

Towards the N³LO evolution of the parton distribution functions

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Theory Challenges in the Precision Era of the Large Hadron Collider

Based on 2203.11181, 2302.07593, 2307.04158, ...

ongoing collaboration with

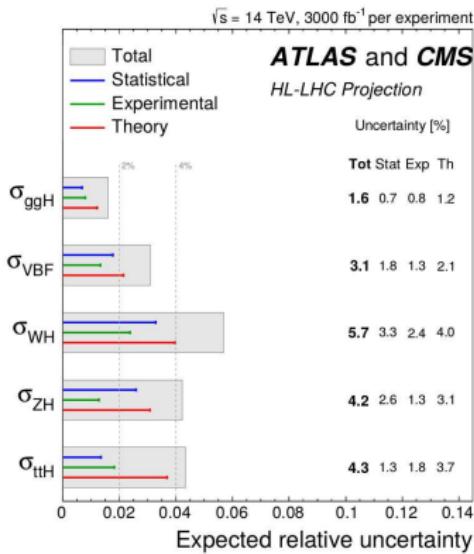
Franz Herzog, Sven Moch, Andreas Vogt and Andrea Pelloni



THE UNIVERSITY *of* EDINBURGH

1%-level phenomenology at the LHC

The HL-LHC will push the precision frontier to the %-level
E.g. Higgs measurements



- Systematic uncertainties (e.g. luminosity determination, resolution) around 1%
- Statistical errors reduced by 20-fold increase of data collected
- Theory will be the most important source of errors.

CERN Yellow report

Theory errors

| | Q [GeV] | $\delta\sigma^{\text{N}^3\text{LO}}$ | $\delta(\text{scale})$ | $\delta(\text{PDF-TH})$ |
|-------------------------------------|-----------|--------------------------------------|------------------------|-------------------------|
| $gg \rightarrow \text{Higgs}$ | m_H | 3.5% | +0.21% -2.37% | $\pm 1.2\%$ |
| $b\bar{b} \rightarrow \text{Higgs}$ | m_H | -2.3% | +3.0% -4.8% | $\pm 2.5\%$ |
| NCDY | 30 | -4.8% | +1.53% -2.54% | $\pm 2.8\%$ |
| | 100 | -2.1% | +0.66% -0.79% | $\pm 2.5\%$ |
| CCDY(W^+) | 30 | -4.7% | +2.5% -1.7% | $\pm 3.2\%$ |
| | 150 | -2.0% | +0.5% -0.5% | $\pm 2.1\%$ |
| CCDY(W^-) | 30 | -5.0% | +2.6% -1.6% | $\pm 3.2\%$ |
| | 150 | -2.1% | +0.6% -0.5% | $\pm 2.13\%$ |

J. Baglio, C. Duhr, B. Mistlberger, R. Szafron 2209.06138

$$\delta(\text{PDF-TH}) = \frac{1}{2} \frac{|\sigma^{\text{NNLO}}(\text{NNLO PDF}) - \sigma^{\text{NNLO}}(\text{NLO PDF})|}{\sigma^{\text{NNLO}}(\text{NNLO PDF})}$$

Towards N³LO PDFs

Scale evolution of the PDFs

(Gribov,Lipatov 1972; Lipatov 1975; Altarelli,Parisi 1977; Dokshitzer 1977)

$$\mu^2 \frac{d}{d\mu^2} f_i(x, \mu^2) = \int_x^1 \frac{dy}{y} P_{ij}(\alpha_s, y) f_j\left(\frac{x}{y}, \mu^2\right), \quad i = g, u, d, s, \dots$$

Flavour decomposition of the quark contribution

$$f_{NS,ik}^\pm = (f_i \pm f_{\bar{i}}) - (f_k \pm f_{\bar{k}}), \quad i, k = u, d, s, \dots$$

$$f_S = \sum_i (f_i + f_{\bar{i}}), \quad i = u, d, s, \dots,$$

Perturbative expansion

$$P_{ij}(\alpha_s, x) = \underbrace{a P_{ij}^{(0)}}_{\text{LO}} + \underbrace{a^2 P_{ij}^{(1)}}_{\text{NLO}} + \underbrace{a^3 P_{ij}^{(2)}}_{\text{NNLO}} + \underbrace{a^4 P_{ij}^{(3)}}_{\text{N}^3\text{LO}}, \quad a = \frac{\alpha_s}{4\pi}$$

Recent progress at N³LO

- Large- n_f limit (Gracey 1994, 1996; Davies,Vogt,Ruijl,Ueda,Vermaseren 2016)
- Flavour non-singlet: complete planar limit and approximate full QCD (Ruijl,Ueda,Vermaseren,Vogt 2017)
- Four Mellin moments of the splitting kernels (Moch,Ruijl,Ueda,Vermaseren,Vogt 2021)
- Approximate N³LO PDF fits (McGowan,Cridge,Harland-Lang,Thorne 2022; Hekhorn, Magni 2023)
- Complete n_f^2 term in $P_{qq}^{(3)}$ (Gehrmann,von Manteuffel,Sotnikov,Yang 2023)

In this talk

- Theory framework to compute the Mellin moments of P_{ij}
- Results for $P_{qq}^{(3)}$ and $P_{qg}^{(3)}$ up to 10 moments.

