Collinear Contributions to QCD Resummations for Hard-Scattering Observables

Prasanna Kumar Dhani

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Introduction & Motivation Fixed Order Contribution Vs Resummed Contribution Transverse Momentum Resummation

Computation of Collinear Contributions

* Perturbative Results up to NNLO

***** Summary

Outline

Parton Model: Short Range Factorizes from the Long Range

For a generic hadro production process *

Additional radiation

$$h_1(p_1) + h_2(p_2) \to F(\{q_i\}) + X$$

$$c_i(x_1p_1) + c_j(x_2p_2) \to F(q \equiv \sum_i q_i) \left| q_\mu q^\mu = Q^2 = M^2 \right|$$

 \mathcal{X}_{i} : momentum fraction

$$\sigma_{h_1,h_2} = \sum_{i,j} \int dx_1 dx_2$$





Differential Cross Section: Analytic Structure

Theoretical separation

$$\frac{d\sigma_F}{dq_T} = \frac{d\sigma_F^{\text{sing.}}}{dq_T}$$

$$1 \ln^n (M^2) = \delta(q^2)$$

Spoils the reliability of FO predictions $\left\{ \left[\frac{1}{q_T^2} \ln^n \left(\frac{M^2}{q_T^2} \right) \right]_+, \, \delta(q_T^2) \right\}$

Process independent w.r.t. the system F

Good news: singular contributions are of infrared (soft and collinear) origin. Hence, they have some degree of universality * that leads to Resummation of these terms.

 $[g(\mathbf{q_T})]_+ =$



$$\lim_{D \to 0} \left[\theta(q_T - q_0)g(\mathbf{q_T}) - \delta^{(2)}(\mathbf{q_T}) \int d^2 \mathbf{k_T} g(\mathbf{k_T})\theta(\mu_0 - k_T)\theta(k_T - q_0) \right]$$





Why Resummation? Fixed Order Vs Resummed Predictions



Low transverse momentum region: while the FO results are not reliable, the resummed results are smooth. * *

For a relative order 'n' compared to Born

$$\alpha_{s}^{n} \underbrace{\frac{1}{q_{T}^{2}} \ln^{2n-1}\left(\frac{M^{2}}{q_{T}^{2}}\right), \frac{1}{q_{T}^{2}} \ln^{2n-2}\left(\frac{M^{2}}{q_{T}^{2}}\right), \dots, \frac{1}{q_{T}^{2}} \ln\left(\frac{M^{2}}{q_{T}^{2}}\right), \frac{1}{q_{T}^{2}}}}{\ln^{2n}\left(\frac{M^{2}b^{2}}{b_{0}^{2}}\right)}$$
pact parameter space
injugate to q_{T}

$$\ln^{2n}\left(\frac{M^{2}b^{2}}{b_{0}^{2}}\right)$$

$$\text{NLL:} \quad \alpha_{s}^{n} \ln^{n+1}\left(\frac{M^{2}b^{2}}{b_{0}^{2}}\right)$$

$$b_{0} = 2e^{-\gamma_{E}}$$

$$\text{NNLL:} \quad \alpha_{s}^{n} \ln^{n-1}\left(\frac{M^{2}b^{2}}{b_{0}^{2}}\right)$$
Next-to-next-to-leading logarithmic accuracy

High transverse momentum region: significant contributions are from the regular parts, making resummation ineffective.



Resummed Prediction Vs Data



Chen, Gehrmann, Glover, Huss, Monni, Rottoli, Re, Torrielli (2203.01565); Camarda, Cieri, Ferrera (2103.04974)

Bulk of the events are produced in the low transverse momentum region.

Theory is consistent with the data within the uncertainties.

From NNLO+NNLL to N3LO+N3LL: reduction in uncertainties and improvement in data agreement.

Improved theory predictions helps in precise determination of the mass of the W-boson and other crucial SM parameters.



How are these large logarithms resummed?









Transverse Momentum Resummation

Using the formalism of qT-resummation [1], singular part of the differential cross section in b-space has the * following structure Sudakov soft contribution Born contribution

$$[d\sigma_{F}] = \frac{M^{2}}{s} \sum_{c=q,\bar{q},g} \left[d\sigma_{c\bar{c},F}^{(0)} \right] \int \frac{d^{2}\mathbf{b}}{(2\pi)^{2}} e^{i\mathbf{b}\cdot\mathbf{q_{T}}} S_{c}(M,b)$$

$$\times \sum_{a_{1},a_{2}} \int_{x_{1}}^{1} \frac{dz_{1}}{z_{1}} \int_{x_{2}}^{1} \frac{dz_{2}}{z_{2}} \left[H^{F}C_{1}C_{2} \right]_{c\bar{c};a_{1}a_{2}} f_{a_{1}/h_{1}} \left(x_{1}/z_{1}, b_{0}^{2}/b^{2} \right) f_{a_{2}/h_{2}} \left(x_{2}/z_{2}, b_{0}^{2}/b^{2} \right)$$

$$\operatorname{Hard} \operatorname{collinear factor} \rightarrow \operatorname{This Talk!} \xrightarrow{x_{2}p_{2} + f_{2}p_{2}} F_{a_{2}/h_{2}} \left(x_{2}/z_{2}, b_{0}^{2}/b^{2} \right)$$

