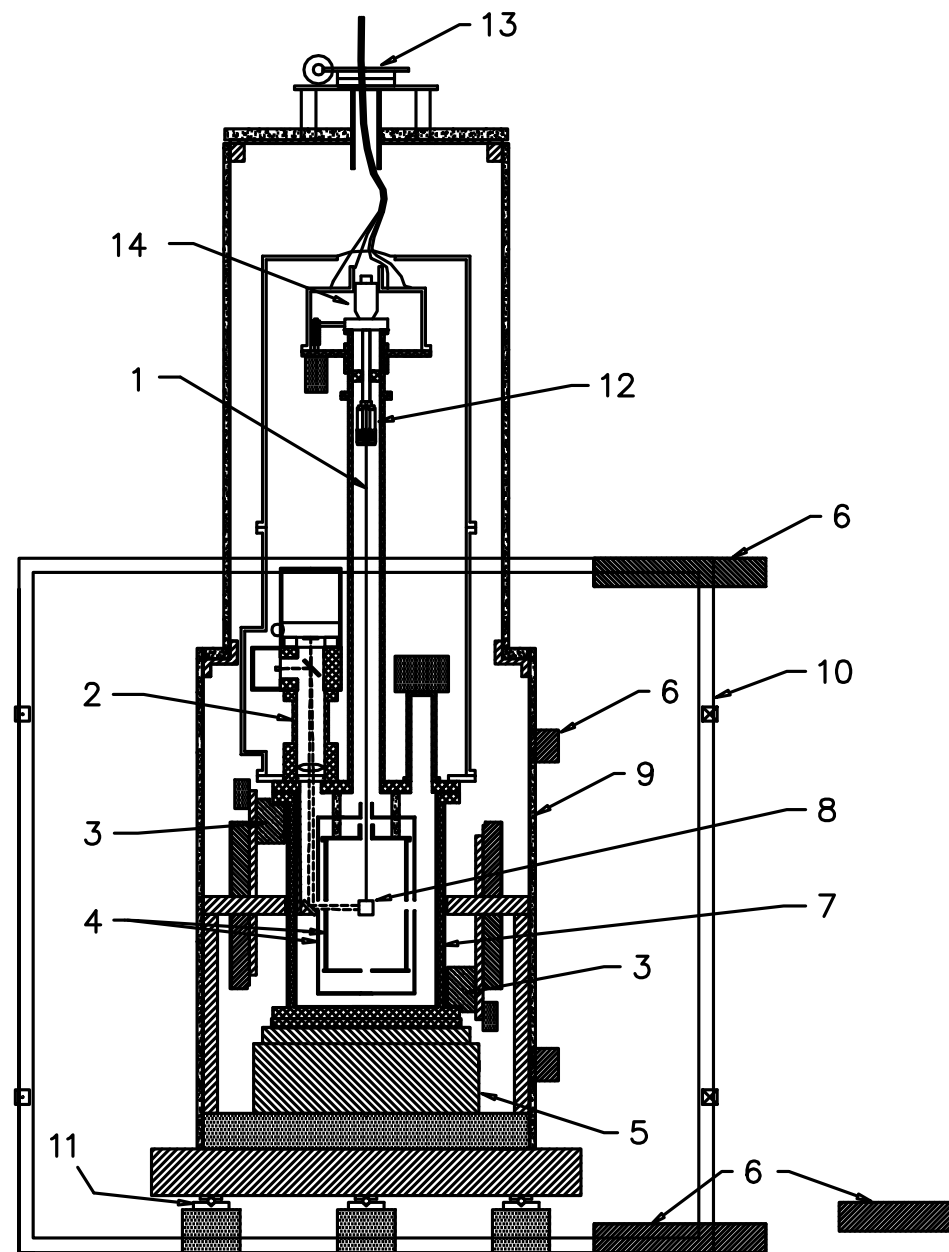


Making a Spin Polarized Pendulum

- Alternating Alnico and SmCo5 magnets
 - At 1 Tesla, Alnico has a spin density of $\sim 7.8 \cdot 10^{22}$ spins/cc
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- Contains a Spin Gradient but no magnetic field!



CPT, fifth force tests



Quantum Gyroscope:

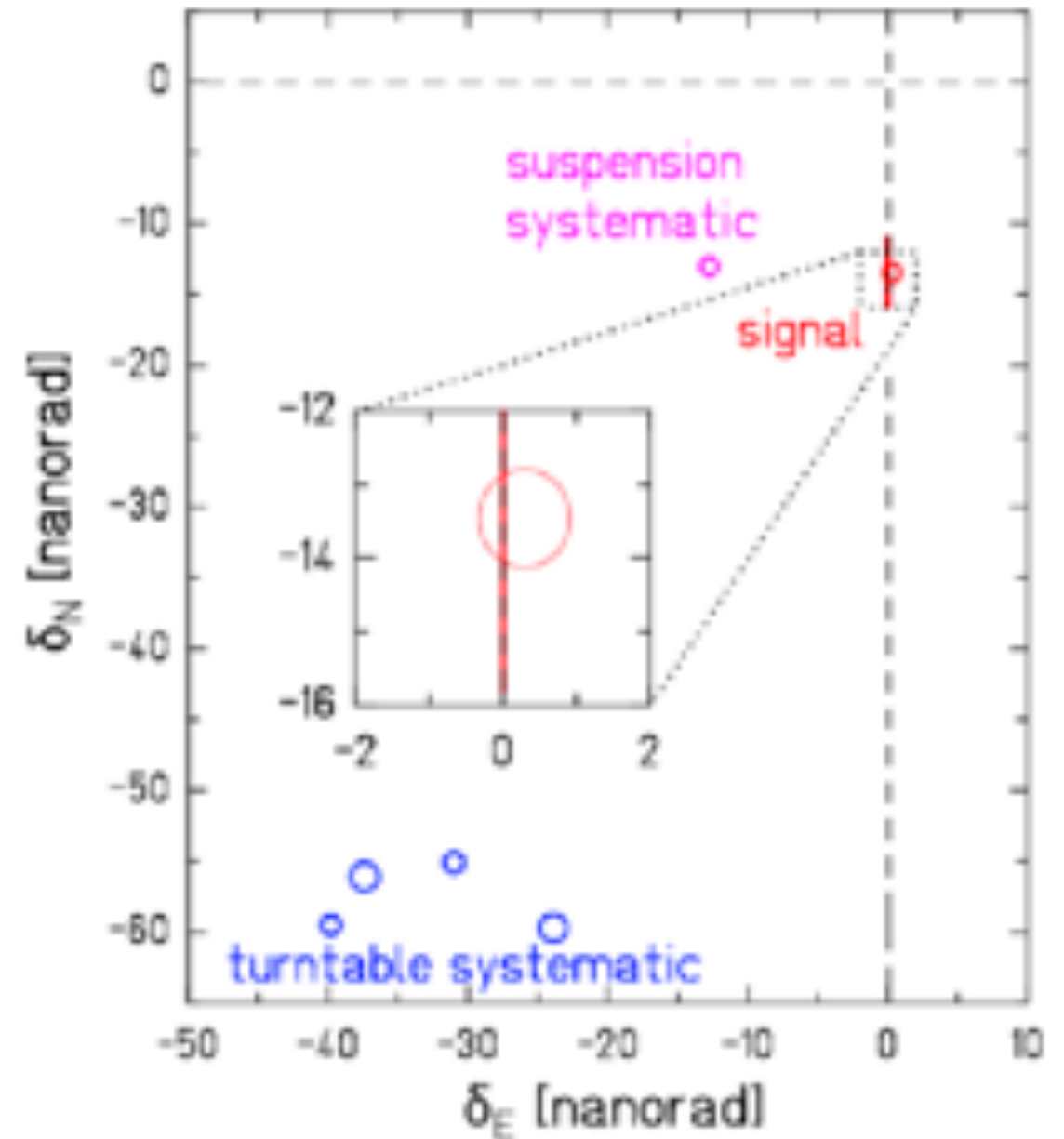
$$B_z(\text{SmCo}_5)/\mu_B = g_s S_z(\text{SmCo}_5) + g_l L_z(\text{SmCo}_5)$$

$$B_z(\text{Alnico})/\mu_B = g_s S_z(\text{Alnico}) + g_l L_z(\text{Alnico}) ,$$

$$2\langle S_z^{\text{tot}} \rangle + \langle L_z^{\text{tot}} \rangle = 0$$

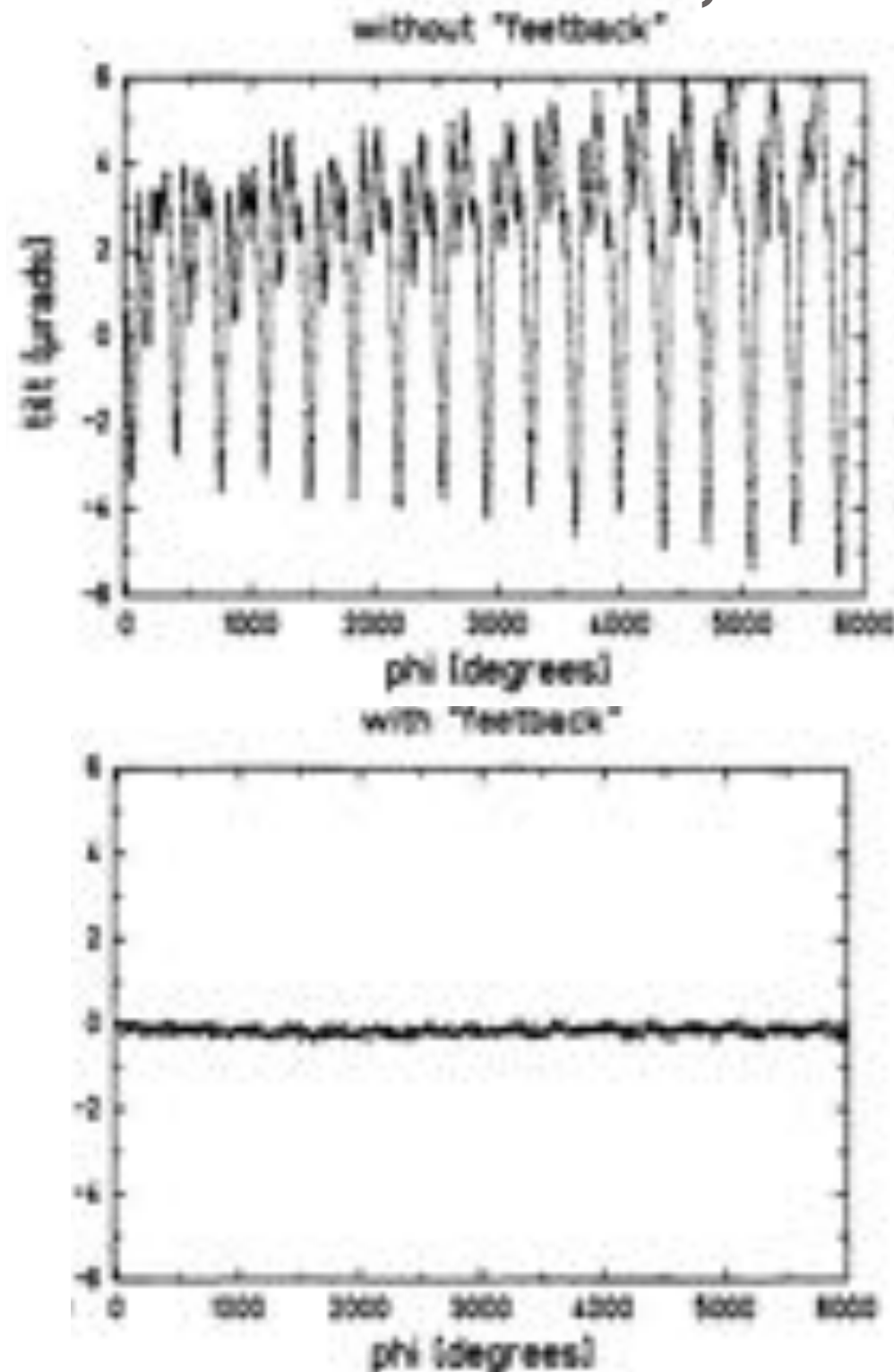
$$\langle J_z^{\text{tot}} \rangle = -\langle S_z^{\text{tot}} \rangle$$

$$\mathbf{T} \cdot \hat{\mathbf{n}} = |\boldsymbol{\Omega}_{\oplus} \times \mathbf{J} \cdot \hat{\mathbf{n}}|$$