Operators of leading twist

The Mellin moments of the PDFs are operator matrix elements
 (Collins,Soper 1981)

$$\langle H(P) | \mathcal{O}_{i;+,\dots,+}^{(N),\text{bare}} | H(P) \rangle = (P^+)^N \int_0^1 dx x^{N-1} f_i^{\text{bare}}(x)$$

$|H(P)\rangle$ proton state of momentum P , $\mathcal{O}_{i;\mu_1\dots\mu_N}^{(N)}$ operators of leading twist

$$\mathcal{O}_{g;\mu_1\dots\mu_N}^{(N)} = \frac{1}{2} \mathcal{S}_T \left\{ F_{\rho\mu_1}^{a_1} D_{\mu_2}^{a_1 a_2} \dots D_{\mu_{N-1}}^{a_{N-2} a_{N-1}} F^{a_{N-1}\rho}_{\mu_N} \right\},$$

$$\mathcal{O}_{q;\mu_1\dots\mu_N}^{(N)} = \frac{1}{2} \mathcal{S}_T \left\{ \bar{\psi}_{i_1} \gamma_{\mu_1} D_{\mu_2}^{i_1 i_2} \dots D_{\mu_N}^{i_{N-1} i_N} \psi_{i_N} \right\},$$

$$\mathcal{O}_{ns;\mu_1\dots\mu_N}^{(N),\rho} = \frac{1}{2} \mathcal{S}_T \left\{ \bar{\psi}_{i_1} (\lambda^\rho) \gamma_{\mu_1} D_{\mu_2}^{i_1 i_2} \dots D_{\mu_N}^{i_{N-1} i_N} \psi_{i_N} \right\},$$

$\lambda^\rho \rightarrow$ generator of $SU(n_f)$.

$\mathcal{S}_T \rightarrow$ symmetrise over $\mu_1 \dots \mu_N$ and remove trace terms.

Scale dependence upon renormalisation

Scale dependence given by a **matrix** of renormalisation constants

$$\mathcal{O}_i^{(N),\text{ren}}(\mu^2) = Z_{ij}^{(N)}(\alpha_s, \mu^2) \mathcal{O}_j^{(N),\text{bare}}$$

The anomalous dimensions of $\mathcal{O}_i^{(N),\text{ren}}$ control the evolution of the PDFs
 (Gross,Wilczek 1974; Politzer,Georgi 1974)

$$\gamma_{ij}^{(N)} \equiv - \left(\mu^2 \frac{d}{d\mu^2} Z_{ik}^{(N)} \right) Z_{kj}^{-1} = - \int_0^1 dx x^{N-1} P_{ij}^{(N)}(x, \alpha_s).$$

In minimal subtraction

$$\gamma_{ij}^{(N)} = a \frac{\partial}{\partial a} Z_{ij}^{(N)} \Big|_{\frac{1}{\epsilon}}$$

Computing the anomalous dimensions

Non-singlet case

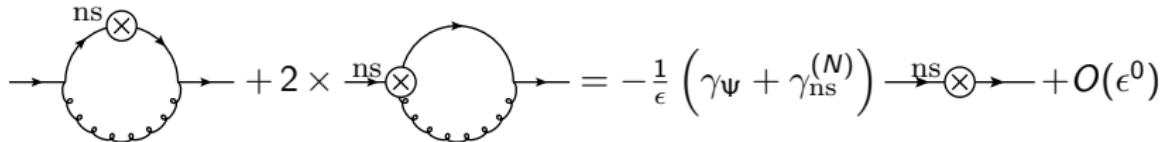
$$\mathcal{O}_{\text{ns}}^{(N),R}(\mu^2) = Z_{\text{ns}}^{(N)}(\mu^2) \mathcal{O}_{\text{ns}}^{(N),\text{bare}}$$

$$\rightarrow \text{loop with } \text{ns} \times \rightarrow + 2 \times \text{loop with } \text{ns} \times \rightarrow = -\frac{1}{\epsilon} \left(\gamma_\Psi + \gamma_{\text{ns}}^{(N)} \right) \text{loop with } \text{ns} \times \rightarrow + O(\epsilon^0)$$

Computing the anomalous dimensions

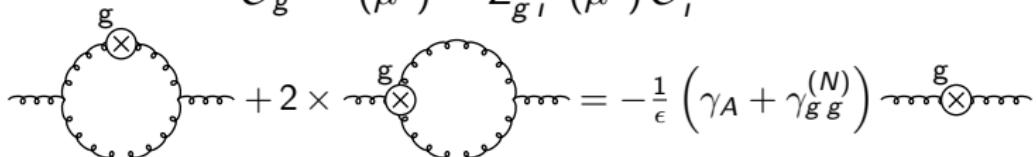
Non-singlet case

$$\mathcal{O}_{\text{ns}}^{(N),R}(\mu^2) = Z_{\text{ns}}^{(N)}(\mu^2) \mathcal{O}_{\text{ns}}^{(N),\text{bare}}$$