[1] Collins, Soper, Sterman (1985); Catani, Cieri, de Florian, Ferrera, Grazzini (1311.1654) [2] J Collins, Foundations of perturbative QCD; Becher, Neubert (1007.4005); Echevarria, Idilbi, Scimemi (1111.4996); Chiu, Jain, Neill, Rothstein (1202.0814)



Other equivalent formulations do exist in the literature that are based either on TMD factorisation or on SCET methods [2]





Hard Collinear Factor: Structure



$$\begin{bmatrix} H^F C_1 C_2 \end{bmatrix}_{gg;a_1 a_2} = H^F_{\mu_1 \nu_1, \mu_2 \nu_2} \begin{pmatrix} z \\ x C^{\mu_1 \nu_1}_{ga_1} (z_1; p_1) \end{pmatrix}$$

where [1]

$$C_{ga}^{\mu
u}\left(z;p_{1},p_{2},\mathbf{b};\alpha_{s}
ight)=d^{\mu
u}\left(p_{1},d^{\mu
u}(p_{1},p_{2})=-g^{\mu
u}+rac{p_{1}^{\mu}p_{2}^{
u}+p_{2}^{\mu}p_{1}^{
u}}{p_{1}.p_{2}}
ight)$$
 in

* For the quark anti-quark annihilation channel, it has a relatively simple structure.

[1] Catani, Grazzini (1011.3918)





Collinear Functions: Review from Literature

Azimuthally independent collinear functions are recently known to N3LO in QCD coupling *

$$C_{ab}(z;\alpha_{\rm S}) = \delta_{ab}\delta(1-z) + \frac{\alpha_{\rm S}}{\pi} C_{ab}^{(1)}(z) + \left(\frac{\alpha_{\rm S}}{\pi}\right)^2 C_{ab}^{(2)}(z) + \left(\frac{\alpha_{\rm S}}{\pi}\right)^3 C_{ab}^{(3)}(z) + \sum_{n=4}^{+\infty} \left(\frac{\alpha_{\rm S}}{\pi}\right)^n C_{ab}^{(n)}(z)$$
[1]
[2]
[3]

Azimuthally dependent collinear functions are recently known to NNLO in perturbation series

$$G_{ga}(z;\alpha_s) = \frac{\alpha_s}{\pi} G_{ga}^{(1)}(z) + \left(\frac{\alpha_s}{\pi}\right)^2 G_{ga}^{(2)}(z) + \sum_{n=3}^{+\infty} \left(\frac{\alpha_s}{\pi}\right) G_{ga}^{(n)}(z)$$
[4]

Similar functions do exist for the processes related by crossing such as SIDIS and production of hadrons from * a pair of leptons and they are called Time-Like collinear functions or Fragmenting Jet functions.

[1] de Florian, Grazzini (0108273); [4] Catani, Grazzini (1011.3918) [2] Catani, Grazzini (1106.4652); Catani, Cieri, de Florian, Ferrera, Grazzini (1209.0158); Gehrmann, Lubbert, Yang (1209.0682, 1403.6451); Echevarria, Scimemi, Vladimirov (1604.07869); Luo, Wang, Xu, Yang, Yang, Zhu (1908.03831); Luo, Yang, Zhu, Zhu (1909.13820) [3] Luo, Yang, Zhu, Zhu (1912.05778); Ebert, Mistlberger, Vita (2006.05329); Luo, Yang, Zhu, Zhu (2012.03256) [5] Luo, Yang, Zhu, Zhu (1909.13820); Gutierrez-Reyes, Leal-Gomez, Scimemi, Vladimirov (1907.03780)





Our Novel Computational Method



Stefano Catani + PKD (2208.05840)



QCD Factorization in the Collinear Limit

*

$$|\mathcal{M}(\{q_i\}; k_1, \dots, k_N)|^2 = \langle \mathcal{M}(\{q_i\}; \tilde{k}_i) \rangle$$
Dependence on non-
collinear partons
Collinear limit of $\sum_{i=1}^N k_i$
The TL collinear region is defined by
$$\{k_i^0\} > 0$$

$$F \longrightarrow HS \longrightarrow Collinear partons$$

- The splitting kernel is process independent and this * property of factorisation is called strict collinear factorisation.
- [1] Catani, de Florian, Rodrigo (1112.4405)

*

The collinear factorisation of hard scattering matrix element having N collinear partons in its most general form is given by



Strict collinear factorisation is instead violated in SL * collinear region [1].





