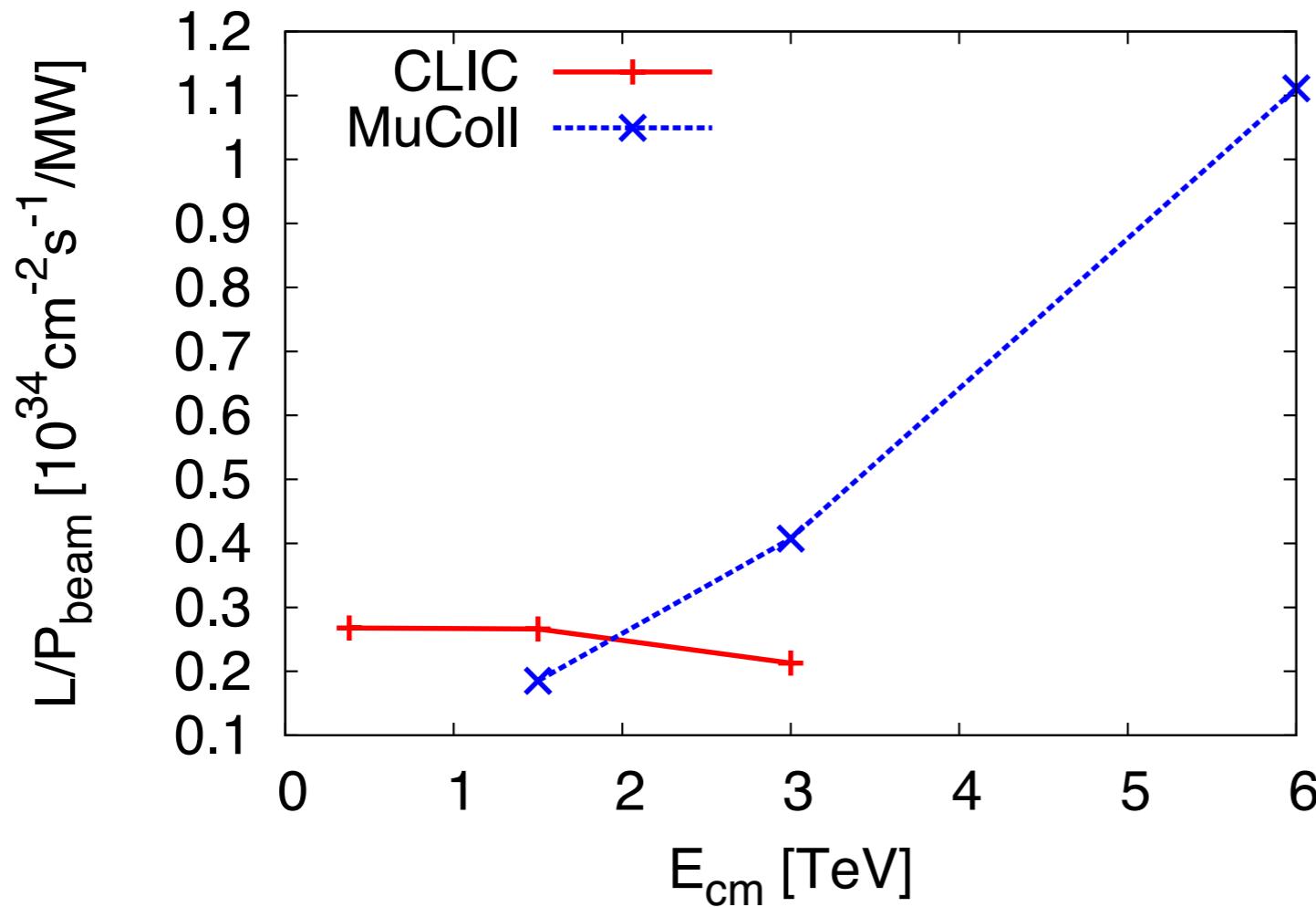


# Muon Collider: physics case

LianTao Wang  
Univ. of Chicago

GGI tea break, March 2, 2022

# The obvious: higher energy, shorter distances

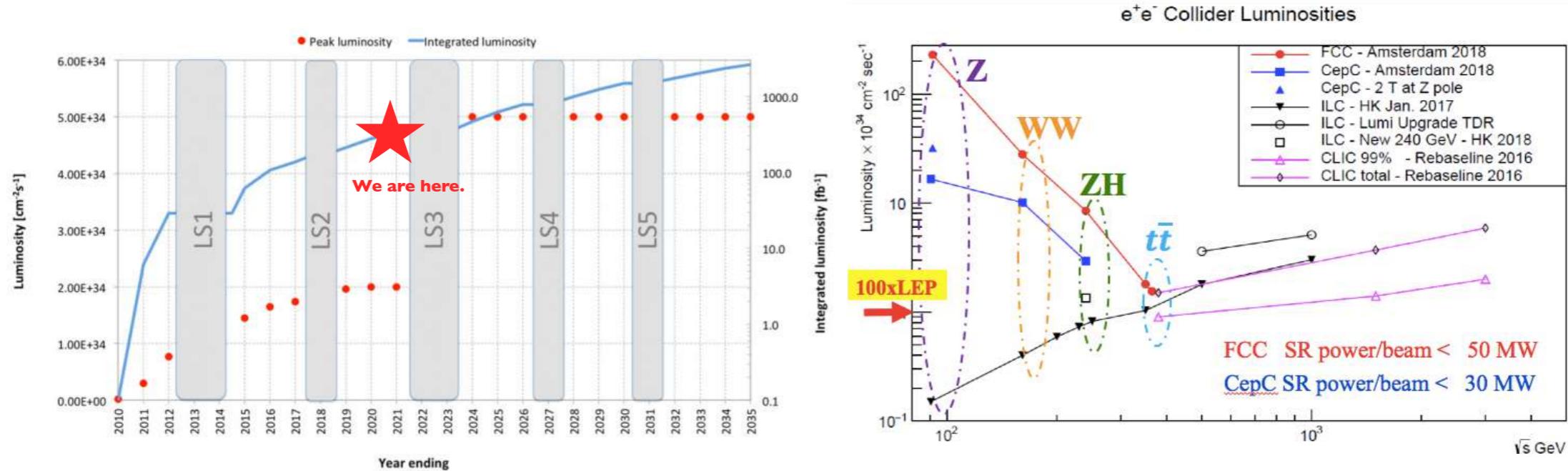


$\sqrt{s}$	$\int \mathcal{L} dt$
3 TeV	1 $\text{ab}^{-1}$
10 TeV	10 $\text{ab}^{-1}$
14 TeV	20 $\text{ab}^{-1}$

Good enough? Yes (for most of us).

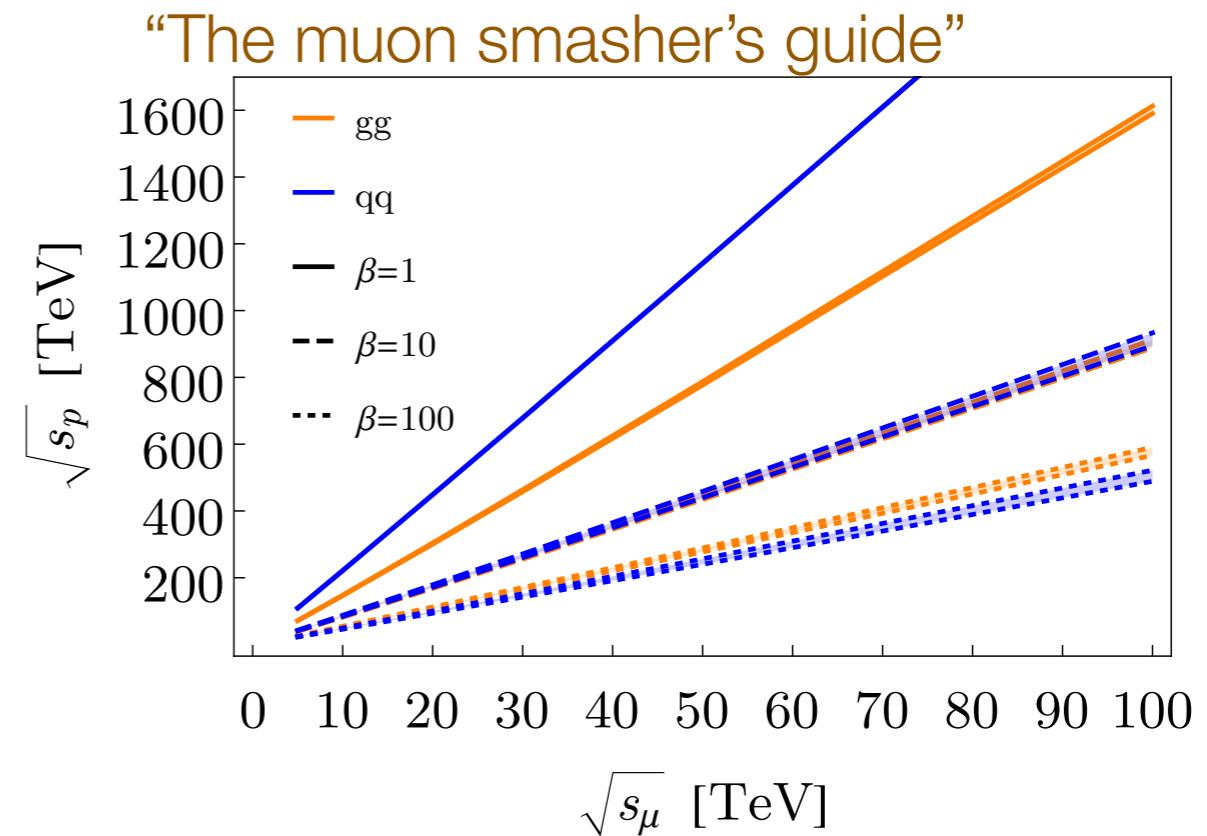
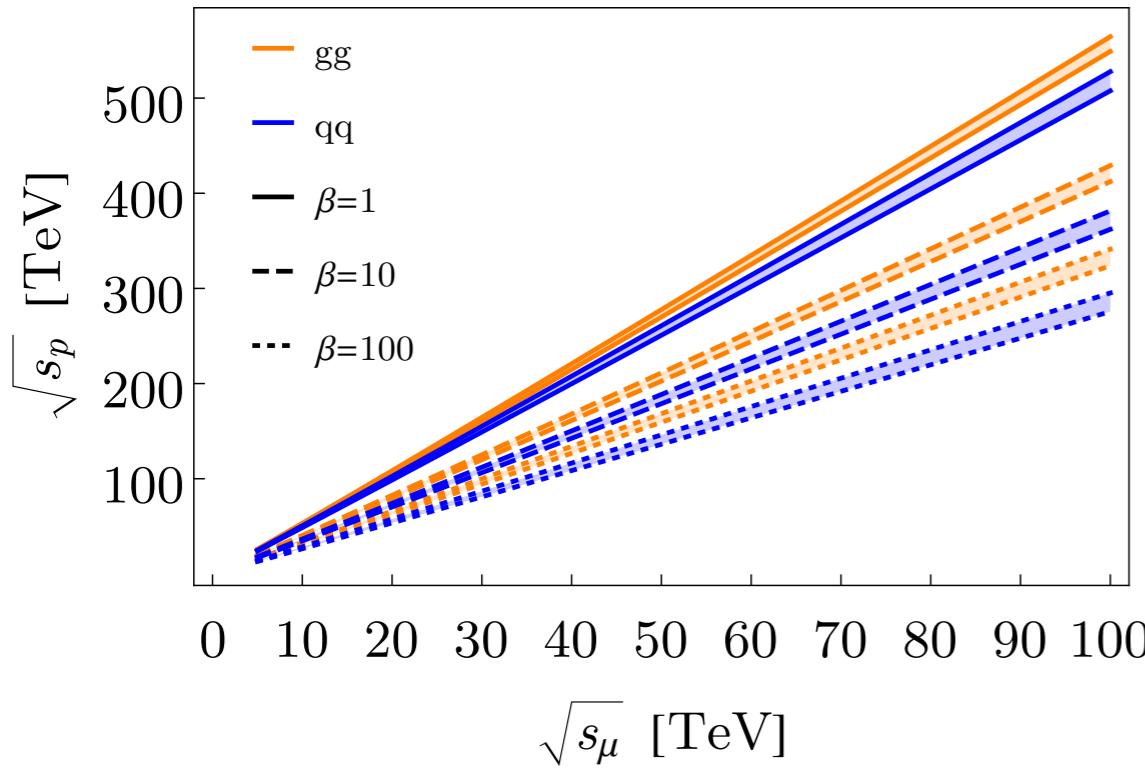
Still, why muon collider at these energies?

# The coming decades



- \* Main “near term” targets: precision, rare processes.
- \* Muon collider, going beyond these options, such as the low energy Higgs factories.

# Comparison with 100-ish TeV pp collider Such as FCC-hh or SPPC



- \* Naively, 100 TeV pp  $\approx$  10+ TeV muon collider.
- \* Lepton collider “cleaner”. Good for precision, difficult channels.

# Physics program at a muon collider

- \* Higgs properties.
- \* New physics at higher energies.

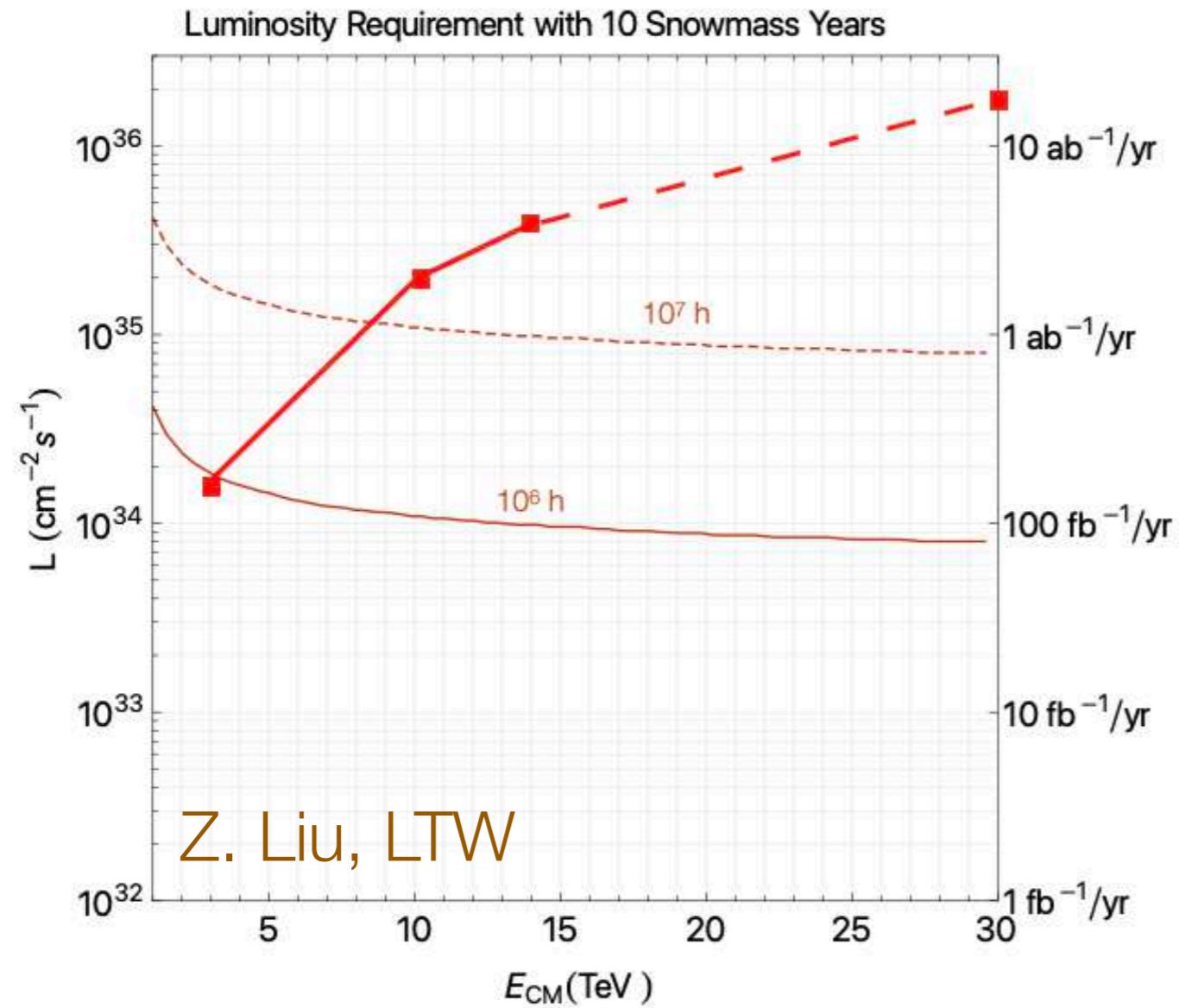
# Higgs properties

- \* Obvious: important to understand the Higgs properties.
  - \* Origin of the weak scale.
  - \* Nature of EW symmetry breaking.
  - \* Portal to dark sector.

# Higgs properties

- \* Obvious: important to understand the Higgs properties.
  - \* Origin of the weak scale.
  - \* Nature of EW symmetry breaking.
  - \* Portal to dark sector.
- \* Not so obvious: what does muon collider bring to the table?

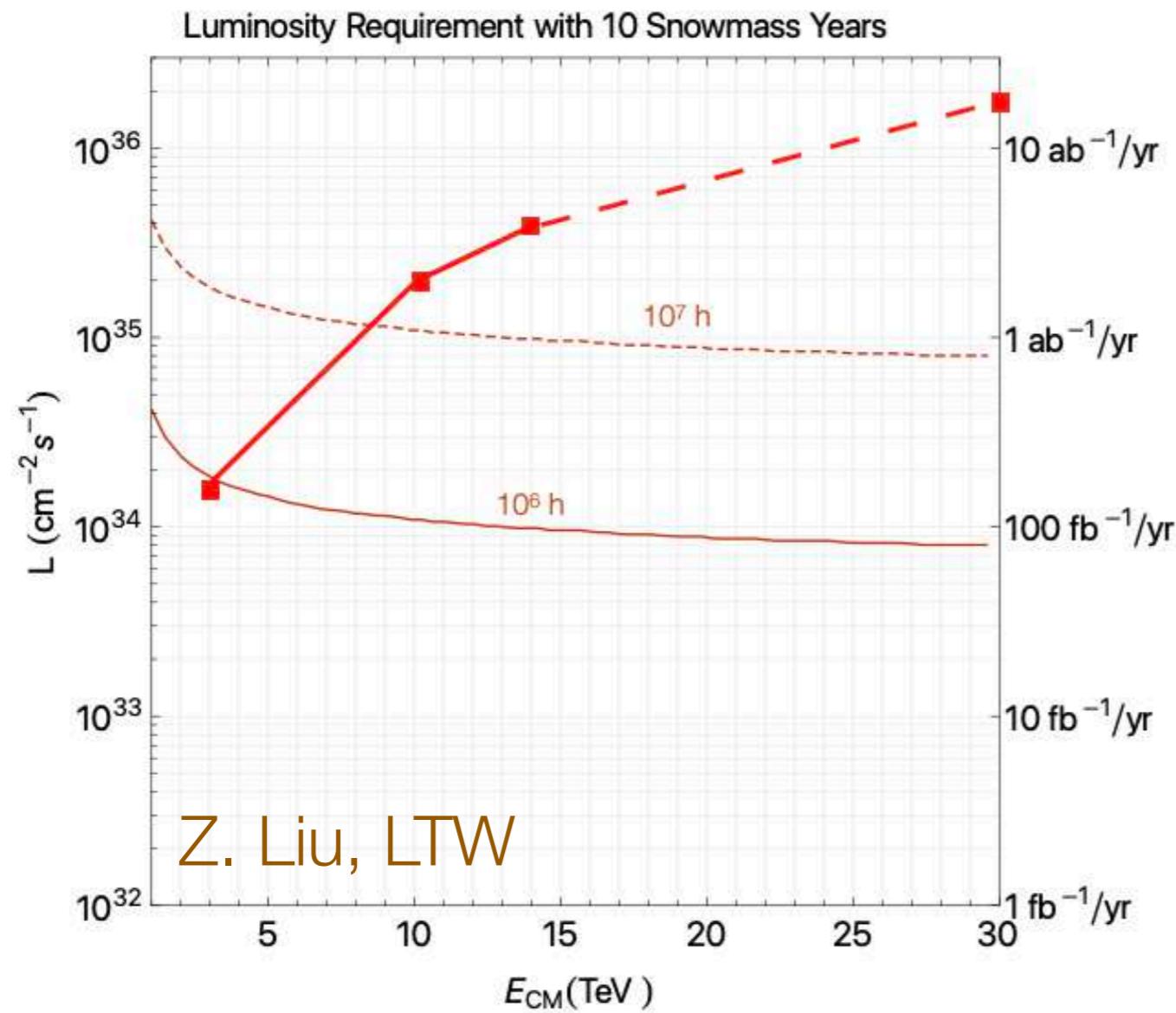
# MC as Higgs factory



In comparison:

low(er) energy Higgs factories  
~  $10^6$  Higgses

# MC as Higgs factory



In comparison:

low(er) energy Higgs factories  
~  $10^6$  Higgses

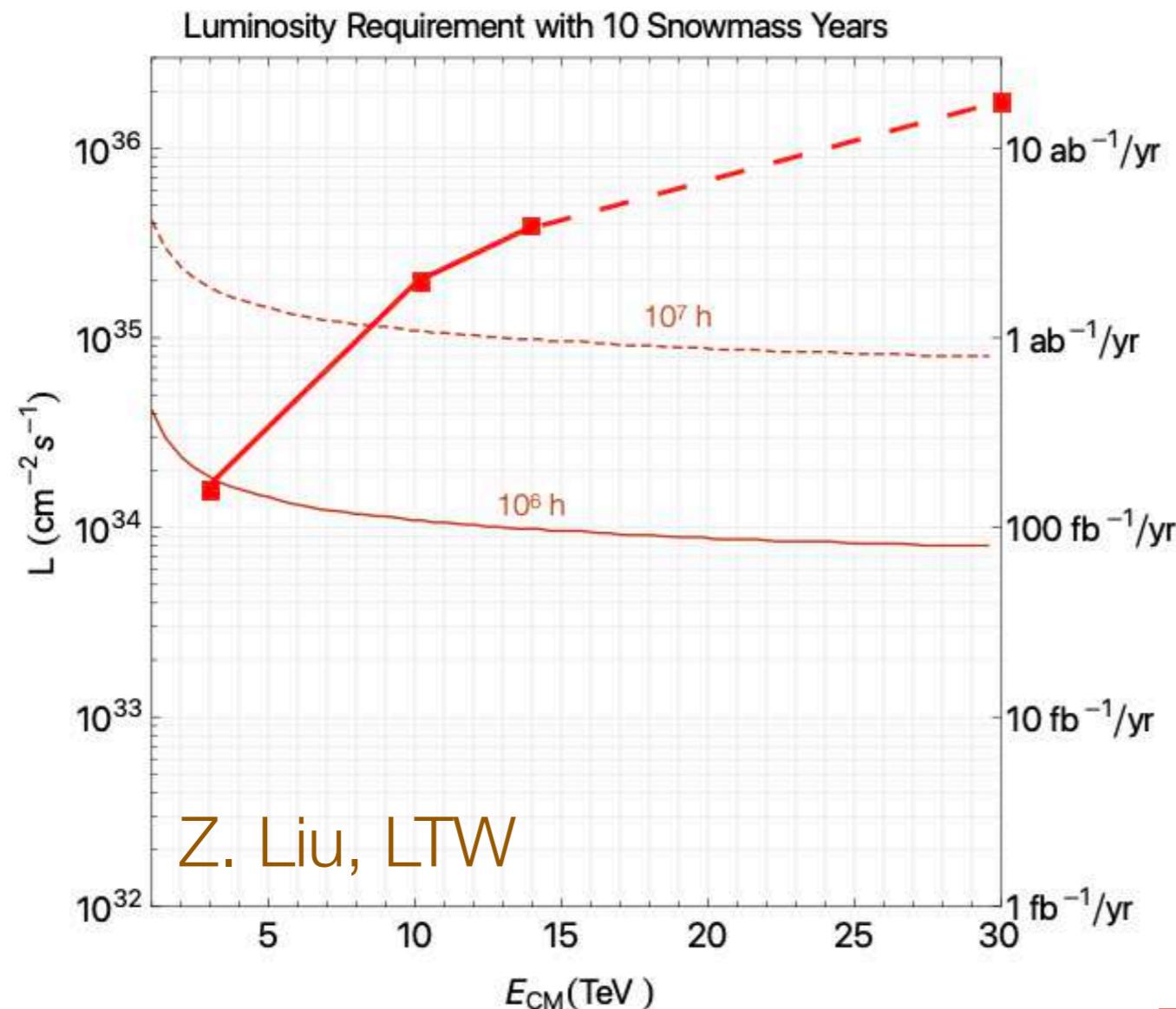
Assuming:

$$\mathcal{L} = \left( \frac{\sqrt{s}}{10 \text{ TeV}} \right)^2 \times 10^{35} \text{ cm}^{-2} \text{s}^{-1}$$

$E_{\text{CM}}$  :    125 GeV    1.5 TeV    3 TeV    6 TeV    10 TeV    30 TeV

# of Higgs/ $10^7$ s:     $\sim 10^{4.5}$      $\sim 4 \times 10^4$      $\sim 2 \times 10^5$      $10^6$      $10^7$      $10^8$

# MC as Higgs factory



Assuming:

$$\mathcal{L} = \left( \frac{\sqrt{s}}{10 \text{ TeV}} \right)^2 \times 10^{35} \text{ cm}^{-2} \text{s}^{-1}$$

$E_{\text{CM}}$  :

125 GeV

1.5 TeV

3 TeV

6 TeV

10 TeV

30 TeV

# of Higgs/ $10^7$ s:

$\sim 10^{4-5}$

$\sim 4 \times 10^4$

$\sim 2 \times 10^5$

$10^6$

$10^7$

$10^8$

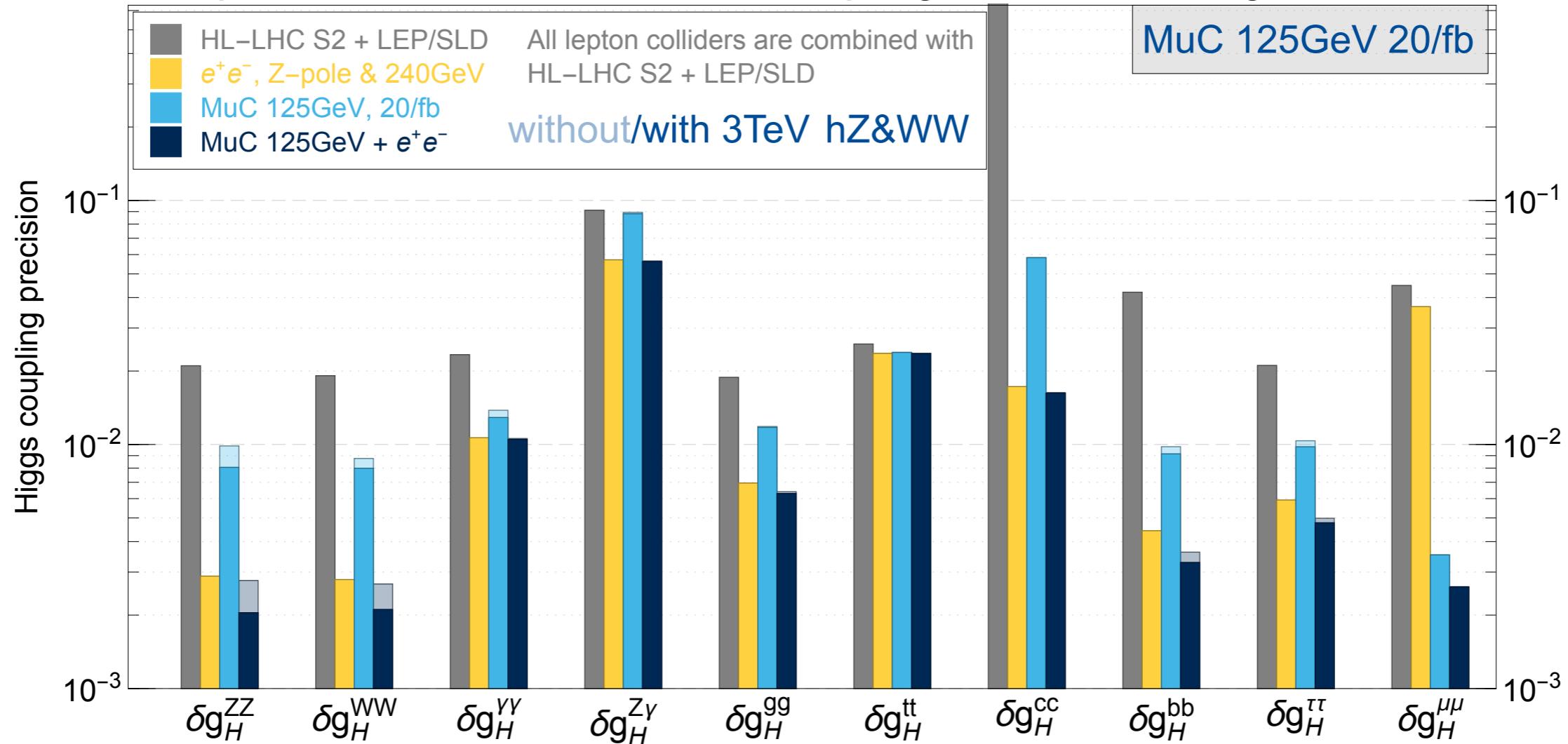


better than low E higgs factories

# Higgs coupling

J. de Blas, J. Gu, Z. Liu, preliminary

precision reach on effective couplings from full EFT global fit

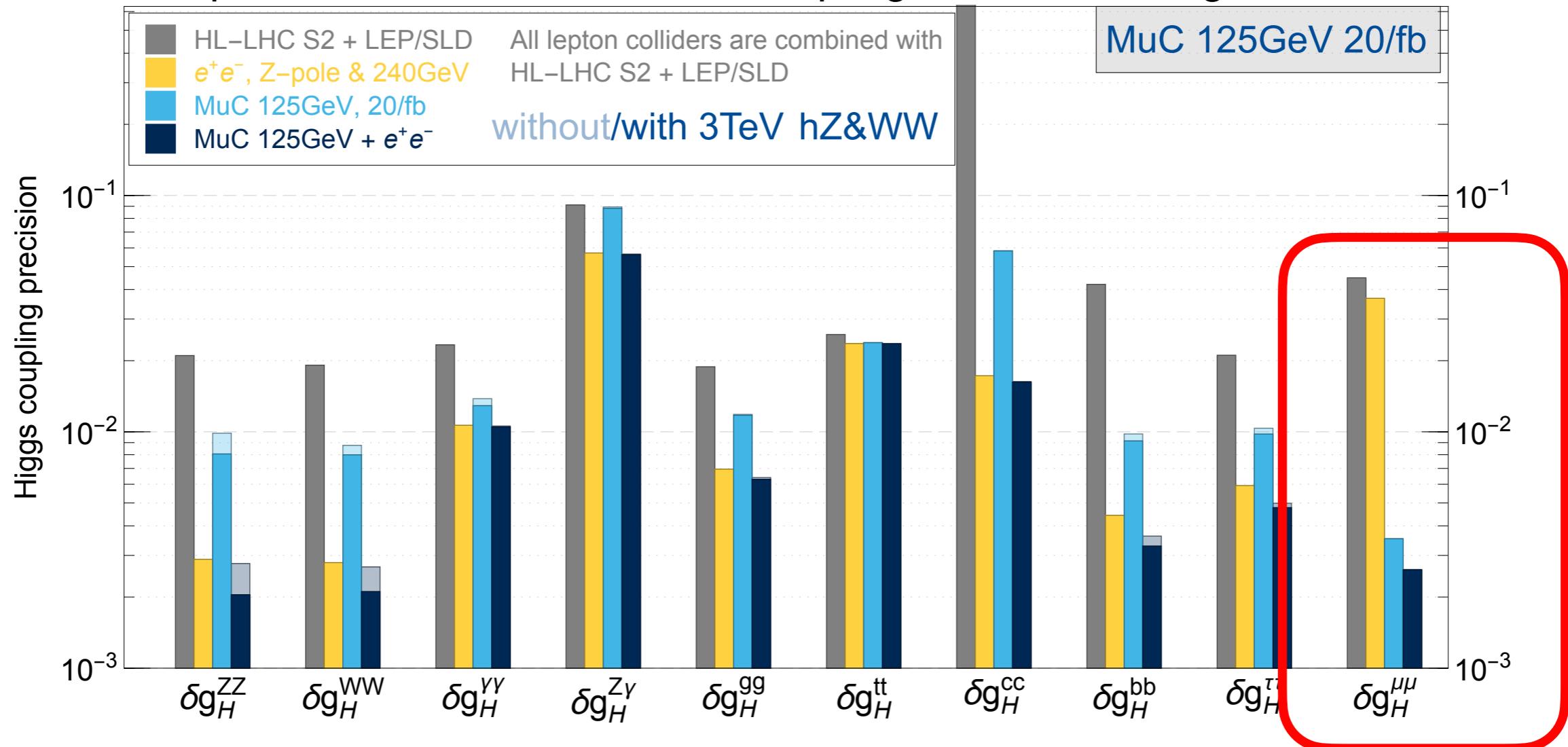


With only 125 GeV MC, less competitive.

# Higgs coupling

J. de Blas, J. Gu, Z. Liu, preliminary

precision reach on effective couplings from full EFT global fit



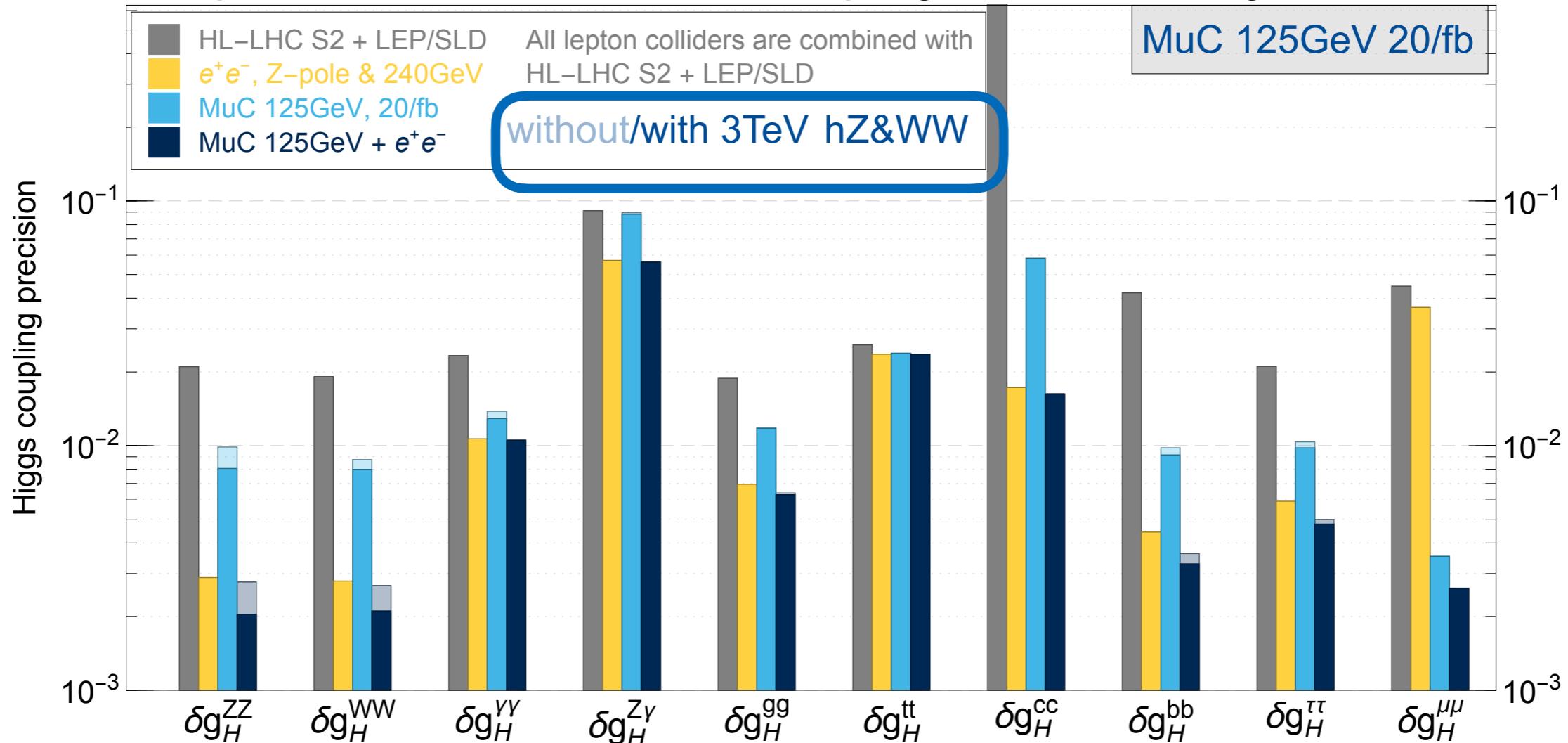
With only 125 GeV MC, not too competitive.

