



# Limit shape phase transitions

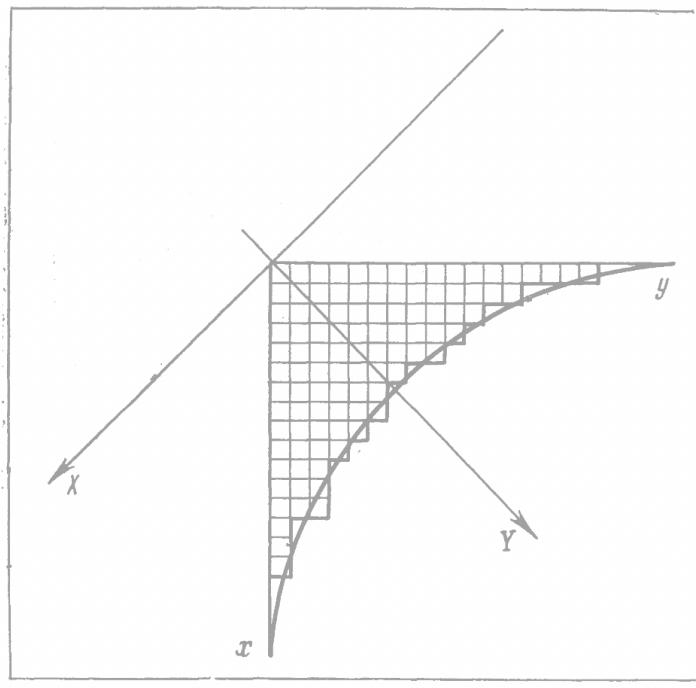
## A merger of Arctic circles

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Stony Brook University

with: **Dmitri Gangardt** and **James Pallister** (Birmingham)

based on: <https://arxiv.org/abs/2203.05269>



# Limit shape phenomenon

The limit shape phenomenon - the formation of a non-random shape in thermodynamic limit of random/statistical systems.

Vershik, Kerov, '77; Logan, Schepp, '77; Pokrovsky, Talapov, '78, '79;  
Elkies, Kuperberg, Larsen, Propp, '92; Jockusch, Propp, '98;  
Prahofer, Spohn; Borodin, Gorin; Nienhuis, Hilhorst, Blote; Cohn, Kenyon, Propp; Kenyon, Okounkov; AGA; Kenyon, Okounkov, Sheffield; Reshetikhin; Allegra, Dubail, Stephan, Viti; Colomo, Pronko, Zinn-Justin, Sportiello; Adler, Johansson, van Moerbeke; Corwin, ...

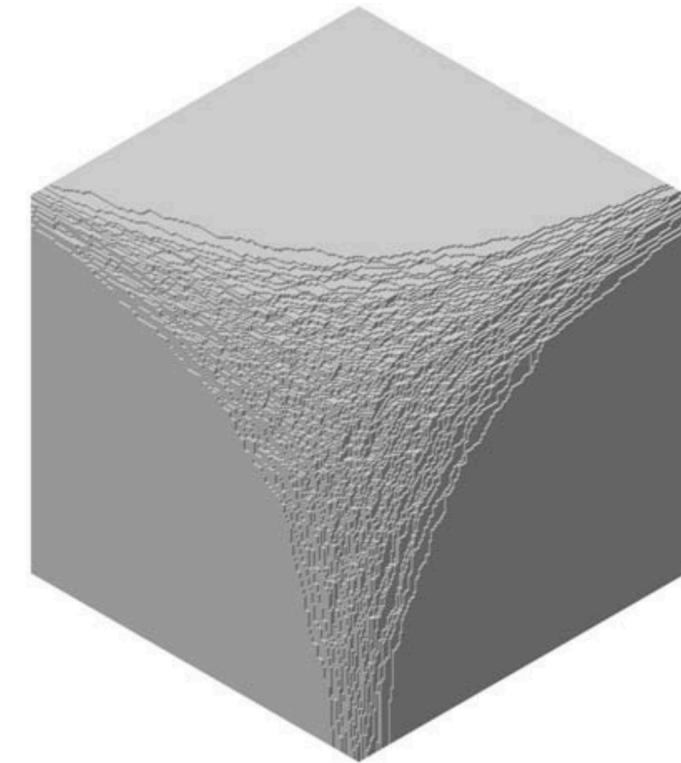
dimer models, polymer models, sorting networks, ASEP (asymmetric exclusion processes), sandpile models, bootstrap percolation models, polynuclear growth models, hydrodynamic flows, free fermions, 6-vertex models, Young diagrams, ...

# Outline

- Examples of limit shapes:
  - Emptiness formation
  - Arctic Circle
- From tilings to fluid dynamics
- Merging Arctic Circles - Gross-Witten-Wadia transition
- Limit shape phase transitions
- Universality of phase transitions

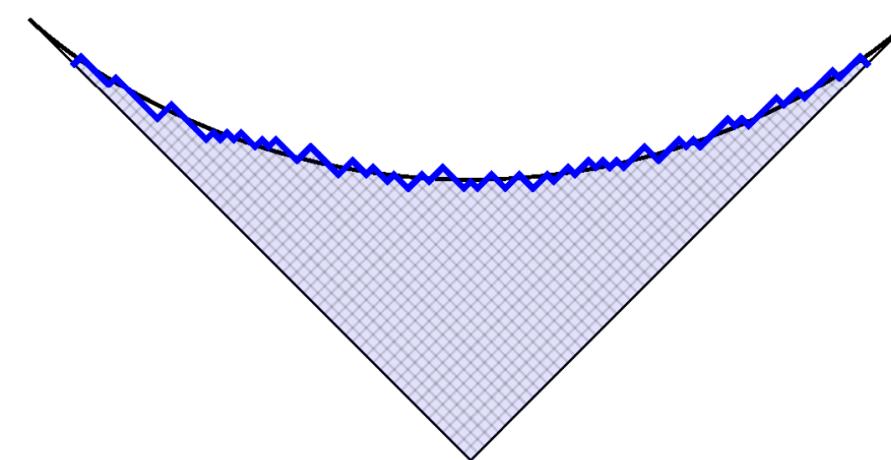
# Examples of limit shapes

- Equilibrium crystal shape



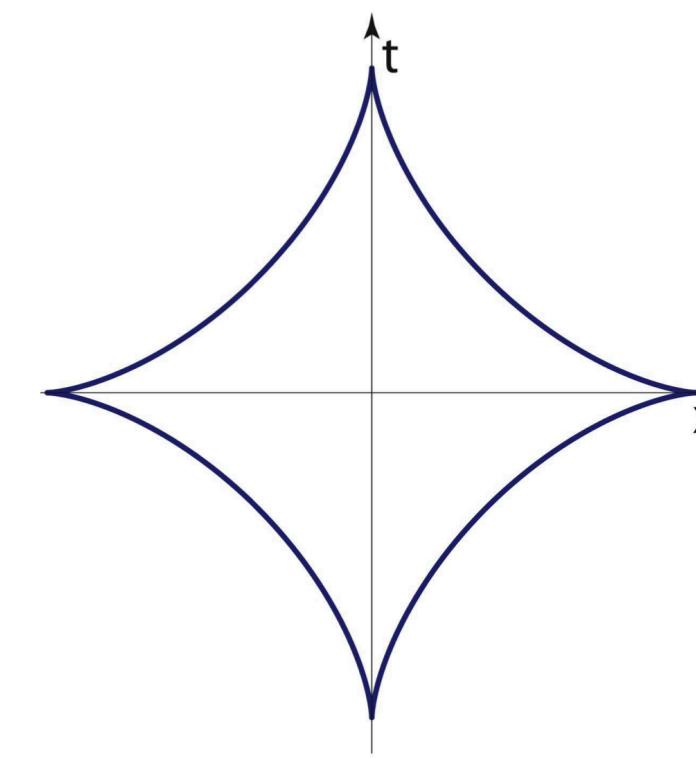
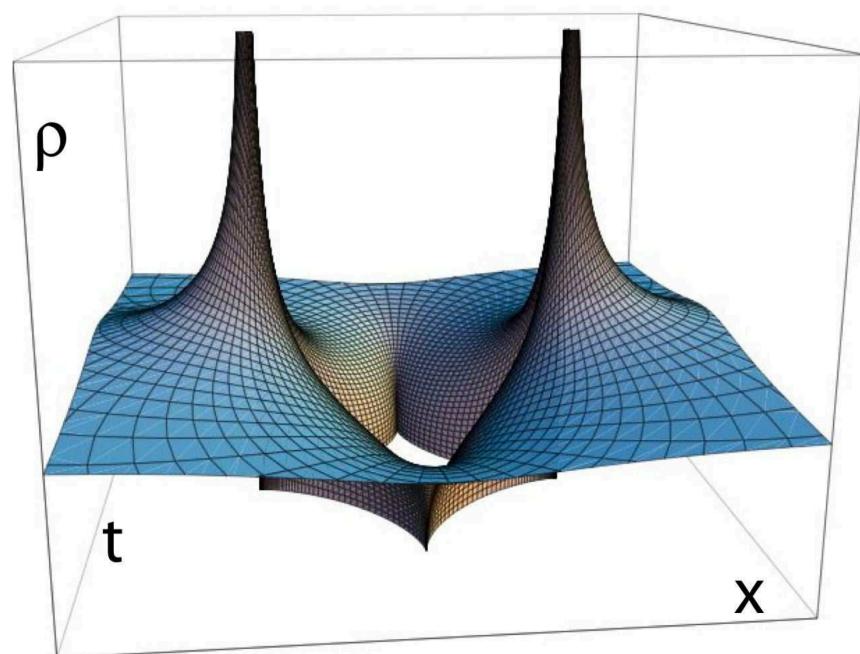
Ferrari, Spohn, 2003

- Random Young tableau



Vershik, Kerov, 1977  
Logan, Shepp, 1977

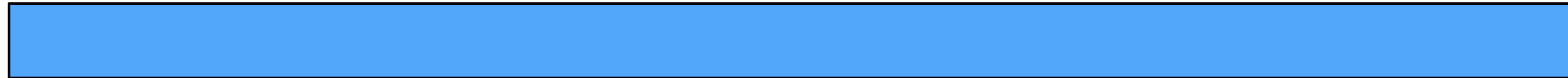
- Emptiness formation in 1d quantum fermions



AGA, 2005

# Emptiness Formation

# Imagine a quantum Fermi gas on a line



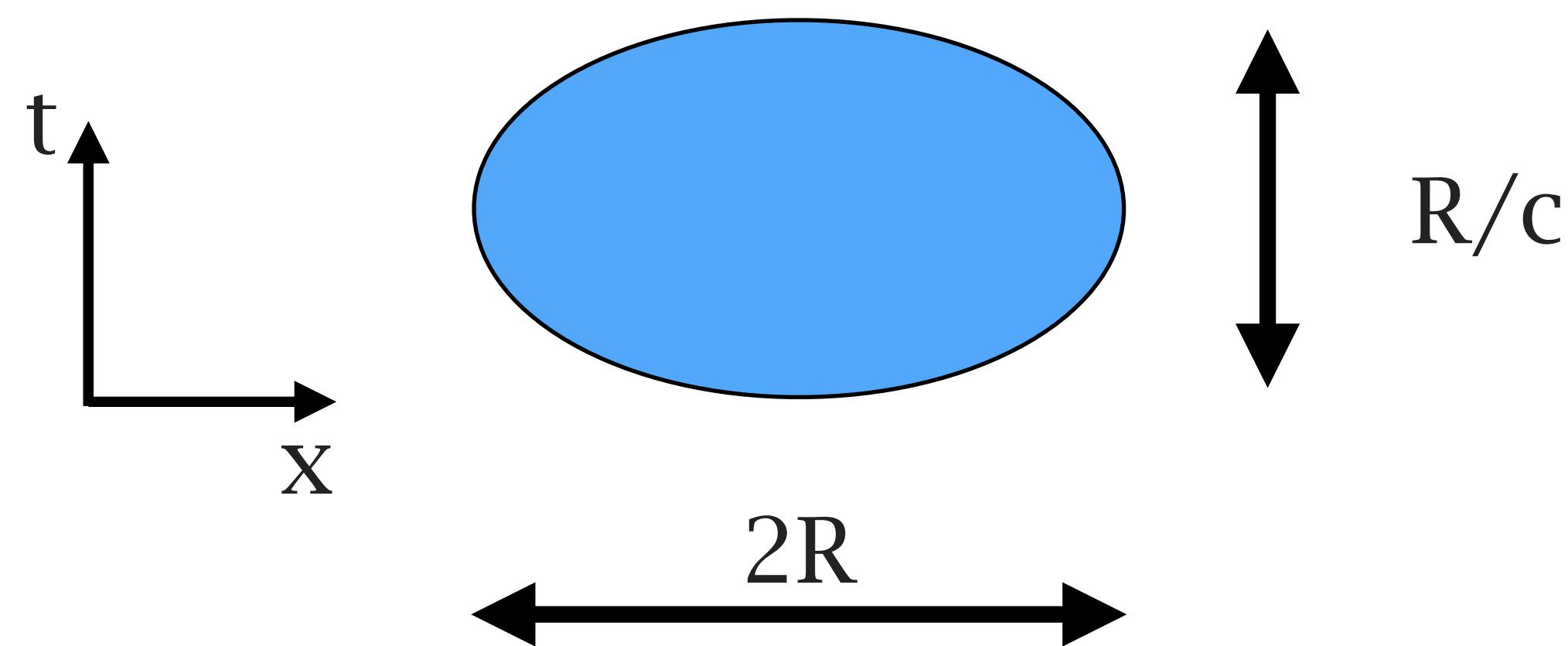
What is the probability that at time  $t=0$ , the gas parts away to form an emptiness of the size  $2R$  at large  $R$ ?



- For random points in one dimension the probability of an empty spot is  $P \sim e^{-2\rho_0 R}$
- Wrong for a quantum system in the ground state
- For large  $R$  - collective dynamics

# Estimate of the instanton's action

- For large  $R$ , the probability is determined by the action of an instanton
- What motion of the gas (in imaginary time) is optimal for the emptiness formation?



$$P \sim e^{-\# R^2}$$

$$P \sim e^{-\frac{1}{2}(k_F R)^2}$$

# Hydrodynamics of free fermions

$$H = \int dx \rho \left( \frac{v^2}{2} + \frac{\pi^2 \hbar^2 \rho^2}{2} \right)$$

fluid's kinetic  
energy

Pauli's internal  
energy



$$\partial_\tau \rho + \partial_x (\rho v) = 0$$

$$\partial_\tau v + v \partial_x v = \partial_x \frac{\pi^2 \rho^2}{2}$$



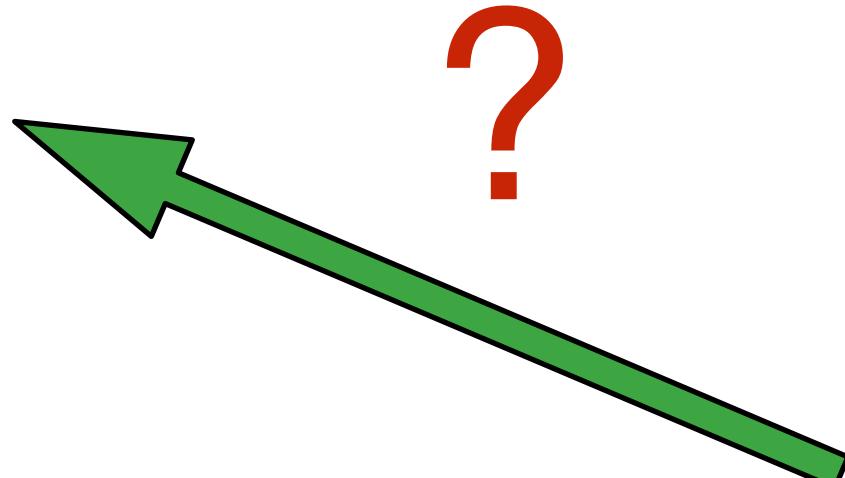
$$ik_\tau + kk_x = 0$$

$$k = \pi \rho \pm iv$$

Find  $F(k)$  from boundary  
conditions and compute the  
action on the solution



$$P \sim e^{-S}$$



$$x + ik\tau = F(k) \quad - \text{analytic}$$

# Emptiness solution

AGA, 2005

General solution:

$$z \equiv \boxed{x + ik\tau = F(k)}$$

$$\begin{aligned} ik_\tau + kk_x &= 0 \\ k &= \pi\rho + iv \\ k, \bar{k}, \rho, v(x, \tau) &\longrightarrow P \sim e^{-S} \end{aligned}$$

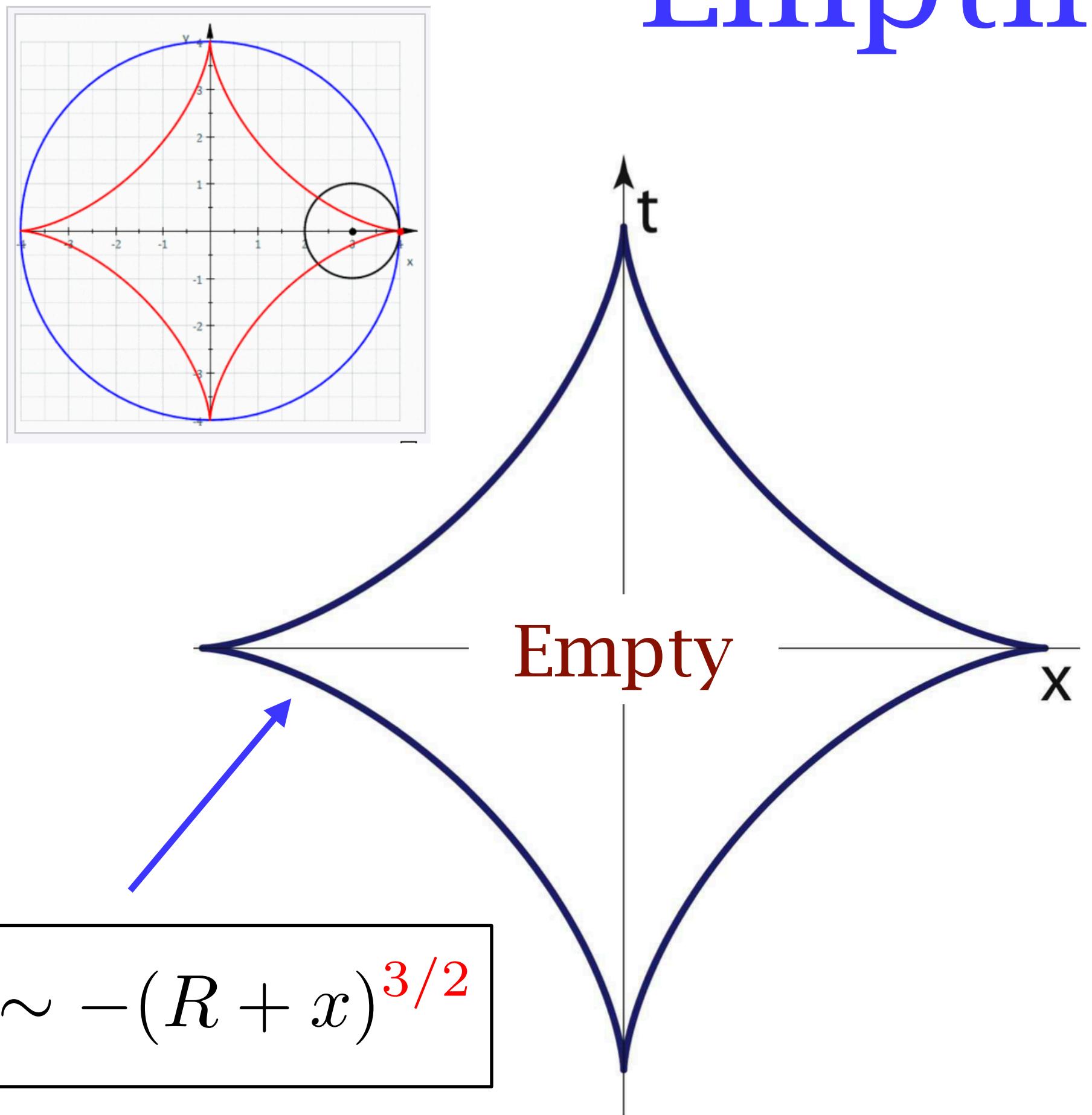
$F(k)$  - any analytic function

Emptiness:  $F(k)$  - to be found from boundary conditions

$$\boxed{F(k) = R \frac{k}{\sqrt{k^2 - k_F^2}}}$$

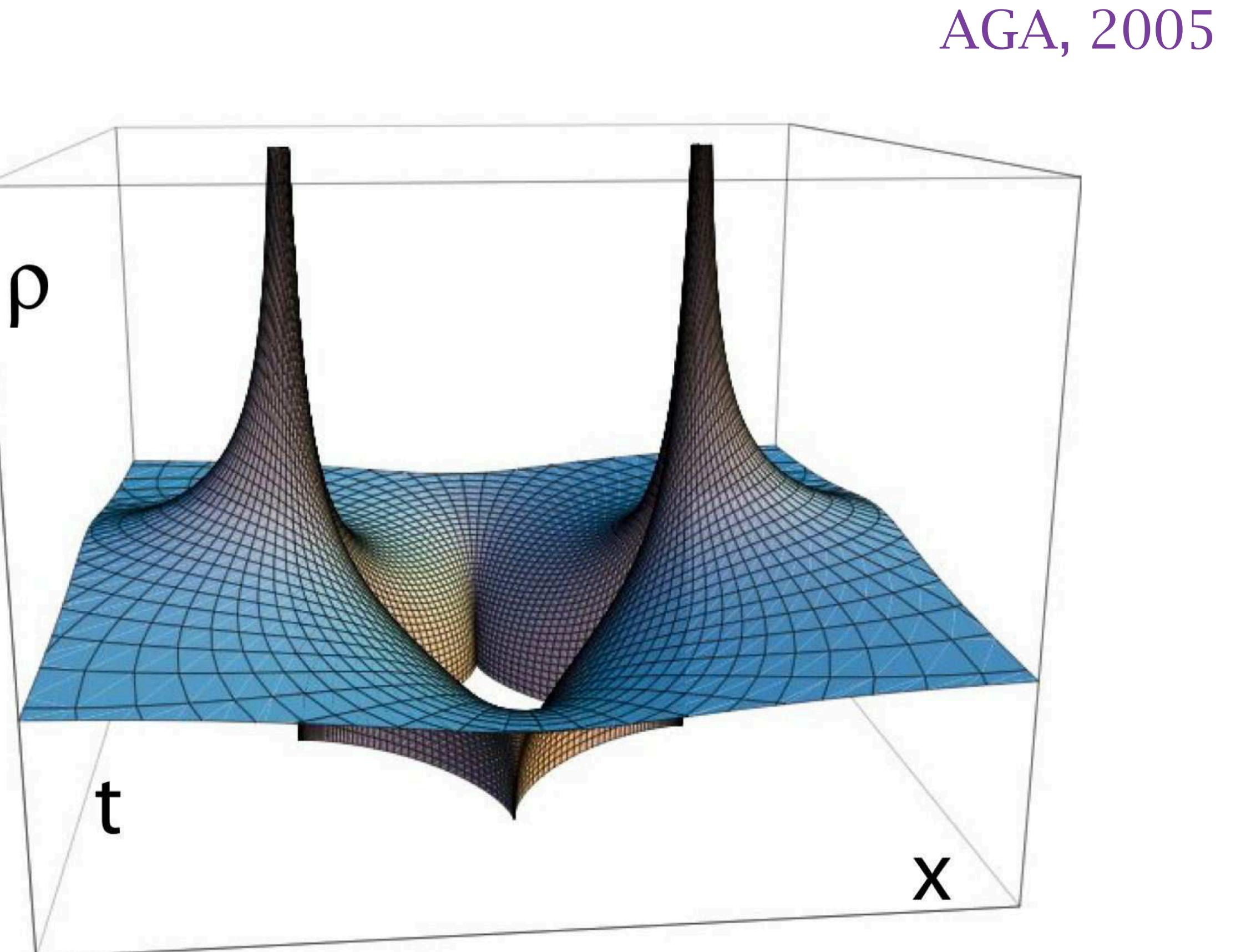
$$\begin{aligned} k &= k_F \frac{z}{\sqrt{z^2 - R^2}} \\ \tau = 0, z = x & \\ \rightarrow \rho = 0 \text{ for } x^2 - R^2 < 0 & \\ z \rightarrow \infty & \\ \rightarrow k \rightarrow k_F & \end{aligned}$$

# Emptiness and Astroid



The shape of the empty  
space - Astroid!

