Sep 29, 2025 - Oct 03, 2025

New Physics from Galaxy Clustering

GGI

Ivana Nikolac
Fabian Schmidt
Beatriz Tucci



Connecting Fieldlevel and Summary Statistics from Galaxy Clustering

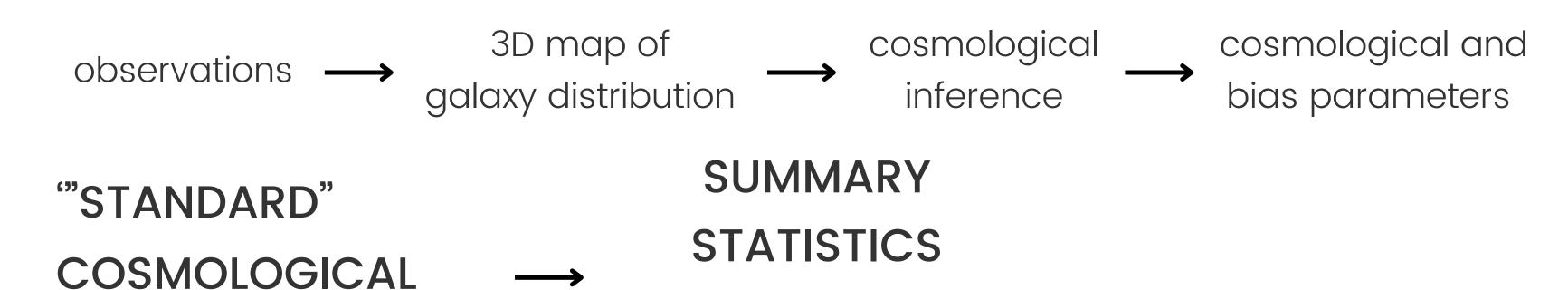
Seeing the Universe Voxel by Voxel: The Field-Level Approach

observations \longrightarrow 3D map of cosmological inference inference cosmological bias parameters

"STANDARD"
COSMOLOGICAL
INFERENCE

Seeing the Universe Voxel by Voxel: The Field-Level Approach

INFERENCE



2- and 3- point correlation functions

Seeing the Universe Voxel by Voxel: The Field-Level Approach

"STANDARD"
COSMOLOGICAL
INFERENCE

SUMMARY STATISTICS

2- and 3- point correlation functions

FIELD LEVEL

$$\delta_g(x) = rac{n_g(x)}{ar{n}_g} - 1$$

Seeing the Universe Voxel by Voxel: The Field-Level Approach

How much reliable cosmological information can be extracted?

"STANDARD"
COSMOLOGICAL
INFERENCE

SUMMARY STATISTICS

2- and 3- point correlation functions

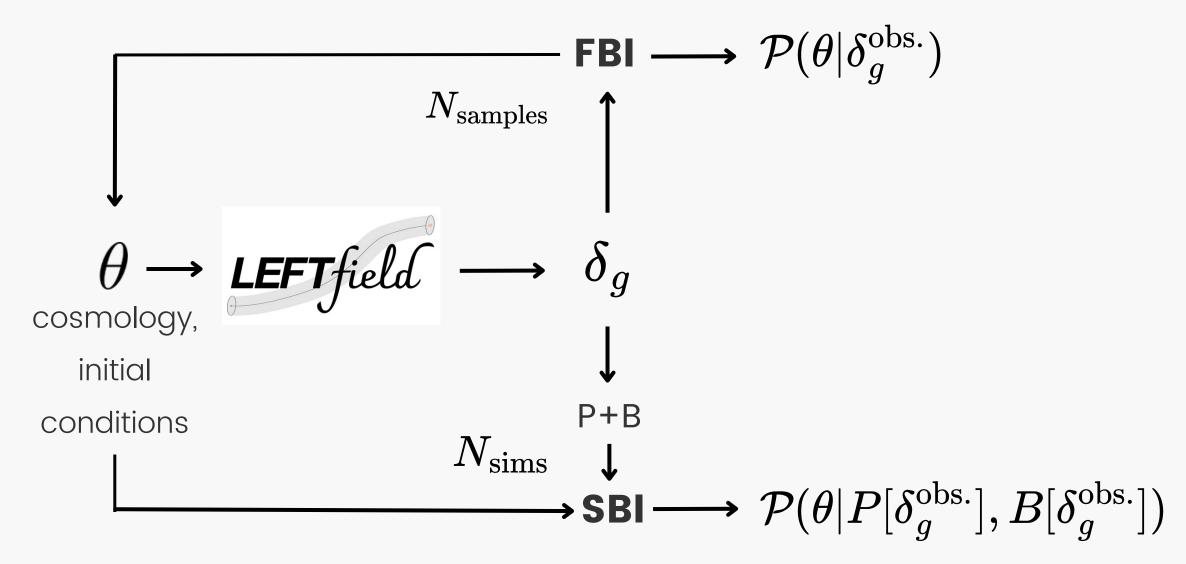
FIELD LEVEL

$$\delta_g(x) = rac{n_g(x)}{ar{n}_g} - 1$$

HOW MUCH INFORMATION CAN BE EXTRACTED FROM GALAXY CLUSTERING AT THE FIELD LEVEL?

