

Waveforms from the KMOC formalism and coherent states

Riccardo Gonzo based on
work in progress with A.Cristofoli, D.Kosower and D.O'Connell



SAGEX
Scattering Amplitudes:
from Geometry to Experiment



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17 May 2021

- 1 Coherent states: Motivation and examples
- 2 Wave scattering observables
- 3 Waveforms and localized observers
- 4 Conclusion

Motivation (I)

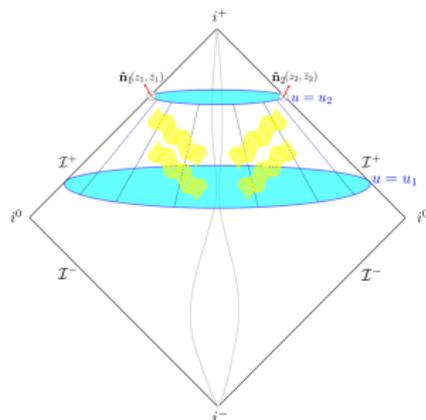
- Scattering amplitudes are useful for the conservative regime of the inspiral phase of the BBH/BNS/BHNS system¹, using a suitable analytic continuation between bound and unbound orbits[Kälin,Porto].
Can we extend this to the dissipative case?

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How is the classical gravitational radiation produced in the two-body scattering problem represented at the quantum level?
What is the expression of the waveform in terms of scattering amplitudes?



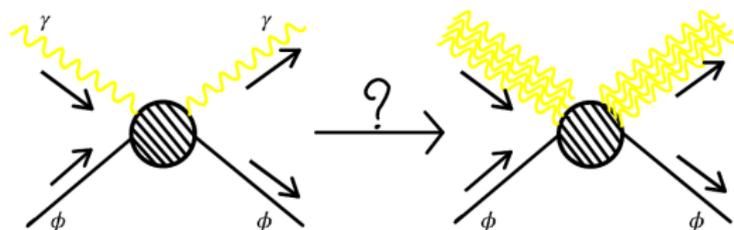
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What are
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Bonus: Toy model for gravity!

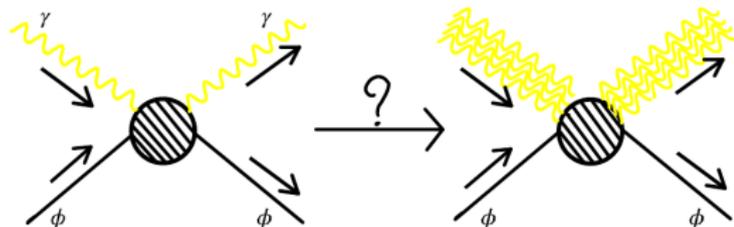


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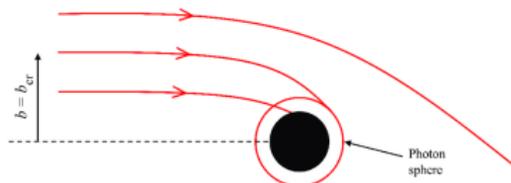
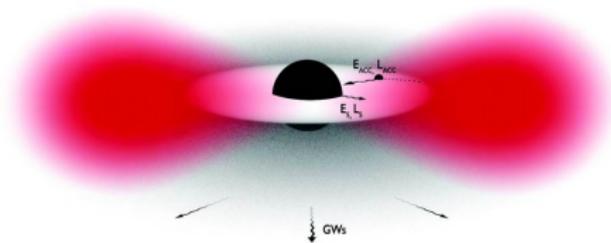
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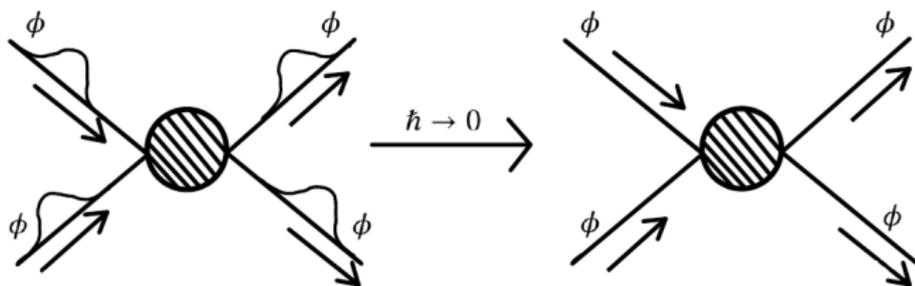


- Phenomenologically relevant for gravitational physics: **Light deflection** around a classical object, **superradiance**, ...



