

# Strange mesons in nuclei and neutron stars



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## Strange mesons: K and K

An (anti-)kaon is a meson containing one (strange)anti-strange quark



| $\begin{array}{c} \textbf{Mesons } q\overline{q} \\ \textbf{Mesons are bosonic hadrons} \\ \textbf{These are a few of the many types of mesons.} \end{array}$ |        |               |                 |                            |      |  |  |
|---|--------|---------------|-----------------|----------------------------|------|--|--|
| Symbol  | Name   | Quark content | Electric charge | Mass<br>GeV/c <sup>2</sup> | Spin |  |  |
| π+  | pion   | ud            | +1              | 0.140                      | 0    |  |  |
| K-  | kaon   | sū            | -1              | 0.494                      | 0    |  |  |
| ρ+  | rho    | ud            | +1              | 0.776                      | 1    |  |  |
| $\mathbf{B}^0$  | B-zero | db            | 0               | 5.279                      | 0    |  |  |
| η <sub>c</sub>  | eta-c  | cē            | 0               | 2.980                      | 0    |  |  |



## Where to find strange mesons?









FINUDA, DISTO, OBELIX, J-PARC..

# Outline

- $\overline{K}N$  interaction:  $\Lambda(1405)$  resonance
- KNN bound state
- Kaons and Antikaons in matter
- Experiments and observations: from atoms to stars
- Bibliography

# $\overline{K}N$ interaction: the $\Lambda(1405)$

•  $\overline{\text{KN}}$  scattering in the I=0 channel is governed by the presence of the  $\Lambda(1405)$  resonance, located only 27 MeV





 $\Lambda^{*}(1405)$ 

K=

s=-1

- 50's: idea originally proposed by Dalitz and Tuan
- since 90's: the study of KN scattering has been revisited by means of unitarized theories using meson-exchange models or chiral Lagrangians

#### meson-exchange models

Mueller-Groeling, Holinde and Speth '90; Buettgen, Holinde, Mueller-Groeling, Speth and Wyborny '90; Hoffmann, Durso, Holinde, Pearce and Speth '95; Haidenbauer, Krein, Meissner and Tolos '11..

#### chiral Lagrangian

Kaiser, Siegl and Weise, '95; Oset and Ramos '98; Oller and Meissner '01; Lutz, and Kolomeitsev '02; Garcia-Recio et al. '03; Jido et al. '03; Borasoy, Nissler, and Weise '05; Oller, Prades, and Verbeni '05; Oller '06; Borasoy, Nissler and Weise '05; Khemchandani, Martinez-Torres, Nagahiro, Hosaka '12 Feijoo, Magas and Ramos '19....

more channels, next-to-leading order, Born terms beyond WT (s-channel, u-channel), fits including new data

....

## **K**N interaction: meson-exchange model

#### The One Boson Exchange model

to consider the exchange of bosons between baryon-baryon (or meson-baryon or meson-meson) within quantum field theory in terms of perturbation theory using Feynman diagrams

## Juelich meson-exchange model

Mueller-Groeling, Holinde and Speth '90



Mueller-Groeling, Holinde and Speth '90; Buettgen, Holinde, Mueller-Groeling, Speth and Wyborny '90; Hoffmann, Durso, Holinde, Pearce and Speth '95; Haidenbauer, Krein, Meissner and Tolos '11..