Except this!

# Higgs coupling

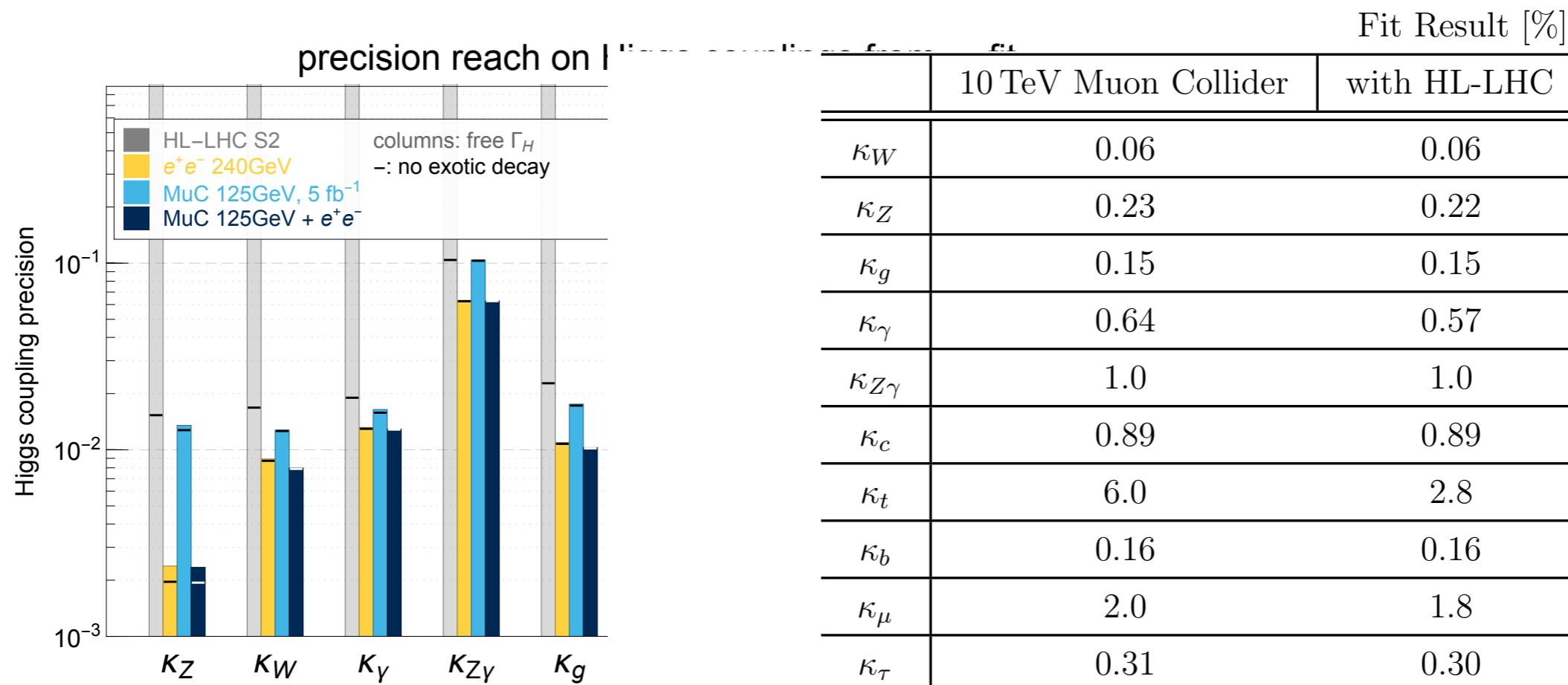
J. de Blas, J. Gu, Z. Liu, preliminary

precision reach on effective couplings from full EFT global fit



Higher energy helps.

# Higher energy

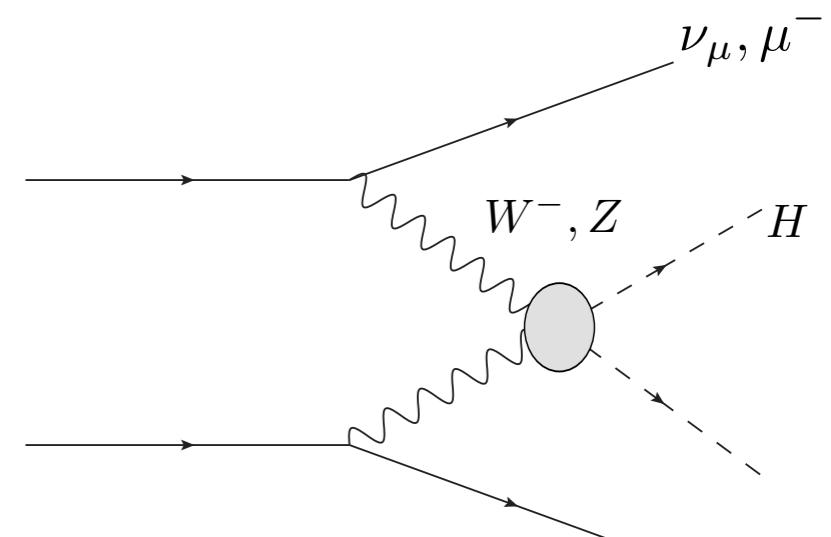
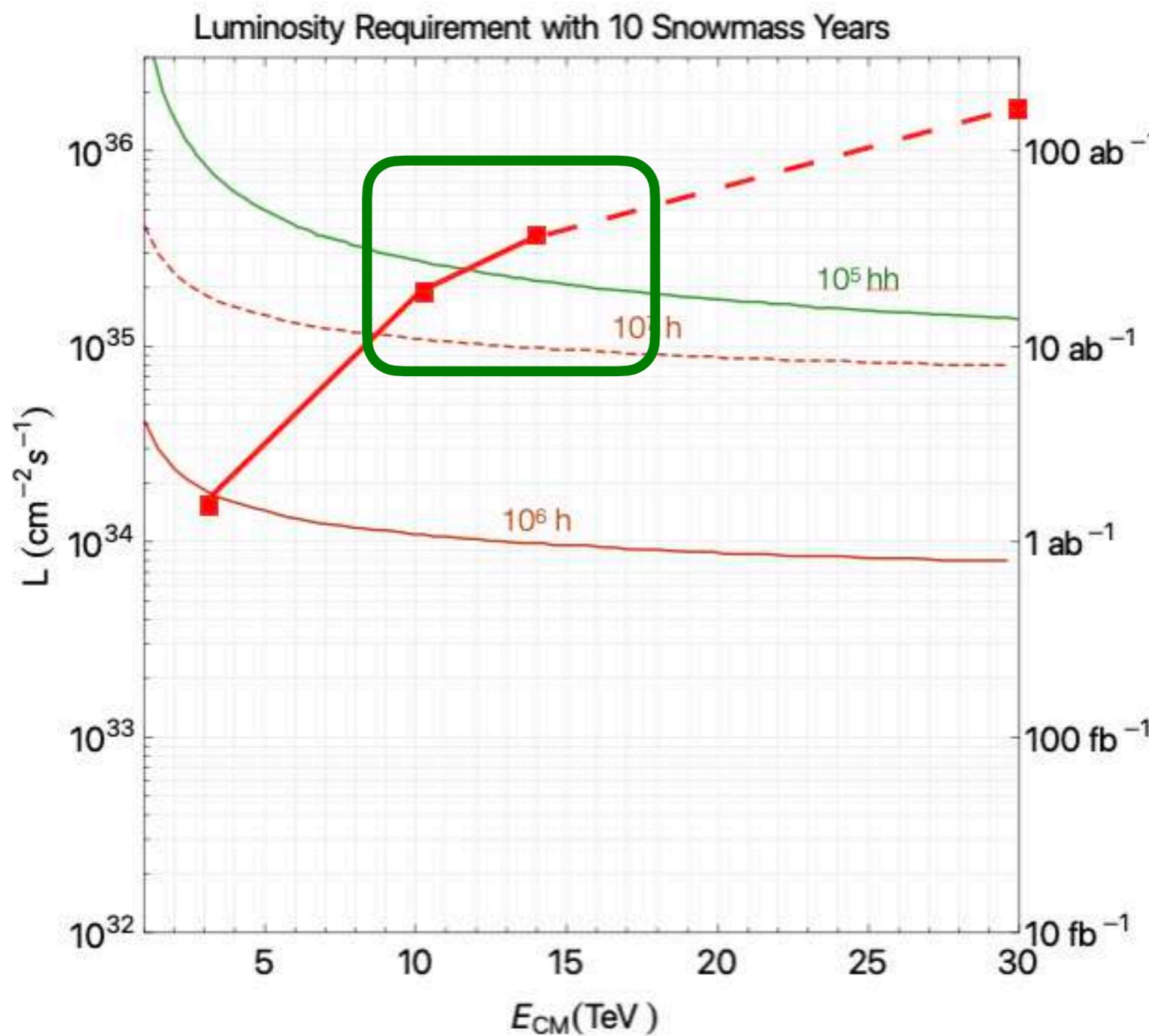


J. de Blas, J. Gu, Z. Liu, preliminary

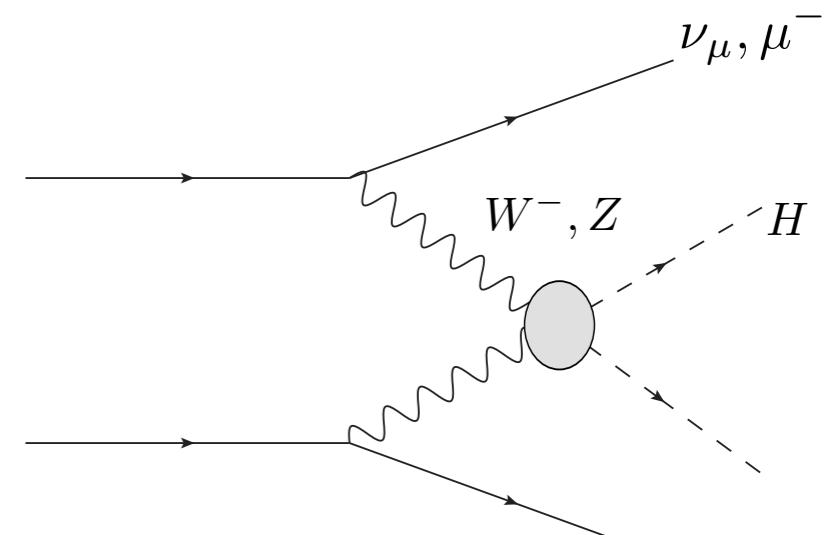
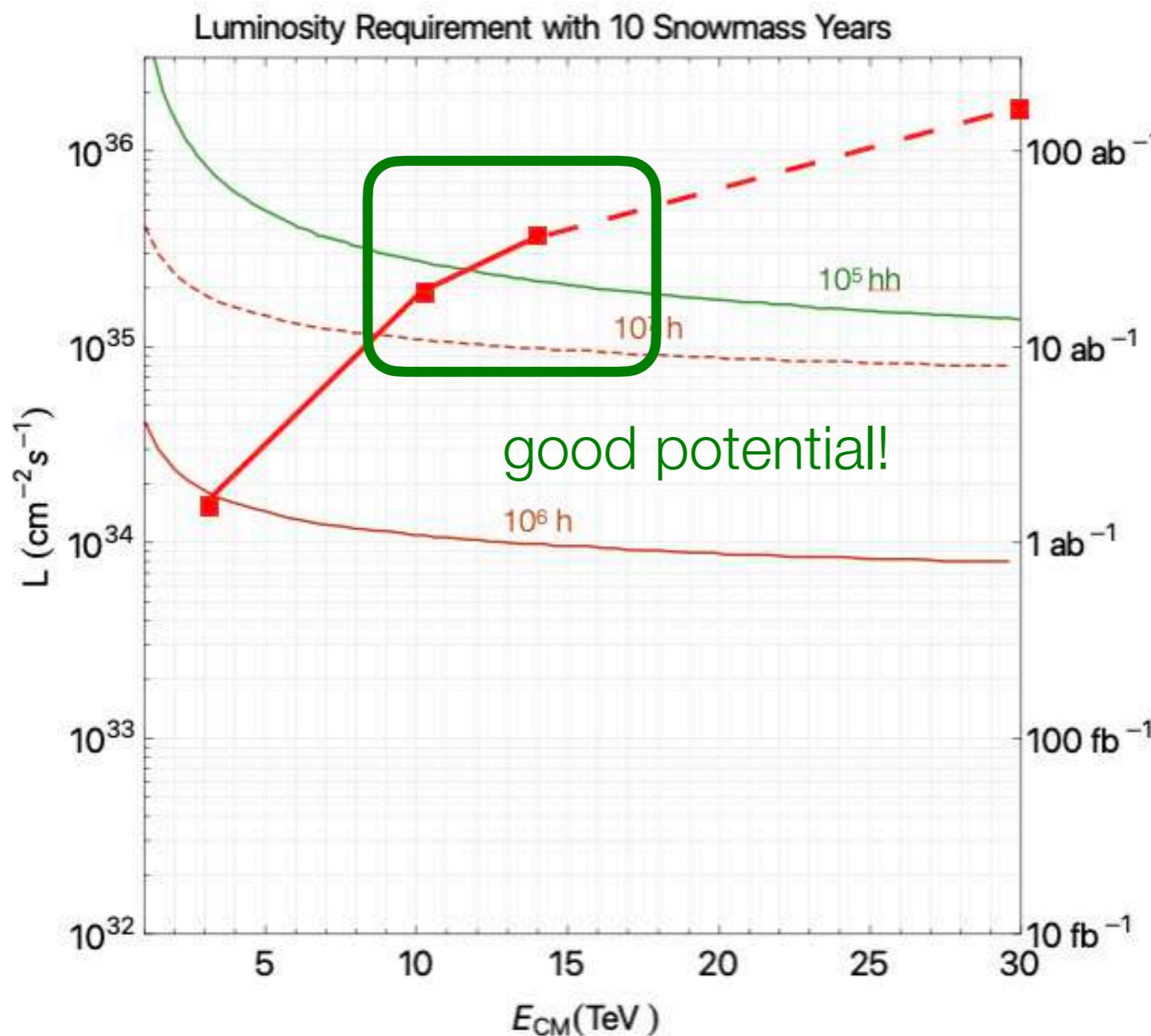
Estimates in “muon smasher’s guide”

0.1% level or better measurement possible at higher energies

# Double Higgs



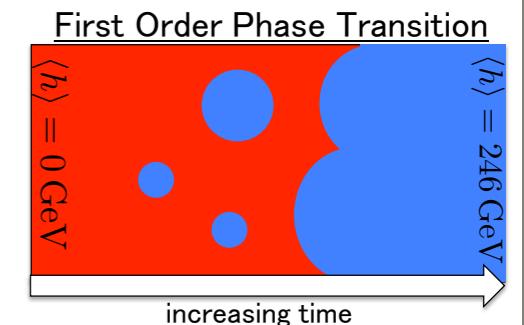
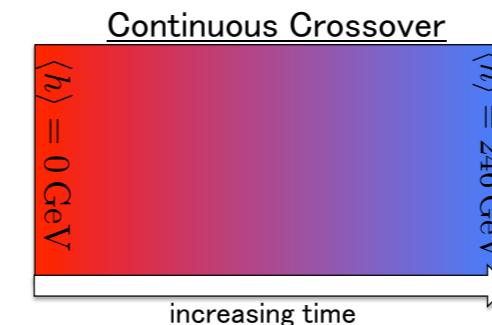
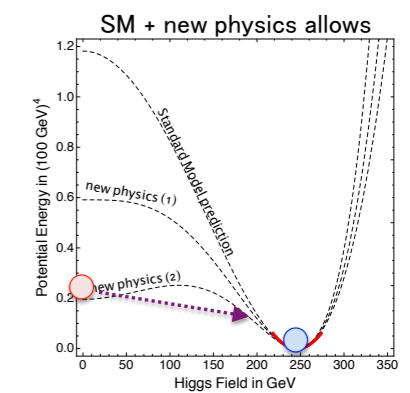
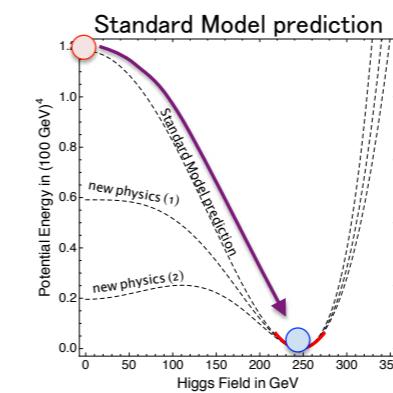
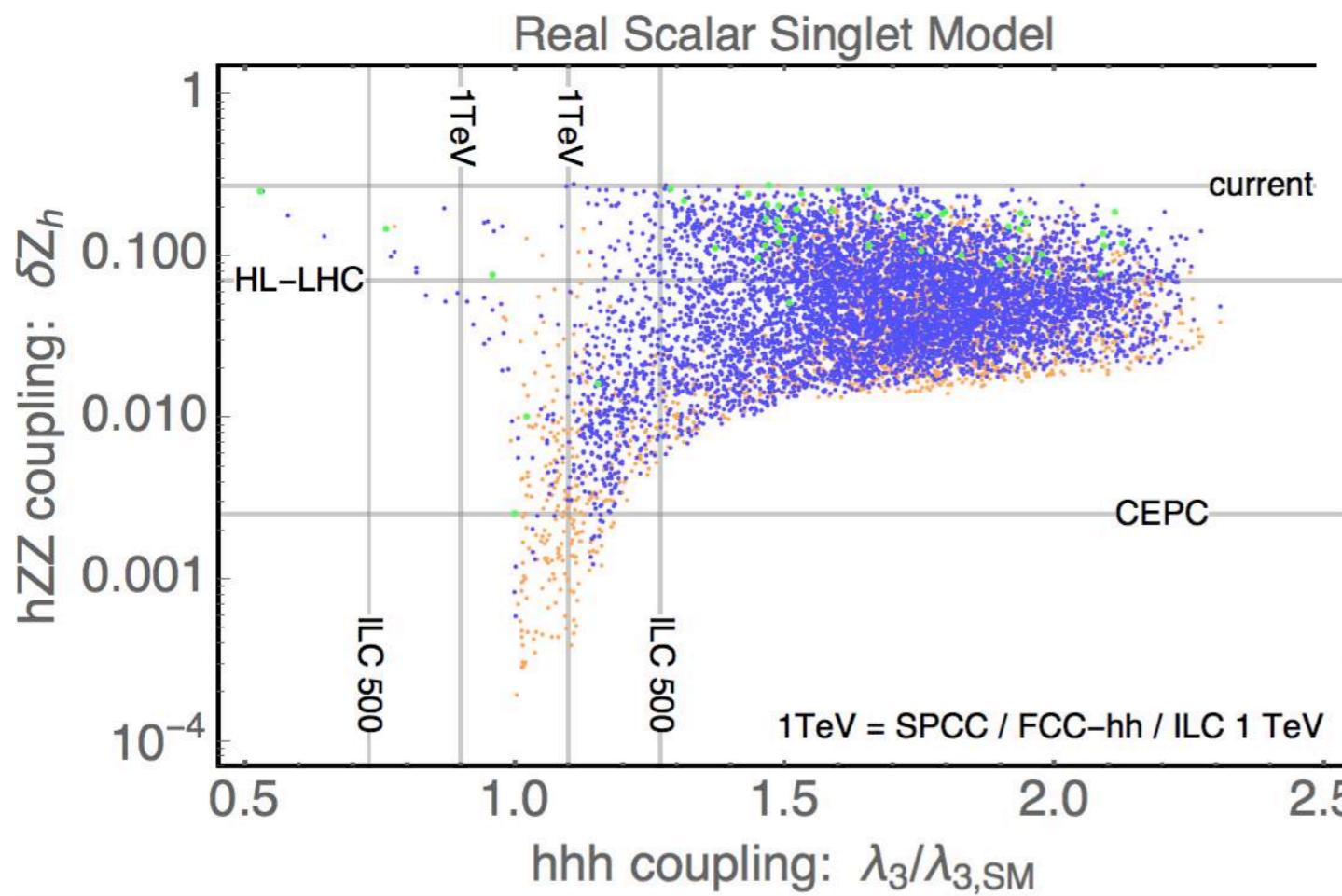
# Double Higgs



New physics effects also larger at higher energies.

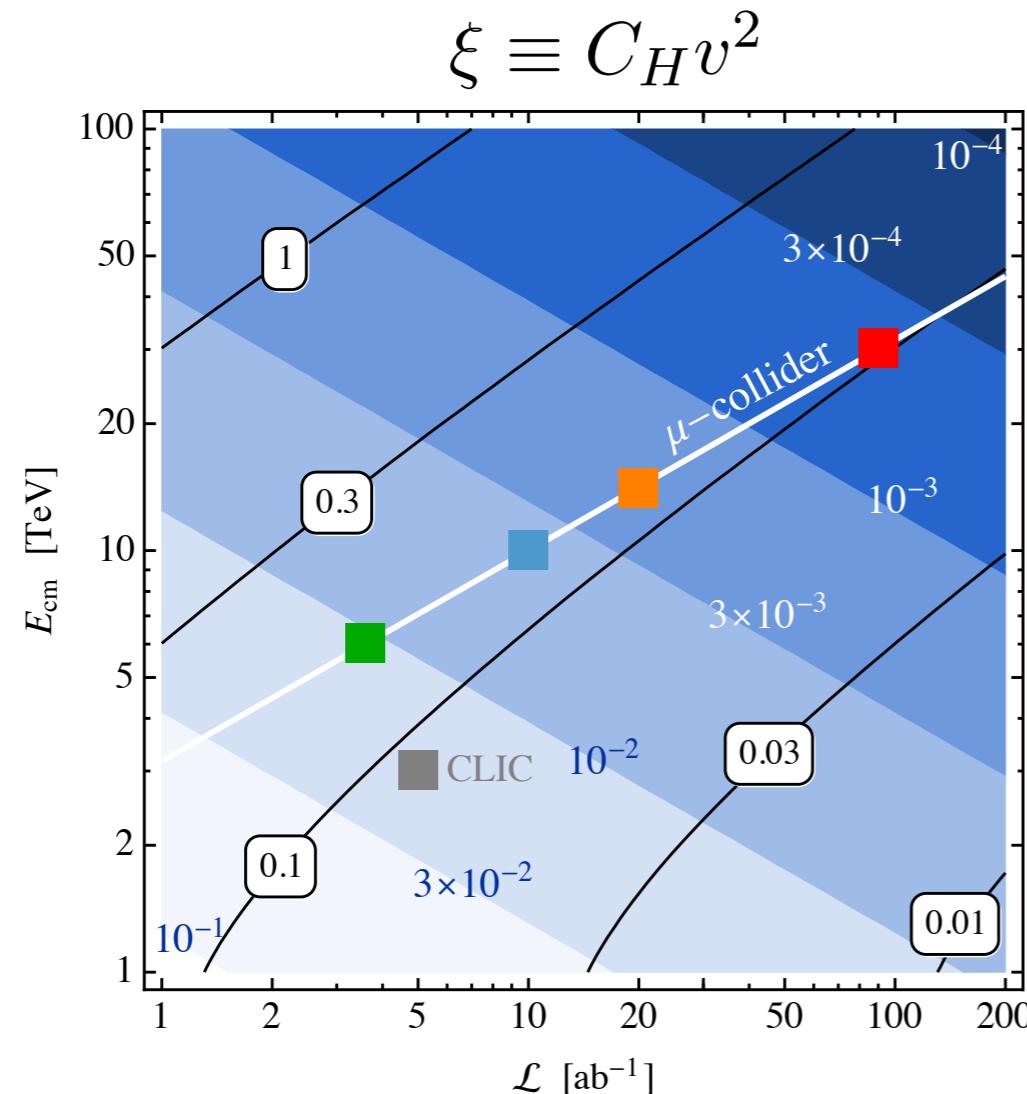
# Nature of EW phase transition

Which one is the right picture?



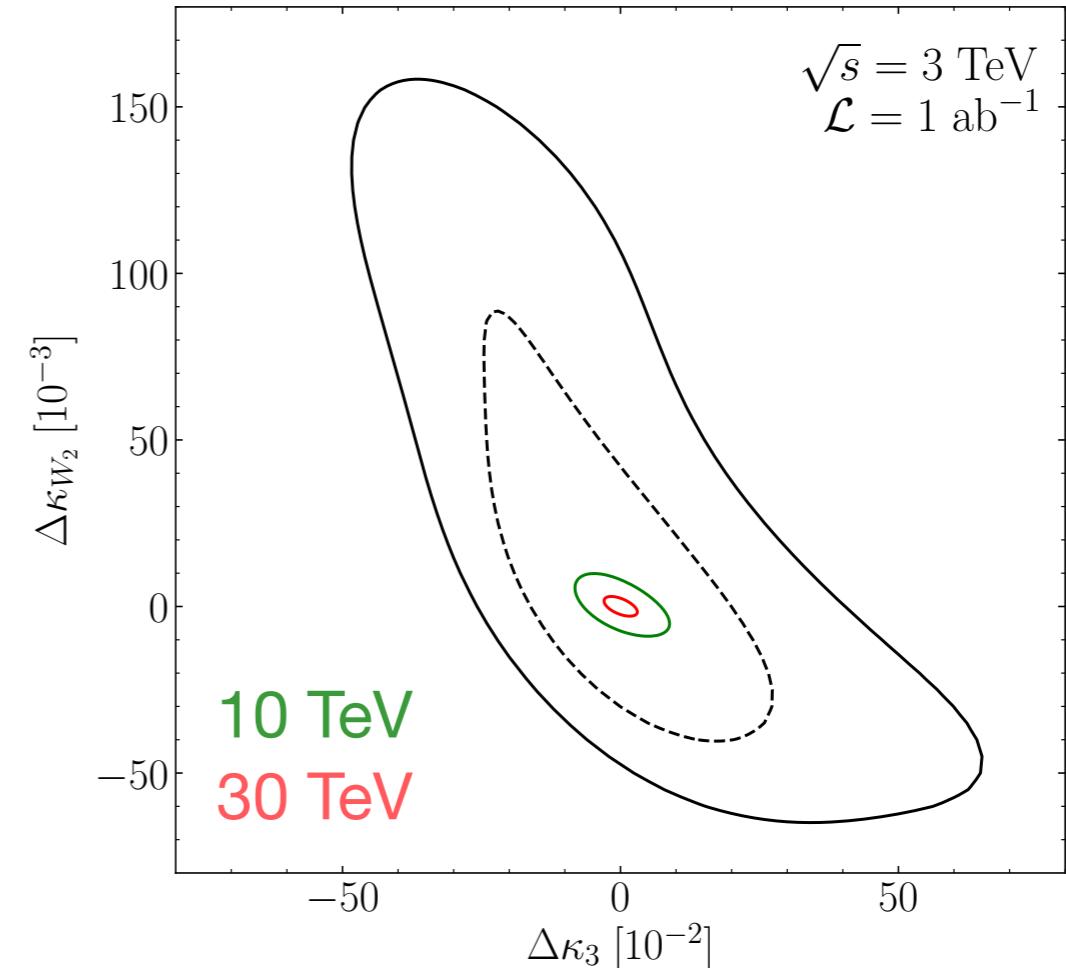
Precision Higgs measurements, self coupling and beyond, can reveal a lot.

# Self coupling modification



Buttazzo, Franceschini, Wulzer, 2012.11555

$$\mathcal{O}_H = \frac{1}{2} \left( \partial_\mu (H^\dagger H) \right)^2$$



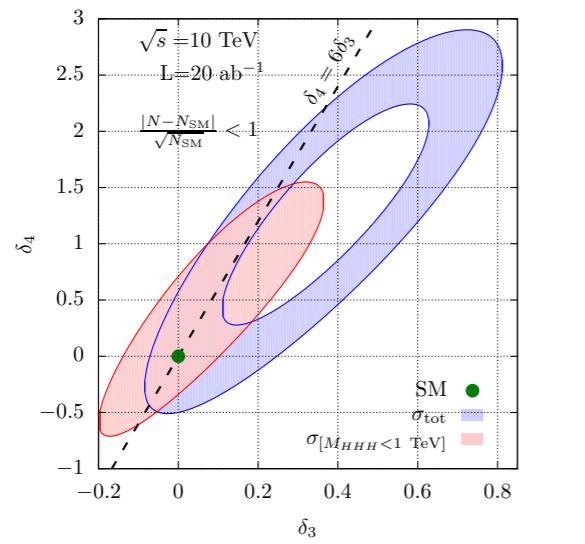
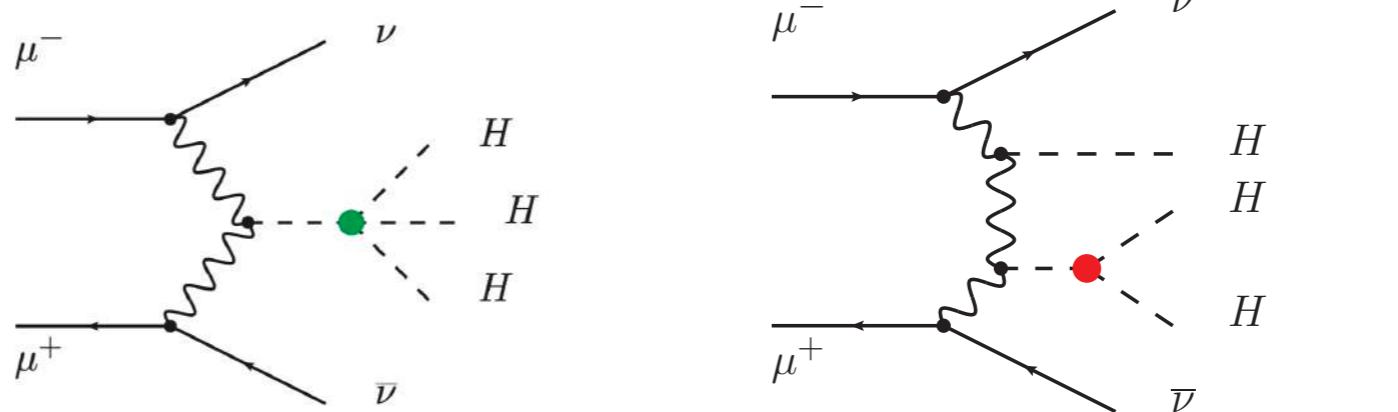
Han, Liu, Low, Wang, 2008.12204

$$\mathcal{L} \supset -\frac{m_H^2}{2v} \left( \kappa_3 H^3 + \frac{1}{4v} \kappa_4 H^4 \right)$$

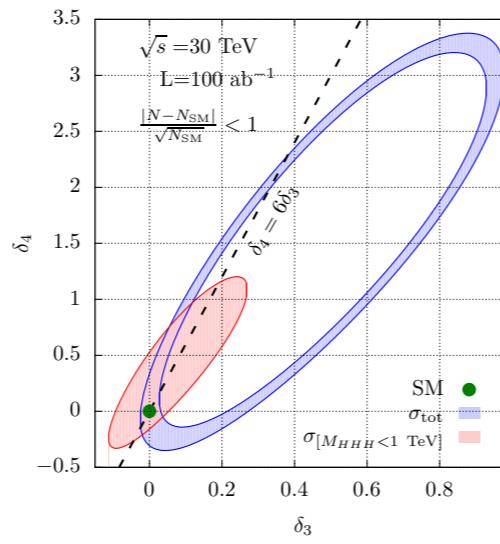
Precision better than CLIC, FCC-hh for  $E_{\text{CM}} = 10 \text{ TeV}$

# Higgs quartic coupling

3 Higgs final state



10 TeV  $\delta_4 \sim [-0.4, 0.7]$



30 TeV  $\delta_4 \sim [-0.2, 0.5]$

$$\frac{S}{\sqrt{B}} = \frac{|\mathcal{L} \cdot (\sigma - \sigma_{SM})|}{\sqrt{\mathcal{L} \cdot \sigma_{SM}}} \leq 1$$

Other colliders:

ILC  $\sim [-10, 10]$   
CLIC  $\sim [-5, 5]$   
FCC  $\sim [-2, 4]$

A 10 TeV muon collider would do better than FCC

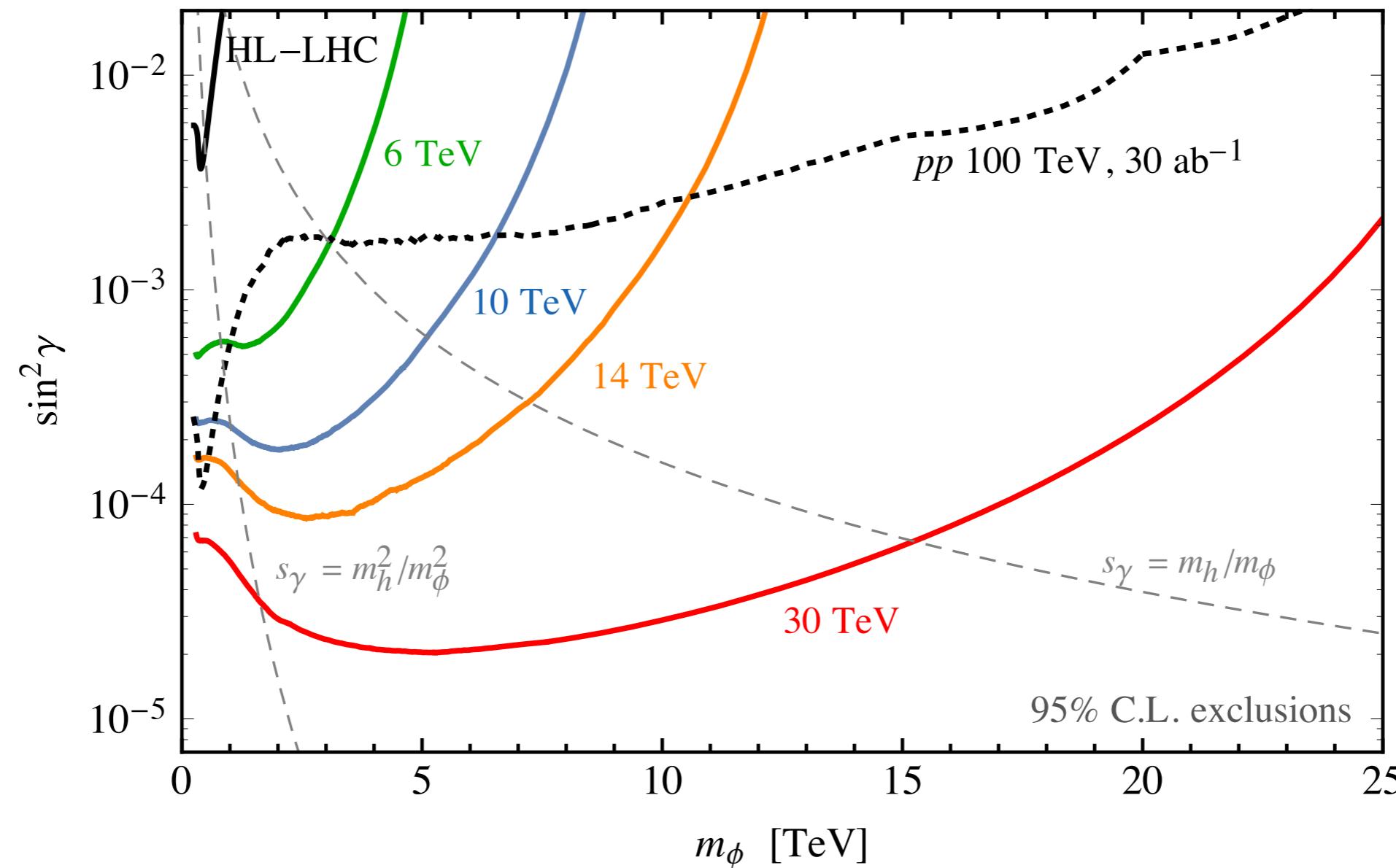
Slide of E. Vryonidou

$$\lambda_3 = \lambda_{SM}(1 + \delta_3) = \kappa_3 \lambda_{SM}$$

$$\lambda_4 = \lambda_{SM}(1 + \delta_4) = \kappa_4 \lambda_{SM}$$

# Higgs's friends

D. Buttazzo, D. Redigolo, F. Sala, A. Tesi, 1807.04743



Singlet scalar mixing with the Higgs.