Classical limit for massive particles

The classical limit for external massive particles can be achieved by using appropriate **minimum-uncertainty wavefunctions** which are **localized on the classical trajectory** [Kosower, Maybee, O'Connell; Krivitski, Tsytovich]



$$\phi(p) = \mathcal{N} m^{-1} \exp\left[-\frac{p \cdot u}{\hbar \ell_c / \ell_w^2}\right] \xrightarrow{\text{rest frame}} \mathcal{N}' \exp\left(-\frac{p^2}{2m^2 \ell_c^2 / \ell_w^2}\right)$$

where p is the 4-momentum, ℓ_c is the Compton wavelength and ℓ_w is related to the intrinsic spread of the wavefunction.

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Solution:

Realize that a single massless particle is *never* classical: only an infinite superposition of them can be! (\rightarrow **Coherent state**)

Why coherent states?(I)

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Equation of motion $\square\phi(x) = \lambda j(x)\phi(x)$:

$$\begin{aligned}\phi(x) &= \phi_{\text{in}}(x) + \lambda \int d^4x' G_{\text{ret}}(x-x')j(x') \\ &= \phi_{\text{out}}(x) + \lambda \int d^4x' G_{\text{adv}}(x-x')j(x')\end{aligned}$$

where $G_{\text{ret}}(x)/G_{\text{adv}}(x)$ are the retarded/advanced propagators and $\phi_{\text{in}}(x)$ (resp. $\phi_{\text{out}}(x)$) is the field configuration at $t \rightarrow -\infty$ (resp. $t \rightarrow +\infty$). We have then

$$\phi_{\text{out}}(x) = \phi_{\text{in}}(x) + \lambda \int d^4x' G_{\text{F}}(x-x')j(x')$$

where $G_{\text{F}}(x)$ is the Feynman propagator.

Why coherent states?(II)

There must exist a unitary transformation which connects in and out states

$$\phi_{\text{out}}(x) = S^\dagger \phi_{\text{in}}(x) S \quad |out\rangle = S |in\rangle$$

Solution:

$$S = \exp \left(-i\lambda \int d^4x \phi_{\text{in}}(x) j(x) \right)$$

This means that if we start at $t \rightarrow -\infty$ with a vacuum state (i.e. no incoming wavepacket) we will create a coherent state

$$\begin{aligned} |out\rangle &= \exp \left(-i\lambda \int d^4x \phi_{\text{in}}(x) j(x) \right) |0\rangle = \\ &= \exp \left(-\frac{1}{2} \int d\Phi(k) \lambda^2 |j(k)|^2 \right) \underbrace{\exp \left(-i\lambda \int d\Phi(k) j(k) a^\dagger(k) \right)}_{\text{Infinite superposition of massless modes}} |0\rangle \end{aligned}$$

Glauber Sudarshan P-representation and S-matrix

- **Theorem(Glauber,1963):** Every quantum state of radiation in QFT (i.e. every density matrix) can be written as a superposition of coherent states

$$\hat{\rho}_{\text{radiation}} = \int P(\alpha) |\alpha\rangle \langle \alpha| d\Re(\alpha) d\Im(\alpha)$$

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- **(Hillery,1985)** The superposition of pure states is trivial, i.e. a classical pure state can be written as a single coherent state

$$P(\alpha) = c_{\star} \delta^2(\alpha - \alpha_{\star})$$

Example: Coherent states from soft expansion

- At leading order in the soft expansion, one can study the interaction hamiltonian of massive scalar particles minimally coupled with gravity in asymptotic limit $|t| \rightarrow +\infty$ [Faddeev,Kulish; Ware,Saotome,Akhoury]

$$H(t) = H_0 + V^{\text{asy}}(t) = H_0 - \int d^3x h^{\mu\nu}(t, \vec{x}) T_{\mu\nu}^{\text{asy}}$$