Singlet operators → the alien issue (Gross,Wilczek 1974)

$$\mathcal{O}_g^{(N),R}(\mu^2) = Z_{g i}^{(N)}(\mu^2) \mathcal{O}_i^{(N),\text{bare}}$$



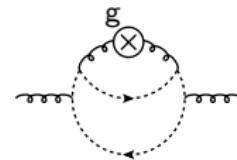
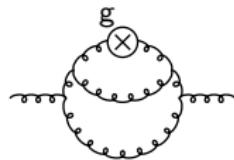
$$-\sum_{i \neq g} \frac{\gamma_{g i}^{(N)}}{\epsilon} \sim \text{Alien operators}$$

Multiloop renormalisation

Π_g^{ren} , i.e. renormalised 2-point functions with an insertion of $\mathcal{O}_g^{(N)}$, is finite

$$Z_3(Z_{g,i} \Pi_i(g_{\text{bare}}(g), \xi_{\text{bare}}(g, \xi))) = \text{finite}$$

Diagrammatically

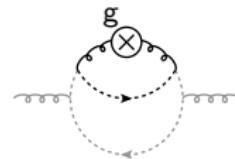
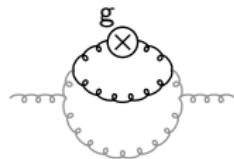


Multiloop renormalisation

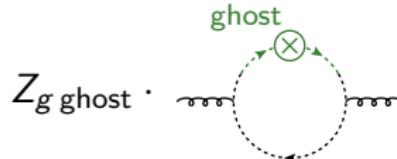
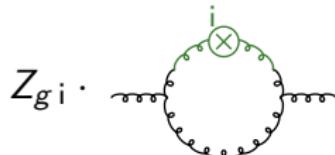
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Diagrammatically



Alien operators, including **ghost** operators enter in subdivergences



Facts about aliens

Required aliens at 2-loop level (Dixon and Taylor 1974). Defining

$$\partial \equiv \partial_+ = \partial_\mu \Delta^\mu, \quad D = D_\mu \Delta^\mu, \quad A^a = A_\mu^a \Delta^\mu, \quad F_\nu^a = F_{\nu\mu}^a \Delta^\mu,$$

$$\mathcal{O}_A^{(N)} = \eta F^{a;\alpha} D_\alpha^{ab} \partial^{N-2} A^b - g f^{abc} F_\alpha^a \sum_{i=1}^{N-2} \kappa_i \partial^\alpha \left[(\partial^{i-1} A^b) (\partial^{N-2-i} A^c) \right] + O(g^2),$$

$$\mathcal{O}_c^{(N)} = -\eta \bar{c}^a \partial^N c^a - g f^{abc} \sum_{i=1}^{N-2} \eta_i(\eta, \kappa_i) \bar{c}^a \partial \left[(\partial^{N-2-i} A^b) (\partial^i c^c) \right] + O(g^2),$$

- η, κ_i chosen to cancel the divergences (Hamberg, van Neerven 1993)
- Joglekar and Lee (1976): generalisation of BRST implies that aliens are
 - Operators proportional to the equation of motion
 - BRST-exact operators

Reminder: BRST invariance of the Yang-Mills lagrangian

$$\mathcal{L} = \underbrace{-\frac{1}{4} F^{a;\mu\nu} F_{\mu\nu}^a}_{\mathcal{L}_0} + \underbrace{s \left[\bar{c}^a \left(\partial^\mu A_\mu^a - \frac{\xi_L}{2} b^a \right) \right]}_{\text{Gauge fixing} + \text{ghost}}$$

- \mathcal{L}_0 invariant under

$$\delta A_\mu^a = (D_\mu \omega)^a,$$

with ω^a scalar function.

- s is the BRST transformation obtained by $\omega^a \rightarrow c^a$

$$s A_\mu^a = (D_\mu c)^a,$$

c^a , \bar{c}^a and b^a transform such that

$$s^2(\text{anything}) = 0.$$

Introducing leading twist operators

$$\widetilde{\mathcal{L}} = \underbrace{\mathcal{L}_0 + c_g \mathcal{O}_g^{(N)} + \mathcal{O}_{\text{EOM}}^{(N)}}_{\mathcal{L}_{\text{GGI}}} + \underbrace{\mathbf{s}' \left[\bar{c}^a \left(\partial^\mu A_\mu^a - \frac{\xi_L}{2} b^a \right) \right]}_{\text{Gauge fixing + ghost}}$$

