

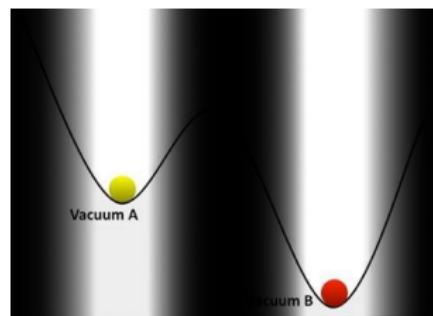
On Vacuum Stability without Supersymmetry

Brane dynamics, bubbles and holography

Ivano Basile | SNS, Pisa | *Cortona Young, May 2020*

based on:

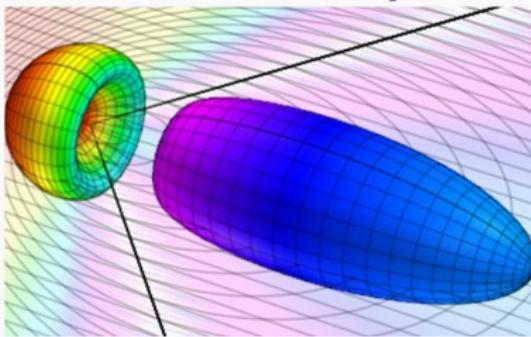
- [hep-th/1811.11448](#)
with J. Mourad and A. Sagnotti
- [hep-th/1908.04352](#)
with R. Antonelli
- [hep-th/1806.02289](#)
with R. Antonelli and A. Bombini



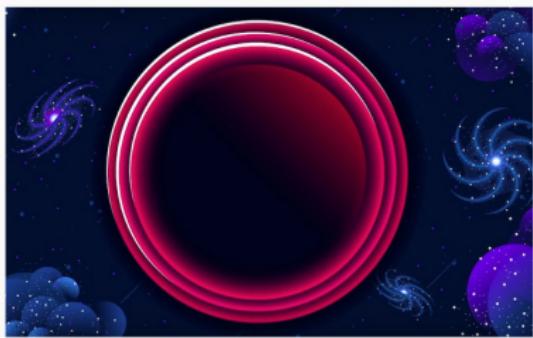
Swampland



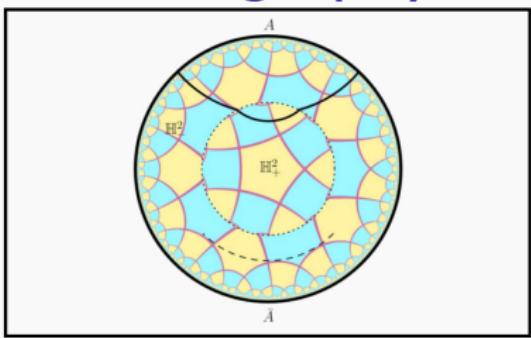
Instability



Vacuum Bubbles



Holography



[credits to Kurzgesagt ↑]

Non-SUSY 10d strings: vantage point?

1. Heterotic $SO(16) \times SO(16)$: **NO SUSY** (Alvarez-Gaume, Ginsparg, Moore, Vafa, 1986)
 2. $U(32)$: Type 0'B **closed + open**: **NO SUSY** (Sagnotti, 1995)
 3. $USp(32)$: **closed + open** (Sugimoto, 1999)
“Brane SUSY Breaking” (Antoniadis, Dudas, Sagnotti, 1999)
- 10D couplings** (Dudas, Mourad; Pradisi, Riccioni, 2000)

Features

- no tachyons
- branes + orientifolds → **residual tension**

$$V(\phi) = V_0 e^{\gamma \phi} \quad \longrightarrow \quad \text{NO flat vacuum!}$$

- \exists $AdS \times S$ **flux backgrounds** (Mourad, Sagnotti, 2016)

are they stable?