Nguyen, Schmidt, Tucci et al. (2024)

field-level Bayesian inference



simulation-based inference



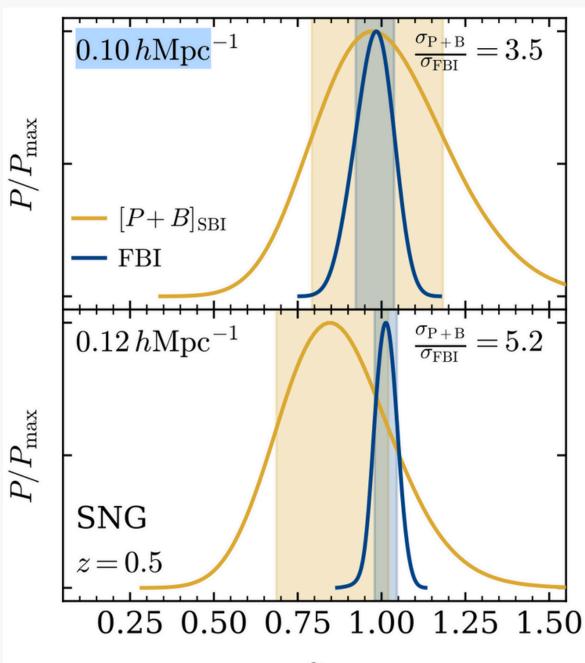
Nhat-Minh Nguyen IPMU, University of Tokyo



Fabian Schmidt MPA



Beatriz Tucci MPA-Stanford



$$\delta_{g, ext{det}}(oldsymbol{x}) = b_\delta \delta^{(1)}(oldsymbol{x}) + b_{\delta^2} [\delta^{(1)}(oldsymbol{x}) - \langle \delta^{(1)}
angle]^2 + arepsilon(oldsymbol{x})$$

as discussed in <u>Cabass et al.(2023)</u> and <u>Schmidt (2025)</u>

$$\delta_{g, ext{det}}(oldsymbol{x}) = oldsymbol{b}_{\delta} \delta^{(1)}(oldsymbol{x}) + oldsymbol{b}_{\delta^2} [\delta^{(1)}(oldsymbol{x}) - \langle \delta^{(1)}
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assumptions:
$$k_{
m max}=\Lambda$$

$$P_arepsilon o 0$$

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MAP expression for b_{δ^2} : $b_{\delta^2} = \frac{N[\delta_g]}{D[\delta_g]}$ with (at field level)

$$N[\delta_g] = rac{1}{b_\delta^2} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (\delta_g^2) (-m{k}) \delta_g(m{k}) - 2 \int_{m{x}} \delta_g(m{x}) dg$$

$$m{D}[m{\delta_g}] = rac{1}{b_\delta^4} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (\delta_g^2) (-m{k}) (\delta_g^2) (m{k}) - rac{8}{b_\delta^2} \int_{m{x}} \delta_g^2 (m{x}) ds$$

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m L}(k)} rac{(\delta_g^2)(-m{k})\delta_g(m{k})}{P_{
m L}(k)} - 2 \int_{m{x}} \delta_g(m{x})$$

$$egin{aligned} oldsymbol{D}[oldsymbol{\delta}_g] &= rac{1}{b_\delta^4} \int_{oldsymbol{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (oldsymbol{\delta}_g^2) (-oldsymbol{k}) (oldsymbol{\delta}_g^2) (oldsymbol{k}) - rac{8}{b_\delta^2} \int_{oldsymbol{x}} \delta_g^2(oldsymbol{x}) \end{aligned}$$

MAP expression for
$$b_{\delta^2}:\ b_{\delta^2}=rac{N[\delta_g]}{D[\delta_g]}$$
 $\ ext{with}$