$$x^{2/3} + t^{2/3} = R^{2/3}$$



The density of gas  
around the emptiness

# Remarks

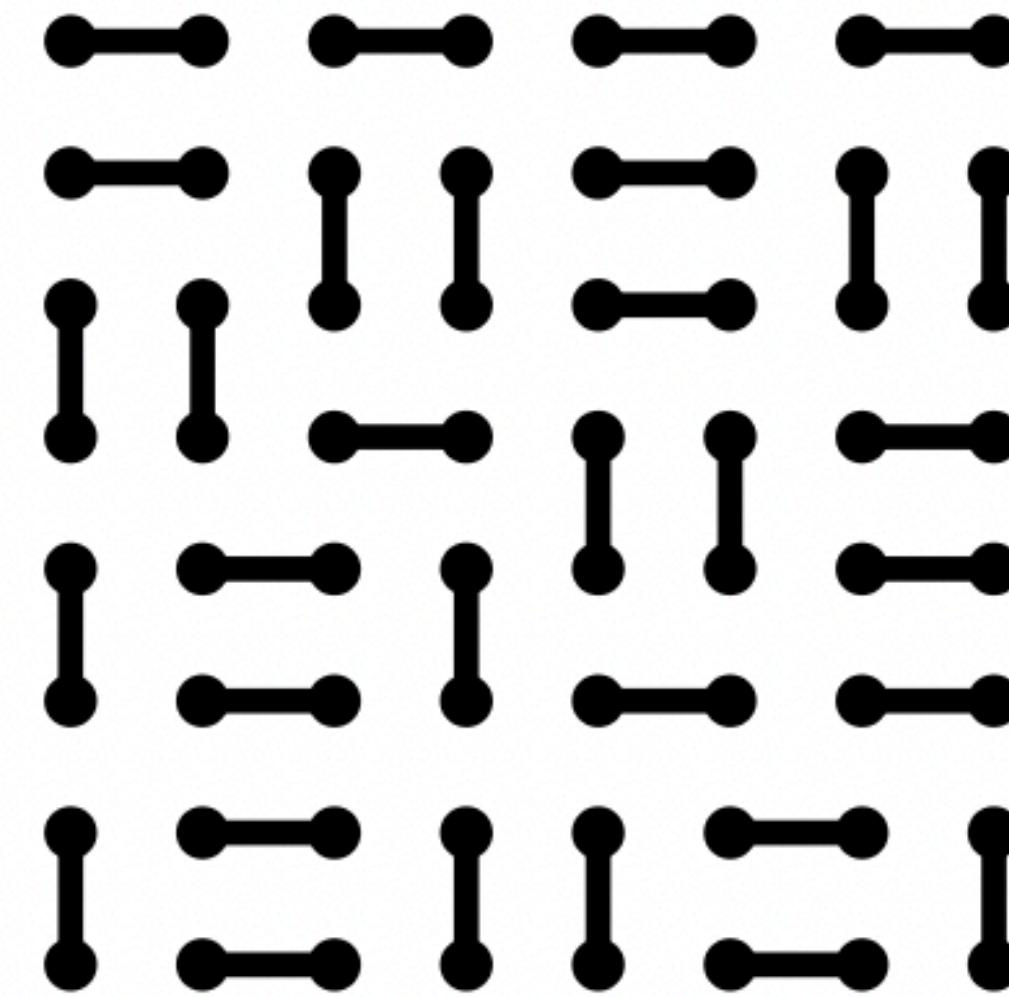
- The probability of large fluctuation is determined by the action of an instanton
- Instanton - solution of classical hydrodynamic equations in imaginary time
- Frozen regions minimize action and are allowed as parts of an instanton solution
- Globally: combination of the solution of hydro equations and frozen regions
- Free fermions: find  $F(k)$  from the boundary conditions imposed in spacetime
- Technical: the action computed on equations of motion - integral over the boundary of the frozen region
- Emptiness formation probability can also be computed from the ground state wave function without fluid dynamics

# Domino Tilings and Arctic Circle Theorem

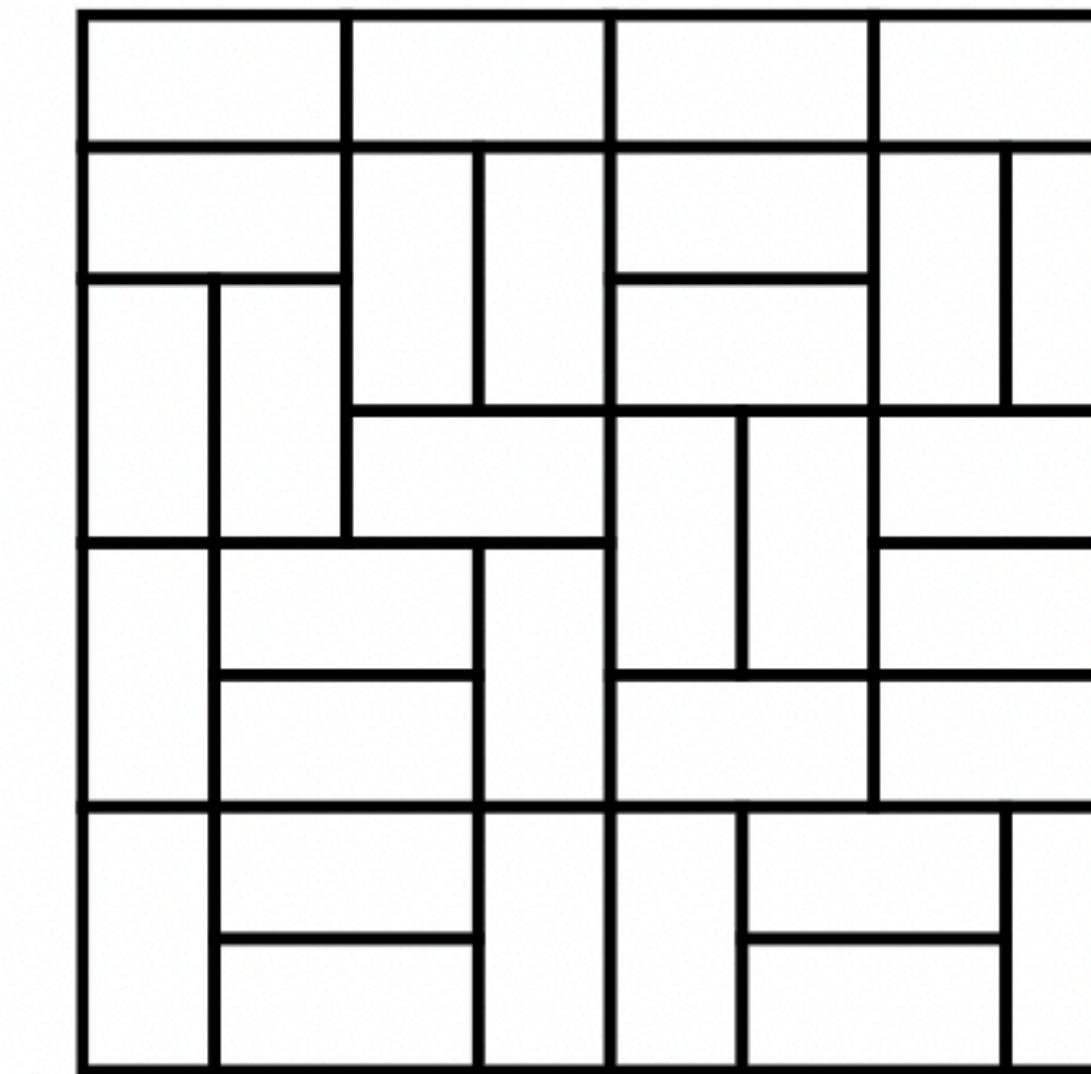
# Domino tiling

Problem: In how many ways one could tile the  $8 \times 8$  chessboard by dominos of the size  $2 \times 1$ ?

Kasteleyn, 1963



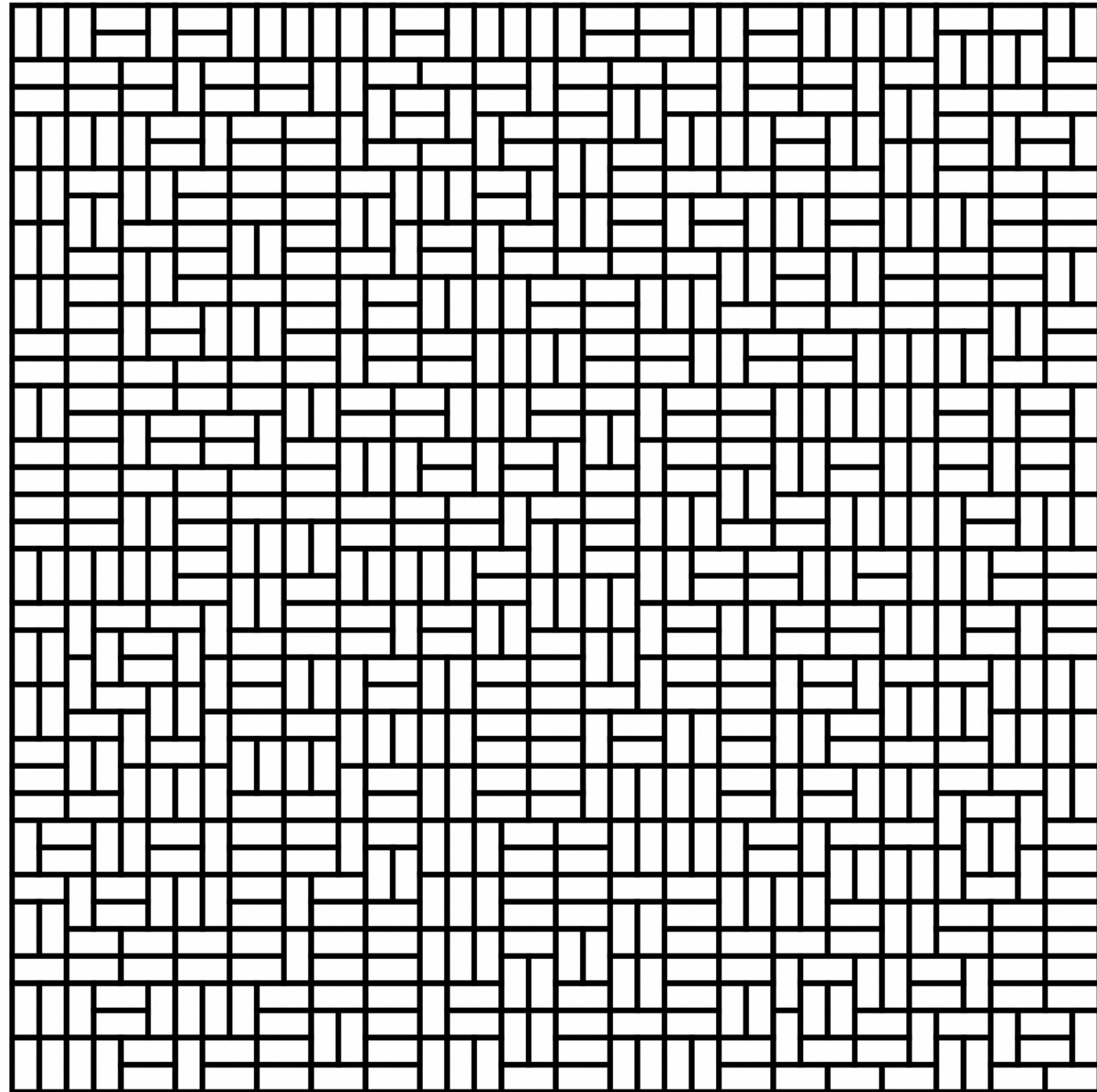
Dimer covering of  
the lattice  $8 \times 8$



Example of a  
domino tiling  $8 \times 8$

Motivation: random dimer coverings in problems of absorption of diatomic molecules by a crystal surface.

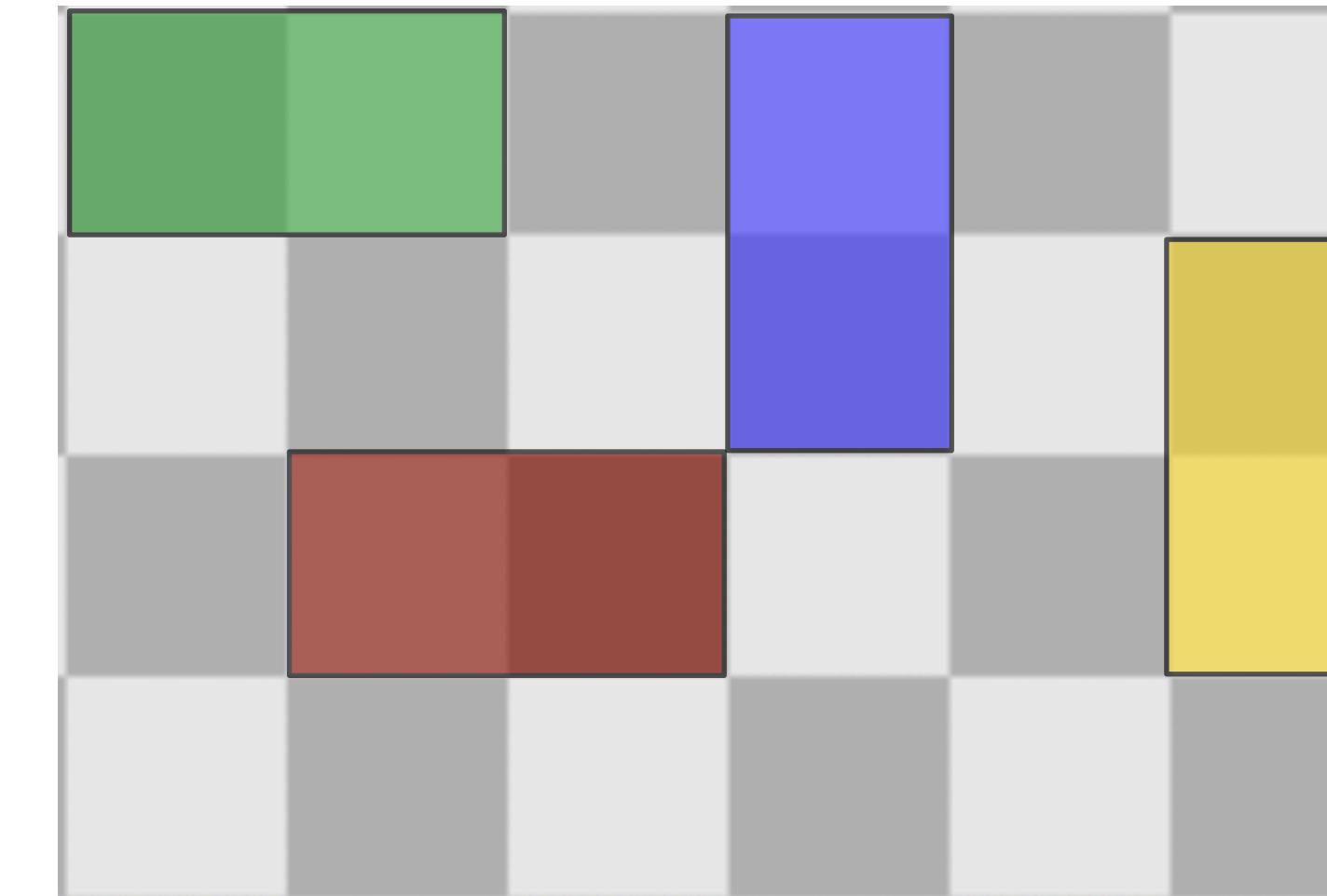
# Colored dominos



Random tiling of the board 40 x 40  
(totally about  $10^{197}$  tilings)

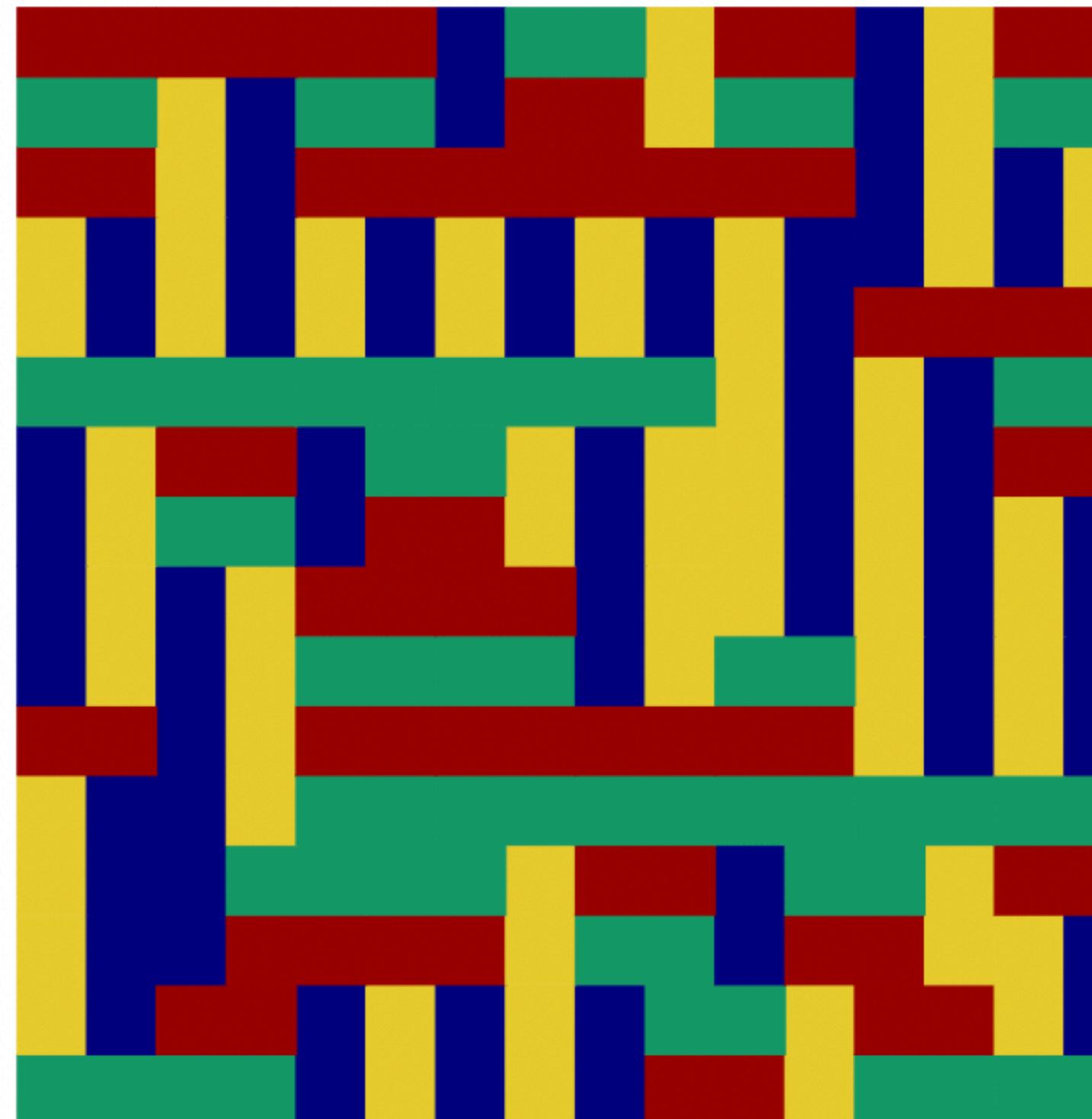
Dominos can be horizontal,  
vertical, ...

Let's color them!