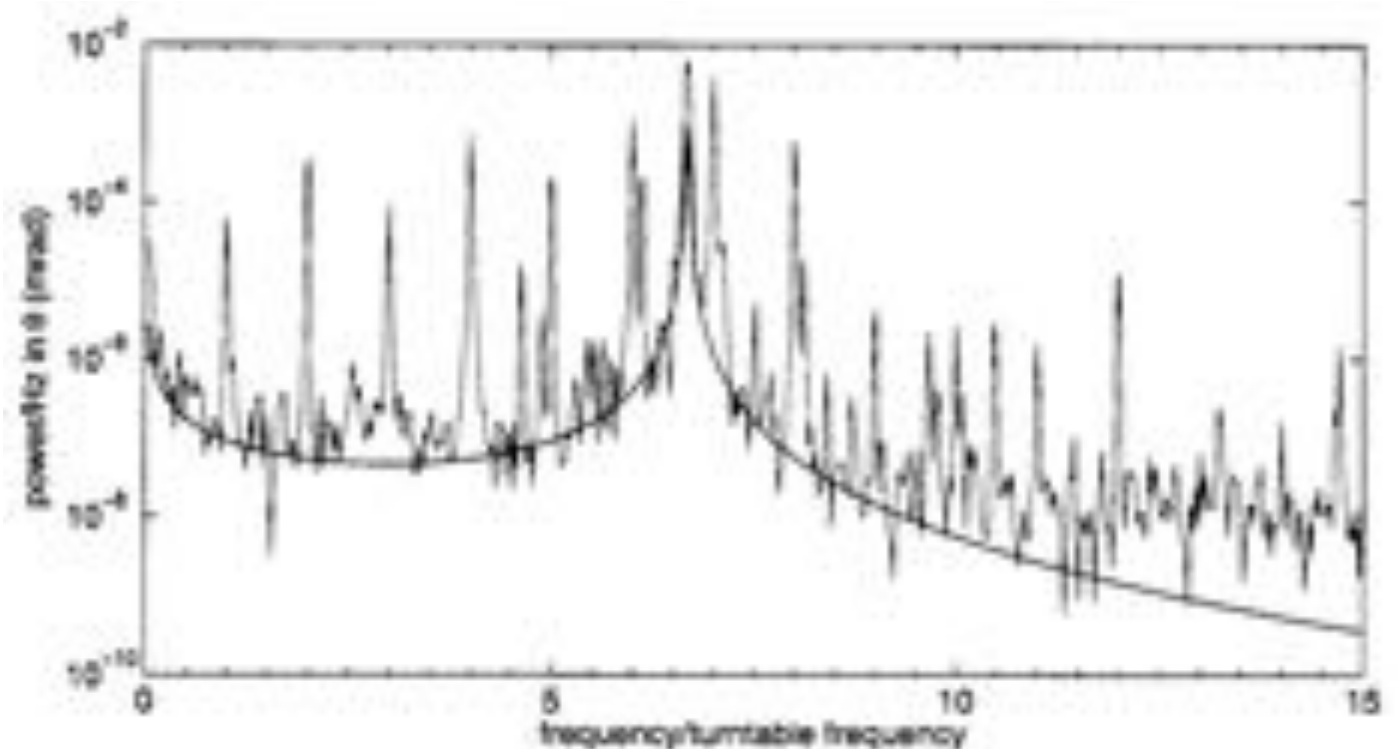


triply-modulated signal

Turntable; Earth rotation; Axion mass



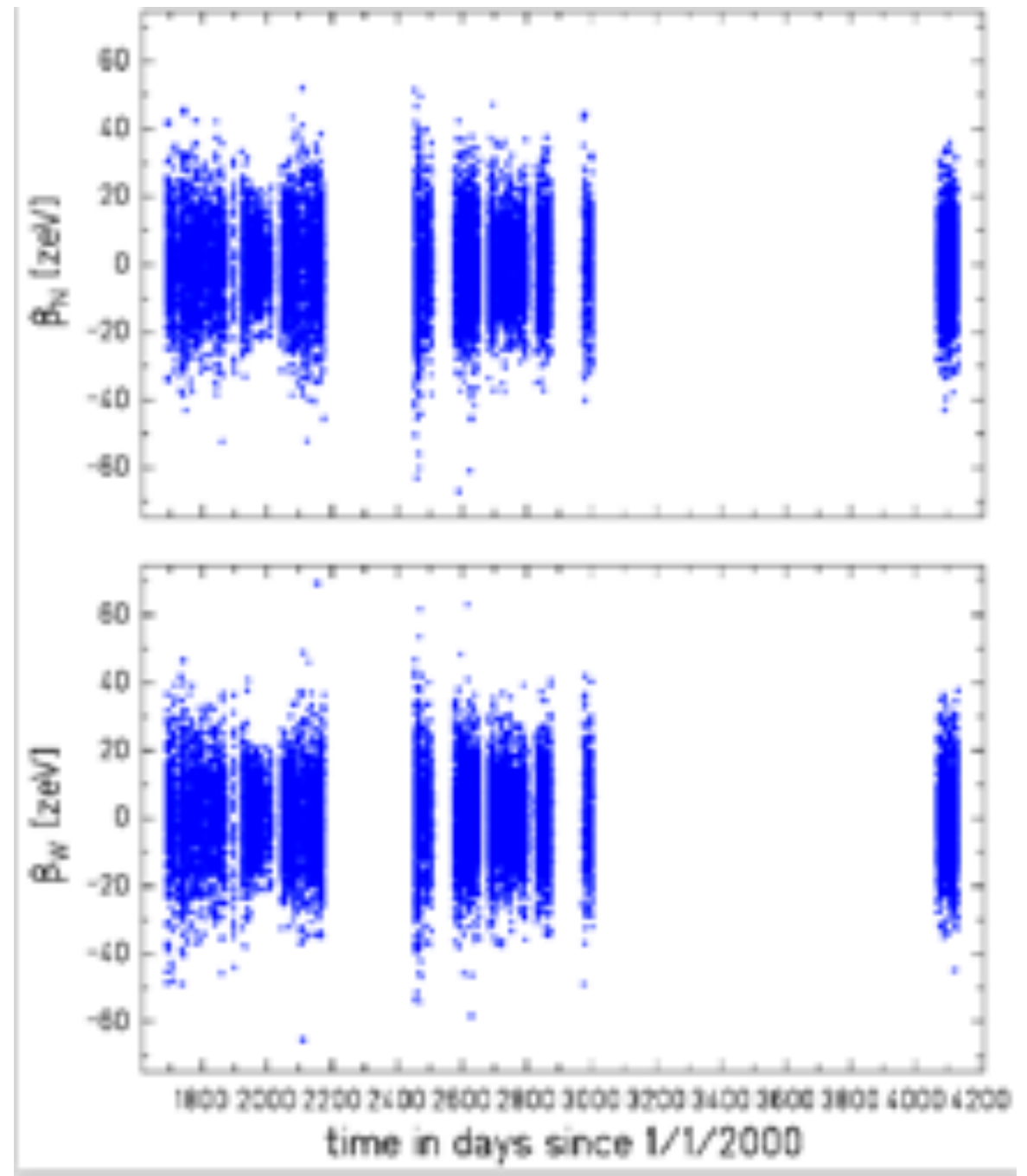
- Turntable reduces $1/f$ noise
- sidereal allows long-term stability



Analysis strategy

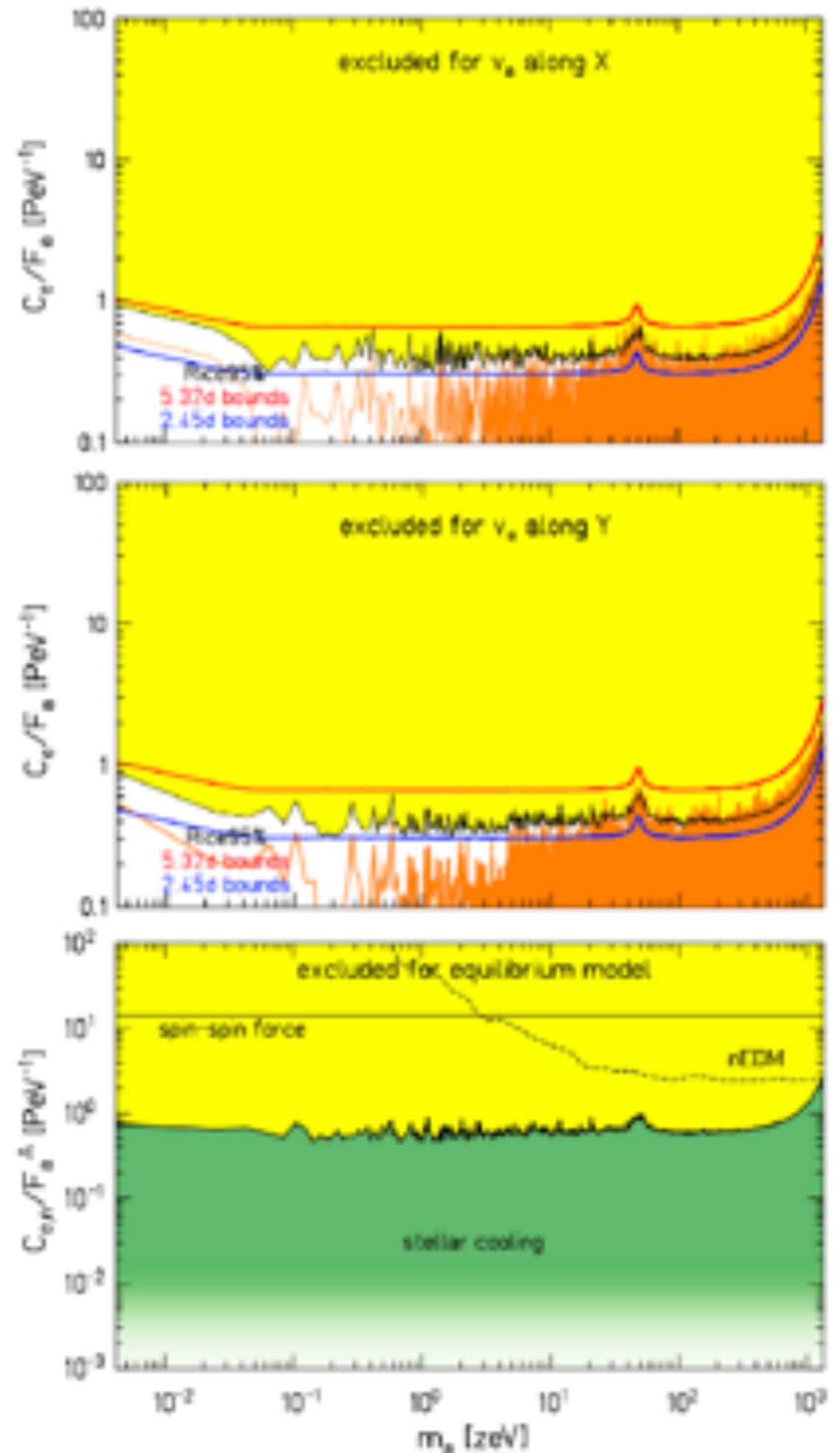
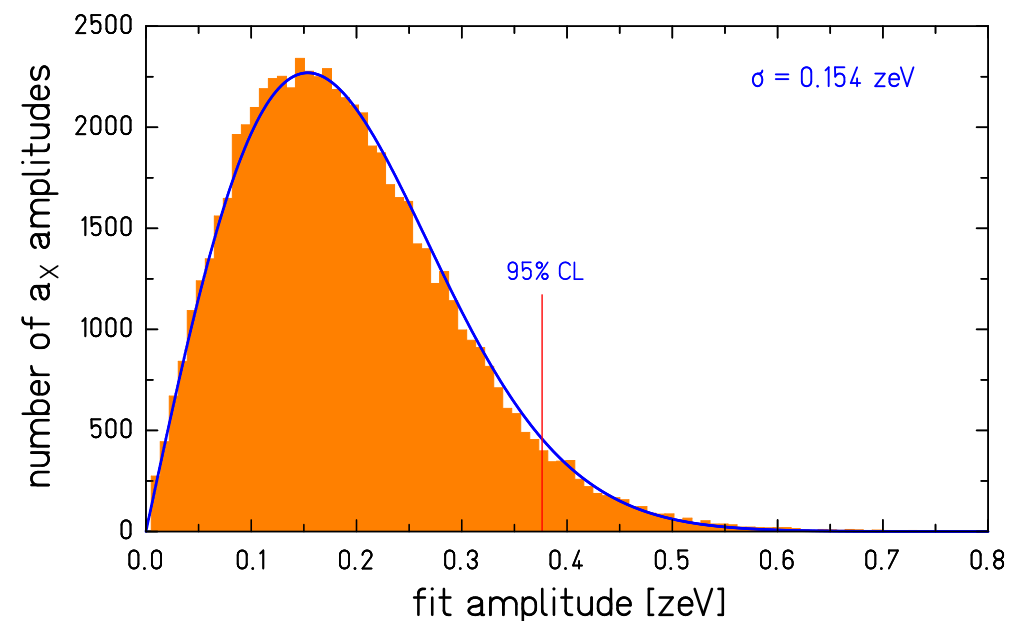
1. Extract pendulum torque at turntable frequency using 6.7 year span of data
2. Linear fit of resulting data to basis functions that include sidereal and axion oscillations
3. Repeat for 67,200 axion frequencies in our band (10^{-9} Hz — $3.2 \cdot 10^{-4}$ Hz)

$$\beta_N^i = b_{XN\cos}^i a_{X\cos} + b_{XN\sin}^i a_{X\sin}$$



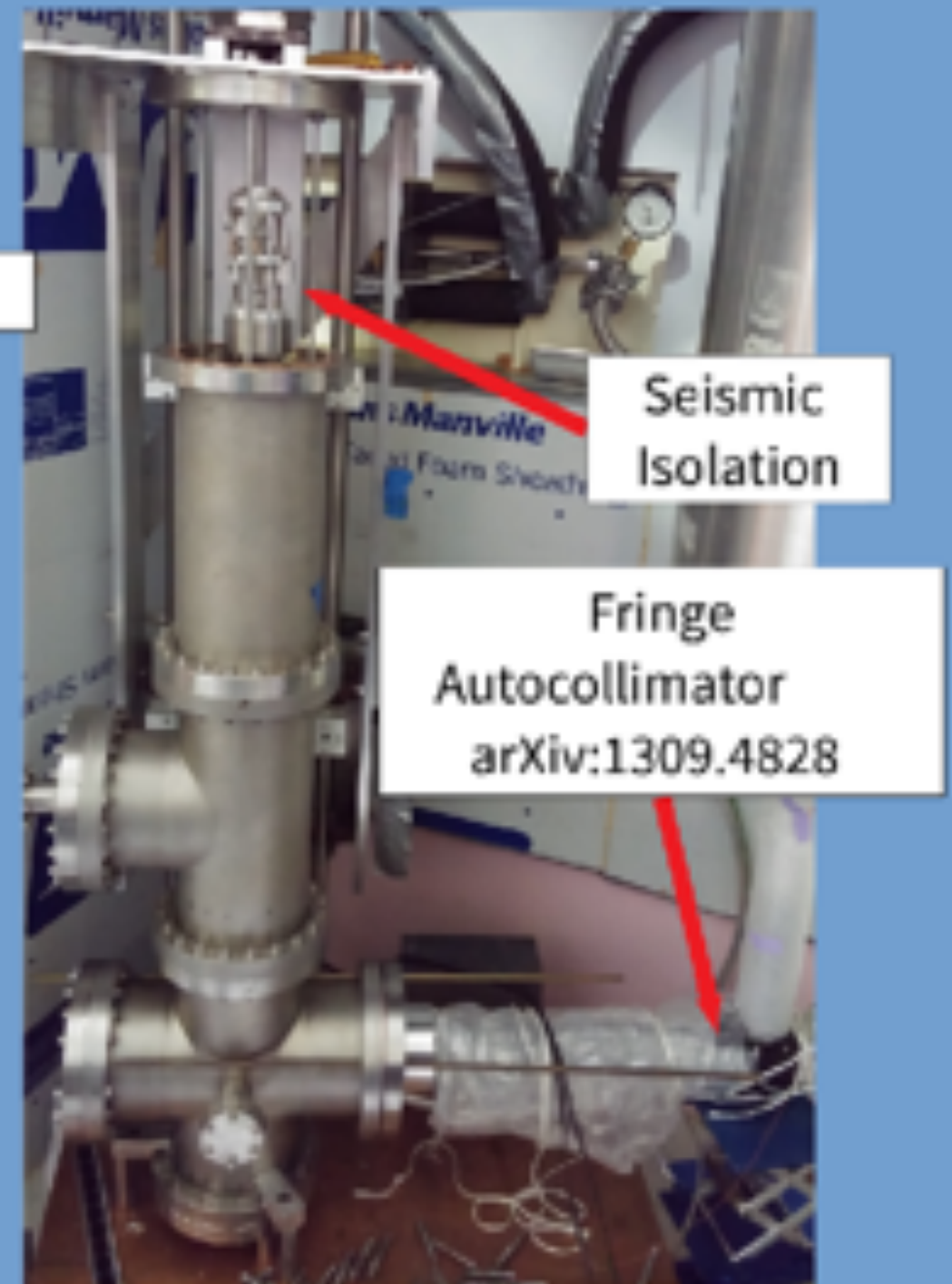
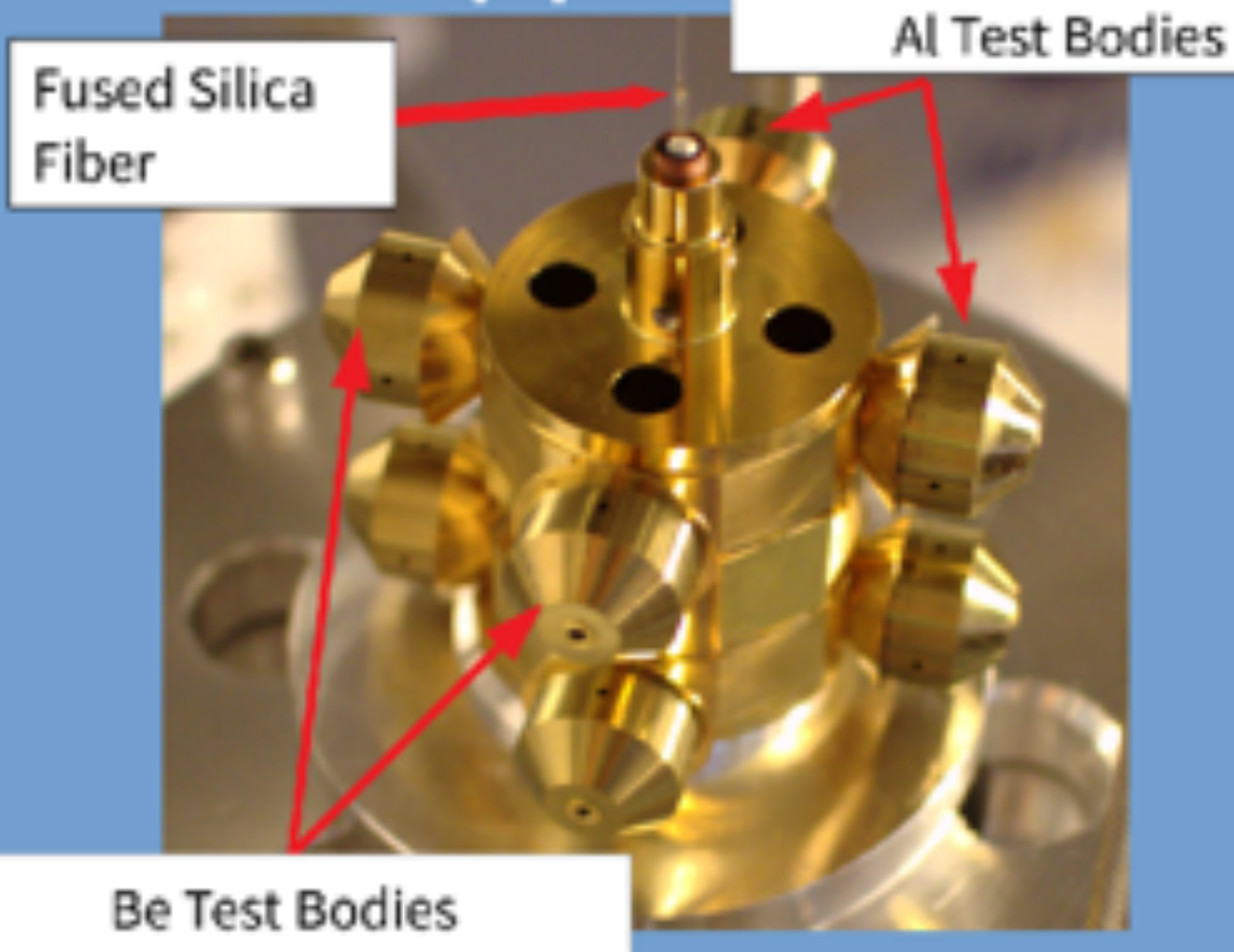
Results:

Error bar set by scatter of the 67,200 measurements



B—L (Erik Shaw)

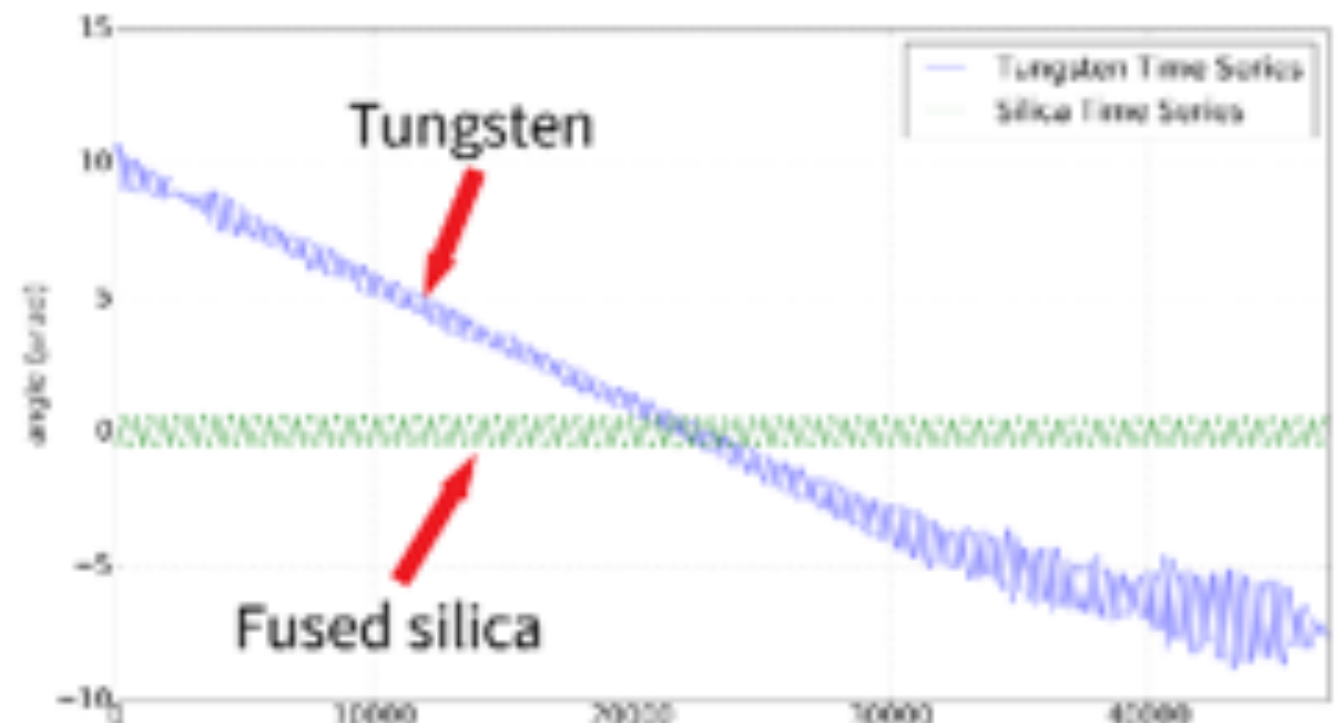
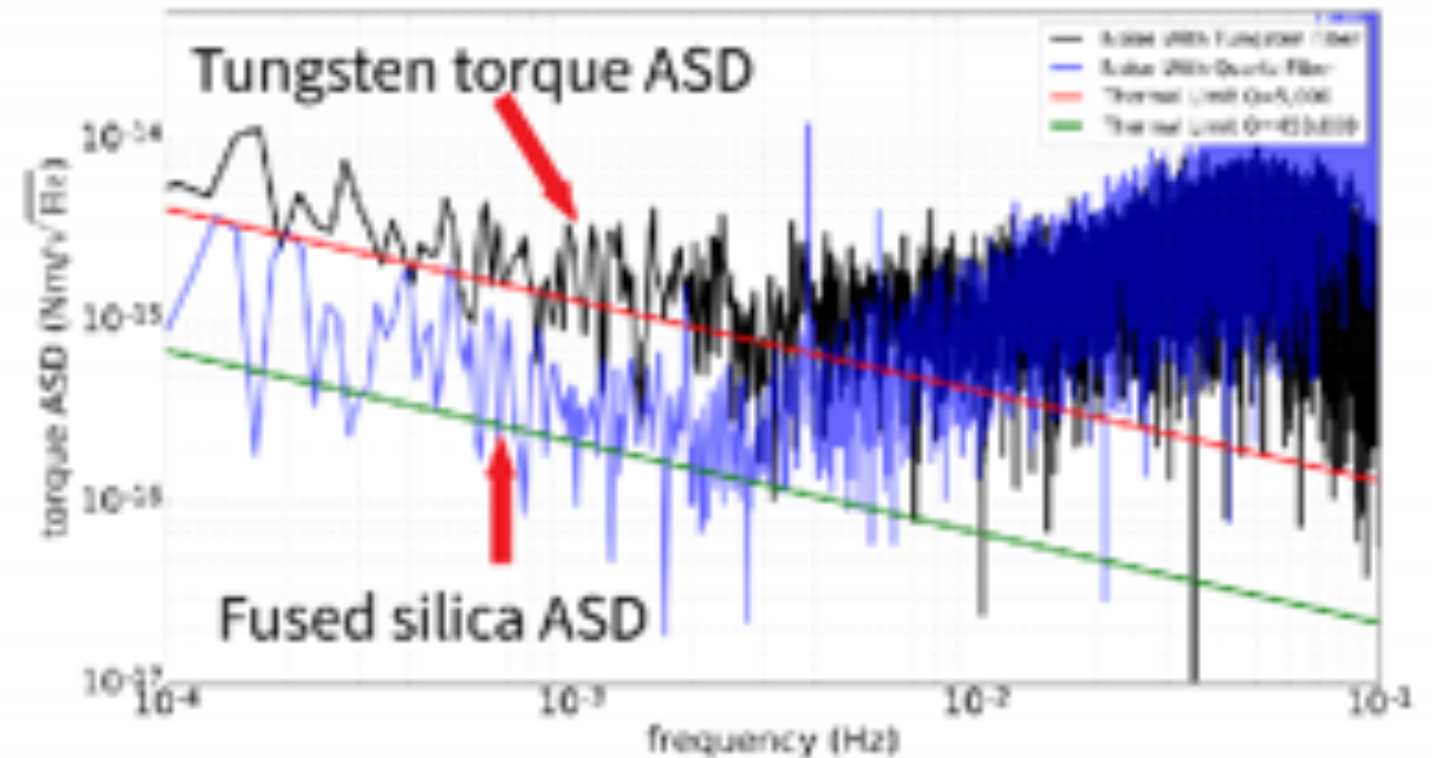
Non-Rotating Apparatus



B—L (Erik Shaw)

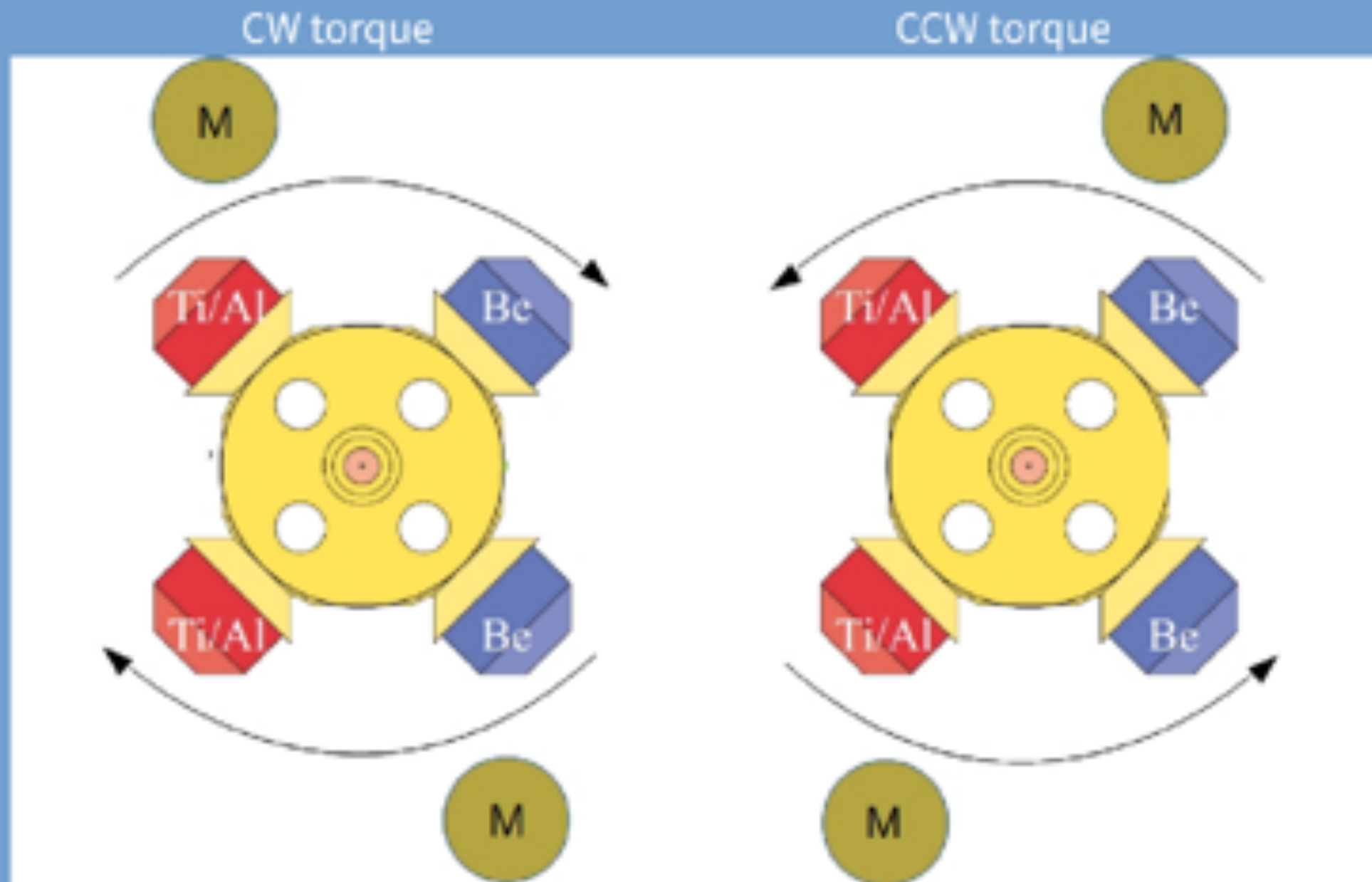
Silica fibers

- High Quality factors improve sensitivity 5x
- Noise near thermal limit
- Fiber drifts reduced more than 10x
- More improvement is likely