$\mathcal{O}_{\text{EOM}}^{(N)}$ takes care of *gluonic* divergent (sub)diagrams

$$\begin{aligned} \mathcal{O}_{\text{EOM}}^{(N)} &= (D^\mu F_\mu)^a \left[\underbrace{\eta \partial^{N-2} A^a}_{\mathcal{O}_g^I} + g f^{aa_1 a_2} \sum_{i_1+i_2=N-3} \underbrace{\kappa_{i_1 i_2} (\partial^{i_1} A^{a_1})(\partial^{i_2} A^{a_2})}_{\mathcal{O}_g^{II}} \right. \\ &+ g^2 \sum_{\substack{i_1+i_2+i_3 \\ N-4}} \left(\underbrace{\kappa_{i_1 i_2 i_3}^{(1)} f^{aa_1 z} f^{a_2 a_3 z} + \kappa_{i_1 i_2 i_3}^{(2)} d_4^{aa_1 a_2 a_3} + \kappa_{i_1 i_2 i_3}^{(3)} d_{4\tilde{f}}^{aa_1 a_2 a_3}}_{\mathcal{O}_g^{III}} \right) (\partial^{i_1} A^{a_1})..(\partial^{i_3} A^{a_3}) \\ &\left. + g^3 \sum_{\substack{i_1+..+i_4 \\ N-5}} \left(\underbrace{\kappa_{i_1 \dots i_4}^{(1)} (f f f)^{aa_1 a_2 a_3 a_4} + \kappa_{i_1 \dots i_4}^{(2)} d_{4f}^{aa_1 a_2 a_3 a_4}}_{\mathcal{O}_g^{IV}} \right) (\partial^{i_1} A^{a_1})..(\partial^{i_4} A^{a_4}) + O(g^4) \right] \end{aligned}$$

Generalised Gauge and BRST transformations

\mathcal{L}_{GGI} is invariant under *generalised gauge transformations*. Given

$$\mathcal{O}_{\text{EOM}}^{(N)} = (D^\mu F_{\mu\nu})^a \mathcal{G}_\nu^a (A^b, \partial A^b, \partial^2 A^b, \dots),$$

the generalised transformation $A_\mu^a \rightarrow A_\mu^a + \delta' A_\mu^a$ is shown to be

$$\delta' A_\mu^a = \delta A_\mu^a - \delta \mathcal{G}_\mu^a + g f^{abc} \mathcal{G}_\mu^b \omega^c$$

This defines immediately the **generalised BRST transformations**

$$s'(A_\mu^a) = s(A_\mu^a) - s(\mathcal{G}_\mu^a) + g f^{abc} \mathcal{G}_\mu^b c^c \equiv s(A_\mu^a) + s_\Delta(A_\mu^a)$$

such that $s'^2(\text{anything}) = 0$.

Generalised BRST symmetry at work

The general ansatz of $\mathcal{O}_{\text{EOM}}^{(N)}$ fixes the structure of the aliens

- Example: first moment $N = 2$

$$\begin{aligned} \mathcal{G}_\mu^a &= \eta \Delta_\mu A^a, & s_\Delta [\bar{c}^a \partial^\rho A_\rho^a] &= -\bar{c}^a \partial \left[-\eta (Dc)^a + \eta g f^{abc} A^b c^c \right], \\ \mathcal{O}_g^I &= \eta (D^\nu F_\nu)^a A^a, & \mathcal{O}_c^I &= \eta \bar{c}^a \partial^2 c^a. \end{aligned}$$

There is a **single** alien operator

$$\mathcal{O}_A^I = \eta \left[(D^\nu F_\nu)^a A^a + \bar{c}^a \partial^2 c^a \right].$$

η mixes the physical operator into **gluon** and **ghost** aliens, Z_{galien} .



Quark operators

$\mathcal{O}_q^{(N)}$ has a much simpler mixing structure

$$2 \times \left(\text{---} \circlearrowleft \text{---} + \text{---} \circlearrowright \text{---} \right) = -\frac{1}{\epsilon} \left(\gamma_A + \gamma_{qg}^{(N)} \right) \text{---} \otimes \text{---} + \text{no alien}$$

The quark 2-point functions renormalise easily. Aliens occur only at 3 loops

$$\text{---} \circlearrowleft \text{---} + \dots \simeq Z_{qg} \cdot \text{---} \circlearrowright \text{---}$$

Note: the aliens must include now also a **quark** contribution in the EOM.

Pure singlet: aliens

$$\begin{aligned}
 O_g^I &= \eta (D^\nu F_\nu)^a \partial^{N-2} A^a, & O_g^{II} &= g f^{abc} (D^\nu F_\nu)^a \sum_{i_1+i_2=N-3} \kappa_{i_1 i_2} (\partial^{i_1} A^b) (\partial^{i_2} A^c), \\
 O_q^I &= \eta g (\bar{\psi} \not{A} t^a \psi) \partial^{N-2} A^a, & O_q^{II} &= g^2 (\bar{\psi} \not{A} t^a \psi) \sum_{i_1+i_2=N-3} \kappa_{i_1 i_2} (\partial^{i_1} A^b) (\partial^{i_2} A^c), \\
 O_c^I &= \eta \bar{c}^a \partial^N c^a, & O_c^{II} &= -(\partial \bar{c}^a) \sum_{i_1+i_2=N-3} \eta_{i_1 i_2} (\partial^{i_1} A^b) (\partial^{i_2+1} c^c),
 \end{aligned}$$

BRST and antiBRST symmetry impose relations

$$\eta_{ij} = 2\kappa_{ij} + \eta \binom{N-2}{i} = - \sum_{s=0}^i (-1)^{s+j} \binom{s+j}{s} \eta_{i-s,j+s}$$

The mixing constant are found to factorise up to 2 loops

$$\kappa_{ij} = \frac{\eta(N)}{8} \left[(-1)^i - 3 \binom{N-2}{i} + 3 \binom{N-2}{i+1} \right]$$

η renormalise ghost 2pt functions. Agreement with (Gehmann, von Manteuffel, Yang 2023).