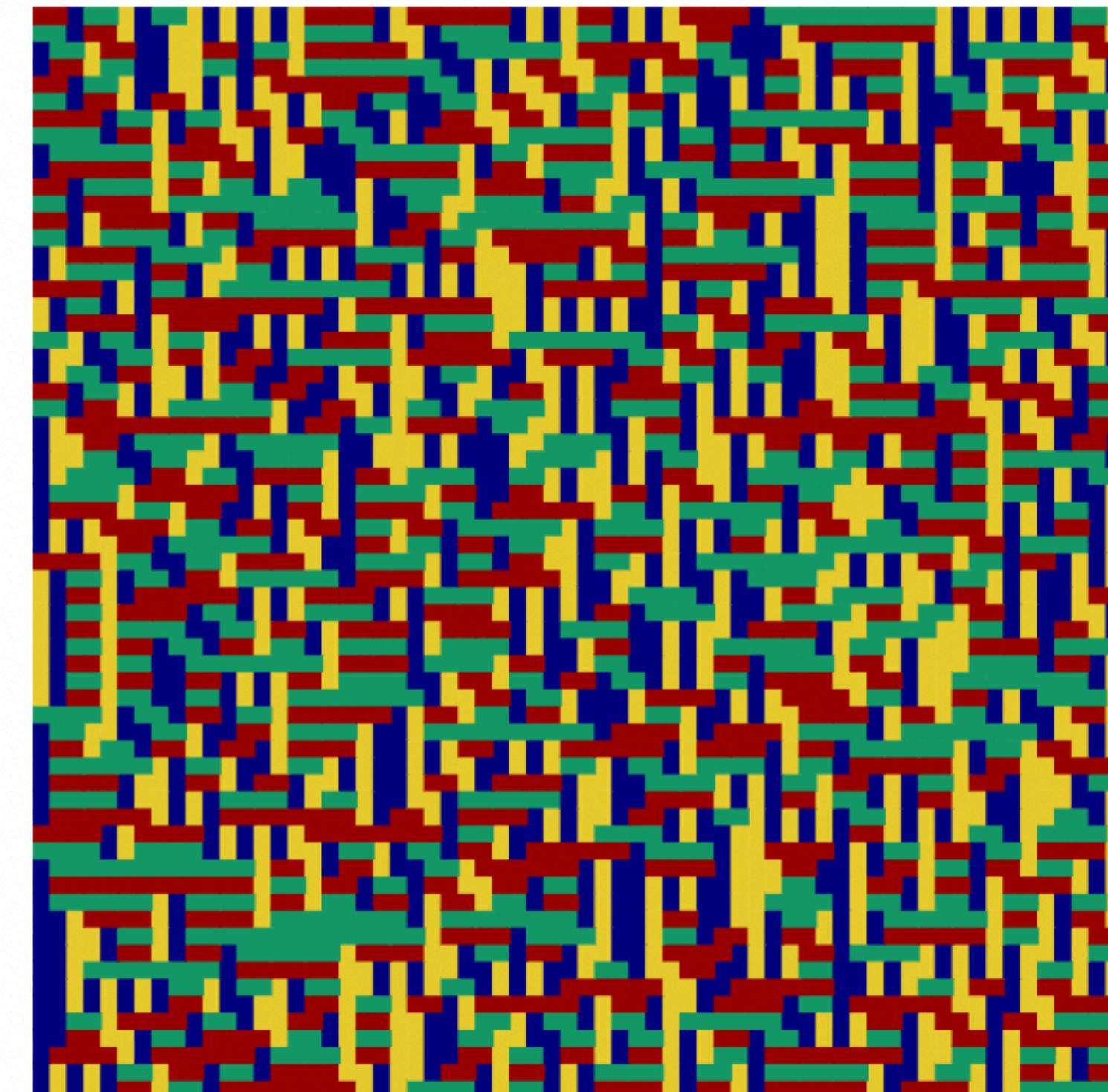


# Colored tilings $L \times L$

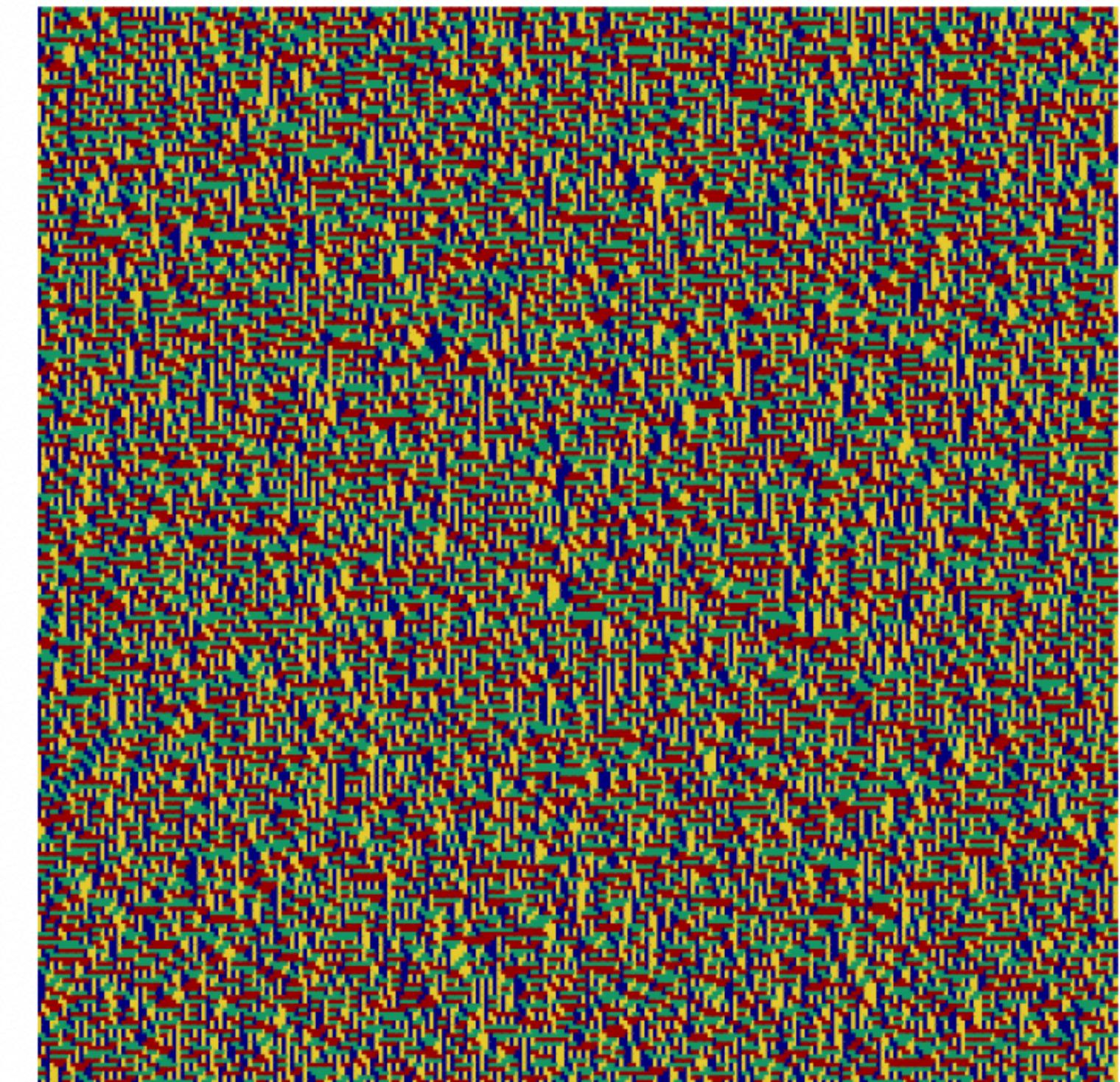
Stéphan, 2020



$L=16$



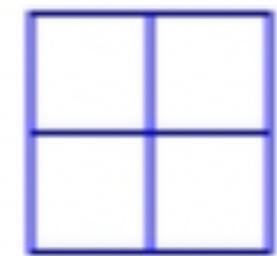
$L=64$



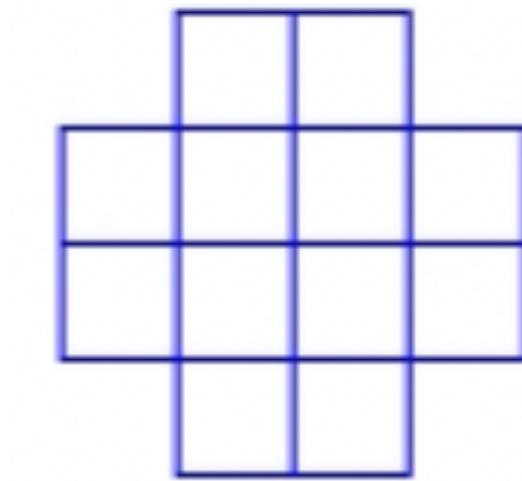
$L=256$

$L \rightarrow \infty$  - Thermodynamic limit

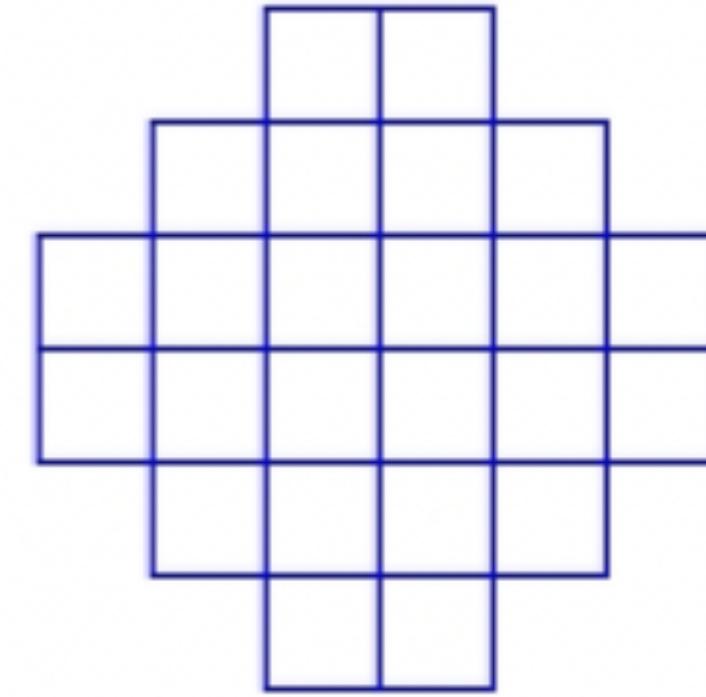
# Aztec Diamond



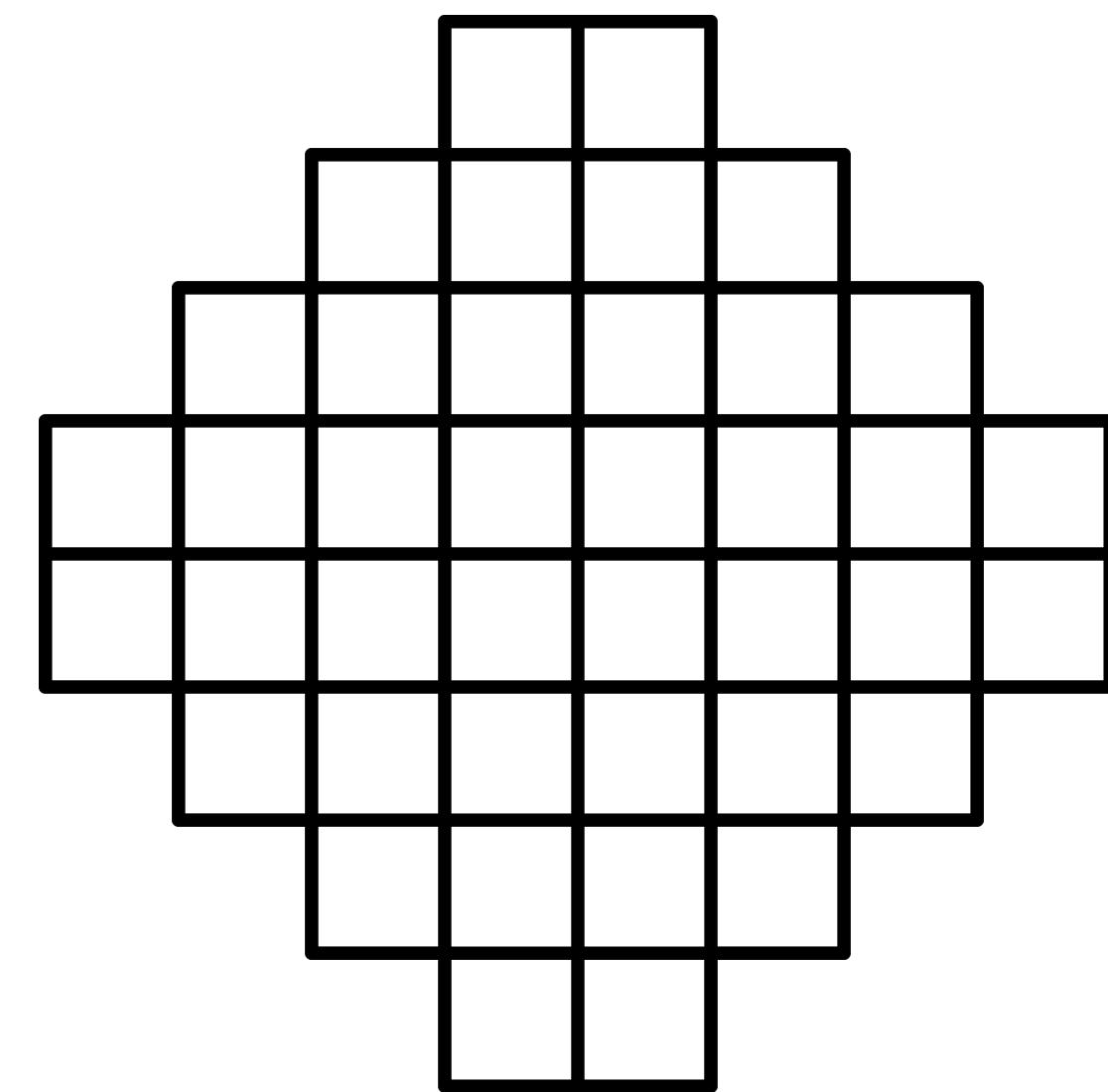
1



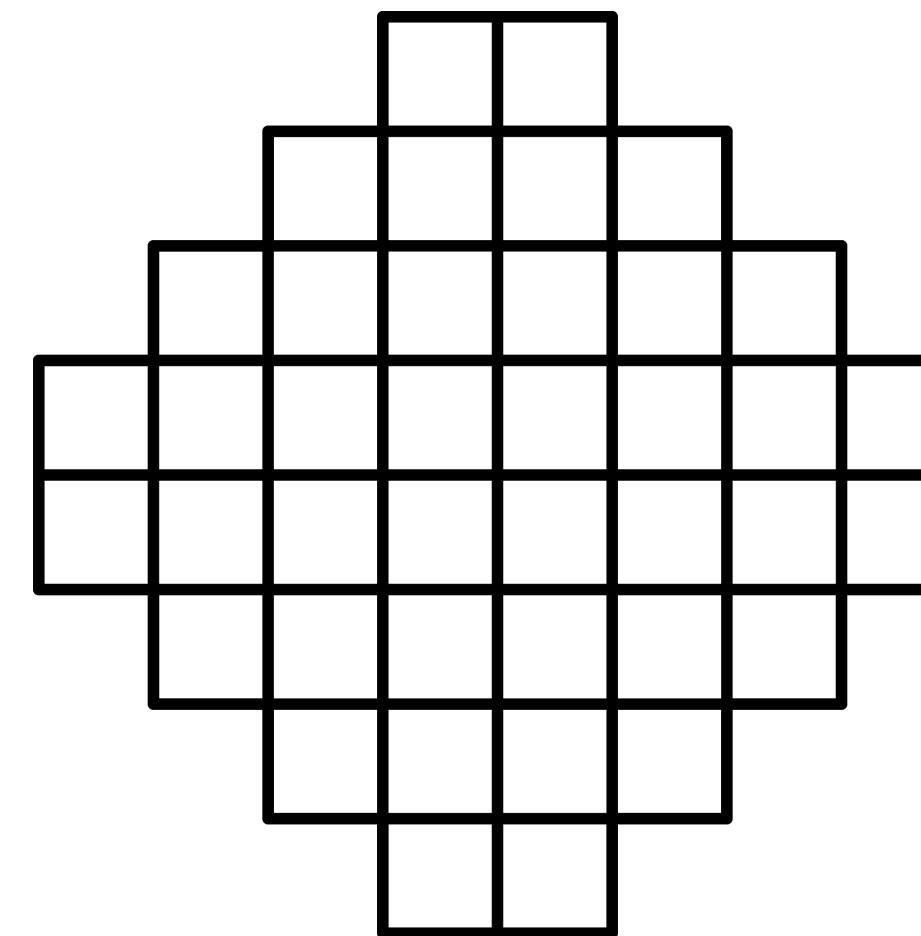
2



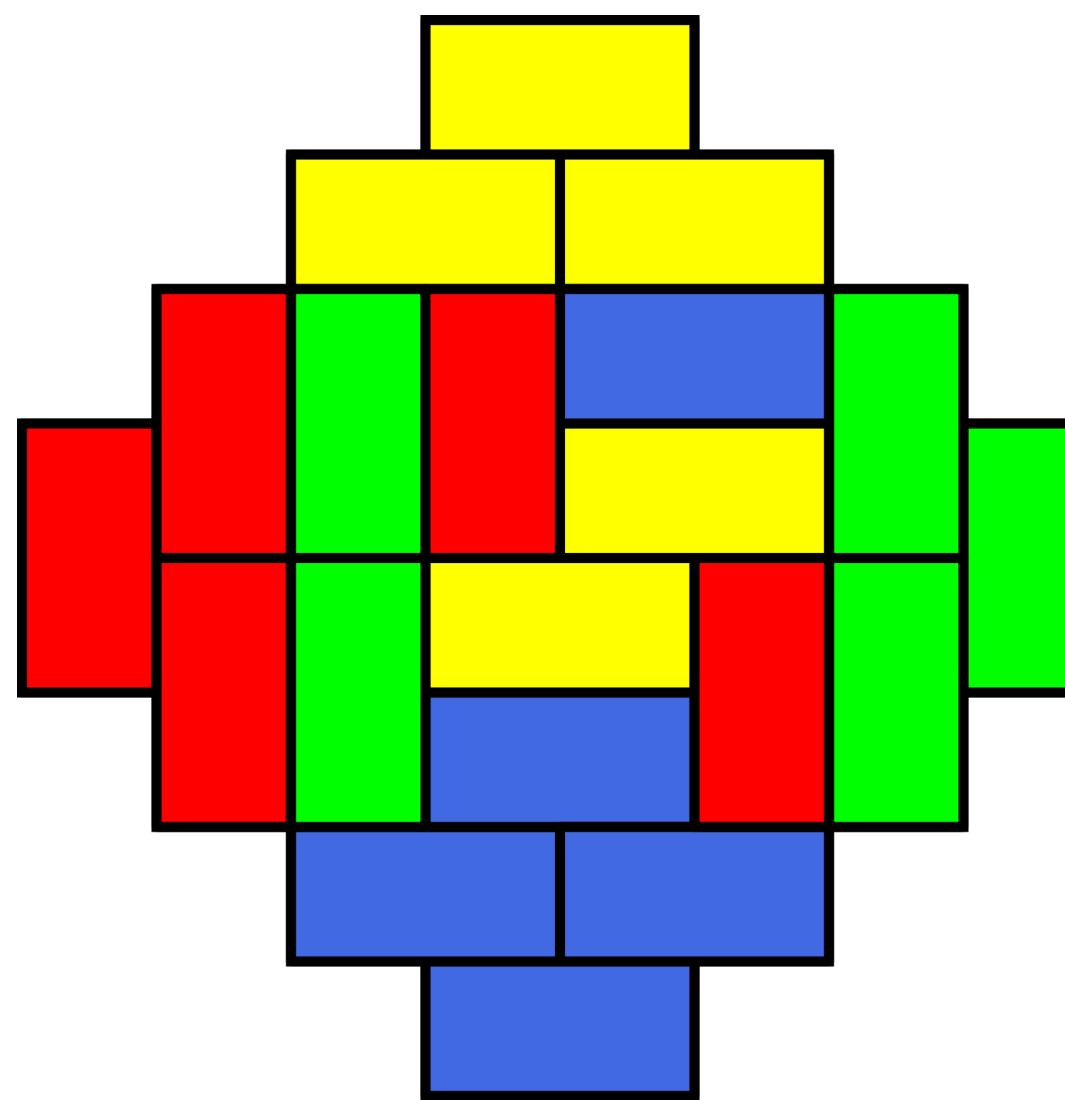
3



Aztec diamond of  
order 4



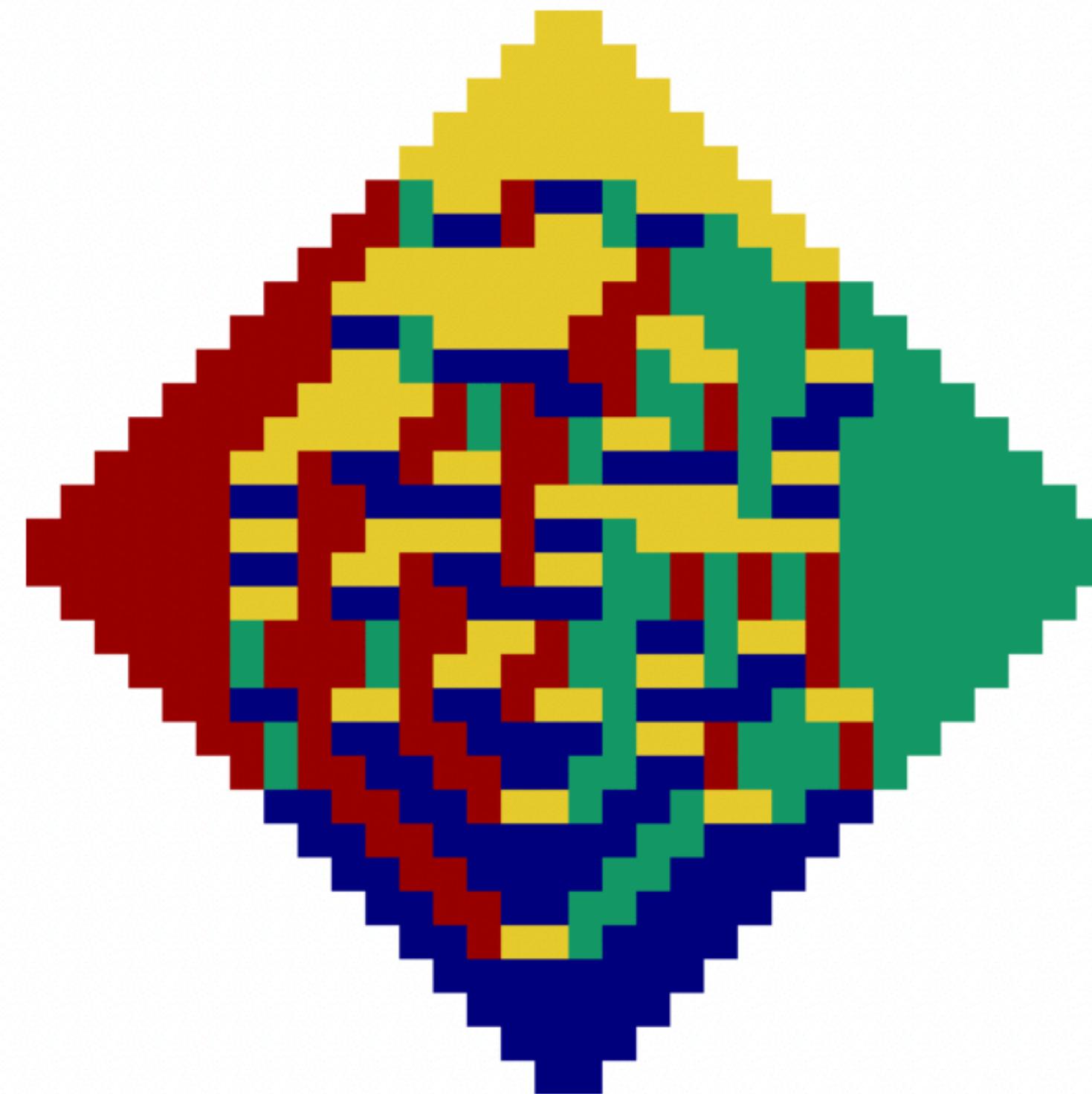
4



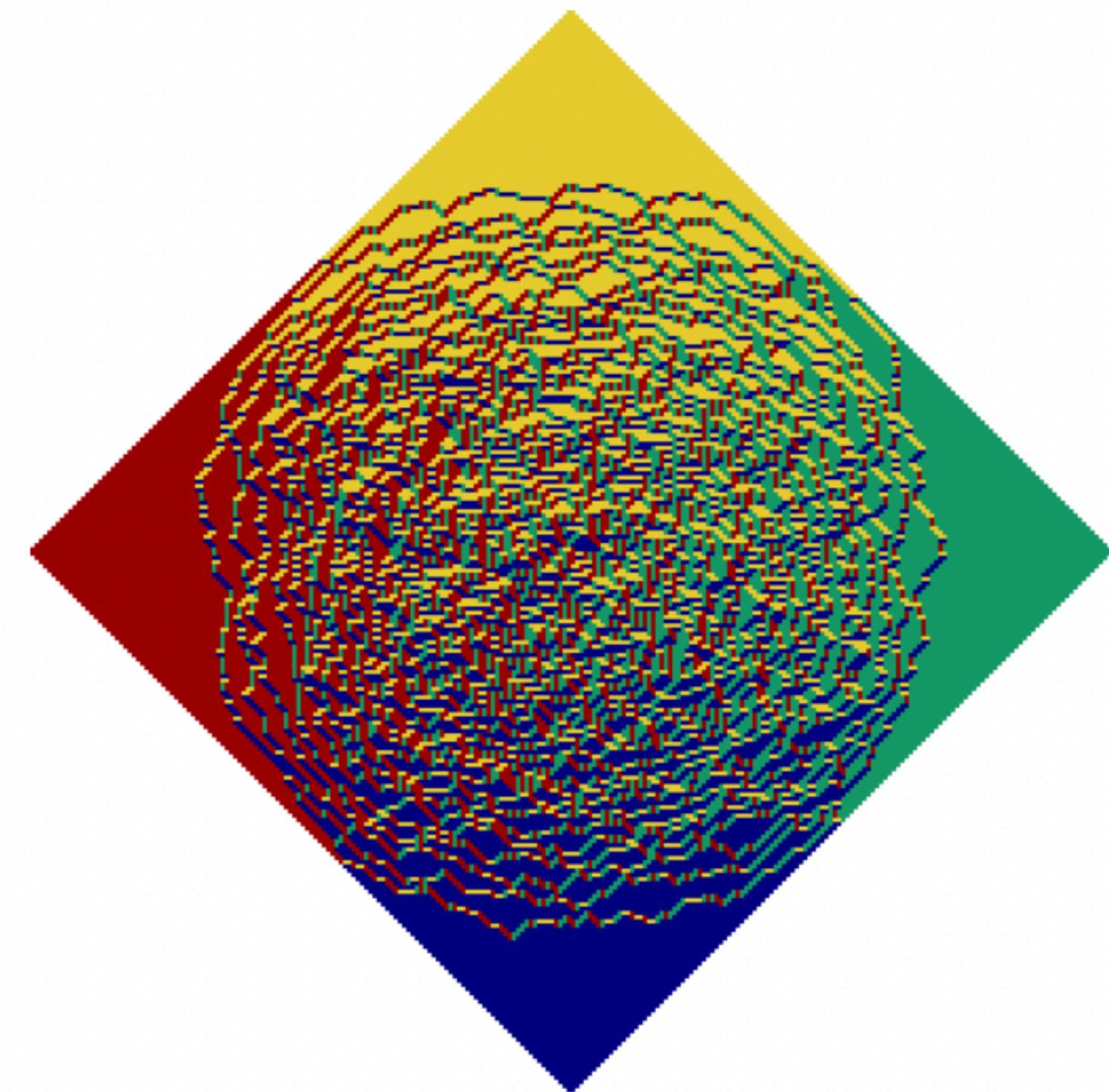
Random tiling of Aztec  
diamond

# Thermodynamic limit

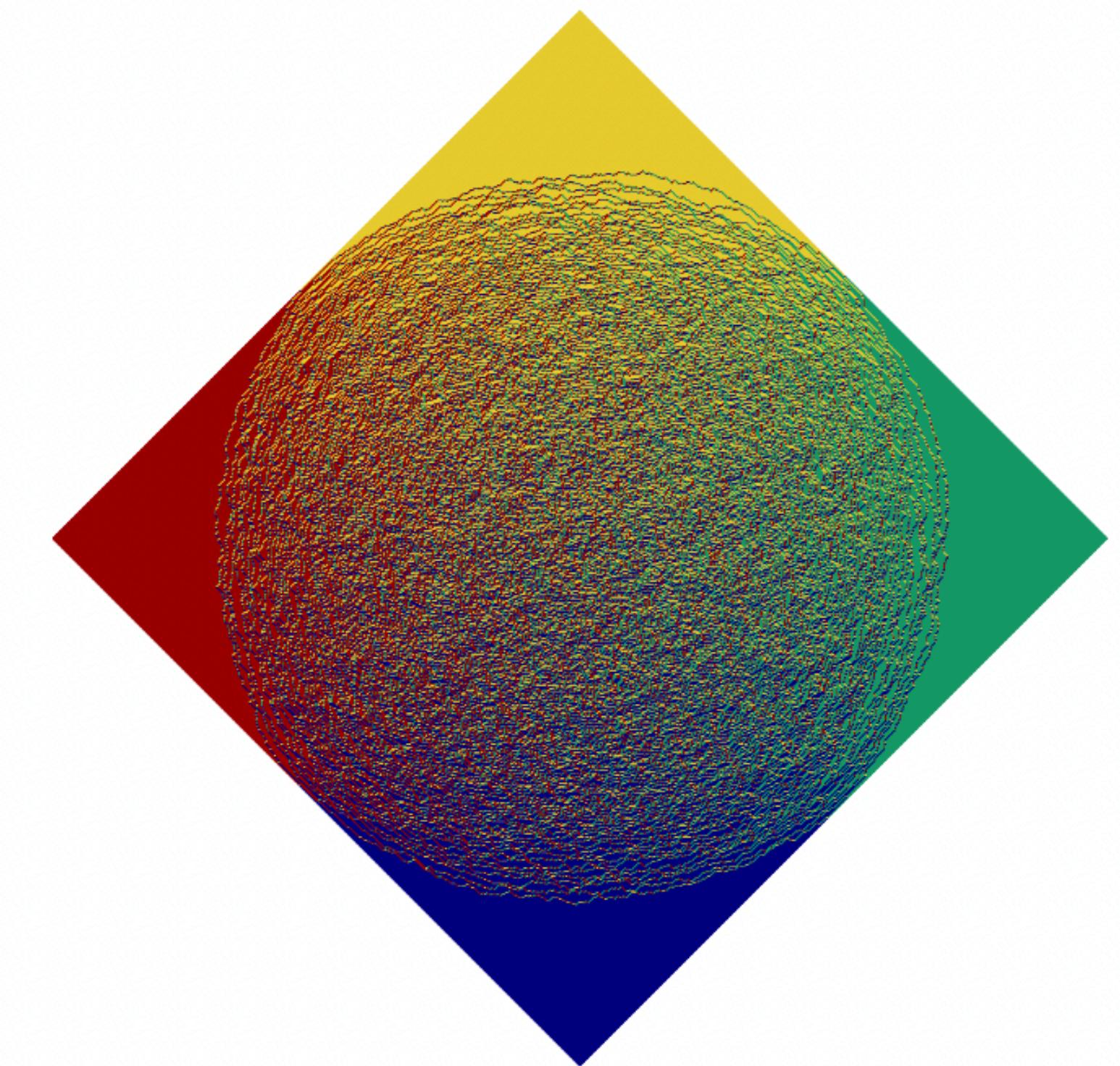
Stéphan, 2020



$L=16$



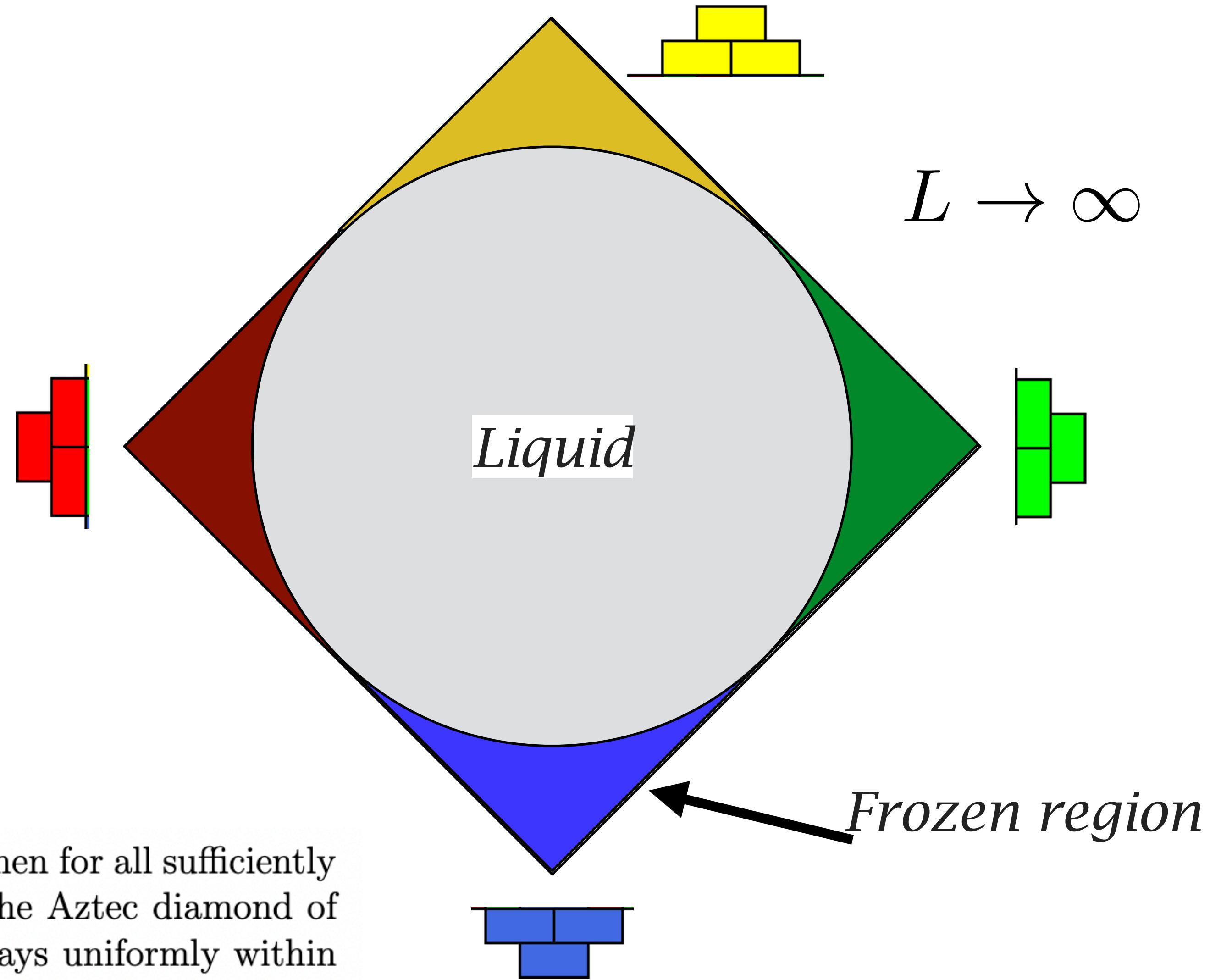
$L=128$



$L=1024$

# Arctic circle theorem

In thermodynamic limit with probability going to 1 the random tiling of Aztec diamond looks like the one shown in picture.