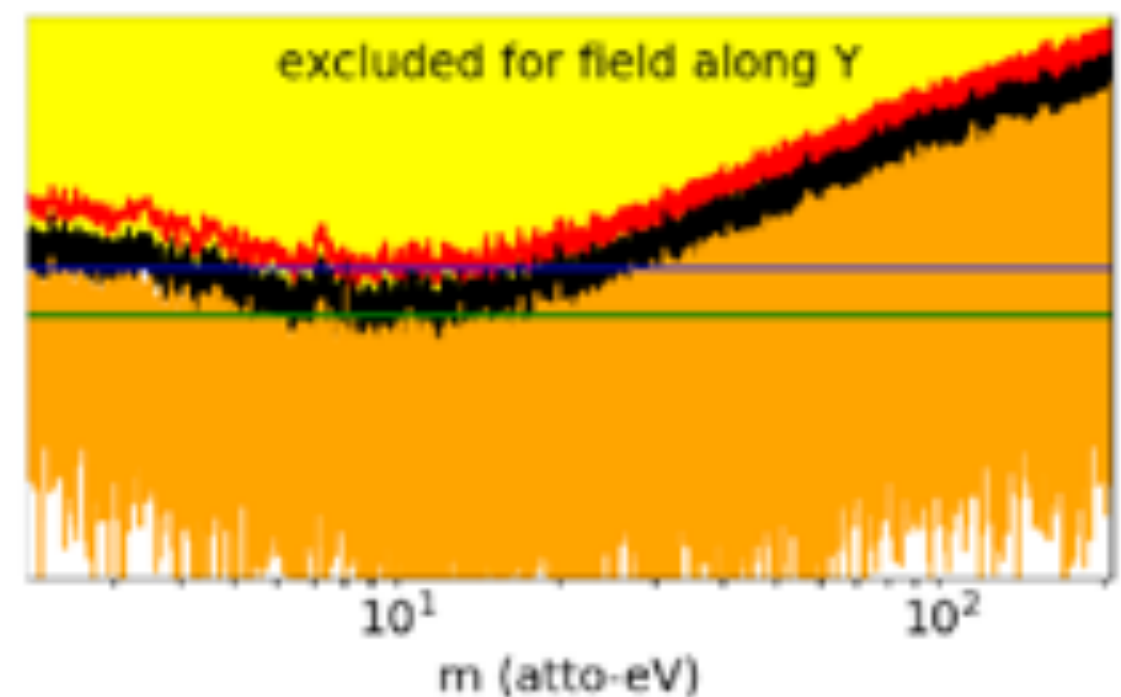
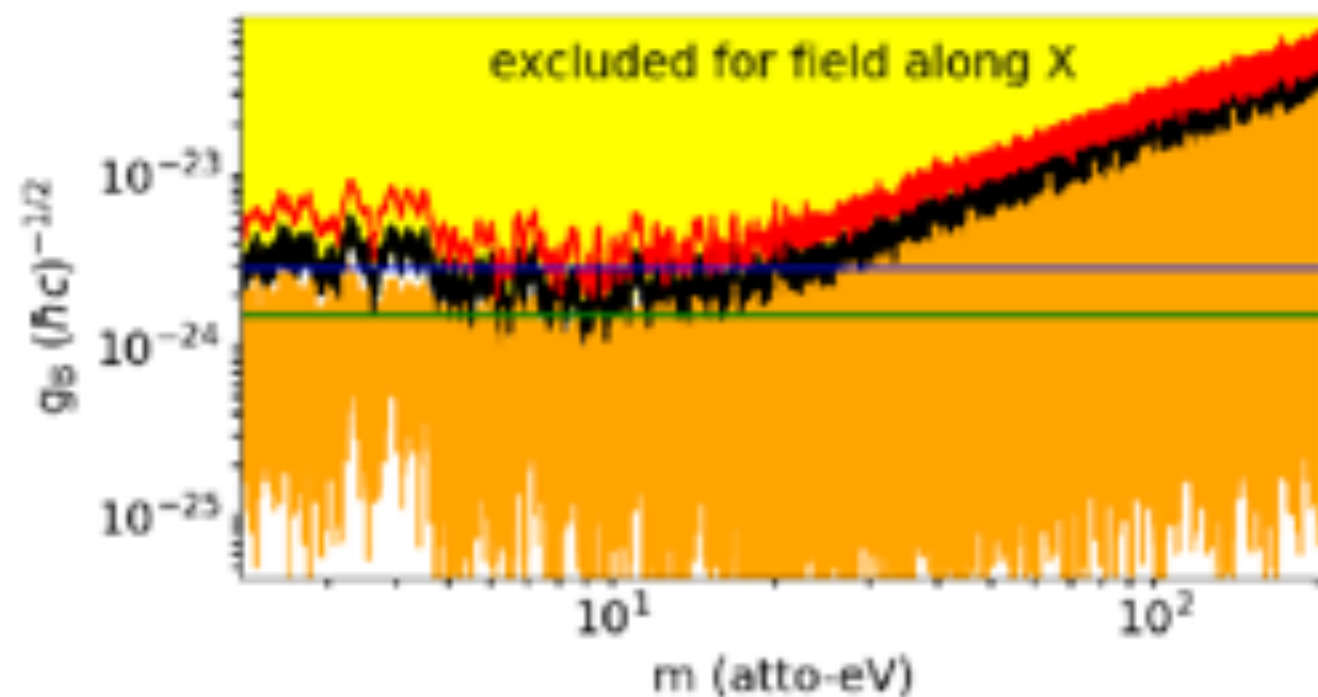
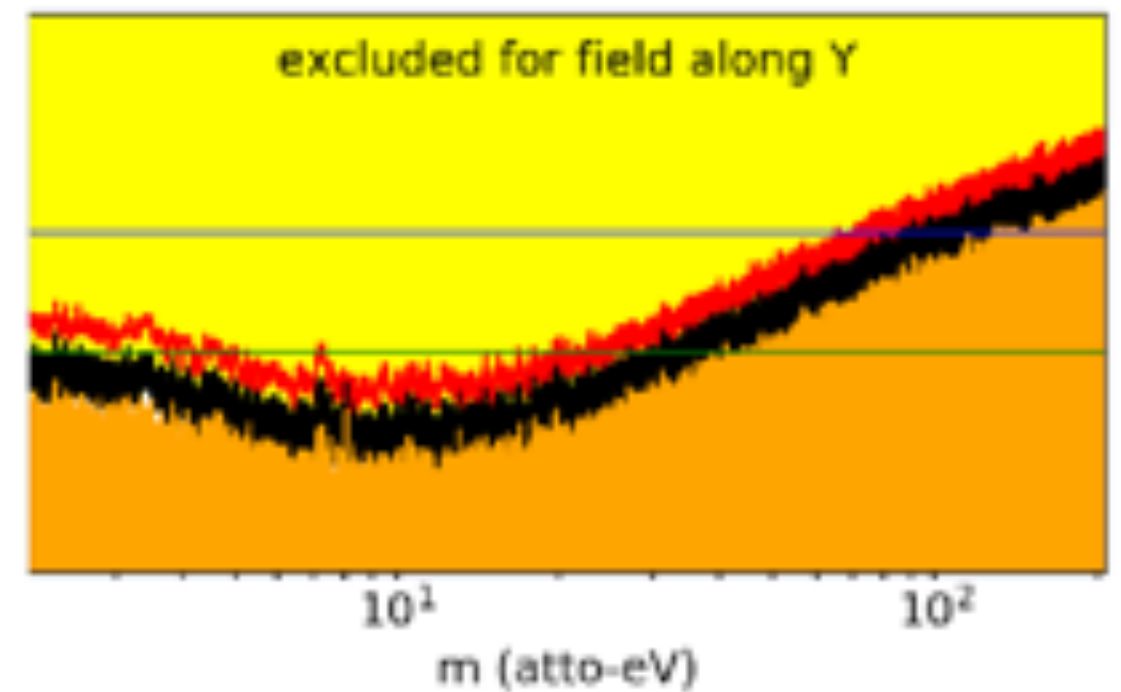
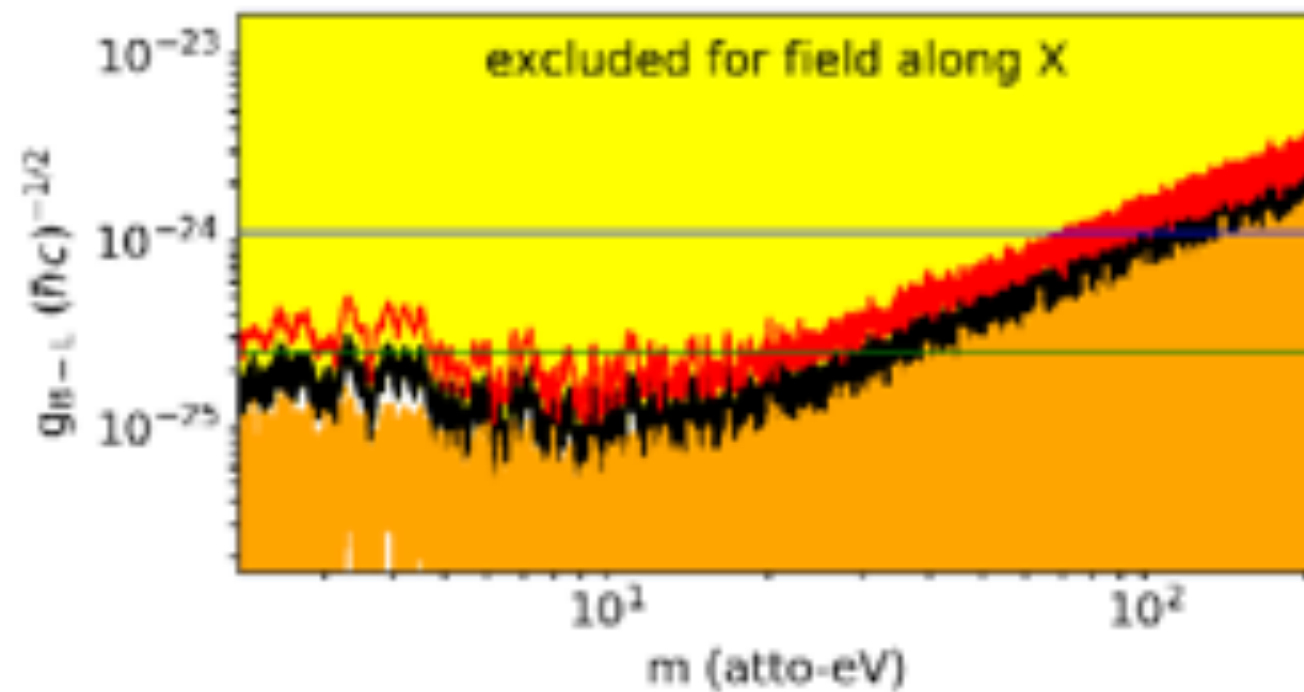


B—L (Erik Shaw)

Gravitational Damping



B—L (Erik Shaw)



B—L (Erik Shaw)

Moving Forward

- There are a variety of ways to expand this search:
 - Improve apparatus sensitivity with improved fibers and angular readout
 - Use pendulum with better composition dipole
 - Use rotating apparatus to extend search to lower masses

Nuclear-spin interactions

- Much larger spin/magnetic ratio
- Probes higher energy scales for given coupling

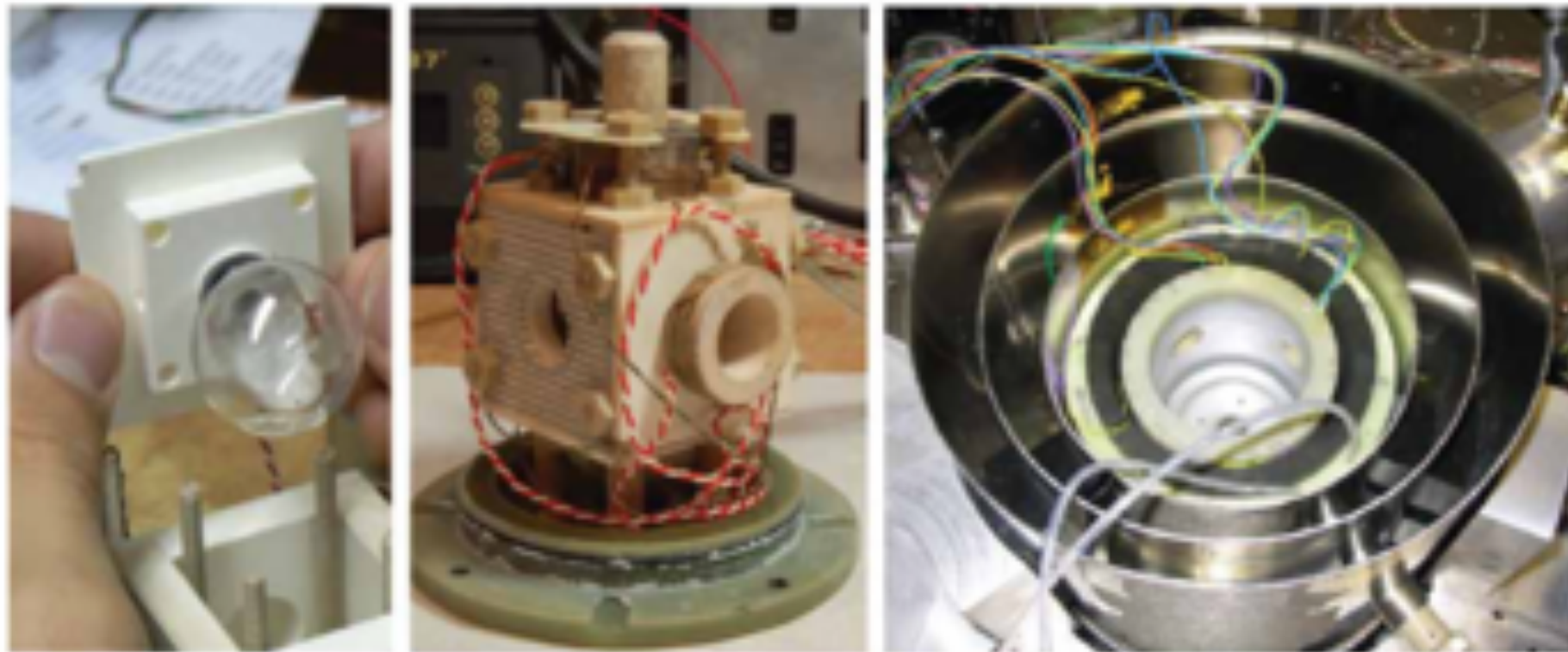
$$g_p = \frac{m_\psi}{F}$$

- Thermal polarizations (earths field) $\sim 10^{-10}$
- hyper-polarization provides $\sim O(10-50\%)$

