

Neutrinos, where BSM physics begins (I)



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Neutrino Frontiers, GGI
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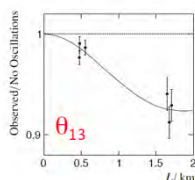
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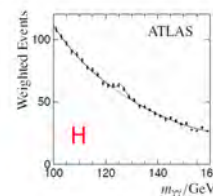
CSIC

~~two~~ 2012 One major discovery ~~ies~~ in particle physics

- A SM-like Higgs boson (ATLAS, CMS)
The key to EWSB and a possible window to



- $\theta_{13} \sim 10^\circ$ (T2K, MINOS, Daya Bay, RENO)
about as large as it could have been !
The door to CP Violation in the leptonic sector



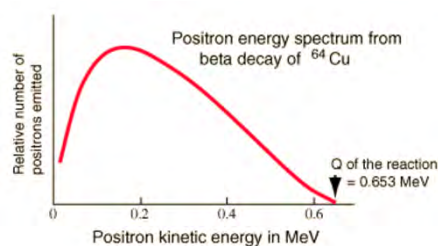
Summer Schools (if existed) were VERY short

β decay was supposed to be a two body decay



$$E_e = \frac{m_n^2 + m_e^2 - m_p^2}{2 m_n}$$

Studies of β decay revealed a continuous energy spectrum.



Another anomaly was the fact that the nuclear recoil was not in the direction opposite to the momentum of the electron.

The emission of another particle was a probable explanation of this behaviour, but searches found no evidence of either mass or charge.

...desperate remedy to save the law of conservation of energy...

Neutron Decay:

$$n \rightarrow p + e^- + \bar{\nu}_e$$



Fermi postulated a theory for β decay in terms of spinors

$$H_{ew} = \frac{G_F}{\sqrt{2}} \bar{\Psi}_p \gamma_\mu \Psi_n \bar{\Psi}_e \gamma^\mu \Psi_\nu$$

A Dirac field is described by a four component spinor

$$\begin{pmatrix} e_L \\ e_R \\ \hat{e}_L \\ \hat{e}_R \end{pmatrix}$$

Standard Model of Particle Physics

Gauge Theory based on the group:

$$SU(3) \times SU(2) \times U(1)$$

$SU(3) \Rightarrow$ Quantum Chromodynamics

Strong Force (Quarks and Gluons)

$SU_L(2) \times U(1) \Rightarrow$ ElectroWeak Interactions broken to $U_{EM}(1)$
by HIGGS

$$\underline{SU_L(2) \times U_Y(1) \Rightarrow U_{EM}(1)}$$

Force Carriers: W^\pm , Z^0 and γ masses: 80, 91 and 0 GeV

quark, $SU(2)$ doublets: $\begin{pmatrix} u \\ d \end{pmatrix}_L, \begin{pmatrix} c \\ s \end{pmatrix}_L, \begin{pmatrix} t \\ b \end{pmatrix}_L$

up-quark, $SU(2)$ singlets: u_R, c_R, t_R

down-quark, $SU(2)$ singlets: d_R, s_R, b_R

lepton, $SU(2)$ doublets: $\begin{pmatrix} \nu_e \\ e \end{pmatrix}_L, \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}_L, \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}_L$

neutrino, $SU(2)$ singlets: — — —

charge lepton, $SU(2)$ singlets: e_R, μ_R, τ_R

Electron mass

comes from a term of the form

$$\bar{L}\phi e_R$$

Absence of ν_R

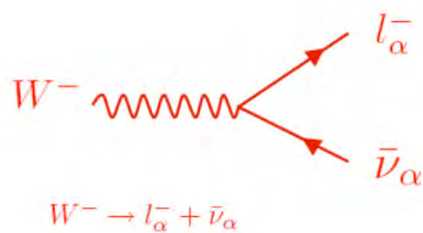
forbids such a mass term (dim 4)

for the Neutrino

Therefore in the SM neutrinos are massless
and hence travel at speed of light.

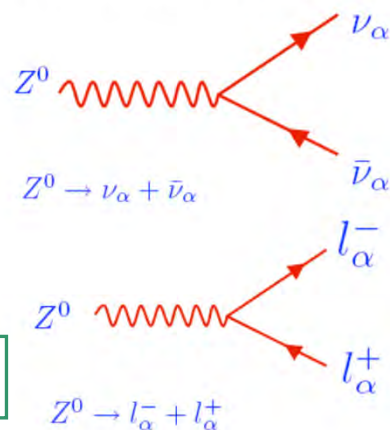
Interactions:

Charge Current (CC)



$$\Gamma(Z^0 \rightarrow f + \bar{f}) = K \frac{g_Z^2 M_Z}{48\pi} [|c_V^f|^2 + |c_A^f|^2]$$

Neutral Current (NC)

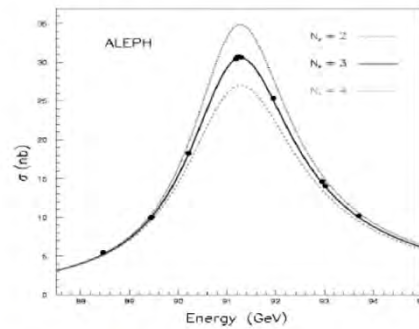


$\alpha = e, \mu, \text{ or } \tau$

Invisible width of Z plus other data from LEP:

$$Z^0 \rightarrow \nu \bar{\nu}$$

$$\text{Implies } N_\nu = 2.99 \pm 0.01$$



Three Active Neutrinos!!!

