

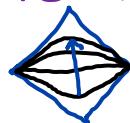
Symmetries of Gravitational Scattering

GGI Workshop
April 2021

Andy Strominger
Harvard University

I'd like to draw attention to 3 unsolved problems in GR scattering in AF spacetimes:

1. How is $\gamma^- \rightarrow \gamma^+$ scattering defined?
More subtle than



2. What are the non-trivial symmetries/conservation laws? $E, \vec{P}, \vec{J}^+ \dots \infty 100^+$ years \geq Einstein still unknown!
3. What are the observational signals (memory $\gamma^- \rightarrow \gamma^+$) of these symmetries?

This talk:

> $\frac{1}{2}$ review last 5-10 yrs progress*

< $\frac{1}{2}$ recent paper on GR
symmetries with

Alfredo Guevara, Mina Himwich,
Monica Pate & AS 2103.03961

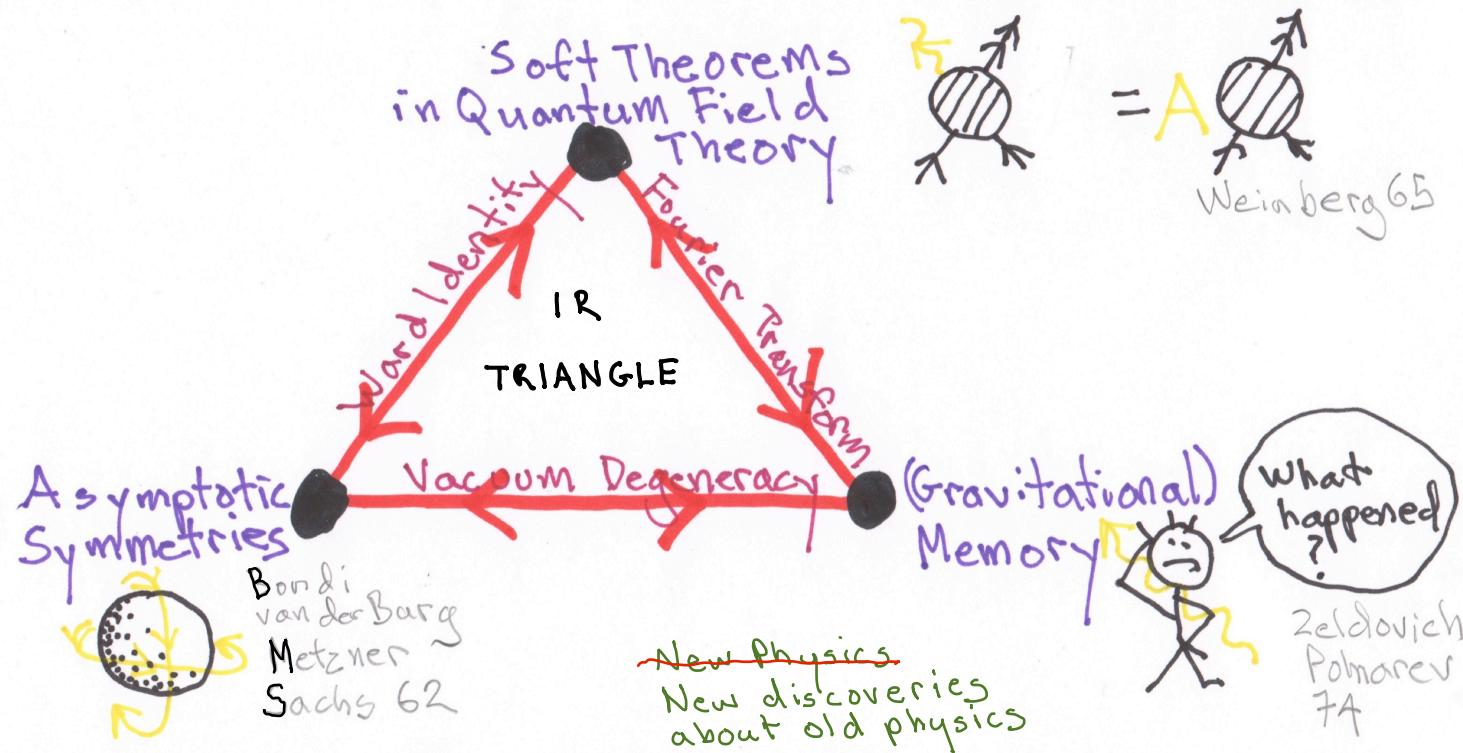
* "Lectures on the IR Structure
of Gravity & Gauge Theory"

Princeton Univ. Press 2018

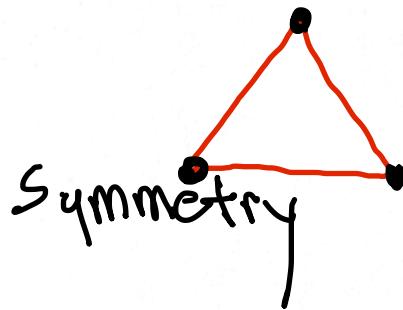
hep-th 1703.09488

I had many wonderful collaborators
referred to herein.