MAP expression for
$$b_{\delta^2}$$
: $b_{\delta^2} = \frac{N[\delta_g]}{D[\delta_g]}$ with
$$D[\delta_g] = \frac{1}{b_{\delta}^2} \int_{\boldsymbol{k}}^{\Lambda} \frac{1}{P_{\mathrm{L}}(k)} (\delta_g^2) (-\boldsymbol{k}) \delta_g(\boldsymbol{k}) - 2 \int_{\boldsymbol{x}} \delta_g(\boldsymbol{x})$$
$$D[\delta_g] = \frac{1}{b_{\delta}^4} \int_{\boldsymbol{k}}^{\Lambda} \frac{1}{P_{\mathrm{L}}(k)} (\delta_g^2) (-\boldsymbol{k}) (\delta_g^2) (-\boldsymbol{k}) (\delta_g^2) (\boldsymbol{k}) - \frac{8}{b_{\delta}^2} \int_{\boldsymbol{x}} \delta_g^2(\boldsymbol{x})$$

$$D[\delta_g] = rac{1}{b_\delta^4} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (\!\delta_g^2) (\!-\!m{k}) (\!\delta_g^2) (m{k}) - rac{8}{b_\delta^2} \int_{m{x}} \delta_g^2 (m{x}) ds$$

new summaries:
$$\{\langle O^{(n)}[\delta_g]\,|\, O^{(m)}[\delta_g]
angle\}$$

MAP expression for
$$b_{\delta^2}:\ b_{\delta^2}=rac{N[\delta_g]}{D[\delta_q]}$$
 with

$$N[\delta_g] = rac{1}{b_\delta^2} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (m{\delta}_g^2) (-m{k}) m{\delta}_g(m{k}) - 2 \int_{m{x}} m{\delta}_g(m{x})$$
 $D[\delta_g] = rac{1}{b_\delta^4} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (m{\delta}_g^2) (-m{k}) (m{\delta}_g^2) (m{k}) - rac{8}{b_\delta^2} \int_{m{x}} m{\delta}_g^2(m{x})$

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new summaries:
$$\{\langle O^{(n)}[\delta_g]\,|\, O^{(m)}[\delta_g]
angle\}$$

 \longrightarrow correlators of powers of δ_q :

OC 2nd, LO:
$$\{\langle \delta_g \delta_g \rangle, \langle \delta_g \delta_g^2 \rangle\}$$

OC 2nd, full:
$$\{\langle \delta_g \delta_g \rangle, \langle \delta_g \delta_q^2 \rangle, \langle \delta_q^2 \delta_q^2 \rangle\}$$

MAP expression for
$$b_{\delta^2}:\ b_{\delta^2}=rac{N[\delta_g]}{D[\delta_q]}$$
 with

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m L}(k)} (\!\delta_g^2) (\!-\!m{k}) (\!\delta_g^2) (\!m{k}) - rac{8}{b_\delta^2} \int_{m{x}} \delta_g^2(m{x}) ds$$

new summaries: $\{\langle O^{(n)}[\delta_q] \,|\, O^{(m)}[\delta_q] \rangle\}$

 \longrightarrow correlators of powers of δ_q :

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OC 2nd, full: $\{\langle \delta_g \delta_g \rangle, \langle \delta_g \delta_q^2 \rangle, \langle \delta_g^2 \delta_q^2 \rangle\}$

P + tree-level bispectrum add part of LO trispectrum

MAP expression for
$$b_{\delta^2}:\ b_{\delta^2}=rac{N[\delta_g]}{D[\delta_g]}$$
 with

$$egin{aligned} N[\delta_g] &= rac{1}{b_\delta^2} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (m{\delta}_g^2) (-m{k}) m{\delta}_g(m{k}) - 2 \int_{m{x}} m{\delta}_g(m{x}) \ D[\delta_g] &= rac{1}{b_\delta^4} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (m{\delta}_g^2) (-m{k}) (m{\delta}_g^2) (m{k}) - rac{8}{b_\delta^2} \int_{m{x}} m{\delta}_g^2(m{x}) \ \end{array}$$

$$D[\delta_g] = rac{1}{b_\delta^4} \int_{m{k}}^{\Lambda} rac{1}{P_{
m L}(k)} (\!\delta_g^2) (\!-\!m{k}) (\!\delta_g^2) (\!m{k}) - rac{8}{b_\delta^2} \int_{m{x}} \delta_g^2(m{x}) ds$$

new summaries: $\{\langle O^{(n)}[\delta_q] \,|\, O^{(m)}[\delta_q] \rangle\}$

 \longrightarrow correlators of powers of δ_q :

OC 2nd, LO: $\{\langle \delta_g \delta_g \rangle, \langle \delta_g \delta_g^2 \rangle\}$