Comments on the Use of Auxiliary Vector

* momenta of the form [1]

$$s_{ij} = 2k_i k_j$$

* convenient for direct computations of the splitting kernels. However, we emphasise that one can also set

$$n^2 \neq$$

The change only affects the non-singular/power-suppressed terms in the collinear limit, and hence neglected. *

[1] Catani, Grazzini (9908523)

In addition to its dependence on spin and colour indices, the splitting kernel depends on scalar functions of collinear

$$\frac{x_i}{x_j} = \frac{nk_i}{nk_j}$$

In the literature, the splitting kernels are usually calculated using a light-like auxiliary vector. Indeed this choice is very

0 This work — Time-like
$$(n^2 > 0)$$







Differential Collinear Functions

We define the differential collinear functions for the gluon channel in the TL region as follows [1] *



- TL splitting kernels for various splitting processes are fully known to second [2] and third order [3] in QCD. *
- SL Splitting kernels for various splitting processes are fully known to second order and partially known to third order * [4] in the QCD strong coupling.

[1] Stefano Catani + PKD; [4] Catani, de Florian, Rodrigo, Forshaw, Seymour, Siodmok, Dixon, Herrmann, Yan, Zhu [2] Bern, Del Duca, Kilgore, Schmidt, Catani, Grazzini, Campbell, Glover, Kosower, Uwer, Sborlini, de Florian, Rodrigo Rodrigo, Badger, Buciuni, Peraro, Bern, Dixon, Kosower, Gehrmann, Jaquier, Czakon, Sapeta

[3] Catani, de Florian, Rodrigo, Del Duca, Frizzo, Maltoni, Birthwright, Glover, Khoze, Marquard, Duhr, Haindl, Lazopoulos, Michel, Sborlini,



Integrated Collinear Functions: Time-Like Region

We define the transverse momentum dependent collinear functions for the gluon case as follows *

For the quark channel, we have only the azimuthal-independent type contributions. *

Stefano Catani + PKD (2208.05840)







Integrated Collinear Functions: Space-Like Region

We define the transverse momentum dependent collinear functions for both quark and gluon splitting as follows *

$$\mathbf{F}_{ca}(\{q_i\}; z; zp, \mathbf{q_T}; n) = \mathbf{1} \ \delta(1-z) \ \delta^{(d-2)}(\mathbf{q_T}) \ \delta_{ca} \\ + z \int d^d k \ \delta^{(d-2)}(\mathbf{k_T} + \mathbf{q_T}) \ \delta\left(\frac{k^+}{p^+} - 1 + z\right) \ \mathcal{F}_{ca}(\{q_i\}; p, k; n) \\ \mathcal{F}_{ca}(\{q_i\};$$

- Note that until now our framework is quite general i.e. without restricting to any perturbative order. *
- * dealing with are as follows

$$F_{ga}^{\mu\nu}(z;zp,\mathbf{q_T};n) , \quad F_{ga,\,\mathrm{az.in.}}\left(z;\mathbf{q_T}^2, \frac{n^2 \mathbf{q_T}^2}{(2zpn)^2}\right) , \quad F_{ga,\,\mathrm{corr.}}\left(z;\mathbf{q_T}^2, \frac{n^2 \mathbf{q_T}^2}{(2zpn)^2}\right), \quad F_{ca}\left(z;\mathbf{q_T}^2, \frac{n^2 \mathbf{q_T}^2}{(2zpn)^2}\right)$$

Stefano Catani + PKD (2208.05840)



Up to NNLO, SL collinear functions are also process independent. Hence, like TL region the functions we will be



Other Applications: Zero Jettiness Beam Functions

- Our method of computing collinear functions can be extended and applied to other observables of interest.
- * as follows

$$\begin{split} \boldsymbol{\mathcal{B}}_{ca}(\{q_i\}; z; zp, t; n) &= \mathbf{1} \ \delta(1-z) \ \delta(t) \ \delta_{ca} \\ &+ z \int d^d k \ \delta(t-2zpk) \ \delta\left(\frac{k^+}{p^+} - 1 + z\right) \ \boldsymbol{\mathcal{F}}_{ca}(\{q_i\}; p, k; n) \end{split}$$