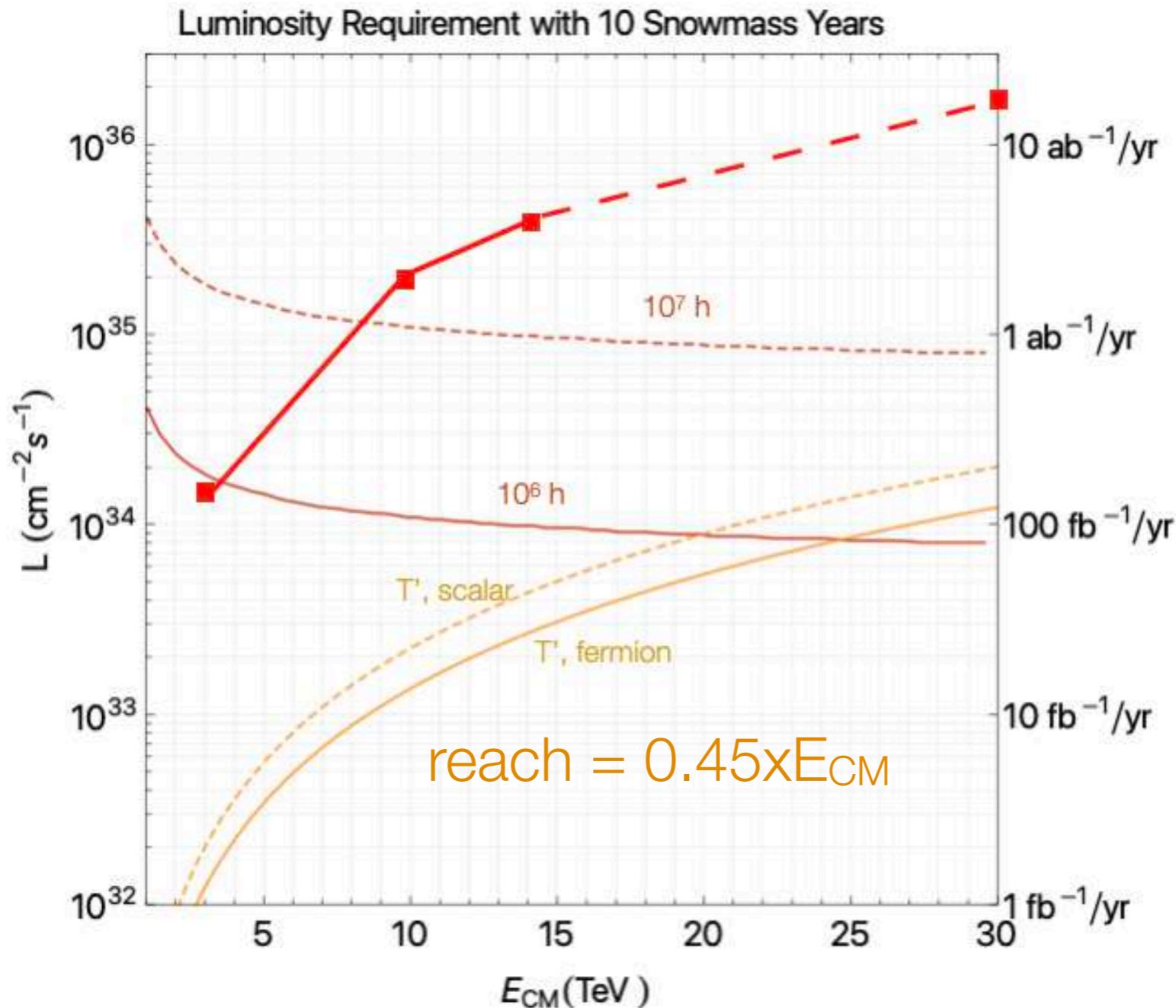
# Origin of the weak scale

— — — — .      New physics at  
TeV<sub>s</sub>-10 TeV ?

—       $\Lambda_{EW}$

The usual naturalness, not discovered at the LHC, just a little delayed?

# Top partner



Pair production

$$\mu^+\mu^- \rightarrow T'\tilde{T}'$$

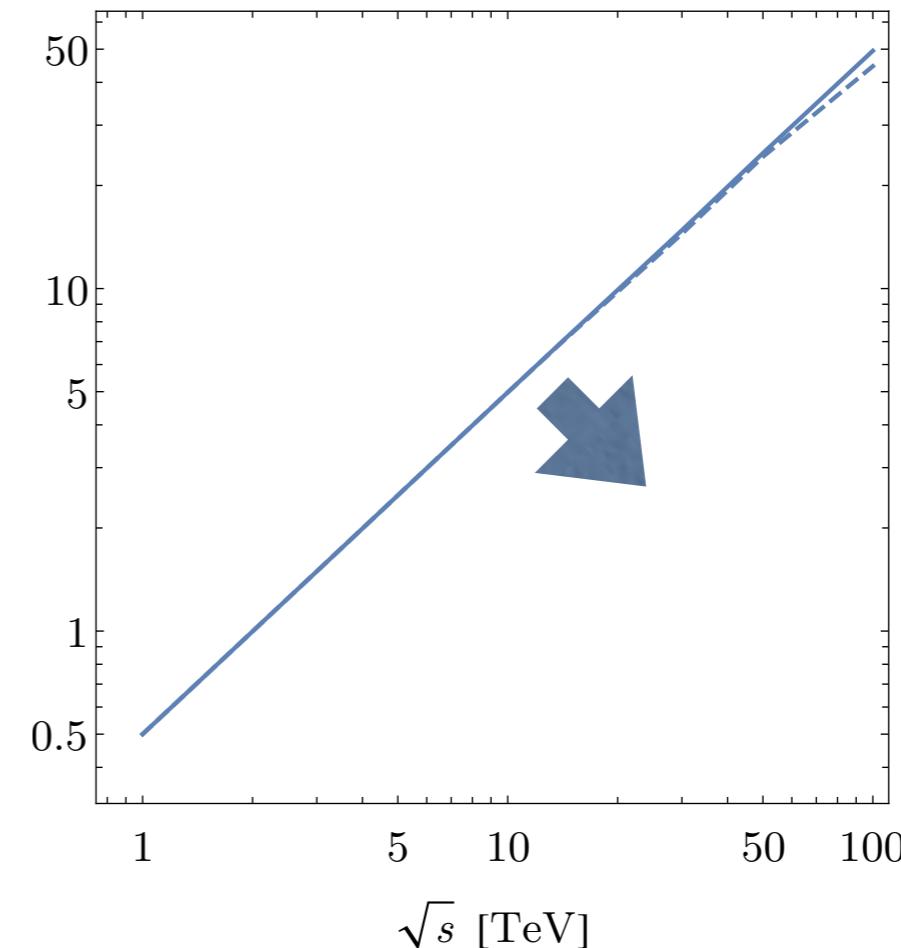
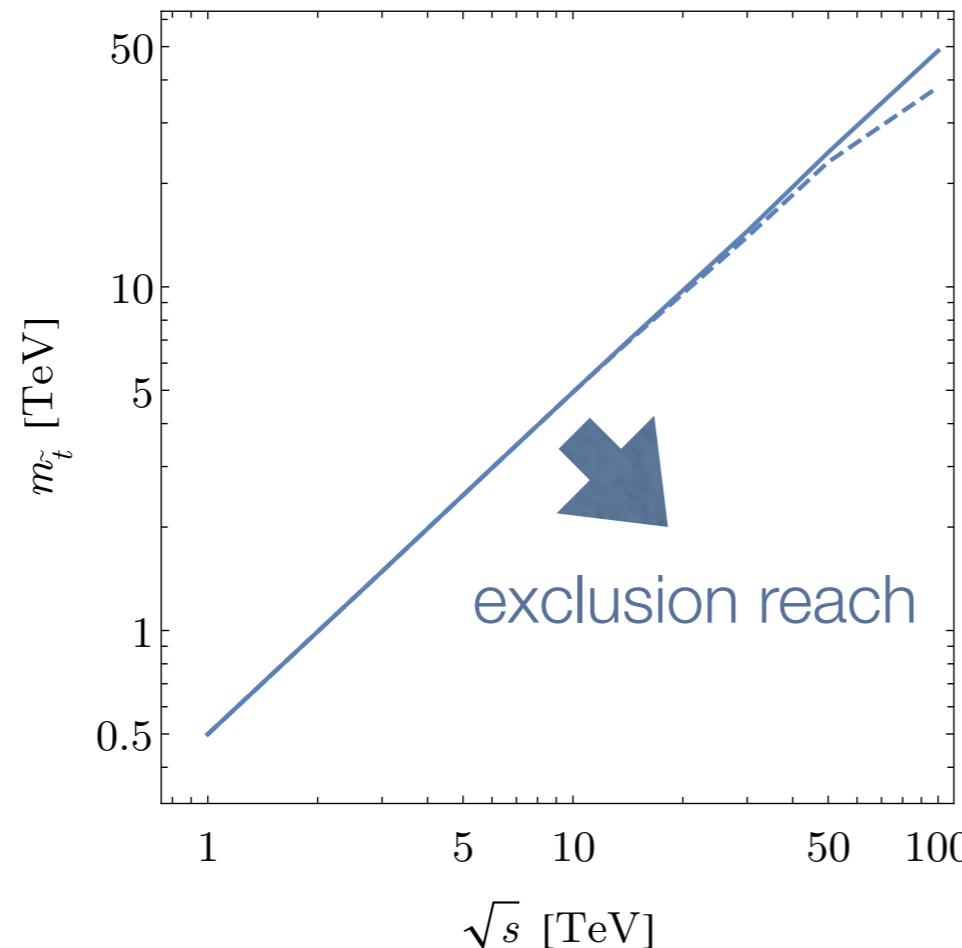
Spectacular signal, 10s event needed for discovery

# SUSY top squark

Estimates in “muon smasher’s guide”

$$\mu^+ \mu^- \rightarrow \tilde{t}_R \tilde{t}_R \rightarrow t\bar{t} + \chi\chi$$

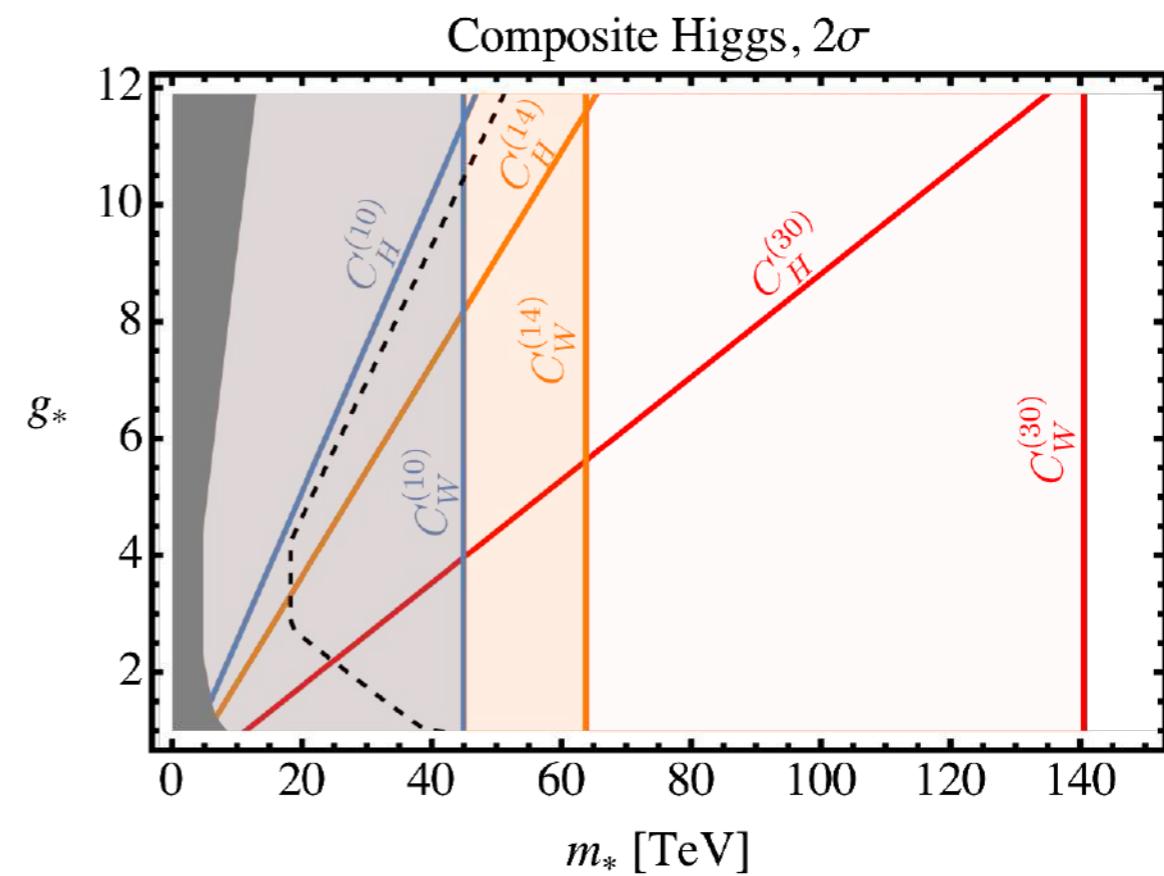
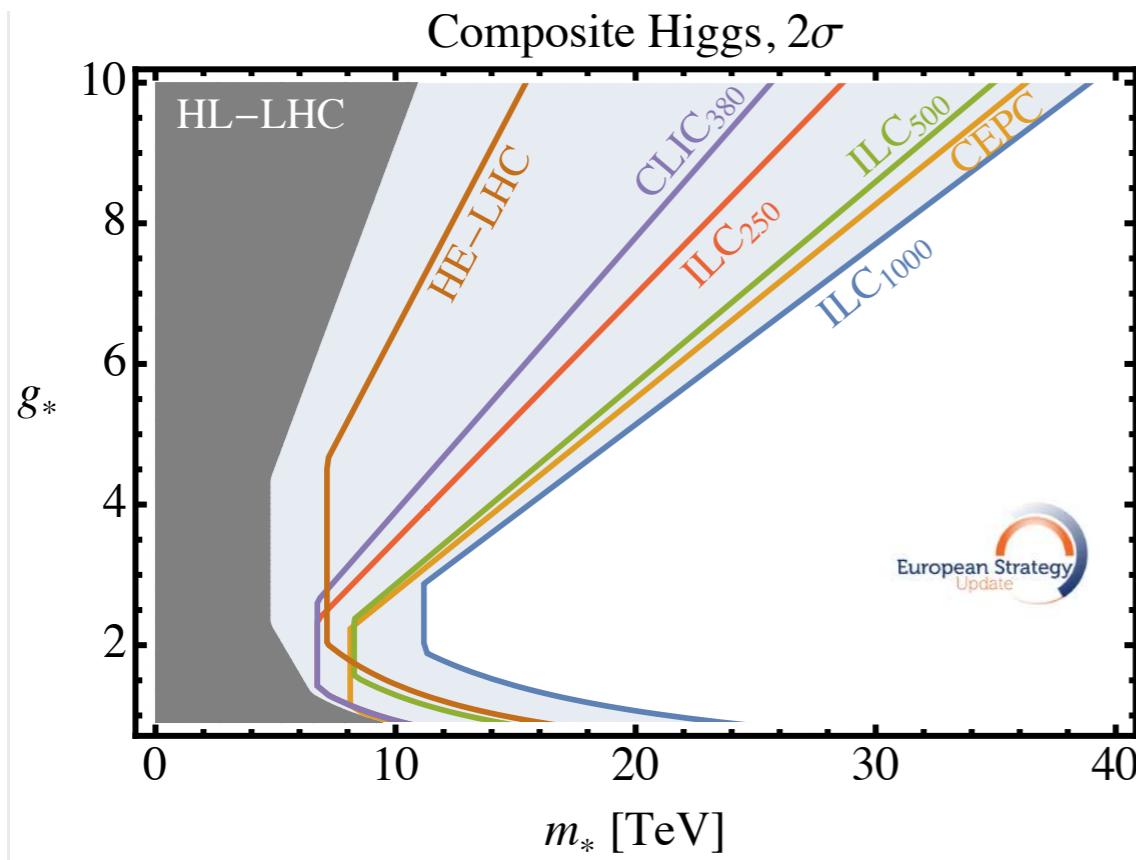
$$\mu^+ \mu^- \rightarrow \tilde{t}_L \tilde{t}_L \rightarrow t\bar{t} + \chi\chi$$



Reach quite close to  $0.5 \times E_{CM}$

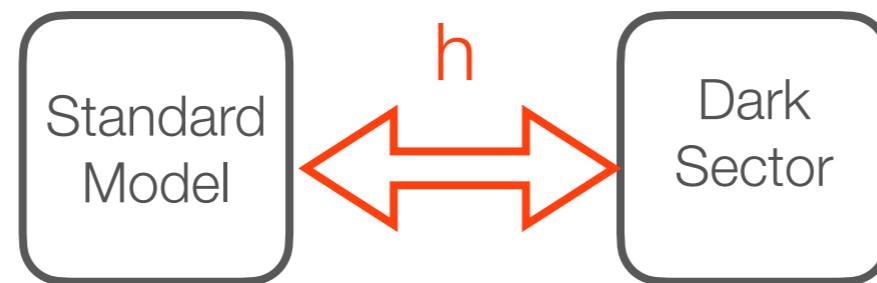
Spectacular signal, 10s event needed for discovery

# Composite



Through precision measurement at high energies.

# Higgs portal



- \* Dark sector coupling to the SM

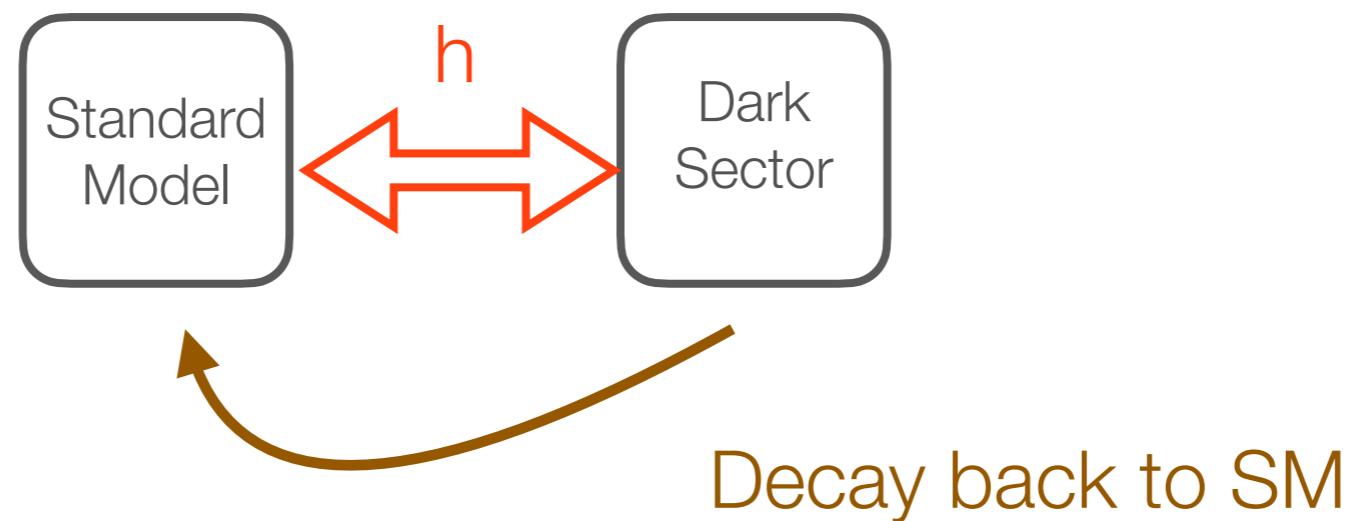
$$O_{\text{SM}} \cdot O_{\text{dark}}$$

$O_{\text{SM}}$ : gauge inv. SM operator

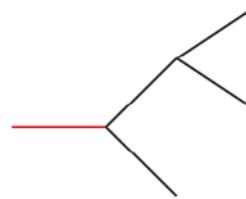
$O_{\text{dark}}$ : dark sector operator

- \* More relevant coupling  $\Leftrightarrow$  lowest dim operator
  - \* Unique choice:  $O_{\text{SM}} = H H^\dagger$ . Higgs portal.

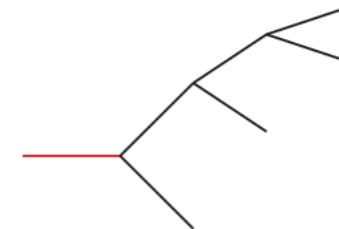
# Higgs rare decay



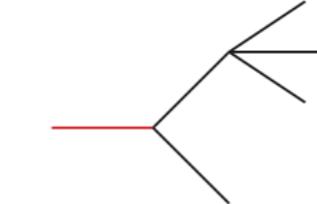
$h \rightarrow 2 \rightarrow 3$



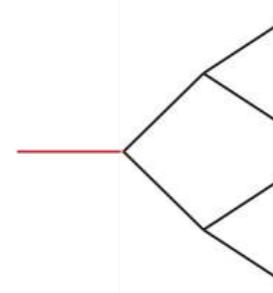
$h \rightarrow 2 \rightarrow 3 \rightarrow 4$



$h \rightarrow 2 \rightarrow (1+3)$

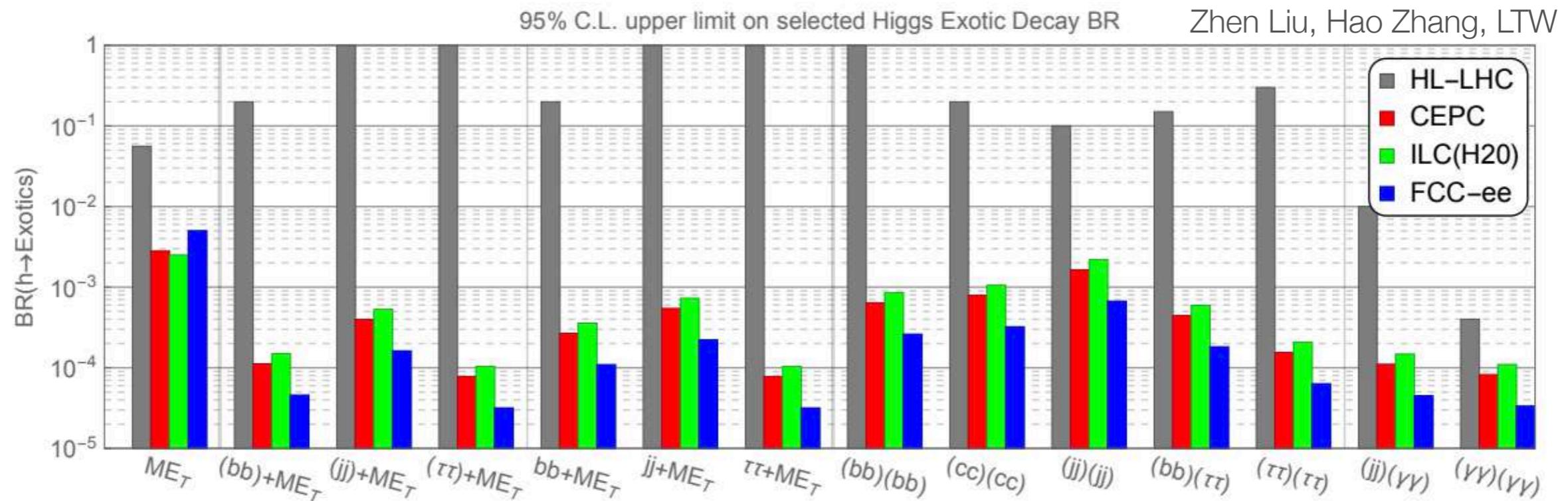


$h \rightarrow 2 \rightarrow 4$



....

# Higgs exotic decay



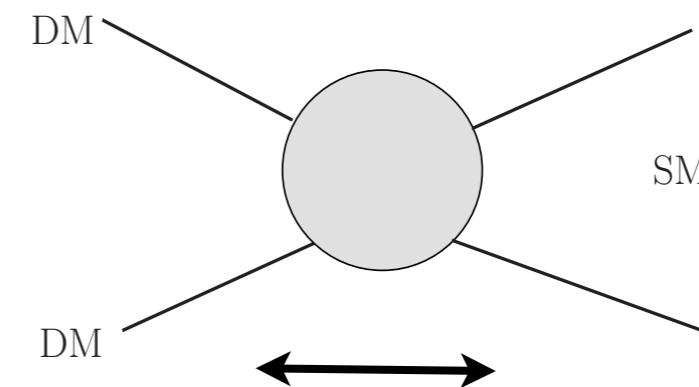
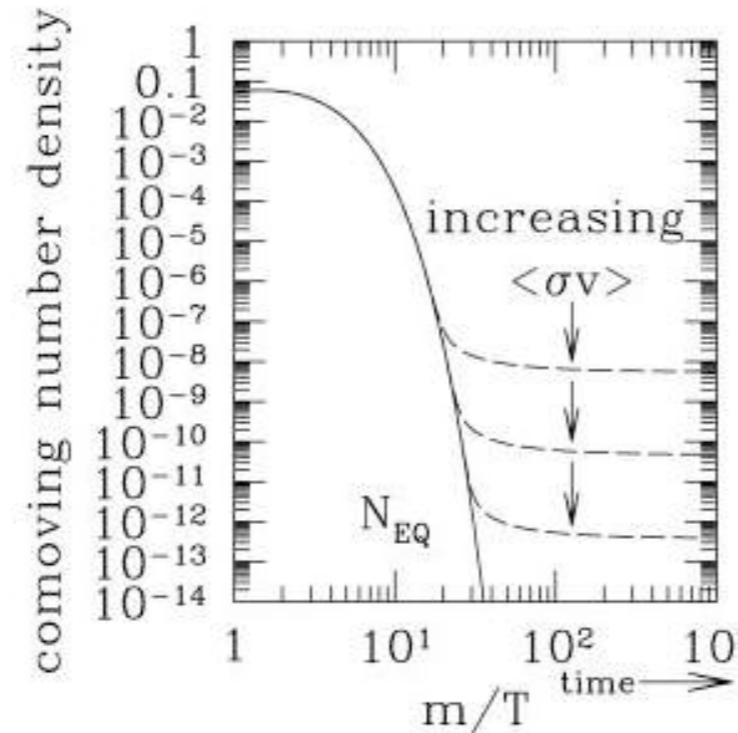
Complementary to hadron collider searches

Studies done for low E Higgs factory with  $10^6$  Higgses.  
High energy muon colliders, with  $10^7+$  Higgses at 10+ TeV,  
will do even better.

# (Other) new physics at high scale

- \* Dark matter (WIMP)
- \*  $Z'$
- \* Flavor/CP
- \* Lepto-quark
- \* ...

# WIMP:

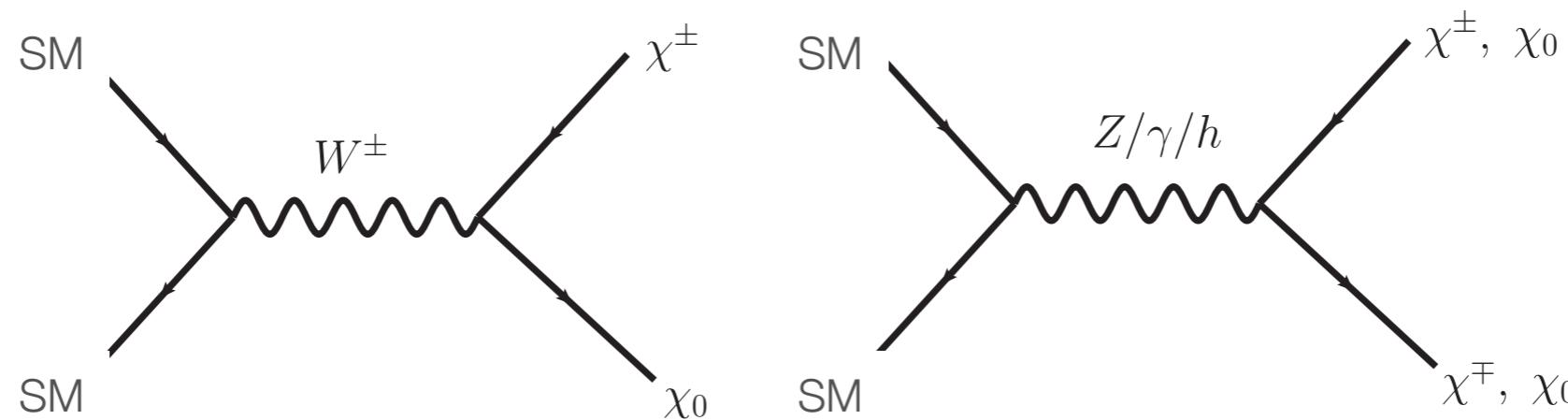


$$\langle \sigma v \rangle \sim \frac{g_D^4}{m_{DM}^2}$$

- \* Simple assumption: DM in thermal eq. with the SM in early universe

# Simplest model: part of an EW multiplet

“Minimal dark matter”, Cirelli, Fornengo and Strumia, hep-ph/0512090, 0903.3381



- \* Simplicity: there is no additional new mediator.
- \* Mediated by  $W/Z/h$ . Very predictive.
- \* In SUSY, there are two such examples
  - \* Higgsino: doublet. Wino: triplet.

# DM part of a EW multiplet

$\text{DM} \in (1, n, Y)$  of  $SU(3)_C \times SU(2)_L \times U(1)_Y$

- \* Consider first the fermionic mulitplets.
  - \* Only couplings at the renormalizable level are the gauge interactions.
  - \* The only free parameter at this level is the mass,  $m_\chi$ .
  - \* Very predictive.

# DM part of a EW multiplet

$\text{DM} \in (1, n, Y) \text{ of } \text{SU}(3)_C \times \text{SU}(2)_L \times \text{U}(1)_Y$

- \* **n odd. Fermionic.**
  - \*  $n > 7$ , Landau pole close to  $M_{\text{DM}}$ .
  - \* After EWSB, mass splitting (minimally) generated at 1-loop.
  - \* Choose  $Y=0$ . Lightest member electric neutral. Potential DM candidate.