OC 2nd, full: $\{\langle \delta_g \delta_g \rangle, \langle \delta_g \delta_g^2 \rangle, \langle \delta_g^2 \delta_g^2 \rangle\}$

OC 3rd, LO: $\{\langle \delta_g \delta_g \rangle, \langle \delta_g \delta_q^2 \rangle, \langle \delta_g^2 \delta_q^2 \rangle, \langle \delta_g \delta_q^3 \rangle\}$

OC 3rd, full: $\{\langle \delta_g \delta_g \rangle, \langle \delta_g \delta_q^2 \rangle, \langle \delta_g^2 \delta_q^2 \rangle, \langle \delta_g \delta_q^3 \rangle, \langle \delta_g^2 \delta_q^3 \rangle, \langle \delta_g^3 \delta_q^3 \rangle\}$

P + tree-level bispectrum

add part of LO trispectrum

P + tree-level B + tree-level T

add part of 5- and 6-point functions

FIELD-LEVEL POSTERIOR

$$\mathcal{P}(\boldsymbol{\theta}, \delta_{\mathrm{in}}, \{b_O\}, \{\sigma_{\varepsilon}\} | \delta_g^{\mathrm{obs.}}) \longrightarrow \mathcal{P}(\{b_O\}, \sigma_{\varepsilon} | \delta_g^{\mathrm{obs.}})$$

FIELD-LEVEL POSTERIOR

$$\mathcal{P}(\boldsymbol{\theta}, \delta_{\text{in}}, \{b_O\}, \{\sigma_{\varepsilon}\} | \delta_g^{\text{obs.}}) \longrightarrow \mathcal{P}(\{b_O\}, \sigma_{\varepsilon} | \delta_g^{\text{obs.}})$$

EFT likelihood

$$\ln \mathcal{L}_{EFT}(\delta_g^{obs.}|\boldsymbol{\theta}, \delta_{in}, \{b_O\}, \{\sigma_{\varepsilon}\}) =$$

$$-\frac{1}{2} \sum_{k>0}^{|\boldsymbol{k}| < k_{\text{max}}} \left(\ln 2\pi \sigma_{\varepsilon}^{2}(k) + \frac{1}{\sigma_{\varepsilon}^{2}(k)} |\delta_{g}^{\text{obs.}}(\boldsymbol{k}) - \sum_{O} b_{O}O[\boldsymbol{\theta}, \delta_{\text{in}}](\boldsymbol{k})|^{2} \right)$$

FIELD-LEVEL POSTERIOR

$$\mathcal{P}(\boldsymbol{\theta}, \delta_{\text{in}}, \{b_O\}, \{\sigma_{\varepsilon}\} | \delta_g^{\text{obs.}}) \longrightarrow \mathcal{P}(\{b_O\}, \sigma_{\varepsilon} | \delta_g^{\text{obs.}})$$