Over the last 7 years, an exact mathematical equivalence has been discovered of 3 previously disparate phenomena, each studied for half a century:



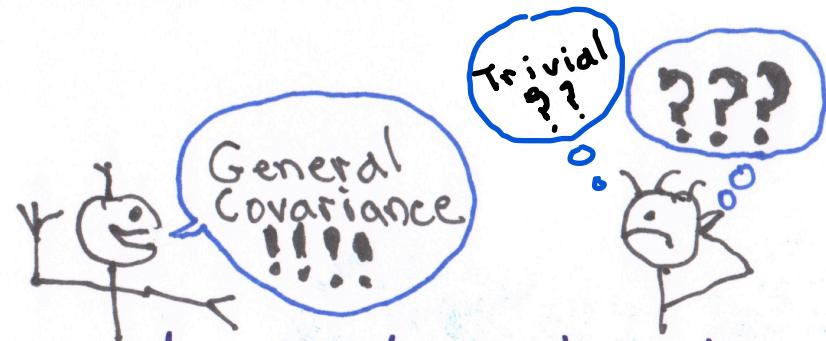
This has led to surprising new insights into the low-energy structure of gravitational & electromagnetic theories, implications for quantum black hole information & novel experiments at hadron accelerators, the EIC, LHC, LIGO, LISA; IR divergences & flatspace holography.



Symmetry



What are
the symmetries
of GR ???



This subtle & often ill-posed question has been the source of consternation & confusion for 100 years. Our best answer has changed many times significantly, right up to the present.

In Special Relativity, Poincare symmetries lead, via Noether's theorem, to measurable conserved charges such as Energy, Momentum and Angular Momentum. For example, Energy conservation follows from time translation symmetry

$$t \rightarrow t + \text{constant}$$

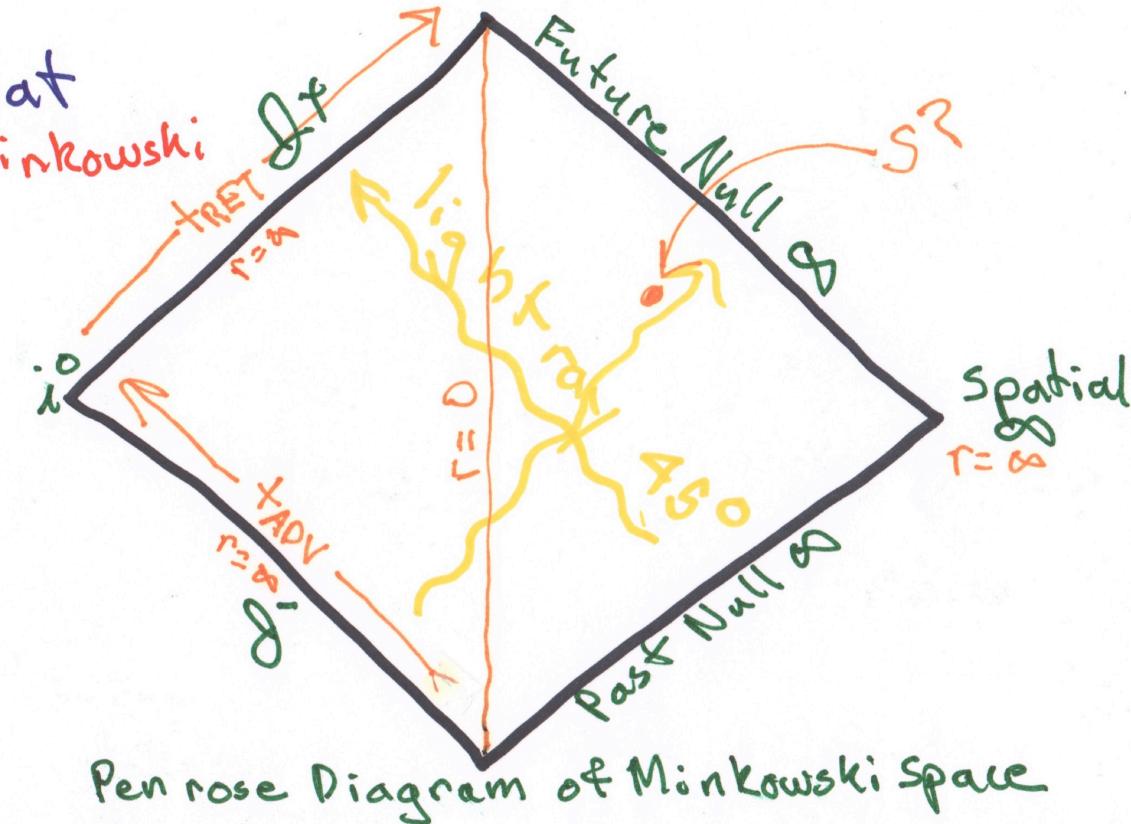
t is a time coordinate, and this is a coordinate transformation. Bondi, van der Burg, Metzner & Sachs (BMS) in 1962 asked "How do Poincare symmetries/charges appear in GR?"

The answer was a big surprise.

KEY IDEAS

1. Asymptotically Flat
Spacetimes "like" Minkowski

near \mathcal{J}^\pm
Very subtle!
Arnowitt
Deser
Misner



2. Asymptotic Symmetries
 $\frac{\text{Coord. Trans. preserving boundary cond.}}{\text{Trivial Coord. Trans.}}$

$=$
 \supseteq
 \subseteq

Poincare **NO!!!**

ON \mathcal{J}^+

ASYMPTOTIC SYMMETRY = ∞ -DIMENSIONAL BMS $^+$ SYMMETRY

1962

comprised of asymptotic boosts,
rotations & SUPERTRANSLATIONS

$$+_{\text{RET}} \rightarrow +_{\text{RET}} + f(\theta, \phi)$$

for any f . These act
non-trivially on outgoing
physical data:

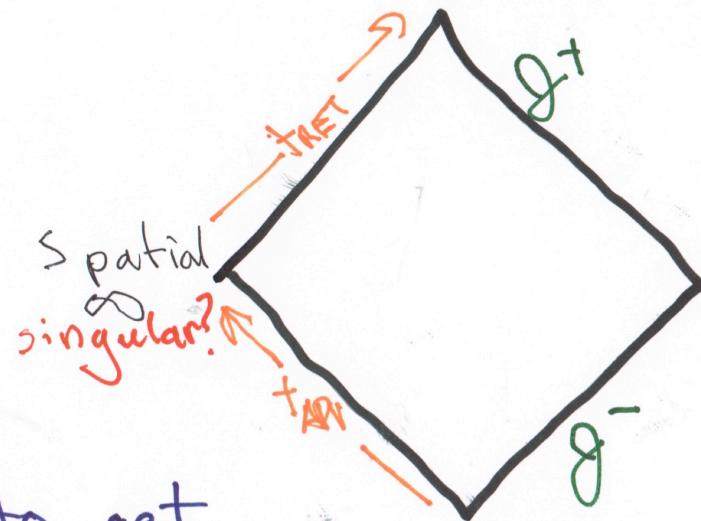


There is also BMS $^-$ on \mathcal{J}^-

$$+_{\text{ADV}} \rightarrow +_{\text{ADV}} + \tilde{f}(\theta, \phi)$$

So what do we do with this fundamental symmetry of GR?

For 50 yrs after BMS, it was widely held that BMS^+ and BMS^- acted separately on g^+ and g^- and could **not** be tied together at spatial ∞ to get a symmetry of gravitational scattering. This was in part due to a misunderstanding about the structure of spatial ∞ which was not cleared up until the seminal 1992 work of Christodoulou & Klainerman.



Using Christodoulou & Klainerman, it was recently shown that the combined δ^+ and δ^- supertranslations

$$t_{RET} \rightarrow t_{RET} + f(\theta, \phi)$$

$$t_{ADV} \rightarrow t_{ADV} + f(\pi - \theta, \phi + \pi) \quad \text{Antipodal}$$

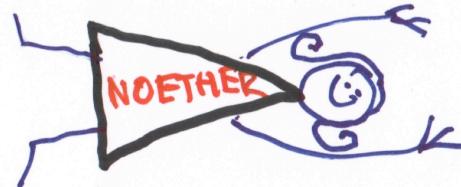
are an ∞ number of exact, nontrivial symmetries of classical gravitational scattering. As 2013

In the language of the S -matrix:

$$B S - S B = 0$$

where the operator B generates infinitesimal supertranslations and acts non-trivially on physical states. These are true "Symmetries of GR". But not all of them!

SYMMETRIES

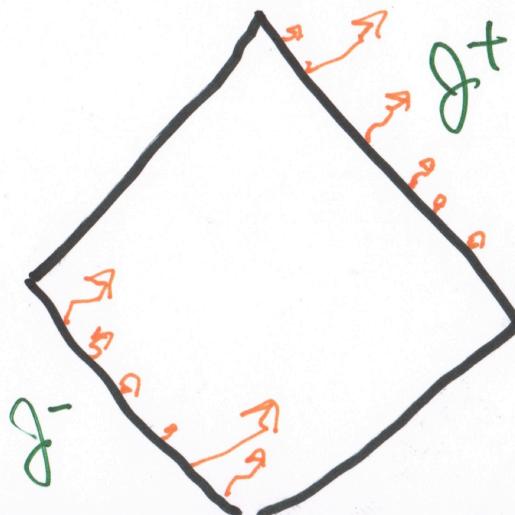


CONSERVATION LAWS.

For $f = \text{constant}$ the conservation law is $E_{\text{in}} = E_{\text{out}}$.
More generally, we find

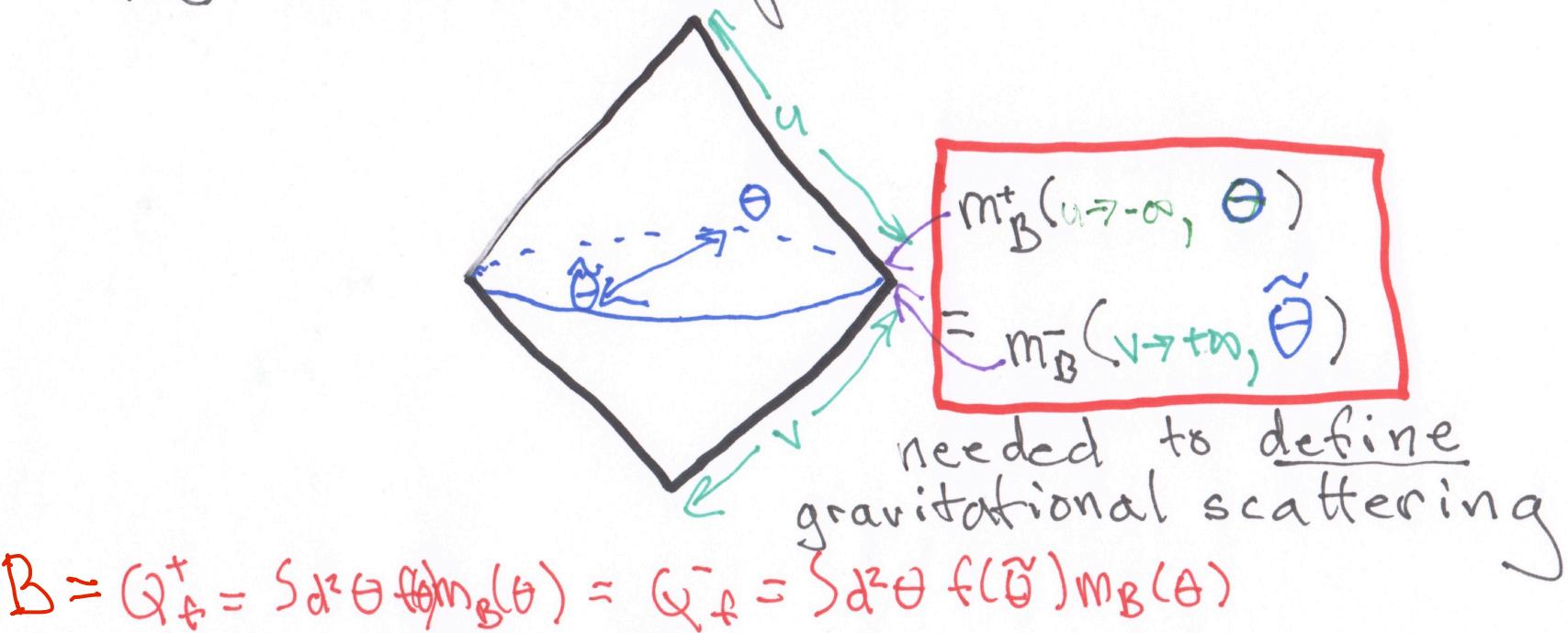
Integrated total incoming energy* flux
from angle θ, ϕ