Sterile Neutrinos don't couple to Z^0

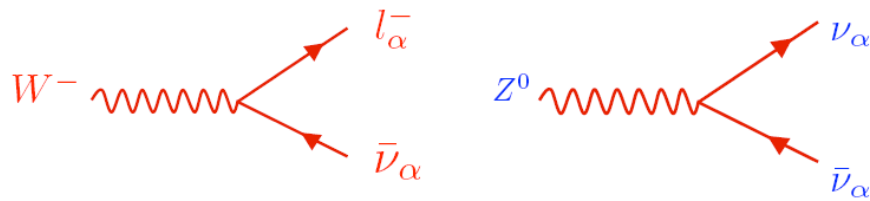
Note That

$$W^- \rightarrow l_\alpha^- + \bar{\nu}_\alpha$$

Implies

$$\begin{array}{ccc} \nu_e & , & \nu_\mu & , & \nu_\tau \\ \updownarrow & & \updownarrow & & \updownarrow \\ e & , & \mu & , & \tau \end{array}$$

Standard Model

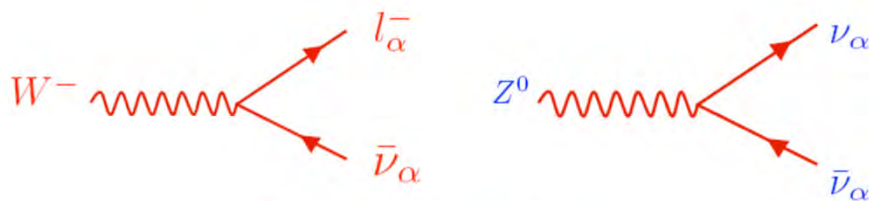


couplings conserve the **Lepton Number L**
defined by—

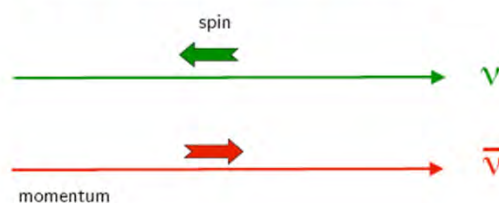
$$L(\nu) = L(\ell^-) = -L(\bar{\nu}) = -L(\ell^+) = 1.$$

Actually L_e , L_μ , and L_τ
separately

Left Handed Nature of The Neutrino



Produce Left-Handed Neutrinos
and Right-Handed Anti-Neutrinos



What about the RH neutrinos and LH anti-neutrino ????

There exist three fundamental and discrete transformations in nature:

- Parity \mathcal{P} $\vec{x} \rightarrow -\vec{x}$
- Time reversal \mathcal{T} $t \rightarrow -t$
- Charge conjugation \mathcal{C} $q \rightarrow -q$

\mathcal{P} , \mathcal{T} and \mathcal{C} are conserved in the classical theories of mechanics and electrodynamics!

$CPT \leftrightarrow$ Lorentz invariance \oplus unitarity: is an essential building block of field theory

CPT : L particle \leftrightarrow R antiparticle

Neutrinos in the MSM are massless and exist only in two states: particle with negative helicity and antiparticle with positive one: Weyl fermion

Summary of ν 's in SM:

Three flavors of massless neutrinos

$$W^- \rightarrow l_{\alpha}^- + \bar{\nu}_{\alpha}$$

$$W^+ \rightarrow l_{\alpha}^+ + \nu_{\alpha}$$

$$\alpha = e, \mu, \text{ or } \tau$$

Anti-neutrino, $\bar{\nu}_{\alpha}$, has +ve helicity, Right Handed

Neutrino, ν_{α} , has -ve helicity, Left Handed

ν_L and $\bar{\nu}_R$ are CPT conjugates

massless implies helicity = chirality

Beyond the SM

What if Neutrino have a MASS?

speed is less than c therefore time can pass

and

Neutrinos can change character!!!

What are the stationary states?

How are they related to the interaction states?