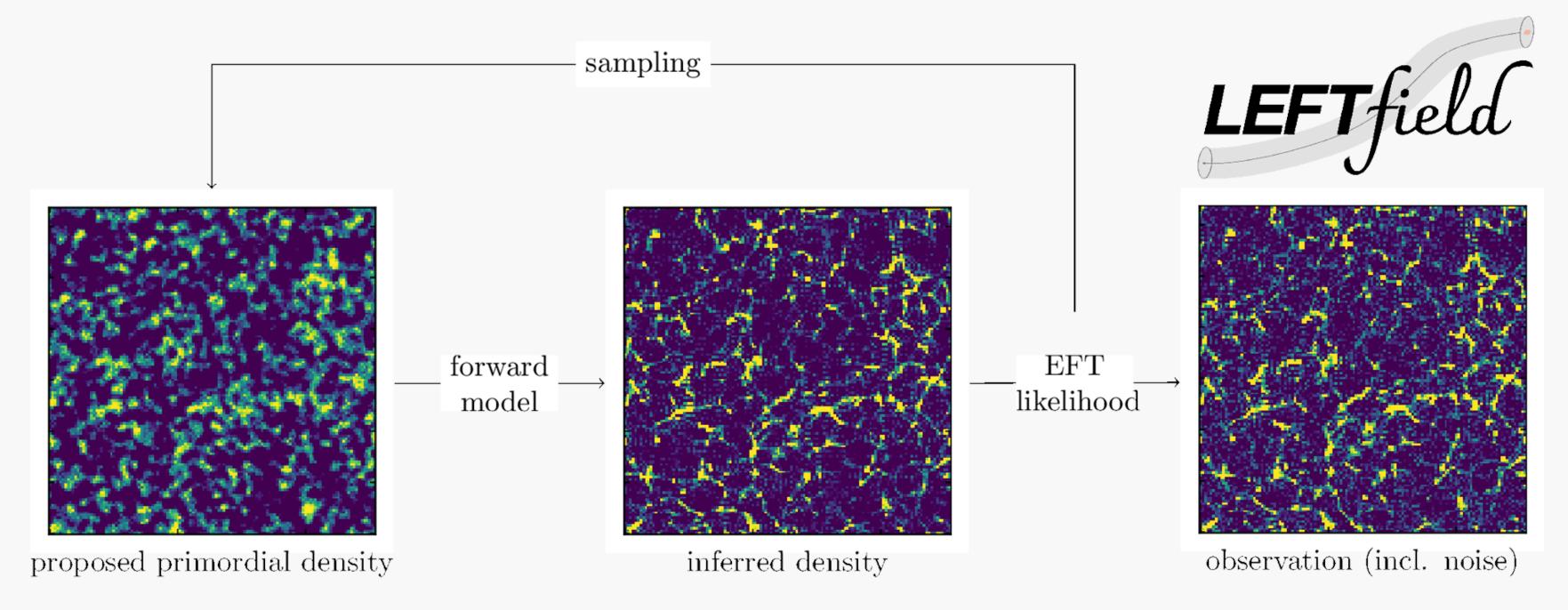
EFT likelihood

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FORWARD MODEL

FIELD-LEVEL BAYESIAN INFERENCE (FBI)



<u>Schmidt (2021)</u>

Credits: Julia Stadler

SIMULATION-BASED INFERENCE (SBI)

parameters summary drawn from statistics prior $\theta \rightarrow \textit{LEFT field} \rightarrow \delta_g \rightarrow \mathbf{X} \rightarrow$ density estimator posterior $\Rightarrow \mathcal{P}(\boldsymbol{\theta}, \{b_O\}, \{\sigma_{\varepsilon}\} | \mathbf{x}^{\text{obs}})$

OC convergence test

$$egin{align} k_{ ext{max}} &= 0.14 h \, ext{Mpc}^{-1} \ L_{ ext{box}} &= 2000 h^{-1} \, ext{Mpc} \ \Delta k pprox 2 k_f \ \end{gathered}$$

size of data vector:

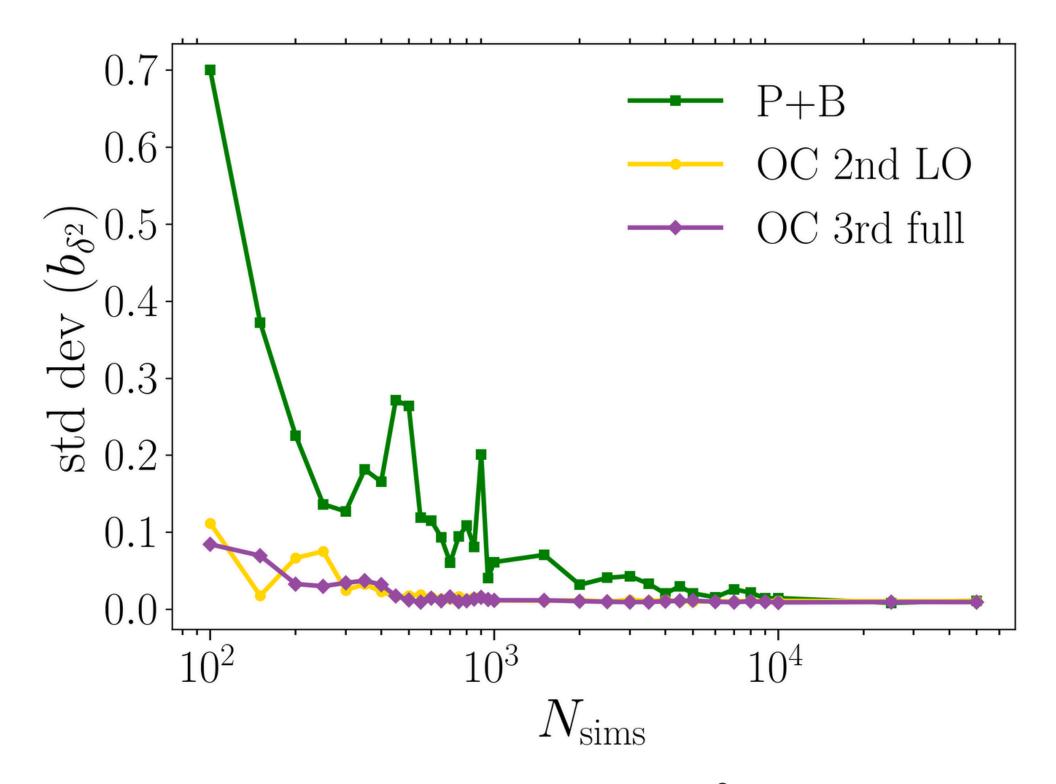
PB: 945 (PBT: ~ 15 000)