= Integrated total outgoing energy
flux through angle $\pi - \theta, \phi + \pi$!



* Energy in GR
has crucial term linear in
metric fluctuation. (Bondi)

An alternate derivation of the laws of conservation follows from the b.c. on the Bondi mass aspect



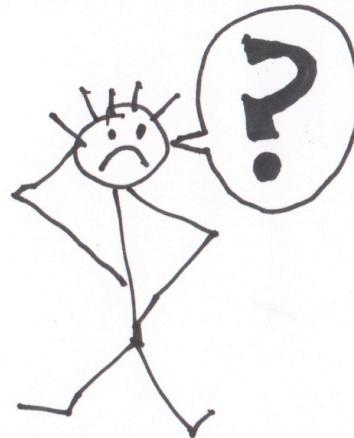
NOT THE END!

The story so far follows from antipodal matching of the leading $\frac{1}{r}$ component of the metric & leading soft thm.

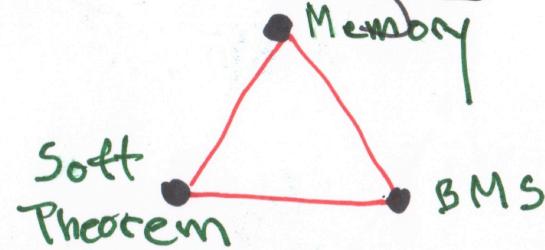
\exists an ∞ hierarchy of subleading subsubleading structures.

More later.

Now I will turn to what seems a completely different topic.



In fact we will see shortly that it is exactly the same topic in different mathematical language; a corner of the triangle



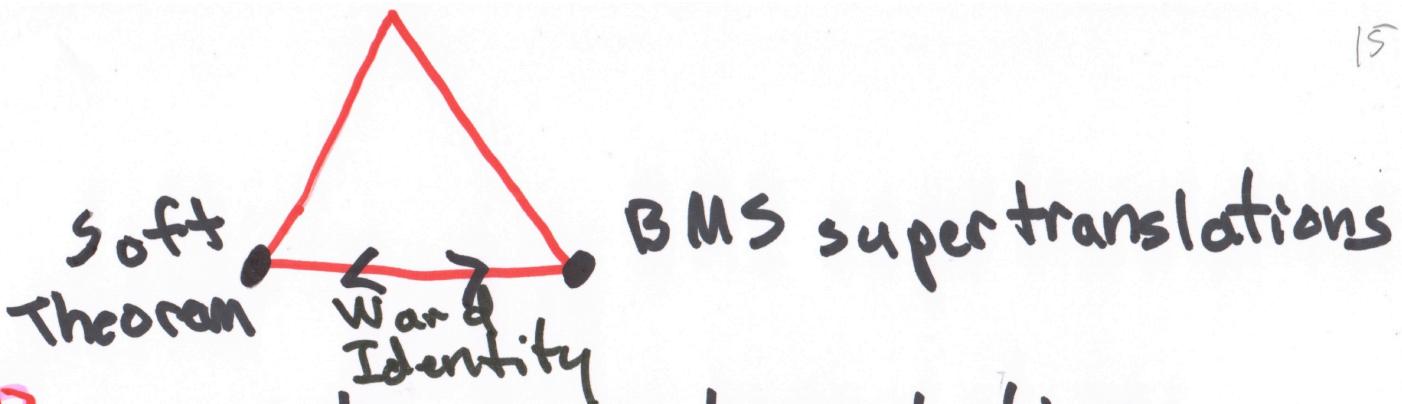


Soft Graviton Theorem

Weinberg
1965

$$= \left(\sum_{k=1}^n \frac{\epsilon_{\mu\nu} p_k^\mu p_k^\nu}{q \cdot p_k} \right) + \mathcal{O}(q^0)$$

Exact & universal formula with simple diagrammatic derivation. Valid in any 4D consistent classical or quantum theory of gravity.



Let B generate super translations.
Then

$$B\delta - \delta B = 0$$

moreover

$$\langle \text{out} | B\delta - \delta B | \text{in} \rangle$$

is precisely Weinberg's soft graviton theorem. α of associated conserved quantities is total energy at each incoming angle.

GR Vacuum is α -Degenerate

Classical

$g_{\mu\nu} = \eta_{\mu\nu}$ NOT preserved by translations
changes angular momentum!
~~"GR problem of ang. momentum"~~

Quantum

$$[H, a_{\text{grav}}^{\dagger}] = 0; [J, a_{\text{grav}}^{\dagger}] \neq 0$$

$$\Rightarrow a_{\text{grav}}^{\dagger} |0\rangle \neq |0\rangle$$

= new vacuum

Supertranslation symmetry is "spontaneously broken".