NEUTRINO OSCILLATIONS:

Two Flavors

flavor eigenstates \neq mass eigenstates

$$\begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

W's produce ν_μ and/or ν_τ 's

but ν_1 and ν_2 are the states

that change by a phase over time, mass eigenstates.

$$|\nu_j\rangle \rightarrow e^{-ip_j \cdot x} |\nu_j\rangle \quad p_j^2 = m_j^2$$

$\alpha, \beta \dots$ flavor index

$i, j \dots$ mass index

Production:

$$|\nu_\mu\rangle = \cos\theta|\nu_1\rangle + \sin\theta|\nu_2\rangle$$

Propagation:

$$\cos\theta e^{-ip_1 \cdot x}|\nu_1\rangle + \sin\theta e^{-ip_2 \cdot x}|\nu_2\rangle$$

Detection:

$$|\nu_1\rangle = \cos\theta|\nu_\mu\rangle - \sin\theta|\nu_\tau\rangle$$

$$|\nu_2\rangle = \sin\theta|\nu_\mu\rangle + \cos\theta|\nu_\tau\rangle$$

$$\begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = |\cos\theta(e^{-ip_1 \cdot x})(-\sin\theta) + \sin\theta(e^{-ip_2 \cdot x})\cos\theta|^2$$

$$P(\nu_\mu \rightarrow \nu_\tau) = |\cos\theta(e^{-ip_1 \cdot x})(-\sin\theta) + \sin\theta(e^{-ip_2 \cdot x})\cos\theta|^2$$

$$\text{Same } E, \text{ therefore } p_j = \sqrt{E^2 - m_j^2} \approx E - \frac{m_j^2}{2E}$$

$$e^{-ip_j \cdot x} = e^{-iEt} e^{-ip_j L} \approx e^{-i(Et - EL)} e^{-im_j^2 L/2E}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2\theta \cos^2\theta |e^{-im_2^2 L/2E} - e^{-im_1^2 L/2E}|^2$$

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

$$\delta m^2 = m_2^2 - m_1^2 \text{ and } \frac{\delta m^2 L}{4E} \equiv \Delta \text{ kinematic phase:}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = |\cos \theta (e^{-ip_1 \cdot x})(-\sin \theta) + \sin \theta (e^{-ip_2 \cdot x}) \cos \theta|^2$$

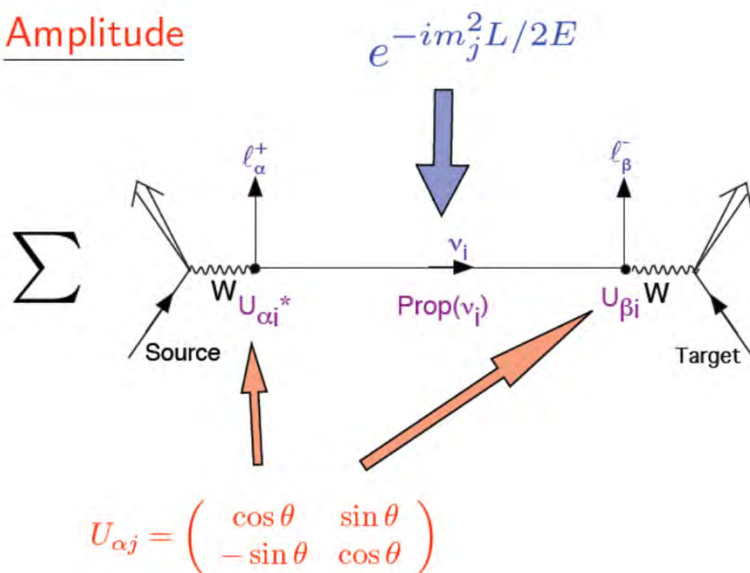
Same E, therefore $p_j = \sqrt{E^2 - m_j^2} \approx E - \frac{m_j^2}{2E}$

$$e^{-ip_j \cdot x} = e^{-iEt} e^{-ip_j L} \approx e^{-i(Et - EL)} e^{-im_j^2 L / 2E}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 \theta \cos^2 \theta |e^{-im_2^2 L / 2E} - e^{-im_1^2 L / 2E}|^2$$

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \sin^2 \left(\frac{\delta m^2 L}{4E} \frac{c^4}{\hbar c} \right)$$

Amplitude



Appearance:

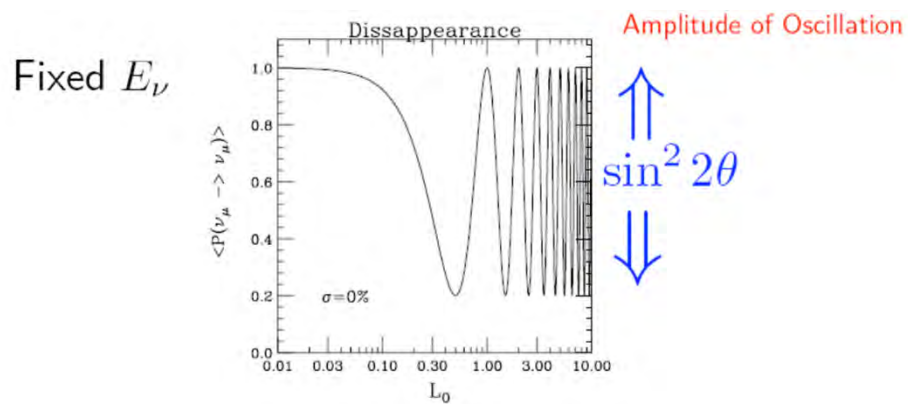
$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