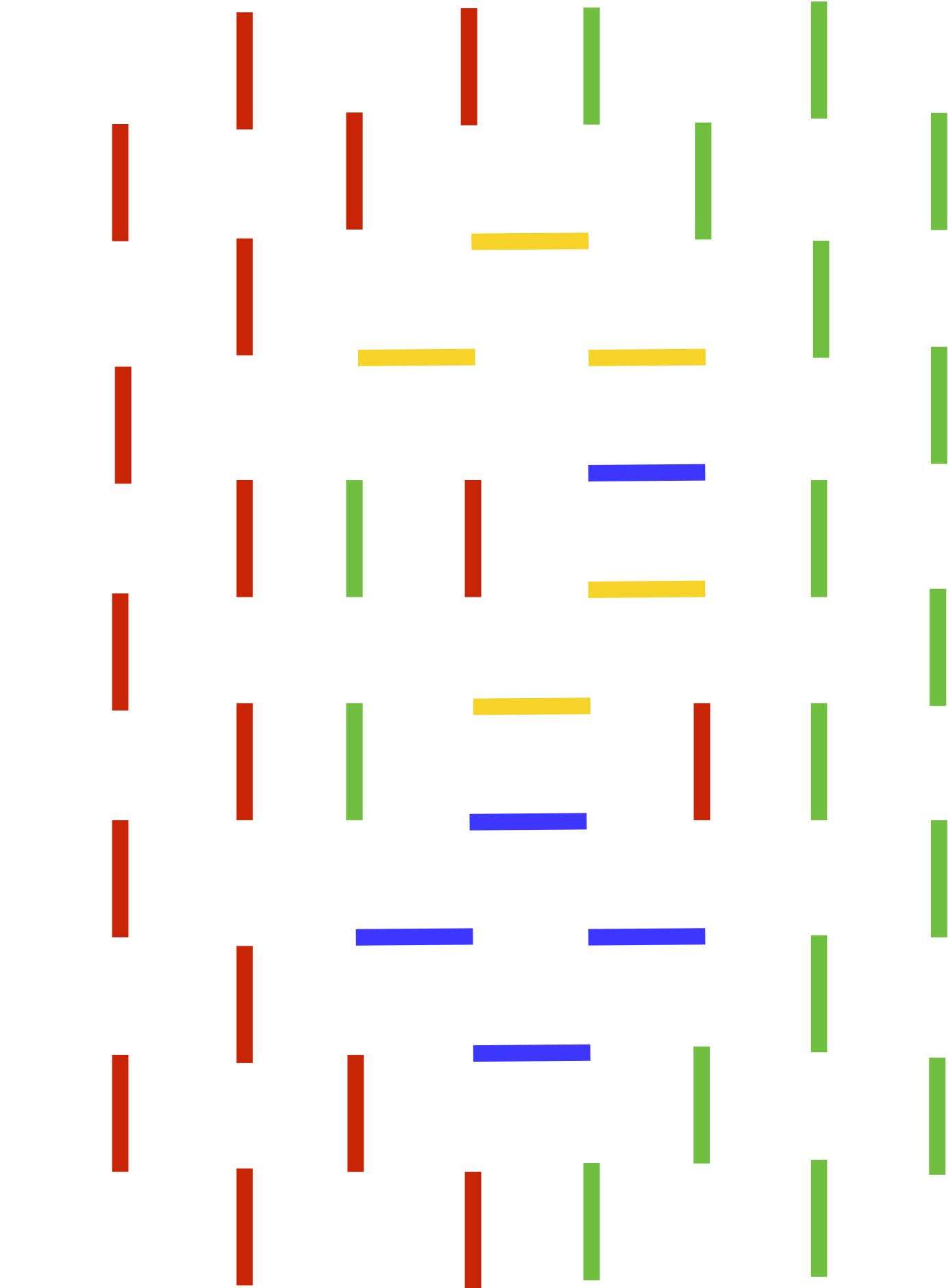
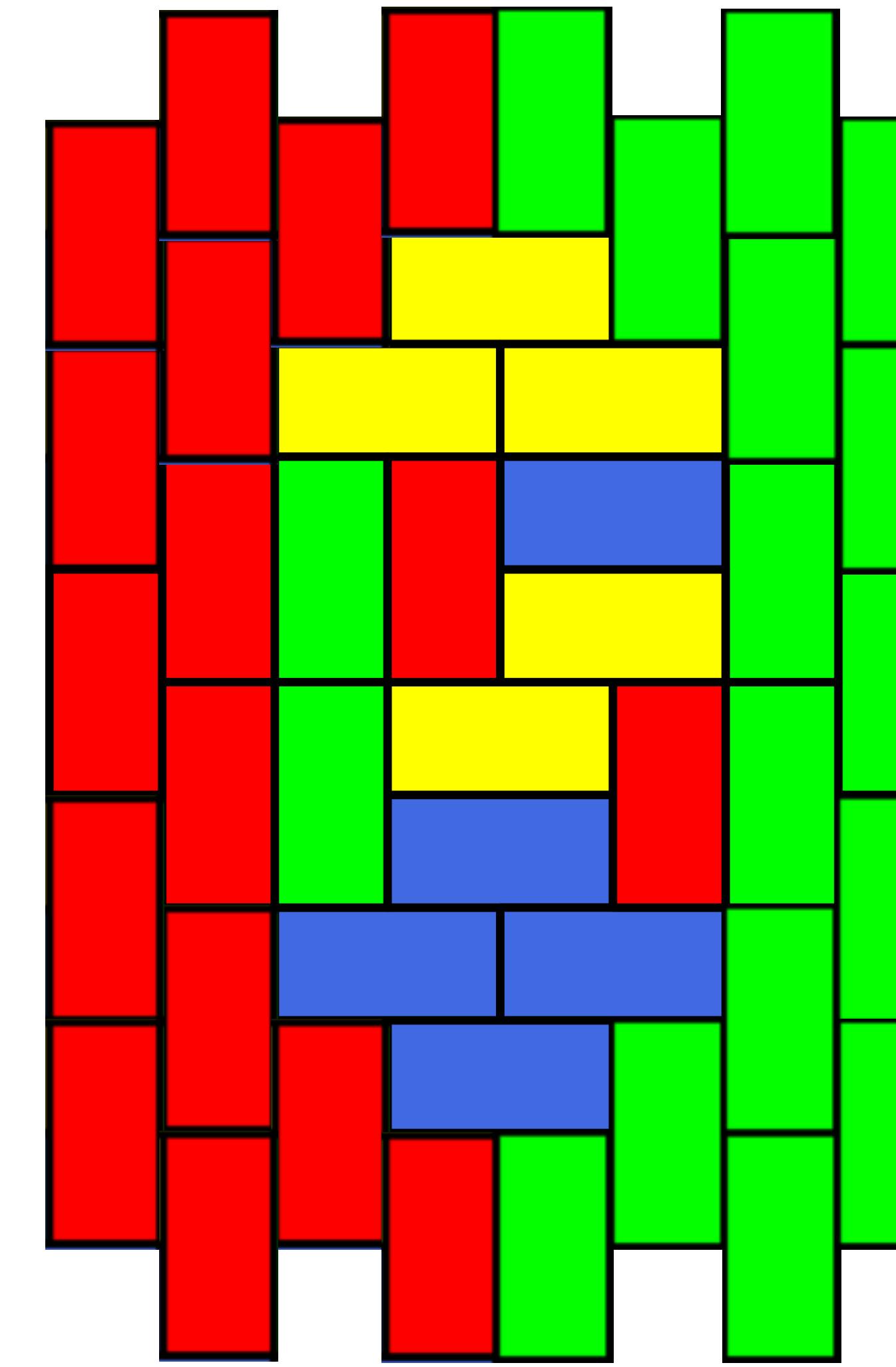
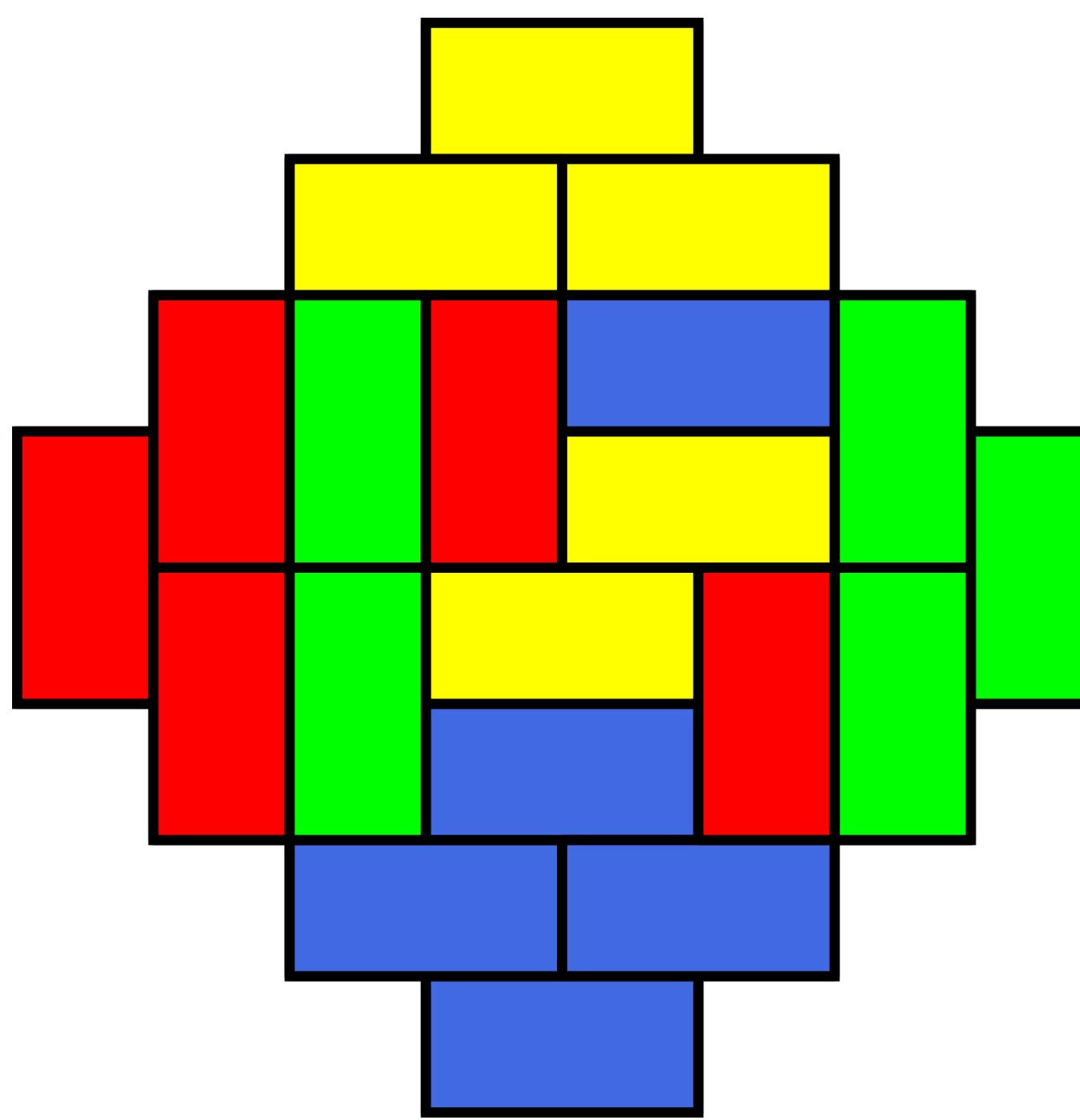


William Jockusch, James Propp, and Peter Shor.  
"Random domino tilings and the arctic circle theorem."  
*arXiv preprint math/9801068* (1998).

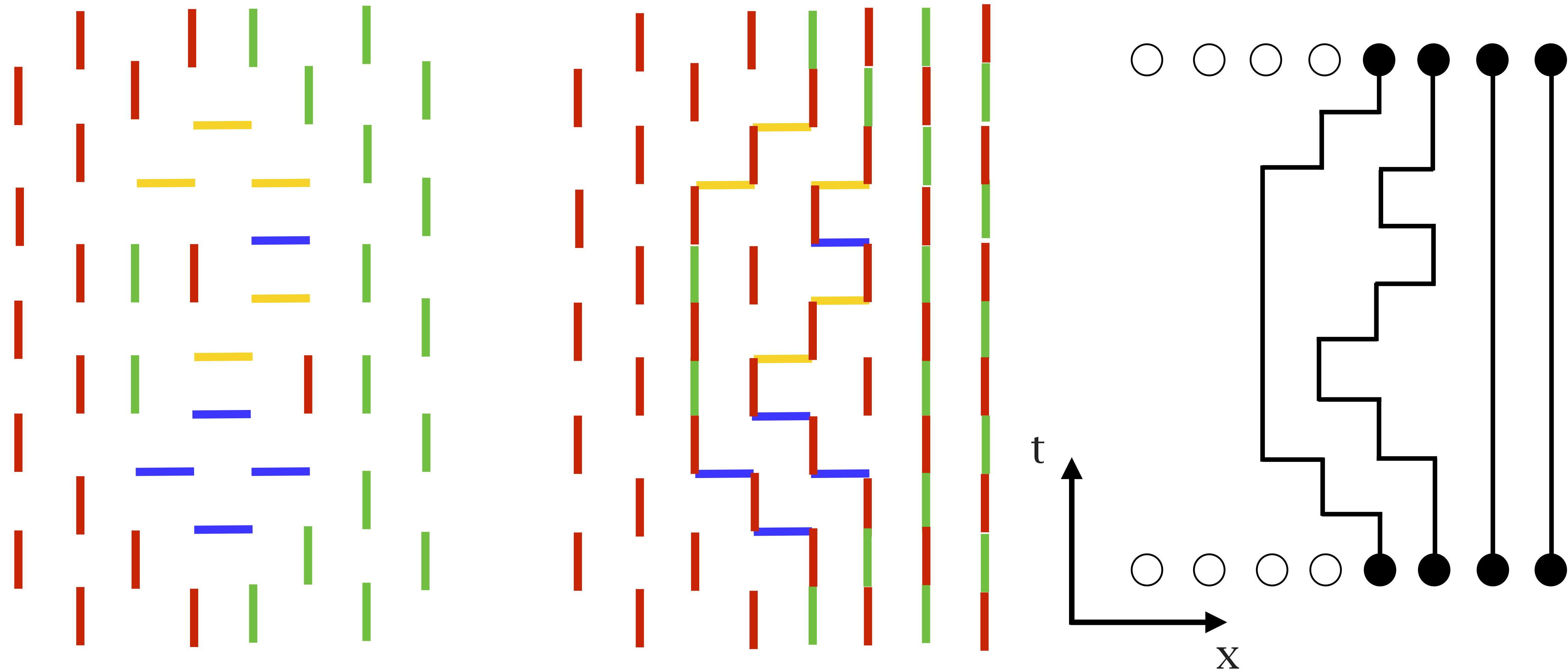
**THEOREM 1** (the Arctic Circle Theorem): Fix  $\epsilon > 0$ . Then for all sufficiently large  $n$ , all but an  $\epsilon$  fraction of the domino tilings of the Aztec diamond of order  $n$  will have a temperate zone whose boundary stays uniformly within distance  $\epsilon n$  of the inscribed circle.

# Tilings as flows

# From dominos to dimers

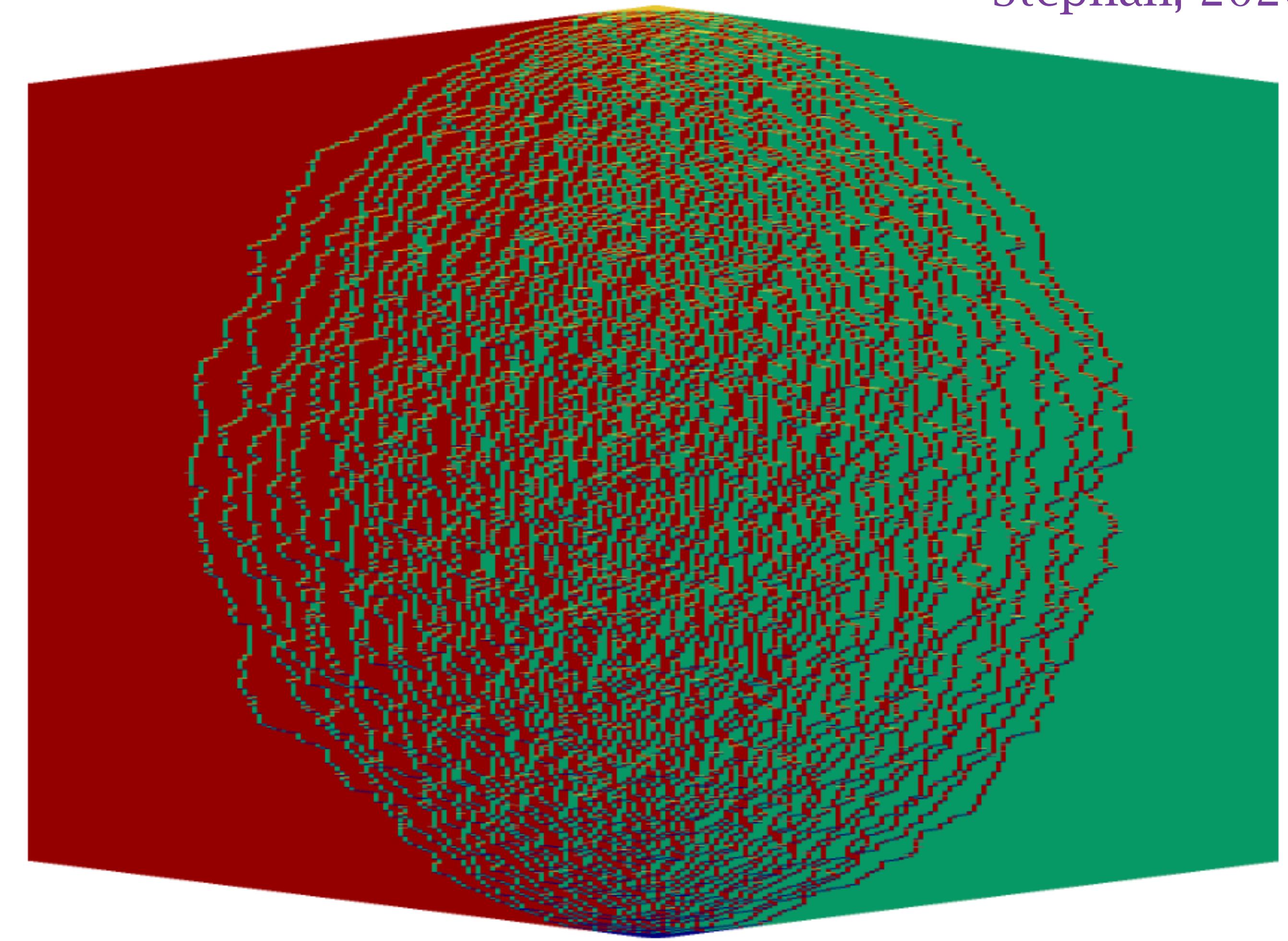
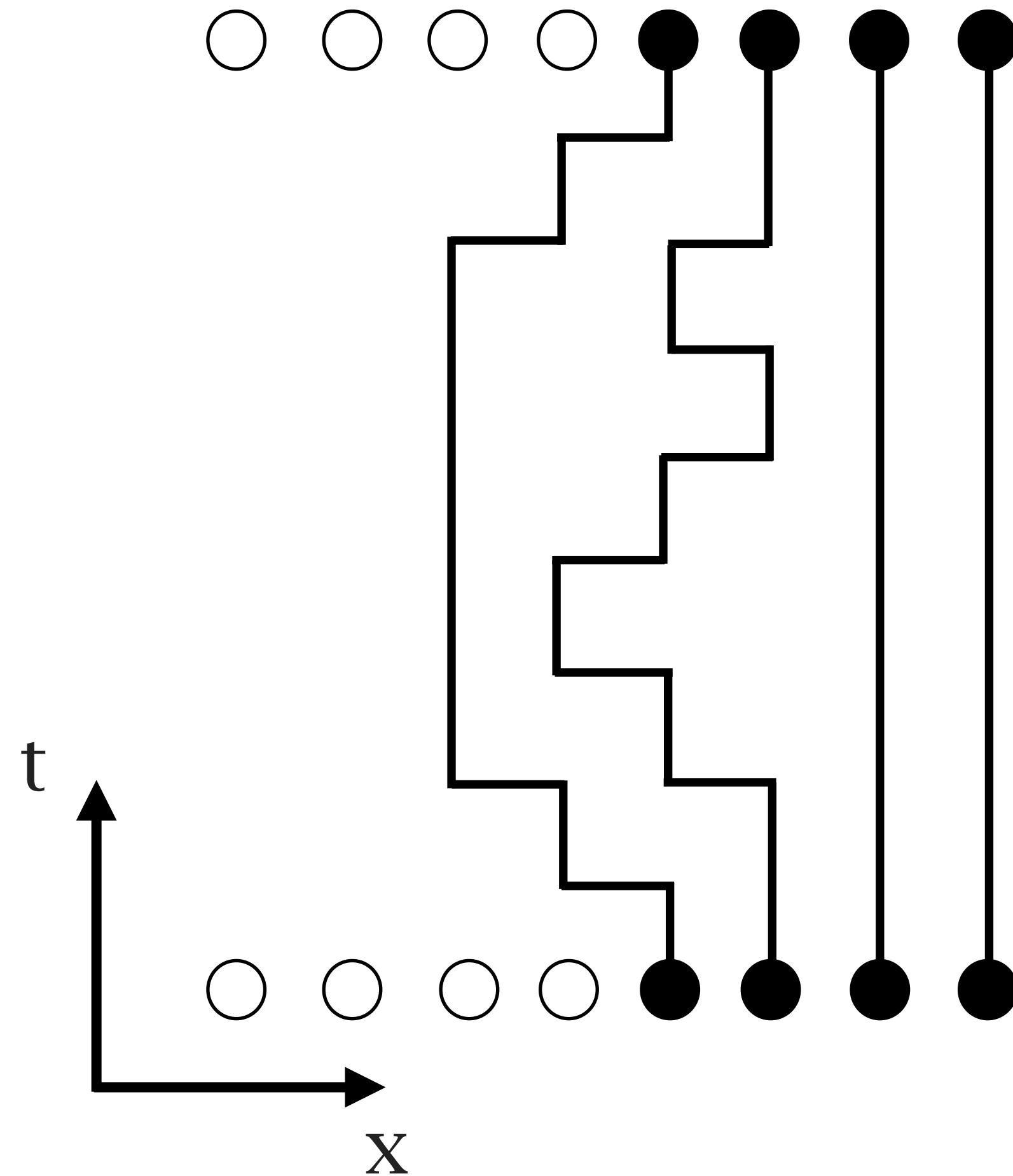


# From dimers to worldlines



# From lines to gas

Stéphan, 2020

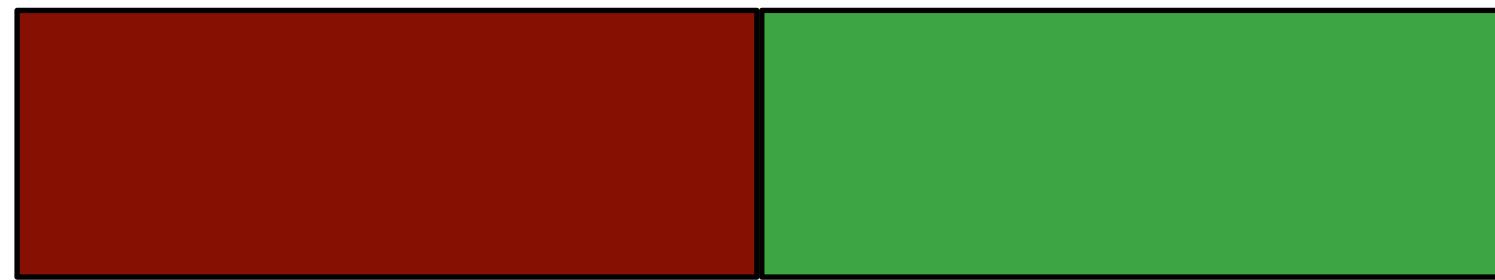


Green lines - particle trajectories

# Optimal gas flow

Empty

Filled



$t=2R$

?