Stewart, Tackmann, Waalewijn (1002.2213); Berger, Marcantonini, Stewart, Tackmann, Waalewijn (1012.4480); Ritzmann, Waalewijn (1407.3272); Gaunt, Stahlhofen, Tackmann (1401.5478); Gaunt, Stahlhofen, Tackmann (1405.1044); Ebert, Mistlberger, Vita (2006.03056); Baranowski, Behring, Melnikov, Tancredi, Wever (2211.05722)

For example, we can define the zero-jettiness partonic beam functions for both quark and gluon splitting in the SL region

Transverse virtuality ~ Zero-jettiness







Comments on Collinear Functions in SCET

* way using auxiliary Wilson line operators along light-like directions.

- * vector [1].
- *

[1] Ritzmann, Waalewijn (1407.3272)

There are related definitions of TMD collinear functions from SCET. These functions are defined in a process independent

For TL case at the partonic level, SCET functions are equivalent to our collinear functions by using a light-like auxiliary

Same equivalence holds true for the SL region is only up to second order in strong coupling. This is due to the fact that ours results are in general process dependent and this dependency goes away if we are within NNLO in perturbation theory.









TMD Functions & Rapidity Divergences

- * terms.
- * space to evaluate only the dominant terms of the cross section in the small qT-region.



- regulators that exist in the literature and they are introduced at the integrated level.
- *

Order-by-order perturbative computations of qT differential distributions at any values of qT can be carried out in exact form without encountering rapidity divergences and small-qT behaviour can be obtained by simply neglecting sub-dominant

Rapidity divergences are artefact of a priori approximations that are introduced in matrix element and qT dependent phase

Only if the differential collinear function is well behaved in the limit $k^- \to \infty$

For the computation of individual components such as soft functions, collinear functions etc. there are many rapidity

We avoid rapidity divergences in our computation by introducing a time-like auxiliary vector at the matrix element level.











Consider the singular term *

$$\frac{1}{1-z_n} = \frac{np}{nk} = \frac{p^+}{k^+ + \frac{n^2}{2np}\frac{k^-}{n^-}p^+} =$$

Using a time-like auxiliary vector *

$$\begin{split} \left(\frac{1}{1-z+\frac{\lambda}{1-z}} &= \left(\frac{1-z}{(1-z)^2+\lambda}\right)_+ + \delta(1-z) \int_0^1 dz' \frac{1-z'}{(1-z')^2+\lambda} \\ &= \frac{1}{2} \ln\left(\frac{1}{\lambda}\right) \,\delta(1-z) + \left(\frac{1}{1-z}\right)_+ + \mathcal{O}(\sqrt{\lambda}) \\ &\lambda = \frac{n^2 \mathbf{q_T}^2}{(2np)^2} & \text{Sub-dominant in the limit } \mathbf{q_T} \to 0 \end{split}$$

Stefano Catani + PKD (2208.05840)

No Rapidity Divergent Terms!





Differential Collinear Functions

SL differential collinear functions up to NNLO have the following perturbative expansion *

$$\mathcal{F}(p,k;n) = \mathcal{F}^{(1R)}(p,k;n) + \left[\mathcal{F}^{(2R)}(p,k;n) + \mathcal{F}^{(1R1V)}(p,k;n) \right] + \mathcal{O}(\alpha_{\mathrm{S}}^{3})$$
Single real radiation
Double real radiation
$$\mathcal{F}^{(1R)}_{ca, \, \mathrm{az.in.}}(p,k;n) = \frac{\alpha_{\mathrm{S}}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon \gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2})}{pk} \frac{1}{z_{n}} \widehat{\mathcal{F}}^{(a}(z_{n};\epsilon)$$
Real contributions to LO
AP splitting functions
$$\mathcal{F}^{(1R)}_{gg, \, \mathrm{corr.}}(p,k;n) = -\frac{\alpha_{\mathrm{S}}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon \gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2}) C_{A}}{pk} \frac{1-z_{n}}{z_{n}^{2}},$$

$$\mathcal{F}^{(1R)}_{gg, \, \mathrm{corr.}}(p,k;n) = -\frac{\alpha_{\mathrm{S}}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon \gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2}) C_{A}}{pk} \frac{1-z_{n}}{z_{n}^{2}},$$

$$\mathcal{F}^{(1R)}_{gg, \, \mathrm{corr.}}(p,k;n) = -\frac{\alpha_{\mathrm{S}}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon \gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2}) C_{F}}{pk} \frac{1-z_{n}}{z_{n}^{2}},$$