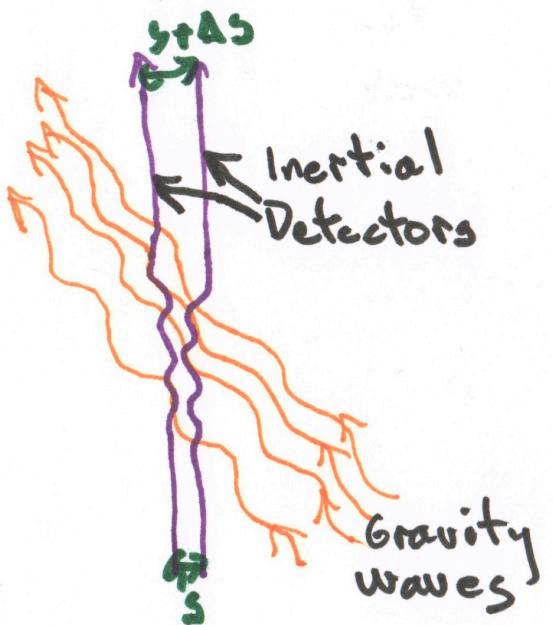
Its fascinating that these two
essentially equivalent discoveries were
made 3 years apart.... perhaps an example
of how ideas percolate unrecognized through
our community...

Now for a second apparent topic
change!



Gravitational Memory

Zeldovich & Polnarev '74
Braginsky & Thorne '87
Christodoulou '81
[17]



A F Minkowski metric:

$$ds^2 = -du^2 - 2 du dr + r^2 \gamma_{\bar{z}\bar{z}} d\bar{z} d\bar{z}$$

$$+ r C_{\bar{z}\bar{z}} d\bar{z}^2 + r (\bar{z}\bar{z} d\bar{z}^2) + \dots$$

Flat iff $C_{\bar{z}\bar{z}} = -2 D_z^2(\bar{z}, \bar{z})$ gravity waves

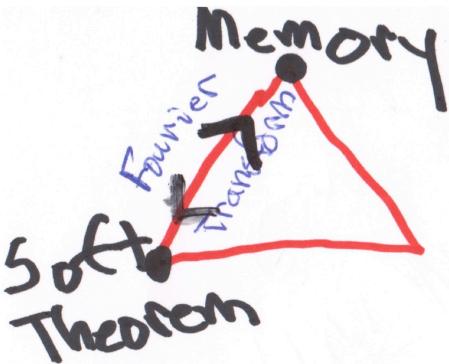
Integrated geodesic deviation equation:

$$\Delta s^{\bar{z}} = \frac{1}{2} r \gamma^{\bar{z}\bar{z}} \Delta C_{\bar{z}\bar{z}} s^{\bar{z}}$$

final - initial = change in $C_{\bar{z}\bar{z}}$

$$\Delta C_{\bar{z}\bar{z}} = \frac{1}{2} D_z^2 f = \text{BMS transf. A.S. Zhiboedov}$$

QC effect. Quite possibly measured in coming decades. Pulsar timing array? LIGO? eLISA?



"Infrared Photons & Gravitons"
S. Weinberg, 1965

The dominance of the $1/(p \cdot q)$ pole in (2.5) implies that the effect of attaching one soft-graviton line to an arbitrary diagram is to supply a factor equal to the sum of (2.5) over all external lines in the diagram

Sdw

$$(8\pi G)^{1/2} \sum_n \eta_n p_n^\mu p_n^\nu / [p_n \cdot q - i\eta_n \epsilon]. \quad (2.7)$$

"Gravitational Wave-Burst with Memory and Experimental Prospects"
V. Braginsky and K.S. Thorne 1987

permanent change in the gravitational-wave field (the burst's memory) δh_{ij}^{TT} is equal to the 'transverse, traceless (TT) part'³⁶ of the time-independent, Coulomb-type, $1/r$ field of the final system minus that of the initial system. If \mathbf{P}^A is the 4-momentum of mass A of the system and P_j^A is a spatial component of that 4-momentum in the rest frame of the distant observer, and if \mathbf{k} is the past-directed null 4-vector from observer to source, then δh_{ij}^{TT} has the following form:

$$\delta h_{ij}^{TT} = \delta \left(\sum_A \frac{4P_i^A P_j^A}{\mathbf{k} \cdot \mathbf{P}^A} \right)^{TT} \quad (1)$$