#### $\phi = (\pi, K, \eta)$ KN interaction: SU(3) chiral Lagrangian $B = (N, \Lambda, \Sigma, \Xi)$ $\mathcal{L}_{\phi B}^{eff} = \mathcal{L}_{\phi B}^{(1)} + \mathcal{L}_{\phi B}^{(2)} ,$ $[D_{\mu}, B] = \partial_{\mu}B + [\Gamma_{\mu}, B]$ $\Gamma_{\mu} = \frac{1}{2} [u^{\dagger}, \partial_{\mu} u]$ $\mathcal{L}_{\phi B}^{(1)} = i \langle \bar{B} \gamma_{\mu} [D^{\mu}, B] \rangle - M_0 \langle \bar{B} B \rangle - \frac{1}{2} D \langle \bar{B} \gamma_{\mu} \gamma_5 \{ u^{\mu}, B \} \rangle \qquad u_{\mu} = i u^{\dagger} \partial_{\mu} U u^{\dagger} U u^{\dagger}$ $U(\phi) = u^2(\phi) = \exp\left(\sqrt{2}i\phi/f\right)$ LO f weak decay constant order $-\frac{1}{2}F\langle \bar{B}\gamma_{\mu}\gamma_{5}[u^{\mu},B]\rangle$ , $g_A = D + F = 1.26$ $M_0$ common baryon octet mass $\mathcal{L}_{\phi B}^{(2)} = b_D \langle \bar{B} \{ \chi_+, B \} \rangle + b_F \langle \bar{B} [\chi_+, B] \rangle + b_0 \langle \bar{B} B \rangle \langle \chi_+ \rangle$ $\chi_{+} = 2B_{0}(u^{\dagger}\mathcal{M}u^{\dagger} + u\mathcal{M}u)$ $\mathcal{M} = \text{diag}(m_u, m_d, m_s)$ $+ d_1 \langle \overline{B} \{ u_\mu, [u^\mu, B] \} \rangle + d_2 \langle \overline{B} [u_\mu, [u^\mu, B]] \rangle$ NLO $B_0 = -\langle 0|\bar{q}q|0\rangle/f^2$ order $b_D, b_F, b_0$ and $d_i$ (*i* = 1, ..., 4) $+ d_3 \langle \bar{B} u_\mu \rangle \langle u^\mu B \rangle + d_4 \langle \bar{B} B \rangle \langle u^\mu u_\mu \rangle$ . NLO low-energy constants



## **KN interaction:** SU(3) chiral Lagrangian $\phi = (\pi, K, \eta)$ $B = (N, \Lambda, \Sigma, \Xi)$



$$B = (N, \Lambda, \Sigma, D)$$
$$[D_{\mu}, B] = \partial_{\mu}B + [\Gamma_{\mu}, B]$$
$$\Gamma_{\mu} = \frac{1}{2}[u^{\dagger}, \partial_{\mu}u]$$
$$u_{\mu} = iu^{\dagger}\partial_{\mu}Uu^{\dagger}$$
$$U(\phi) = u^{2}(\phi) = \exp(\sqrt{2}i\phi/f)$$
$$f \text{ weak decay constant}$$
$$g_{A} = D + F = 1.26,$$
$$M_{0} \text{ common baryon octet mass}$$

$$\chi_{+} = 2B_{0}(u^{\dagger}\mathcal{M}u^{\dagger} + u\mathcal{M}u)$$
  

$$\mathcal{M} = \text{diag}(m_{u}, m_{d}, m_{s}),$$
  

$$B_{0} = -\langle 0|\bar{q}q|0\rangle/f^{2}$$
  

$$b_{D}, b_{F}, b_{0} \text{ and } d_{i} (i = 1, ..., 4)$$
  
NLO low-energy constants



WT most important contribution, Born terms and NLO contribution needed for fine tuning Fine tuning crucial for description of shift and width of the 1s state kaonic hydrogen

## **Unitarization in coupled channels**



## Shift and width of the 1s state of the kaonic hydrogen

The SIDDHARTA collaboration at DAΦNE collider has determined the most precise values of the **shift and width of the 1s state of the kaonic hydrogen**, clarifying the discrepancies between KEK and DEAR results



SIDDHARTA results provide important constraints on theoretical descriptions

Ikeda, Hyodo and Weise '12 Guo and Oller '13 Mai and Meissner '13 Feijoo, Magas and Ramos '15

## Double-pole structure of Λ(1405)

 $\Lambda(1405)$  results from the superposition of two poles in the complex plane,



with different coupling to  $\pi\Sigma$  and  $\overline{K}N$  states

Pole positions for the  $\Lambda(1405)$  coming from recent chiral effective models including the SIDDHARTA constraint.