$$\mathcal{F}^{(1R)}_{gg, \, \mathrm{corr.}}(p,k;n) = -\frac{\alpha_{\mathrm{S}}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon \gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2}) C_{F}}{pk} \frac{1-z_{n}}{z_{n}^{2}},$$

$$\mathcal{F}^{(1R)}_{gg, \, \mathrm{corr.}}(p,k;n) = -\frac{\alpha_{\mathrm{S}}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon \gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2}) C_{F}}{pk} \frac{1-z_{n}}{z_{n}^{2}},$$

$$\mathcal{F}^{(1R)}_{gg, \, \mathrm{corr.}}(p,k;n) = -\frac{\alpha_{\mathrm{S}}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon \gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2}) C_{F}}{pk} \frac{1-z_{n}}{z_{n}^{2}}},$$

* Azi

$$\mathcal{F}^{(1R)}(p,k;n) + \left[\begin{array}{c} \mathcal{F}^{(2R)}(p,k;n) + \mathcal{F}^{(1R1V)}(p,k;n) \end{array} \right] + \mathcal{O}(\alpha_{\rm S}^{3})$$
Double real radiation
Contributions @ $\mathcal{O}(\alpha_{\rm S})$

$$\mathcal{F}^{(1R)}_{ca,\,{\rm az.in.}}(p,k;n) = \frac{\alpha_{\rm S}^{u} \mu_{0}^{2\epsilon} S_{\epsilon}}{\pi} \frac{e^{\epsilon\gamma_{E}}}{\pi^{1-\epsilon}} \frac{\delta_{+}(k^{2})}{pk} \frac{1}{z_{n}} \widehat{\mathcal{F}}^{ca}(z_{n};\epsilon)$$
Real contributions to LO
AP splitting functions
$$\mathcal{O}(\alpha_{\rm S})$$

$$\mathcal{O}(\alpha_{\rm S}$$

* Azi





Collinear Functions in Momentum Space

At the bare level, azimuthally independent collinear functions at NLO are obtained as follows *

$$F_{ca, \text{az.in.}}^{(1R)}\left(z; \mathbf{q_T}^2, \frac{n^2 \mathbf{q_T}^2}{(2zpn)^2}\right) = \frac{\alpha_S^u \,\mu_0^{2\epsilon} \,S_\epsilon}{\pi} \, \frac{e^{\epsilon\gamma_E}}{\pi^{1-\epsilon} \,\mathbf{q_T}^2} \left\{ \widehat{P}_{ca}^{\text{reg}}(z; \epsilon) + \delta_{ca} \, A_c^{(1)} \left[\left(\frac{1}{1-z}\right)_+ -\frac{1}{2} \ln\left(\frac{n^2 \mathbf{q_T}^2}{(2zpn)^2}\right) \,\delta(1-z) \right] \right\}$$
$$\widehat{P}_{ca}(x; \epsilon) = \widehat{P}_{ca}^{\text{sing}}(x) + \widehat{P}_{ca}^{\text{reg}}(x; \epsilon) \quad | \quad \widehat{P}_{ca}^{\text{sing}}(x) = \delta_{ca} \, \frac{A_c^{(1)}}{1-x}$$

Azimuthally correlated collinear functions at $\mathcal{O}(\alpha_{\rm S})$ are obtained as follows *

$$F_{ga,\,\text{corr.}}^{(1R)}\left(z;\mathbf{q_T}^2,\frac{n^2\mathbf{q_T}^2}{(2zpn)^2}\right) = -\frac{\alpha_{\mathrm{S}}^u\,\mu_0^{2\epsilon}\,S_{\epsilon}}{\pi}\,\frac{e^{\epsilon\gamma_E}}{\pi^{1-\epsilon}\,\mathbf{q_T}^2}\,C_a\,\frac{1-z}{z}$$

Formulae for the Fourier transformation from momentum space to conjugate impact parameter space *