$t=0$

Empty

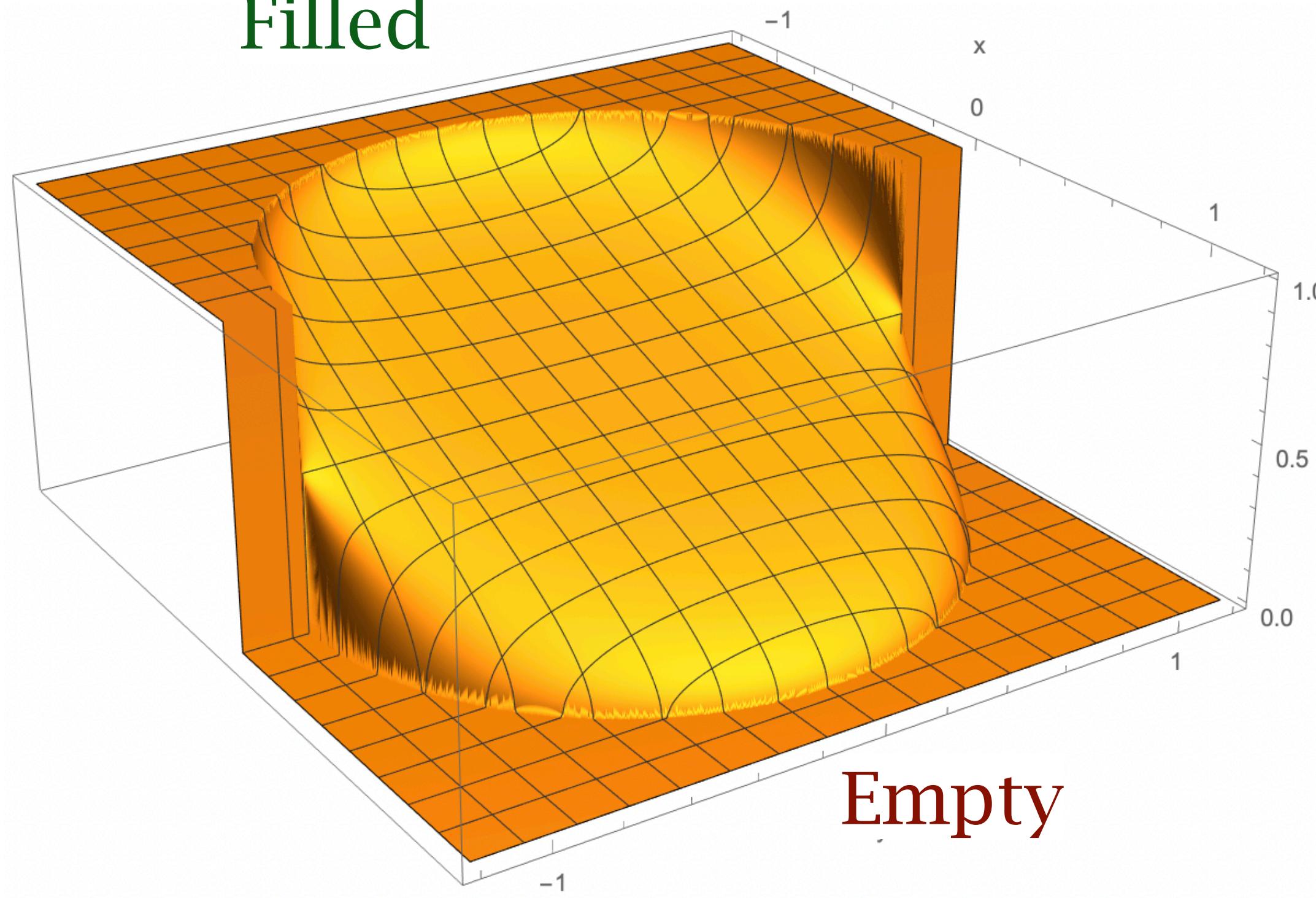
Filled

What is the optimal fluctuation of the gas in space and time so that at  $t=0$  and at  $t=2R$  the left half line is empty and the right one is filled?

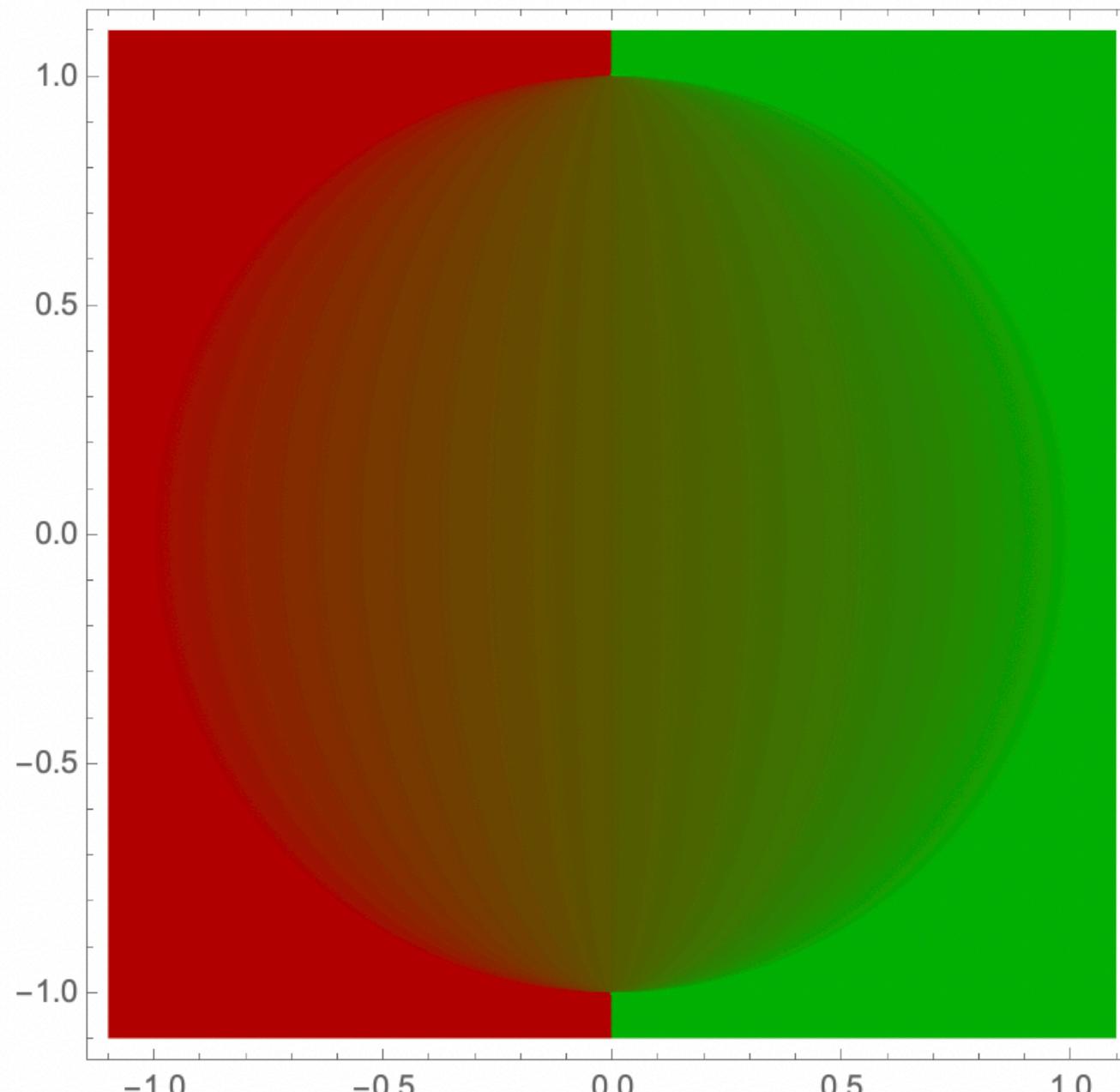
There are many particles ( $R$  is big), and one can think of gas as of continuous fluid and use hydrodynamic equations.

# The solution for the optimal motion

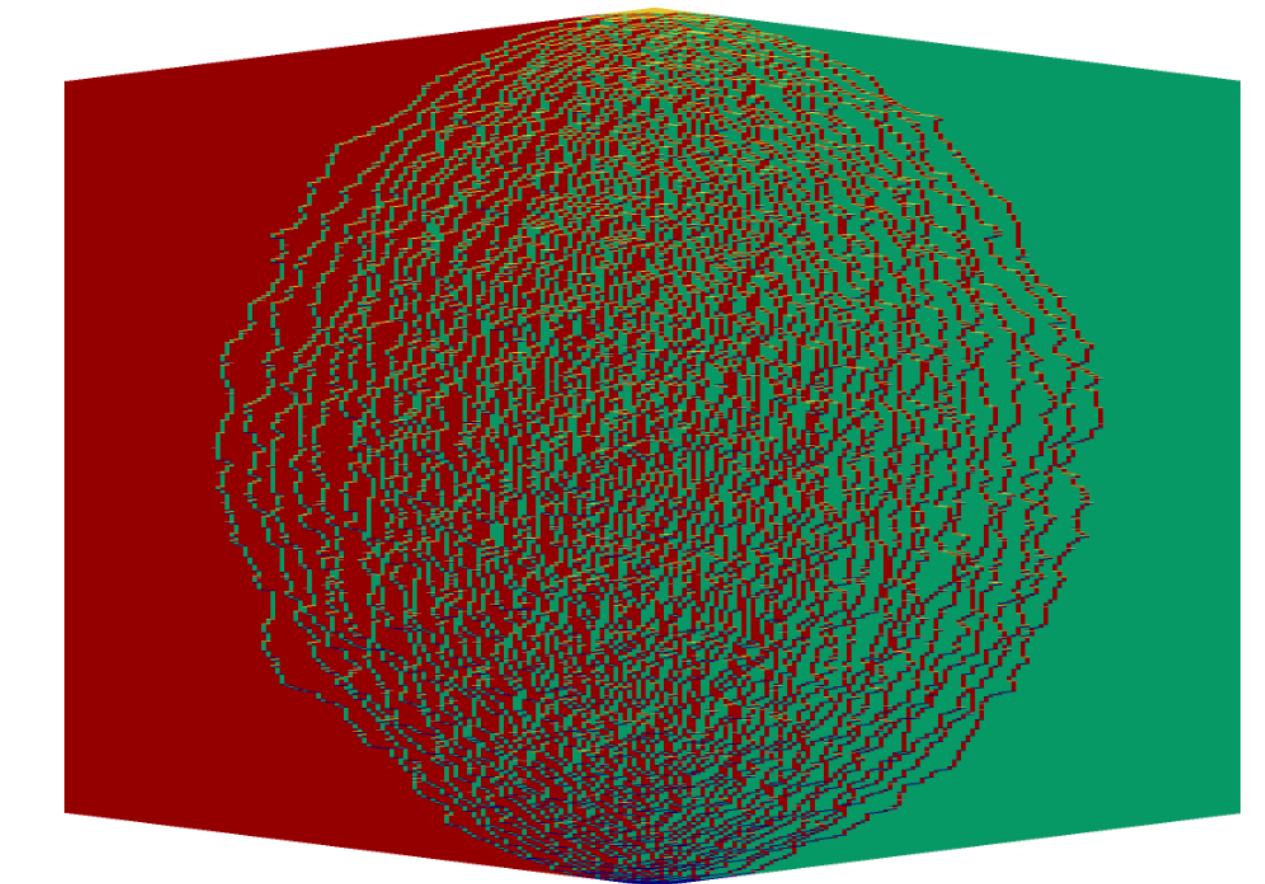
Filled



Empty



Empty



Filled

The density of the gas  
inside the Arctic circle

# Arctic Circle hydro solution

$$ik_\tau + \epsilon'(k)k_x = 0$$

$$z \equiv x + i\epsilon'(k)\tau = F(k)$$

$$\epsilon(k) = -\cos k$$

$$F(k) = R \cos k$$

$$x + i\tau \sin k = R \cos k$$
$$k = \pi\rho + iv$$

Free fermion hydro

General solution

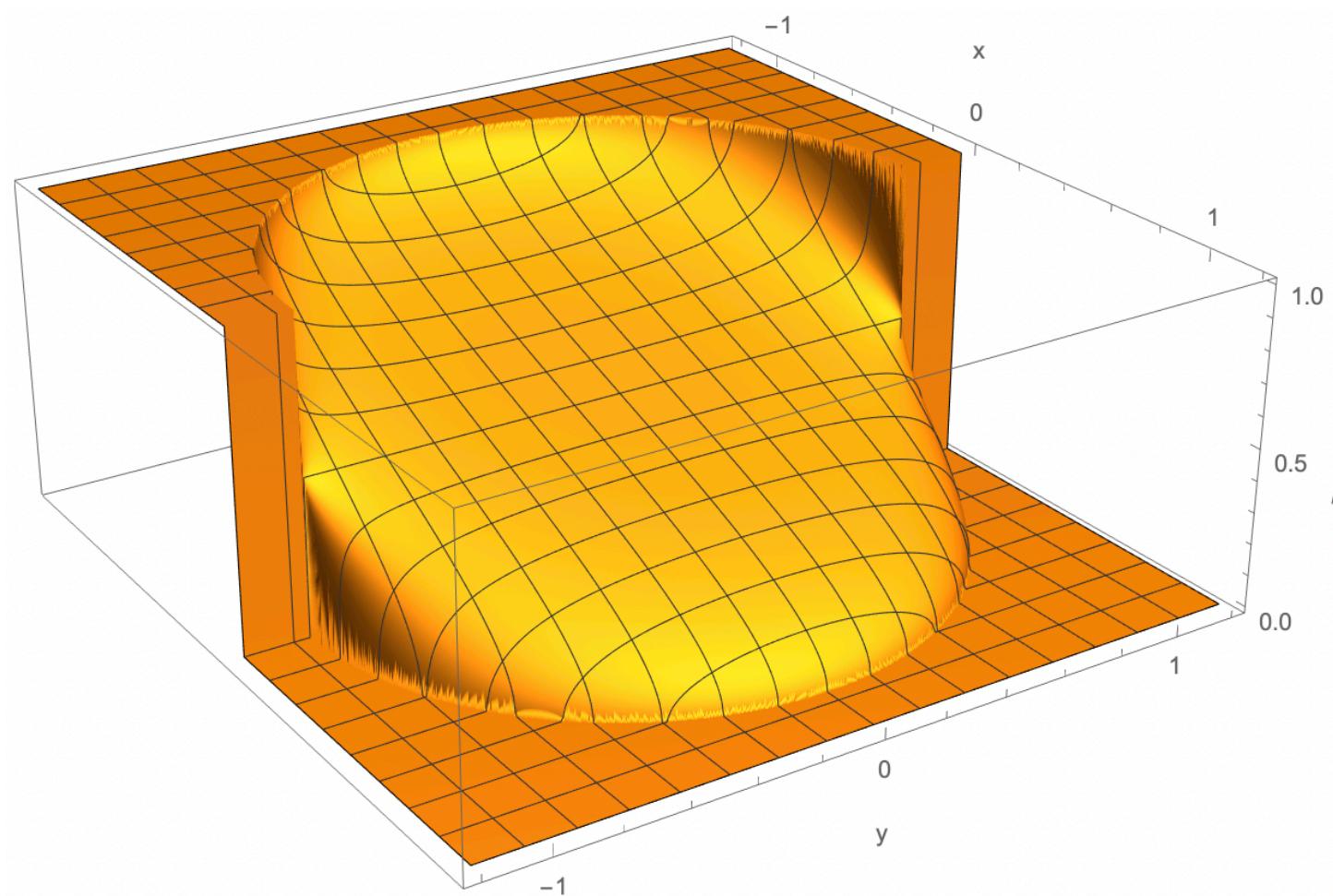
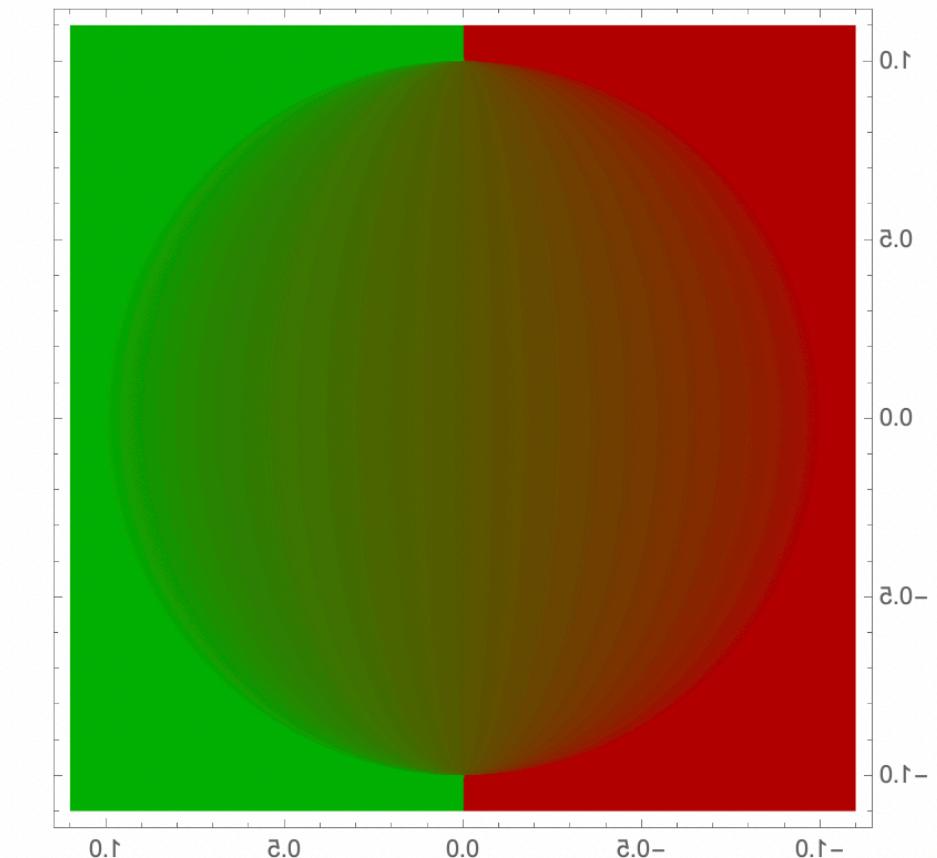
Dispersion

Arctic Circle solution

$$\tau = -R$$

$$x = Re^{i\pi\rho-v}$$

$$\rho = \theta(-x)$$



see: Allegra, Dubail, Stéphan, Viti, 2016 for general dispersion

# Limit shape phase transitions Merging Arctic Circles

with James Pallister and Dimitri Gangardt

based on: <https://arxiv.org/abs/2203.05269>

# Gross-Witten-Wadia model

Gross, Witten, 1980  
Wadia, 1980

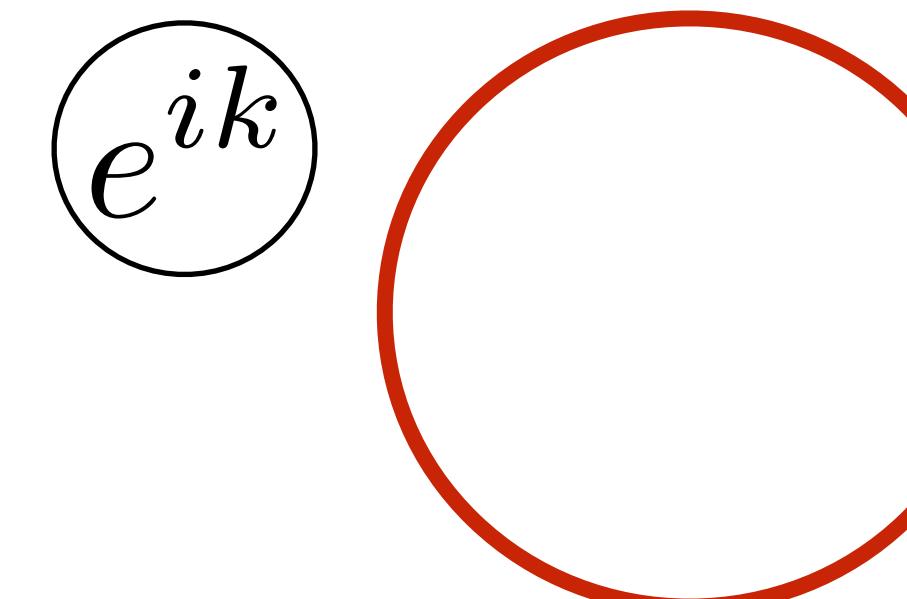
Partition function for 2d U(N) lattice gauge theory was reduced to:

$$Z_N = \int dU \exp \left\{ \frac{1}{\lambda} \text{Tr} (U + U^\dagger) \right\}$$

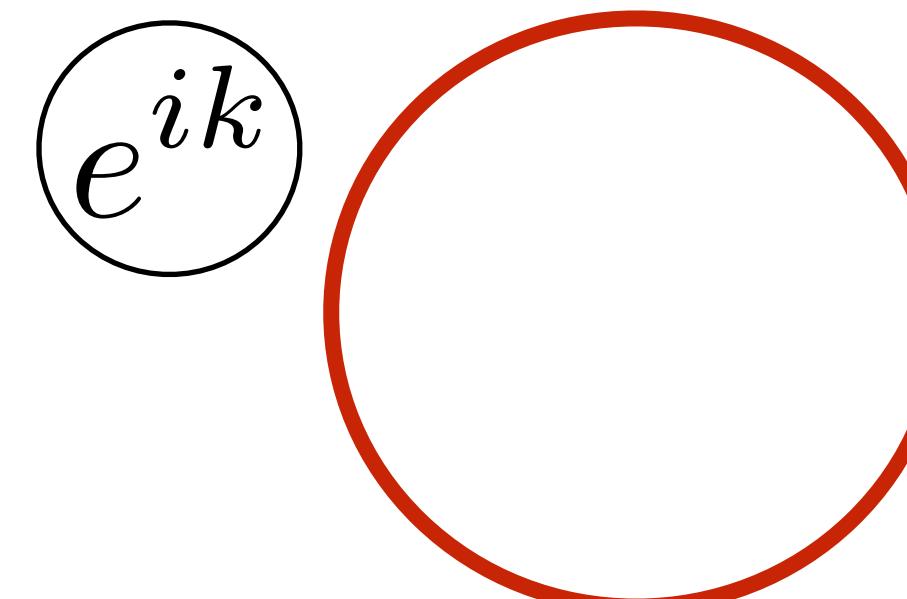
$$U = V \text{diag} \{ e^{ik_j} \} V^\dagger$$

$$Z_N = \prod_{i=1}^N \int_{-\pi}^{\pi} dk_i \prod_{i < j} \sin^2 \left( \frac{k_i - k_j}{2} \right) \exp \left( \frac{2N}{\lambda} \cos k_i \right)$$

Third-order weak-strong coupling phase transition at  
the t'Hooft coupling  $\lambda = 2$  in  $N \rightarrow \infty$  limit



weak coupling  $\lambda < 2$



strong coupling  $\lambda > 2$

# Electrostatic interpretation

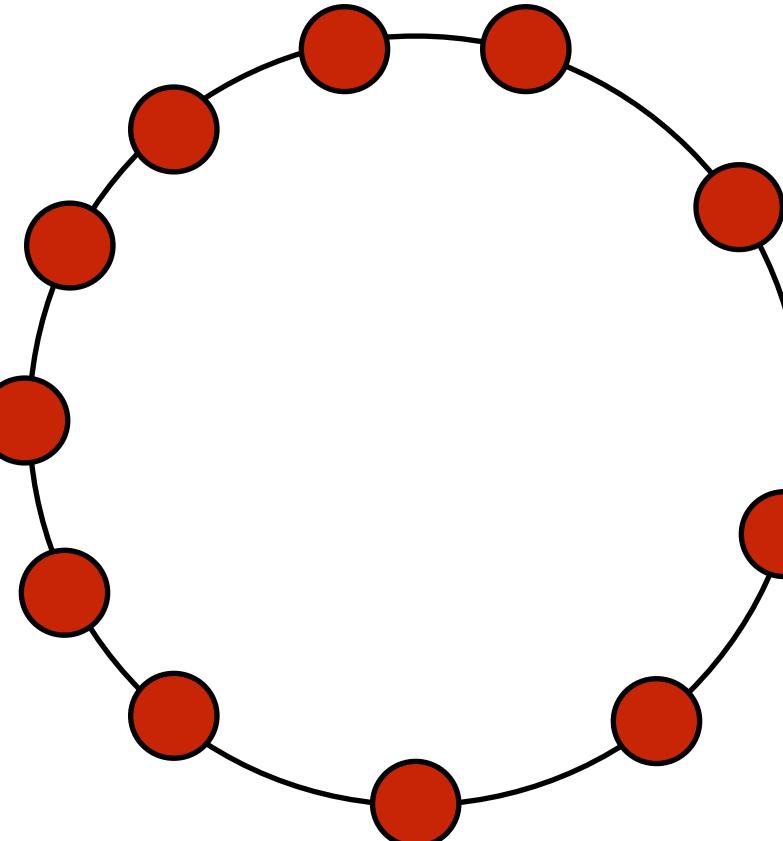
Partition function for charges on unit circle

$$Z_N = \prod_{i=1}^N \int_{-\pi}^{+\pi} dk_i e^{-E_N}$$

logarithmic repulsion

$$E_N = -2 \sum_{i < j} \ln |e^{ik_i} - e^{ik_j}| - \sum_i \frac{2N}{\lambda} \cos k_i$$

charges



external electric field  $E$



small  $\lambda$  means large  $E$

# Large N solution

Density of charges:

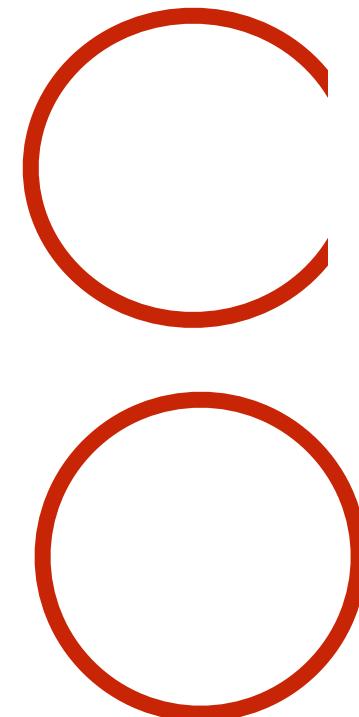
$$\sigma(k) = \frac{1}{N} \sum_i \delta(k - k_i)$$

$$2\pi\sigma(k) = \begin{cases} \frac{4}{\lambda} \cos \frac{k}{2} \sqrt{\frac{\lambda}{2} - \sin^2 \frac{k}{2}} & \lambda \leq 2 \\ 1 + \frac{2}{\lambda} \cos k & \lambda \geq 2 \end{cases}$$

$$-E(\lambda) = N^2 \begin{cases} \frac{2}{\lambda} + \frac{1}{2} \ln \frac{\lambda}{2} - \frac{3}{4} & \lambda \leq 2 \\ \frac{1}{\lambda^2} & \lambda \geq 2 \end{cases}$$

$\frac{\partial^3 E}{\partial \lambda^3}$  is discontinuous at  $\lambda = 2$

Gross, Witten, 1980  
Wadia, 1980



# Free fermions

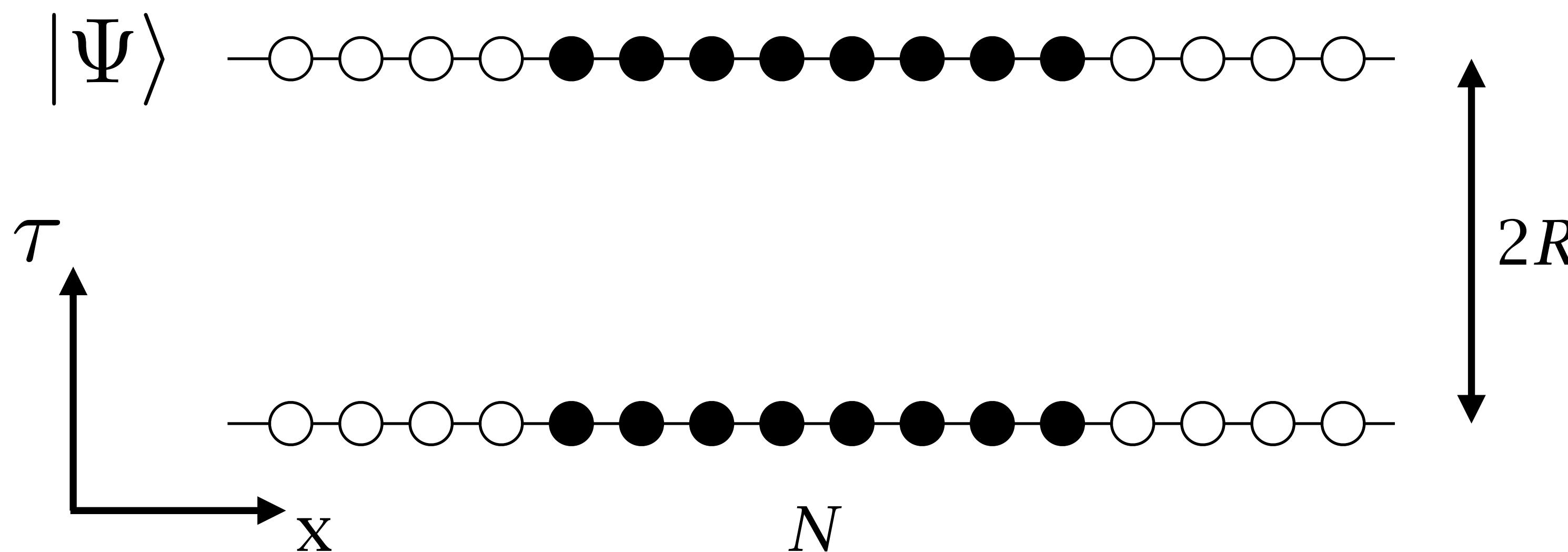
Hamiltonian of free fermions

$$H = \int_{-\pi}^{\pi} \frac{dk}{2\pi} \varepsilon(k) c^\dagger(k) c(k)$$

Diagonal in k-space

$$Z_N = \langle \Psi | e^{-2RH} | \Psi \rangle$$

$|\Psi\rangle$  Diagonal in x-space



$$N, R \rightarrow \infty$$

$$\lambda = \frac{N}{R} \quad - \text{fixed}$$

# Partition function

$$Z_N(R) = \langle N | e^{-2RH} | N \rangle \quad |N\rangle = c_1^\dagger \dots c_N^\dagger |0\rangle$$



$$Z_N(R) = \frac{1}{N!} \int_{-\pi}^{\pi} \prod_{i=1}^N \frac{dk_i}{2\pi} \langle N | \{k_i\} \rangle e^{-2R \sum_l \varepsilon(k_l)} \langle \{k_i\} | N \rangle$$

$$Z_N = \frac{1}{N!} \int \frac{d^N k}{(2\pi)^N} |\Delta(e^{ik})|^2 e^{-2R \sum_l \varepsilon(k_l)}$$

$$\Delta(e^{ik}) = \prod_{i < j} (e^{ik_i} - e^{ik_j}) \quad \text{- Vandermonde determinant}$$

# Electrostatic interpretation

for the dispersion  $\varepsilon(k) = -\cos k$  and  $\lambda = \frac{N}{R}$

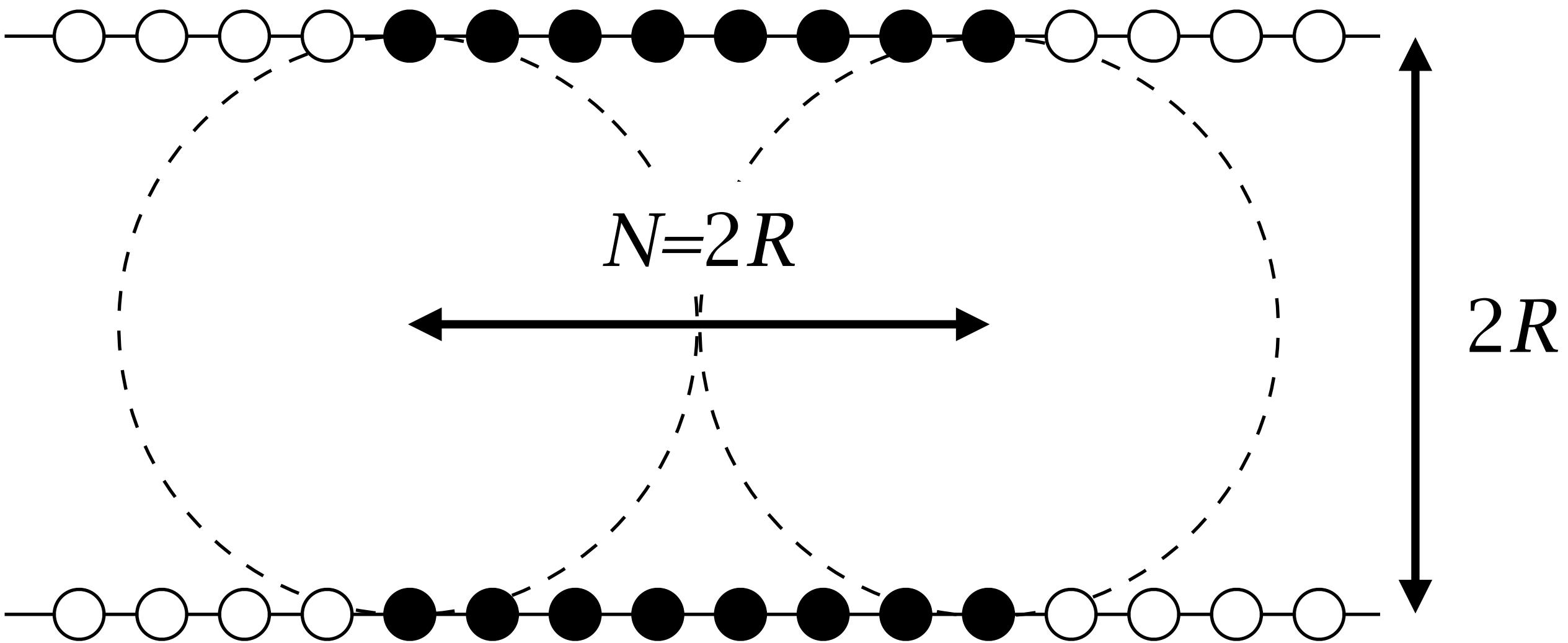
$$Z_N = \prod_{i=1}^N \int_{-\pi}^{+\pi} dk_i e^{-E_N}$$

$$E_N = -2 \sum_{i < j} \ln |e^{ik_i} - e^{ik_j}| - \sum_i \frac{2N}{\lambda} \cos k_i$$

$\lambda = 2$  - phase transition (as in Gross-Witten-Wadia model)

What is the picture of this transition in spacetime?

# Spacetime picture



$N, R \rightarrow \infty$

$$\lambda = \frac{N}{R} \quad \text{- fixed}$$

Arctic circles touch at  $\lambda = 2$

# From k-space to spacetime

Density of charges

$$\sigma(k) = \frac{1}{N} \sum_i \delta(k - k_i)$$

$$2\pi\sigma(k) = \begin{cases} \frac{4}{\lambda} \cos \frac{k}{2} \sqrt{\frac{\lambda}{2} - \sin^2 \frac{k}{2}} & \lambda \leq 2 \\ 1 + \frac{2}{\lambda} \cos k & \lambda \geq 2 \end{cases}$$

$$F(\pi\lambda\sigma(k), k) = 0$$

$$F(z, k) = \left( z - \frac{\lambda}{2} - \cos k \right) \left( z + \frac{\lambda}{2} + \cos k \right) + \left( 1 - \frac{\lambda}{2} \right)^2 \Theta(2 - \lambda)$$

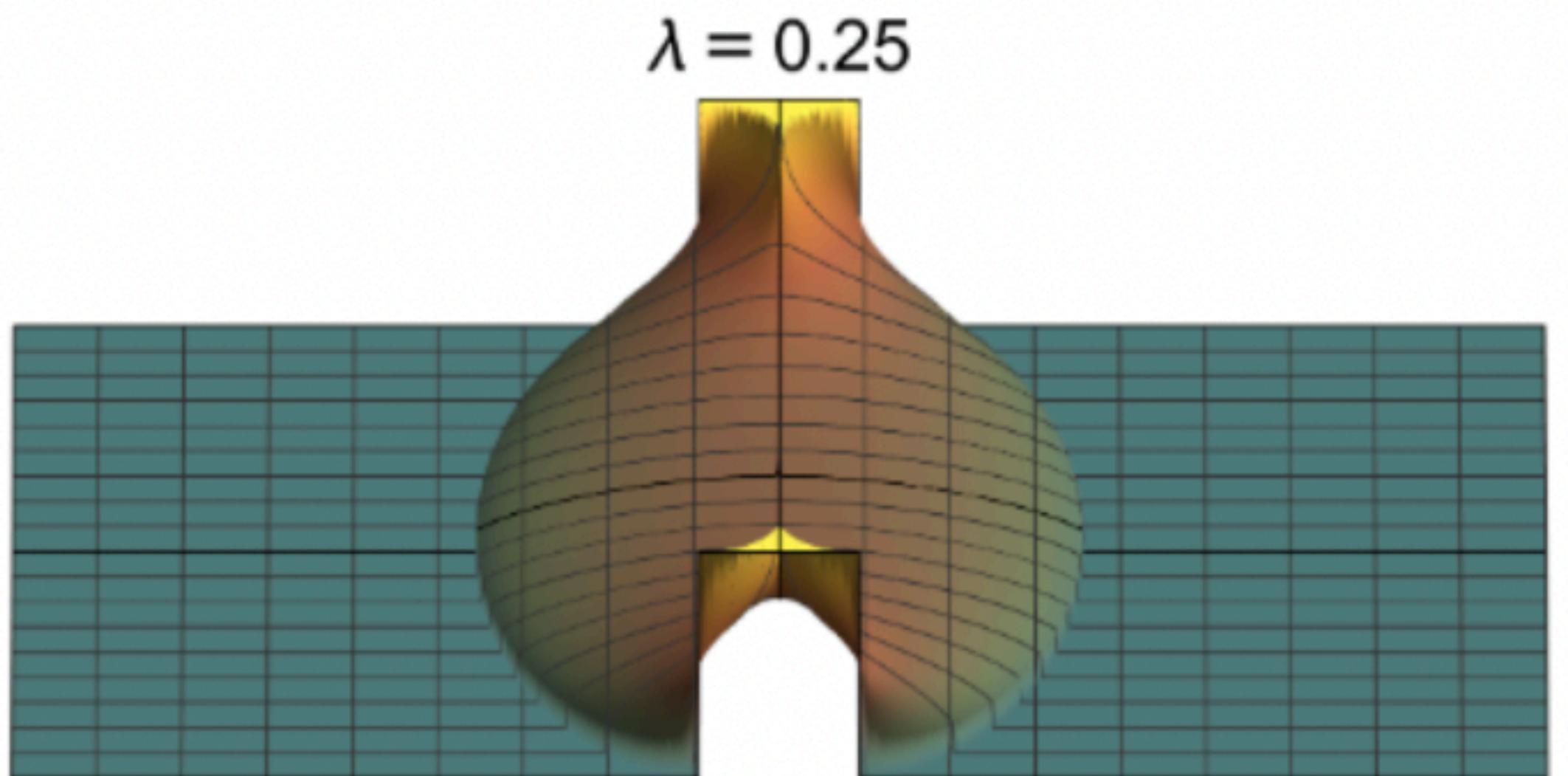
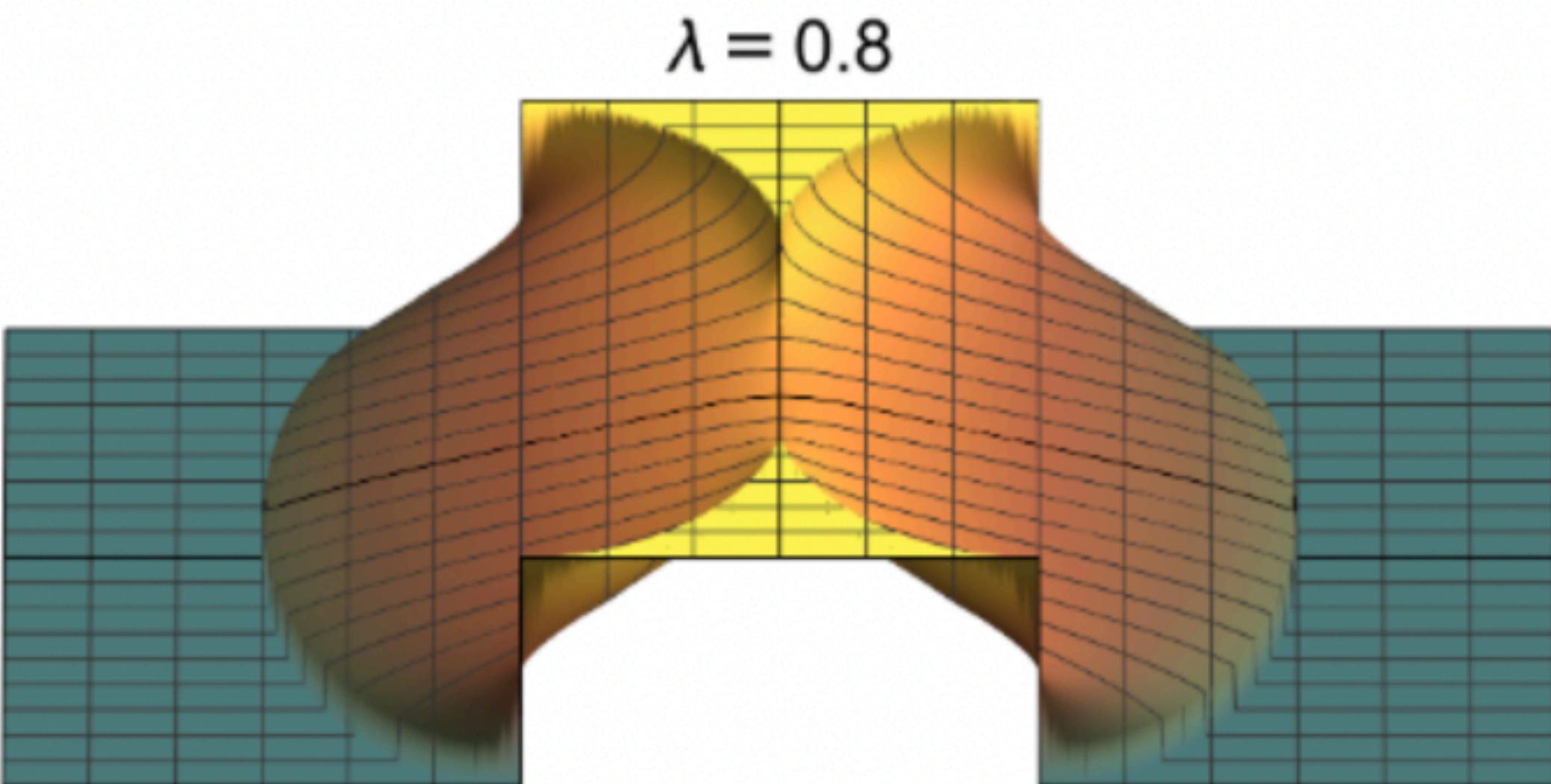
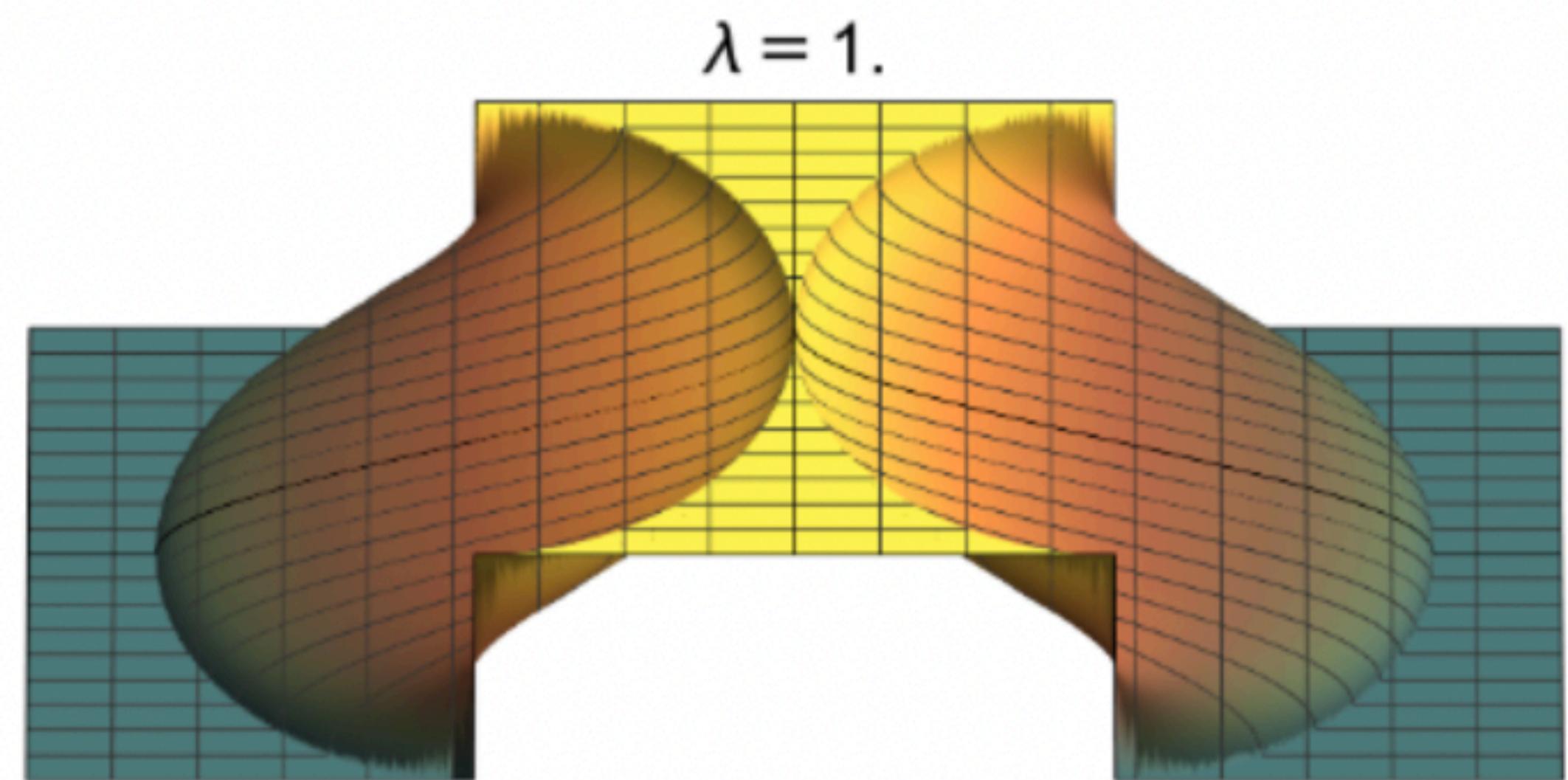
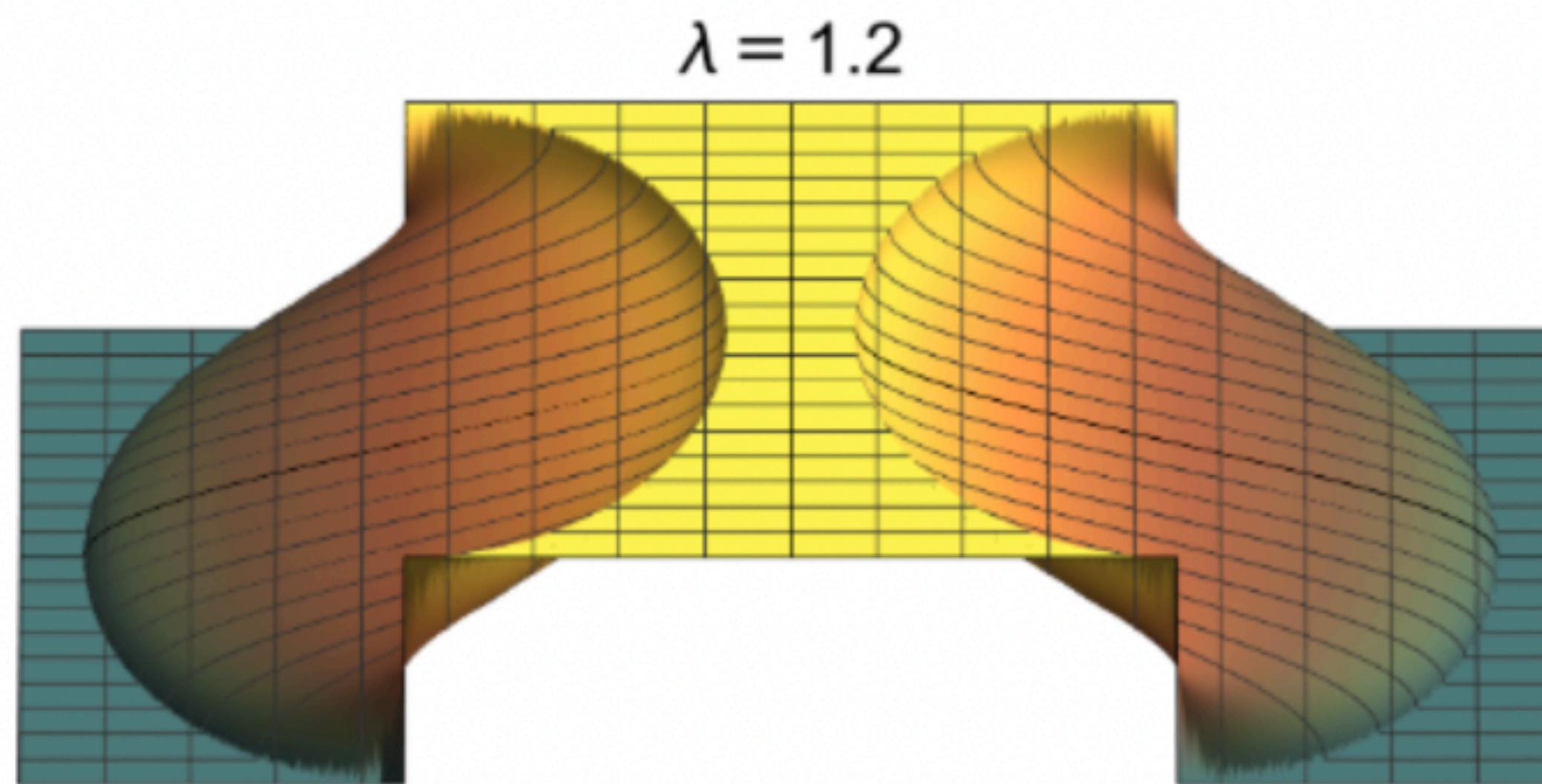
complex curve  
encodes large N  
solution!