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- Solving the asymptotic evolution equation for the potential $V^{\text{asy}}(t)$ gives an evolution operator which generates a coherent state of soft gravitons

$$\mathbb{C}_{\alpha,\sigma} |0\rangle = |\alpha_p^\sigma\rangle = \exp \left[\int d\Phi(p) (\alpha_p(p) a_\sigma^\dagger(p) - \text{h.c.}) \right] |0\rangle$$

See also [Addazi,Bianchi,Veneziano;Ciafaloni,Colferai,Veneziano; Monteiro,O'Connell,Veiga,Sergola]

Classical light beam and coherent states (I)

- Let's choose the waveshape $\alpha(k)$ in such a way to have a localized beam of light moving along the z-direction, and symmetrical around the z-axis

$$\alpha(k) = \frac{1}{\hbar^3} |\vec{k}| (2\pi)^3 A_0 \sqrt{2\hbar} \delta_{\sigma_{\parallel}}(\omega - k^z/\hbar) \delta_{\sigma_{\perp}}(k^x/\hbar) \delta_{\sigma_{\perp}}(k^y/\hbar)$$

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- The expectation value of the number operator in a coherent state

$$N_{\gamma} = \langle \alpha^+ | N_{\gamma} | \alpha^- \rangle = \int d\Phi(k) |\alpha(k)|^2 \stackrel{k=\hbar\bar{k}, \bar{\alpha}=\alpha\hbar^{\frac{3}{2}}}{=} \frac{1}{\hbar} \int d\Phi(\bar{k}) |\bar{\alpha}(\bar{k})|^2.$$

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- The momentum carried by the classical wave

$$k_{\gamma}^{\mu} = \langle \alpha^+ | \mathbb{K}^{\mu} | \alpha^- \rangle = \int d\Phi(k) |\alpha(k)|^2 k^{\mu} \stackrel{k=\hbar\bar{k}, \bar{\alpha}=\alpha\hbar^{\frac{3}{2}}}{=} \int d\Phi(\bar{k}) |\bar{\alpha}(\bar{k})|^2 \bar{k}^{\mu}.$$

is finite in the classical limit.

Classical light beam and coherent states (II)

- The expectation value of the field is

$$\langle \alpha^+ | \hat{A}_\mu(x) | \alpha^- \rangle = \frac{1}{\sqrt{\hbar}} \int d\Phi(k) [\alpha(k) \varepsilon_\mu^{+,*}(k) e^{-ik \cdot x / \hbar} + \alpha^*(k) \varepsilon_\mu^+(k) e^{+ik \cdot x / \hbar}]$$
$$\stackrel{\sigma_{\parallel} \rightarrow 0}{=} \sqrt{2} A_0 \Re \left[e^{-i\omega(t-z)} \varepsilon_\mu^{+,*}(\bar{k}_0) e^{-(x^2+y^2)/(4\ell_{\perp}^2)} \right]$$

where we require that the transverse momentum components should be subdominant (analogue of the [Goldilocks conditions for a localized wave](#))

$$\omega = \lambda^{-1} \gg \sigma_{\perp} = \ell_{\perp}^{-1}$$

Factorization and Glauber complete coherence condition

- For a **single coherent state**, **Glauber's complete coherence condition** should hold

$$\langle \alpha^+ | \mathbb{F}^{\mu\nu}(x) \mathbb{F}^{\rho\sigma}(y) | \alpha^- \rangle \simeq \langle \alpha^+ | \mathbb{F}^{\mu\nu}(x) | \alpha^- \rangle \langle \alpha^+ | \mathbb{F}^{\rho\sigma}(y) | \alpha^- \rangle$$

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- Check: for a constant wavelshape ($\sigma_{\parallel} \rightarrow 0, \sigma_{\perp} \rightarrow 0$) we have

$$\begin{aligned} \langle \alpha^+ | \mathbb{F}^{\mu\nu}(x) \mathbb{F}^{\rho\sigma}(y) | \alpha^- \rangle &= \langle \alpha^+ | \mathbb{F}^{\mu\nu}(x) | \alpha^- \rangle \langle \alpha^+ | \mathbb{F}^{\rho\sigma}(y) | \alpha^- \rangle \\ &\quad + 4\hbar \partial^{[\mu} \eta^{\nu][\sigma} \partial^{\rho]} \int d\Phi(\bar{k}) e^{-i\bar{k}\cdot(x-y)}, \end{aligned}$$

so that asking the factorization implies exactly $\hbar \rightarrow 0$.