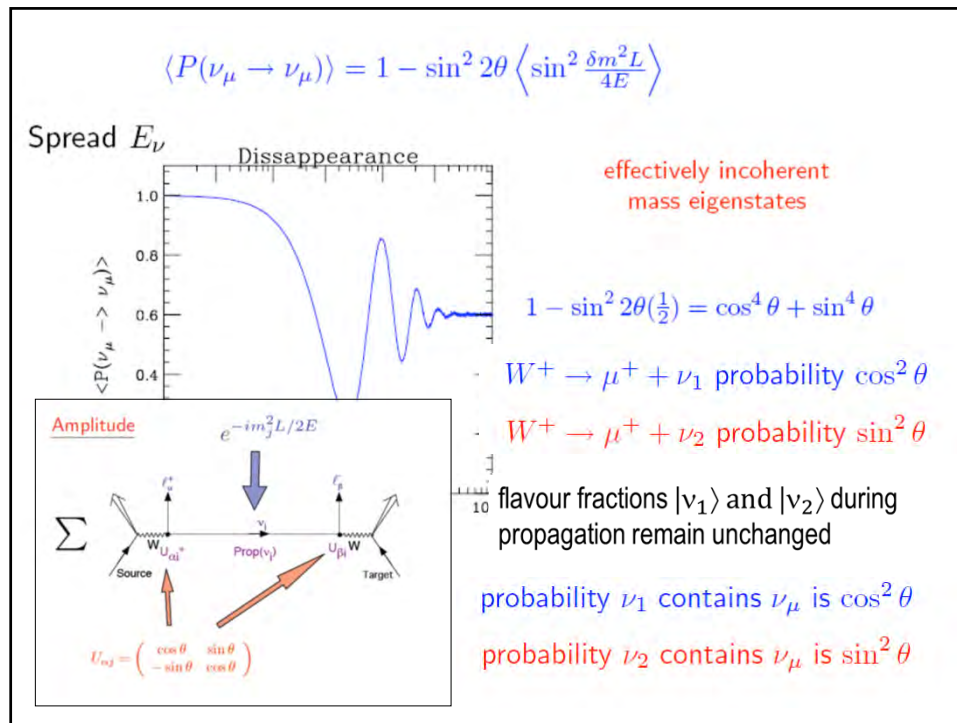
Disappearance:

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

Oscillation Length $L_0 = 4\pi E / \delta m^2$





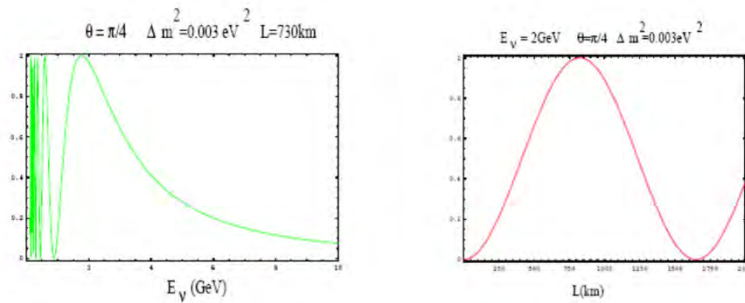
Using the unitarity of the mixing matrix: ($W_{\alpha\beta}^{jk} \equiv [V_{\alpha j} V_{\beta j}^* V_{\alpha k} V_{\beta k}^*]$)

$$P(\nu_\alpha \rightarrow \nu_\beta) = \delta_{\alpha\beta} - 4 \sum_{k>j} \text{Re}[W_{\alpha\beta}^{jk}] \sin^2 \left(\frac{\Delta m_{jk}^2 L}{4E_\nu} \right) \pm 2 \sum_{k>j} \text{Im}[W_{\alpha\beta}^{jk}] \sin \left(\frac{\Delta m_{jk}^2 L}{2E_\nu} \right)$$

For 2 families: $V_{MNS} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$

$$P_{\alpha\beta} = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E_\nu} \right) \rightarrow \text{appearance}$$

$$P_{\alpha\alpha} = 1 - P_{\alpha\beta} < 1 \rightarrow \text{disappearance}$$



Oscillation probabilities show the expected **GIM** suppression of any flavour changing process: they vanish if the neutrinos are degenerate