Low-energy description

$$S_{\text{eff}} = \int d^{10}x \sqrt{-g} \left(\mathcal{R} - \frac{1}{2}(\partial\phi)^2 - V(\phi) - \frac{e^{\alpha\phi}}{12} H_3^2 + \dots \right)$$

AdS flux compactifications:

- Constant dilaton
- $AdS_3 \times S^7$ (BSB \simeq 0'B), $AdS_7 \times S^3$ (heterotic)
- electric vs magnetic flux N of $H_3 = dB_2$

$$\boxed{N \gg 1 \quad \longrightarrow \quad e^\phi, (\alpha' \mathcal{R}) \ll 1}$$

In orientifold models: $\text{AdS}_3 \times S^7$

Parameters: $V = T e^{\frac{3}{2}\phi}$, coupling $\alpha = 1$ to R-R 3-form

from disk amplitude: $T = 2k_{10}^2 \times \begin{cases} 64 T_{D9} & (\text{BSB}) \\ 32 T_{D9} & (0'B) \end{cases}$

- electric flux

$$N = \int_{S^7} \star e^\phi H_3$$

- (super)gravity regime

$$L_3, R_7 \propto N^{3/16} \quad e^\phi \propto N^{-1/4}$$

- ratio of radii

$$\frac{L_3^2}{R_7^2} = \frac{1}{6}$$

In the heterotic model: $\text{AdS}_7 \times S^3$

Parameters: $V = \Lambda e^{\frac{5}{2}\phi}$, coupling $\alpha = -1$ to NS-NS 3-form

from 1-loop torus amplitude: $\Lambda = (\text{modular integral}) = \frac{\mathcal{O}(1)}{\alpha'}$

- magnetic flux

$$N = \int_{S^3} H_3$$

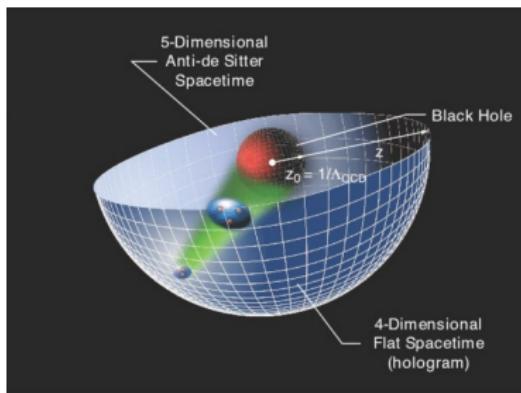
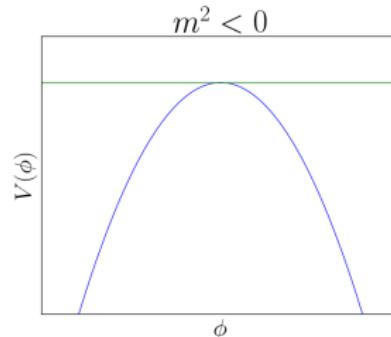
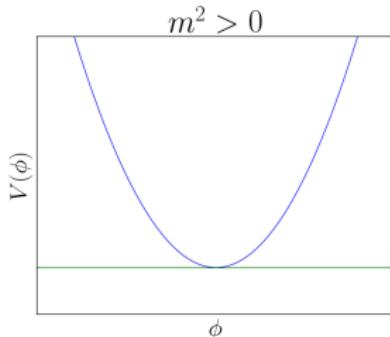
- (super)gravity regime

$$L_7, R_3 \propto N^{5/8} \quad e^\phi \propto N^{-1/2}$$

- ratio of radii

$$\frac{L_7^2}{R_3^2} = 12$$

Perturbative instabilities: Minkowski vs AdS



$m^2 < 0 \rightarrow$ modes can **grow**

AdS: *BF bound* (Breitenlohner, Freedman, 1982)

$$m_{\text{scalar}}^2 \geq -\frac{(d-1)^2}{4L_{AdS}^2}$$

- **some tachyons allowed!**
- **extends to general fluctuations**

AdS vacua: unstable scalar KK modes

- BSB & 0'B \longrightarrow $\ell = 2, 3, 4$ $\text{AdS}_3 \times S^7$
- Heterotic \longrightarrow $\ell = 1$ $\text{AdS}_7 \times S^3$

Dudas-Mourad vacua (Dudas, Mourad, 2000): pert. stable...

- 9d static: ...but large corrections
- 10d cosmology: ...except isotropy?