OC 2nd, LO: 40

OC 2nd, full: 60

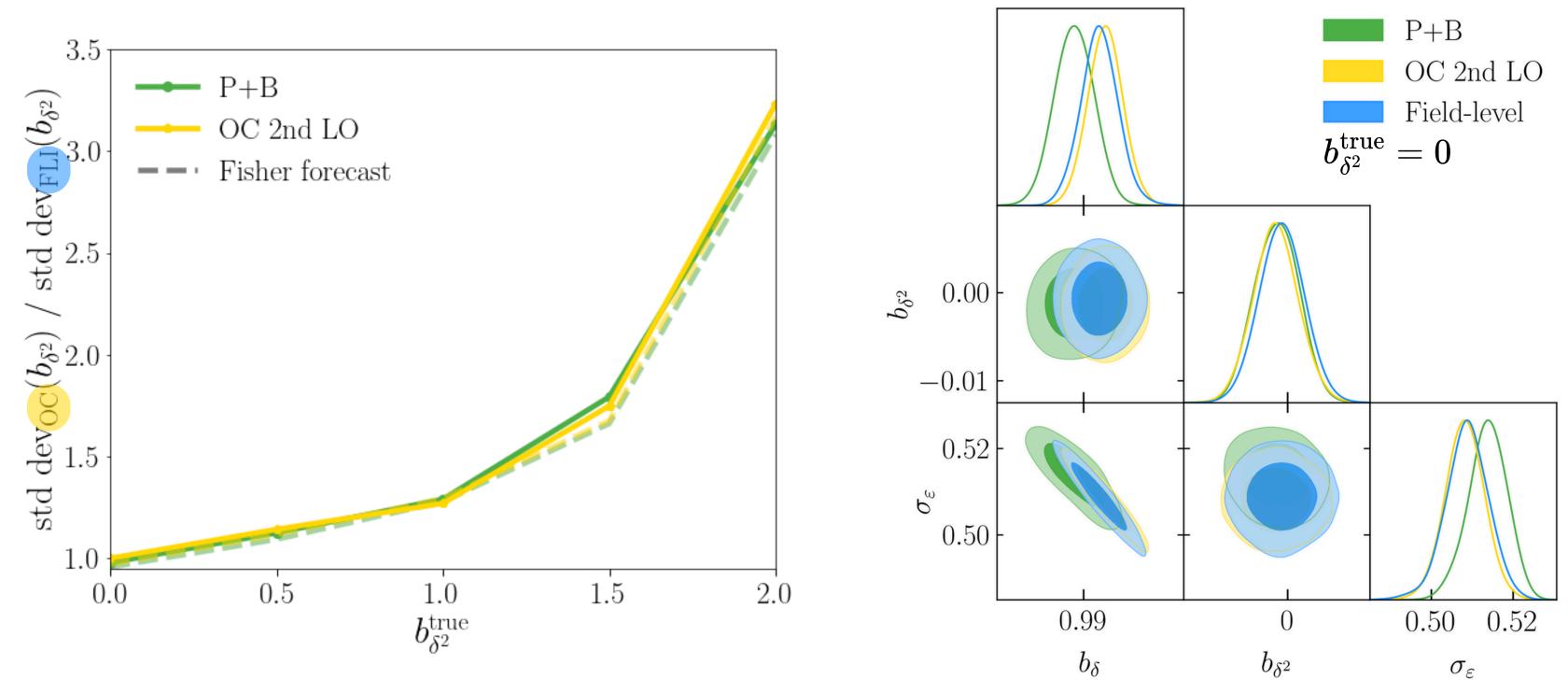
OC 3rd, LO: 80

OC 3rd, full: 120

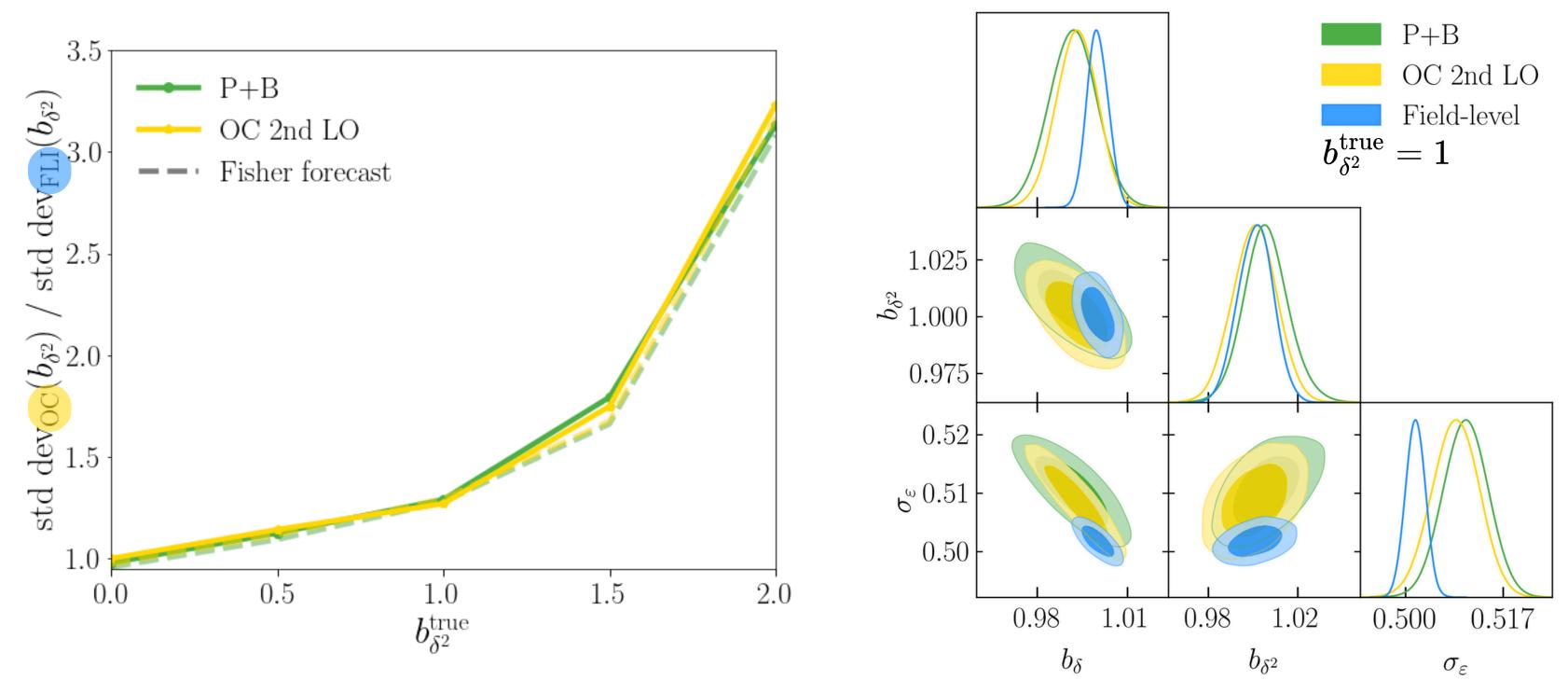