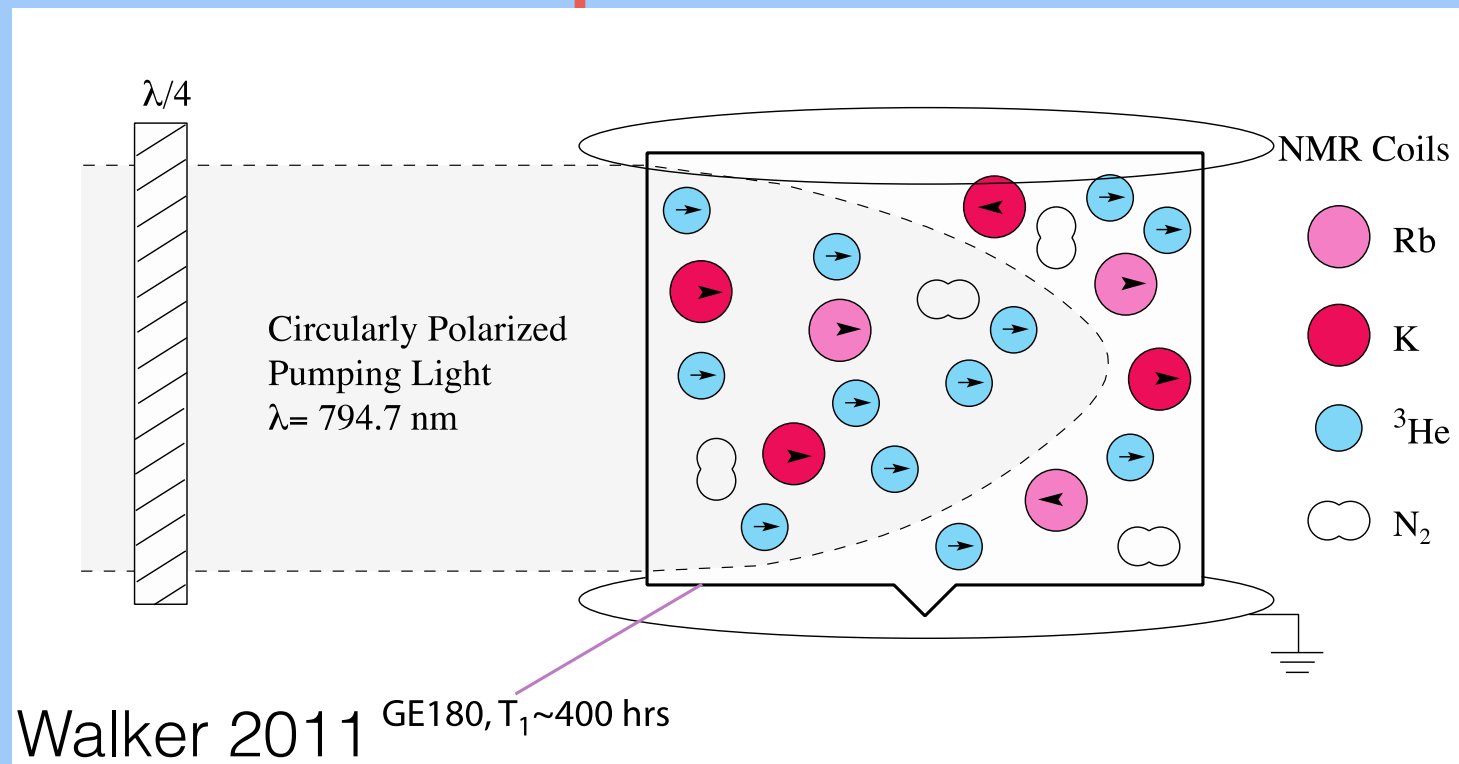


Princeton Axion Project

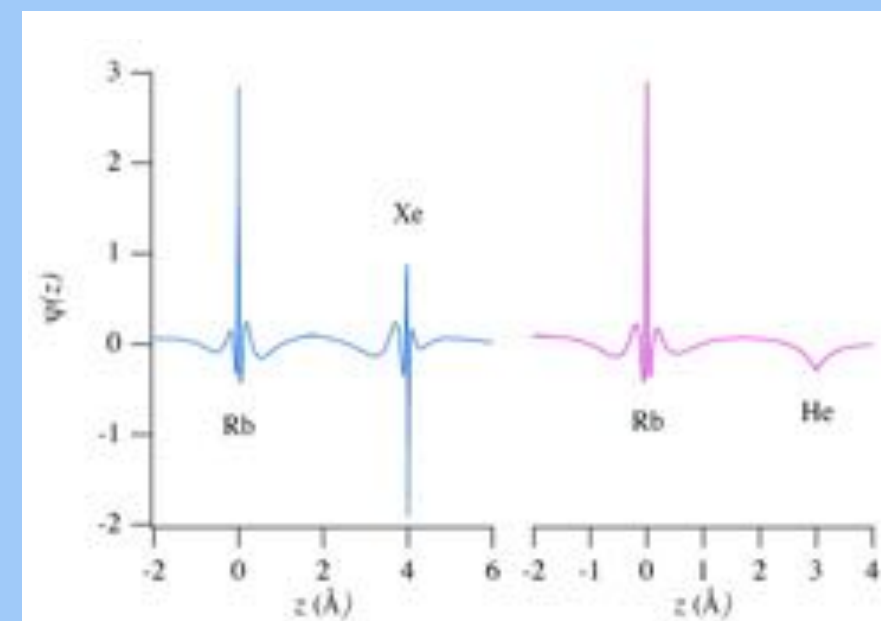
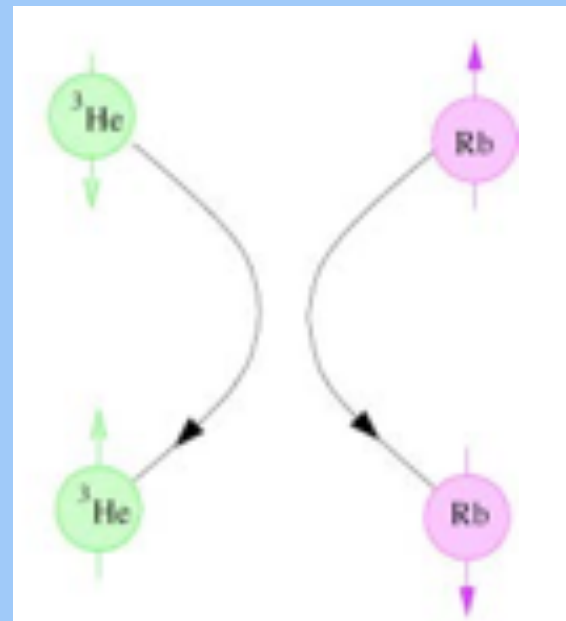
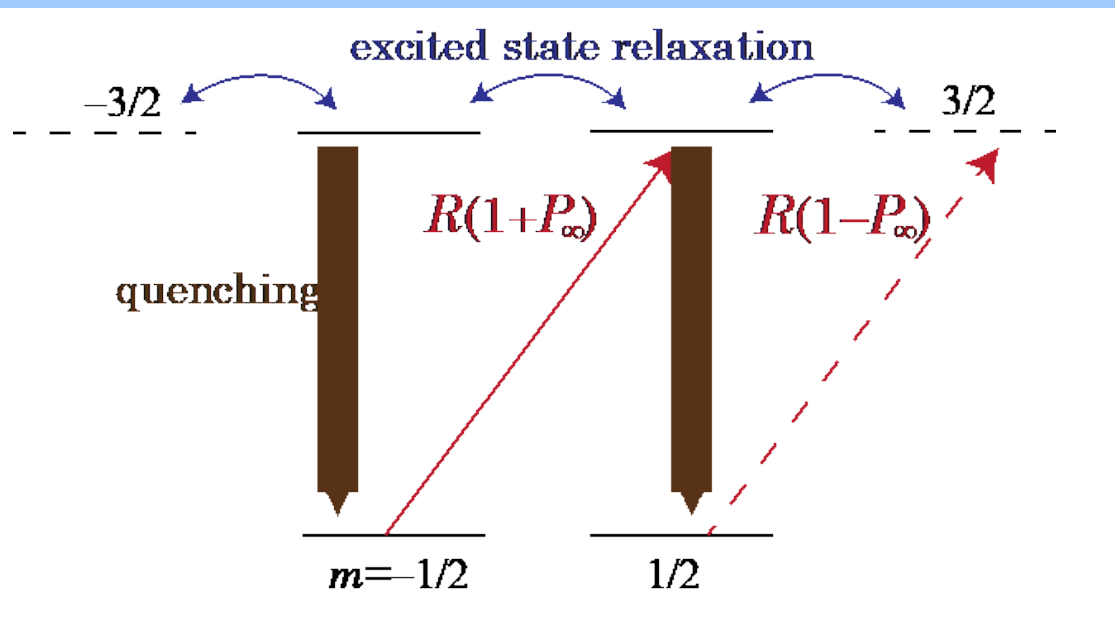
- Main issues:
 1. cancel magnetic perturbations
 2. keep sensitivity to non-magnetic interactions
- I. Compare two nuclear spins in same volume
- II. Compare electron-spin to nucleon spin



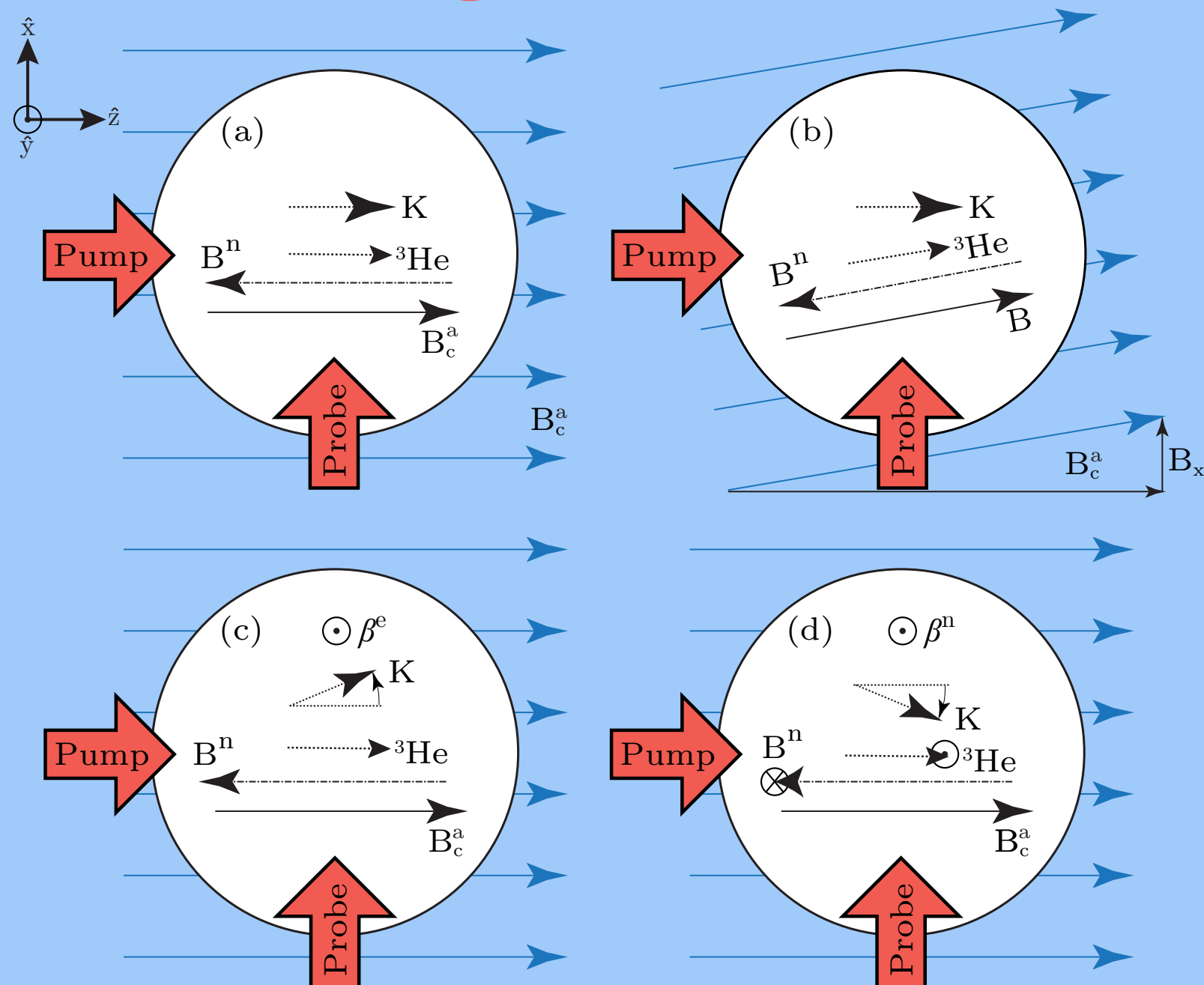
Nuclear spin hyperpolarization



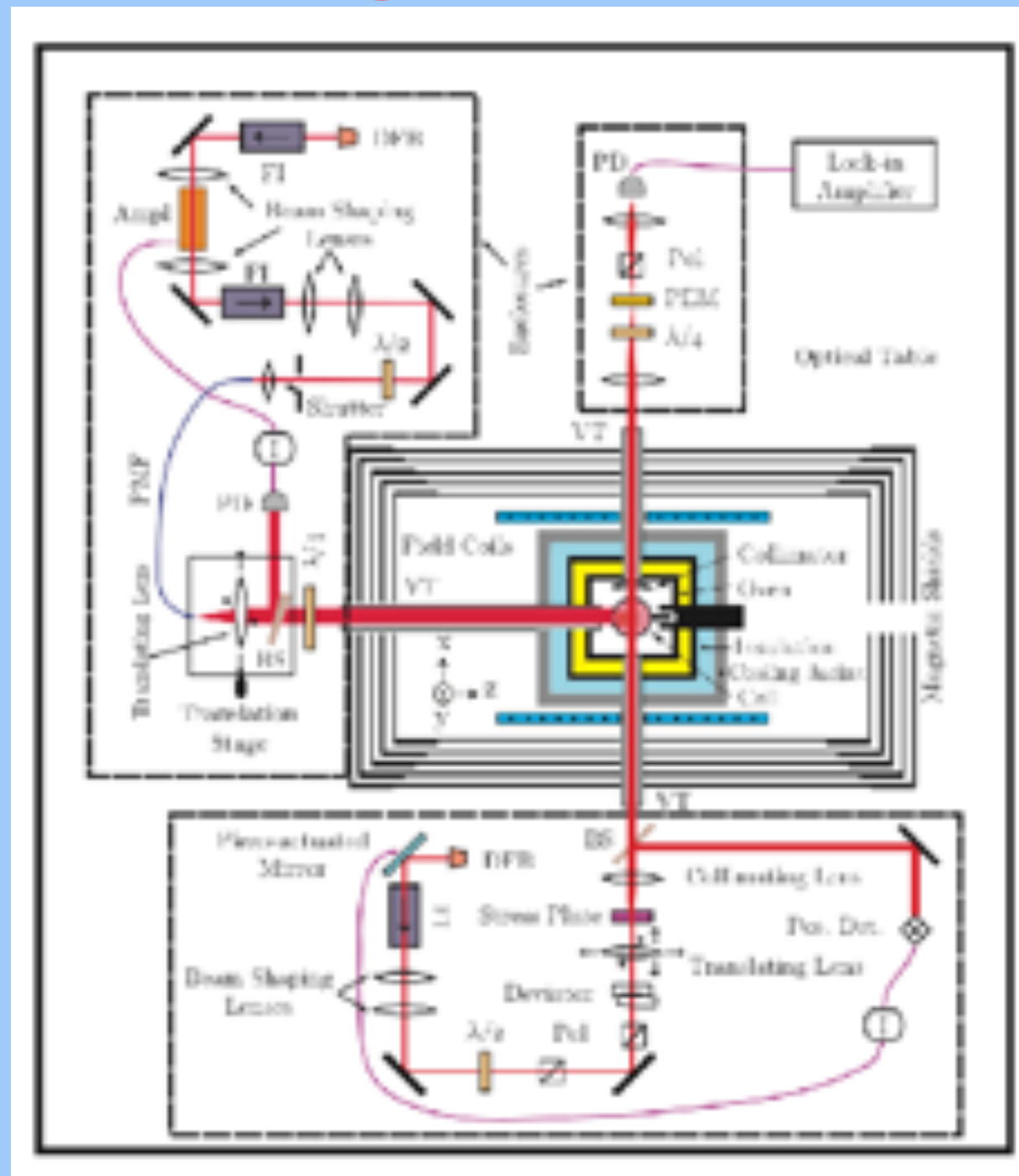
angular momentum
from laser \rightarrow
electron \rightarrow
nucleon