Probability for Neutrino Oscillation in Vacuum

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\text{Amp}(\nu_\alpha \rightarrow \nu_\beta)|^2 =$$

$$P_{\alpha\beta} = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4E_\nu} \right) \rightarrow \text{appearance}$$

$$P_{\alpha\alpha} = 1 - P_{\alpha\beta} < 1 \rightarrow \text{disappearance}$$

Probability for Neutrino Oscillation in Vacuum

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\text{Amp}(\nu_\alpha \rightarrow \nu_\beta)|^2 =$$

$$P_{\alpha\beta} = \sin^2 2\theta \sin^2 \left(\frac{\Delta m^2 L}{4 E} \right)$$

$$P_{\alpha\alpha} = 1 - P_{\alpha\beta}$$

$$\left(1.27 \frac{\Delta m^2 (eV^2) L(km)}{E(GeV)} \right)$$

L/E becomes crucial !!!

Evidence for Flavor Change:

*** Atmospheric and Accelerator Neutrinos with L/E = 500 km/GeV

*** Solar and Reactor Neutrinos with L/E = 15 km/MeV

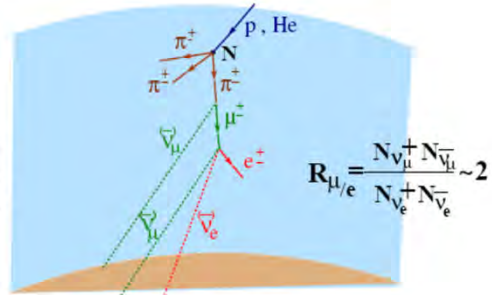
Neutrinos from Stopped muons L/E= 2m/MeV (Unconfirmed)

Atmospheric neutrinos

- Atmospheric neutrinos are produced by the interaction of *cosmic rays* (p, He, \dots) with the Earth's atmosphere:

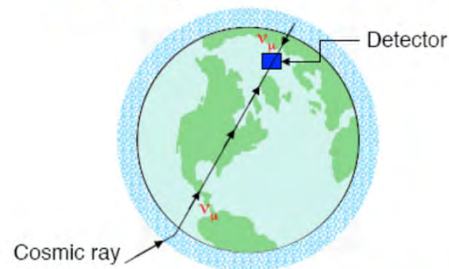
- 1 $A_{\text{cr}} + A_{\text{air}} \rightarrow \pi^{\pm}, K^{\pm}, K^0, \dots$
- 2 $\pi^{\pm} \rightarrow \mu^{\pm} + \nu_{\mu}$
- 3 $\mu^{\pm} \rightarrow e^{\pm} + \nu_e + \nu_{\mu}$

- at the detector, some ν interacts and produces a **charged lepton**, which is observed.



A deficit was observed in the ratio μ/e events: **Soudan2, IMB, Kamiokande**

Atmospheric Neutrinos



Isotropy of the $\geq 2 \text{ GeV}$ cosmic rays + Gauss' Law + No ν_{μ} disappearance

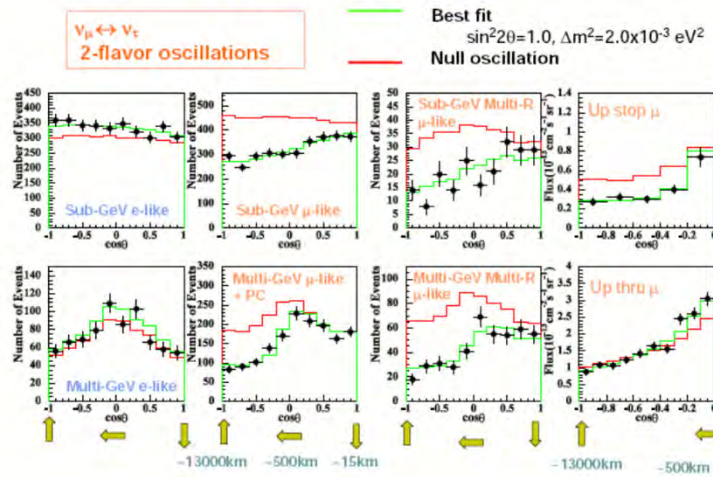
$$\Rightarrow \frac{\phi_{\nu_{\mu}}(\text{Up})}{\phi_{\nu_{\mu}}(\text{Down})} = 1.$$

But Super-Kamiokande finds for $E_{\nu} > 1.3 \text{ GeV}$

$$\frac{\phi_{\nu_{\mu}}(\text{Up})}{\phi_{\nu_{\mu}}(\text{Down})} = 0.54 \pm 0.04.$$



Zenith angle distributions



Half of the upward-going, long-distance-traveling ν_μ are disappearing.