Here we use units with $G = c = 1$. In the observer's local Car-

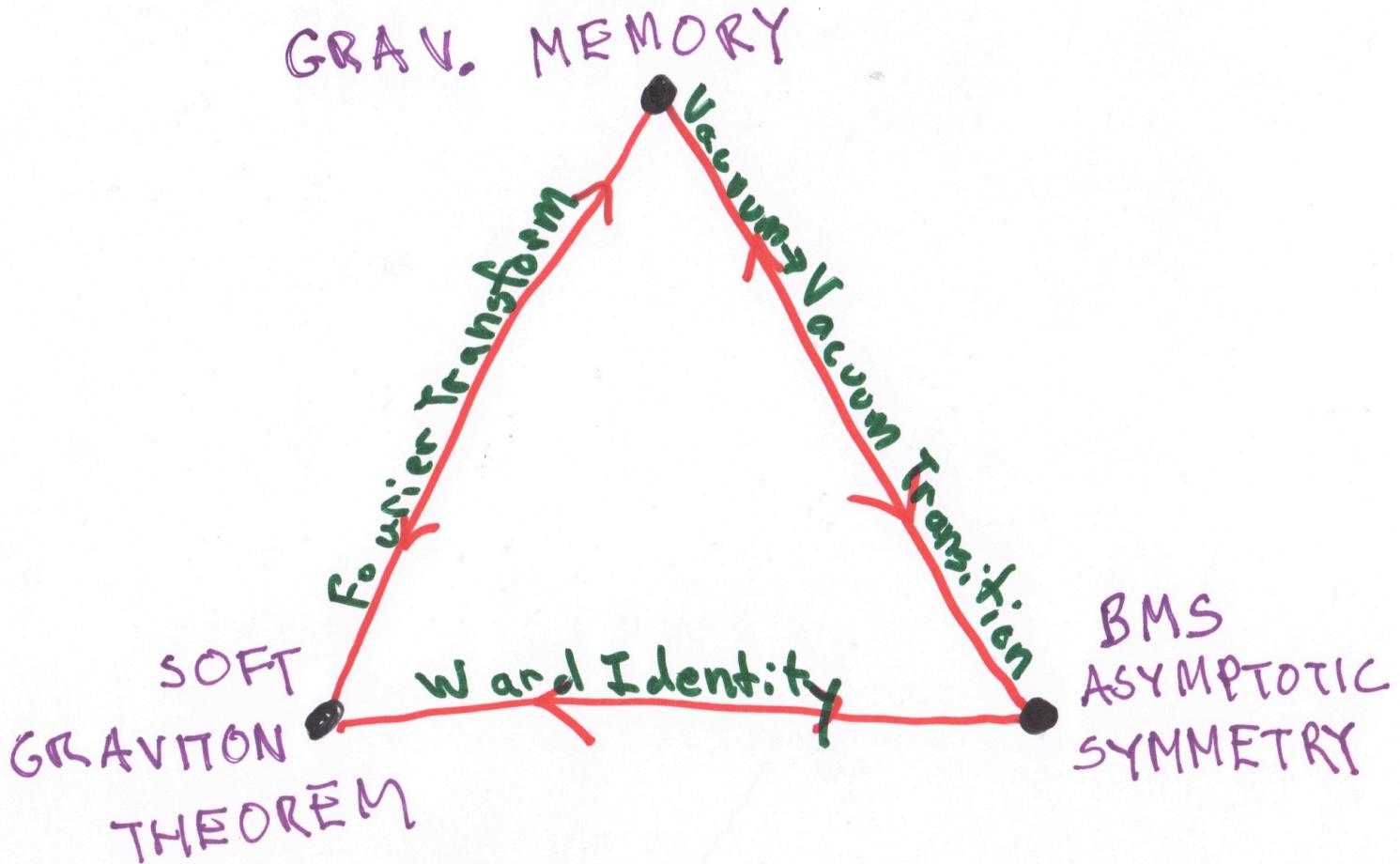
!!!

Changing notations, using $\Theta(\omega) = S \frac{d\omega}{\omega} \frac{e^{i\omega u}}{2\pi i \delta}$ and
elementary particles \rightarrow black holes, stars!

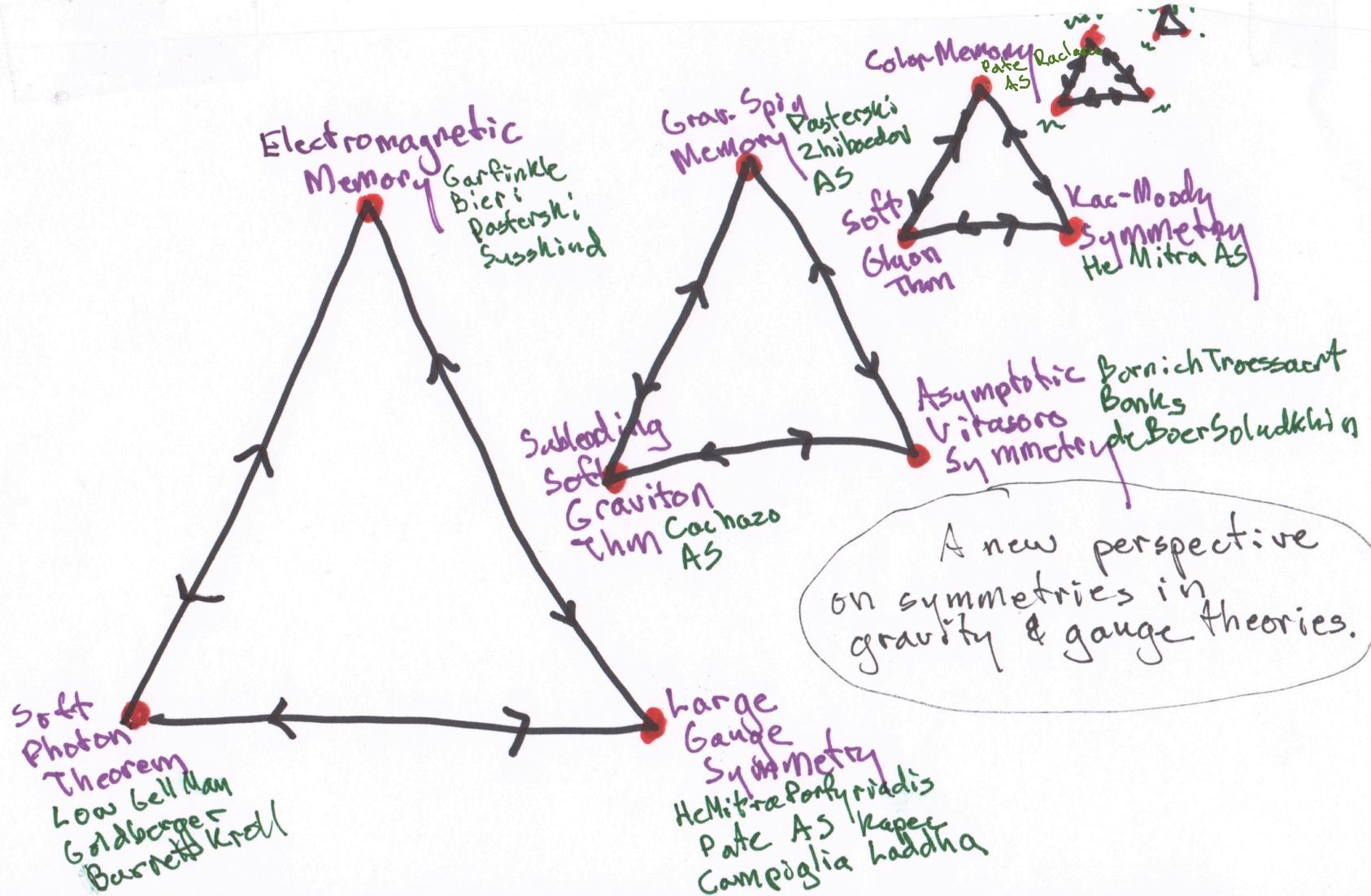
Memory effect provides physical method
to measure soft gravitons.