$$F(x + it \sin k, k) = 0 \longrightarrow k(x, t) \text{ - complex function of } x, t$$

$$k(x, t) = \pi\rho(x, t) + iv(x, t)$$

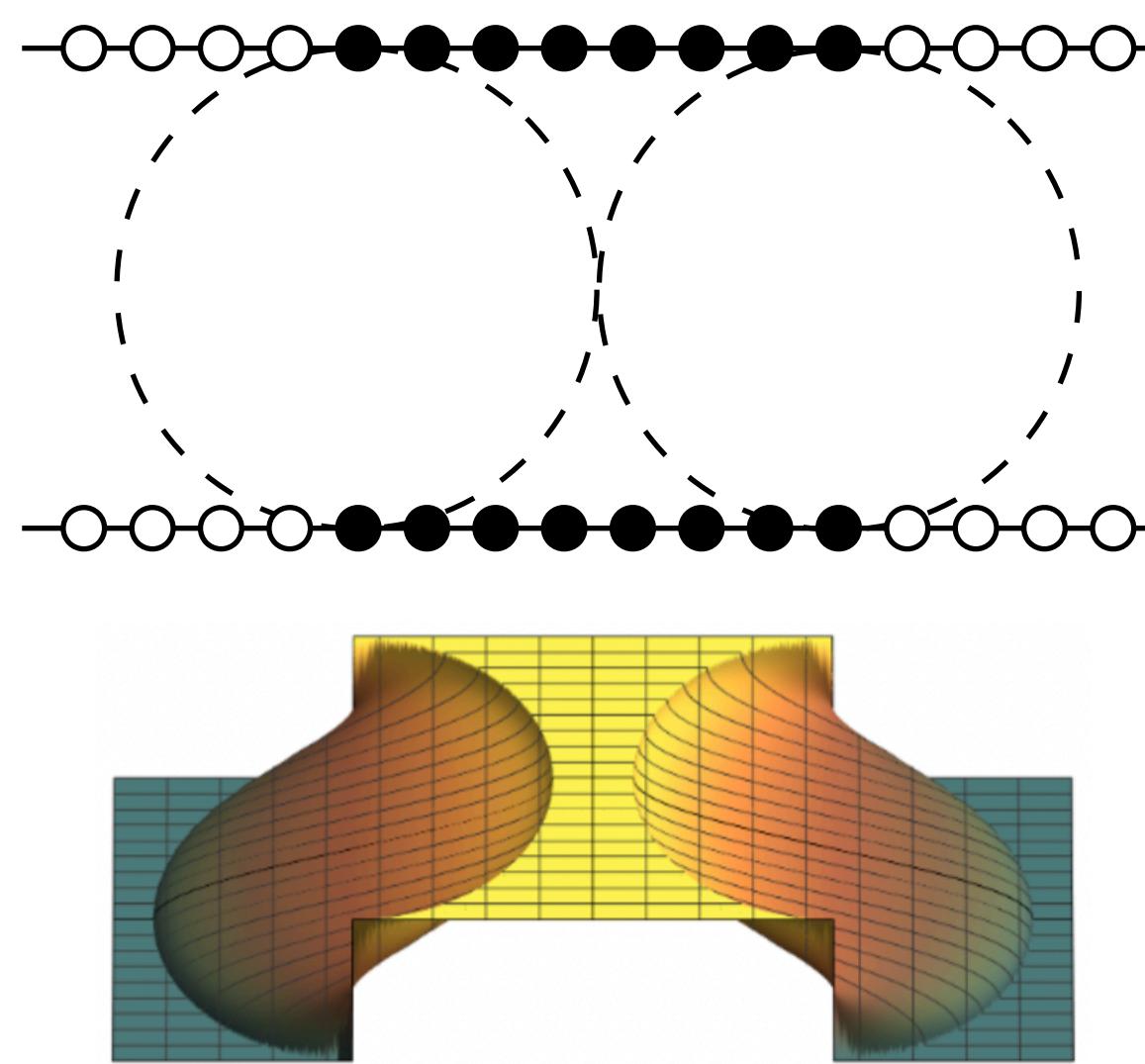
gives density and velocity of  
fluid for optimal fluctuation

# Density profiles



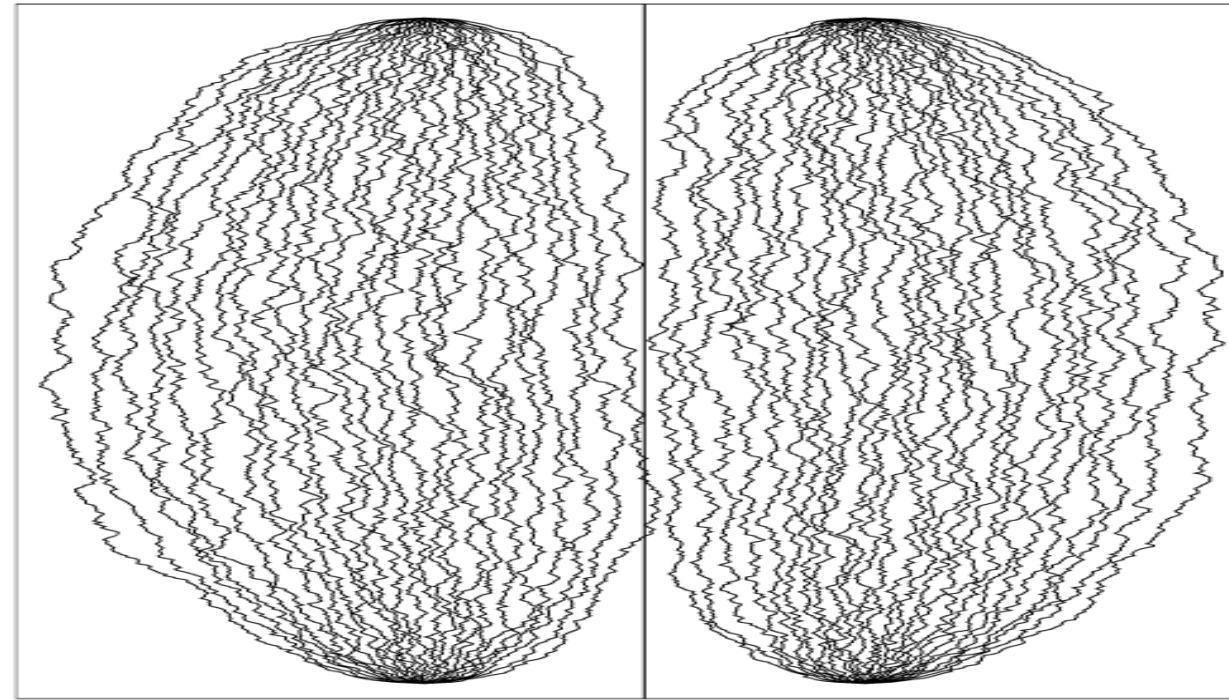
# Other equivalent systems

1.

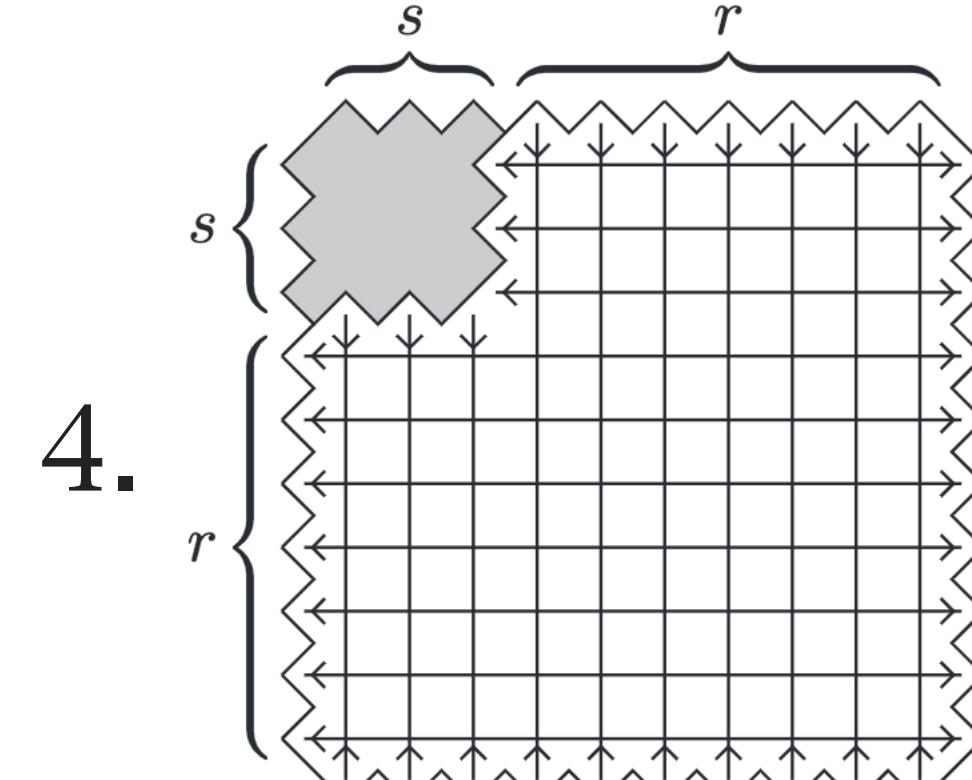


1. Free fermion (+hydro)
2. Nonintersecting random walkers
3. Domino tilings
4. Six-vertex models with domain wall bc
5. Random Young diagrams

2.

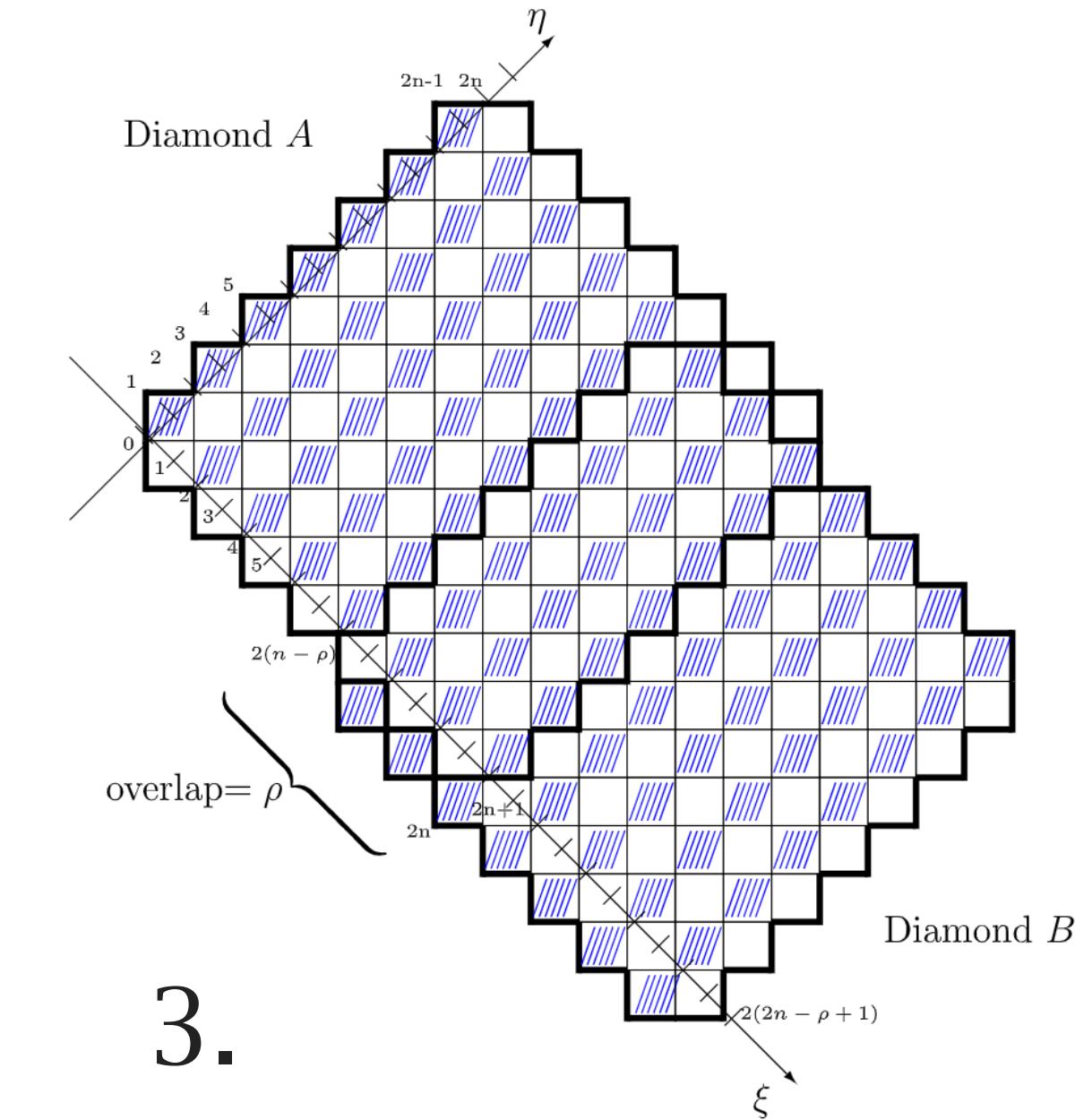


Adler, Ferrari, van Moerbeke, 2013

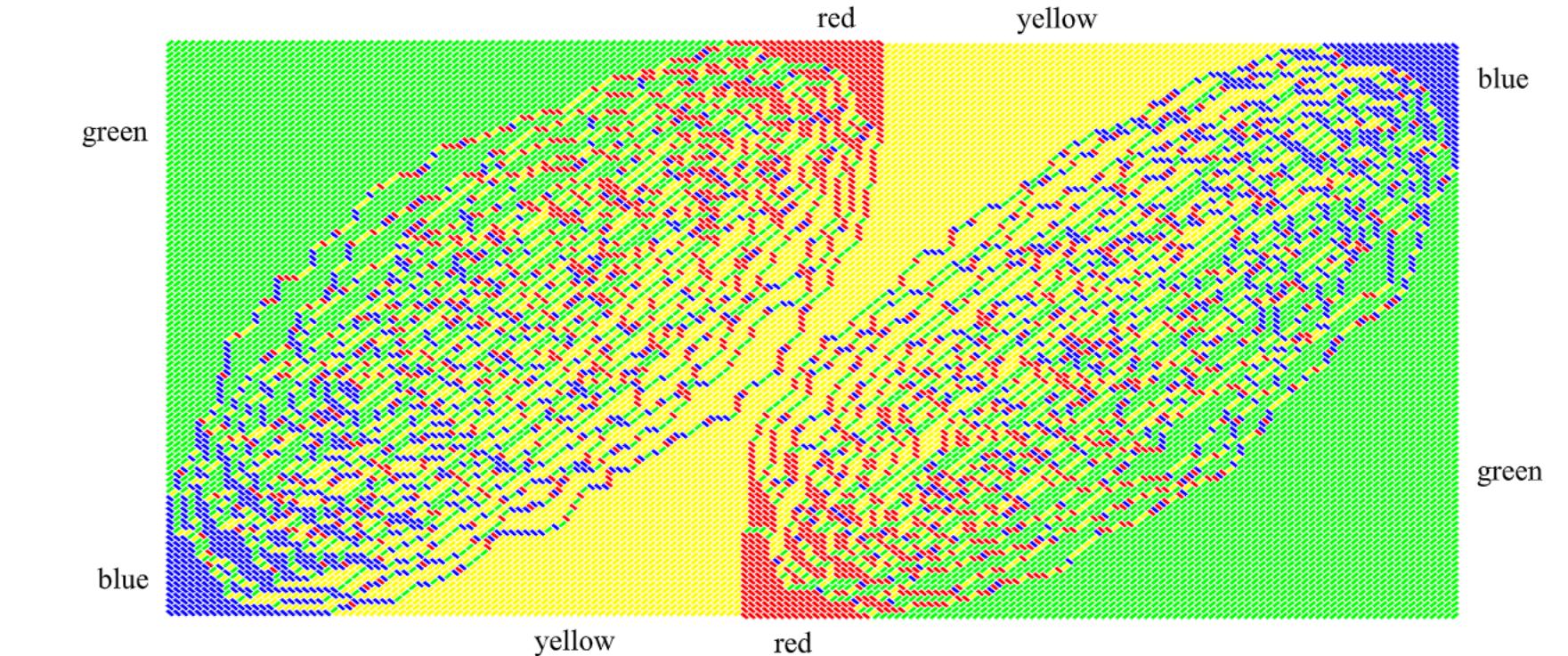


Colomo, Pronko, 2013

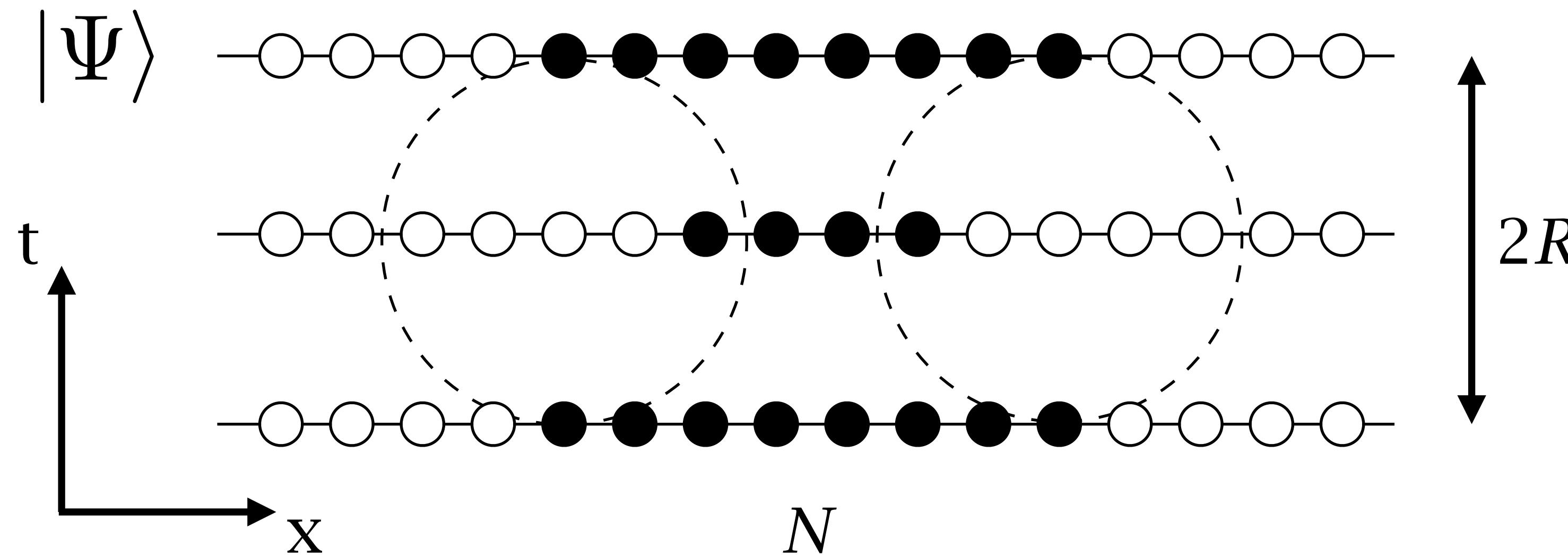
Adler, Johanson, van Moerbeke, 2014



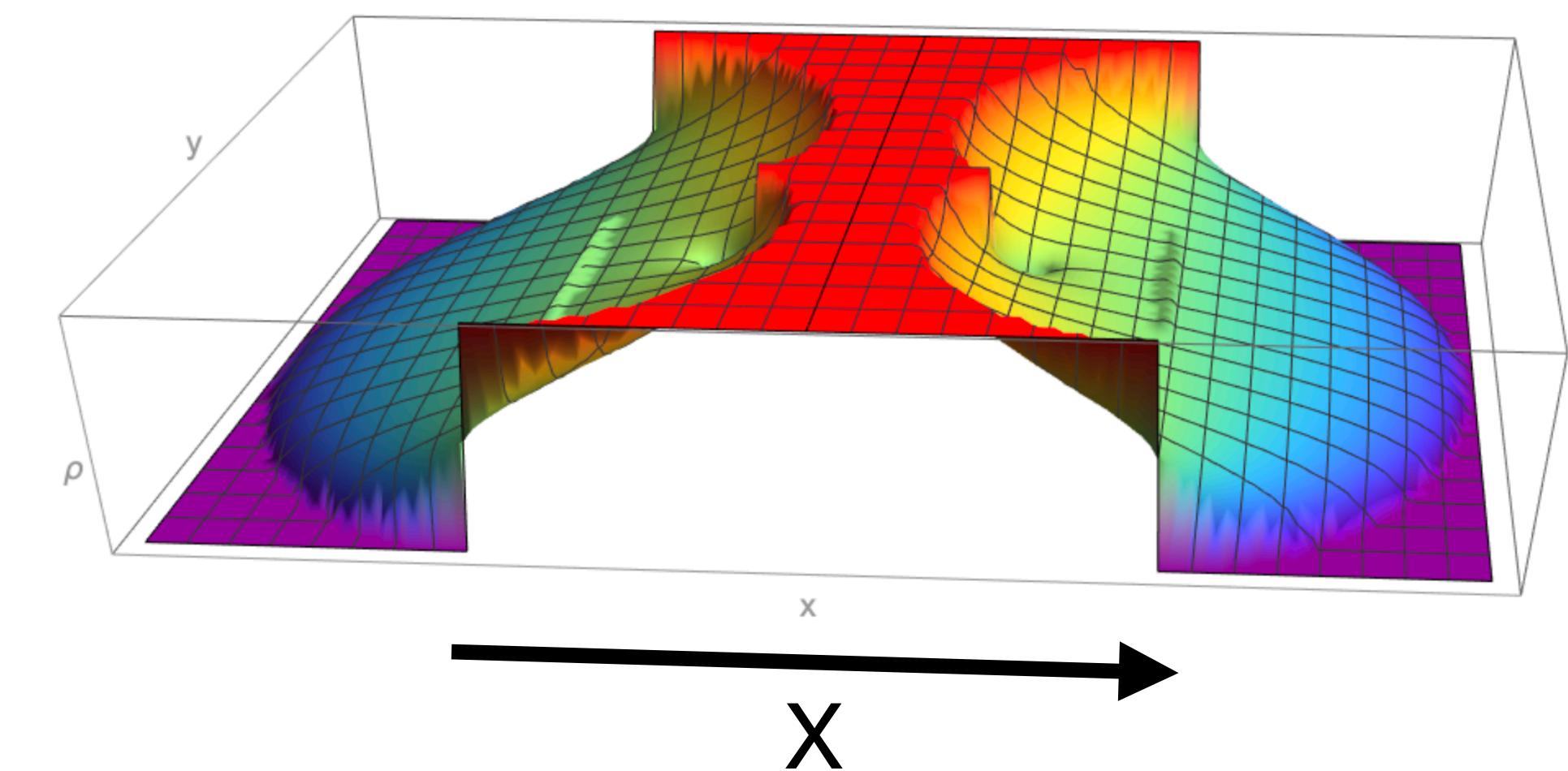
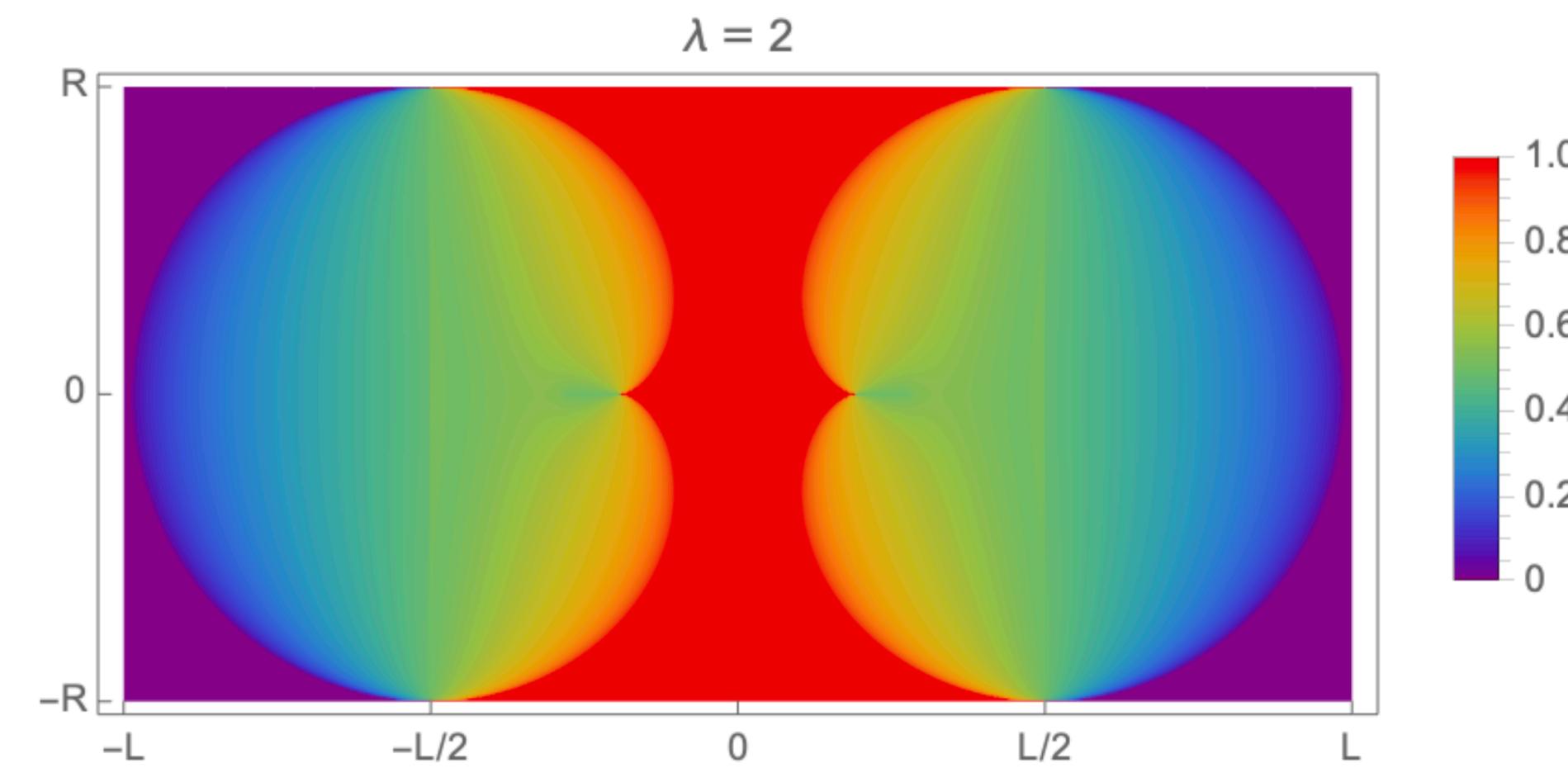
3.