| Model          |                                  | First Pole [MeV]                     | Second Pole [MeV]                    |
|----------------|----------------------------------|--------------------------------------|--------------------------------------|
| NLO            | Ikeda, Hyodo and Weise '12       | $1424^{+7}_{-23} - i26^{+3}_{-14}$   | $1381^{+18}_{-6} - i81^{+19}_{-8}$   |
| Fit II         | Guo and Oller '13                | $1421^{+3}_{-2} - i19^{+8}_{-5}$     | $1388^{+9}_{-9} - i114^{+24}_{-25}$  |
| Solution       | <b>Nr. 2</b> Mai and Meissner '1 | $_{5}1434^{+2}_{-2} - i10^{+2}_{-1}$ | $1330^{+4}_{-5} - i56^{+17}_{-11}$   |
| Solution Nr. 4 |                                  | $1429^{+8}_{-7} - i  12^{+2}_{-3}$   | $1325^{+15}_{-15} - i90^{+12}_{-18}$ |



comparison of the measured spectra of the  $\Sigma\pi$  final states associated to the  $\Lambda(1405)$  for kaon- and pion-induced reactions supports the double-pole structure of the  $\Lambda(1405)$ 

Magas, Oset and Ramos '05

## **Λ(1405) production**



Fit to photoproduction data, where  $\pi^0\Sigma^0$  is in red,  $\pi^-\Sigma^+$  in blue and  $\pi^+\Sigma^-$  in green Roca and Oset '13 Data from CLAS: Moriva et al '13

## **Photo-induced reactions**

**CLAS** Moriya et al '13  $\gamma + p \rightarrow \Lambda(1405)K^+ \rightarrow (\Sigma\pi)^0 K^+$ 

Theory: Nacher et al. '99, Roca and Oset '13, Nakamura and Jido '14, Mai and Meissner '15...  $\Lambda(1405)$  photoproduction reactions most sensitive to high-energy pole

Kaon-induced reactions sensitive to high-energy pole Magas, Oset and Ramos '05, whereas Pion-induced reactions most sensitive to the low-energy pole Hyodo et al. '03

#### Kaon-induced reactions in deuteron J-PARC (E31) Naruki et al. '12

 $K^-d \to \pi \Sigma n$ 

**Theory:** Jido, Oset and Sekihara '09 '13, Miyagawa and Haidenbauer '12, Ohnishi et al '16, Miyagawa, Haidenbauer and Kamada '18..

## Proton-proton collisions



## **KNN bound state**

if the KN interaction is so attractive, the K-nuclear clusters may form  $\rightarrow$  The KNN (I=1/2) state



## thoroughly addressed theoretically

Akaishi, Yamazaki, Shevchenko, Gal, Mares, Revai, Ikeda, Sato, Kamano, Dote, Hyodo, Weise, Wycech, Green, Bayar, Oset, Ramos, Yamagata-Sekihara, Barnea, Liverts, Dote, Inoue, Myo, Uchino, Hyodo, Oka..

initial claims by FINUDA, DISTO and OBELIX, that could find alternative conventional explanation Ramos et al '08 or not be reproduced Agakishiev et al [HADES] '15

more recent experiments did not find any Tokiyasu et al. [Spring8/LEPS] '14; Hashimoto et al [JPARC E15] '15; Vazquez-Doce et al. [AMADEUS] '16 or if found Ichikawa et al [J-PARC E27] '15; Nagae et al [J-PARC E27] '16 may have other interpretation Garcilazo et al '13

J-PARC E15 has found a structure near KNN threshold <u>Sa</u>da et al [J-PARC E15] '16 being interpreted as KNN bound state Sekihara et al '16



Binding energy and width of K<sup>-</sup>pp for different chiral and phenomenological calculations using variational, Faddeev or ccCSM+Feshbach methods. Tolos and Fabbietti '20

| Work              | B [MeV]   | Г [MeV] | Method      | Type of potential       |
|-------------------|-----------|---------|-------------|-------------------------|
| Barnea et al.     | 16        | 41      | Variational | Chiral                  |
| Dote et al.       | 17–23     | 40-70   | Variational | Chiral                  |
| Dote et al.       | 14–50     | 16-38   | ccCSM       | Chiral                  |
| Ikeda et al.      | 9–16      | 34-46   | Faddeev     | Chiral                  |
| Bayar et al.      | 15-30     | 75-80   | Faddeev     | Chiral                  |
| Sekihara et al.   | 15–20     | 70–80   | Faddeev     | Chiral                  |
| Yamazaki et al.   | 48        | 61      | Variational | phenomenological        |
| Shevchenko et al. | 50–70     | 90-110  | Faddeev     | Phenomenological        |
| Ikeda et al.      | 60-95     | 45-80   | Faddeev     | Phenomenological        |
| Wycech et al.     | 40-80     | 40-85   | Variational | phenomenological        |
| Dote et al.       | 51        | 32      | ccCSM       | Phenomenological        |
| Revai et al.      | 32/ 47–54 | 50–65   | Faddeev     | Chiral/phenomenological |