Pure singlet anomalous dimensions

The required 2pt functions computed with FORCER ([Ruijl,Ueda,Vermaseren 2017](#)) for moments up to $N = 20$

$$\begin{aligned}\gamma_{\text{ps}}^{(3)}(N=2) &= -691.5937093 n_f + 84.77398149 n_f^2 + 4.466956849 n_f^3, \\ \gamma_{\text{ps}}^{(3)}(N=4) &= -109.3302335 n_f + 8.776885259 n_f^2 + 0.306077137 n_f^3, \\ \gamma_{\text{ps}}^{(3)}(N=6) &= -46.03061374 n_f + 4.744075766 n_f^2 + 0.042548957 n_f^3, \\ &\dots \\ \gamma_{\text{ps}}^{(3)}(N=20) &= -0.442681568 n_f + 0.805745333 n_f^2 - 0.020918264 n_f^3.\end{aligned}$$

- Agreement with results up to $N = 8$ ([Moch,Ruijl,Ueda,Vermasersen,Vogt 2021](#)), extended up to $N = 12$ ([Moch,Ruijl,Ueda,Vermaseren,Vogt to appear](#)).
- Leading terms in the large- n_f limit agree with ([Davies,Moch,Ruijl,Ueda,Vermaseren,Vogt 2016](#))
- Terms n_f^2 agree with ([Gehrmann,von Manteuffel,Sotnikov,Yang 2023](#))

Approximations of $P_{qq}^{(3)}(x)$ (I)

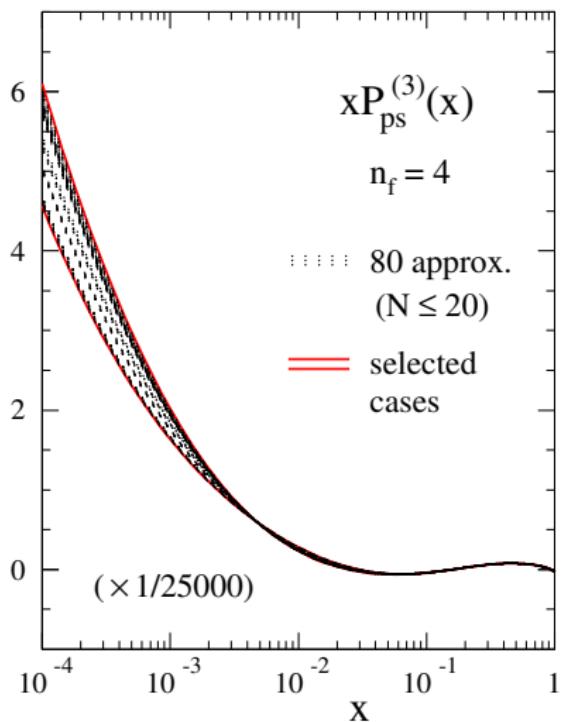
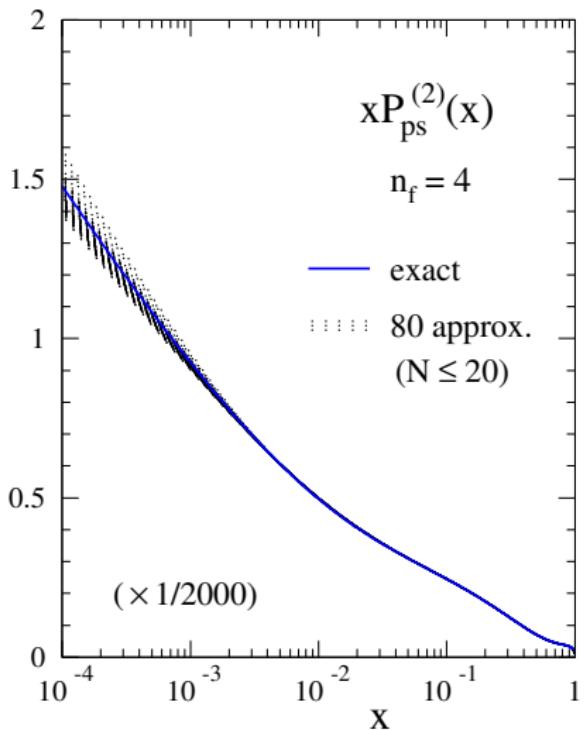
Following (Moch,Ruijl,Ueda,Vermaseren,Vogt 2017): approximations of the x-space results from 80 trial functions matching