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- Wave deflection angle $\Delta\Theta_{\gamma}$
- Energy event shapes $\mathcal{E}(\hat{n}) = \int_{-\infty}^{+\infty} du \lim_{r \rightarrow \infty} r^2 T_{uu}(u, r, \hat{n})$

[Belitsky, Korchemsky, Sterman; G., Pokraka]

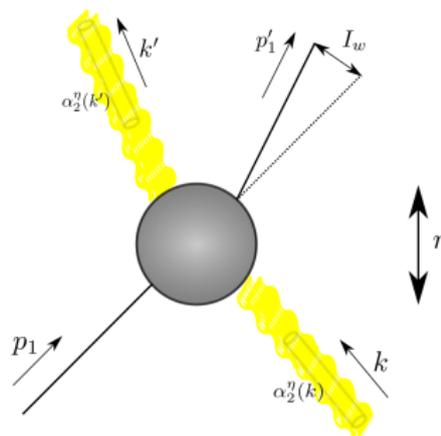
→ connected to the amplitude of the waveform!

Classical impulse

Theory: **Scalar QED**

- Using the **KMOC formalism** we can write the **impulse for the massive charged particle** as

$$\begin{aligned} & \langle \Delta p_1^\mu \rangle \\ &= \langle \psi_w | i[\mathbb{P}_1^\mu, T] | \psi_w \rangle + \langle \psi_w | T^\dagger [\mathbb{P}_1^\mu, T] | \psi_w \rangle \\ &= I_{w(1)}^\mu + I_{w(2)}^\mu \end{aligned}$$



where

$$|\psi_w\rangle_{\text{in}} = \int d\Phi(p_1) \phi_1(p_1) e^{ib \cdot p_1 / \hbar} |p_1 \alpha_2^\eta\rangle_{\text{in}}.$$

Classical impulse in wave scattering (II)

- The properties of the **coherent state as a displacement operator**

$$\mathbb{C}_{\alpha,\eta}^\dagger a_\rho(k) \mathbb{C}_{\alpha,\eta} = a_\eta(k) + \delta_{\eta\rho} \alpha(k), \quad \mathbb{C}_{\alpha,\eta}^\dagger a_\rho^\dagger(k) \mathbb{C}_{\alpha,\eta} = a_\eta^\dagger(k) + \delta_{\eta\rho} \alpha(k)$$

allow us to formulate the problem as an **expansion around the background**

$$\begin{aligned} \langle \Delta p_1^\mu \rangle &:= \int d\Phi(p_1) d\Phi(p'_1) \phi_1(p_1) \phi_1^*(p'_1) e^{-i\frac{q \cdot b}{\hbar}} \langle p'_1 | i[\mathbb{P}_1^\mu, T(A + A_\sigma^{\text{class}})] | p_1 \rangle + \\ &+ \int d\Phi(p_1) d\Phi(p'_1) \phi_1(p_1) \phi_1^*(p'_1) e^{-i\frac{q \cdot b}{\hbar}} \langle p'_1 | T^\dagger(A + A_\sigma^{\text{class}}) [\mathbb{P}_1^\mu, T(A + A_\sigma^{\text{class}})] | p_1 \rangle \end{aligned}$$

where $q := p'_1 - p_1$ and $T(A + A_\sigma^{\text{class}})$ is a short notation for the **background field expansion at the level of the interaction lagrangian**.

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- Note: $I_{w(1)}^\mu$ is always proportional to the **two point function in the classical background $\alpha(k)$** !

Classical field strength in Thomson scattering

- Classically, we solve **perturbatively** the equations of motion:

$$\frac{dp^\mu}{d\tau} = g F^{\mu\nu}(x(\tau))u_\nu(\tau) \quad x(0) = u(0)\tau$$
$$\partial^\mu F_{\mu\nu} = \int d\tau g u_\nu \delta^4(x - x(\tau))$$

where $F_{\mu\nu} = \partial_{[\mu}A_{\nu]}$ and $A^{\mu,(-)}(x) = \int d\Phi(k_1) \alpha(k_1) (\varepsilon_\mu^{-,*}(k_1) e^{-ik_1 \cdot x} + h.c.)$.