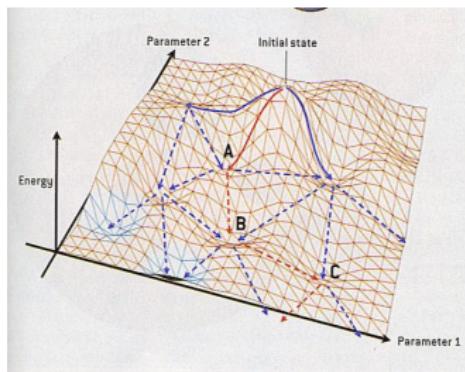
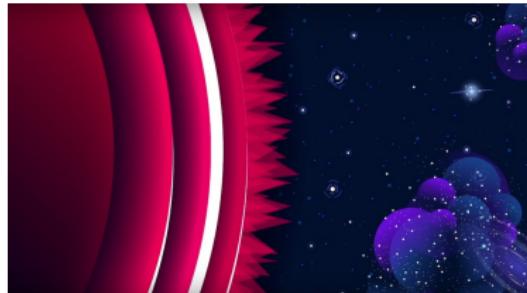
$$\delta g_{ij}(\mathbf{k} = 0) \sim A_{ij} + B_{ij} \log \eta$$

Non-perturbative instabilities

expect instantons...

$$\Gamma_{\text{decay}} \sim (\det) \times e^{-\overbrace{(S_{\text{inst.}} - S_0)}^B}$$

(Coleman, Callan, 1977), (Coleman, De Luccia, 1980)



BUT: no **global** knowledge
no **SUSY!** → any criteria?

Non-perturbative instabilities: brane picture

(Antonelli, IB, 2019)

AdS vacua → *flux tunneling* (Brown, Teitelboim, 1987-1988), (Blanco-Pillado, Schwartz-Perlov, Vilenkin, 2009)

$$\mathcal{E}_{\text{vac}} \propto -N^{-3} \quad \text{or} \quad -N^{-2}$$

$$N \longrightarrow N - \delta N \quad : \quad \text{out of EFT}$$

- Instantons ↔ branes (D1 or NS5)? right charge & dim.
- AdS → **near-horizon** of brane stack...

→ **brane-antibrane** nucleation

$$S_{\text{brane}}^E = \left[\tau_p \text{Area} - \frac{N \mu_p}{e^{\alpha\phi} R^q} \text{Vol} \right]_{\text{extremum}} = B_{\text{CdL}}$$

Consistency: the right branes

$$S_{\text{brane}}^E = \tau_p \Omega_{p+1} L^{p+1} \left[\frac{1}{(\beta^2 - 1)^{\frac{p+1}{2}}} - \frac{p+1}{2} \beta \int_0^{\frac{1}{\beta^2-1}} \frac{u^{\frac{p}{2}}}{\sqrt{1+u}} du \right]$$

Consistency:

- *existence:* nucl. parameter $\beta \equiv v_0 \left(\frac{\mu_p}{\tau_p} \right) g_s^{-\frac{\alpha}{2}} > 1$
- *semi-classical:* $\beta = \mathcal{O}(N^0) \longrightarrow \tau_p = T_p g_s^{-\frac{\alpha}{2}}$

→ relation for fundamental (and exotic) branes! (Bergshoeff, Riccioni et al.)

$$\tau_p^{\text{string}} = \frac{T_p}{g_s^\sigma}, \quad \boxed{\sigma = 1 + \frac{1}{2} \alpha_{\text{electric}}^{\text{string}}}$$

After tunneling: Lorentzian evolution

probe p/\bar{p} -brane in (Poincaré) AdS throat at pos. Z :

$$V_{\text{probe}} = \tau_p \left(\frac{L}{Z} \right)^{p+1} \left(1 \pm v_0 \frac{\mu_p}{T_p} \right)$$

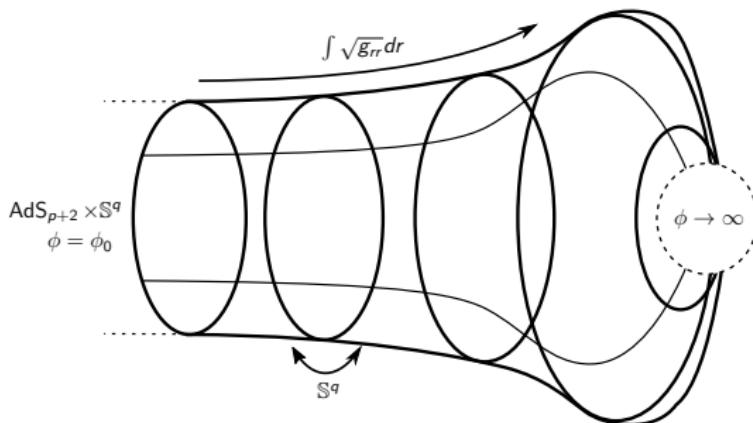
for our string models: $v_0 > 1$!

$$\boxed{\text{WGC} : \nearrow \left(\frac{\text{charge}}{\text{tension}} \right)_{\text{eff}} \nearrow}$$

→ these non-SUSY branes **feel the right forces**

Back-reaction: pinch-off at finite distance

$SO(1, p) \times SO(q)$ geometry



$$\text{geodesic length: } \int \sqrt{g_{rr}} dr < \infty$$

$p = 8$: recover 9d Dudas-Mourad

2 asymptotic free parameters: extremal tuning?