# DM part of a EW multiplet

$\text{DM} \in (1, n, Y) \text{ of } \text{SU}(3)_C \times \text{SU}(2)_L \times \text{U}(1)_Y$

- \* **n even. Fermionic**
  - \* Choose  $Y=(n-1)/2$  ensures lightest member is neutral.
  - \* Direct detection rules out the minimal case due to tree level Z exchange.
    - \* Can be avoided to introduce a small splitting,  $\delta m > 10^2 \text{ keV}$ , of the neutral states (for example, from a dim-5 operator). Not quite minimal (additional model dependence).
  - \* Famous example: Higgsino  $(1,2)_{1/2}$

# DM part of a EW multiplet

$\text{DM} \in (1, n, Y)$  of  $SU(3)_C \times SU(2)_L \times U(1)_Y$

- \* **Scalar (real and complex)**
- \* Minimal case: mass splitting, stability discussion parallel to that of the fermionic multiplets.
- \* Addition couplings of the form  $H^\dagger H X^\dagger X$ . More parameters involved in a full analysis.
- \* **More focus on the fermion case (so far).**

# Thermal targets

Model (color, $n$ , $Y$ )		Therm. target
(1,2,1/2)	Dirac	1.1 TeV
(1,3,0)	Majorana	2.8 TeV
(1,3, $\epsilon$ )	Dirac	2.0 TeV
(1,5,0)	Majorana	14 TeV
(1,5, $\epsilon$ )	Dirac	6.6 TeV
(1,7,0)	Majorana	48.8 TeV
(1,7, $\epsilon$ )	Dirac	16 TeV

Correct relic abundance  
⇒ Thermal targets

Reach up to thermal target  
≈  
complete coverage for WIMP candidate

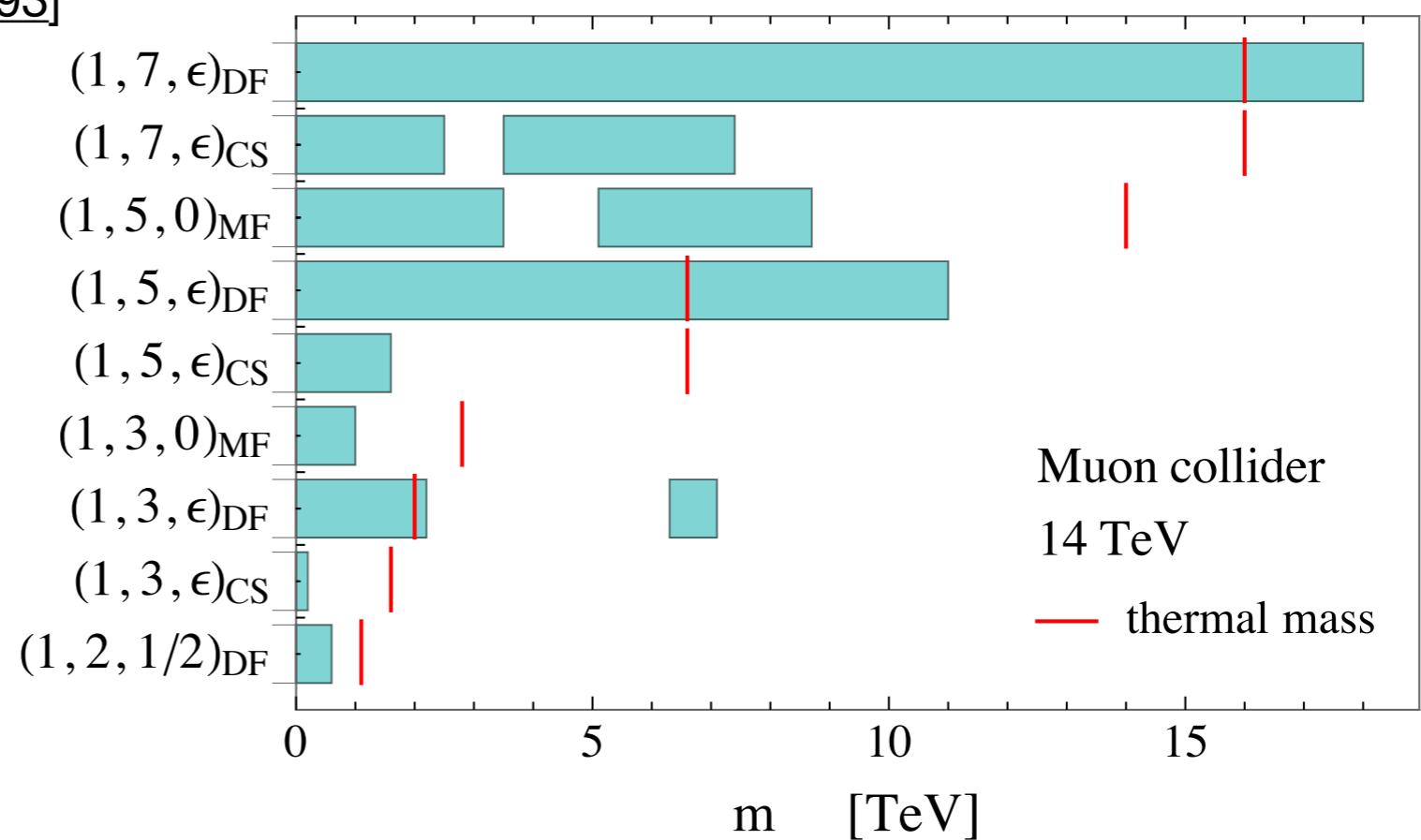
Mitridate, Redi, Smirnov, Strumia, 1702.01141

S. Bottaro, D. Buttazzo, M. Costa, R. Franceschini, P. Panci, D. Redigolo, L. Vittorio, 2107.09688

# “indirect”, from precision measurement

**indirectly** [1810.10993]

Di Luzio, Grober, Panico, 1810.10993

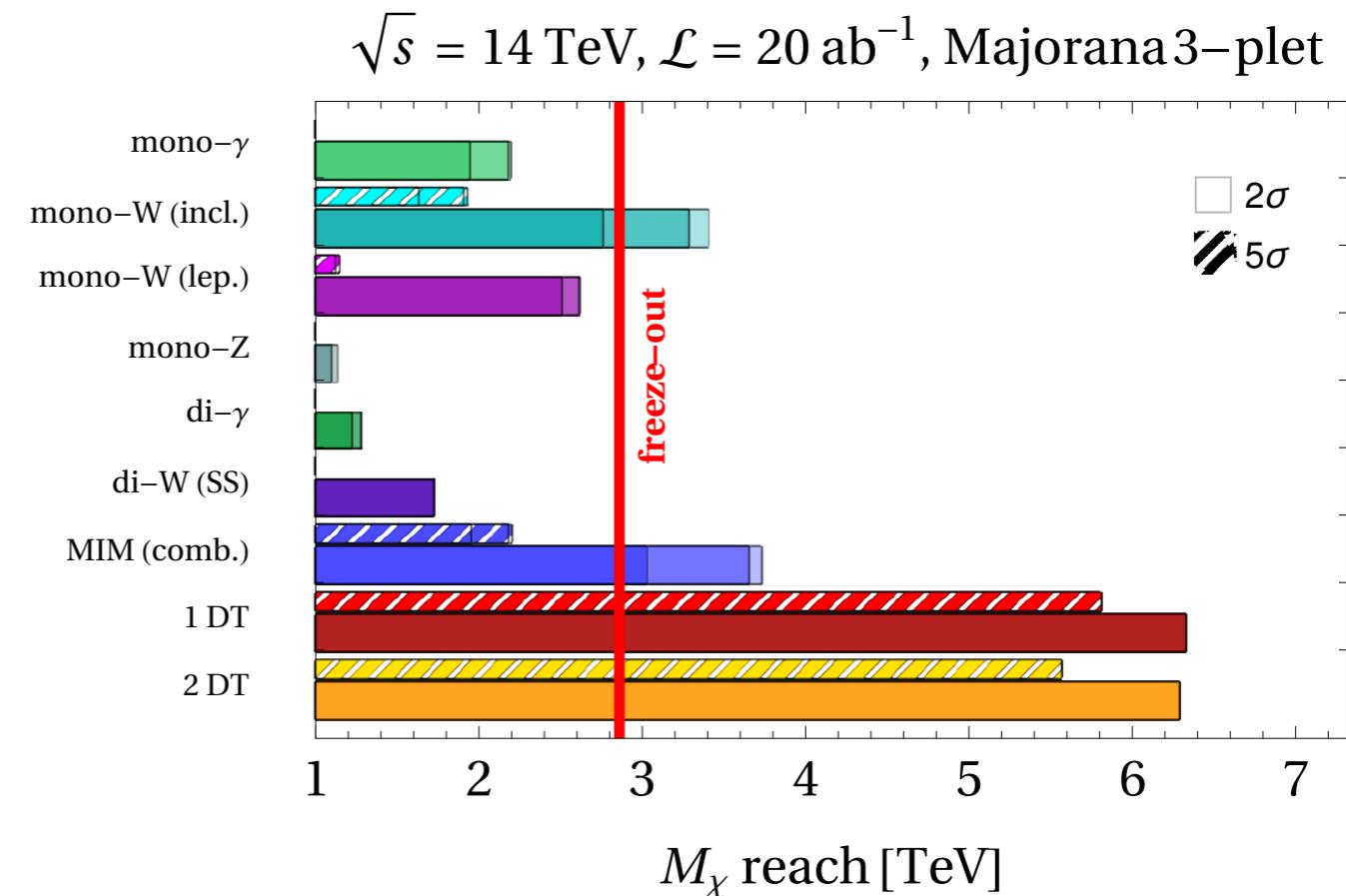
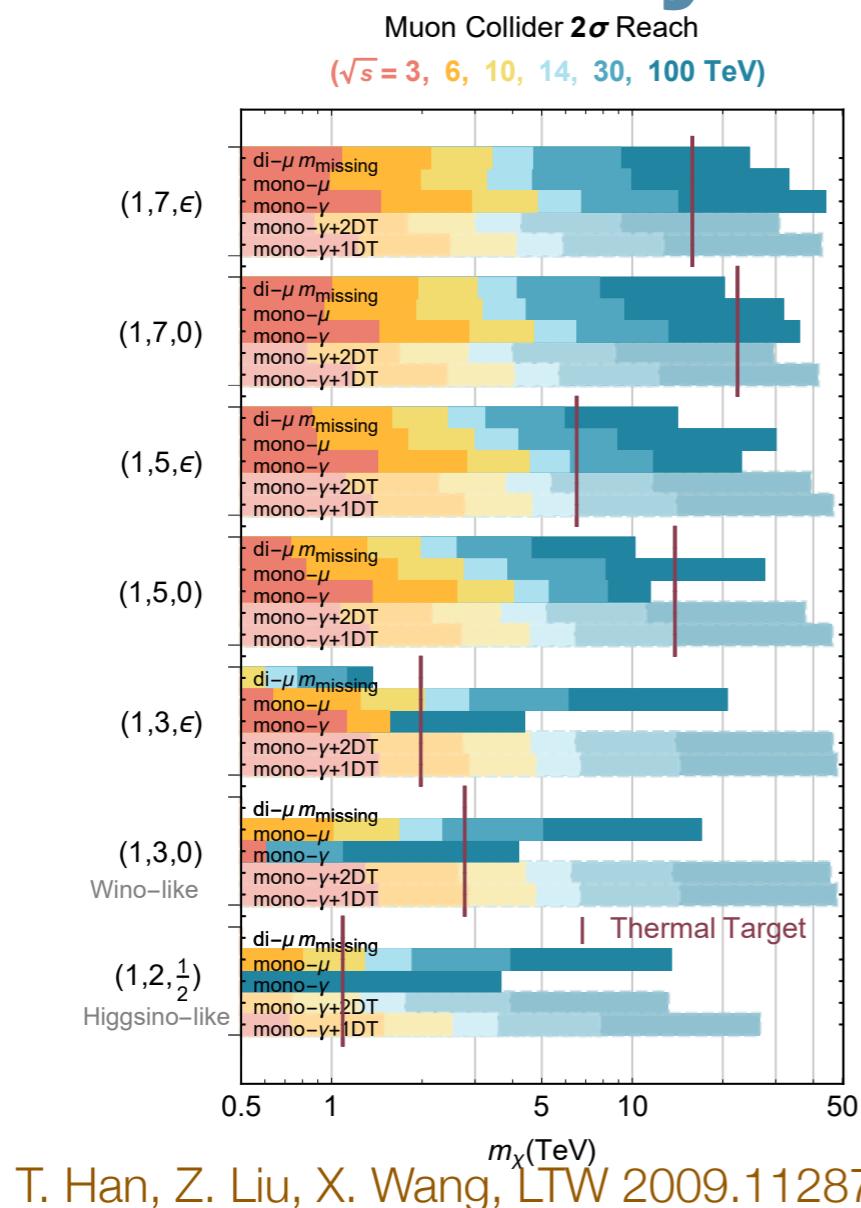


At loop level, modifying the  $q\bar{q}$  (or  $\ell^+\ell^-$ )  $\rightarrow f\bar{f}$  amplitude

# Two classes of “direct” DM signals at colliders

- \* Production of dark matter particle.
  - \* Inclusive search for X+MET
    - \* similar to mono-jet at hadron colliders.
- \* Small EW induced mass splitting, charged member long-lived.
  - \* Disappearing track

# Reach by channel

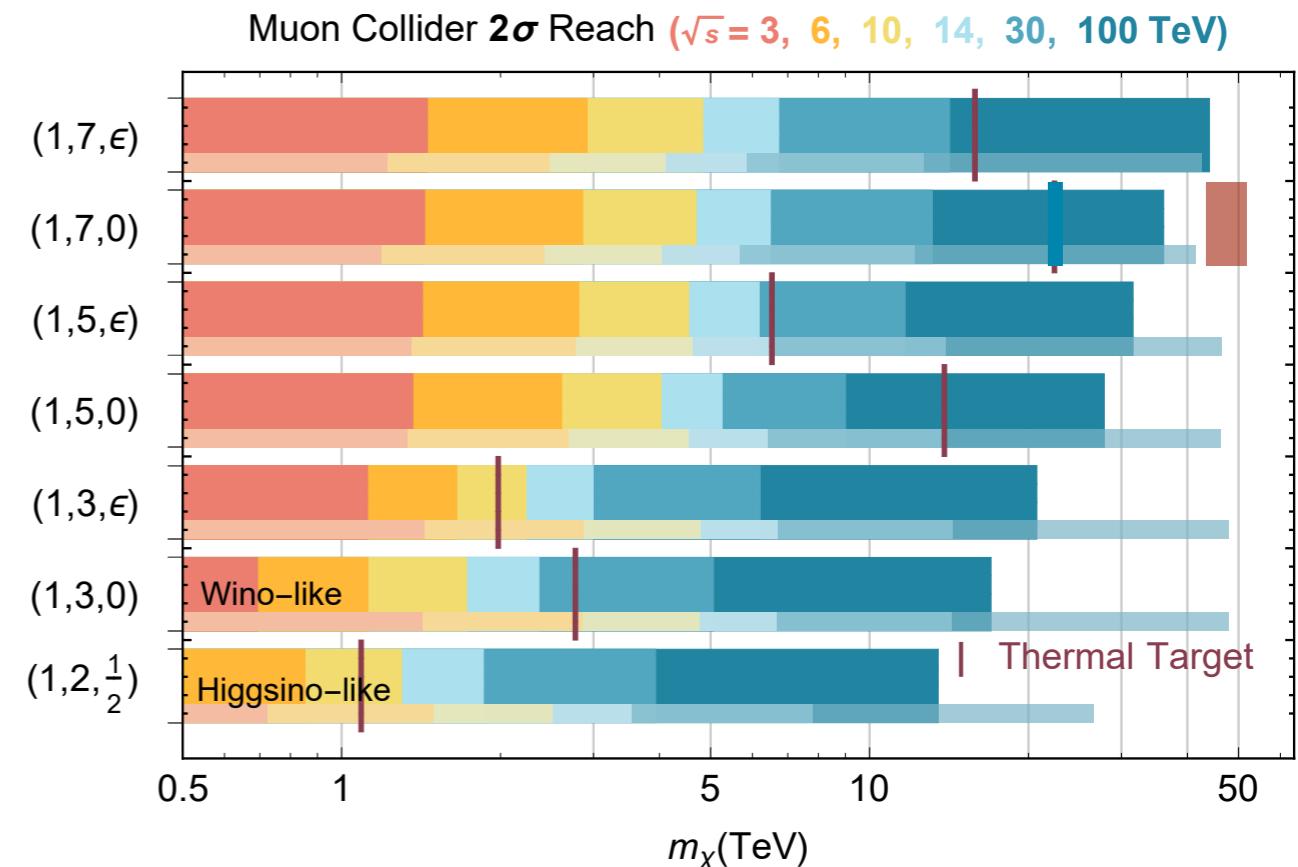
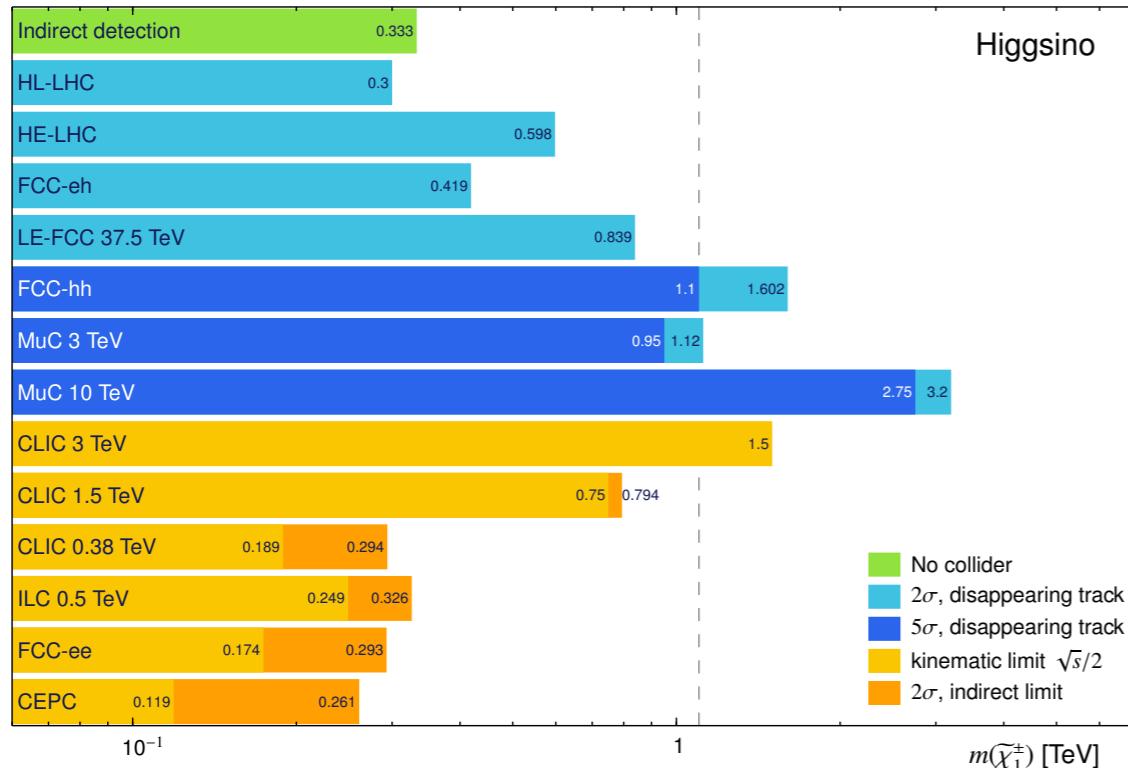


S. Bottaro, D. Buttazzo, M. Costa, R. Franceschini, P. Panci, D. Redigolo, L. Vittorio, 2107.09688

mono-X, more generic model independent. Interesting channels: muon-mu, mono-W.

Disappearing track. Some model dependence. Important to have the right BIB estimates.

# Reach



R. Capdevilla F. Meloni, J. Zurita, 2102.11292

T. Han, Z. Liu, X. Wang, LTW 2009.11287

With inclusive signal:  $E_{CM} \approx 14$  TeV enough to cover  $n \leq 3$  multiplets.  
Higher energy needed to cover higher multiplets.

If we have disappearing track: potential to reach almost  $m_\chi \approx 1/2 E_{CM}$

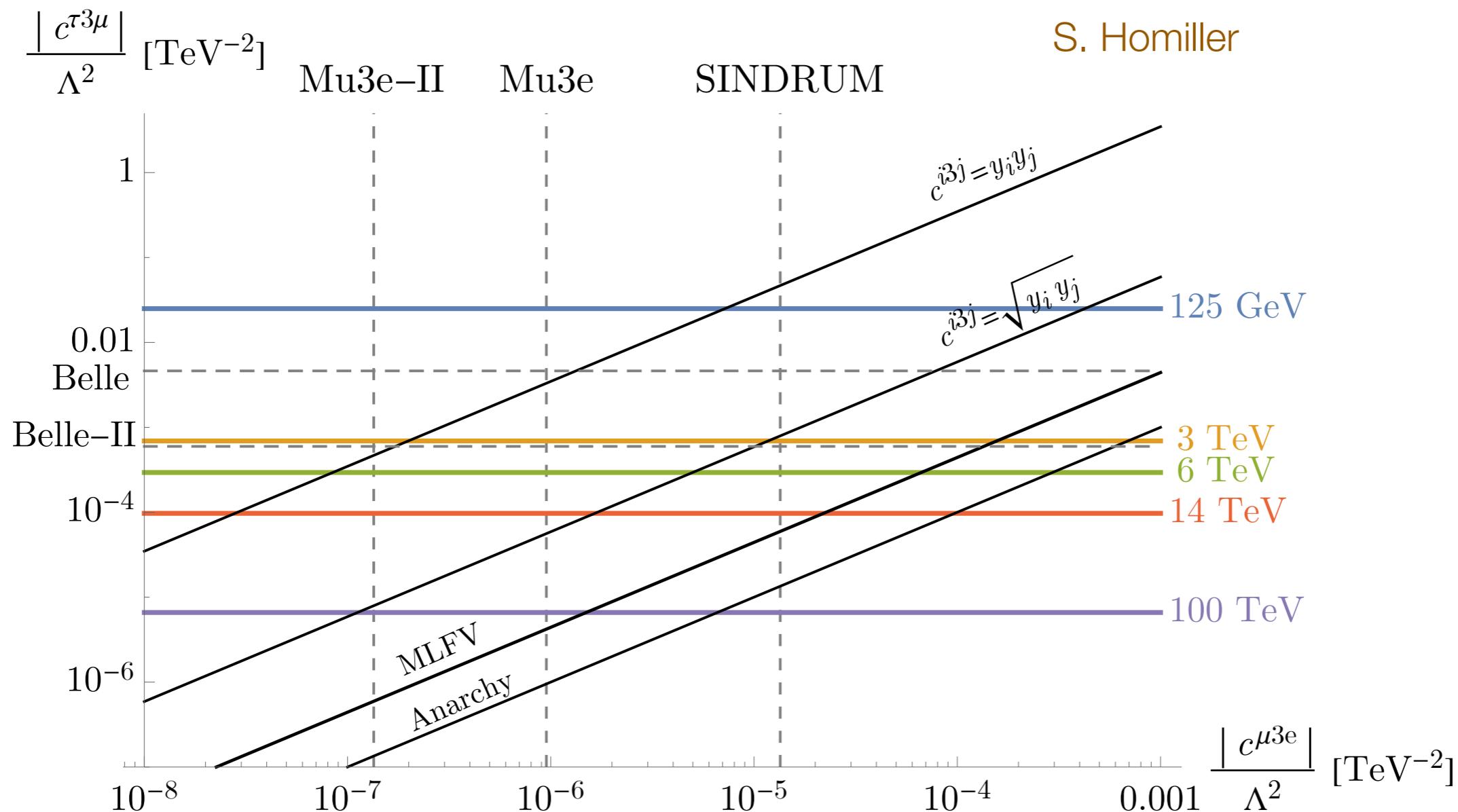
# (Other) new physics at high scale

- \* Dark matter (WIMP)
- \*  $Z'$
- \* Flavor/CP
- \* Lepto-quark
- \* ...

# Flavor (CP), complementarity

- \* The absence of (confirmed) BSM flavor (CP) signal point to a higher scale for flavor violation
  - \*  $\Lambda_{\text{flavor}} > 10(s) \text{ TeV}$
- \* Muon collider can (both directly and indirectly) probe this scale.

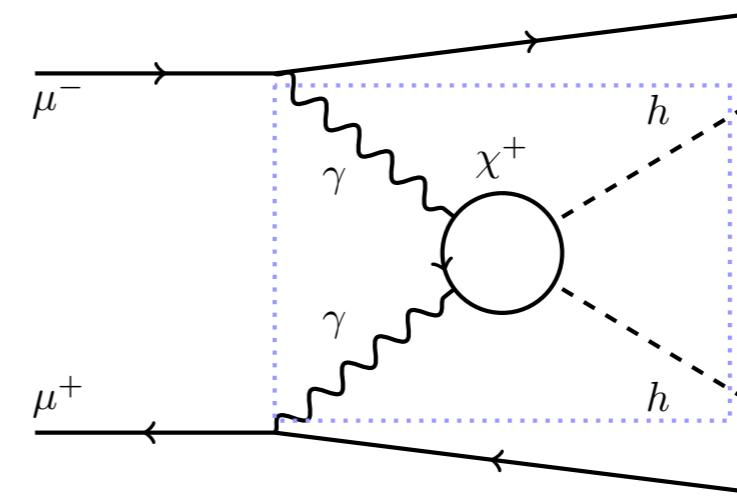
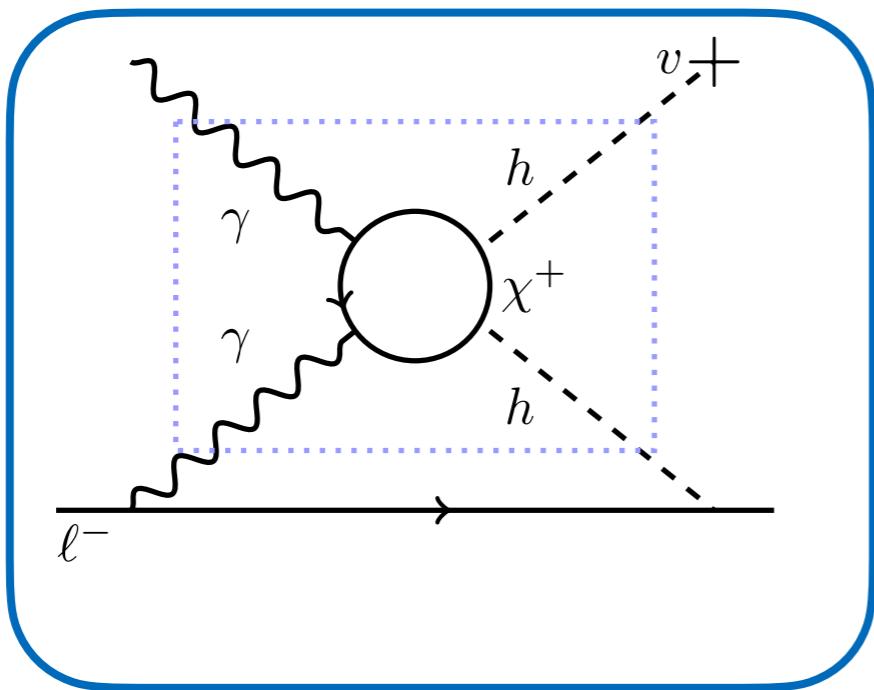
# Probing lepton flavor violation



See also (b anomaly motivated)

P. Asadi, R. Capdeville, C. Cesarotti, S. Homiller, 2014.05720

# EDM

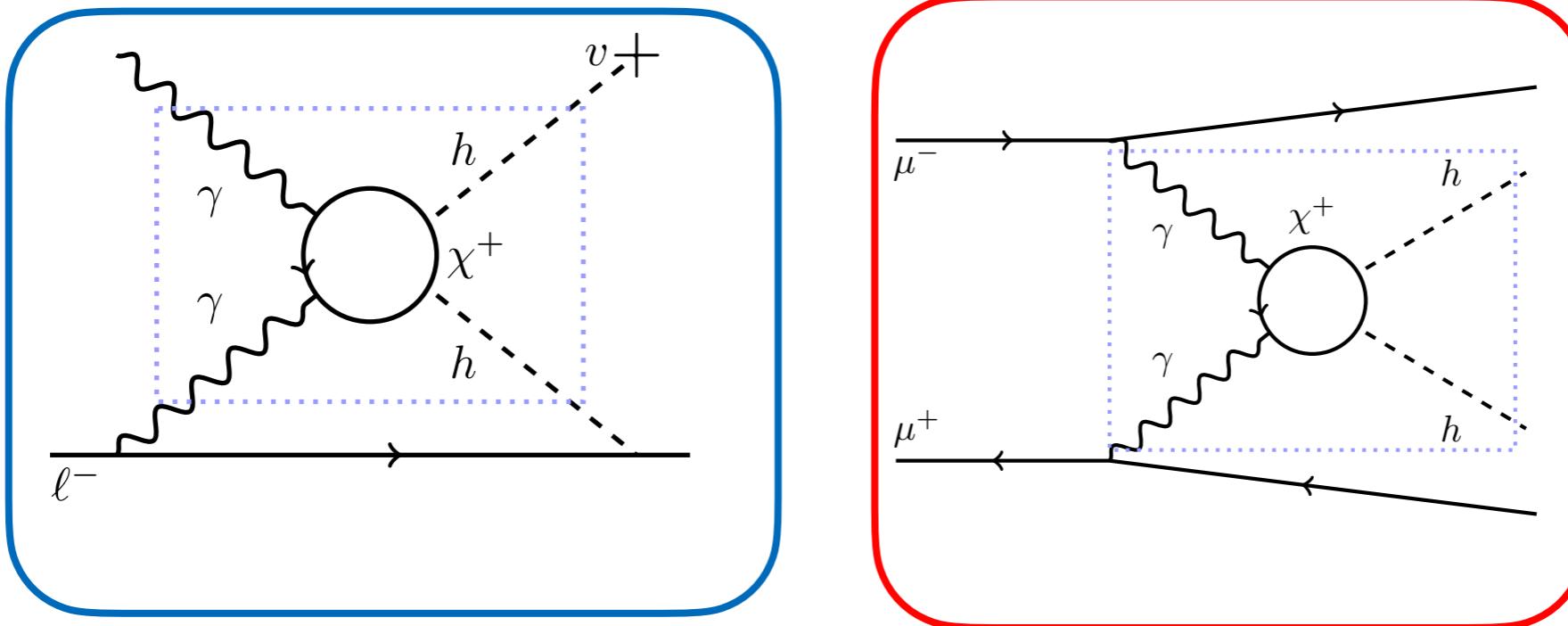


$$d_e \sim \sin(\delta_{\text{CP}}) \frac{e m_e}{M^2} \left( \frac{\alpha}{4\pi} \right)^k \simeq 10^{-32} e \text{ cm} \sin(\delta_{\text{CP}}) \times \begin{cases} (1 \text{ PeV}/M)^2 & \text{for } k = 1 \\ (20 \text{ TeV}/M)^2 & \text{for } k = 2 \end{cases}$$

2-loop Barr-Zee

Potential sensitivity of next generation exp.

# EDM



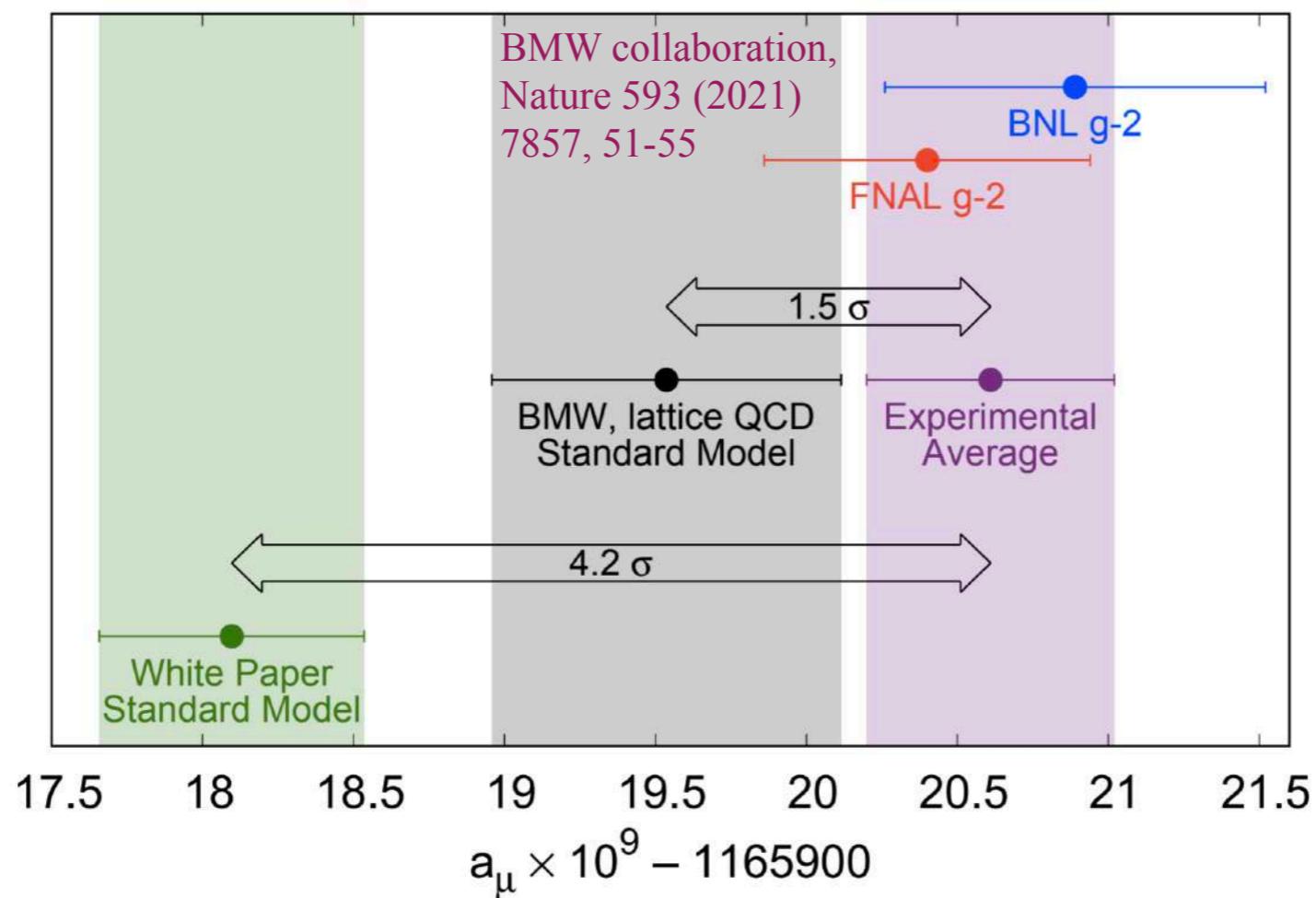
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2-loop Barr-Zee

Potential sensitivity of next generation exp.

Same process probed by 10(s) TeV muon collider!

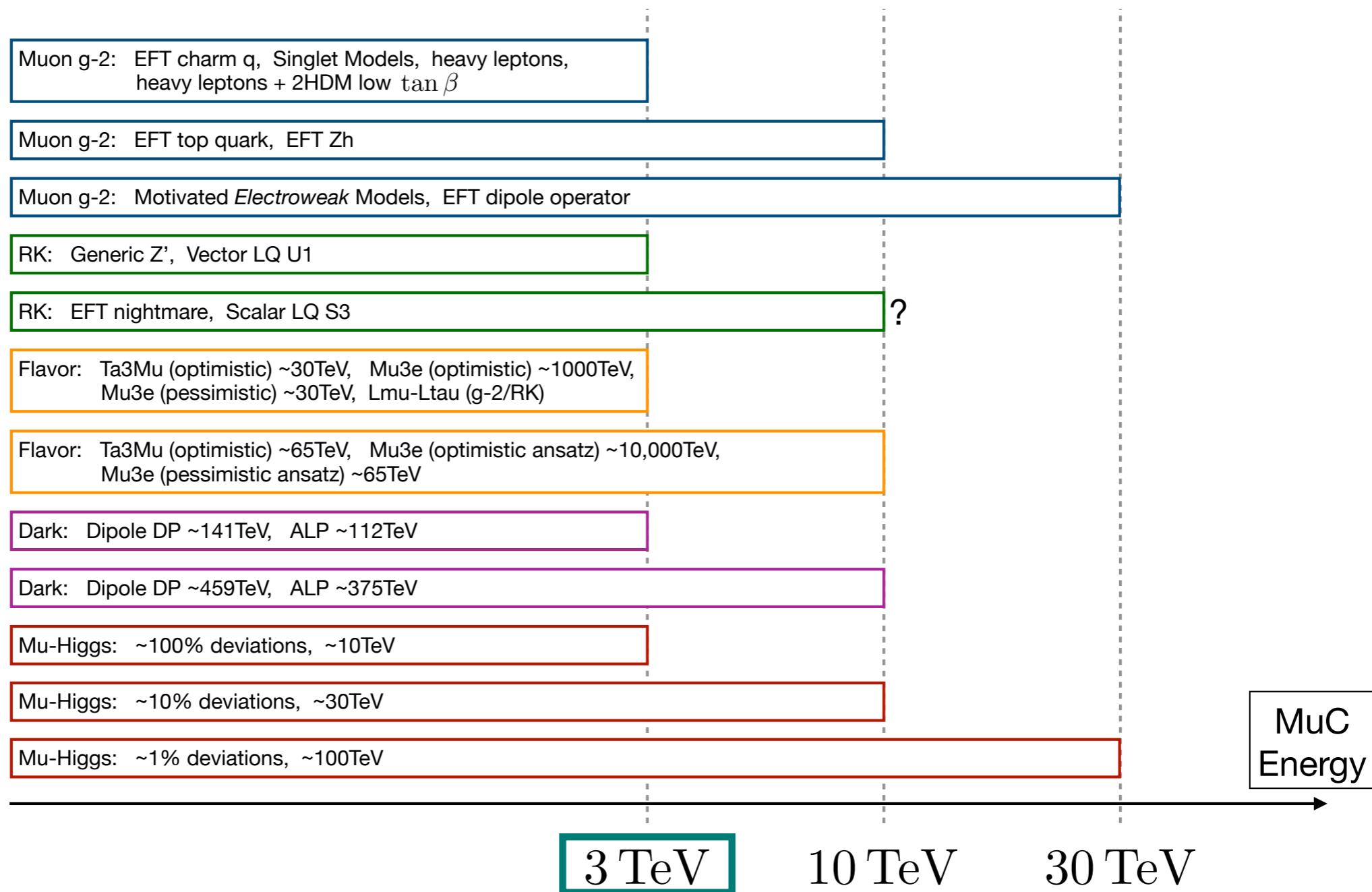
# What if new physics just like muon



Will be a new era, with a new definitive target.  
Much of the current would have to be re-done.