convergence bellow 10³ sims

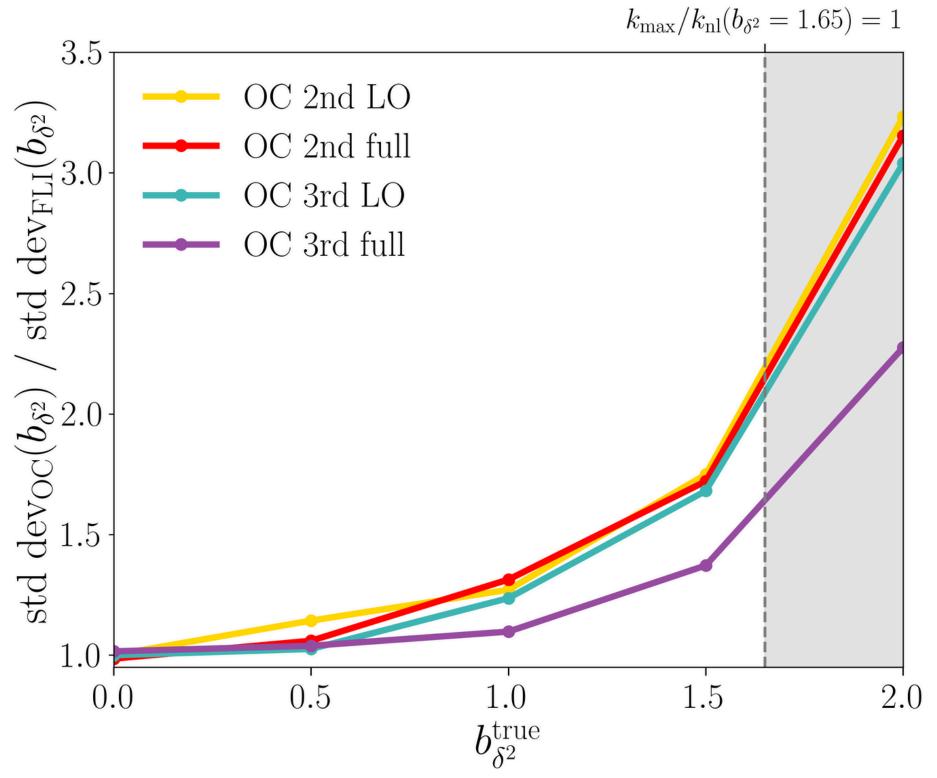
FBI to SBI comparison - OC2ndLO (vs P+B)