AS + Zhi-boedov
2014

SUMMARY: IR TRIANGLE



REVERBERATING TRIANGLES



What are the fundamental symmetries of SM + GR?

- i) Consider AF approximation.
- ii) Diffeos mostly trivial.
- iii) Non-trivial includes diagonal subgroup of $BMS^+ \otimes BMS^-$
- iv) Certainly more, at present no proposal for complete answer

Important basic question w/ observational content. Surprisingly not yet answered!

☰ many possible formalisms

Canonical, covariant S-matrix, Lagrangian,
soft thms, asympt. analysis...

Should all be equivalent. Esp. efficient

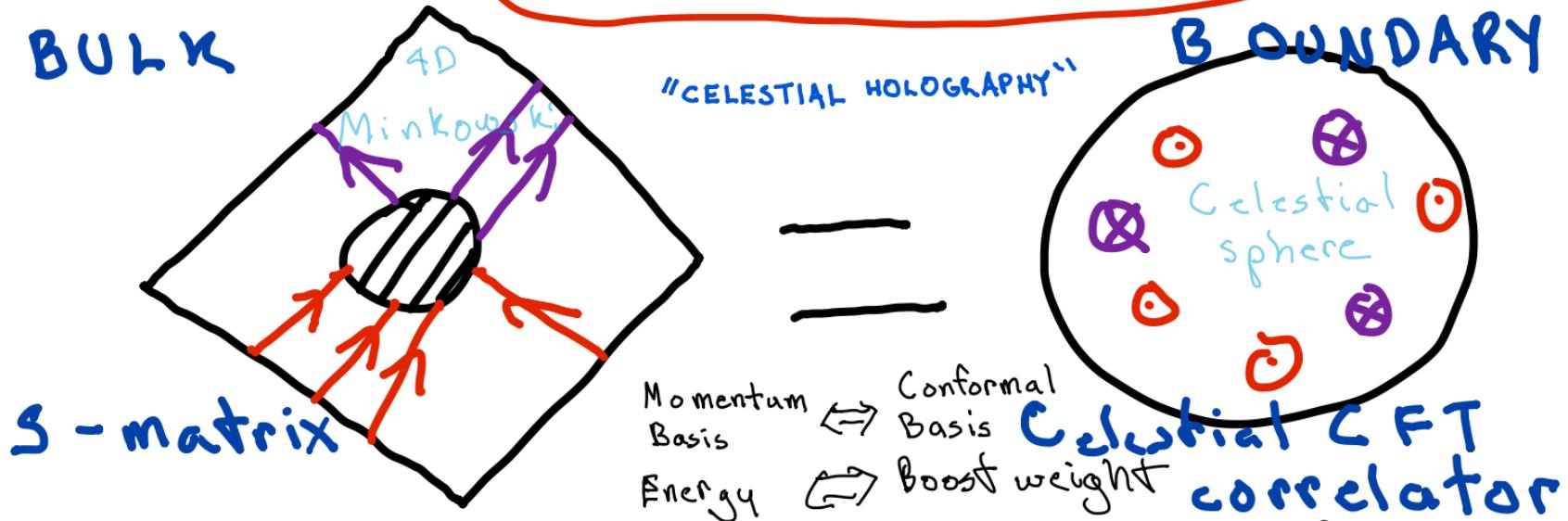
"Generalized currents on celestial sphere"

Gauge choice boundary conditions

Recent progress

A. Guevara, M. Himmich, M. Pate + AS
hep-th/2103.03961

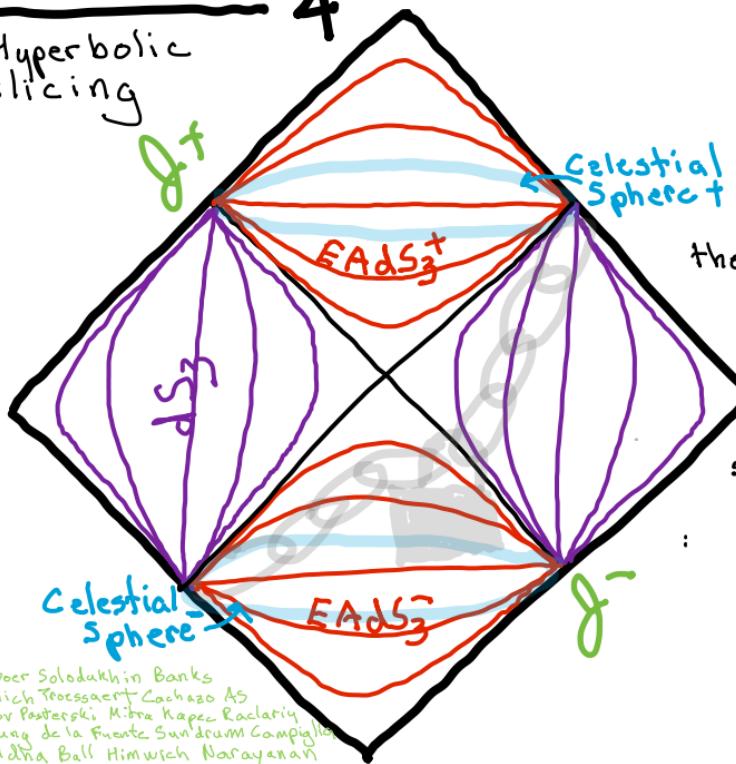
Basic Idea



$SL(2, \mathbb{C})$ Lorentz = global 2D conformal
Superrotations = local 2D conformal
Supertranslations $\mathcal{O}_\Delta(z, \bar{z}) \rightarrow \mathcal{O}_{\Delta+1}(z, \bar{z})$!
... more ...