$$\begin{split} \int d^{d-2}\mathbf{q}_{\mathbf{T}} \, e^{-i\mathbf{b}\cdot\mathbf{q}_{\mathbf{T}}} \frac{\ln^{m}(\mathbf{q}_{\mathbf{T}}^{2})}{(\mathbf{q}_{\mathbf{T}}^{2})^{1+\delta}} & \stackrel{d=4-2\epsilon}{\longrightarrow} \frac{d^{m}}{d\rho^{m}} \Big|_{\rho=0} \pi^{1-\epsilon} \left(\frac{\mathbf{b}^{2}}{4}\right)^{\epsilon+\delta-\rho} \frac{\Gamma(\rho-\epsilon-\delta)}{\Gamma(1+\delta-\rho)}, \\ \int d^{d-2}\mathbf{q}_{\mathbf{T}} \, e^{-i\mathbf{b}\cdot\mathbf{q}_{\mathbf{T}}} \frac{\ln^{m}(\mathbf{q}_{\mathbf{T}}^{2})}{(\mathbf{q}_{\mathbf{T}}^{2})^{1+\delta}} D^{\mu\nu}(p,n;\mathbf{q}_{\mathbf{T}}) & \stackrel{d=4-2\epsilon}{\longrightarrow} -\frac{d^{m}}{d\rho^{m}} \Big|_{\rho=0} \pi^{1-\epsilon} \left(\frac{\mathbf{b}^{2}}{4}\right)^{\epsilon+\delta-\rho} \frac{\Gamma(1+\rho-\epsilon-\delta)}{\Gamma(2+\delta-\rho)} D^{\mu\nu}(p,n;\mathbf{b}). \end{split}$$





Factorisation of IR Divergences

Infrared factorisation in b-space

$$\widetilde{F}_{ca,\text{az.in.}}\left(z;\frac{\mathbf{b}^{2}}{b_{0}^{2}},\frac{n^{2}b_{0}^{2}}{(2zpn)^{2}\mathbf{b}^{2}}\right) = Z_{c}\left(\alpha_{\mathrm{S}}(b_{0}^{2}/\mathbf{b}^{2}),\frac{n^{2}b_{0}^{2}}{(2zpn)^{2}\mathbf{b}^{2}}\right)\sum_{b}\int_{z}^{1}\frac{dx}{x}\widetilde{C}_{cb}\left(x;\alpha_{\mathrm{S}}(b_{0}^{2}/\mathbf{b}^{2}),\epsilon,\frac{n^{2}b_{0}^{2}}{(2zpn)^{2}\mathbf{b}^{2}}\right)\widetilde{\Gamma}_{ba}(z/x;b_{0}^{2}/\mathbf{b}^{2})$$

$$\widetilde{F}_{ga,\text{corr.}}\left(z;\frac{\mathbf{b}^{2}}{b_{0}^{2}},\frac{n^{2}b_{0}^{2}}{(2zpn)^{2}\mathbf{b}^{2}}\right) = Z_{g}\left(\alpha_{\mathrm{S}}(b_{0}^{2}/\mathbf{b}^{2}),\frac{n^{2}b_{0}^{2}}{(2zpn)^{2}\mathbf{b}^{2}}\right)\sum_{b}\int_{z}^{1}\frac{dx}{x}\widetilde{G}_{gb}\left(x;\alpha_{\mathrm{S}}(b_{0}^{2}/\mathbf{b}^{2}),\epsilon,\frac{n^{2}b_{0}^{2}}{(2zpn)^{2}\mathbf{b}^{2}}\right)\widetilde{\Gamma}_{ba}(z/x;b_{0}^{2}/\mathbf{b}^{2})$$

$$\widetilde{C}_{cb}\left(z;\alpha_{\rm S},\epsilon=0,\frac{n^2b_0^2}{(2zpn)^2\mathbf{b}^2}\right) = C_{cb}(z;\alpha_{\rm S}), \widetilde{G}_{gb}\left(z;\alpha_{\rm S},\epsilon=0,\frac{n^2b_0^2}{(2zpn)^2\mathbf{b}^2}\right) = G_{gb}(z;\alpha_{\rm S})$$
$$\widetilde{\Gamma}_{ba}(z;\mu_F^2) = \delta_{ba}\,\delta(1-z) - \frac{\alpha_{\rm S}(\mu_F^2)}{\pi}\,\frac{P_{ba}^{(1)}(z)}{\epsilon} + \mathcal{O}(\alpha_{\rm S}^2)$$