# Muon collider can do it, of course.

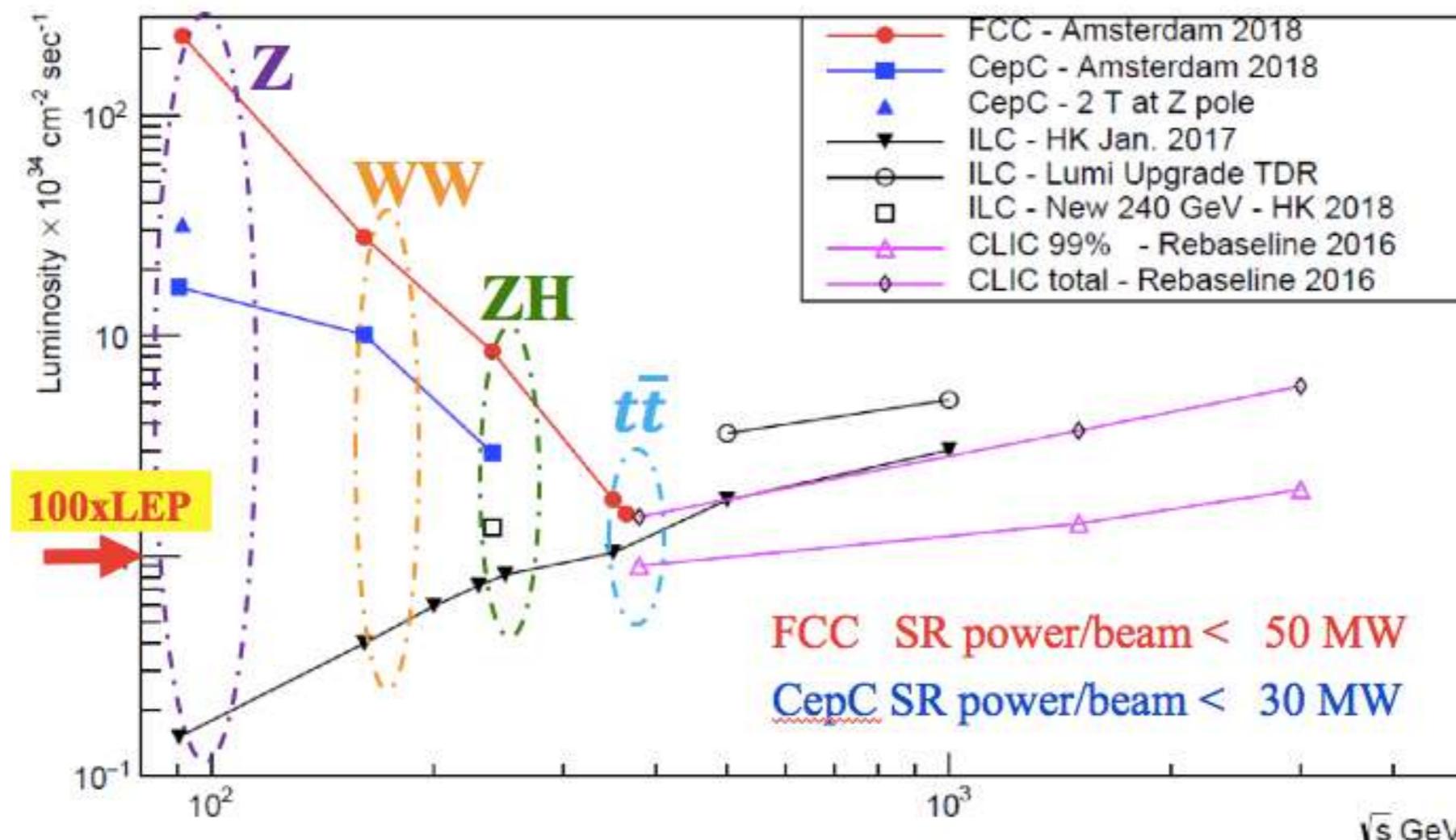
Slide of R. Capdevilla



# Conclusions

- \* Muon collider is a powerful machine at the energy frontier, with a rich physics program.
- \* Great Higgs measurements for  $E_{CM} > 3 \text{ TeV}$ .
- \* Exciting reaches for a broad spectrum of new physics in the range multiple - 10(s) TeV.
  - \* Tests of WIMP DM, naturalness, flavor violations, ...

## e<sup>+</sup>e<sup>-</sup> Collider Luminosities



e<sup>+</sup> e<sup>-</sup> colliders will play an essential role!

Offer rich program, covering 90 eV - TeV(s).

# Muon Collider

Daniel Schulte for the Muon Collider Collaboration

# Why Muon Collider?

Most of the muon collider R&D has been done by MAP in the US

- Experimental programme at MICE in the UK
- alternative LEMMA concept considered mainly at INFN
- Still the basis for our designs

Interest in Europe started with last Strategy Update

- **Change of goals:** Started looking for very high energy high-luminosity lepton collider
  - The champion is CLIC at 3 TeV, which has been optimised over decades for this
    - 18 GCHF, 590 MW power consumption
  - The muon collider promises to be able to go to 10 TeV or higher
- **Technology and design advances** since MAP
  - e.g. superconducting magnet technology (HTS)
  - e.g. rectilinear cooling channel
  - Progress on specific technologies for the muon collider
  - Expect competitive cost and power consumption

CERN allocated budget to muon collider and initiated collaboration

Muon collider is part of European Accelerator R&D Roadmap

- evaluated in detail, with **lots of help from US experts**

# Proton-driven Muon Collider Concept



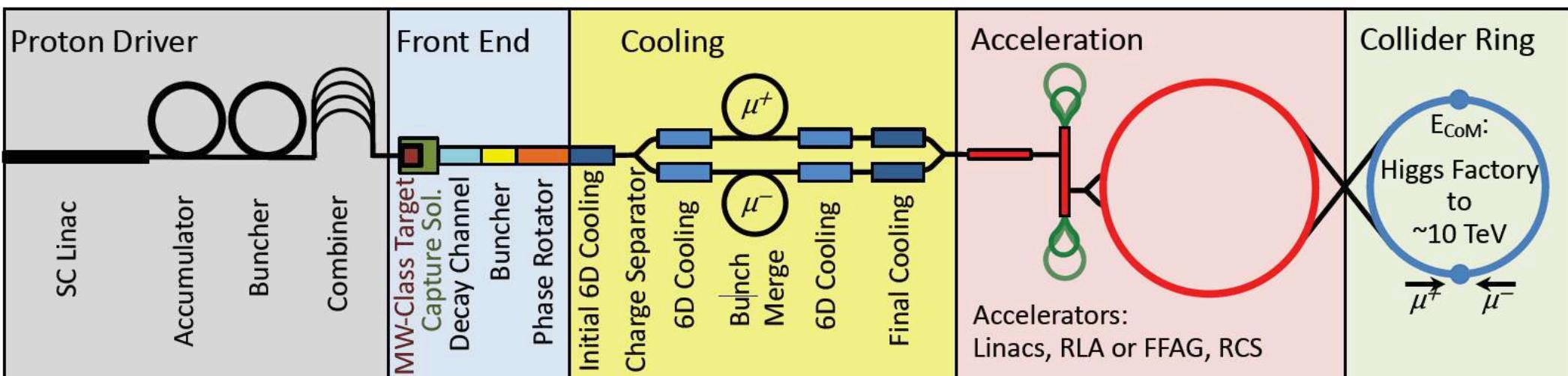
Some exploration long time ago in Europe

MAP study in US developed the concept, MICE experiment in UK (but US did not go for high-energy collider)

Alternative developed at INFN (but needs consolidation)

With last Strategy for Particle Physics interest in Europe started

Also interest in US and elsewhere again



Produce short,  
intense proton  
bunch

Ionisation cooling of  
muon in matter

Acceleration to  
collision energy

Collision

Protons in target  
produce pions which  
decay into muons  
muons are captured

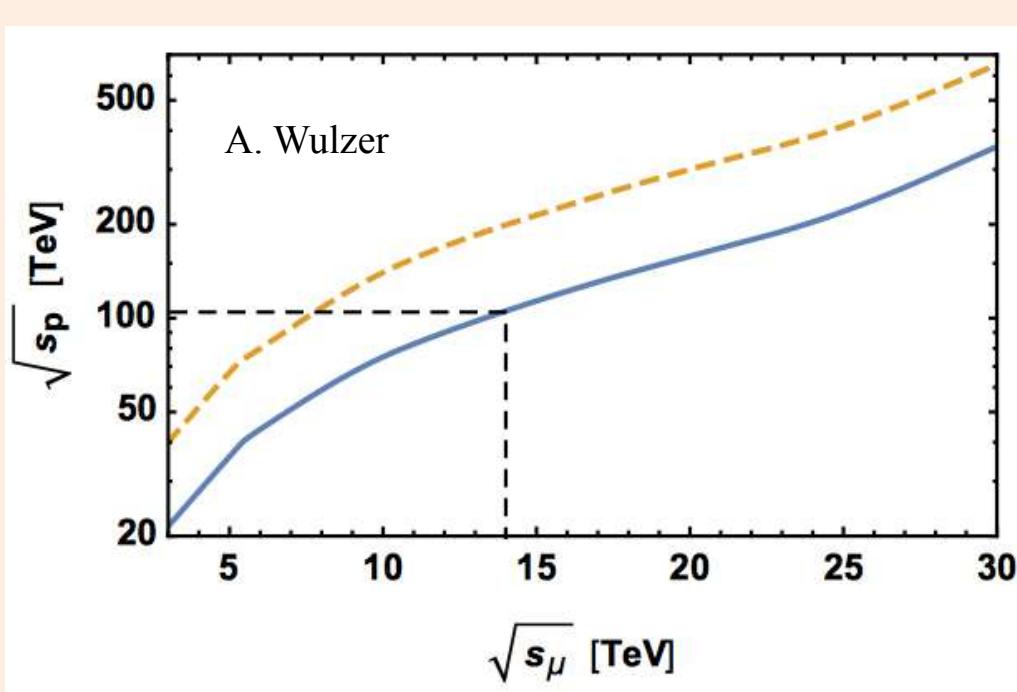
Would be easy if the muons did not decay  
Lifetime is  $\tau = \gamma \times 2.2 \mu\text{s}$

# Lepton Physics at High Energy

High energy lepton colliders are precision and discovery machines

$$V = \frac{1}{2} m_h^2 h^2 + (1 + k_3) \lambda_{hhh}^{SM} v h^3 + (1 + k_4) \lambda_{hhhh}^{SM} h^4$$

Chiesa, Maltoni, Mantani,  
Mele, Piccinini, Zhao  
[Muon Collider -](#)  
[Preparatory Meeting](#)



Precision potential

Measure  $k_4$  to some 10%  
With 14 TeV,  $20 \text{ ab}^{-1}$

Discovery reach

14 TeV lepton collisions are comparable  
to 100 TeV proton collisions for  
production of heavy particle pairs

Luminosity goal

(Factor O(3) less than CLIC at 3 TeV)  
 $4 \times 10^{35} \text{ cm}^{-2} \text{s}^{-1}$  at 14 TeV

$$L \gtrsim \frac{5 \text{ years}}{\text{time}} \left( \frac{\sqrt{s}_\mu}{10 \text{ TeV}} \right)^2 2 \cdot 10^{35} \text{ cm}^{-2} \text{s}^{-1}$$

# Muon Collider Promises



CLIC is at the limit of what one can do (decades of R&D)

- No obvious way to improve luminosity

Luminosity per beam power increases with energy in muon collider

- **power efficient**

Site is **compact**

- 10 TeV comparable to 3 TeV CLIC

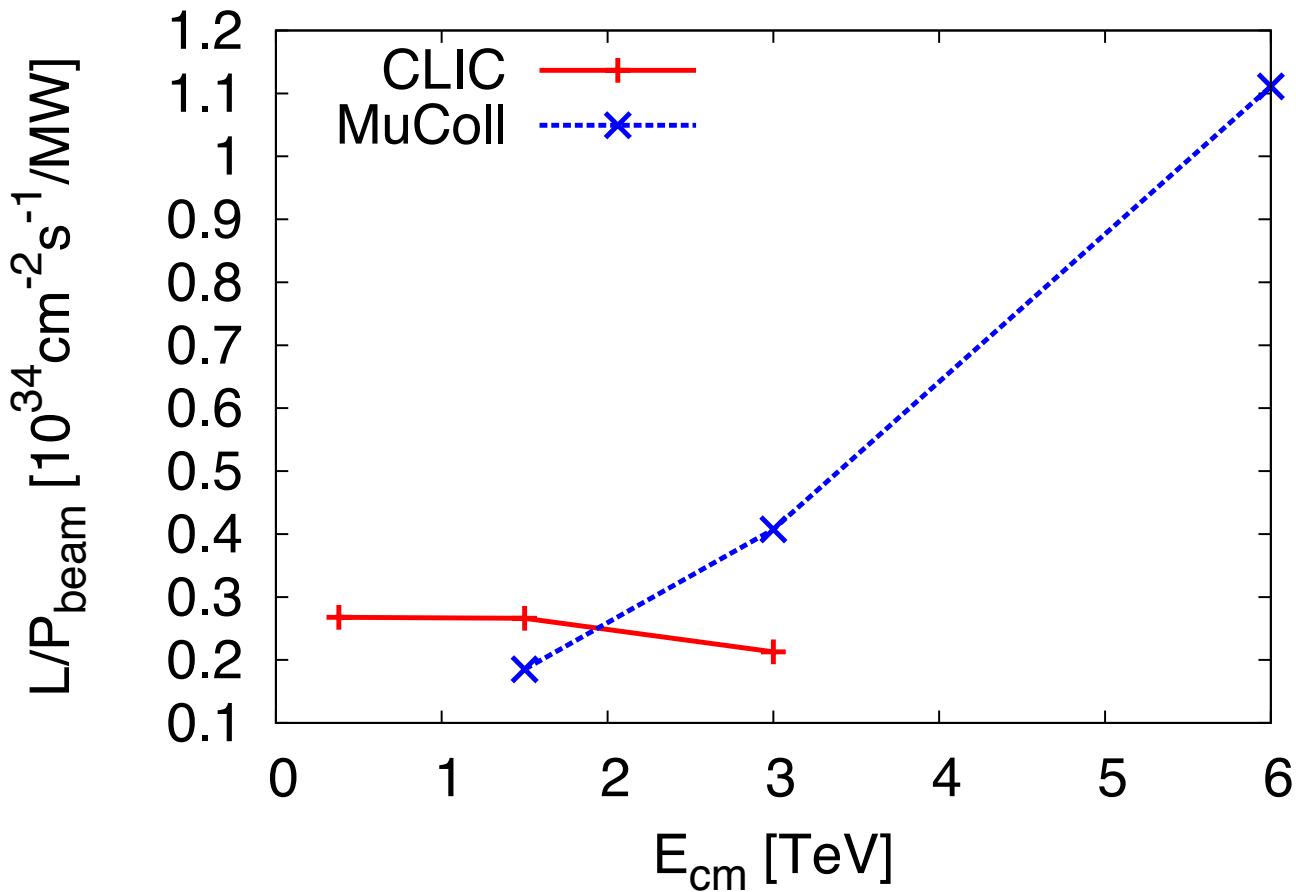
**Staging** is natural

- Each ring accelerates by a factor of a few

Promises **cost effectiveness**

- but need detailed study

Other **synergies** (neutrino/higgs)



Muon collider promises unique opportunity for a **high-energy, high-luminosity lepton collider**

# Luminosity Goals

Target integrated luminosities

$\sqrt{s}$	$\int \mathcal{L} dt$
3 TeV	1 ab $^{-1}$
10 TeV	10 ab $^{-1}$
14 TeV	20 ab $^{-1}$

Note: currently consider 3 TeV and either 10 or 14 TeV

- Tentative parameters achieve goal in 5 years
- FCC-hh to operate for 25 years
- Might integrate some margins
- Aim to have two detectors

Now study if these parameters lead to realistic design with acceptable cost and power

Tentative target parameters  
Scaled from MAP parameters

Parameter	Unit	3 TeV	10 TeV	14 TeV
L	$10^{34} \text{ cm}^{-2}\text{s}^{-1}$	1.8	20	40
N	$10^{12}$	2.2	1.8	1.8
f <sub>r</sub>	Hz	5	5	5
P <sub>beam</sub>	MW	5.3	14.4	20
C	km	4.5	10	14
<B>	T	7	10.5	10.5
$\varepsilon_L$	MeV m	7.5	7.5	7.5
$\sigma_E / E$	%	0.1	0.1	0.1
$\sigma_z$	mm	5	1.5	1.07
$\beta$	mm	5	1.5	1.07
$\varepsilon$	$\mu\text{m}$	25	25	25
$\sigma_{x,y}$	$\mu\text{m}$	3.0	0.9	0.63

Comparison:  
CLIC at 3 TeV: 28 MW

# International Muon Collider Collaboration



## Memorandum on Cooperation for the Muon Collider (MC) Study

THE INSTITUTES, LABORATORIES, UNIVERSITIES AND FUNDING AGENCIES AND OTHER SIGNATORIES OF THIS MEMORANDUM ON COOPERATION AND CERN AS THE HOST ORGANIZATION ("the Participants")

CERN is the initial host organization for the MC Study.

Goal is to

- [...] establish whether the investment into a full Conceptual Design Report ("CDR") and demonstrator for a muon collider is scientifically justified. [...]
- [...] focus on the high-energy frontier and consider options with a centre-of-mass energy of 3 TeV and of 10 TeV or more.
  - **10+ TeV option** is the reason to study muon colliders,
  - **3 TeV option** might be initial energy step with technologies available in 10-20 years
- Potential synergies [...] shall be explored [...].
  - Neutrino facilities and potentially higgs factories

[muon.collider.secretariat@cern.ch](mailto:muon.collider.secretariat@cern.ch)

## Partners

(in some cases formal procedure is not yet completed)

- INFN
- Commissariat à l'Energie Atomique (CEA)
- Iowa State University
- University of Huddersfield
- University of Warwick
- Institute of High Energy Physics
- University of Rostock
- University of Oxford (JAI)
- Tampere University
- Uppsala University
- TU Darmstadt
- Peking University
- Sun Yat-Sen University
- European Spallation Source (ESS)
- Rutherford Appleton Laboratory (RAL)
- University de Genève

# Muon Collider Luminosity Scaling



Fundamental limitation

Assumes no emittance growth after source and no technical limitation

Applies to MAP and LEMMA scheme

$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_{\delta} \frac{N_0}{\epsilon \epsilon_L} f_r N_0 \gamma$$

High energy

High field in collider ring  
= small ring  
= many collisions

Large energy acceptance  
= short bunch  
= small betafunction

Dense beam

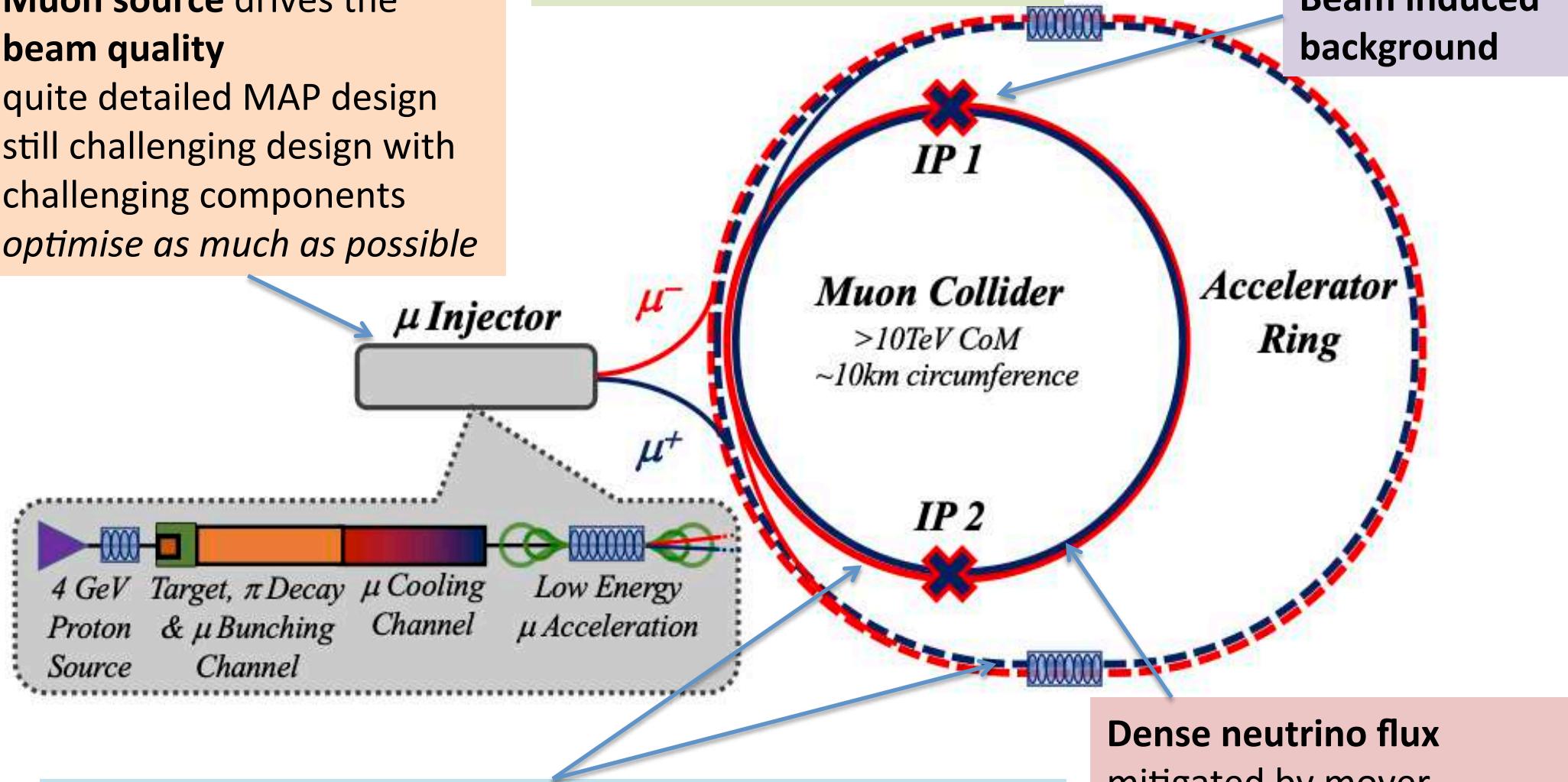
High beam power

Note: emittances are normalised

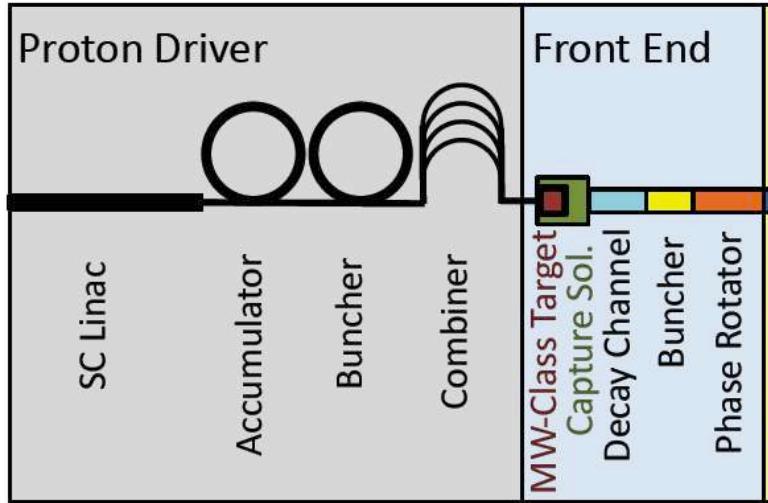
# Key Challenges

Muon source drives the beam quality  
 quite detailed MAP design  
 still challenging design with challenging components  
*optimise as much as possible*

## Physics potential assessment

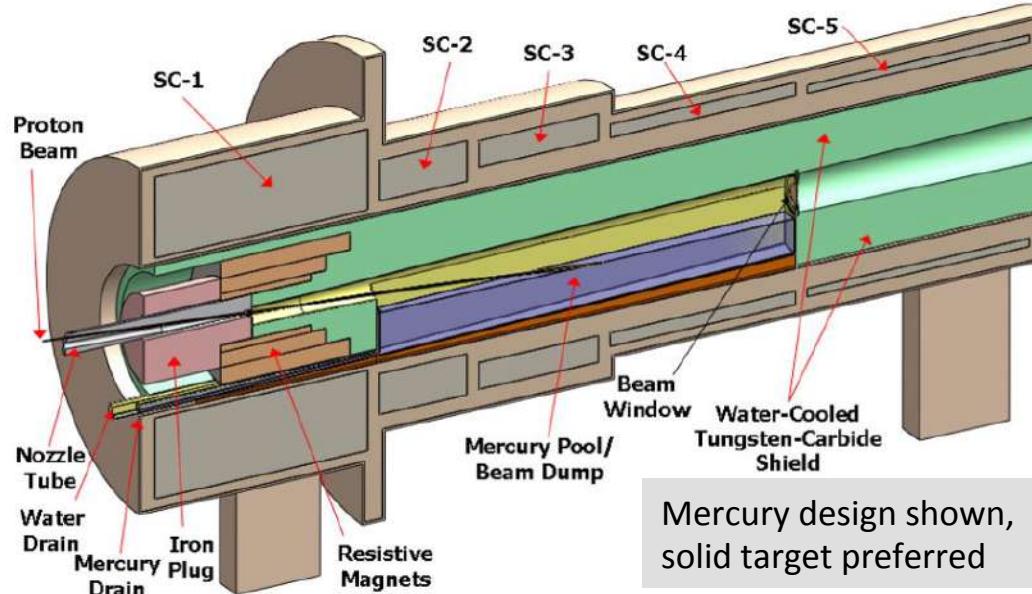


# Proton Complex and Target Area



Proton beam power is no issue, some look required at  
**H- source and accumulator and combiner complex**

**2 MW proton beam**  
requires radiation protection



Mercury design shown,  
solid target preferred

**High field** to efficiently collect pions/  
muons: 20 T, then tapering  
Using copper solenoid in  
superconducting solenoid

**Large aperture**  $O(1.2\text{m})$   
to allow shielding

# Target Design

Two approaches:

- Superconducting O(15 T) outer solenoid boosted by resistive inner solenoid
- O(20 T) HTS solenoid

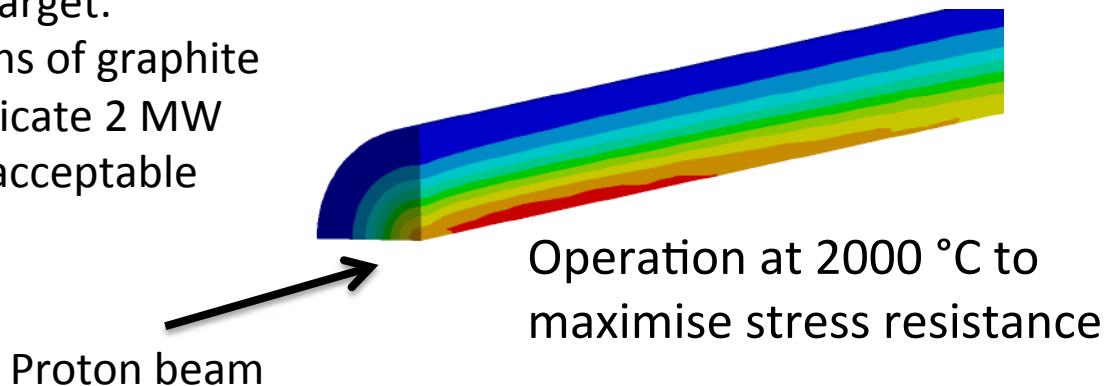
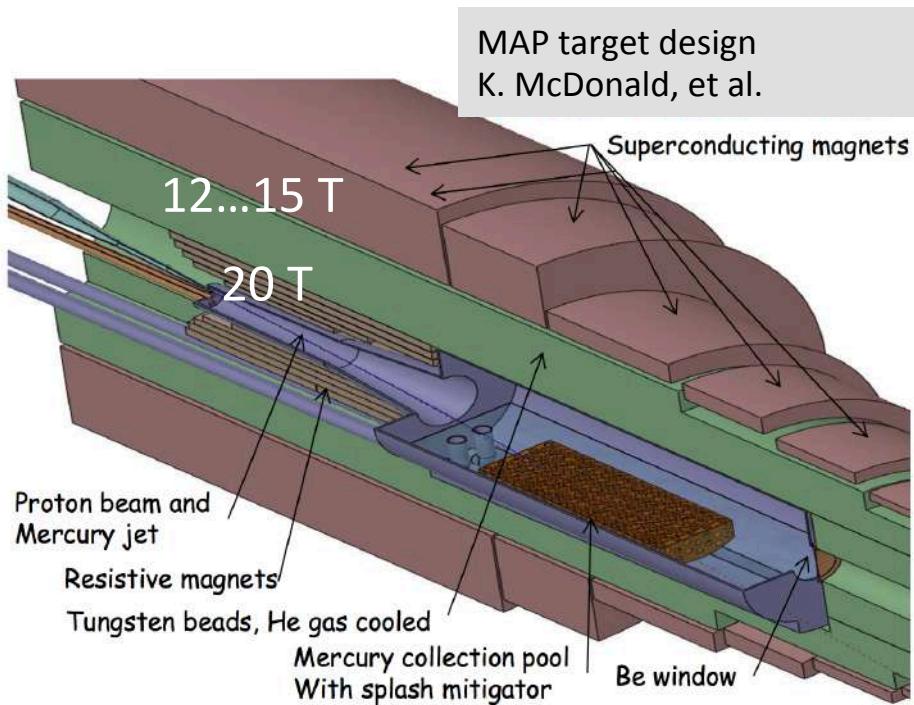
Need to shield superconducting solenoid  
 ⇒ larger aperture

Past simulations showed

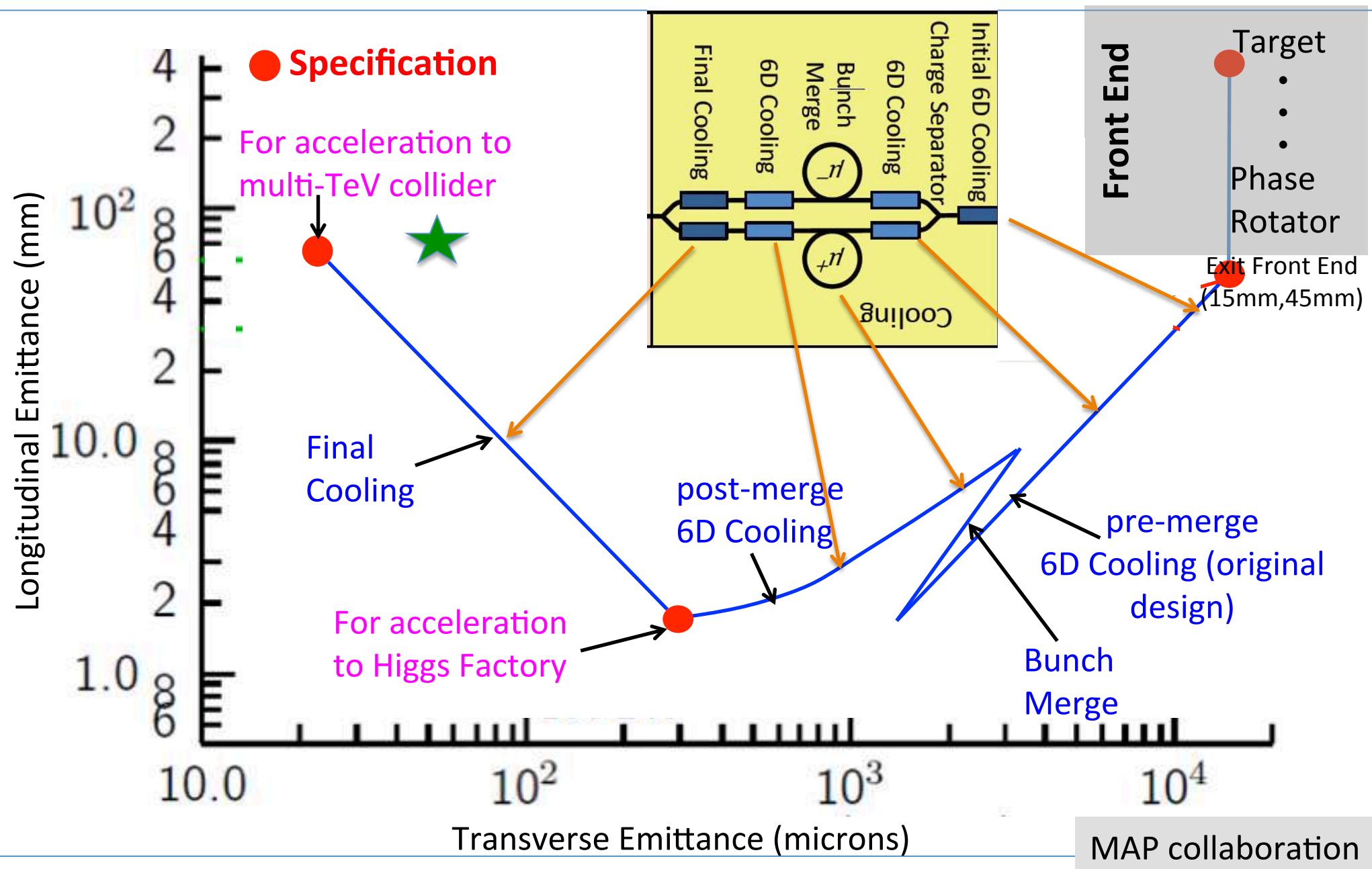
- 5 kW to outer (O(2 m) aperture)
- 100 kW to inner



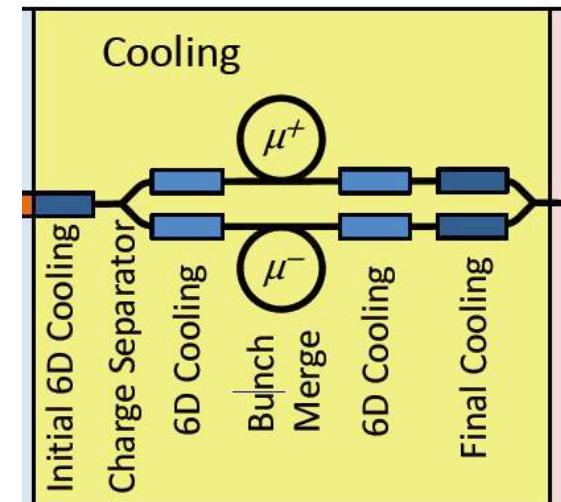
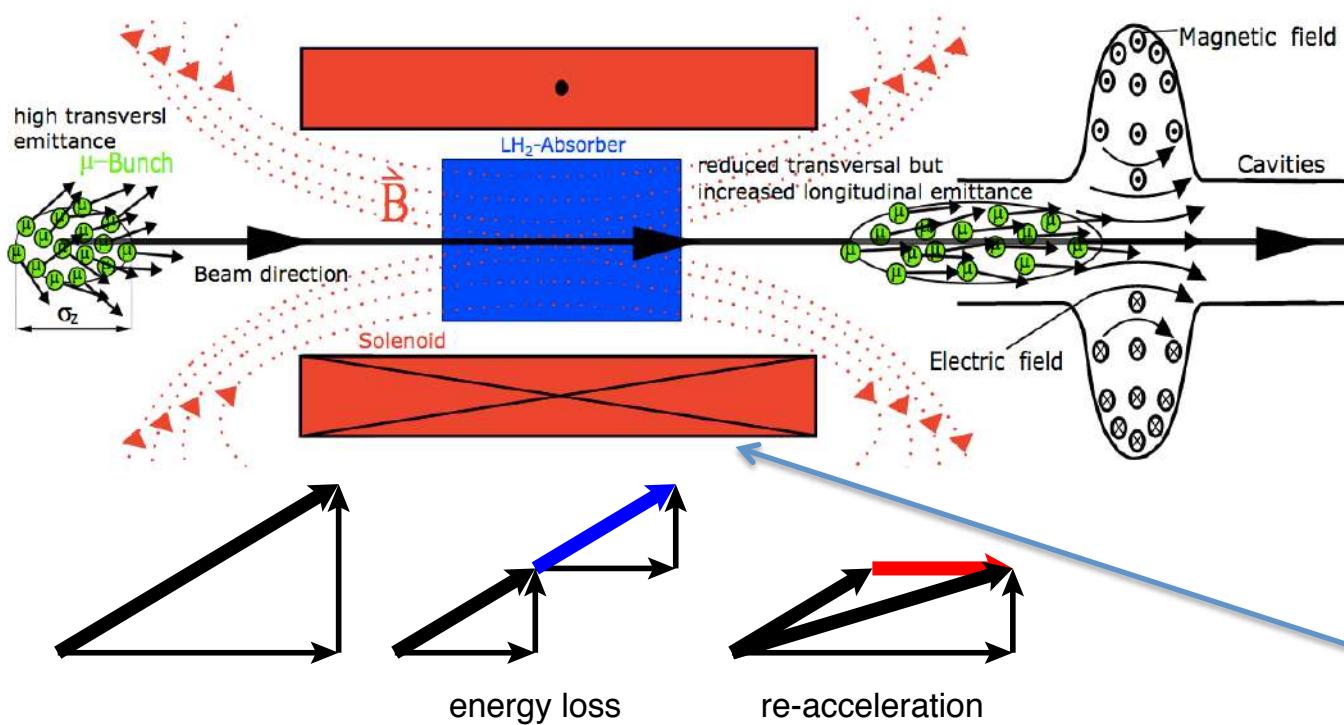
Shock in target:  
 Simulations of graphite  
 target indicate 2 MW  
 could be acceptable