Voluminous atmospheric neutrino data are well described by —

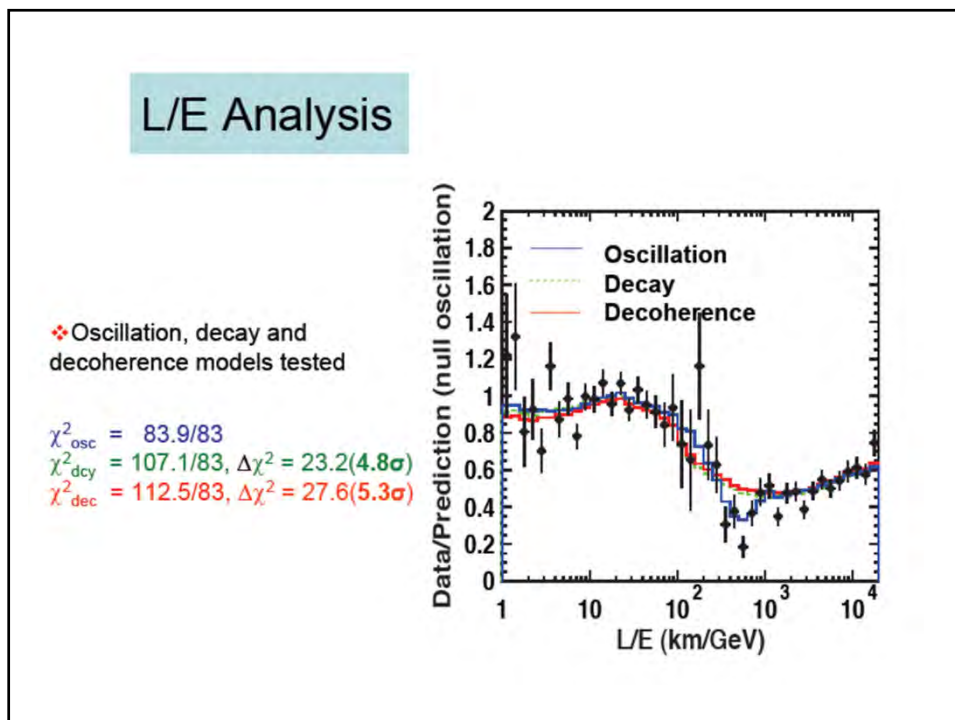
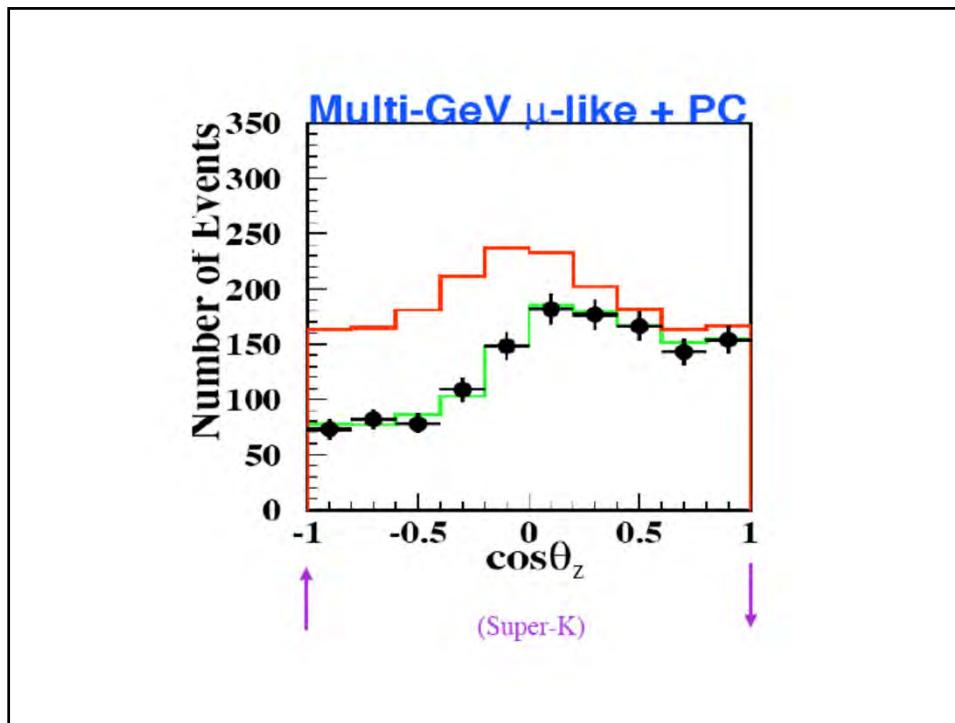
$$\nu_\mu \longrightarrow \nu_\tau$$