Infrared factorisation factors for both gluon and quark are as follows

$$Z_{g} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{A}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right) + \frac{\beta_{0}}{\epsilon} - C_{A} \frac{\pi^{2}}{24} \right] + \mathcal{O}(\alpha_{\rm S}^{2}) \quad Z_{q} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{F}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right) + \frac{3}{4} \frac{C_{F}}{\epsilon} - C_{F} \frac{\pi^{2}}{24} \right] + \mathcal{O}(\alpha_{\rm S}^{2}) \quad Z_{q} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{F}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right) + \frac{3}{4} \frac{C_{F}}{\epsilon} - C_{F} \frac{\pi^{2}}{24} \right] + \mathcal{O}(\alpha_{\rm S}^{2}) \quad Z_{q} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{F}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right) + \frac{3}{4} \frac{C_{F}}{\epsilon} - C_{F} \frac{\pi^{2}}{24} \right] + \mathcal{O}(\alpha_{\rm S}^{2}) \quad Z_{q} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{F}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right) + \frac{3}{4} \frac{C_{F}}{\epsilon} - C_{F} \frac{\pi^{2}}{24} \right] + \mathcal{O}(\alpha_{\rm S}^{2}) \quad Z_{q} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{F}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right] + \frac{3}{4} \frac{C_{F}}{\epsilon} - C_{F} \frac{\pi^{2}}{24} \right] + \mathcal{O}(\alpha_{\rm S}^{2}) \quad Z_{q} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{F}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right] + \frac{3}{4} \frac{C_{F}}{\epsilon} - C_{F} \frac{\pi^{2}}{24} \right] + \mathcal{O}(\alpha_{\rm S}^{2}) \quad Z_{q} = 1 + \frac{\alpha_{\rm S}}{\pi} \left[\frac{C_{F}}{2} \left(\frac{1}{\epsilon^{2}} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right] + \frac{1}{\epsilon} \left[\frac{1}{\epsilon} + \frac{1}{\epsilon} \ln\left(\frac{n^{2}b_{0}^{2}}{(2znp)^{2}b^{2}}\right) \right] + \frac{1}{\epsilon} \left[\frac{1}{\epsilon} + \frac{1}{\epsilon} \ln\left(\frac{1}{\epsilon} + \frac{1}{\epsilon} \ln\left(\frac{1}{\epsilon} + \frac{1}{\epsilon} \ln\left(\frac{1}{\epsilon} + \frac{1}{\epsilon} + \frac{1}{\epsilon} + \frac{1}{\epsilon} \ln\left(\frac{1}{\epsilon} + \frac{1}{\epsilon} + \frac{1}{\epsilon} + \frac{1}{\epsilon} + \frac{1}{\epsilon} \ln\left(\frac{1}{\epsilon} + \frac{1}{\epsilon} + \frac{1}{\epsilon$$







Perturbative Results at NLO

Collinear functions in the Space-Like Region [1] *

$$egin{aligned} &C_{gq}^{(1)}(z) = &rac{1}{2}C_F z, \ &C_{qg}^{(1)}(z) = &T_R z (1-z)\,, \ &C_{qq}^{(1)}(z) = &rac{1}{2}C_F (1-z)\,. \end{aligned}$$

Collinear functions in the Time-Like Region [2] *

$$C_{ca}^{\mathrm{TL}\,(1)}(z) = -\widehat{P}_{ac}'(z;\epsilon=0) + 2P_{ac}^{(1)}(z)\,\ln z$$
$$G_{gg}^{\mathrm{TL}\,(1)}(z) = C_A \, z(1-z) \,, \ G_{gq}^{\mathrm{TL}\,(1)}(z) = -T_R \, z(1-z)$$

Our results are in full agreement with those in the literature. *

[1] de Florian, Grazzini (0108273); Catani, Grazzini (1106.4652) [2] Luo et al. (1908.03831,1909.13820); Echevarria et al. (1604. 07869); Nadolsky et al. (9906280)

$$egin{aligned} G_{gq}^{(1)}(z) = & C_F rac{1-z}{z} \ G_{gg}^{(1)}(z) = & C_A rac{1-z}{z} \ \end{aligned}$$



Perturbative Results at NNLO: Space-Like Region

Azimuthally correlated collinear functions are obtained as follows *

$$\begin{split} G_{gg}^{(2)}(z) = & C_A^2 \left\{ -\frac{37}{36z} + \frac{31}{18} - \frac{13z}{12} + \frac{11z^2}{36} - \ln(z) \left[\frac{1}{z} + \frac{19}{12} \right] + \frac{\ln^2(z)}{2} + \frac{1-z}{z} \left[\text{Li}_2(z) - \frac{\pi^2}{6} \right] \right\} \\ & + C_F N_f \left\{ \frac{(1-z)^3}{2z} - \frac{1}{4} \ln^2(z) \right\} + C_A N_f \left\{ -\frac{17}{36z} + \frac{4}{9} + \frac{z}{12} + \frac{z^2}{36} - \frac{1}{6} \ln(z) \right\} \\ G_{gq}^{(2)}(z) = & C_F^2 \left\{ -\frac{1-z}{2} + \frac{5}{4} \ln(z) - \frac{1}{4} \ln^2(z) - \frac{1-z}{2z} \left[\ln(1-z) + \ln^2(1-z) \right] \right\} \\ & + C_F N_f \left\{ -\frac{1-z}{3z} \left[\frac{2}{3} + \ln(1-z) \right] \right\} + C_A C_F \left\{ -\frac{11}{18z} + \frac{10}{9} - \frac{z}{2} - \ln(z) \left[\frac{1}{z} + \frac{5}{2} \right] \right. \\ & \left. + \frac{1}{2} \ln^2(z) + \frac{1-z}{z} \left[\frac{5}{6} \ln(1-z) + \frac{1}{2} \ln^2(1-z) + \text{Li}_2(z) - \frac{\pi^2}{6} \right] \right\} \end{split}$$