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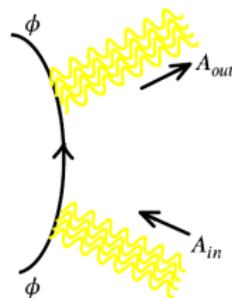
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- The solution at order g^2 is

$$A^\nu(x) = -\frac{g^2}{m} \int d\Phi(k_1) \alpha(k_1) \frac{\delta(k_1 \cdot u - p_1 \cdot u)}{k_1^2 + i\epsilon} \left[-\varepsilon^\nu(k_1) \right. \\ \left. + \frac{(u \cdot \varepsilon(k_1)) p_1^\nu + (p_1 \cdot \varepsilon(k_1)) u^\nu}{p_1 \cdot u + i\epsilon} - \frac{u^\nu [(p_1 \cdot k_1) (\varepsilon(k_1) \cdot u)]}{(p_1 \cdot u + i\epsilon)^2} \right]$$



Classical field strength in Compton scattering (I)

- The incoming state is

$$|\psi\rangle_{\text{in}} = \int d\Phi(p_1) \phi_1(p_1) e^{ib \cdot p_1 / \hbar} |p_1 \alpha^-\rangle$$

and using the KMOC formalism

$$\langle \Delta F^{\mu\nu(+)}(x) \rangle = i \langle \psi | [\mathbb{F}^{\mu\nu(+)}, T] | \psi \rangle + \langle \psi | T^\dagger [\mathbb{F}^{\mu\nu(+)}, T] | \psi \rangle$$

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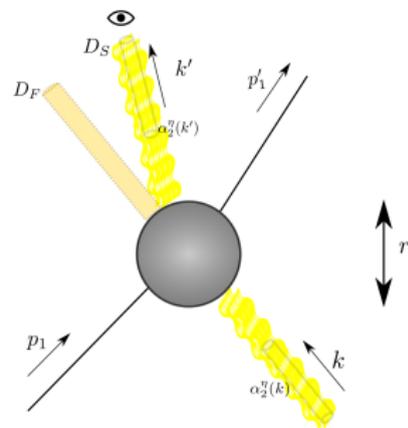
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- we get two types of terms up to g^2

$$\langle \Delta F^{\mu\nu(+)}(x) \rangle_1 = \frac{1}{\hbar^{3/2}} 2\Re \left\{ \int d\Phi(k) d\Phi(p_1) d\Phi(p'_1) \phi(p_1) \phi^*(p'_1) e^{-i \frac{b \cdot q}{\hbar}} \right. \\ \left. \times [k^\mu \varepsilon^{(+)\nu}]^* \langle p'_1 \alpha_2^+ | a_+^\dagger(k) T | p_1 \alpha_2^- \rangle e^{-ik \cdot x / \hbar} \right\}$$

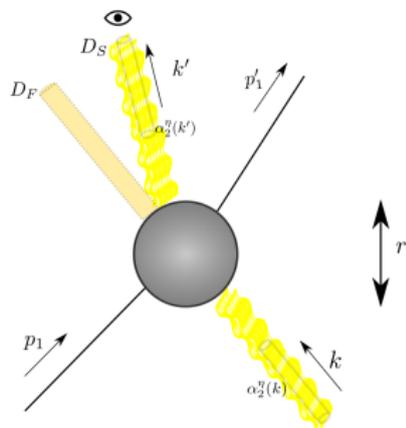
Classical field strength in Compton scattering (II)

We cannot observe a **forward scattered photon** [Frantz;Moncrief]: therefore we should restrict the domain of $\mathbb{F}^{\mu\nu}$ to the **scattered region**!

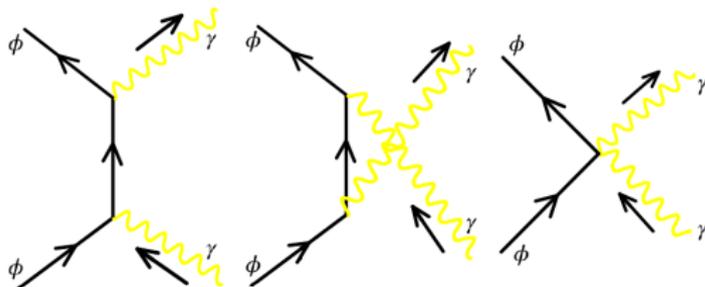


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Using the Compton scattering amplitude for the evaluation of $\langle \Delta F^{\mu\nu(+)}(x) \rangle_{1, D_S}$



Agreement in the classical limit with the Thomson scattering!