electron-nucleon comagnetometer

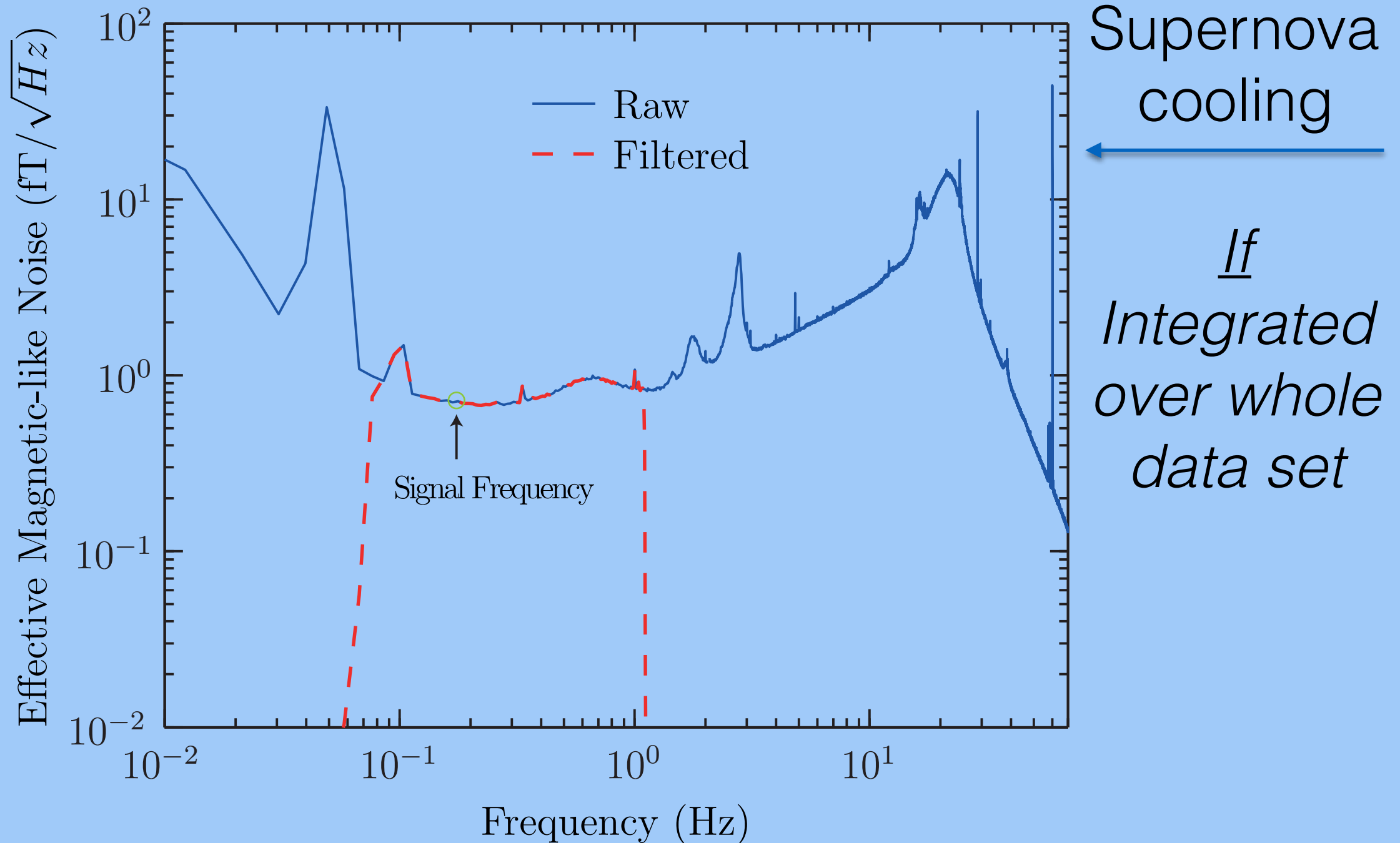


electron-nucleon comagnetometer



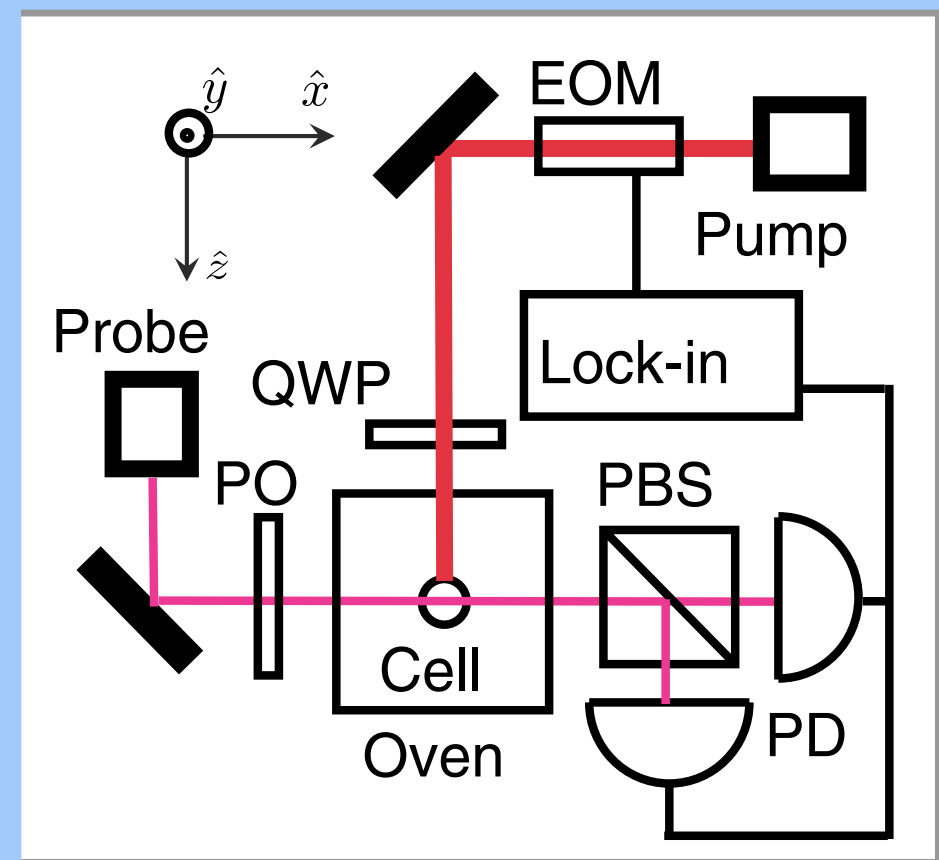
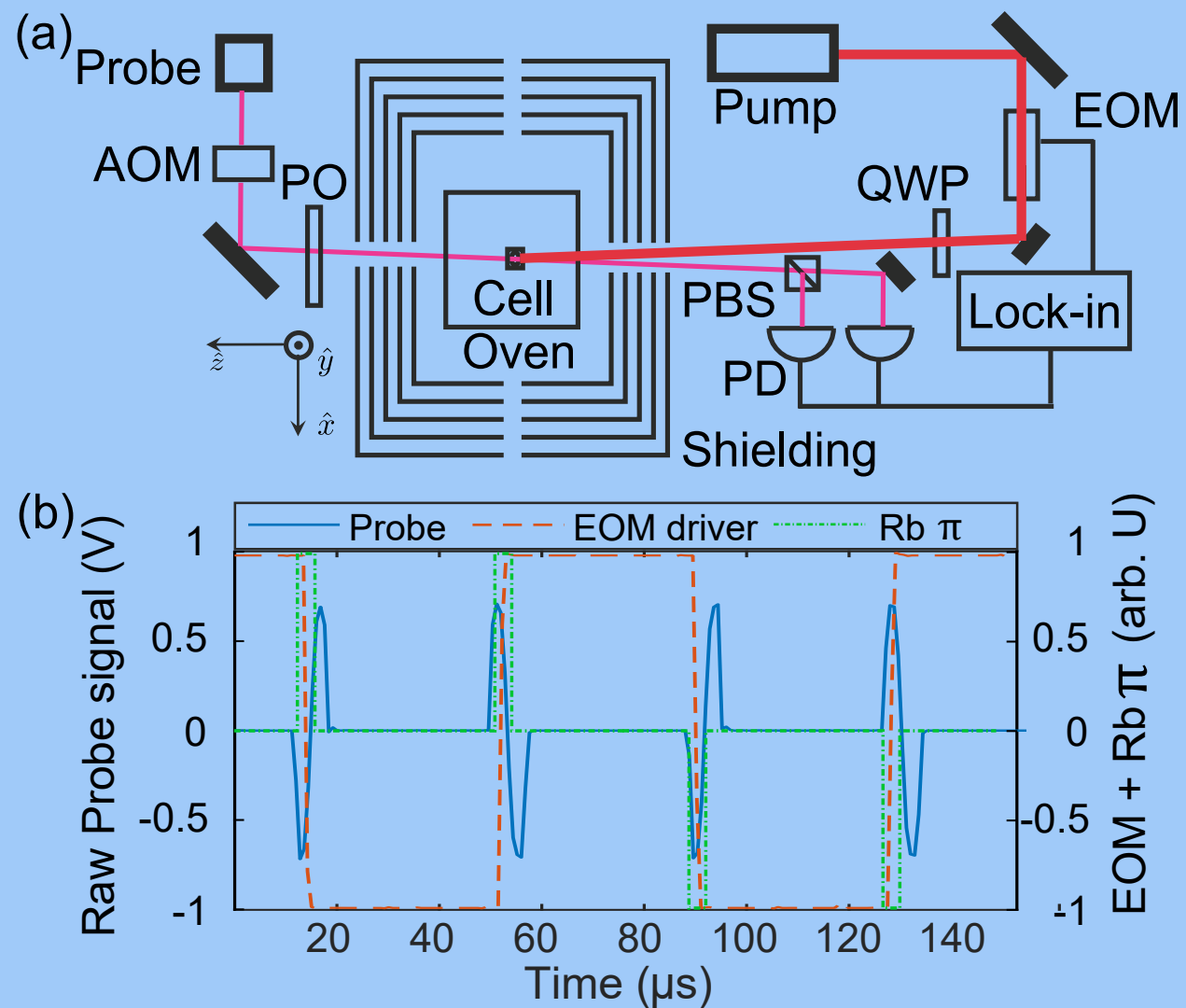
So simple!

Existing data



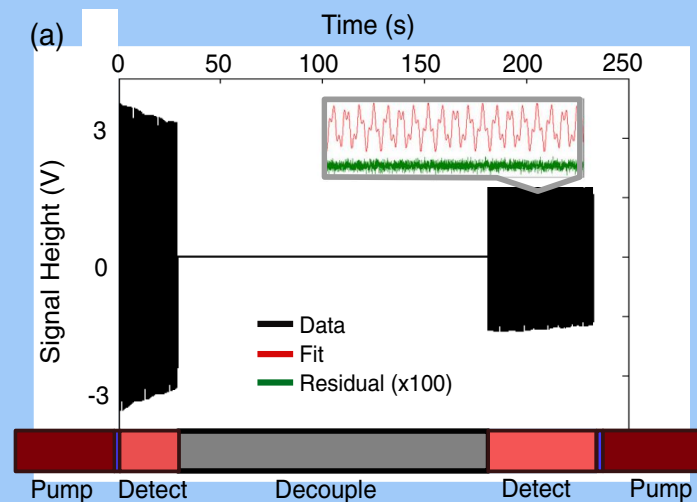
He-? gyroscope with Rb pump & Rb read-out

Developed with Xenon, now using Neon

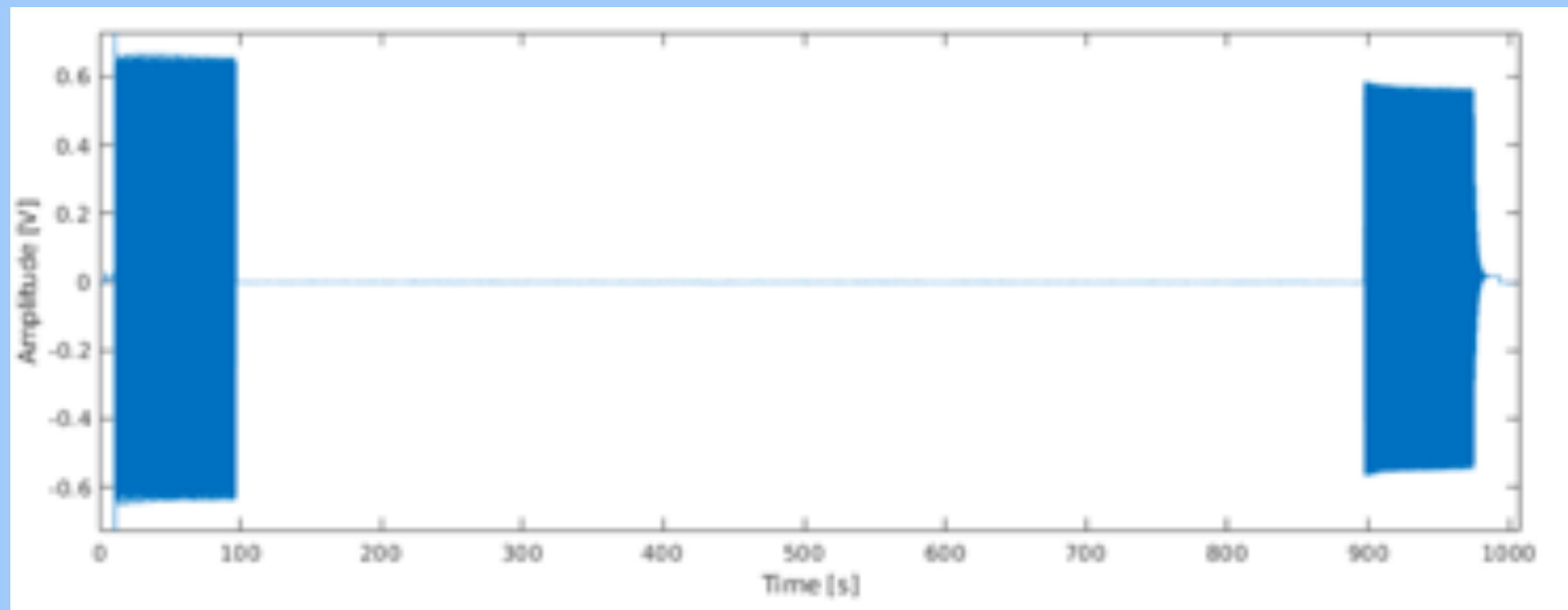


Limes et al 2017

Current status

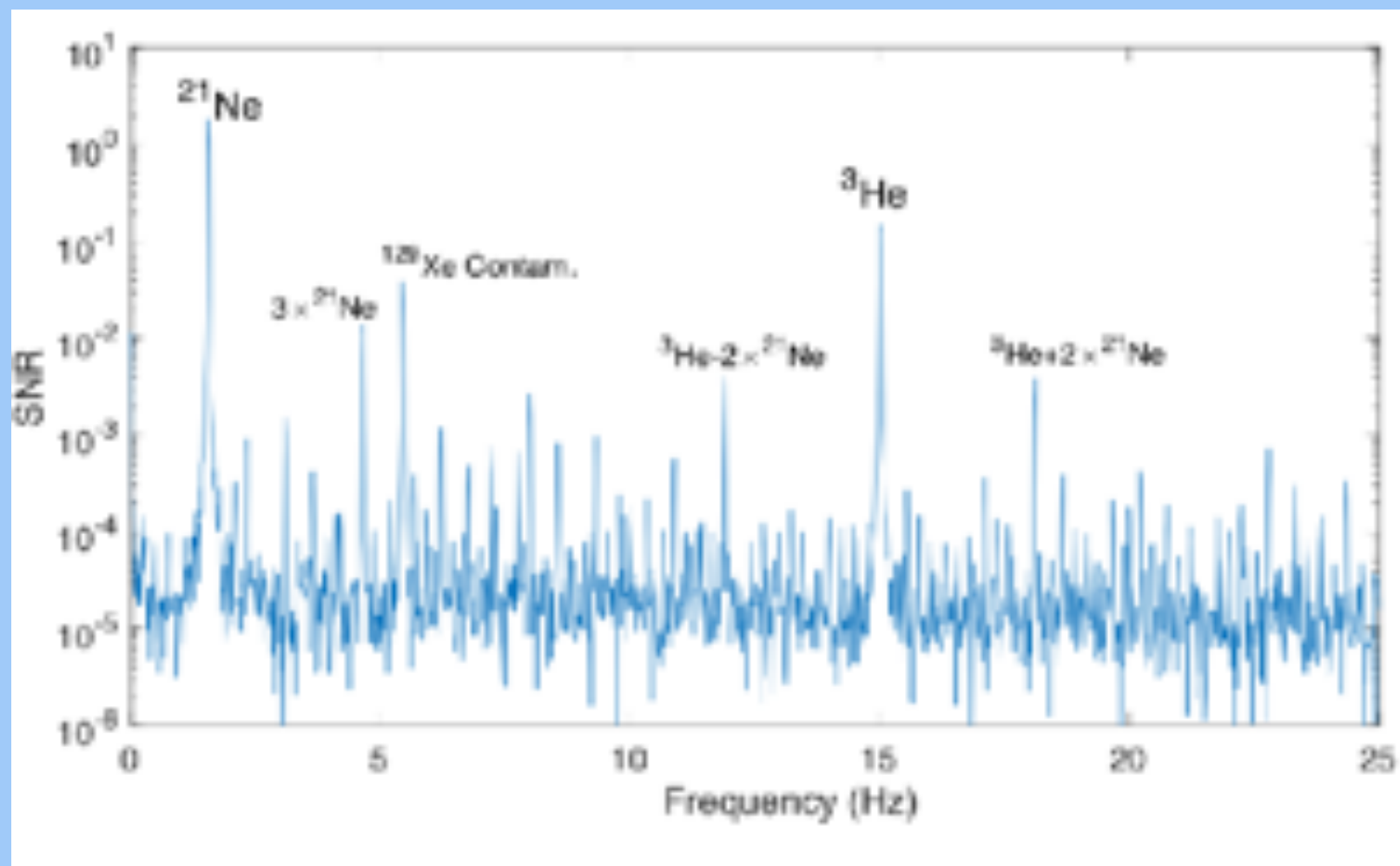


Replaced Xenon with Neon
(lower Rb interactions)