Our results are in full agreement with those in the literature [1]. *

[1] Luo, Yang, Zhu, Zhu (1909.13820); Gutierrez-Reyes, Leal-Gomez, Scimemi, Vladimirov (1907.03780)



Perturbative Results at NNLO: Time-Like Region

Azimuthally correlated collinear functions are obtained as follows *

$$\begin{split} G_{gg}^{\mathrm{TL}(2)}(z) = & C_F N_f \left\{ \frac{1}{18z} + \frac{1}{2} + z - \frac{14z^2}{9} + \ln(z) \left[-\frac{1}{3z} + 1 + \frac{3z}{2} \right] + \frac{3z}{4} \ln^2(z) \right\} \\ & + C_A N_f \left\{ -\frac{1}{36z} - \frac{1}{12} - \frac{4z}{9} + \frac{17z^2}{36} - \frac{z}{6} \ln(z) \right\} + C_A^2 \left\{ -\frac{1}{36z} - \frac{5}{12} - \frac{20z}{9} + \frac{11z^2}{4} \right. \\ & + \ln(z) \left[\frac{1}{3z} - 1 - \frac{67z}{12} + z^2 \right] + z(1-z) \left[\ln(z) \ln(1-z) - \mathrm{Li}_2(1-z) \right] - \ln^2(z) \left[3z - \frac{z^2}{2} \right] \right\} \\ G_{gq}^{\mathrm{TL}(2)}(z) = C_F \left\{ -\frac{1}{8} + \frac{3z}{4} - \frac{5z^2}{8} - \ln(z) \left[\frac{1}{4} + \frac{3z}{8} - \frac{z^2}{4} \right] + \ln^2(z) \left[\frac{3z}{8} - \frac{3z^2}{4} \right] \right. \\ & + z(1-z) \left[\frac{1}{4} \ln(1-z) - \frac{3}{2} \ln(z) \ln(1-z) + \frac{1}{4} \ln^2(1-z) - \mathrm{Li}_2(z) - \frac{\pi^2}{12} \right] \right\} \\ & + N_f \left\{ z(1-z) \left[\frac{1}{9} - \frac{1}{6} \ln(z) + \frac{1}{6} \ln(1-z) \right] \right\} + C_A \left\{ \ln(z) \left[\frac{1}{4} + \frac{13z}{6} - \frac{17z^2}{12} \right] \right. \\ & + \ln^2(z) \left[\frac{3z}{4} + \frac{z^2}{2} \right] + z(1-z) \left[-\frac{25}{36} - \frac{5}{12} \ln(1-z) - \frac{1}{2} \mathrm{Li}_2(1-z) - \frac{1}{4} \ln^2(1-z) + \frac{\pi^2}{4} \right] \right\} \end{split}$$

Our results are in full agreement with those in the literature [1]. * [1] Luo, Yang, Zhu, Zhu (1909.13820)



- * phase space.
- *

- * factorisation of scattering matrix elements in soft and collinear limits.
- * for various algorithms which aim to compute fully differential cross sections for various observables.



Fixed order predictions can be plagued with large logarithmic corrections stemming from soft and collinear regions of the

For better prediction, one needs to resum these large logarithms systematically to all orders in the perturbation theory.

Large logarithmic contributions to an observable can be obtained from the singular elements associated with the QCD

Singular terms associated with the QCD factorisation and the corresponding integrated functions are important elements



Summary

- * kernels for the scattering matrix element.
- * observable definition to obtain specific collinear functions for transverse momentum resummation.
- * correlated case, we have results up to NNLO in perturbation theory.
- * dependency comes from splitting kernels to collinear functions.
- * like auxiliary vector to avoid them at the matrix element level.

I presented our method to compute both SL and TL collinear functions for QCD resummations using respective splitting

To compute these functions, we defined a differential version at the intermediate level and integrated them using proper

For the azimuthally independent collinear functions, we have presented results up to NLO and for the azimuthally

In our computation, we have stressed on the point that SL collinear functions, in general, can be process dependent and this

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Thank You !