- Moments up to $N = 20$
- Small-x limits
 - Coefficients of $\frac{\log^2 x}{x}$ (Catani,Hautmann 1994)
 - Coefficients of $\log^k x$ with $k = 6, 5, 4$ (Davies,Kom,Moch,Vogt 2022)
- Large-x limits
 - Coefficients of $(1 - x)^j \log^k(1 - x)$ with $k = 4, 3$ and $\forall j \geq 1$ (Soar,Moch,Vermaseren,Vogt 2010)

while unknown coefficients are fitted

- Small-x: $\frac{\log x}{x}, 1/x, \log^k x$ with $k = 3, 2, 1$
- Large-x: $(1 - x) \log^k x, k = 2, 1$

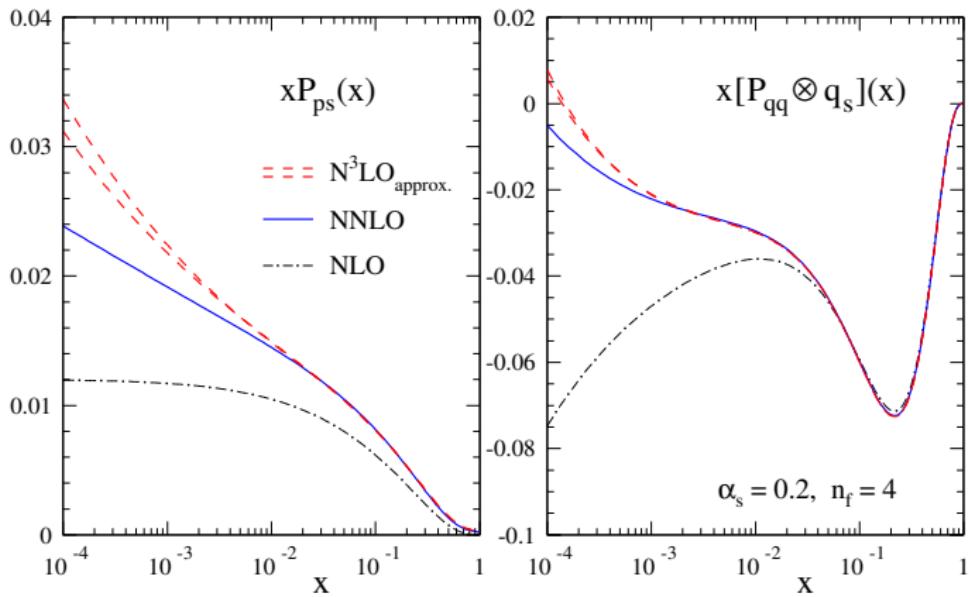
Approximations of $P_{qq}^{(3)}(x)$ (II)



Impact of the N³LO corrections

$P_{qq}(x)$ including approximate N³LO corrections for fixed $\alpha_s = 0.2$ (left).
 $P_{qq} \otimes f_S$ (right), where

$$x f_S(x) = 0.6 x^{-0.3} (1-x)^{3.5} (1 + 5.0 x^{0.8})$$



Results for $\gamma_{qg}^{(3)}$

Moments up to $N = 20$ of $\gamma_{qg}^{(3)}$ were computed in the same approach

$$\begin{aligned}\gamma_{qg}^{(3)}(N=2) &= -654.4627782 n_f + 245.6106197 n_f^2 - 0.924990969 n_f^3, \\ \gamma_{qg}^{(3)}(N=4) &= 290.3110686 n_f - 76.51672403 n_f^2 - 4.911625629 n_f^3, \\ \gamma_{qg}^{(3)}(N=6) &= 335.8008046 n_f - 124.5710225 n_f^2 - 4.193871425 n_f^3, \\ &\dots \\ \gamma_{qg}^{(3)}(N=20) &= 52.24329555 n_f - 109.3424891 n_f^2 - 2.153153725 n_f^3.\end{aligned}$$

- Agreement with moments up to $N = 8$ computed in
[\(Moch,Ruijl,Ueda,Vermaseren,Vogt 2021\)](#), extended to $N = 10$
[\(Moch,Ruijl,Ueda,Vermaseren,Vogt to appear\)](#)
- Agreement with the large- n_f limit [\(Davies,Vogt,Ruijl,Ueda,Vermaseren 2016\)](#)

Approximations of $P_{qg}^{(3)}$ (I)

The trial functions for $P_{qg}^{(3)}$ are constrained by the limits at

- Small- x :