(Top-down) holography?

Dual “CFT”: (IR of) world-volume gauge theory?

$$c_{D1} \propto N^{3/2}$$

Toward small N : bubble RG?

Recent developments: “de Sitter on a brane”?

(Banerjee, Danielsson, Dibitetto, Giri, Schillo, 2018)

e.g. N NS5-branes in $SO(16) \times SO(16)$ \longrightarrow

$$\Lambda_{6d} \propto \frac{1}{N^2}$$

AdS vacua:

- weak coupling
- discrete ($N \rightarrow g_s, \mathcal{R}$)
- non-SUSY

Flux tunneling:

- vacuum bubbles
- branes
- toward UV



Brane picture

- AdS \longleftrightarrow IR world-volume theory?
- tunneling \longleftrightarrow renormalization group flows?
- de Sitter on a brane?

Take-home message

SUSY breaking



spontaneous dynamics

Backup slides

Perturbations and mixings

Linearized analysis: AdS tensors + angular momenta ℓ (IB, Mourad, Sagnotti, 2018)

- **Tensors:** no mixing \rightarrow stable ✓
- **Vectors:** mixing, still stable ✓

$$\delta g_{\mu i} \quad \delta B_{\mu i}$$

- **Scalars:** Einstein eqs. \rightarrow 2 constraints!

$$\delta \phi \quad \delta B_2 = \star_3 d \textcolor{blue}{B}$$

$$\delta g_{\mu\nu} = \textcolor{blue}{A} g_{\mu\nu}^{(0)}$$

$$\delta g_{ij} = \textcolor{blue}{C} g_{ij}^{(0)}$$

$$\delta g_{\mu i} = \nabla_\mu \nabla_i \textcolor{blue}{D}$$

Linearized scalar equations: orientifold case

$$L_3^2 \square A - \left[4 + 3\sigma_3 + \frac{\ell_3}{3}(\sigma_3 - 1) \right] A + \frac{7}{2}\alpha\sigma_3 \delta\phi - \frac{\ell_3}{2}(\sigma_3 - 1) B = 0$$

$$L_3^2 \square \delta\phi + 2\alpha\sigma_3 A - \left[2\alpha^2\sigma_3 + \tau_3 + \frac{\ell_3}{3}(\sigma_3 - 1) \right] \delta\phi + \alpha \frac{\ell_3}{3}(\sigma_3 - 1) B = 0$$

$$L_3^2 \square B - 8\sigma_3 A + 4\alpha\sigma_3 \delta\phi - \frac{\ell_3}{3}(\sigma_3 - 1) B = 0$$

where:

$$\ell_3 = \ell(\ell + 6), \quad \sigma_3 = 1 + 3 \frac{L_3^2}{R_7^2}, \quad \tau_3 = L_3^2 V_0''$$

Linearized scalar equations: heterotic case

$$L_7^2 \square A - [\ell_7 (\sigma_7 - 3) + 5\sigma_7 + 12] A + \frac{5}{2} \alpha \sigma_7 \delta\phi - \frac{3\ell_7}{2} (\sigma_7 - 3) B = 0$$

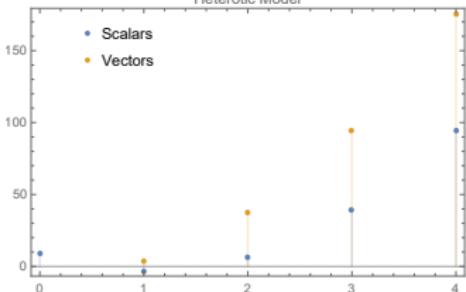
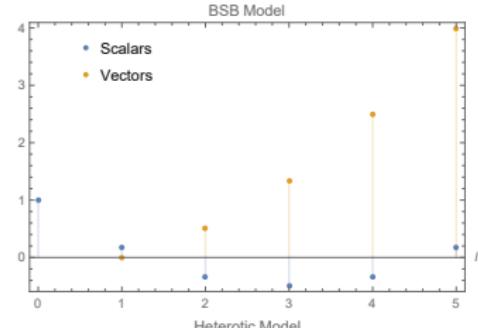
$$L_7^2 \square \delta\phi + 6\alpha \sigma_7 A - [2\alpha^2 \sigma_7 + \tau_7 + \ell_7 (\sigma_7 - 3)] \delta\phi + \alpha \ell_7 (\sigma_7 - 3) B = 0$$