Binding energies B~9-95 MeV with decay widths T~16-110 MeV

## Variety of values due to

- uncertainties in subthreshold extrapolation of the KN interaction

(chiral interactions give lower binding energies than phenomenological ones)
 use of variational or Faddeev calculations introduces certain approximations

(full three-body not account for in variational methods, whereas Faddeev calculations deal with separable two-body interactions), and ccCSM combines merits of variational and Faddeev but high computational cost

## Kaons and Antikaons in matter KN interaction in matter

Since no baryonic resonances with positive strangeness exist (assuming no pentaquarks), the KN interaction at low densities can be described by

$$U_K \sim T_{KN-KN} \rho$$
,

This is the so-called low-density theorem or Tp approximation Different models throughout time

## Nambu-Jona-Lasinio (NJL)

10% change in mass at  $\rho = \rho_0$  Lutz, Steiner and Weise '94

## **Relativistic mean-field (RMF)**

repulsive potential similar to NJL at  $\rho_0$  Schaffner, Bondor and Mishustin '97

## Quark-meson-coupling (QMC)

repulsive potential ~20 MeV at  $\rho_0$  Tsushima et al. '98

**Unitarized coupled-channel approaches** Oset, Ramos '98, Kaiser, Siegel, Weise '95 with SU(3) chiral Lagrangian: 10% or less change in mass at  $\rho = \rho_0$ 



Schaffner-Bielich.

Mishustin and Bondorf '97

## **K**N interaction in matter

## Relativistic mean-field, Quark meson coupling models...

RMF: early works based on mesonexchange picture or the chiral approach for the KN interaction on the mean-field level and fit the parameters to the KN scattering length



## Phenomenological models

density dependent potentials fitted to kaonic atoms



recent K-N scattering amplitudes from  $\chi$ SU(3) EFT supplemented with phenomenological terms for K-multinucleon interactions: kaonic atoms test densities  $\rho < \rho_0$ 

Friedman and Gal '17

## **Unitarized theory in matter:**

selfconsistent coupled-channel procedure



The presence of the  $\Lambda(1405)$  resonance makes the in-medium  $\overline{K}N$  interaction very sensitive to the particular details of the many-body treatment.



Important quantities:

Self-energy (what we calculate!)

$$\Pi_{\bar{K}}(q_0, \vec{q}, T) = \int \frac{d^3p}{(2\pi)^3} n(\vec{p}, T) \left[ T_{\bar{K}N}^{(I=0)}(P_0, \vec{P}, T) + 3T_{\bar{K}N}^{(I=1)}(P_0, \vec{P}, T) \right]$$

Optical potential (non-relativistic approach, BHF type in this case)

Spectral function

$$S_{\bar{K}}(q_0, \vec{q}, T) = -\frac{1}{\pi} \operatorname{Im} D_{\bar{K}}(q_0, \vec{q}, T)$$

## K spectral function in matter





Koch '94; Waas and Weise '97; Kaiser et al '97; Oset and Ramos'98; Lutz '98; Schaffner-Bielich et al '00; Ramos and Oset '00; Lutz et al '02; Tolos et al '01 '02; Jido et al '02 '03; Magas et al '05; Tolos et al '06 '08; Lutz et al '08; Cabrera et al '14..

# $\begin{array}{l} \text{Re } U_{\text{K-}}(\rho_0) \thicksim -50 \text{ to } -80 \text{ MeV} \\ \text{Im } U_{\text{K-}}(\rho_0) \gtrsim \text{Re } U_{\text{K-}}(\rho_0) \end{array}$

# **•s-wave** $\overline{K}N$ interaction governed by $\Lambda(1405)$ :

attraction due to modified  $\Lambda(1405)$  in the medium using a self-consistent coupled-channel approach

## p-wave (and beyond)

contributions to KN interaction: not important for atoms but important for heavy-ion collisions due to large momentum