# Other boundary conditions

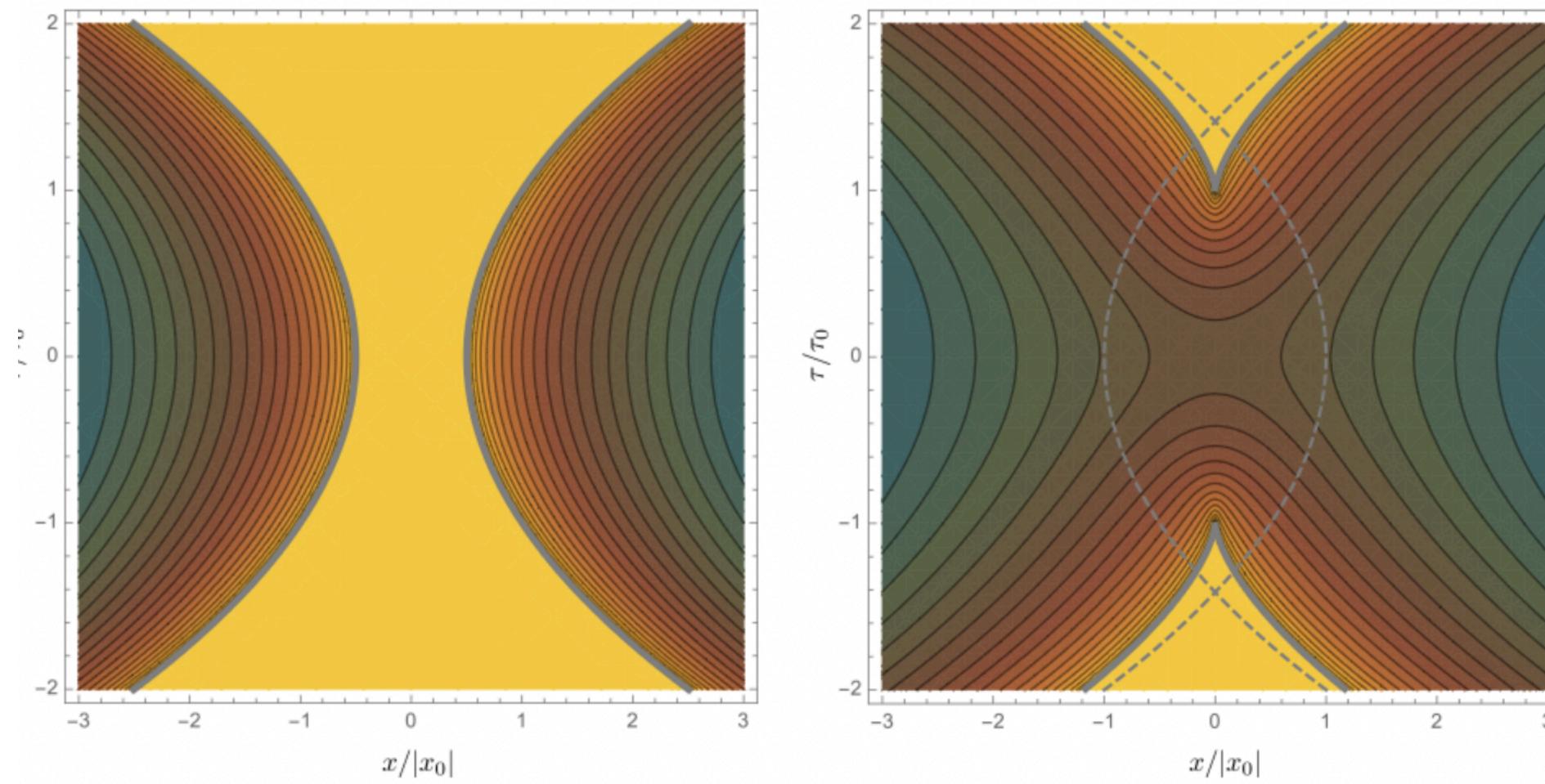
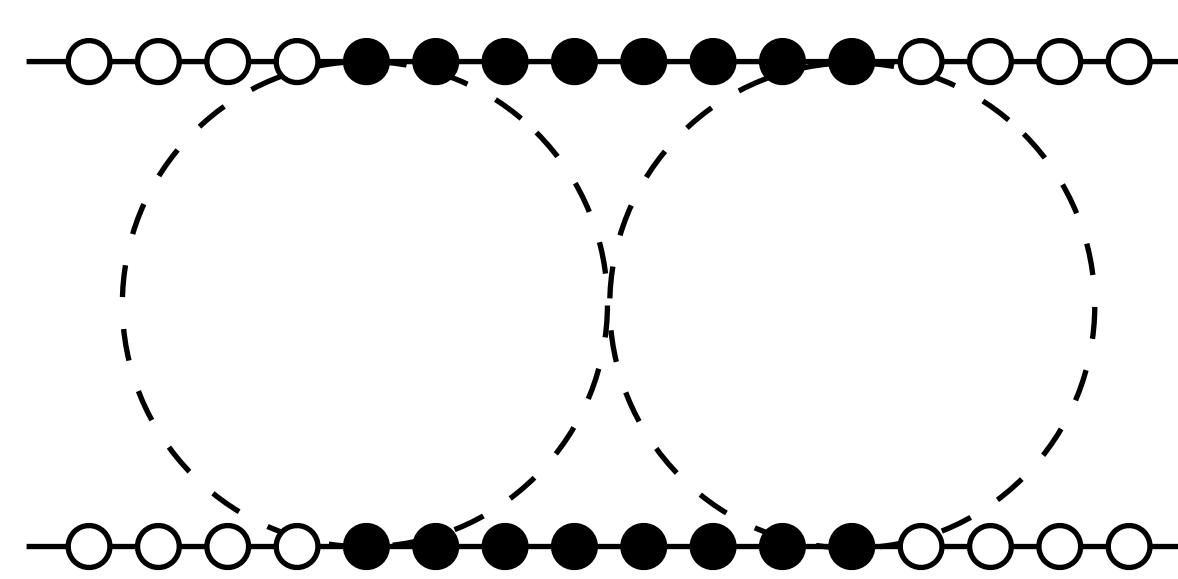


Insertion of  
“Fullness” at  $t=0$



# On the universality of the third-order phase transition

# Interacting problems

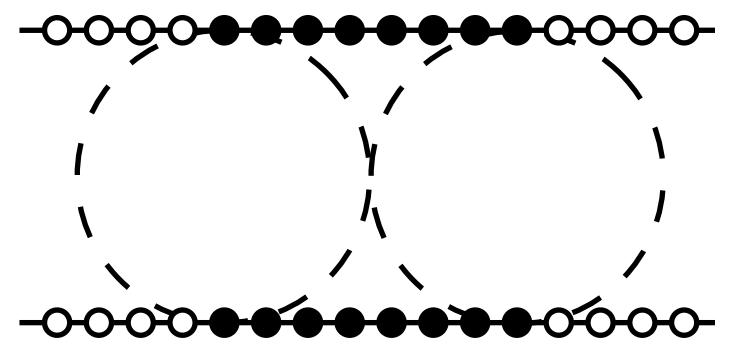


Density profiles

Lattice fermions with interactions or XXZ-type models

- Hydrodynamic approach is not limited to free or even integrable systems
- The difficulty is in solving PDEs
- At the point of Arctic Circles merging the density of particles (or holes) is small
- Fermion interactions are not important for the transition
- The transition is of the third order and fluid profiles in the vicinity of the merging point the shape are universal

# Universal profiles



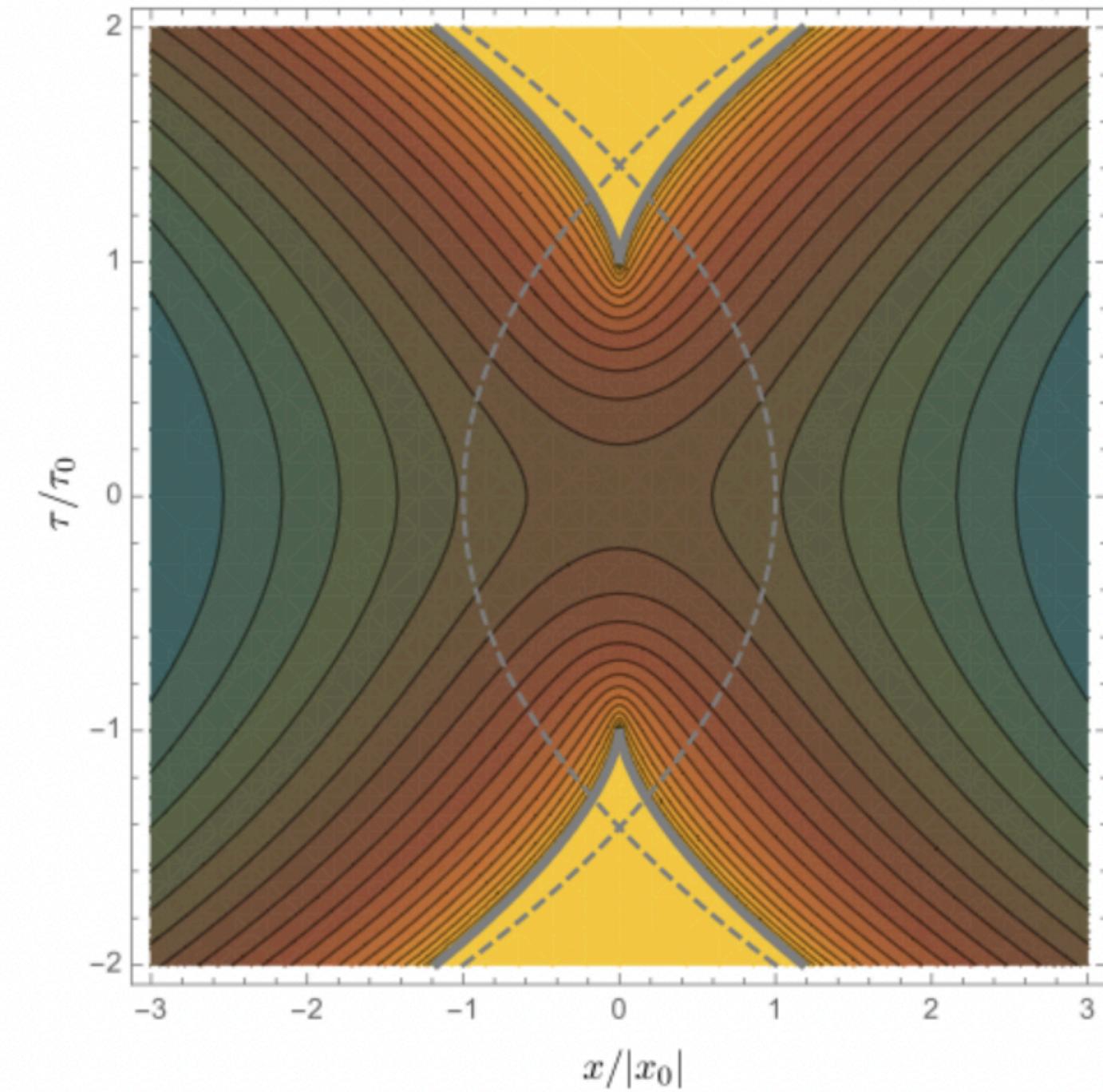
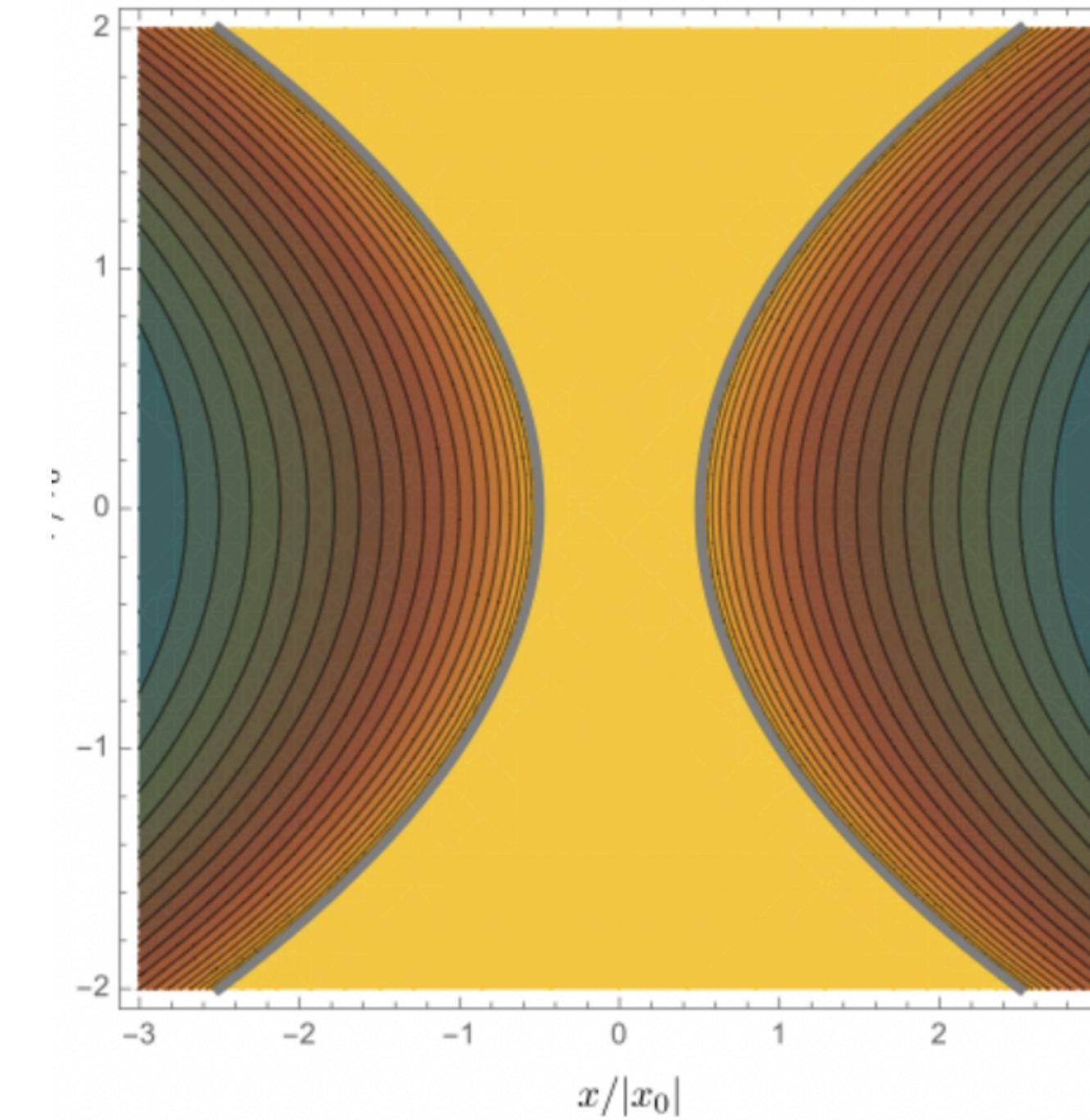
Lattice fermions with interactions or XXZ-type models. P and T symmetries.

$$X^2 = (q^2 + \delta_1)(q^2 + \delta_2)$$

$$X = x - i\tau q \quad q = \pi\rho - iv$$

$$X^2 = \begin{cases} (q^2 + \delta)^2, & \delta > 0, \\ q^4, & \delta = 0, \\ q^2(q^2 + 2\delta), & \delta < 0. \end{cases}$$

$\delta > 0$



$\delta < 0$

# Free energy estimate

$$\delta\rho \sim q \sim \delta^{1/2}$$

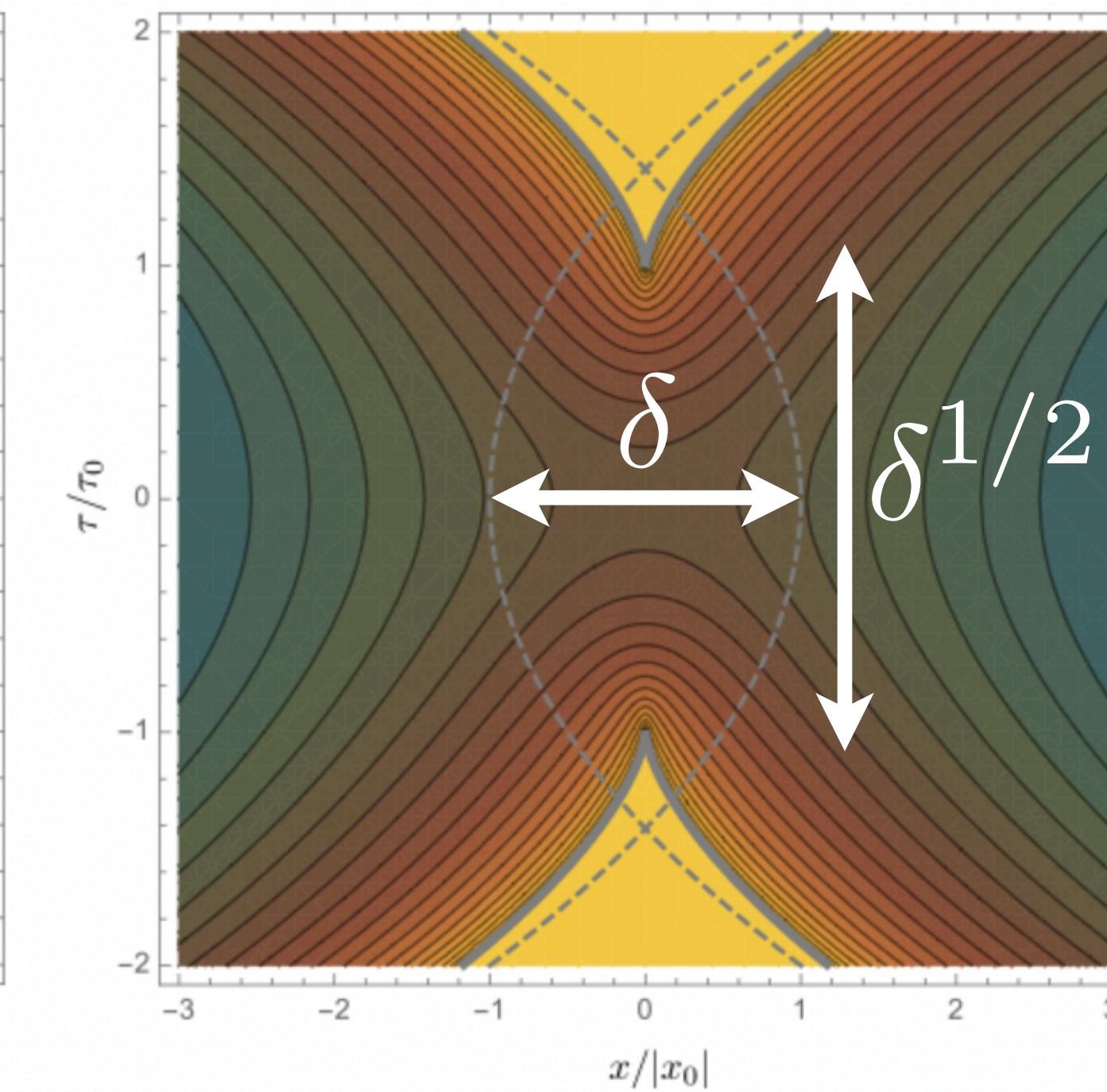
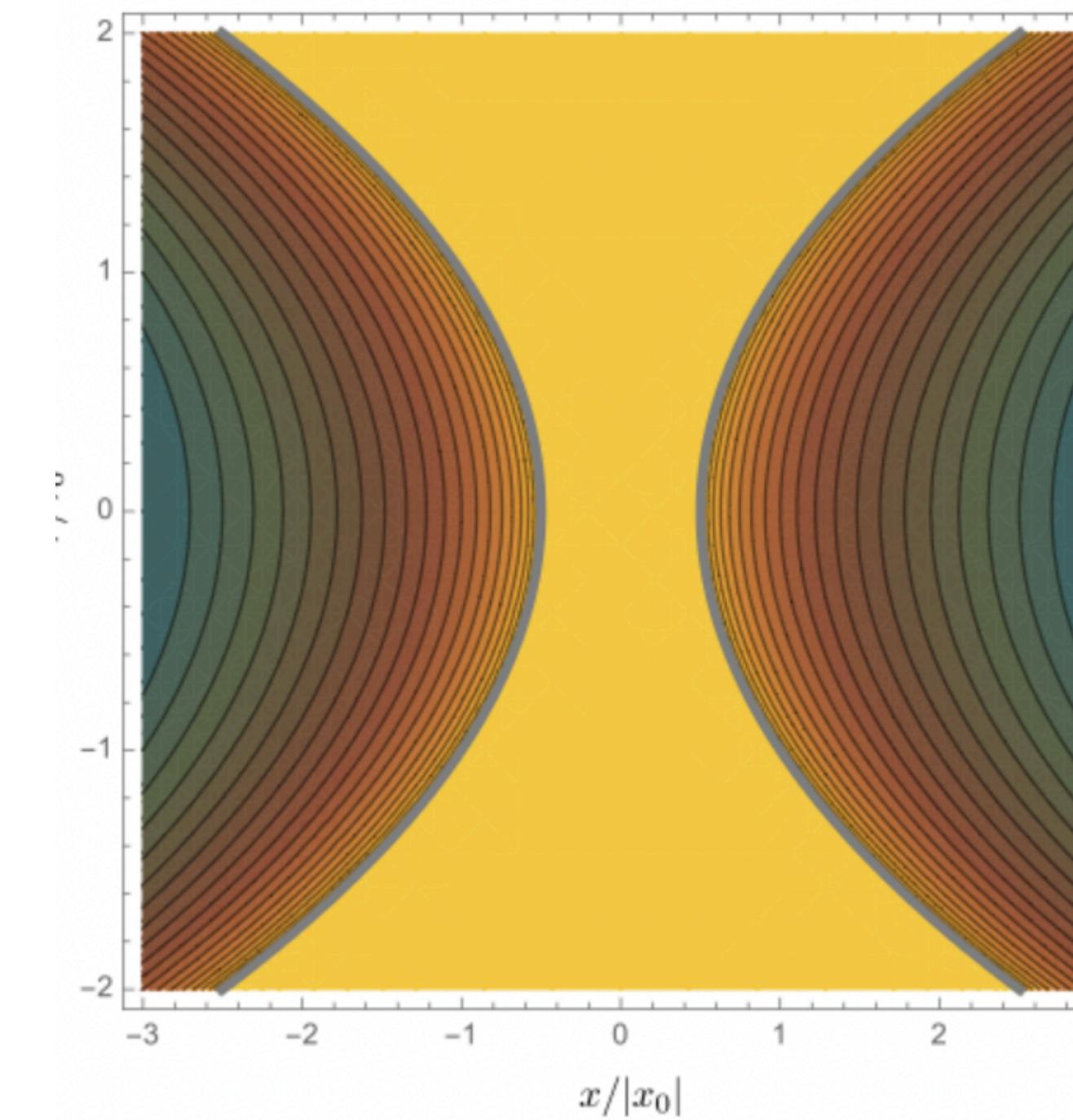
$$\delta E \sim \delta\rho^3 \sim \delta^{3/2}$$

$$\delta S \sim \delta\tau \delta x \delta E \sim \delta^{1/2} \delta \delta^{3/2} \sim \boxed{\delta^3}$$

Discontinuity of  
third derivative

Third order!

$$\delta > 0$$



$$\delta < 0$$

# Conclusions

- Limit shape problems occur in many random/statistical systems
- The Gross-Witten-Wadia transition has been mapped to the problem of Arctic circle merging in dimer and free fermion models.
- The hydrodynamic picture can be generalized for interacting particles.
- Universal (stable to interactions) density and velocity profiles corresponding to the merging of Arctic circles have been found. They are given by solutions of fermion hydrodynamic equations.
- It is conjectured that the transition is of the third order even in the presence of interactions (protected by P and T symmetries).

# Remarks

- Only mean-field (large N) limit has been considered in this talk, not fluctuations.
- In proper scaling limits one can study Tracy-Widom, Airy, tachnoid, ... statistics.
- A lot of connections with algebraic geometry, combinatorics, and representation theory.
- Integrable interacting models (6-vertex etc).

# Reviews

1. R. Kenyon, Lectures on Dimers, [arXiv:0910.3129v1](https://arxiv.org/abs/0910.3129v1) [math.PR]
2. J.-M. Stéphan, Extreme boundary conditions and random tilings, [arXiv:2003.06339v2](https://arxiv.org/abs/2003.06339v2) [cond-mat.stat-mech]
3. YouTube: (4 lectures in Russian) Н. Решетихин, “Предельные формы в статистической механике” (N. Reshetikhin, “Limit shapes in statistical mechanics”).
4. YouTube (popular): Mathologer (B. Polster), “The ARCTIC CIRCLE THEOREM or Why do physicists play dominoes?”.  