Waveform and localized detectors(I)

- In the original KMOC paper *global observables* were introduced

$$R^\mu \equiv \langle k^\mu \rangle =_{\text{in}} \langle \psi | S^\dagger \mathbb{K}^\mu S | \psi \rangle_{\text{in}}$$

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- The radiation's spectral function $f(\omega, \vec{n})$ can be obtained from the waveform via a Fourier transform,

$$f(\omega, \vec{n}; \vec{x}) = \int_{-\infty}^{+\infty} dt W(t, \vec{n}; \vec{x}) e^{i\omega t}$$

Waveform and localized detectors(II)

- Let's consider two massive charged particles scattering each other

$$|\psi\rangle_{\text{in}} = \int d\Phi(p_1)d\Phi(p_2)\phi_1(p_1)\phi_2(p_2)e^{i\frac{b\cdot p_1}{\hbar}} |p_1p_2\rangle$$

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- From $\langle F_{\mu\nu}(x)\rangle \equiv_{\text{in}} \langle\psi|S^\dagger\mathbb{F}_{\mu\nu}(x)S|\psi\rangle_{\text{in}}$ we have, considering the linear terms

$$\langle F_{\mu\nu}(x)\rangle = \frac{1}{\hbar^{\frac{3}{2}}} 2\Re \int d\Phi(k) [k_{[\mu}\tilde{j}_{\nu]}(k)e^{-ik\cdot x}]$$

where

$$\begin{aligned} \tilde{j}_\mu(k) \equiv \sum_{\sigma} \int d\Phi(p'_1)d\Phi(p'_2)d\Phi(p_1)d\Phi(p_2)\phi_1^*(p'_1)\phi_2^*(p'_2)\phi_1(p_1)\phi_2(p_2) \\ \times e^{-i\frac{b\cdot(p'_1-p_1)}{\hbar}} \varepsilon_{\mu}^{*(\sigma)}(k) \langle p'_1p'_2k^\sigma|T|p_1p_2\rangle \end{aligned}$$

Waveform and localized detectors(III)

- Assuming that the measurement distance is much larger than the impact parameter, so that there is a unique and well-defined direction, and using

$$G_{\text{ret}}(x) = i\theta(x^0) \int d\Phi(k) (e^{-ik \cdot x} - e^{ik \cdot x}) = \frac{1}{4\pi|\vec{x}|} \delta(x^0 - |\vec{x}|)$$

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- we get for $x^0 > 0$

$$\begin{aligned} \langle F_{\mu\nu}(x) \rangle &= \int d^4y G_{\text{ret}}(x-y) \left[\frac{\partial}{\partial y^\mu} j_\nu(y) - \frac{\partial}{\partial y^\nu} j_\mu(y) \right] = \\ &= -\frac{i}{4\pi|\vec{x}|} \int \hat{d}\omega e^{-i\omega u} k_{[\mu} \tilde{j}_{\nu]}(k) \Big|_{k=(\omega, \omega \hat{x})} \end{aligned}$$

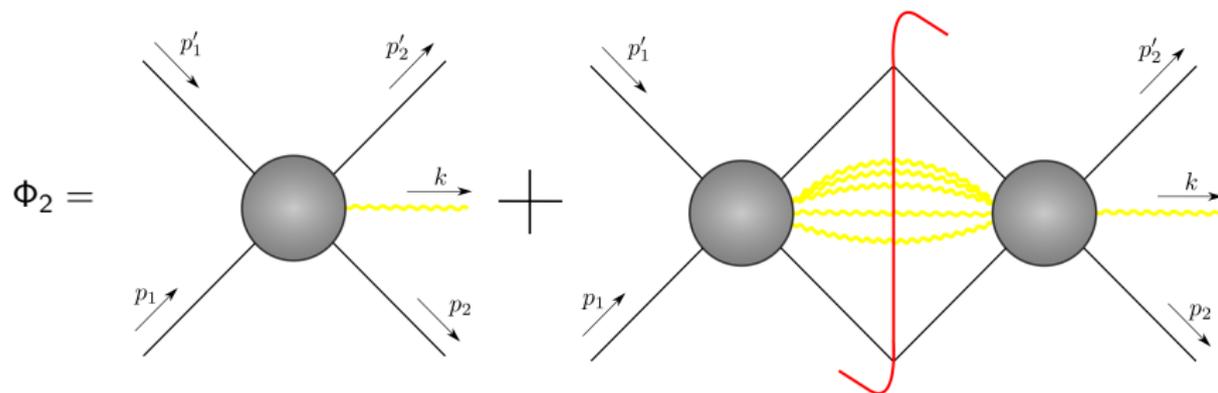
where $u = x^0 - |\vec{x}|$ is the retarded time (\rightarrow Spectral function).