$$L_7^2 \square B - 8\sigma_7 A + 4\alpha \sigma_7 \delta\phi - \ell_7 (\sigma_7 - 3) B = 0$$

where:

$$\ell_7 = \ell(\ell+2), \quad \sigma_7 = 3 + \frac{L_7^2}{R_3^2}, \quad \tau_7 = L_7^2 V_0''$$

Results: violations of BF bounds



BSB Model $(AdS_3 \times S^7)$

- **Scalars:** $\ell = 2, 3, 4$
- **Vectors:** $\ell = 1$ massless (KK)

Heterotic Model $(AdS_7 \times S^3)$

- **Scalars:** $\ell = 1$
- **Vectors:** $\ell = 1$ massless (KK)

Orbifolds: can get rid of unstable modes...

→ **vacuum bubbles?** (Horowitz, Orgera, Polchinski, 2008)

[also **cosmological** vacuum: stable, but **isotropy** breaking?] (IB, Mourad, Sagnotti, 2018)

Non-perturbative instabilities: flux tunneling

- Gravity in $D = p + 2 + q$ dims + fluxes: S^q reduction

$$ds^2 = R^{-\frac{2q}{p}} ds_{p+2}^2 + R^2 d\Omega_q^2$$

- Reduced action: $(p+2)$ -Einstein frame

$$S_{p+2} = \frac{1}{2\kappa_{p+2}^2} \int d^{p+2}x \sqrt{-g_{p+2}} \left(\mathcal{R}_{p+2} - 2\Lambda R^{-\frac{2q}{p}} \right)$$

vacuum energy $\longrightarrow \mathcal{E}_{\text{vac}} \propto -R^{-\frac{2q}{p}-2}$

\mathcal{E}_{vac} depends on flux...

...higher-dim. instantons, flux transitions (Blanco-Pillado, Schwartz-Perlov, Vilenkin, 2009)

Many branes: background geometry

$SO(1, p) \times SO(q)$ symmetry: $\phi(r), v(r), b(r)$

$$ds^2 = e^{\frac{2}{p+1}v - \frac{2q}{p}b} dx_{1,p}^2 + e^{2v - \frac{2q}{p}b} dr^2 + e^{2b} R_0^2 d\Omega_q^2,$$

$$\phi = \phi(r),$$

$$H_{p+2} = c e^{2v - \frac{q}{p}(p+2)b} d^{p+1}x \wedge dr, \quad c \equiv \frac{N}{e^{\alpha\phi} (R_0 e^b)^q}$$

(Constrained) Toda-like system: $(+, -, +)$ kinetic term w/ potential

$$U = -T e^{\gamma\phi + 2v - \frac{2q}{p}b} - \frac{n^2}{2R_0^{2q}} e^{-\alpha\phi + 2v - \frac{2q(p+1)}{p}b} + \frac{q(q-1)}{R_0^2} e^{2v - \frac{2(D-2)}{p}b}$$

Geometry: near-horizon

Recover original $\text{AdS}_{p+2} \times S^q$ with $[r < 0]$

$$\phi = \phi_0, \quad e^\nu = \frac{L}{p+1} \left(\frac{R}{R_0} \right)^{-\frac{q}{p}} \frac{1}{-r}, \quad e^b = \frac{R}{R_0}$$

Radial perturbations: $\delta\phi, \delta\nu, \delta b \propto (-r)^\lambda$

$$\{\lambda\}_{\text{BSB}} = \left\{ -1, \frac{1 \pm \sqrt{13}}{2}, \frac{1 \pm \sqrt{5}}{2} \right\}, \quad \{\lambda\}_{\text{het}} = \left\{ -1, \pm 2\sqrt{\frac{2}{3}}, 1 \pm 2\sqrt{\frac{2}{3}} \right\}$$