# Experiments and observations: from atoms to stars







## **Kaonic atoms**

Kaonic atoms are atoms in which an electron is replaced by a negatively charged antikaon





best fits to kaonic atoms seem to prefer  $U_{K} \sim -200 \text{ MeV at } \rho_{0}$ 

$$\left(-\nabla^2 + m_K^2 + 2m_K U_K(r) + V_{\text{Coul.}}\right)\Psi_K(r) = E_K^2 \Psi_K(r)$$

However, theoretical models based on the chiral  $\overline{KN}$  interaction with many-body effects (albeit an additional moderate phenomenological piece)  $\rightarrow$  moderate attraction for the K<sup>-</sup>-nucleus potential and **kaonic** atoms are also well described!! Kaonic atoms





Hirenzaki et al. '00; Baca, García-Recio and Nieves '00



kaonic data only constrains the K<sup>-</sup> optical potential for  $\rho < 25\%$  (50%)  $\rho_0$ 

## **Strangeness production in HICs**

## strangeness production in matter

is one of the major research domains in heavy-ion collisions from SIS/GSI to LHC and RHIC up to the future FAIR/NICA/BESII/J-PARC-HI

Iow-energy HICs:Zinyuk (FOPI) '14KaoS/SIS18: K+,K-Foerster et al (KaoS) '07Agakishiev et al (HADES) '13 '14FOPI/SIS18: K+,K- $\phi(1020)$ .HADES/SIS18: K+, K\*(892)<sup>0</sup>,  $\phi(1020)$ ,  $\Xi(1321)$ , Ω,...

#### Early Universe Future LHC Experiments Current RHIC Experiments Current RHIC Experiments Quark-Gluon Plasma Future FAIR Experiments Critical Point Hadron Gas Vacuum UMeV 0 MeV 900 MeV Baryon Chemical Potential

high-energy HICs:Ada<br/>AggSTAR/RHIC: K\*(892)<sup>0</sup>, φ(1020), Ω..Kun<br/>KunALICE/LHC: K\*(892)<sup>0</sup>, φ(1020),  $Σ^{+-}(1385)$ ,  $\Xi(1530)^{0}$ ..Ada<br/>Ada

Adams et al. (STAR) '05 Aggarwal et al (STAR) '11 Kumar et al (STAR) '15 Abelev (ALICE) '15 Adam (ALICE) '16 Badala (ALICE) '17..

#### future:

CBM/FAIR BM@N/NICA BESII/RHIC J-PARC-HI

CBM (FAIR) Physics Book '11 NICA: http://theor0.jinr.ru/twiki-cgi/view/NICA Aggarwal et al (BES STAR White Paper) '10 JPARC: http://silver.j-parc.jp/sako/white-paper-v1.21.pdf-HI

#### credit: DOE

# K<sup>-</sup> and K+ at high $\mu_{B}$ (FOPI/HADES @ SIS18)

KaoS: from systematics of the experimental results and detailed comparison to transport model calculations<sub>150</sub> Foerster et al (KaoS) '07

• K<sup>+</sup> probe a soft EoS

• K<sup>+</sup> and K<sup>-</sup> yields are coupled  $NN \rightarrow K^+YN$ by strangeness exchange:  $K^-N \Leftrightarrow \pi Y$ 

- K<sup>+</sup> and K<sup>-</sup> exhibit different freeze-out conditions
- repulsion for K+ and attraction for K- seemed to be confirmed

but, for example, what is the role of  $\phi \rightarrow K^+ K^-$ ?

#### More recent results from HADES and FOPI indicate Zinyuk et al (FOPI)'14; Gasik et al (FOPI) '16; Piasecki et al (FOPI) '16; Adamczewski-Musch et al (HADES) '17,...

- K<sup>+</sup> in-medium potential is repulsive: U<sub>KN</sub> (ρ<sub>0</sub>)≈ 20...40 MeV
- K<sup>-</sup> from Φ decay wash out the effects of the potential (spectra and flow!!)
- separate direct kaons ( $\rightarrow$  COSY)/elementary reactions
- more systematic, high statistic data on K<sup>-</sup> production necessary conclusions from Leifels-SQM2017



Recent results on kaon and antikaon production in HiCs using a PHSD model with in-medium strange mesons compared to KaoS, FOPI and HADES experimental data