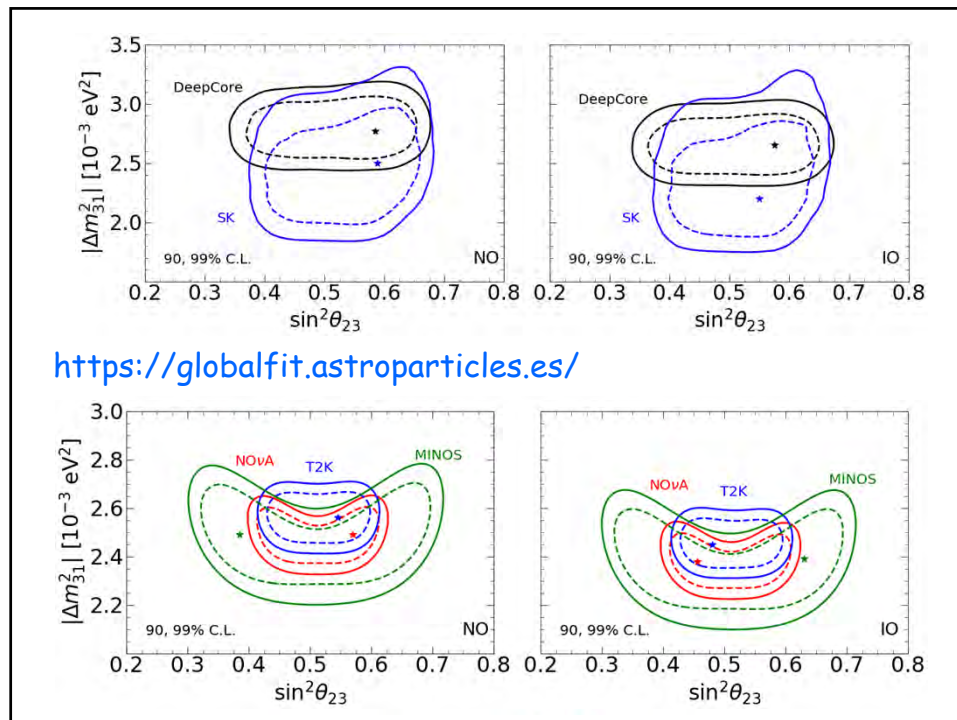
with —

$$\Delta m_{\text{atm}}^2 \cong 2.4 \cdot 10^{-3} \text{ eV}^2$$

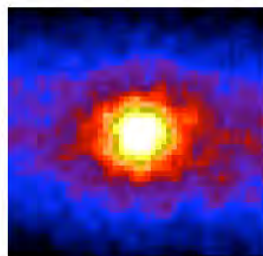
and —

$$\sin^2 2\theta_{\text{atm}} \cong 1$$

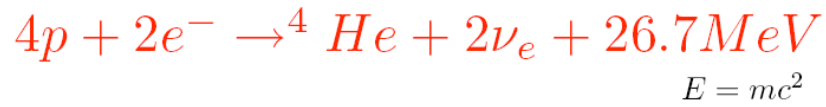




Solar δm^2



Solar Engine:

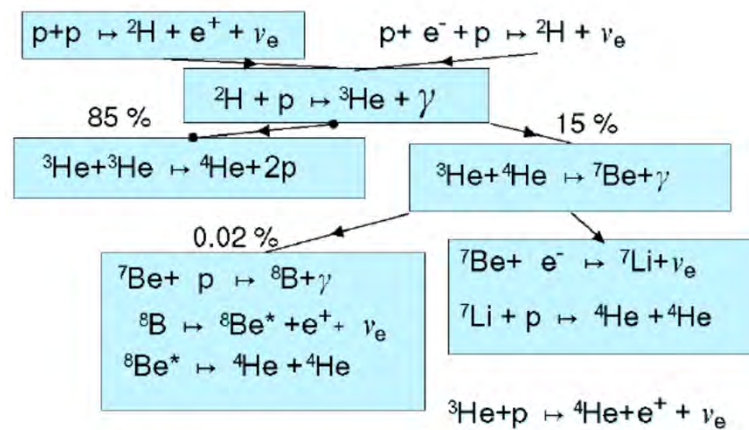


1 ν_e for every 13.4 MeV ($=2.1 \times 10^{-12}$ J)

\mathcal{L}_\odot at earth's surface 0.13 watts/cm²

$$\phi_\nu = \frac{0.13}{2.1 \times 10^{-12}} = 6 \times 10^{10} / \text{cm}^2 / \text{sec}$$

This corresponds to an average of 2 ν 's per cm³
since they are going at speed c.



Solar Spectrum:

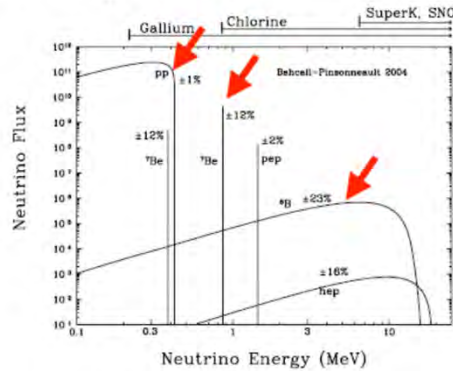
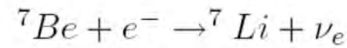