# Cooling: The Emittance Path



# Final Cooling Challenge



High field solenoids minimise beta-function and impact of multiple scattering

Energy loss = cooling

Multiple scattering = heating

$$\frac{d\epsilon_{\perp}}{ds} = -\frac{1}{(v/c)^2} \frac{dE}{ds} \frac{\epsilon_{\perp}}{E} + \frac{1}{2} \frac{1}{(v/c)^3} \left( \frac{14 \text{ MeV}}{E} \right)^2 \frac{\beta \gamma}{L_R}$$

# Equilibrium Emittance, Material and Energy



$$\epsilon_{equ}(E_\mu) \propto \frac{1}{L_R} \frac{dE}{ds} \Big|_{E_\mu} \frac{1}{B_z} \gamma$$

Material property best is hydrogen

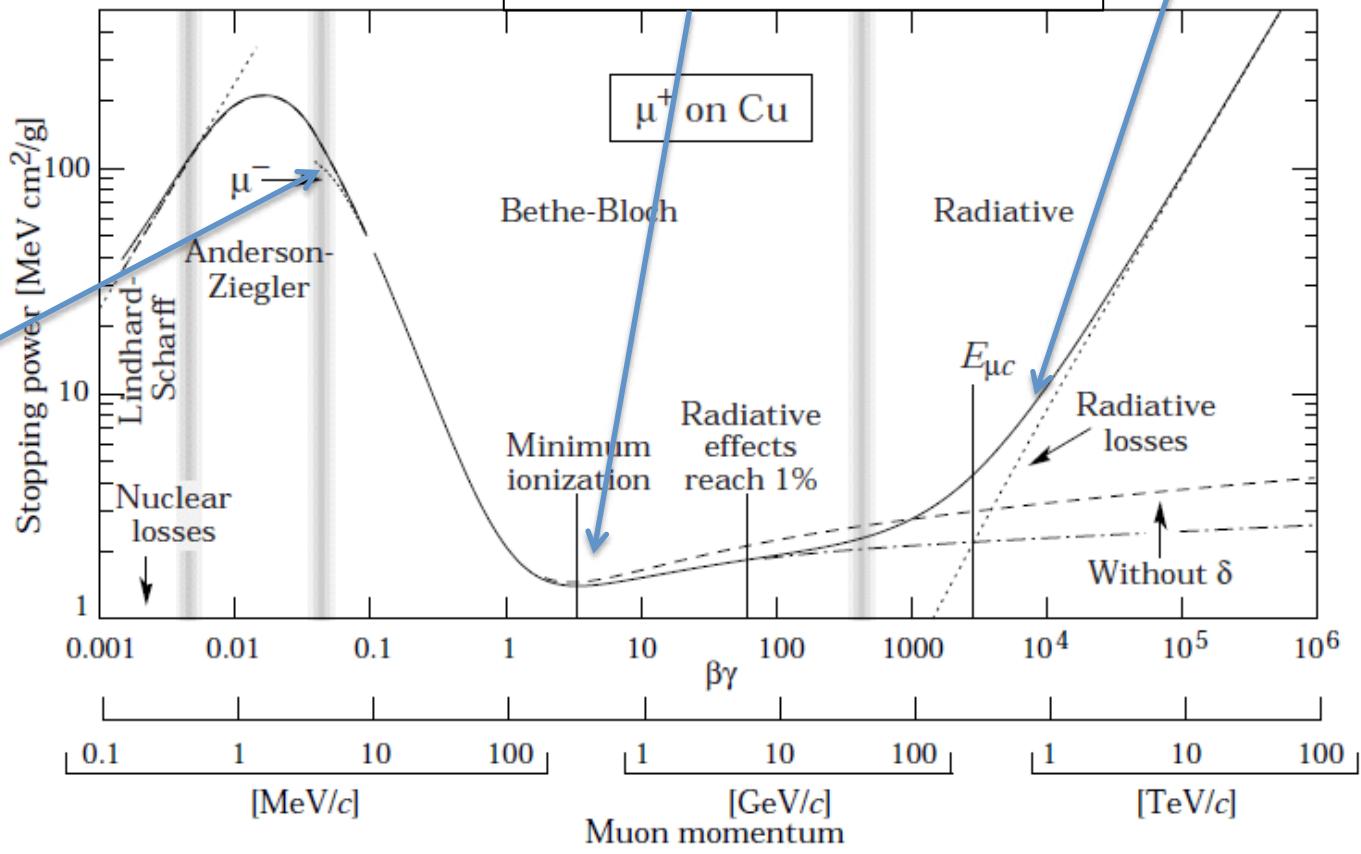
Lithium hydride is good solid material

best at low energies but rapidly increasing energy spread  
 ⇒ use for final cooling

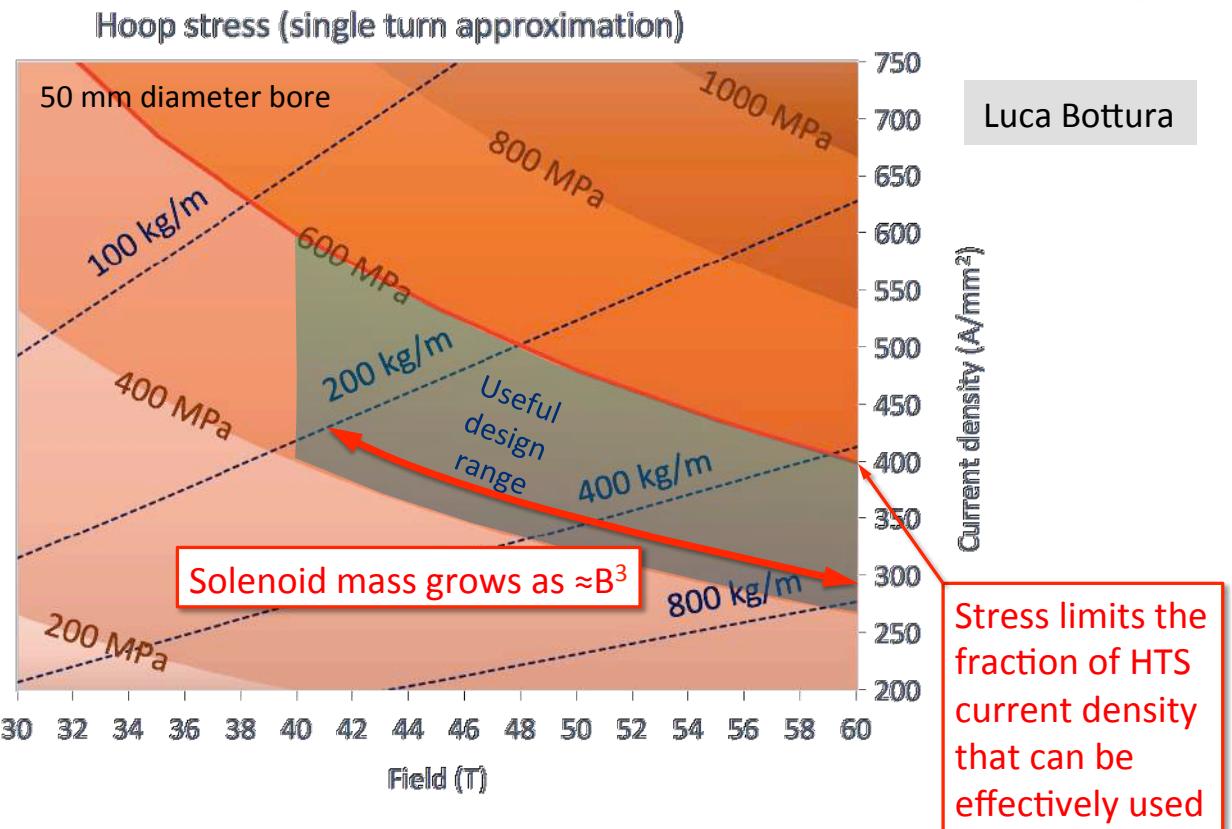
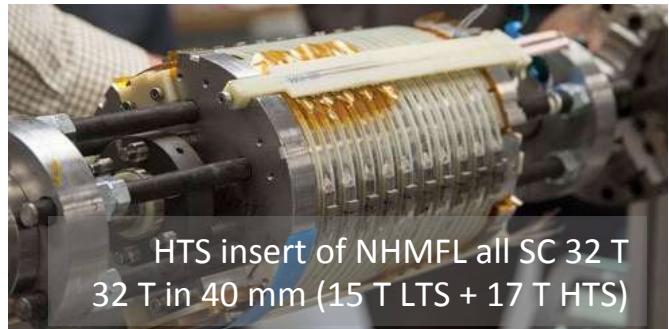
Strong solenoid field helps

Hard stochastic losses

Most of cooling is done here



# Solenoid Fields



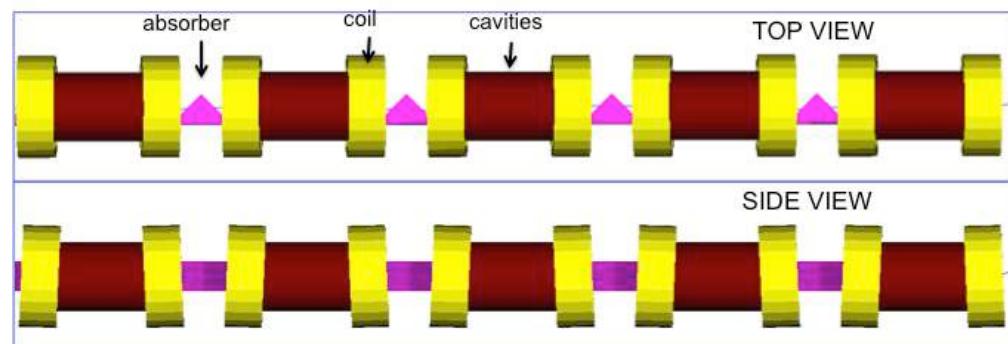
Solenoids with >30 T and aperture of 50 mm exist for high magnetic field science

The limit is the stress in the HTS

Even 60 T might be possible

- but important R&D required to see how far we can push

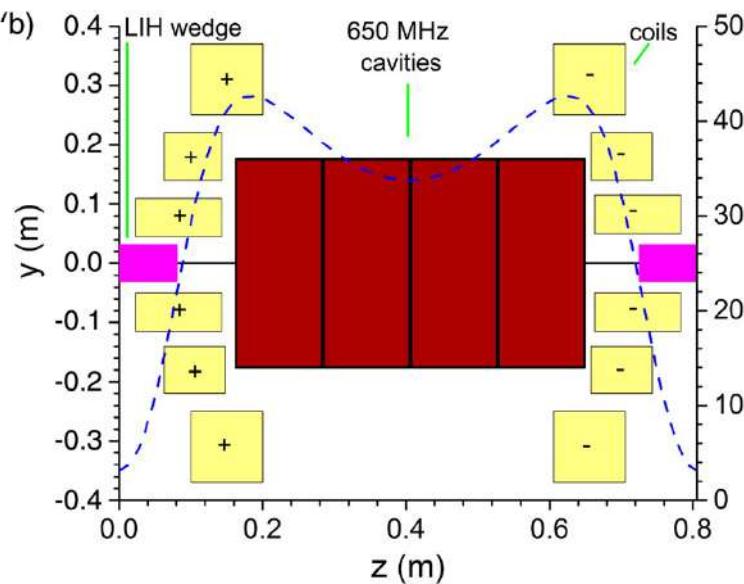
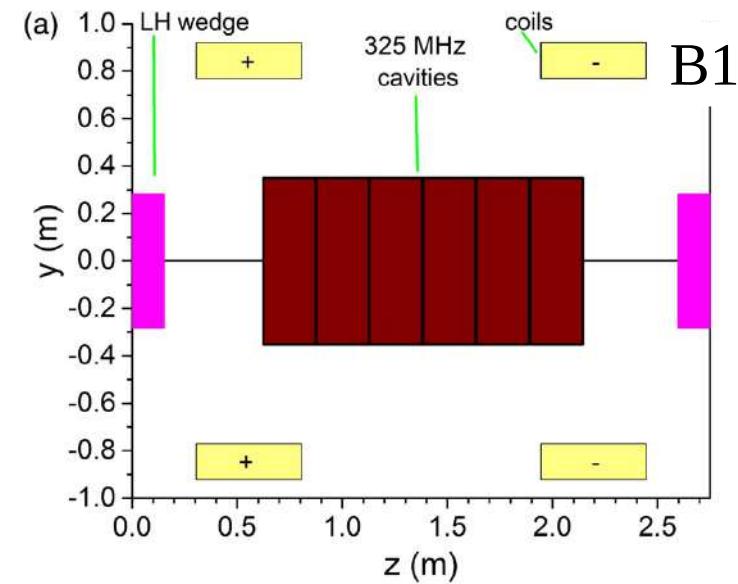
# 6D Cooling Cell Design



**MuCool:** >50 MV/m in 5 T

Two solutions

- H<sub>2</sub>-filled copper cavities
- Cavities with Be end caps



**High-gradient cavities in high magnetic field**

**Tight integration** of solenoids, RF, absorbers, instrumentation, cooling, vacuum, alignment, ...

Will aim for further optimisation  
 This is the **unique** and **novel** system of the muon collider  
 Will need a **test facility**

# RF Cavity Challenge

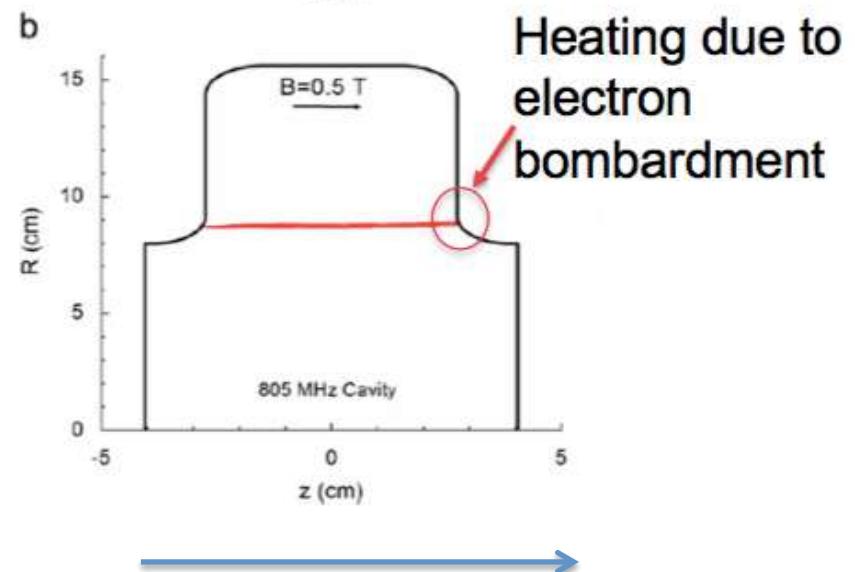
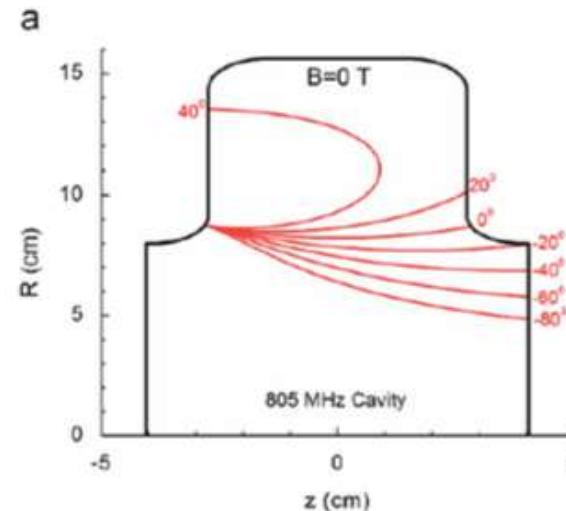
Gradient of normal-conducting RF cavities is typically limited by breakdown

Model:

- electrons start at one emitter site
- electrons are accelerated by cavity field and hit other side

In solenoid field:

- electrons spiral around magnetic field line
- they all hit one spot on the other site
- this can heat spot above plastic deformation stress
- and lead to breakdown
- in any case electrons bounce back and forth between sites



# RF Cavities: Solutions

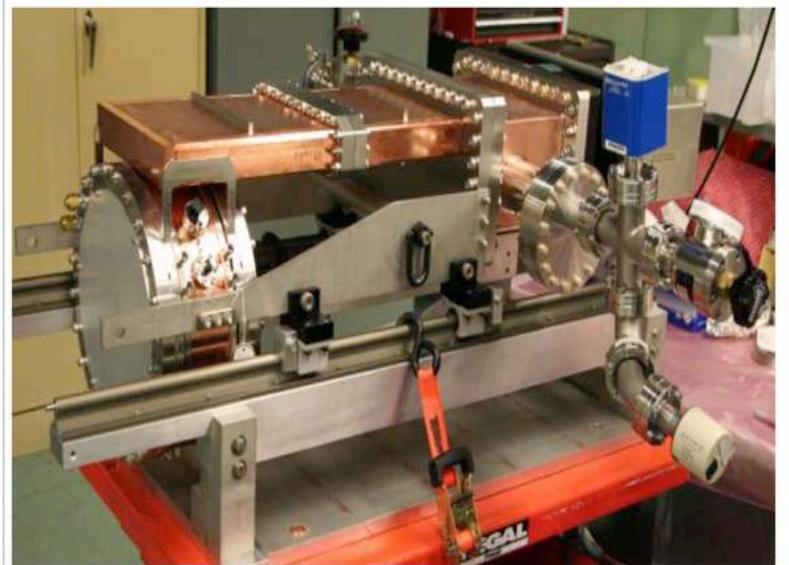
Solution 1: Reduce losses in the wall

- Shorter pulses
- Replace copper with beryllium
  - lower density, smaller Z

Experimental result (MUCOOL):

- Beryllium cavity in 3 T: 50 MV/m with 30  $\mu$ s RF pulse

To be studied: gradient in higher field

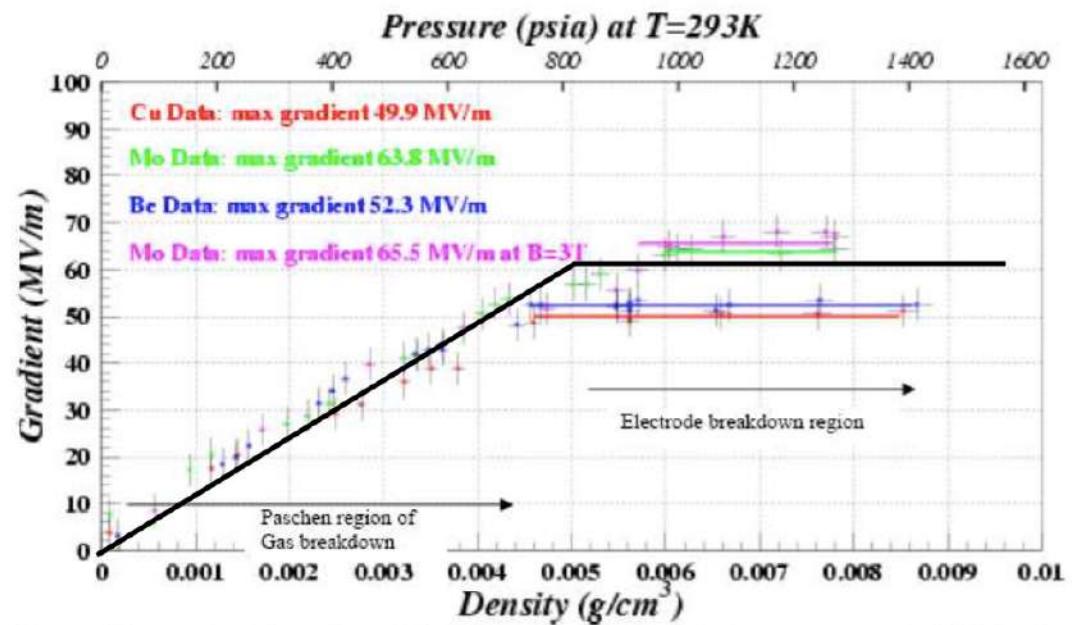


Solution 2: Use high-pressure hydrogen gas

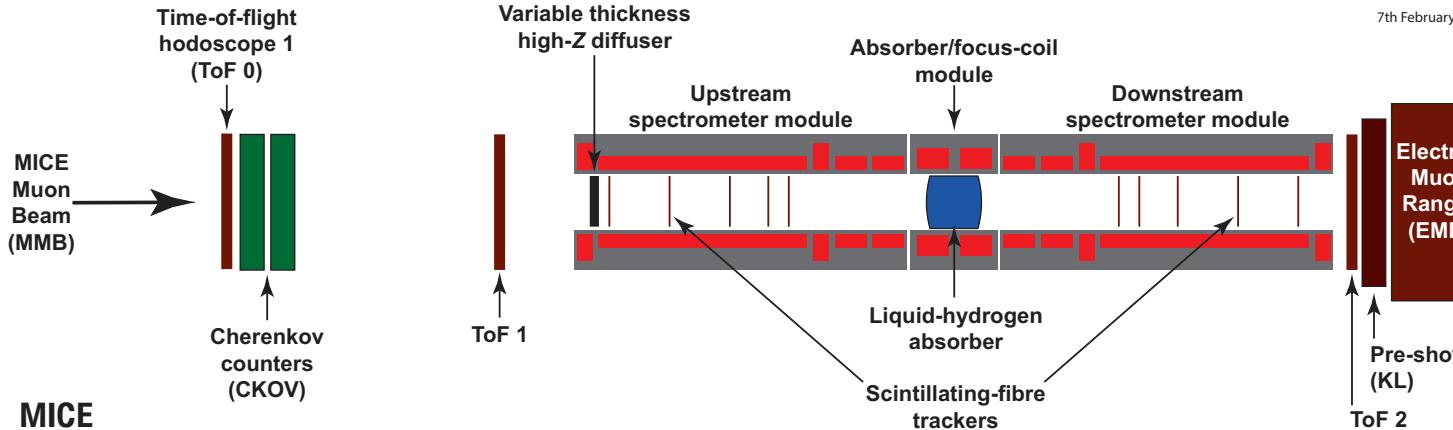
- until mean free path allows less than 15 eV energy gain (ionisation energy)

Experimental results (MUCOOL):

- Molybdenum cavity in 3 T: 65 MV/m
- for low density, limited by gas density (gradient is proportional to density)
- Final value depends on surface property



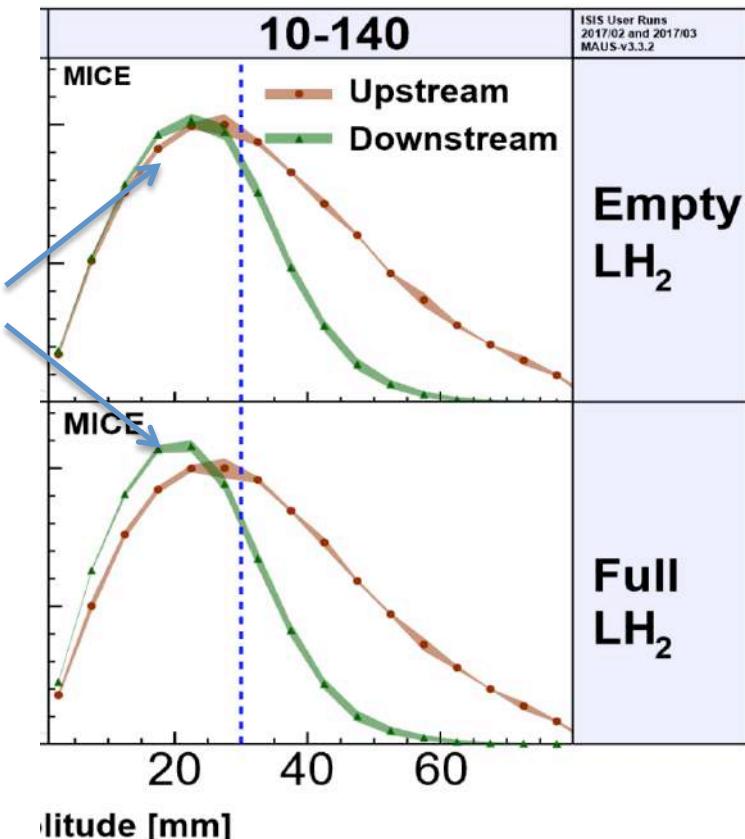
# MICE (in the UK)



MICE

More particles at smaller amplitude after absorber is put in place

Principle of ionisation cooling has been demonstrated

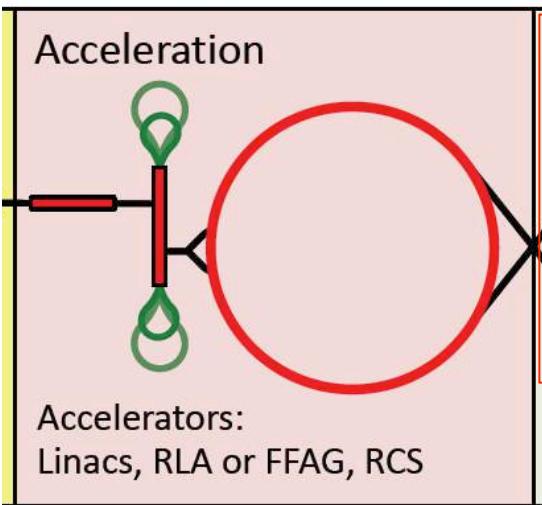


Nature volume 578,  
pages 53-59 (2020)

More complete experiment with higher statistics, more than one stage required

Integration of magnets, RF, absorbers, vacuum is engineering challenge

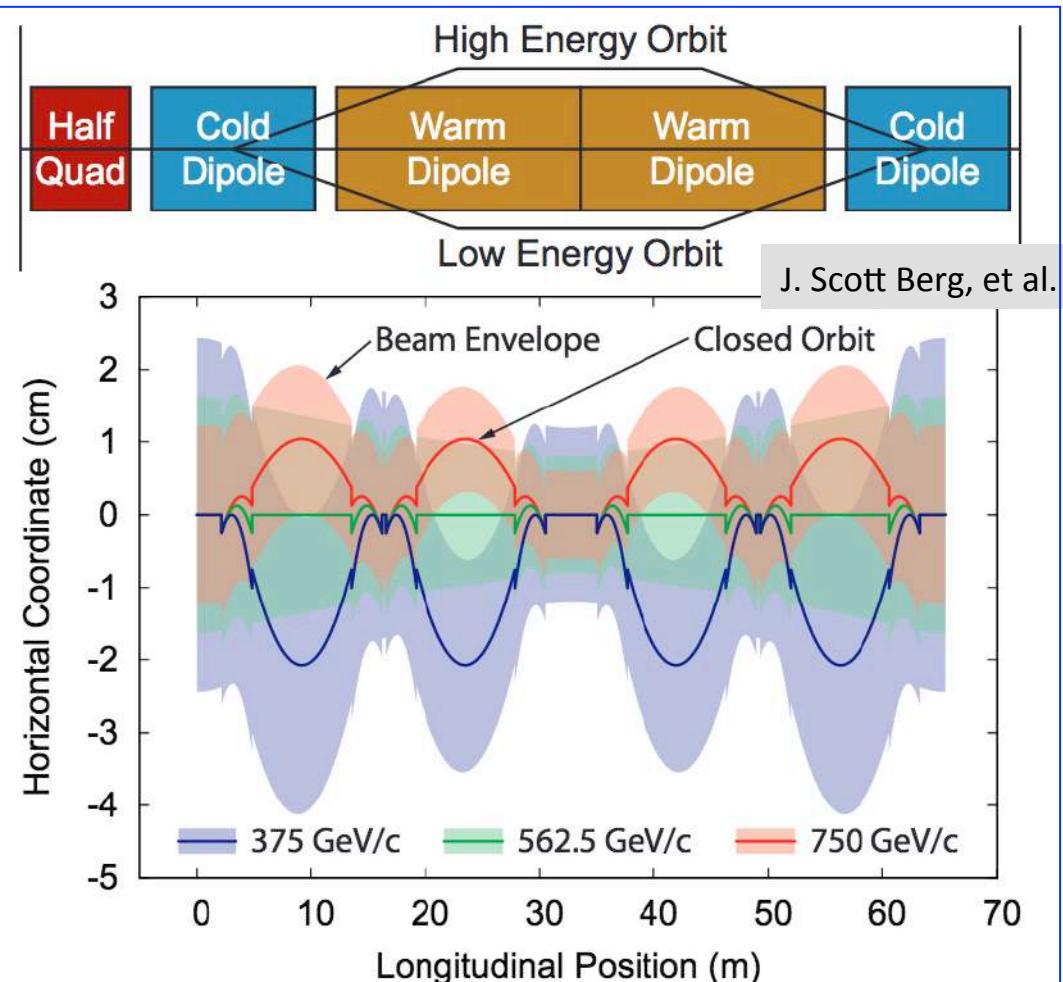
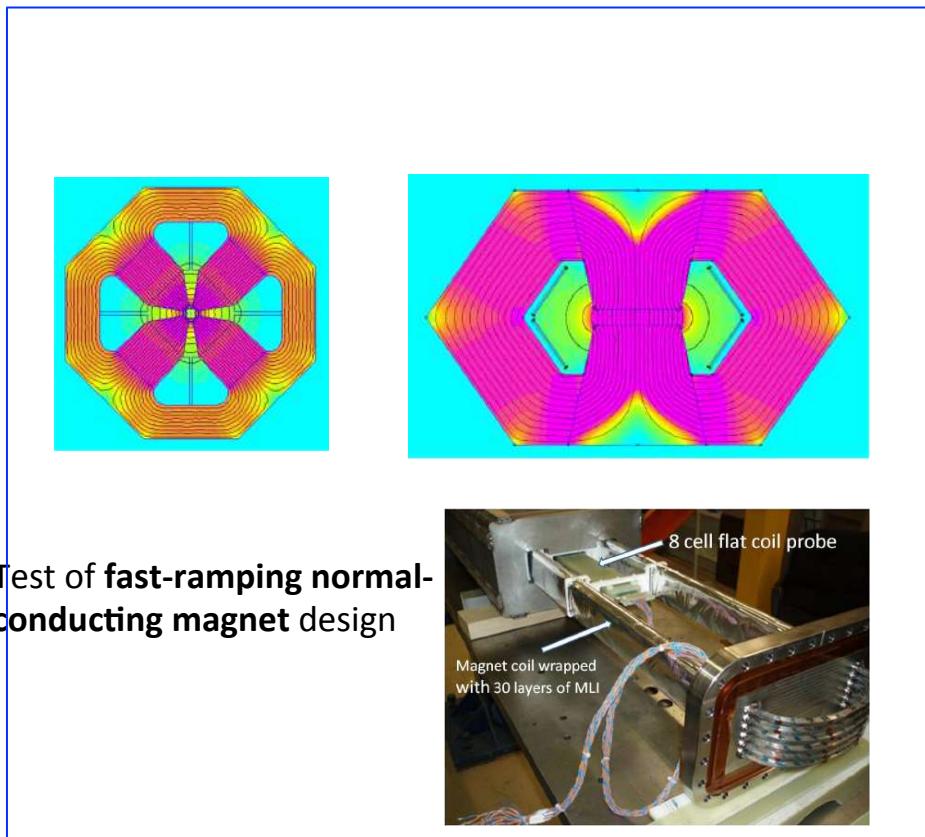
# Acceleration



## Linac

- Recirculating linacs
- Sequence of rings
  - baseline: pulsed synchrotron (RCS)
  - alternative: FFA

Hybrid RCS combines static superconducting magnets and fast-ramping normal-conducting magnets



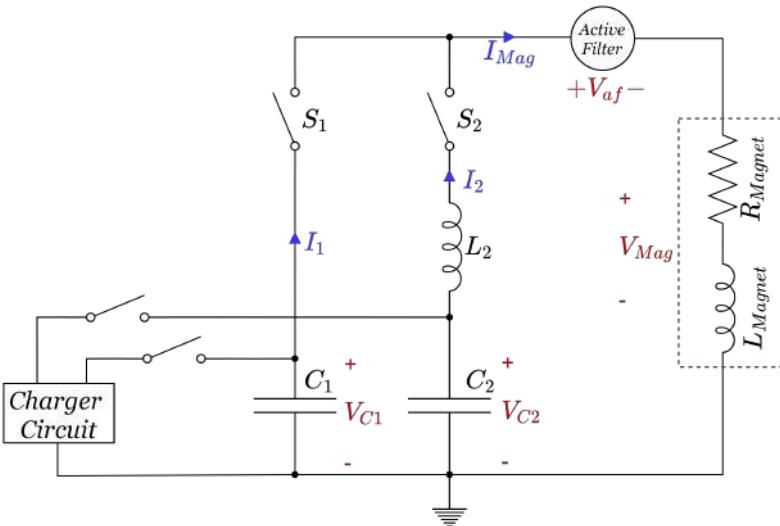
# Power Converter Challenge

Total stored energy in ramped magnets: O(300 MJ), OK

Peak power during ramp (all rings combined) O(200 GW), OK

Average power to magnets 3.3 GW

- novel regime
- requires outstanding energy recovery (98%)
- close interaction between beam physicists and power experts required to limit system cost



**Laser Mega Joule (France)**  
 400 MJ, 480 x 5.76 GW peak power  
 1 pulse every 25min

Synergy with power grid applications



**POPS (CERN)**  
 20MJ, 6 x 10 MW peak power  
 1 pulse every 1.2s;

# Collider Ring Arc Challenge



## Beam loss protection $O(500 \text{ W/m})$

- requires larger aperture for shielding
- will be optimised

30 mm W shielding for 10 years:

- suppresses neutrons very well: DPA  $O(10^{-4})$
  - 30-40 MGy (3, 10 TeV), very local can be reduced by slightly increased shielding
- ⇒ taken care of