- The nuclear effects on (anti)kaon are more prominent in the collision of large nuclei
- (Anti)kaon production is (enhanced)suppressed due to (broadening of spectral function)repulsive kaon potential
- (Anti)kaon spectrum becomes (softer)harder in nuclear matter, whereas y-distribution (shrinks)broadens
- Different behaviour of v1/v2 for antikaons and kaons due to the attractive vs repulsive character of the interaction with nucleons
- A moderate EoS (K~300 MeV) reproduces the experimental HiC data better



Song et al '21

## Kaon condensation in neutron stars





K<sup>-</sup> feels attraction in the medium
 → Kaon condensation in neutron stars? (

$$n \leftrightarrow p \ e^- \ ar{
u}_e \ o \ \mu_n = \mu_p + \mu_{e^-}$$

$$n \leftrightarrow p K^- \rightarrow \mu_n = \mu_p + \mu_{K^-}$$

Kaons are bosons. If  $\mu_{K} \leq \mu_{e}$  for  $\rho \geq \rho_{c}$ , with  $\rho_{c}$  being a feasible density inside neutron stars, kaons will condensate

 $\overline{K}^0/K^0$  more difficult to be produced as  $\mu_{K0} \le 0$ K<sup>+</sup> even more difficult at  $\mu_{K+} \le -\mu_{e-}$ 



#### Glendenning '85 Kaon condensation irrelevant as (anti)kaons have to lower their mass drastically

#### Kaplan and Nelson '86 In-medium effects on (anti)kaons can be pronounced so as to have kaon condensation

Brown, Kubodera, Rho and Thorsson'92; Thorsson, Prakash and Lattimer '94; Fujii, Maruyama, Muto and Tatsumi '96; Li, Lee and Brown '97; Knorren, Prakash and Ellis '95; Schaffner and Mishustin '96; Glendenning and Schaffner-Bielich '98 '99

#### Renewed interest on antikaon-nucleon interaction with effective field theoretical models



# Effects of hyperonization on kaon condensation

Knorren, Prakash and Ellis '95

electron fraction decreases once hyperons appear, thus, the presence of hyperons increases the critical density for kaon condensation



Later on different groups have worked on improved relativistic-mean field models (with density-dependent or higher-order couplings, with or without hyperons, including three-body forces..), so as to fulfill 2M<sub>sun</sub> neutron star mass observations and, in some cases, to study proto-neutron stars, core-collapse supernova or neutron star mergers

Banik and Bandyopadhyay '01 '02 Char and Banik '14 Malik, Banik and Bandyopadhyay '20 '21

RMF effective model with density-dependent couplings for nucleons, hyperons and kaon condensate



Antikaon potential at saturation density deeper than -120 MeV

Gupta and Arumugam '12 '13

RMF effective models with higher-order couplings for nucleons and kaon condensate. No hyperons are considered



density is deeper than -140 MeV



Thapa and Sinha '20 Thapa, Sinha, Li and Sedrakian '21

RMF model (CDF model) for nucleons and kaon condensate; or for nucleons, hyperons,  $\Delta$  and kaon condensate

Antikaon potential at saturation density is deeper than -120 MeV

Muto '08

Muto, Maruyama, Tatsumi and Takatsuka '19 Muto, Maruyama, Tatsumi and Takatsuka '21

RMF effective model for hyperons and kaon condensate with repulsive threebody forces (SJM) including or not TNA

Antikaon potential at saturation density is deeper than ~ - 100 MeV

Thapa, Sinha, Li and Sedrakian '21



Using microscopic unitarized schemes...

The condition  $\mu_{K_{-}} \sim m^*_{K_{-}} \leq \mu_{e_{-}}$  for a given  $\rho_c$ implies that  $m_{K_{-}} - m^*_{K_{-}}(\rho_c) \approx 200, 300$ MeV. However, unitarized schemes based on meson-exchange models or chiral (MeV) Lagrangians predict a moderate attraction in nuclear matter

Lutz '98 Ramos and Oset '00 Tolos, Polls, Ramos '01 Tolos, Ramos and Oset '06 Tolos, Cabrera and Ramos '08 Cabrera, Tolos, Aichelin and Bratkovskaya'14

#### Therefore,

kaon condensation seems very unlikely within microscopic unitarized schemes



## Bibliography

Laura Tolos and Laura Fabbietti, Progress in Particle and Nuclear Physics 112 (2020) 103770

Other references mentioned in the lecture!