This is the structure of the waveform written in terms of scattering amplitudes! (see also [Maggiore; Jakobsen, Mogull, Plefka, Steinhoff; Mougiakakos, Maria Riva, Vernizzi] for a complementary approach)

- Contracting the field strength with an appropriate **tetrad at infinity**

$$K = \frac{k}{\omega} = (1, \hat{n}) \quad M = \varepsilon_+ \quad \bar{M} = (\varepsilon_+)^* = \varepsilon_- \quad K \cdot N = 1 = -M \cdot \bar{M}$$

we can define $\Phi_2(\omega, r, \hat{n}) = F_{\mu\nu}(\omega, r, \hat{n}) \bar{M}^\mu N^\nu$. Schematically



EM memory effect (I)

- The leading **EM memory effect** is related with the **Weinberg soft factor**

$$\begin{aligned} \tilde{j}_\mu(k) &\stackrel{\omega \rightarrow 0}{\simeq} \sum_\sigma \int d\Phi(p_1) d\Phi(p_2) |\phi_1(p_1)|^2 |\phi_2(p_2)|^2 \\ &\times \int \widehat{d^4 q_1} \widehat{d^4 q_2} \hat{\delta}(2p_1 \cdot q_1) \hat{\delta}(2p_2 \cdot q_2) e^{-i \frac{b \cdot q_1}{\hbar}} \varepsilon_\mu^{*,(\sigma)}(k) \hat{\delta}^4(k - q_1 - q_2) \\ &\times g \left[\frac{\varepsilon^{*,(\sigma)}(k) \cdot (p_1 - q_1)}{k \cdot (p_1 - q_1)} - \frac{\varepsilon^{*,(\sigma)}(k) \cdot p_1}{k \cdot p_1} + (1 \leftrightarrow 2) \right] \langle p'_1 p'_2 | T | p_1 p_2 \rangle \end{aligned}$$

- Expanding the soft factor in the **classical limit** $q_1 = \hbar \bar{q}_1$ we get

$$g \left[\underbrace{q_1^\mu}_{\text{Impulse!}} \left(-\frac{1}{k \cdot p_1} \varepsilon_\mu^{*,(\sigma)}(k) + \frac{\varepsilon^{*,(\sigma)}(k) \cdot p_1}{(k \cdot p_1)^2} k_\mu \right) \right]$$

EM memory effect (II)

- We can then compute the leading expectation value of $\langle \Phi_2 \rangle$

$$\langle \Phi_2(\omega \simeq 0, r, \hat{n}) \rangle = \frac{g}{4\pi r} \left[\left(\frac{1}{k \cdot p_1} \right) l_{(1)} \cdot \varepsilon_{\mu}^{*,(\sigma)}(k) - \left(\frac{\varepsilon^{*,(\sigma)}(k) \cdot p_1}{(k \cdot p_1)^2} \right) l_{(1)} \cdot k + (1 \leftrightarrow 2) \right]$$

which is exactly the **electromagnetic memory effect** (we checked that it agrees with the purely classical prediction, see also [Guevara, Bautista]).

Summary and future directions

- Since massless particles have **no rest frame** and due to **Newton-Wigner theorem**, the classical limit requires to work with **coherent states**.
- We can **scatter coherent states** with a definite waveshape and build a notion of **impulse** and of **field strength** from KMOC formalism which connects smoothly to the **classical wave scattering observables**. Interesting IR safe observables defined for localized observers like the energy event shape.
- **The waveform and its spectral function** have a natural expression in terms of **scattering amplitudes**; the **memory effect** is naturally included in our framework.
- For the future: focus purely on the gravitational case. Understand better the role of coherent states, classical exponentiation to calculate more efficiently our observables in terms of on-shell data.