Minkowski 4

Hyperbolic
slicing



Thanks to the subleading soft graviton thm, we now know the ∞ -dimensional 2D conformal symmetry is uplifted to the celestial sphere. But there are ∞ -ly many more from supertranslations, subsubleading soft gravitons,

de Boer Solodukhin Banks
Barnich Roessner Cachazo AS
Lysov Pasterski Mitev Kaptev Radcliffe
Cheung de la Fuente Sundrum Campiglia
Haddad Ball Himwich Narayanam

Why is celestial holography of interest?

1. Flat space version of AdS/CFT, but bottom-up (c.f. Brown-Henneaux) not (so-far) top-down (c.f. Maldacena) and applies to real world.
2. Computationally effective, can apply powerful constraints/methods of 2D CFT to 4D (classical and quantum) gravity. Very useful for symmetry analysis.
3. ...

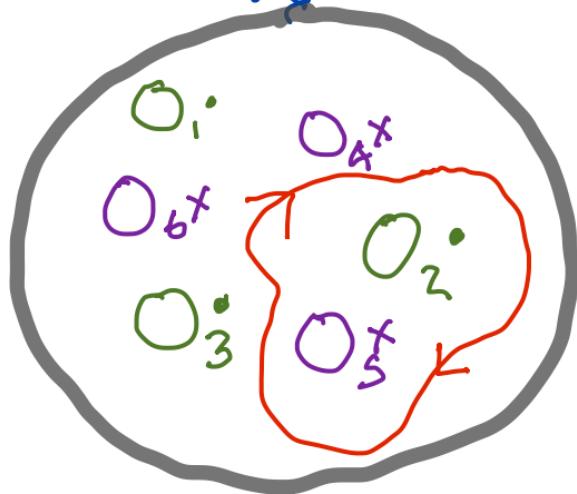
HOW TO REPRESENT A 4D SYMMETRY ON THE CELESTIAL SPHERE

$$ds^2 = -du^2 - 2du dr + r^2 \gamma_{z\bar{z}} dz d\bar{z} + r C_{zz} d\bar{z}^2 + r C_{\bar{z}\bar{z}} d\bar{z}^2 + \dots$$

$N_{zz} \equiv \partial_u C_{zz}$ Bondi News

$$P_z^- \equiv \frac{1}{4G} \left\{ \frac{du}{dr} D^z N_{zz} - [g^-] \right\}, \quad \partial_{\bar{z}} P_z^- = 0$$

Asymptotic constraint
away from
fluxes/
operators



$$\begin{aligned} & \oint_C f(z) \langle P(z) O_1(z_1) \dots O_6(z_6) \rangle \\ &= \sum_{k=1,5} \frac{f(z_k) E_k}{z - z_k} \langle O_1(z_1) \dots O_6(z_6) \rangle \end{aligned}$$

shifts boost weight by one

$f(z)$ = global translation
SUPERTRANSLATIONS

Under $\tilde{z} \rightarrow \frac{az+b}{cz+d}$, \exists SL2R doublet of currents $P^\pm(z)$. Moreover \exists SL2R triplet starting from superrotation generator $\int_{\tilde{z}}^{\infty} du u D^z N_{zz} = 0$. Continuing infinite tower from $\int du u^{1-n} D^z N_{zz}$ denoted $H_n^k(z)$, $k=1, 0, -1 \dots$, $\frac{k-z}{2} \leq n \leq \frac{z-k}{2}$ of currents.

Highly reminiscent of Newman-Penrose conserved quantities from 60s.

SYMMETRY ALGEBRA

2103.03961

Guevara, Himwich, Pate AS

We found an infinite tower of left-moving currents
in finite-dimensional SU_R representations

$$\left[H_n^k, H_{n'}^l \right] = -\frac{\kappa}{2} \left[n'(2-k) - n(2-l) \right] \frac{\left(\frac{2-k}{2}-n+\frac{2-l}{2}-n'-1\right)!}{\left(\frac{2-k}{2}-n\right)!\left(\frac{2-l}{2}-n'\right)!} \frac{\left(\frac{2-k}{2}+n+\frac{2-l}{2}+n'-1\right)!}{\left(\frac{2-k}{2}+n\right)!\left(\frac{2-l}{2}+n'\right)!} H_{n+n'}^{k+l},$$

$\frac{k-2}{2} \leq n \leq \frac{2-k}{2}$

$k, l = 1, 0, -1, \dots$

- (i) Exhibits power of celestial formulation
- (ii) Obey Jacobi, Poincare - fully determined by?
- (iii) Infinite tower of soft thms, symmetries & memories.
- (iv) $K=1, 0$ generate all gauge, $K=1, 0, -1$ all gravity
- (v) These are symmetries of nature w/ in principle observable memory effects. $K=1, 0$ known.

Conclusion

Gravitational scattering is a rich problem theoretically & exp.
Asymptotic symmetries, soft theorems and the celestial reformulation are leading to new insights. There is much yet to be understood.