→ two *extremality-breaking* deformations

[*two asymptotic fine-tunings?*]

Geometry: “far-horizon”

Away from branes [$r > 0$]: assume $U \sim U_T = -T e^{\gamma\phi + 2v - \frac{2q}{P} b}$

Solutions as $r \rightarrow \infty$: $\phi, v, b \propto y(r) + \text{subleading}$

$$y'' \sim \hat{T} e^{\Omega y + L r}$$

$$\frac{1}{2} \Omega y'^2 + Ly' \sim \hat{T} e^{\Omega y + L r} - M$$

where $\Omega = \frac{D-2}{8} \gamma^2 - \frac{2(D-1)}{D-2} = \frac{D-2}{8} (\gamma^2 - \gamma_{\text{crit}}^2)$ (IB, Mourad, Sagnotti, 2018)

- Orientifolds: $\phi, v, b \propto r^2$ (due to $\Omega = 0$)
- Heterotic: $\phi, v, b \propto \log(r_0 - r)$

Holography of vacuum bubbles (“Bubbleography”)

Non-SUSY brane instantons at low energy: **vacuum bubbles**

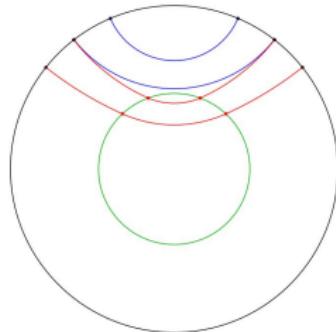
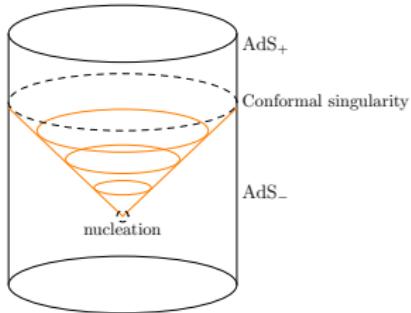
Spontaneous, irreversible process

$$\text{AdS}_3(L_-) \longrightarrow \text{AdS}_3(L_+)$$

dual “central charge” $c_- < c_+$

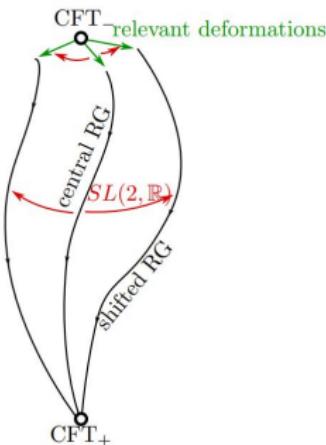
...holographic description? (Antonelli, IB, Bombini, 2018)

- (AdS) vacuum **bubbles** \longleftrightarrow boundary RG flow
- Bubble **nucleation** \longleftrightarrow (**non-local**) RG trigger
- **Displaced** bubble \longleftrightarrow **entanglement** pattern of boundary



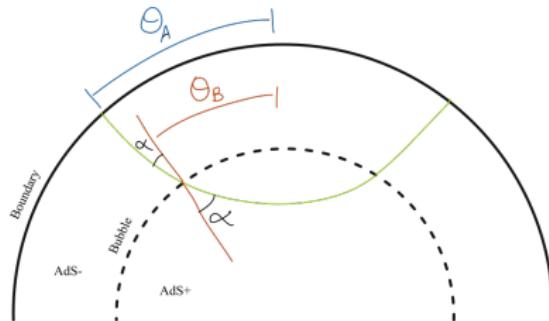
Bulk \rightarrow expanding bubble and geodesics

Boundary \rightarrow relevant deformations and RG



Our check: bubble entanglement entropy (in 3d)

$$ds_{\pm}^2 = L_{\pm}^2 \left(-\cosh^2 \rho_{\pm} d\tau_{\pm}^2 + d\rho_{\pm}^2 + \sinh^2 \rho_{\pm} d\phi_{\pm}^2 \right)$$