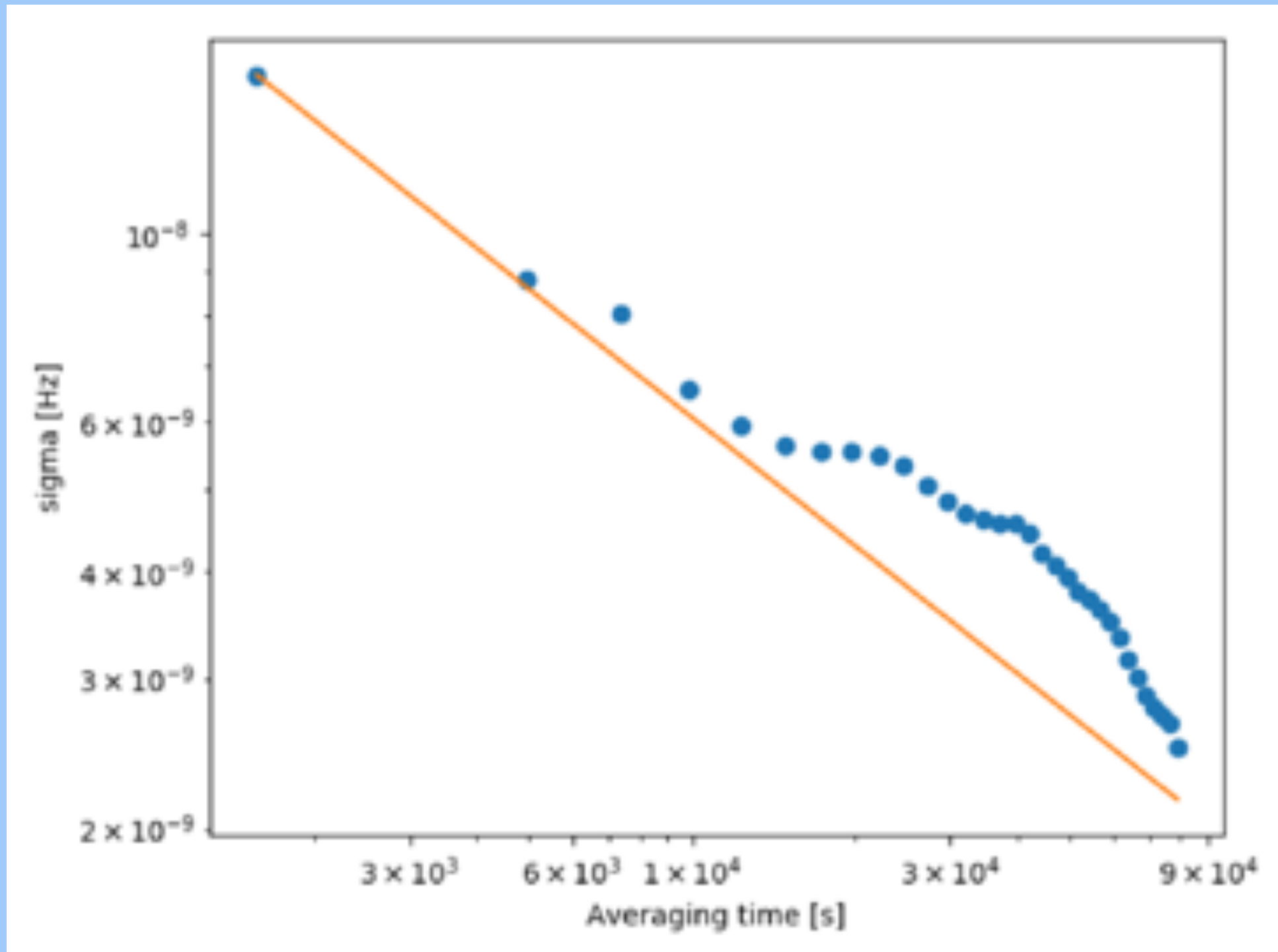


SNR is not a problem

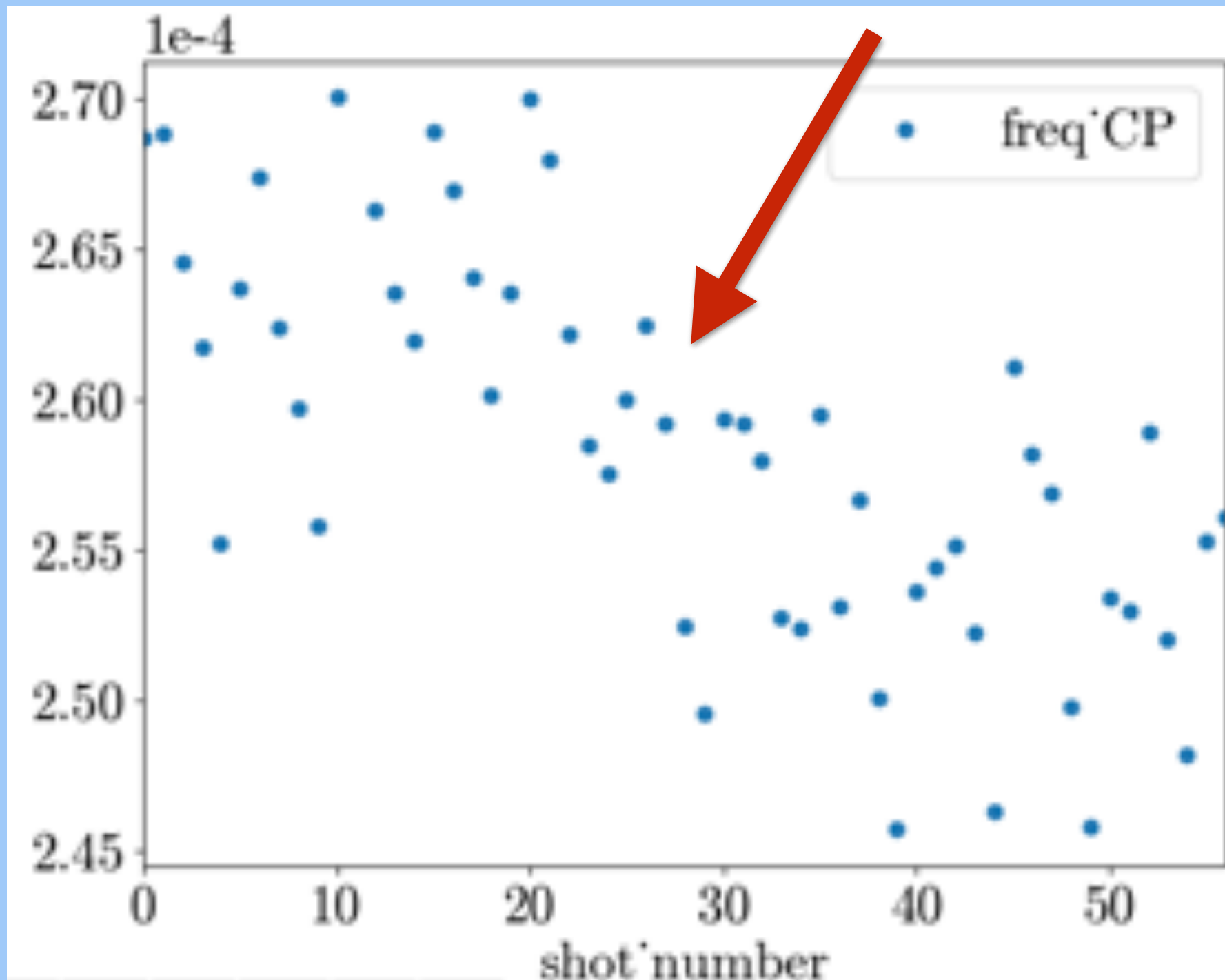
SERF - magnetometer non-linear response



Current stability

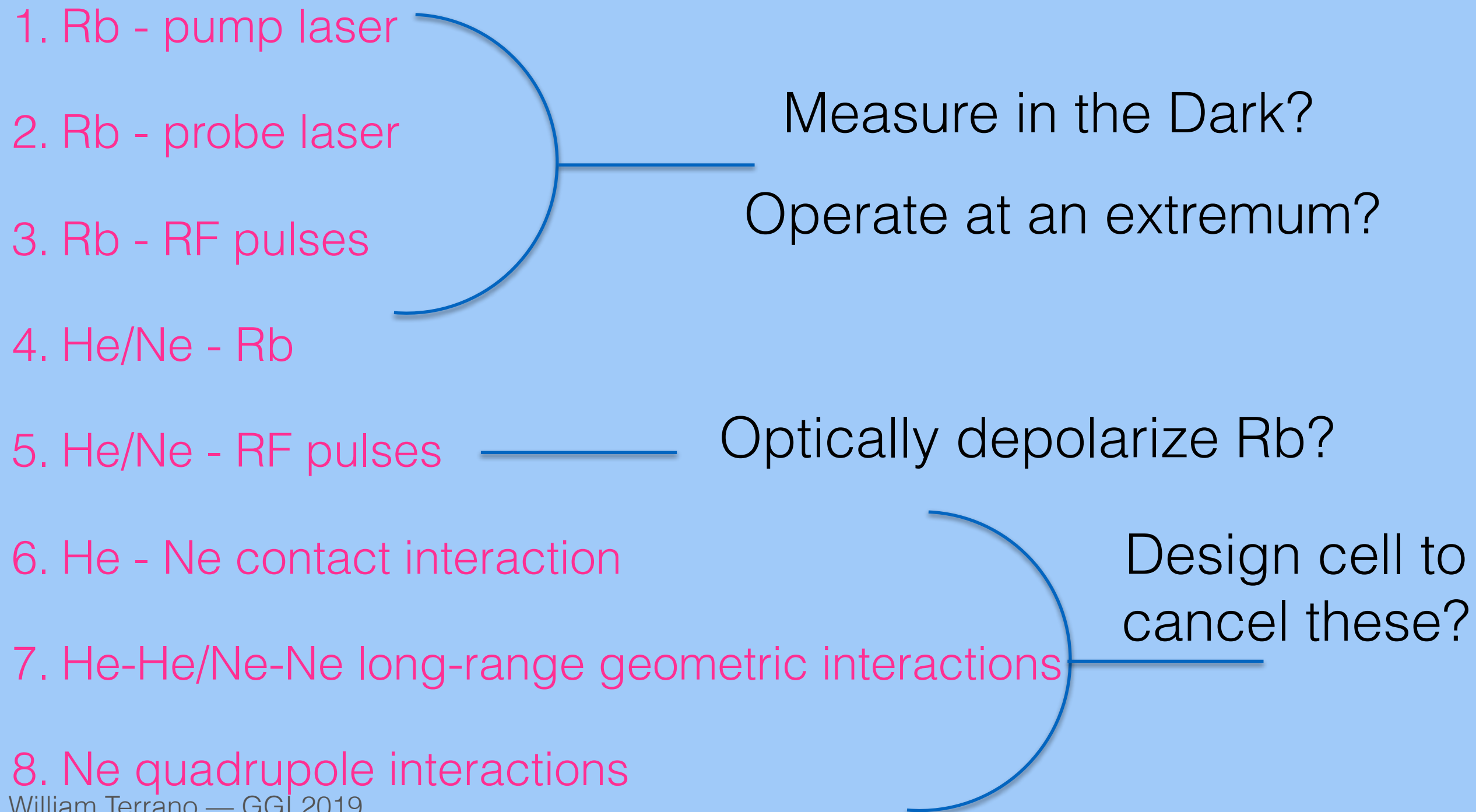


Current Problem

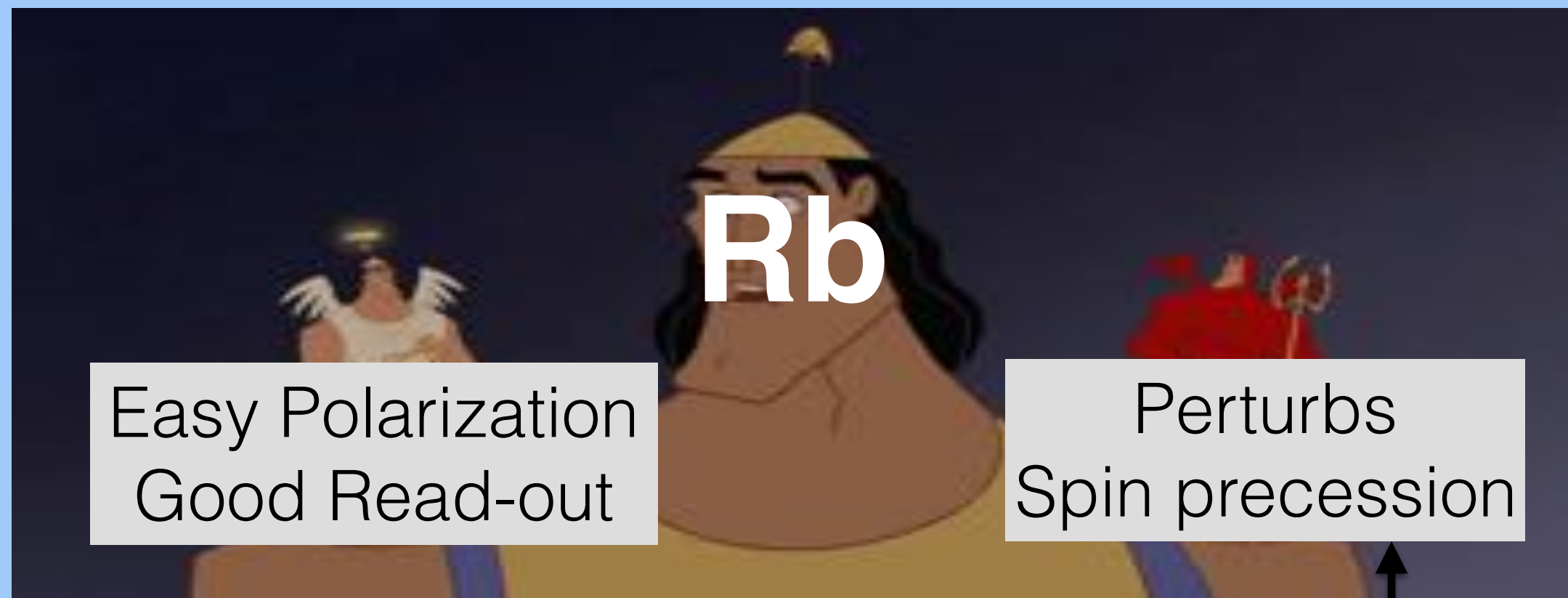


What changed?