## Arc dipoles

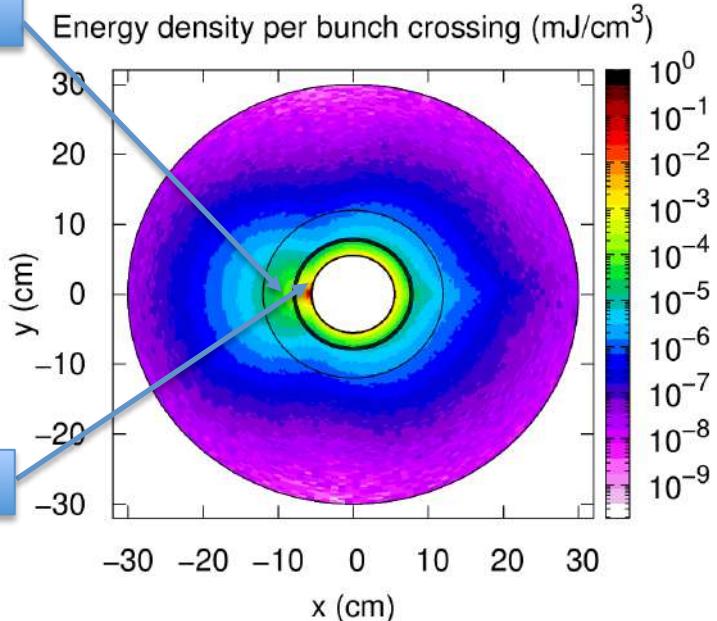
aperture  $O(150 \text{ mm})$ , stress limited

**3 TeV:** 3 km of 11 T

- $\text{Nb}_3\text{Sn}$  dipoles (HL-LHC level)

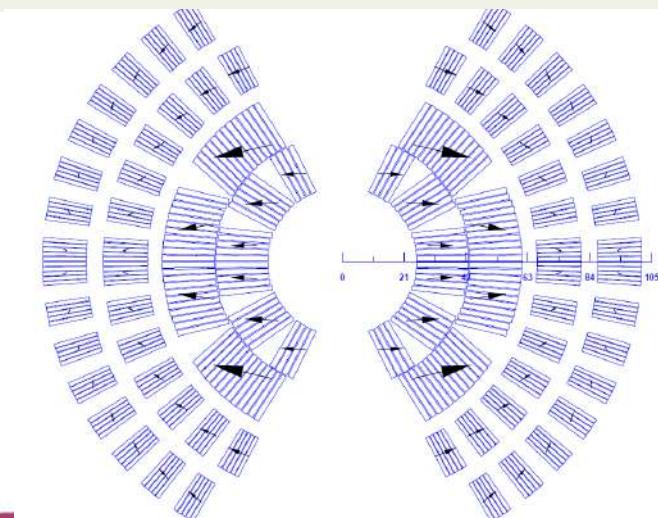
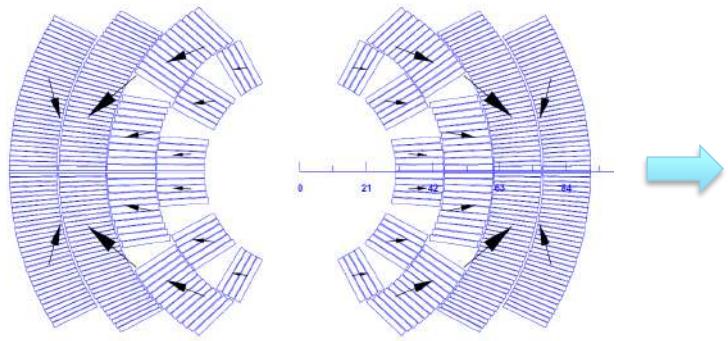
## Shielding

A. Lechner  
D. Calzolari



**10 TeV:** 7 km of 16 T

- stress managed  $\text{Nb}_3\text{Sn}$  dipoles
- or HTS



# Final Focus Magnet Challenge



At 3 TeV:

Up to 12 T in aperture 150 mm

Similar to HL-LHC final focus magnets

Close to state of the art with  $\text{Nb}_3\text{Sn}$

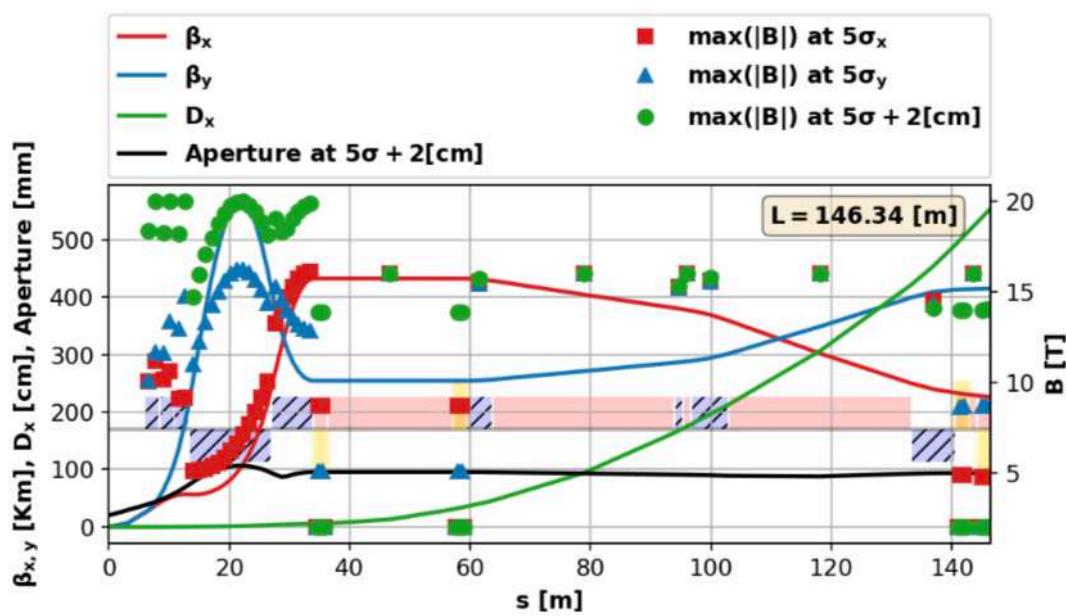
At 10+ TeV:

Aperture of 200+ mm

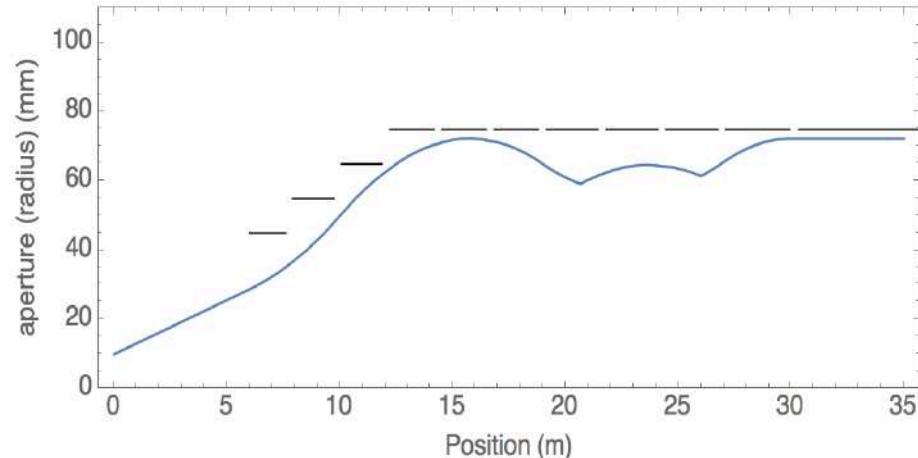
20 T field in aperture

$\Rightarrow$  Stress is significantly higher

$\Rightarrow$  Will need HTS



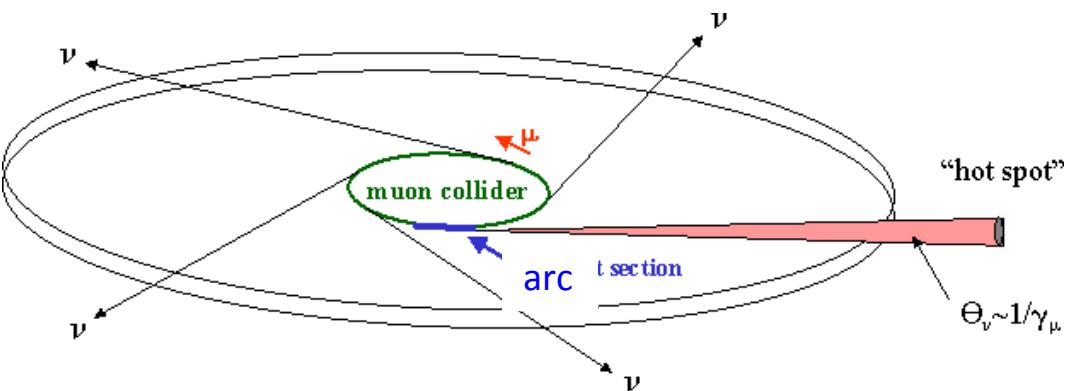
3 TeV FFS Design (MAP)



Parameter	Q1	Q1	Q3	Q4
Aperture (mm)	90	110	130	150
Gradients (T/m)	267	218	-154	-133.5
Peak field (T)	12	12	10+	10+
Dipole field (T)	0	0	2.00	2.00

HTS magnet development is essential for 10 TeV

# Neutrino Flux

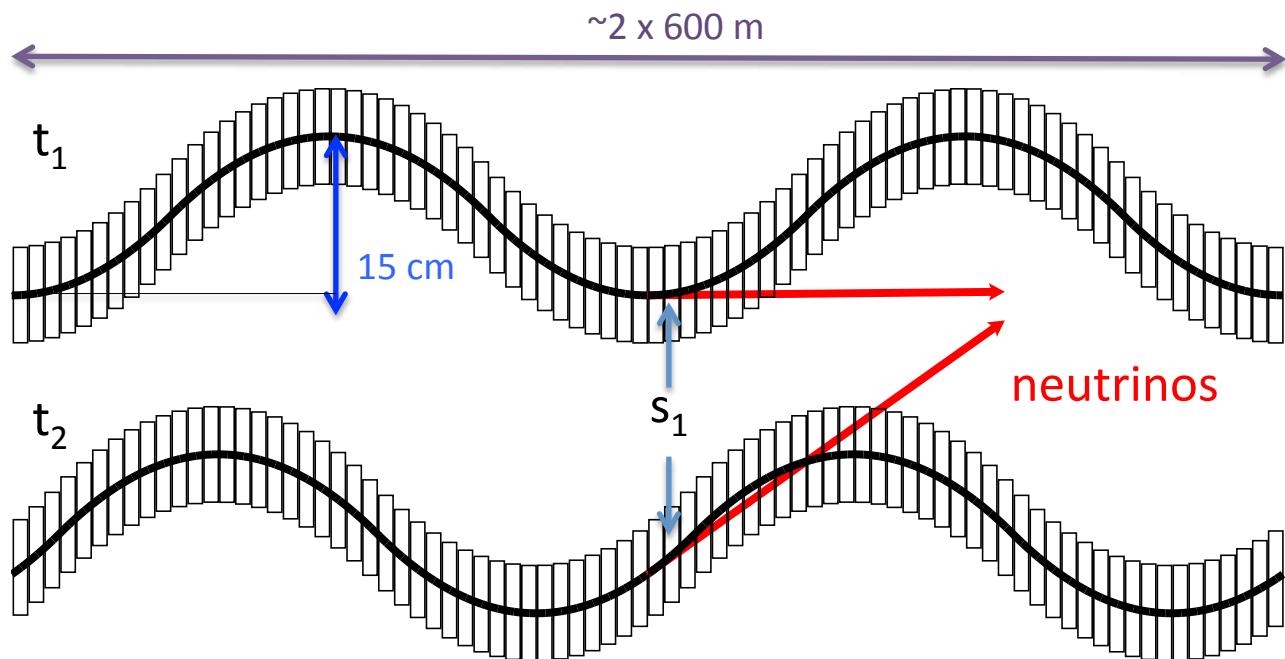


Concentrate neutrino cone from arcs  
can approach legal limits for 14 TeV

**Goal is to reduce to level similar to LHC**

**3 TeV, 200 m deep tunnel is about OK**

**Need mitigation of arcs at 10+ TeV:** idea of Mokhov, Ginneken to move beam in aperture  
Our approach: move collider ring components, e.g. vertical bending with 1% of main field



Opening angle  $\pm 1$  mradian

14 TeV, in 200 m deep tunnel  
comparable to LHC case

**Need to study mover  
system, magnet, connections  
and impact on beam**

**Working on different  
approaches for experimental  
insertion**

# Neutrino Flux Mitigation



Team of RP experts, civil engineers, beam physicist and FLUKA experts

Goal to be **similar to LHC**: i.e. **negligible**, “fully optimised” (10 x better than MAP goal, 100 x better than legal requirements)

- With indirect effects (air, ground water, ...)

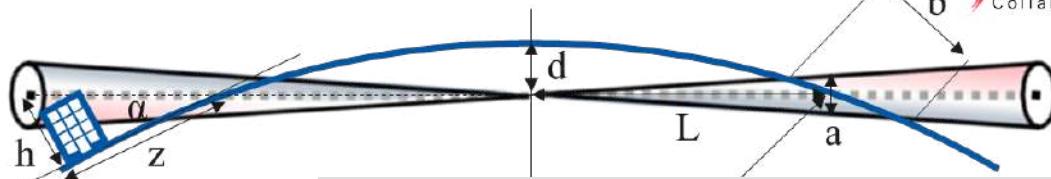
Addressed by:

**Site choice** in direction of experiments

- tools in preparation

**Mechanical mover system** in arcs

- allows 14 TeV in 200 m deep tunnel



C. Ahdida, P. Vojtyla, M. Widorski, H. Vincke

MC simulations  
→ presentation G. Lerner

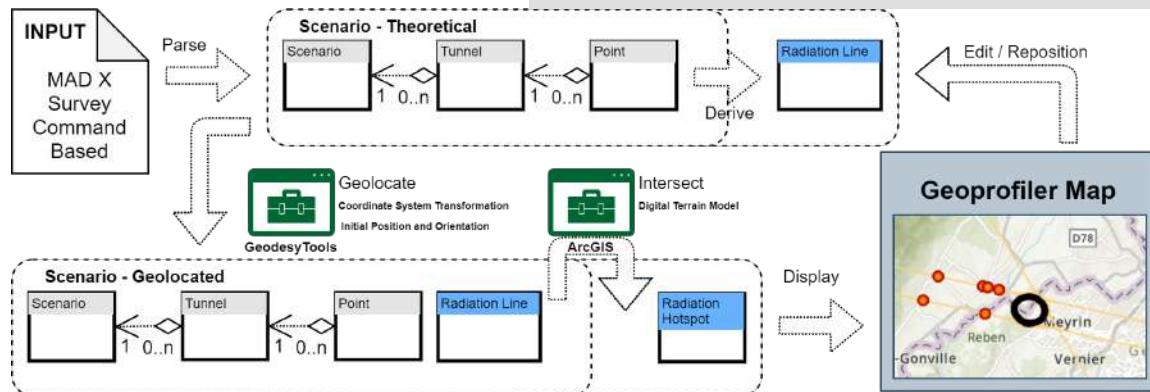
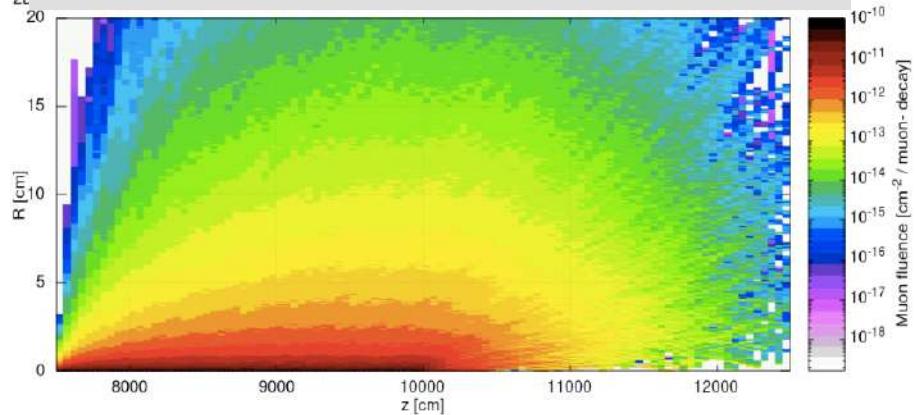
Dose surface map  
→ presentation G. Lacerda



Folding with realistic source term

- Dose assessment
- Sensitivity analysis
- Demonstration of compliance

G. Lerner, D. Calzolari, A. Lechner, C. Ahdida



G. Lacerda, Y. Robert, N. Guilhaudin

Mover system and impact on beam will be addressed in the coming years before end if 2025

# Machine Detector Interface



## Main background sources

- Muon decay products (40,000 muons/m/crossing at 14 TeV)
- Beam-beam background
- Note: background reduces while beam burns off

## Mitigation methods

- masks
- detector granularity
- detector timing
- solenoid field
- event reconstruction strategies
- ...

- Driven by INFN, Padua, (Donatella Lucchesi et al.), contributions from CERN and FNAL
  - interest exists: JAI-Oxford, Sussex, CEA, LIP Lisbon, DESY
- First studies at lower energies (125 GeV and 1.5 TeV are encouraging (D. Lucchesi et al.)
- 10 TeV studies started
- beam-beam started

**Opportunities to help exist**

# Demonstrator Considerations

Muon cooling is the key novel and unique component of muon collider

⇒ need a test in a demonstrator

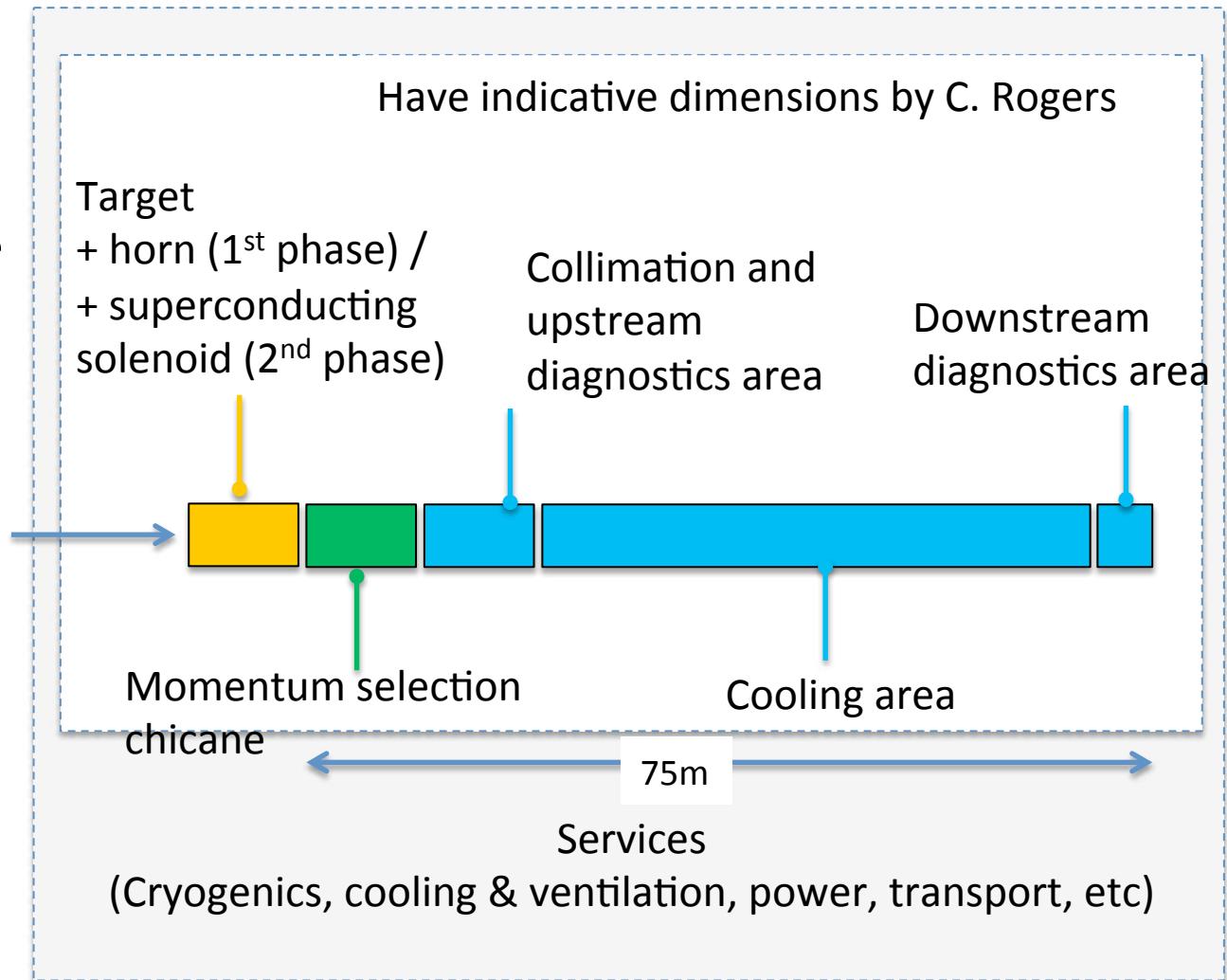
Other technology challenges exist but can be addressed with prototypes (e.g. collider dipoles), also in other locations

Modular approach to demonstrator: start with minimum complex and upgrade as demonstration progresses

Identified components of test facility with approximate dimensions

Will also explore alternative options, if resources permit

- e.g. PIC, parametric ionization cooling

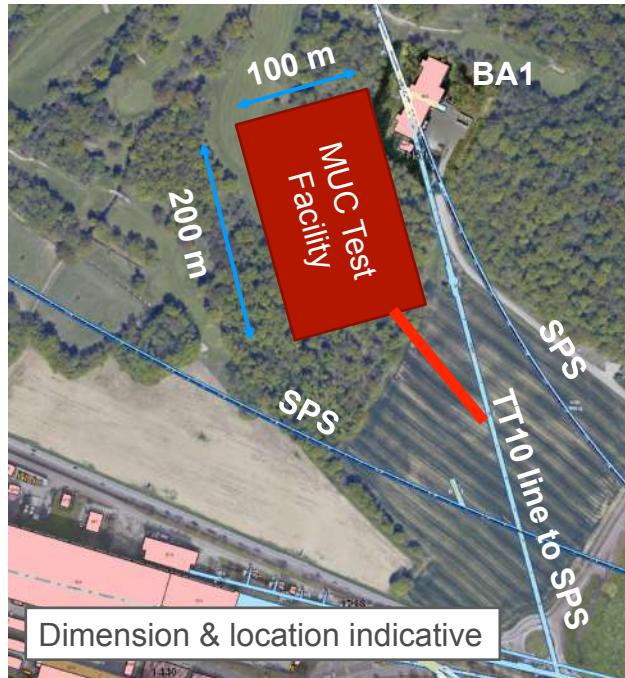


# CERN Site Example

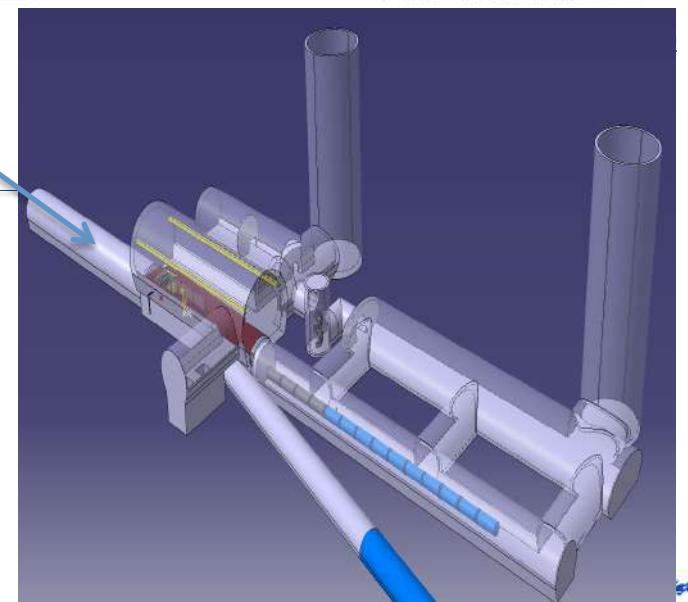
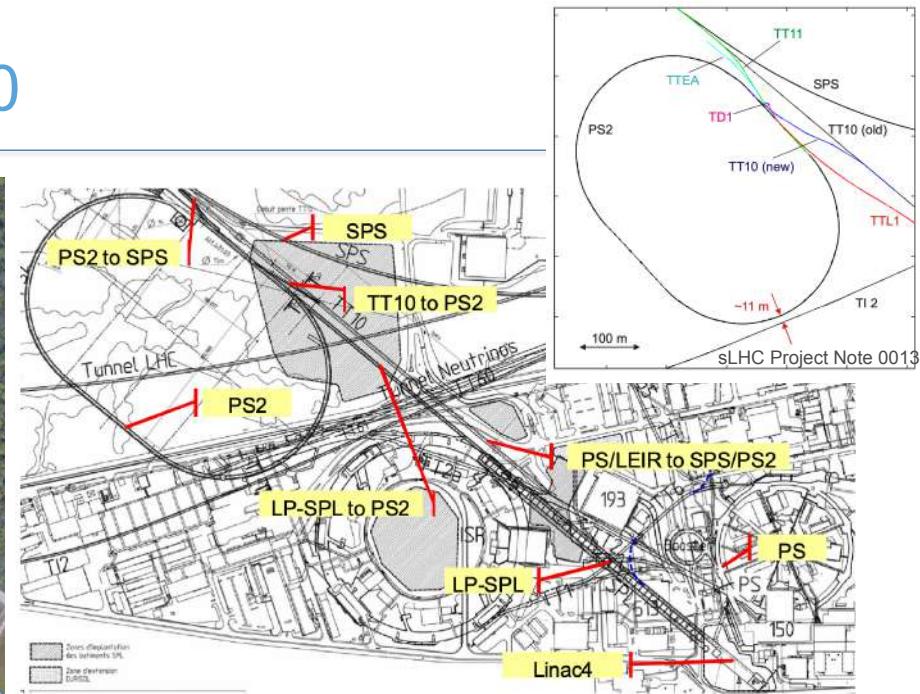


Will consider site proposed by partners and at CERN, but need at least one

## Possibility around TT10



First option considered:  
Could use CERN land close to TT10 and inject beam from PS



- $10^{13}$  26 GeV protons in 7ns, produces a few  $10^{12}$  muons per pulse
- Would be in molasse (no radiation to ground water), could accommodate 4 MW
- Could later upgrade with SPL and accumulator ring to have full power option

# European Accelerator R&D Roadmap



CERN Council charged Laboratory Directors Group (LDG) to develop Roadmap

- reviewed by SPC
- agreed by Council December 2021

CERN Council charged LDG to develop implementation plan by March 2022

Roadmap identifies muon collider challenges and two R&D scenarios to address them

- An **aspirational scenario**
- A **minimal scenario**

Scenario	FTEy	kCHF
Aspirational	445.9	11,875
Minimal	193	2,445

<http://arxiv.org/abs/2201.07895>

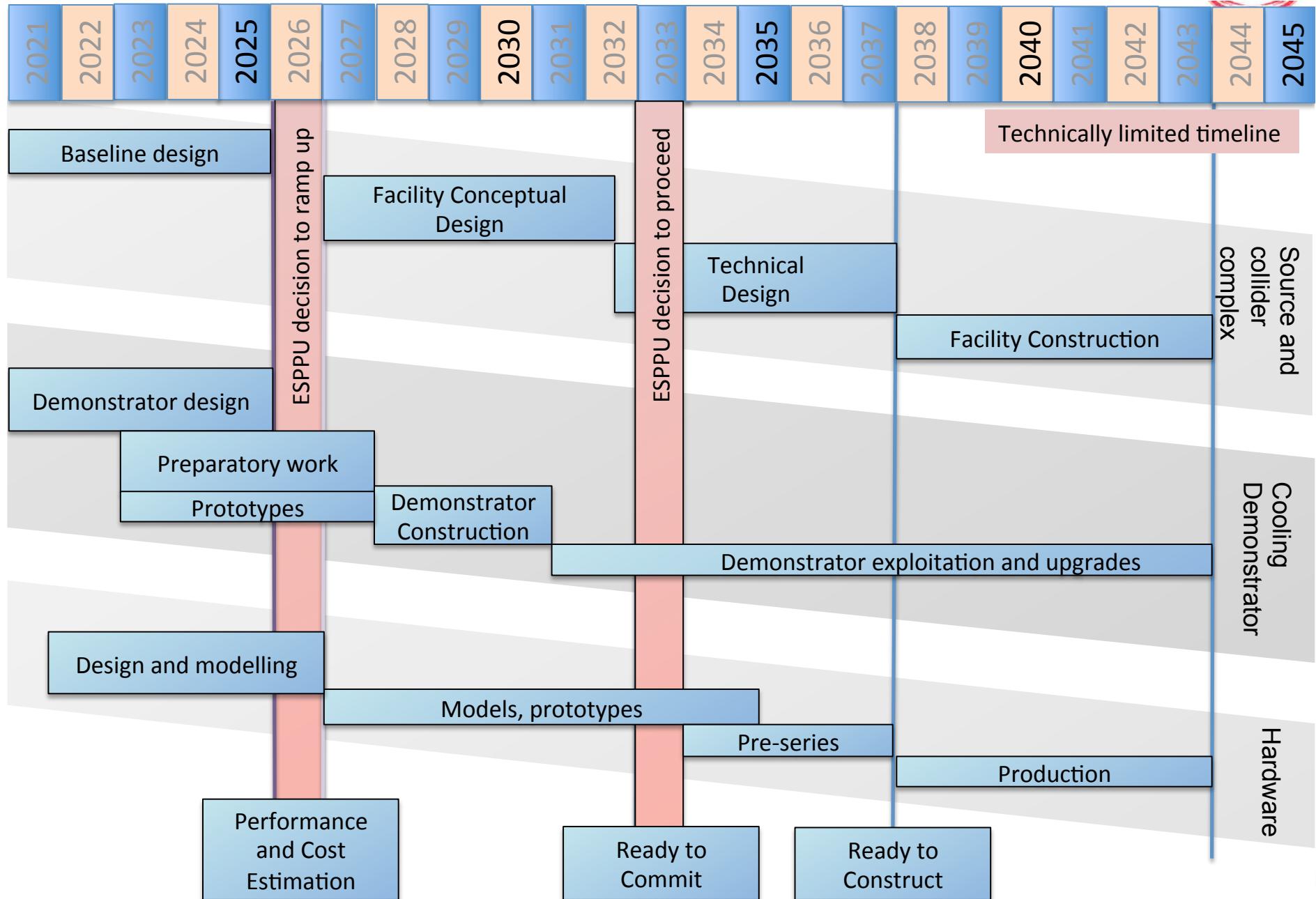
Aspirational scenario = 10 years of MAP (up to 45 FTE)

Label	Begin	End	Description	Aspirational		Minimal	
				[FTEy]	[kCHF]	[FTEy]	[kCHF]
MC.SITE	2021	2025	Site and layout	15.5	300	13.5	300
MC.NF	2022	2026	Neutrino flux mitigation system	22.5	250	0	0
MC.MDI	2021	2025	Machine-detector interface	15	0	15	0
MC.ACC.CR	2022	2025	Collider ring	10	0	10	0
MC.ACC.HE	2022	2025	High-energy complex	11	0	7.5	0
MC.ACC.MC	2021	2025	Muon cooling systems	47	0	22	0
MC.ACC.P	2022	2026	Proton complex	26	0	3.5	0
MC.ACC.COLL	2022	2025	Collective effects across complex	18.2	0	18.2	0
MC.ACC.ALTER	2022	2025	High-energy alternatives	11.7	0	0	0
MC.HFM.HE	2022	2025	High-field magnets	6.5	0	6.5	0
MC.HFM.SOL	2022	2026	High-field solenoids	76	2700	29	0
MC.FR	2021	2026	Fast-ramping magnet system	27.5	1020	22.5	520
MC.RF.HE	2021	2026	High Energy complex RF	10.6	0	7.6	0
MC.RF.MC	2022	2026	Muon cooling RF	13.6	0	7	0
MC.RF.TS	2024	2026	RF test stand + test cavities	10	3300	0	0
MC.MOD	2022	2026	Muon cooling test module	17.7	400	4.9	100
MC.DEM	2022	2026	Cooling demonstrator design	34.1	1250	3.8	250
MC.TAR	2022	2026	Target system	60	1405	9	25
MC.INT	2022	2026	Coordination and integration	13	1250	13	1250
			Sum	445.9	11875	193	2445

**Table 5.5:** The resource requirements for the two scenarios. The personnel estimate is given in full-time equivalent years and the material in kCHF. It should be noted that the personnel contains a significant number of PhD students. Material budgets do not include budget for travel, personal IT equipment and similar costs. Colours are included for comparison with the resource profile Fig. 5.7.

# Aspirational Timeline

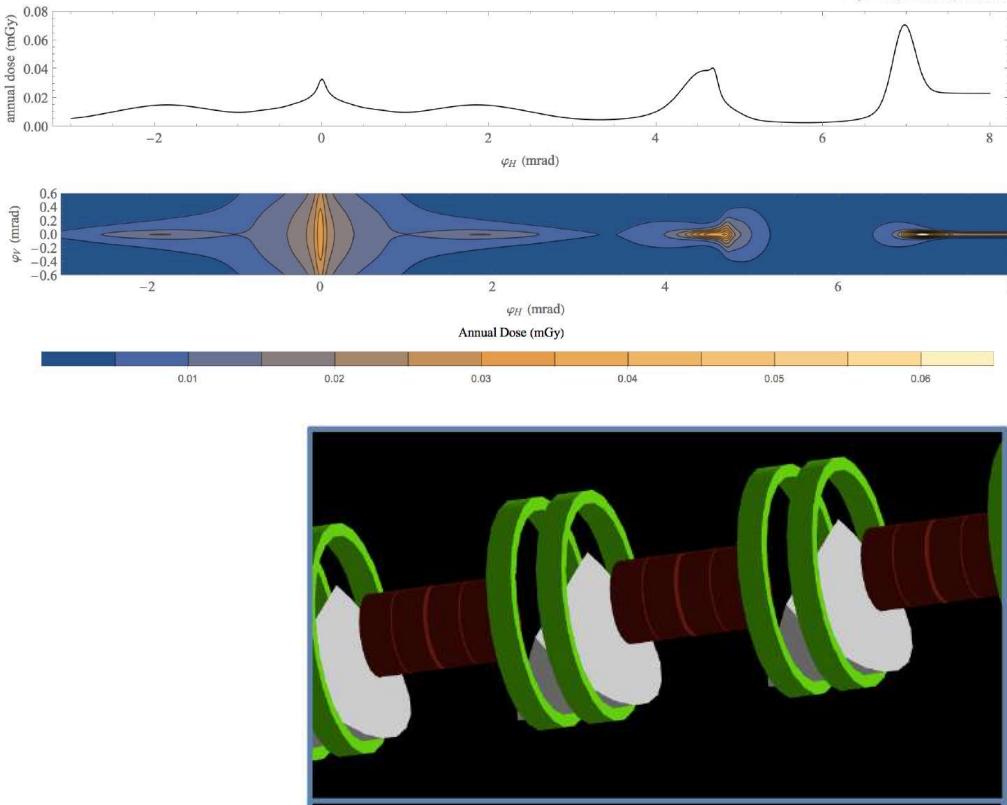
in case muon collider is next project after HL-LHC



# Ongoing Activities



- Identification of challenges and workplan
  - many contributors
- Radiation protection (CERN, FNAL)
- MDI (INFN, Padua, CERN, FNAL)
- Collider ring (CERN)
- Muon cooling system (RAL, CERN)
- Target (CERN)
- RF team is very active (CEA, CERN, Rostock)
- Power converter for fast-ramping magnets
- Demonstrator site and scope identification (all)
- A number of key “small” contributions
  - e.g. RF calculation (Rostock), solenoid stress estimate (KIT), lattice design (JAI students), RF test stand estimates (CEA), ...