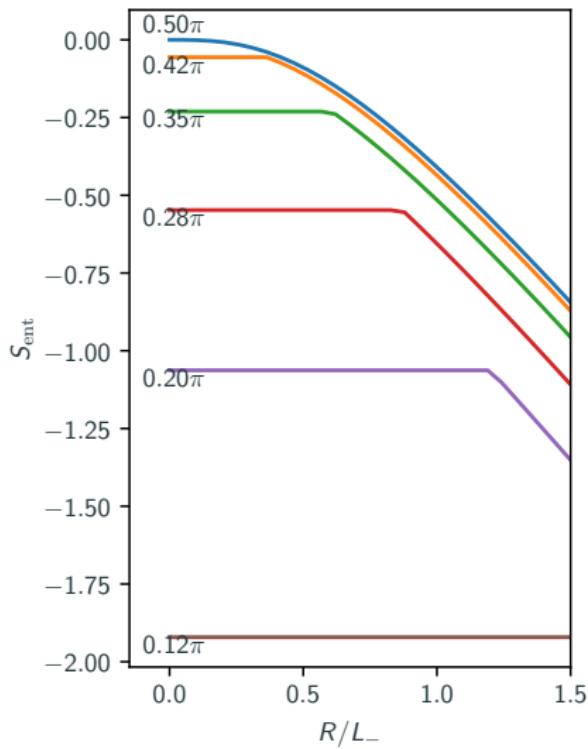
- **Thin-wall:** geodesic is hyperbolic polygonal

$$\begin{aligned} \text{length} &= 2L_- \Lambda + 2L_- \log(\cosh \rho_- - \sinh \rho_- \cos(\theta_B - \theta_A)) \\ &\quad + L_+ \cosh^{-1}(\cosh^2 \rho_+ - \sinh^2 \rho_+ \cos(2\theta_B)) + \mathcal{O}(\Lambda^{-1}) \end{aligned}$$

- Angle at bubble θ_B : **no-kink condition**

gluing condition: $L_- \sinh \rho_- = L_+ \sinh \rho_+$

c -functions and “bubble RG”



Finite part of $S_{\text{ent}}(\theta_A, R)$
decreases with **bubble radius**.



candidate **c -function!**

Check: trace anomaly

(Henningson, Skenderis, 1998)

$$\langle T^{\mu}_{\mu} \rangle = \frac{c_{\text{ent}}(R)}{12} \mathcal{R}$$

Moving the bubble:
can use integral geometry

(Antonelli, IB, Bombini, 2018)

Moving the bubble: pills of integral geometry [backup]

Crofton theorem: space of all lines \mathcal{K} (Crofton, 1968)

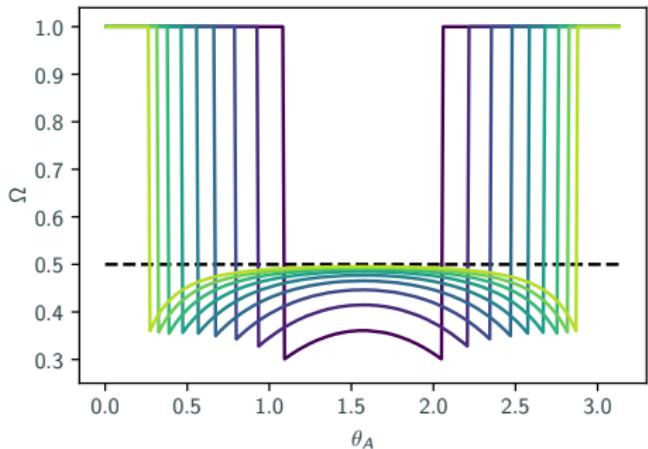
$$\text{length}(\gamma) = \frac{1}{4} \int_{\mathcal{K}} n_{\gamma, \kappa} \omega(\kappa) \leftarrow \text{Crofton form}$$

Asymptotically \mathbb{H}^2 slices

$$\omega_{\text{bubble}} = \Omega \omega_{\mathbb{H}^2}$$

moving bubble via isometry

→ Ω scalar!



Conclusions

Summary

- **Non-SUSY** AdS flux vacua: *controlled for large fluxes*
- **Perturbative instabilities:** violation of BF bounds
- **Non-perturbative instabilities:** flux tunneling
- Brane description \longrightarrow **dual framework** (“*bubble RG*”)

Outlook

- Brane dynamics: **small N potential**, **gauge theory**, ...
- Explore **bubble RG**: structure of RG trigger?
(closer to SUSY: \mathbb{Z}_k orbifolds of $\mathcal{N} = 4$) (Horowitz, Orgera, Polchinski, 2008)
- Time dependence: analog of 10d cosmological vacuum?