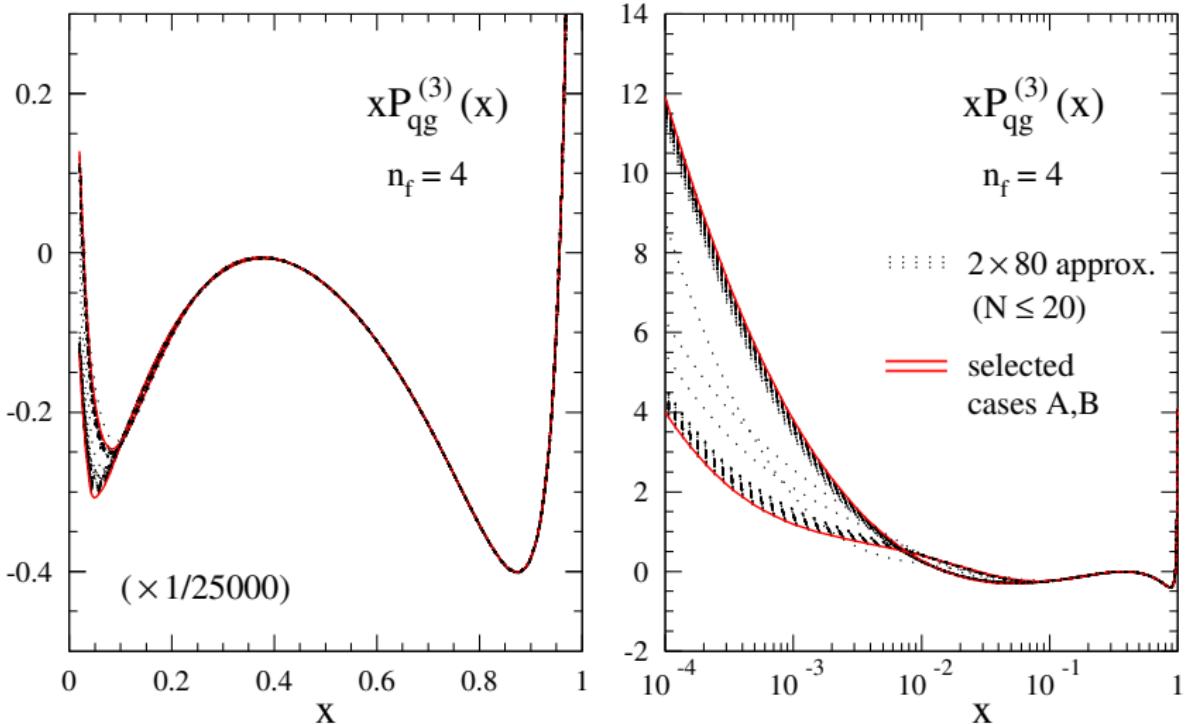
- Coefficients of $\frac{\log^2 x}{x}$ (Catani,Hautmann 1994)
- Coefficients of $\log^k x$ with $k = 6, 5, 4$ (Davies,Kom,Moch,Vogt 2022)

- Large- x :

- Coefficients of $\log^k(1 - x)$ with $k = 6, 5, 4$ (Soar,Moch,Vermaseren,Vogt 2010; Vogt 2010; Almasy,Soar,Vogt 2011)
- Coefficients of $(1 - x) \log^k(1 - x)$ with $k = 6, 5, 4$ (Soar,Moch,Vermaseren,Vogt 2010)

The coefficients of $\log^k(1 - x)$ with $k = 1, 2, 3$ are now **unknown**
 → uncertainties are larger compared to P_{qq} .

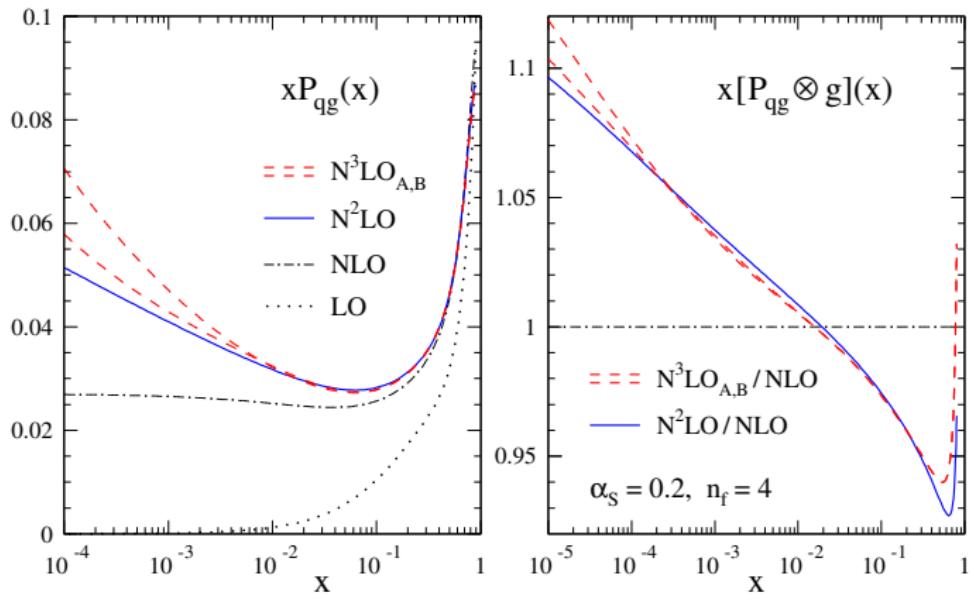
Approximations of $P_{qg}^{(3)}$ (II)