**People started the journey**

# Interest (in part some work done)

- Pulsed synchrotrons (CEA-IRFU)
  - (Recirculating) linacs (IJCLab, BNL, CERN software)
  - Proton complex (ESS, Uppsala)
  - Alternatives (e.g. FFA, final muon cooling) (UK)
- Preparing an EU Design Study
- Target and target area (ESS, Uppsala, RAL, Warwick, ENEA)
  - High-field collider ring magnets (KEK-CERN, Tokyo)
  - Fast-ramping magnet systems (LNCMI/EMFL/Dresden, Darmstadt, KIT, PSI, EPFL, CERN)
  - High-field solenoids (CEA-IRFU, KIT, looking for more)
- Superconducting high-energy RF (Rostock, INFN-Milano)
  - Normal-conducting muon cooling RF (CEA-IRFU, RAL, Cockcroft Institute, Lancaster, Strathclyde, Daresbury, INFN-Catania))
- Integration challenges of muon cooling cell (INFN, ...)
  - Test facility (CERN, all)
  - Ongoing activities still need more resources

**People getting ready to start the journey**

# Thanks for the Roadmap

**Muon Beam Panel:** Daniel Schulte (CERN, chair), Mark Palmer (BNL, co-chair), Tabea Arndt (KIT), Antoine Chance (CEA/IRFU) Jean-Pierre Delahaye (retired), Angeles Faus-Golfe (IN2P3/IJClab), Simone Gilardoni (CERN), Philippe Lebrun (European Scientific Institute), Ken Long (Imperial College London), Elias Metral (CERN), Nadia Pastrone (INFN-Torino), Lionel Quettier (CEA/IRFU), Magnet Panel link, Tor Raubenheimer (SLAC), Chris Rogers (STFC-RAL), Mike Seidel (EPFL and PSI), Diktys Stratakis (FNAL), Akira Yamamoto (KEK and CERN) **Contributors:** Alexej Grudiev (CERN), Roberto Losito (CERN), Donatella Lucchesi (INFN)

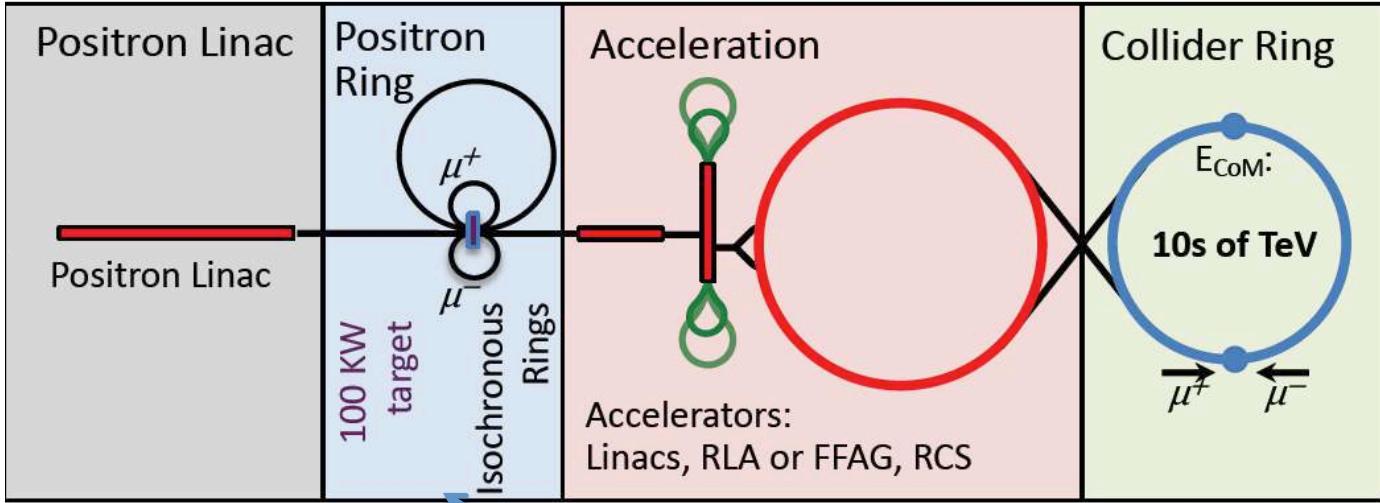
**Community conveners:** *Radio-Frequency (RF):* Alexej Grudiev (CERN), Jean-Pierre Delahaye (CERN retiree), Derun Li (LBNL), Akira Yamamoto (KEK). *Magnets:* Lionel Quettier (CEA), Toru Ogitsu (KEK)<sup>†</sup>, Soren Prestemon (LBNL), Sasha Zlobin (FNAL), Emanuela Barzi (FNAL). *High-Energy Complex (HEC):* Antoine Chance (CEA), J. Scott Berg (BNL), Alex Bogacz (JLAB), Christian Carli (CERN), Angeles Faus-Golfe (IJCLab), Eliana Gianfelice-Wendt (FNAL), Shinji Machida (RAL). *Muon Production and Cooling (MPC):* Chris Rogers (RAL), Marco Calviani (CERN), Chris Densham (RAL), Diktys Stratakis (FNAL), Akira Sato (Osaka University), Katsuya Yonehara (FNAL). *Proton Complex (PC):* Simone Gilardoni (CERN), Hannes Bartosik (CERN), Frank Gerigk (CERN), Natalia Milas (ESS). *Beam Dynamics (BD):* Elias Metral (CERN), Tor Raubenheimer (SLAC and Stanford University), Rob Ryne (LBNL). *Radiation Protection (RP):* Claudia Ahdida (CERN). *Parameters, Power and Cost (PPC):* Daniel Schulte (CERN), Mark Palmer (BNL), Jean-Pierre Delahaye (CERN retiree), Philippe Lebrun (CERN retiree and ESI), Mike Seidel (PSI), Vladimir Shiltsev (FNAL), Jingyu Tang (IHEP), Akira Yamamoto (KEK). *Machine Detector Interface (MDI):* Donatella Lucchesi (University of Padova), Christian Carli (CERN), Anton Lechner (CERN), Nicolai Mokhov (FNAL), Nadia Pastrone (INFN), Sergo R Jindariani (FNAL). *Synergy:* Kenneth Long (Imperial College), Roger Ruber (Uppsala University), Koichiro Shimomura (KEK). *Test Facility (TF):* Roberto Losito (CERN), Alan Bross (FNAL), Tord Ekelof (ESS, Uppsala University).

**And the participants to the community meetings and the study**

# Alternatives: The LEMMA Scheme

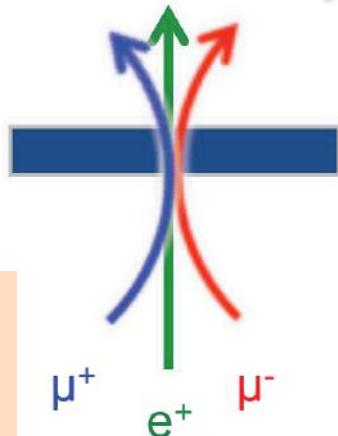
LEMMA scheme (INFN)

P. Raimondi et al.



45 GeV positrons to produce muon pairs  
Accumulate muons from several passages

$$e^+ e^- \rightarrow \mu^+ \mu^-$$



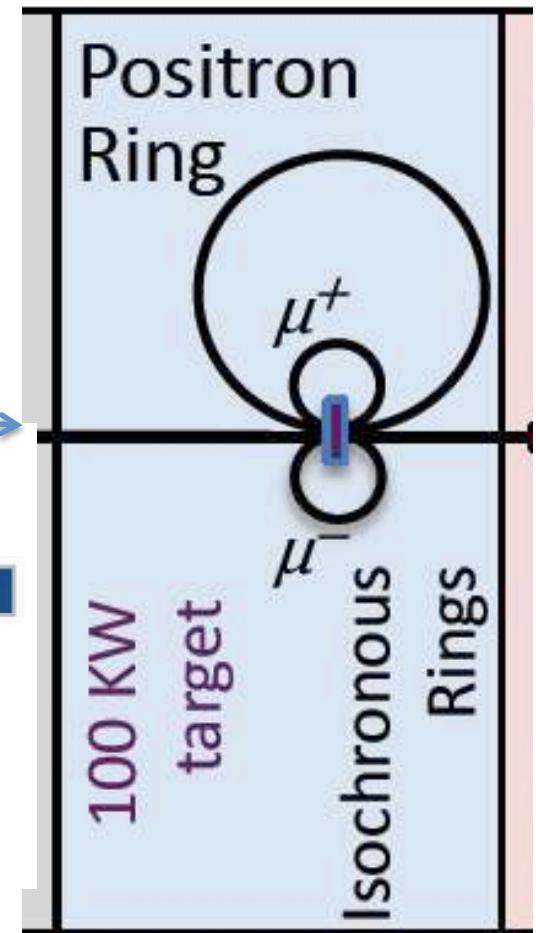
**Excellent idea, but nature is cruel**

Detailed estimates of fundamental limits show that we require a very large positron bunch charge to reach the same luminosity as the proton-based scheme

⇒ Need same game changing invention

Note: New proposal by C. Curatolo and L. Serafini needs to be looked at

- Uses Bethe-Heitler production with electrons



# Conclusion

- Muon colliders are a unique opportunity for a high-energy, high-luminosity lepton collider
  - high luminosity to beam power ratio
  - cost efficiency to be assessed
- Two different options considered
  - 3 TeV collider that can start construction in less than 20 years
  - 10 TeV collider that uses advanced technologies
- Not as mature as ILC or CLIC
  - have to address **important R&D** items
  - but **no showstopper** identified
- Aim to develop concept to a **maturity level** that allows to make **informed choices by the next ESPPU** and other strategy processes
  - Baseline design
  - R&D and demonstration programme
- An important opportunity that we should not miss
- <http://muoncollider.web.cern.ch>

Many thanks to the Muon Beam Panel, the collaboration, the MAP study, the MICE collaboration, and many others

# Reserve

# Muon Collider Luminosity Scaling



Fundamental limitation

Assumes no emittance growth after source and no technical limitation

Applies to MAP and LEMMA scheme

$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_{\delta} \frac{N_0}{\epsilon \epsilon_L} f_r N_0 \gamma$$

High energy → High field in collider ring  
= small ring  
= many collisions

Large energy acceptance  
= short bunch  
= small betafunction

Dense beam → High beam power

Note: emittances are normalised

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# Muon Collider Luminosity Scaling



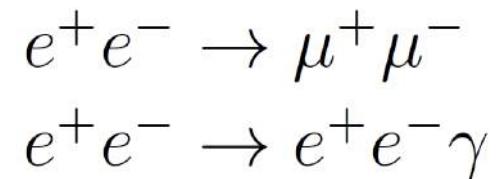
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Same for MAP and LEMMA

O(1%) of proton scheme  
= 100 MW of positrons lost



Bremsstrahlung O( $10^5$ ) times  
more likely than pair production  
 $O(150\text{mb})$ ,  $E_\gamma \geq 0.01 E_p$   
 $O(60\text{mb})$ ,  $E_\gamma \geq 0.1 E_p$

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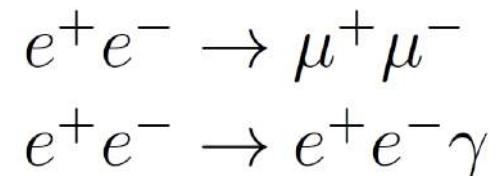
Same for MAP and LEMMA

Each passage in target increases emittance (multiple scattering)  
⇒ Need to produce enough muons per passage for high  $N/\epsilon$

Example to reach luminosity is

- 3 mm BE target, 0.86 mm betafunction (optimum)
- $3 \times 10^{15}$  positrons per bunch (22 MJ)
  - 60 kJ lost in target, temperature jump of MK
- at least 100 bunches per pulse (2 GJ)
  - (only 1% is lost)

O(1%) of proton scheme  
= 100 MW of positrons lost



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Note: Additional beam combination schemes can reduce positron bunch charge but increase energy in pulse

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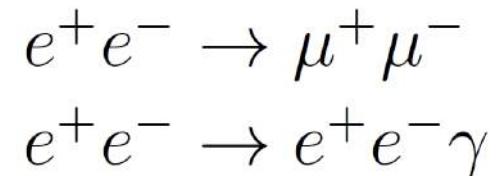
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 $O(60\text{mb})$ ,  $E_\gamma \geq 0.1 E_p$

Unfortunately, seems too hard from fundamental physics  
Need a new game-changing invention

# Muon Collider Luminosity Scaling



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Applies to MAP and LEMMA scheme

$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_\delta \frac{N_0}{\epsilon \epsilon_L} f_r N_0 \gamma$$

Assume 3 mm thick Be target, 0.86 mm beta

⇒ 0.6 nm emittance growth per muon beam passage through target (optimum case)

⇒ Need bunches with **3 x 10<sup>15</sup> positrons** (=22 MJ) to obtain required

⇒ Positron beam energy **2 GJ/burst**, 5 burst per second

⇒ Energy deposition in target **60 kJ per pulse** (minimum ionisation) 4.5 MK temperature rise per bunch (linear approximation)

⇒ Extremely challenging, not sure even a fluid target can do this

LEMMA scheme needs O(0.7 mJ) positrons lost per produced muon pair  
⇒ 100 MW loss yield  $1.4 \times 10^{11} \text{ s}^{-1}$  muon pairs

- (proton case:  $1 \times 10^{13} \text{ s}^{-1}$ )

⇒ Need 70 times denser beam for same luminosity

⇒ Lose  $1.4 \times 10^{16}$  positrons per second

# Note: Stacking

$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_\delta \frac{N_0}{\epsilon \epsilon_L} f_r N_0 \gamma$$

stacking in longitudinal plane does not increase luminosity

bunch length and beta-function increase with the charge

Stacking in transverse plane can help because

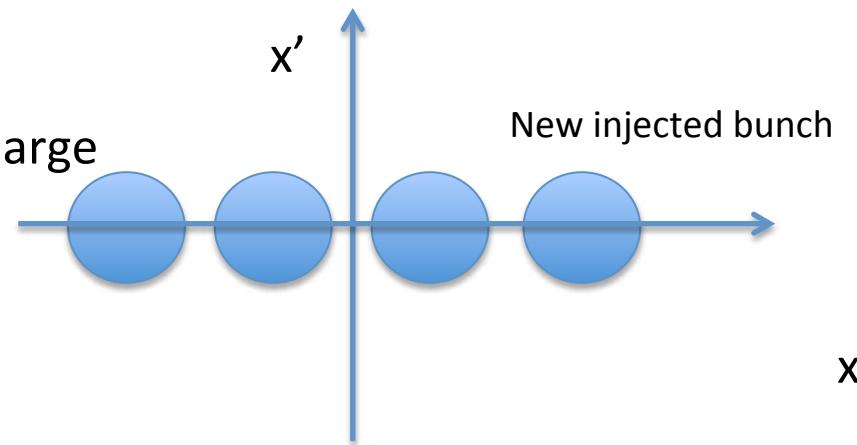
$$\epsilon = \sqrt{\epsilon_x \epsilon_y}$$

stacking  $m^2$  bunches leads to

$$N = m^2 N_1 \quad \epsilon = m \epsilon_1$$

and the luminosity scales as

$$\frac{N}{\epsilon} = m \frac{N_1}{\epsilon_1}$$



$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_\delta \frac{N}{\epsilon \epsilon_L} f_r N \gamma$$

$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_\delta m \frac{N_0}{\epsilon_0 \epsilon_{L,0}} f_{r,0} N_0 \gamma$$

$$\mathcal{L} \propto \gamma \langle B \rangle \sigma_\delta \sqrt{f_{r,0} \tau \gamma} \frac{N_0}{\epsilon_0 \epsilon_{L,0}} f_{r,0} N_0 \gamma$$

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High energy → High field in collider ring  
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Large energy acceptance  
= short bunch  
= small betafunction

Dense beam → High beam power

Note: emittances are normalised

# Some Comments

F. Zimmermann 2018 J. Phys.: Conf. Ser. 1067 022017 claims

$$\begin{aligned}
 L &\approx f_{\text{rev}} \dot{N}_\mu \frac{\dot{N}_\mu}{\varepsilon_N} \frac{1}{3^6} \gamma \tau^2 \frac{1}{4\pi\beta^*} \\
 &= \frac{1}{3^6} \left\{ \left( \frac{eF_{\text{dip}}}{2\pi m_\mu} \right)^3 \frac{\tau_0^2}{4\pi c^2} \right\} [B^3 C^2] \left[ \dot{N}_\mu \frac{\dot{N}_\mu}{\varepsilon_N} \right] \frac{1}{\beta^*}
 \end{aligned}$$

$\mathcal{L} \propto \frac{(f_r N)^2}{\epsilon}$

The paper assumes that muons can be stacked but ignores the associated emittance growth  
 This is wrong, with these assumption LEMMA would be viable

New proposal by C.  
 Curatolo and L.  
 Serafini needs to be  
 looked at  
 Uses Bethe-Heitler  
 production with  
 electrons

scheme	$p\text{-}\gamma$	G.-F. $\mu$	$e^+$	G.-F. $e^+$
	base	LHC/FCC-hh	FCC-ee	FCC
rate $\dot{N}_\mu$ [GHz]	1	400	0.003	100
$\mu/\text{pulse}$ [ $10^4$ ]	0.01	4	0.2	6,000
p. spacing [ns]	100	100	15	15
energy [GeV]	2.5	0.1	22	22
rms en. spread	3%	10%	10%	10%
n. emit. [ $\mu\text{m}$ ]	7	2000	0.04	0.04
$\dot{N}_\mu/\varepsilon_N$ [ $10^{15} \text{ m}^{-1}\text{s}^{-1}$ ]	0.1	0.2	0.1	3,000

the LEMMA scheme

at 14 TeV:  
 9 GW beam power

even 30 times more  
 beam particles

# Physics at Muon Collider

Muon Collider can be the game changer

## Muon collider physics potential

A high-energy muon collider is simply a **dream machine**: allows to probe unprecedented energy scales, exploring many different directions at once!

### Direct searches

Pair production,  
Resonances, VBF,  
Dark Matter, ...

### High-rate measurements

Single Higgs,  
self coupling, rare and  
exotic Higgs decays,  
top quarks, ...

### High-energy probes

Di-boson, di-fermion,  
tri-boson, EFT,  
compositeness, ...

### Muon physics

Lepton Flavor  
Universality,  $b \rightarrow s\mu\mu$ ,  
muon g-2, ...

- Theory input needed: define energy, luminosity and detector performance goals — physics potential of a multi-TeV muon collider
- Great interest in the theory community:

1807.04743 2005.10289 2008.12204 2012.11555 2102.11292 2104.05720  
1901.06150 2006.16277 2009.11287 2101.10334 2103.01617 etc ...  
2003.13628 2007.14300 2012.02769 2102.08386 2103.14043



D. Buttazzo

## The Muon Smasher's Guide

P. Maede

### A Muon Collider is great!

$\kappa_0$ fit	HL-LHC	LHeC	HE-LHC	ILC	CLIC	CEPC	FCC-ee	FCC-ee/ eh/hh	$\mu^+\mu^-$ 10000
		S2	S2'	250 500 1000	380 1500 3000		240 365		
$\kappa_W$ [%]	1.7	0.75	1.4 0.98	1.8 0.29 0.24	0.86 0.16 0.11	1.3	1.3 0.43	0.14	0.06
$\kappa_Z$ [%]	1.5	1.2	1.3 0.9	0.29 0.23 0.22	0.5 0.26 0.23	0.14	0.20 0.17	0.12	0.23
$\kappa_g$ [%]	2.3	3.6	1.9 1.2	2.3 0.97 0.66	2.5 1.3 0.9	1.5	1.7 1.0	0.49	0.15
$\kappa_\gamma$ [%]	1.9	7.6	1.6 1.2	6.7 3.4 1.9	98* 5.0 2.2	3.7	4.7 3.9	0.29	0.64
$\kappa_{Z\gamma}$ [%]	10.	—	5.7 3.8	99* 86* 85*	120* 15 6.9	8.2	81* 75*	0.69	1.0
$\kappa_e$ [%]	—	4.1	— —	2.5 1.3 0.9	4.3 1.8 1.4	2.2	1.8 1.3	0.95	0.89
$\kappa_t$ [%]	3.3	—	2.8 1.7	— 6.9 1.6	— — 2.7	—	— —	1.0	6.0
$\kappa_b$ [%]	3.6	2.1	3.2 2.3	1.8 0.58 0.48	1.9 0.46 0.37	1.2	1.3 0.67	0.43	0.16
$\kappa_\mu$ [%]	4.6	—	2.5 1.7	15 9.4 6.2	320* 13 5.8	8.9	10 8.9	0.41	2.0
$\kappa_\tau$ [%]	1.9	3.3	1.5 1.1	1.9 0.70 0.57	3.0 1.3 0.88	1.3	1.4 0.73	0.44	0.31

7

P. Maede

### Di-Higgs too!

## Double Higgs production

- Reach on Higgs trilinear coupling:  $hh \rightarrow 4b$

B. Franceschini, Wulzer 2012.11555

Costantini et al. 2005.10289

Han et al. 2008.12204

E [TeV]	$\mathcal{L} [\text{ab}^{-1}]$	$N_{\text{rec}}$	$\delta\sigma \sim N_{\text{rec}}^{-1/2}$	$\delta\kappa_3$
3	5	170	~ 7.5%	~ 10%
10	10	620	~ 4%	~ 5%
14	20	1340	~ 2.7%	~ 3.5%
30	90	6'300	~ 1.2%	~ 1.5%

# Challenges and Status

**FNAL**

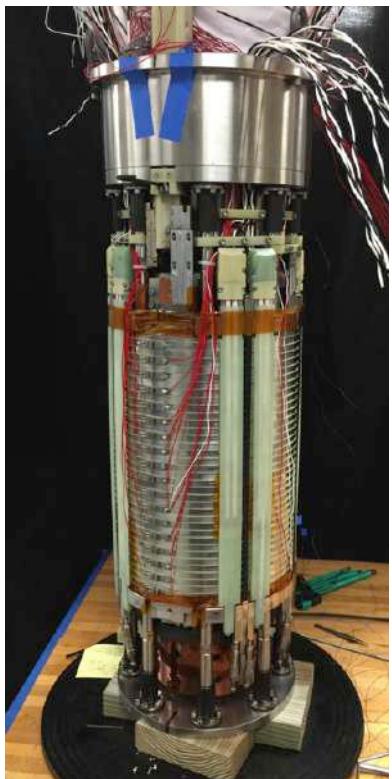
12 T/s HTS

0.6 T max

now 290 T/s



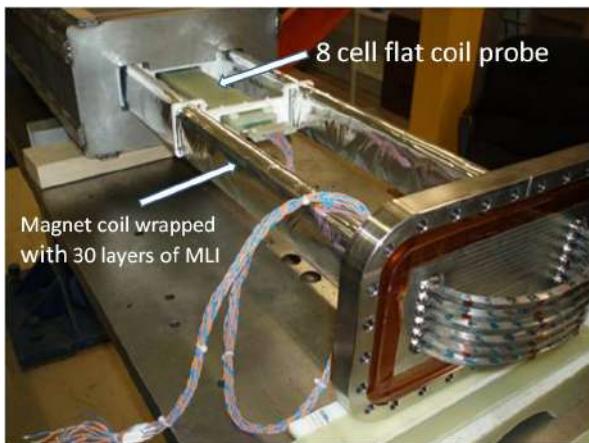
Test of **fast-ramping**  
**normal-conducting**  
**magnet** design



**NHFML**

32 T solenoid  
with HTS

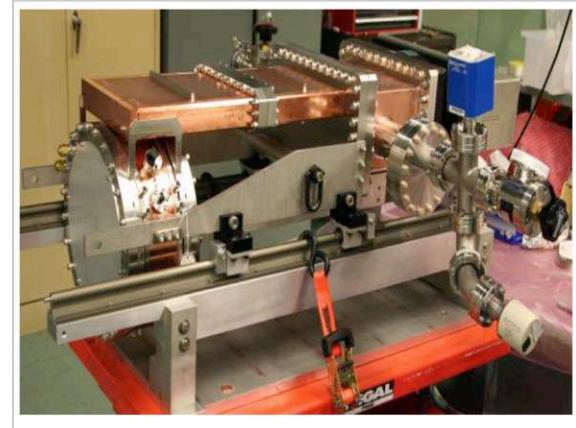
Planned efforts  
to push even  
further



**MuCool:** >50 MV/m in 5 T field

Two solutions

- Copper cavities filled with hydrogen
- Be end caps



**MICE (UK)** Muon cooling principle



# Selected Recent Progress

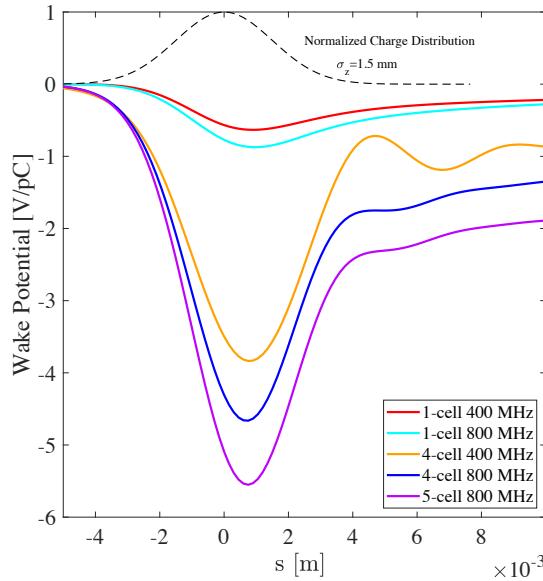


D. Aguglia  
F. Boattini  
G. Brauchli

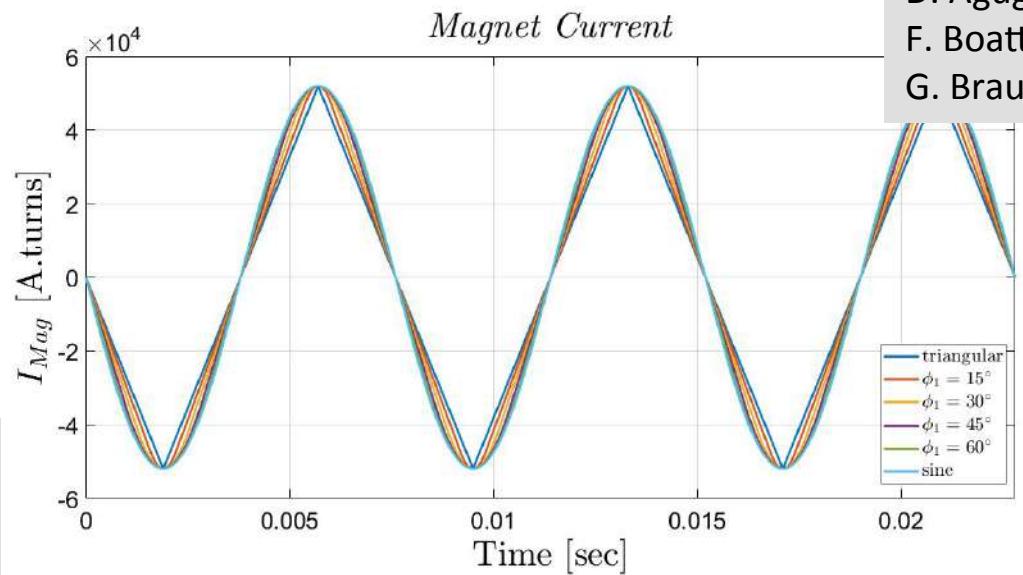
## Ramping magnet challenge

At 14 TeV, energy in field is  $O(200 \text{ MJ})$   
ramped  $5 \times 2$  times per second

Need to recover energy pulse to pulse  
Started to develop **powering scheme**  
with energy recovery



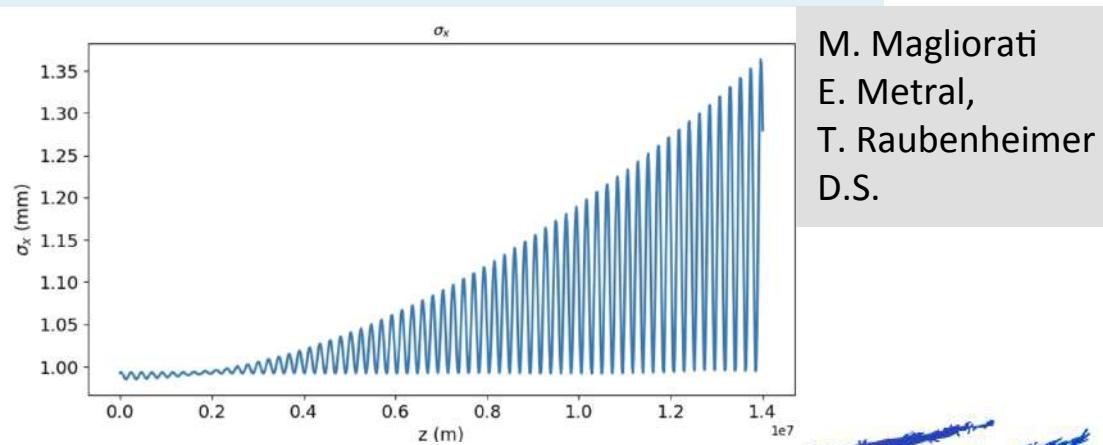
University of Rostock  
S. Zadeh  
U. van Rienen



## RF challenge (also for FFA):

High efficiency for power consumption  
High-charge ( $10 \times \text{HL-LHC}$ ), short, single-bunch beam  
Maintain small longitudinal emittance  
Studies on cavity wakefields and longitudinal dynamics started

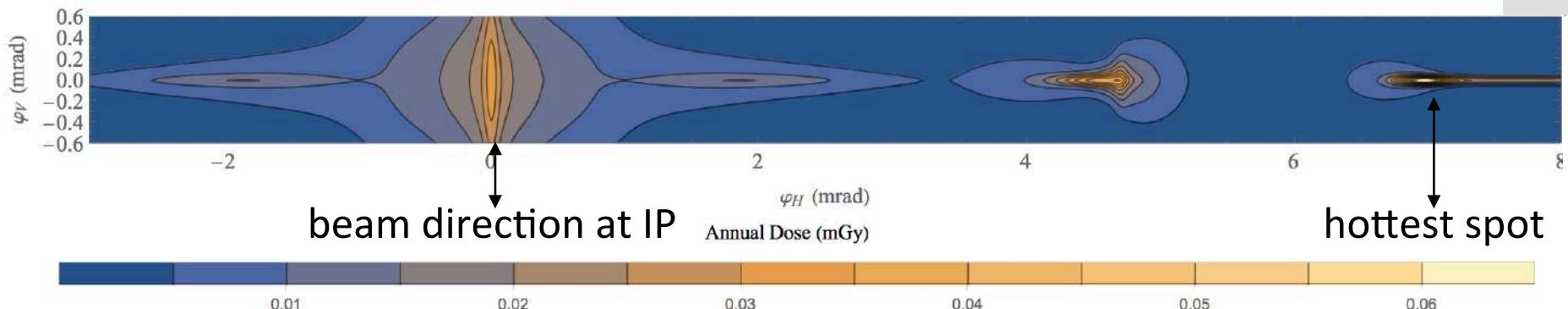
**Collective effects** might be a bottleneck  
Revisiting for higher energies  
Need to develop tools for collective effects in matter



M. Magliorati  
E. Metral,  
T. Raubenheimer  
D.S.



C. Carli



### Collider Ring Lattice Design:

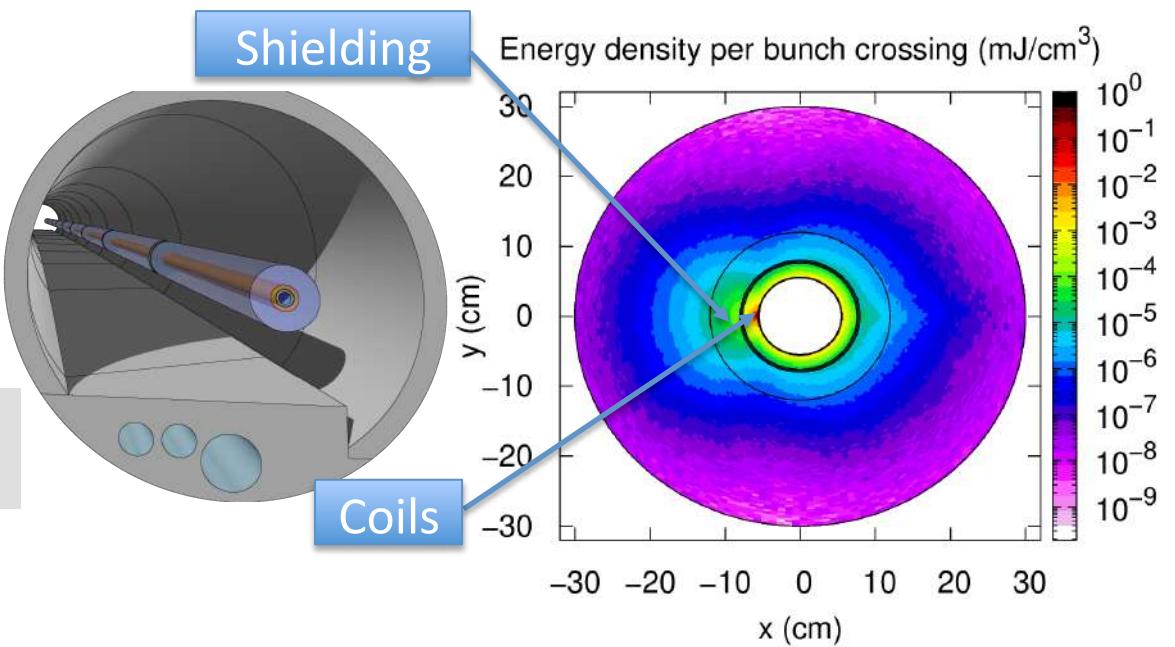
Based on MAP design, lattice design for high energy is starting

Started production of **radiation maps** and identified hot spots around IP and in arcs

Need to include radiation considerations in lattice design

**Loss challenge** in collider ring:  
Loss per unit length is constant  
fewer, but higher energy particles  
Simulations of shielding started

A. Lechner  
D. Calzolari



# High-energy Acceleration



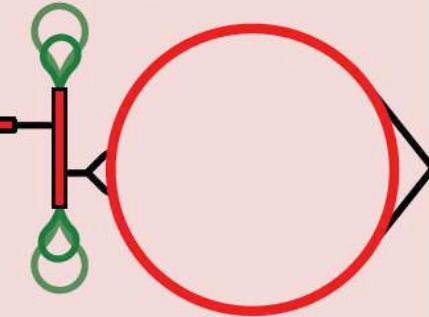
## Rapid cycling synchrotrons (RCS)

- Combine static and ramping magnets
- **Fast-ramping magnets** to follow beam energy
  - normal conducting or novel HTS
  - $O(kT/s)$  required
- **Efficient magnets and power converters** with energy recovery
  - Highest efficiency



FNAL 290 T/s HTS magnet

## Acceleration



Accelerators:  
Linacs, RLA or FFAG, RCS

## RF system

- **Important single-bunch beam loading**
- $2 \times 10^{12}$  particles in O(mm)-long bunch at 5 TeV

## FFA

- Fixed (high-field) magnets but large energy acceptance
- Challenging **lattice design** for large bandwidth and limited cost
- **Complex high-field magnets**
- Challenging beam dynamics

Test of **fast-ramping normal-conducting magnet** design

EMMA proof of FFA principle

Nature Physics 8, 243–247  
(2012)