Plenty of fun to have



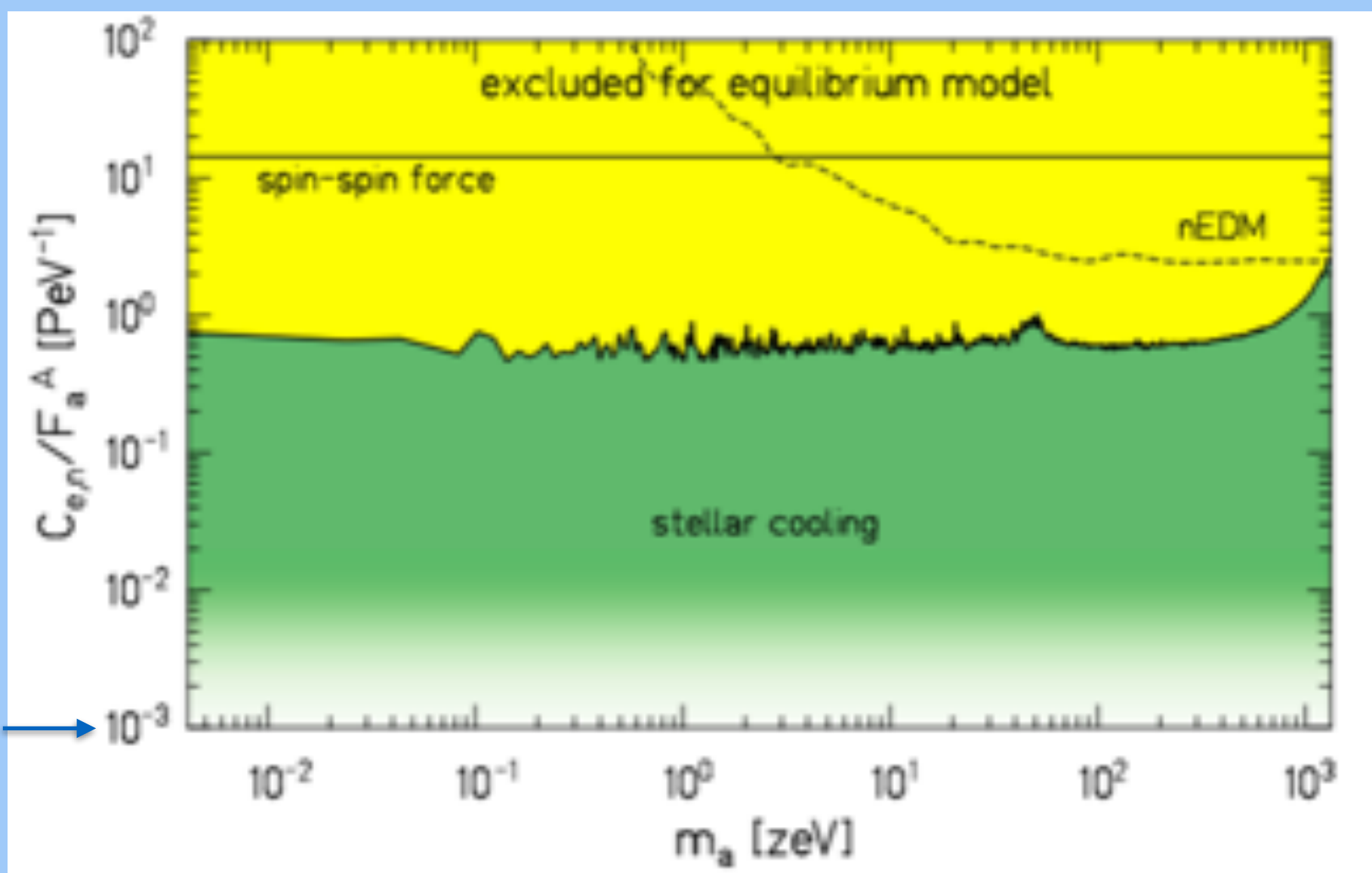
Zooming out on the problems



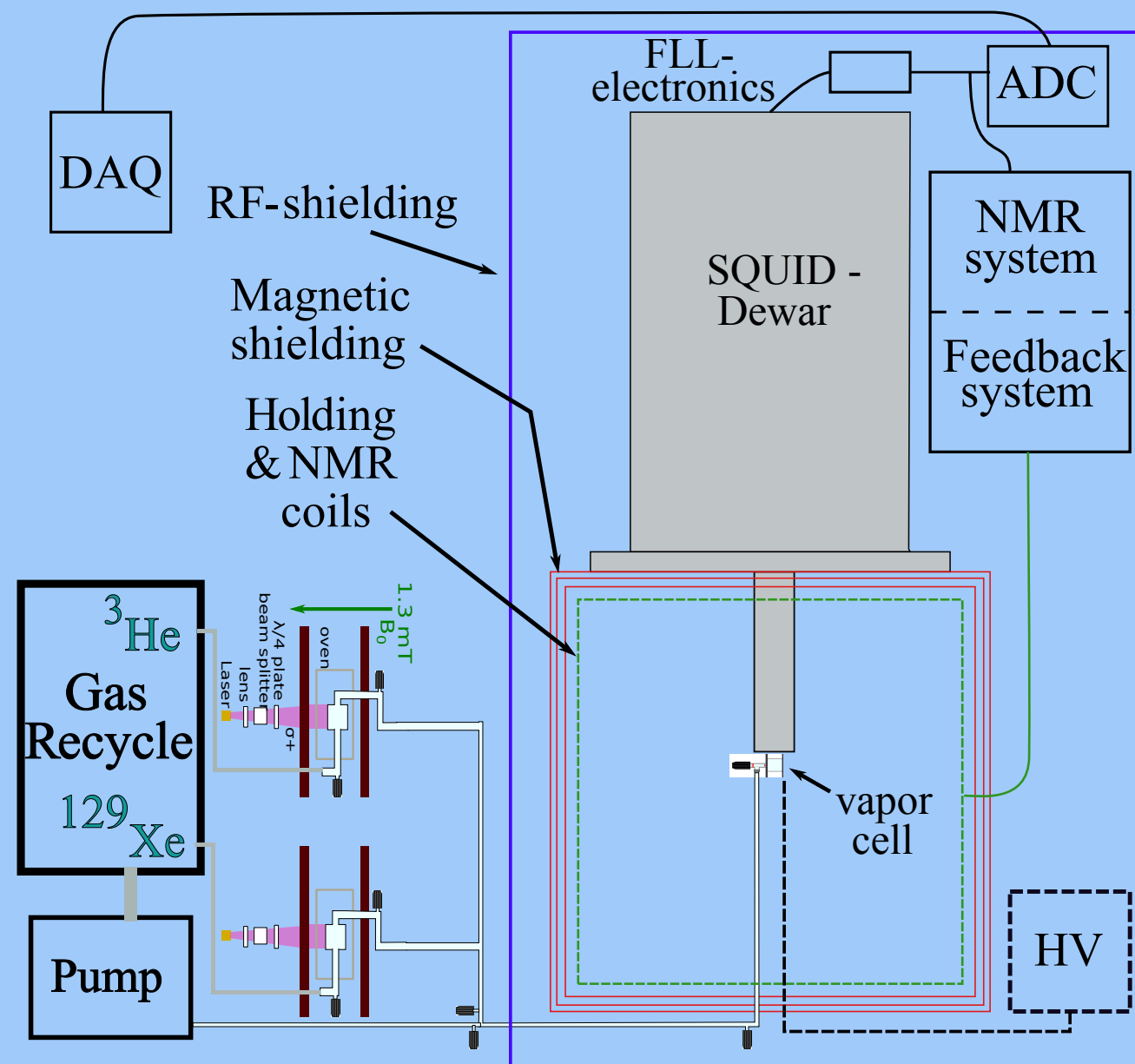
- Self-interacting system:
Energy splitting depends on the superposition you are in. $O(10^{-3})$ Suppression Required

$$\mathcal{H}_{\text{ex}} = - \sum_{i < j} J_{ij} \mathbf{S}_i \cdot \mathbf{S}_j$$

10^9 GeV →



Rb-interaction free system



Potential for Ultra-light Axion search

Astro-physics

Reach F_{DM}
 10^9 GeV

He-Ne Goal $2 \cdot 10^{10} \text{ GeV}$

2nd Gen $1.2 \cdot 10^{14} \text{ GeV}$

2nd Gen
 T2 5hrs: Xenon-dimer
Adjustable volume (ppm) and fidelity (ppm)

Commercial Gear

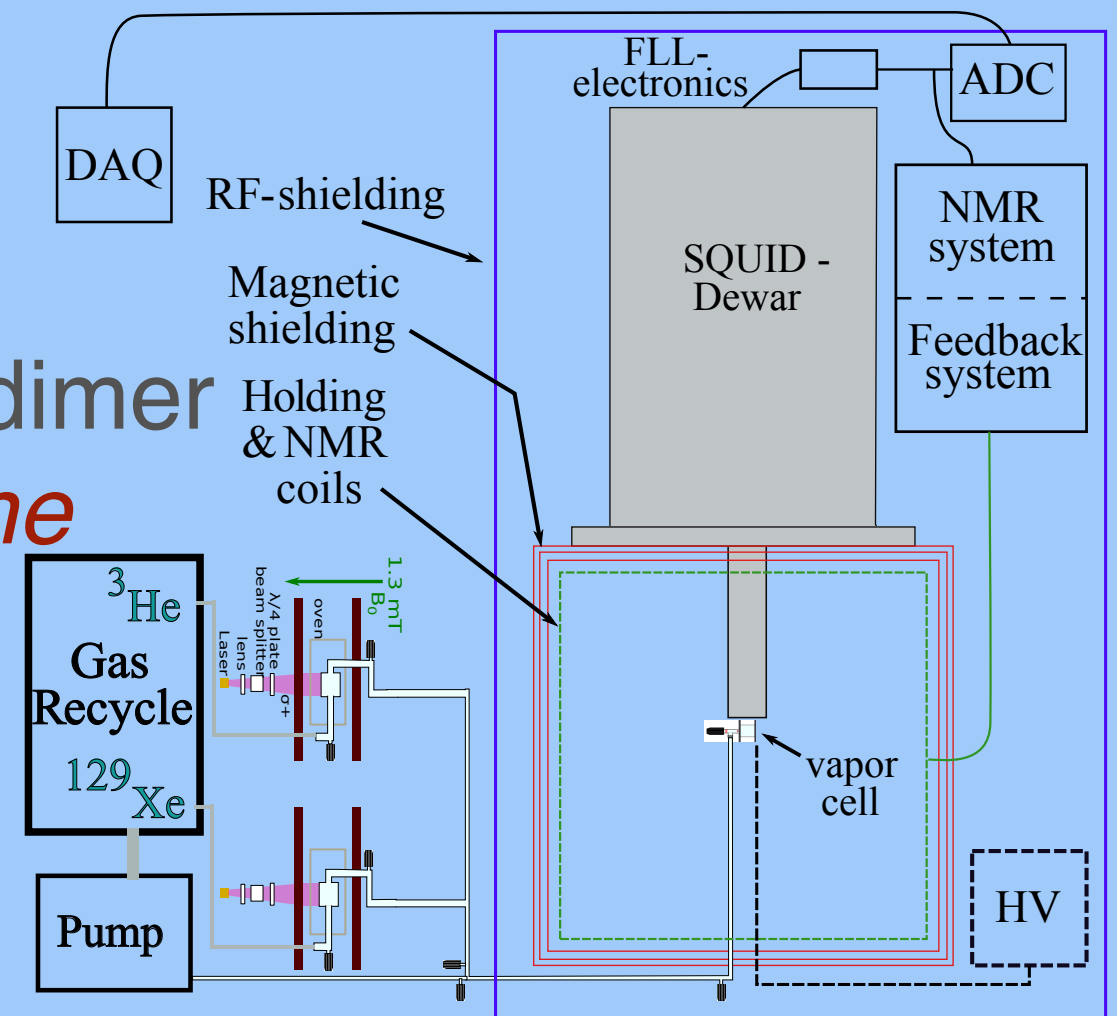
Density: 1 amagat

Polarization: 10%

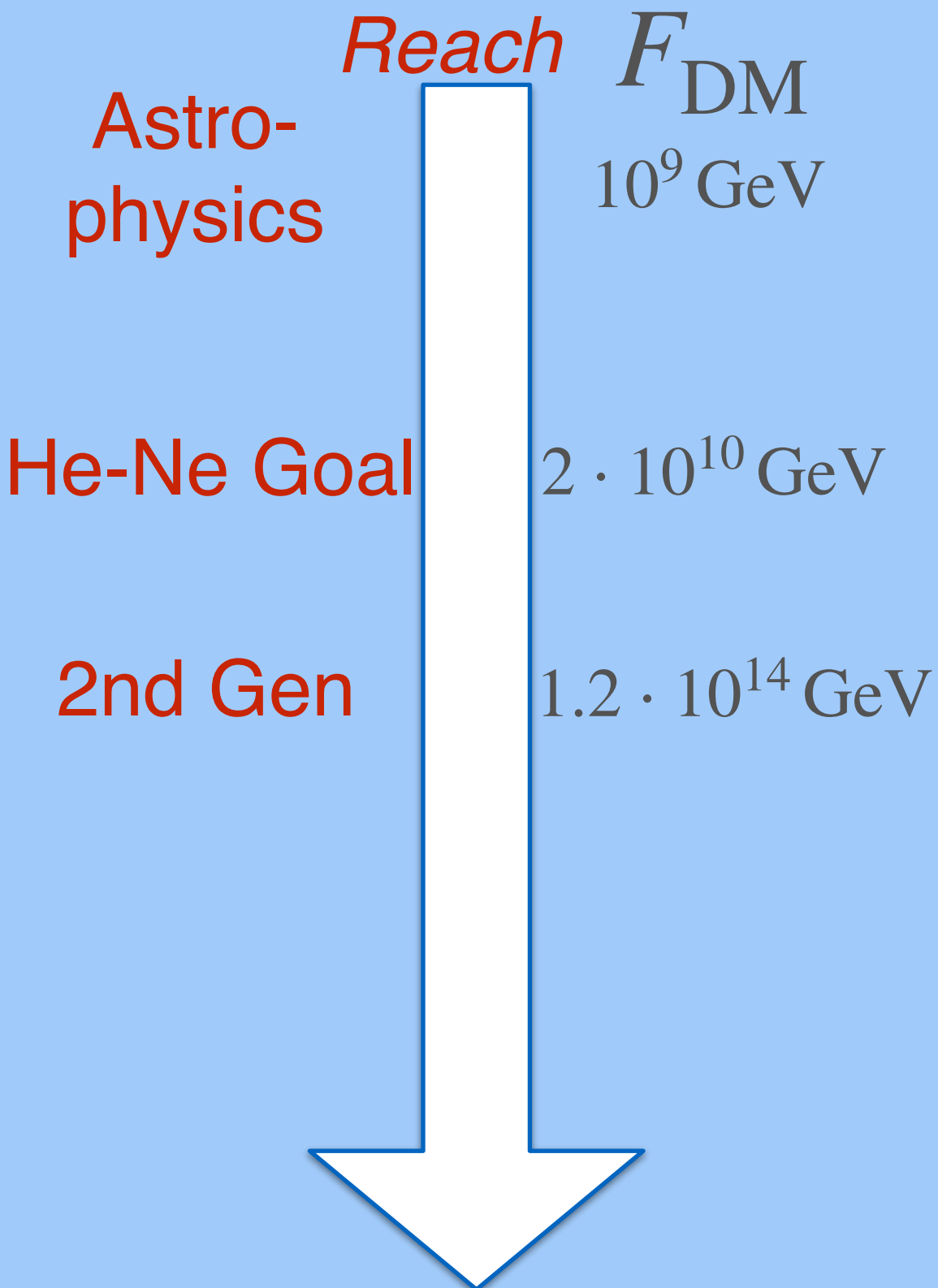
Cell size: 5 x 4.285 cm

SQUID noise: 0.180 fT

SQUID distance: 4 cm



Towards GUT scale



Beyond

Decay time T_2 limited by Xenon-dimers

- *Run hotter ($T_2 \times 10?$) or*
- *Higher pressure ($\times 20$)*

SQUID noise limited by pickup-loop inductance:

- *Custom SQUID coil ($\times 7$) &*
- *Better coupling ($\times 5$)*

Big hammer:

- *50 cm cell ($\times 20$), gradients at $n\text{EDM}$ levels*

Thanks

- Eot-wash group
- Romalis group — DARPA & ONR
- Humboldt foundation & Dicke Fellowship