Impact of the N³LO corrections

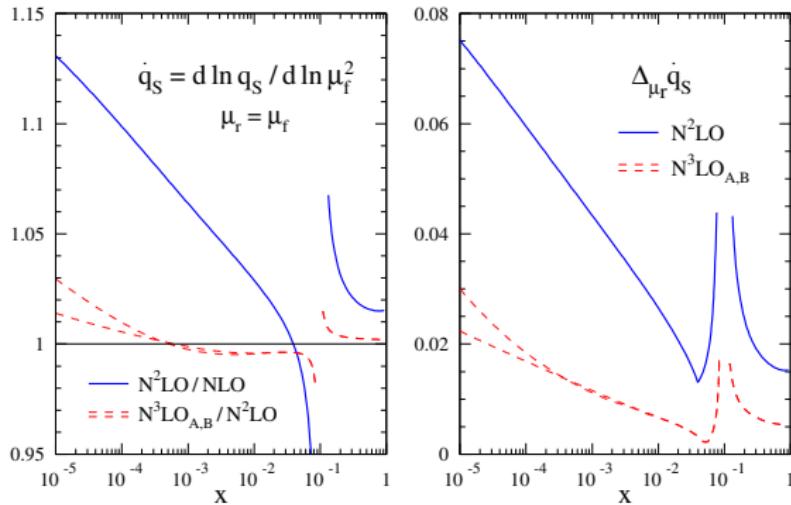
$P_{qg}(x)$ including approximate N³LO corrections fixing $\alpha_s = 0.2$ (left).
 $P_{qg} \otimes f_g$ (right), where

$$x f_g(x) = 1.6 x^{-0.3} (1-x)^{4.5} (1 - 0.6x^{0.3})$$



Scale evolution of the quark distribution

Using approximate P_{qq} and P_{qg} one derives $\mu_f^2 \frac{d}{d\mu_f^2} f_S \equiv \dot{q}_S = P_{qq} \otimes f_S + P_{qg} \otimes f_g$



The stability under variations of the renormalisation scale are quantified via

$$\Delta_{\mu_r} \dot{q}_S = \frac{1}{2} \frac{\max[\dot{q}_S(\mu_r^2 = \lambda \mu_f^2)] - \min[\dot{q}_S(\mu_r^2 = \lambda \mu_f^2)]}{\text{average}[\dot{q}_S(\mu_r^2 = \lambda \mu_f^2)]}, \quad \lambda = \frac{1}{4} \dots 4$$

Conclusions

Theory

- The moments of $P_{ij}(x)$ are computed *efficiently* from the renormalisation of 2-point correlators **provided** we take into account the mixing with **alien operators**.
- A generalised BRST symmetry fixes the structure of the aliens.
 - Classification in towers of contributions: O_k^I , O_k^{II} , O_k^{III} , ... with $k = g, c, q$
 O_k^I include 2-, 3- and 4-point vertices,
 O_k^{II} include 3-, 4- and 5-point vertices, ...
 - The renormalisation of quark operators requires few classes of terms.

Results

- The moments of $P_{qq}^{(3)}$ and $P_{qg}^{(3)}$ were computed up to $N = 20$.
- The approximate expressions of $P_{qq}(x)$ and $P_{qg}(x)$ at N³LO are characterised by
 - Small uncertainties at large-x, growing at small-x. E.g.
 $\delta P_{qg}(x = 10^{-4}) \sim \mathcal{O}(10\%)$
 - The convolution with PDFs dampens the uncertainty at small-x.
- Effect of the N³LO corrections to $\mu^2 \frac{d}{d\mu^2} f_S(x, \mu^2) = \dot{q}_S$
$$\delta_{\text{N}^3\text{LO}} \dot{q}_S(x = 10^{-4}) \lesssim 1\%.$$
- Renormalisation scale uncertainties are small, e.g. $x = 10^{-4}$ 2% vs compared to 5% at NNLO.

Outlook

- Ongoing work to compute the moments of $P_{gq}^{(3)}(x)$ and $P_{gg}^{(3)}(x)$.
- Can we obtain the exact expressions? This requires results for all N .
Only coefficient of Riemann- ζ numbers were reconstructed from the available moments
 - $\gamma_{qq}(N) \rightarrow$ Coefficients of ζ_5 , ζ_4 and of $\zeta_3 n_f \frac{d_{RR}}{n_c}$ and $\zeta_3 n_f^2 C_F^2$
 - $\gamma_{qg}(N) \rightarrow$ Coefficients of ζ_5 and $\zeta_3 n_f \frac{d_{RA}}{n_A}$ and $\zeta_3 n_f^2 \frac{d_{RR}}{n_A}$
 - These do not translate to the same coefficients of ζ in x -space.
- All fits would improve significantly with knowledge of the terms $\sim \frac{\log x}{x}$ in $P_{ij}^{(3)}$.
- Different methods to attack the all- N problem.

Thank you!