Figure 1. The predicted solar neutrino energy spectrum. The figure shows the energy spectrum of solar neutrinos predicted by the BP04 solar model [22]. For continuum sources, the neutrino fluxes are given in number of neutrinos $\text{cm}^{-2} \text{s}^{-1} \text{MeV}^{-1}$ at the Earth's surface. For line sources, the units are number of neutrinos $\text{cm}^{-2} \text{s}^{-1}$. Total theoretical uncertainties taken from column 2 of table 1 are shown for each source. To avoid complication in the figure, we have omitted the difficult-to-detect CNO neutrino fluxes (see table 1).



$$\phi_{pp} = 5.94(1 \pm 0.01) \times 10^{10} \text{cm}^{-2} \text{sec}^{-1}$$



$$\phi_{{}^7\text{Be}} = 4.86(1 \pm 0.12) \times 10^9 \text{cm}^{-2} \text{sec}^{-1}$$

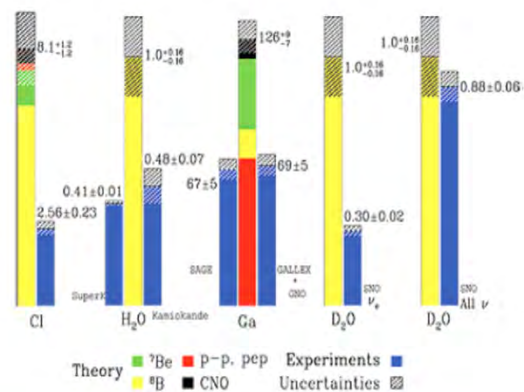


$$\phi_{{}^8\text{B}} = 5.82(1 \pm 0.23) \times 10^6 \text{cm}^{-2} \text{sec}^{-1}$$



Ray Davis & John Bahcall

Total Rates: Standard Model vs. Experiment
Bahcall-Serenelli 2005 [BS05(OP)]



Theory v Exp.

Neutrino Flavor Transitions!!!

Kinematical Phase:

$$\delta m_{\odot}^2 = 8.0 \times 10^{-5} eV^2$$

$$\sin^2 \theta_{\odot} = 0.31$$

$$\Delta_{\odot} = \frac{\delta m_{\odot}^2 L}{4E} = 1.27 \frac{8 \times 10^{-5} eV^2 \cdot 1.5 \times 10^{11} m}{0.1-10 MeV}$$

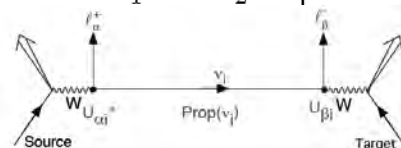
$$\Delta_{\odot} \approx 10^{7 \pm 1}$$

Effectively Incoherent !!!

Vacuum ν_e Survival Probability:

$$\langle P_{ee} \rangle = f_1 \cos^2 \theta_{\odot} + f_2 \sin^2 \theta_{\odot}$$

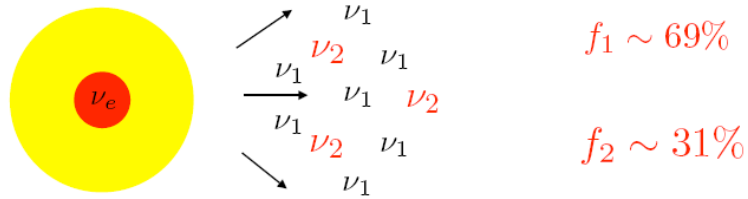
where f_1 and f_2 are the fraction of ν_1 and ν_2 at production.

In vacuum $f_1 = \cos^2 \theta_{\odot}$ \sum 

$$\langle P_{ee} \rangle = \cos^4 \theta_{\odot} + \sin^4 \theta_{\odot} = 1 - \frac{1}{2} \sin^2 2\theta_{\odot}$$

for pp and ${}^7\text{Be}$ this is approximately THE ANSWER.

$$f_1 \sim 69\% \text{ and } f_2 \sim 31\% \text{ and } \langle P_{ee} \rangle \approx 0.6$$

pp and ${}^7\text{Be}$ 

$$\langle P_{ee} \rangle \approx 0.6$$

$$f_3 = \sin^2 \theta_{13} < 4\%$$

What about ${}^8\text{B}$?

SNO's CC/NC

CC: $\nu_e + d \rightarrow e^- + p + p$

NC : $\nu_x + d \rightarrow \nu_x + p + n$

ES: $\nu_\alpha + e^- \rightarrow \nu_\alpha + e^-$

$$\frac{CC}{NC} = \langle P_{ee} \rangle = f_1 \cos^2 \theta_\odot + f_2 \sin^2 \theta_\odot$$

$$f_1 = (\frac{CC}{NC} - \sin^2 \theta_\odot) / \cos 2\theta_\odot$$

$$= (0.35 - 0.31) / 0.4 \approx 10\%$$