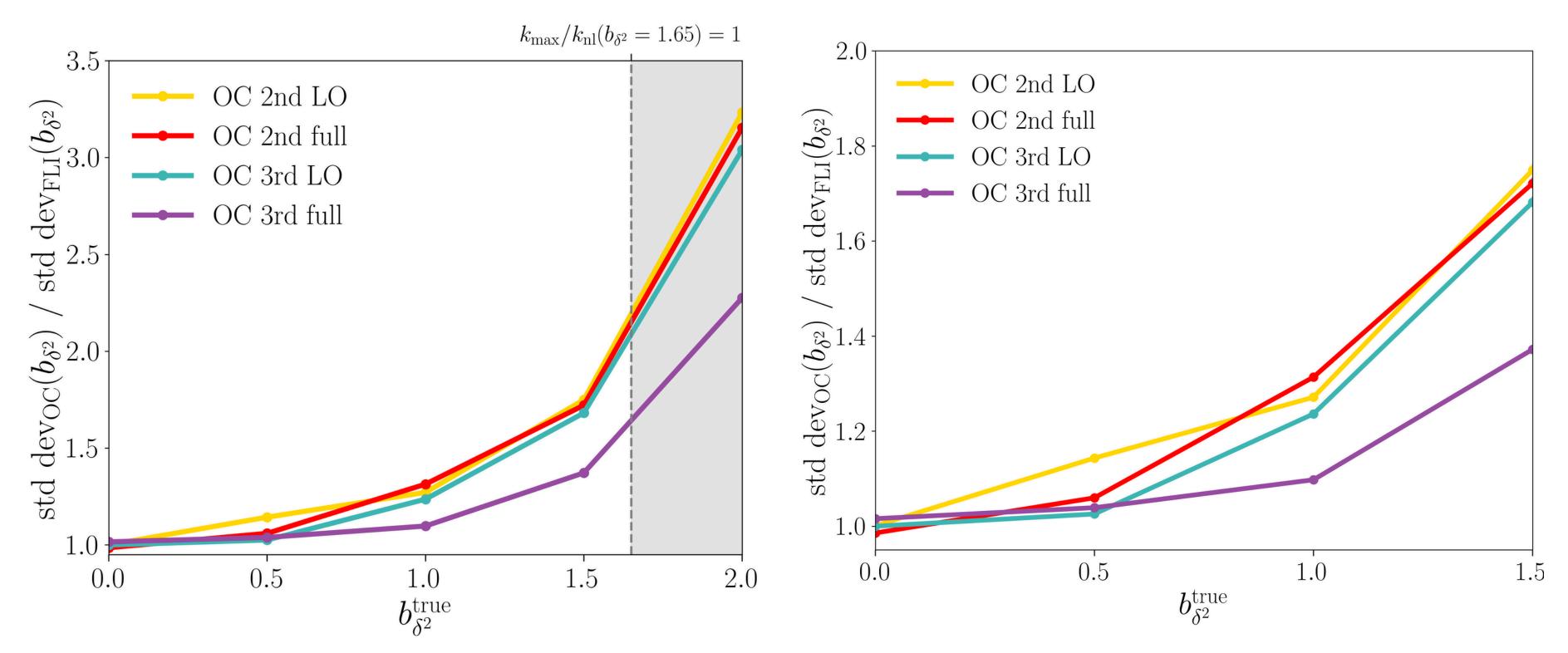
FBI to SBI comparison - OC2ndLO (vs P+B)



FBI to SBI comparison - OC cases



FBI to SBI comparison - OC cases

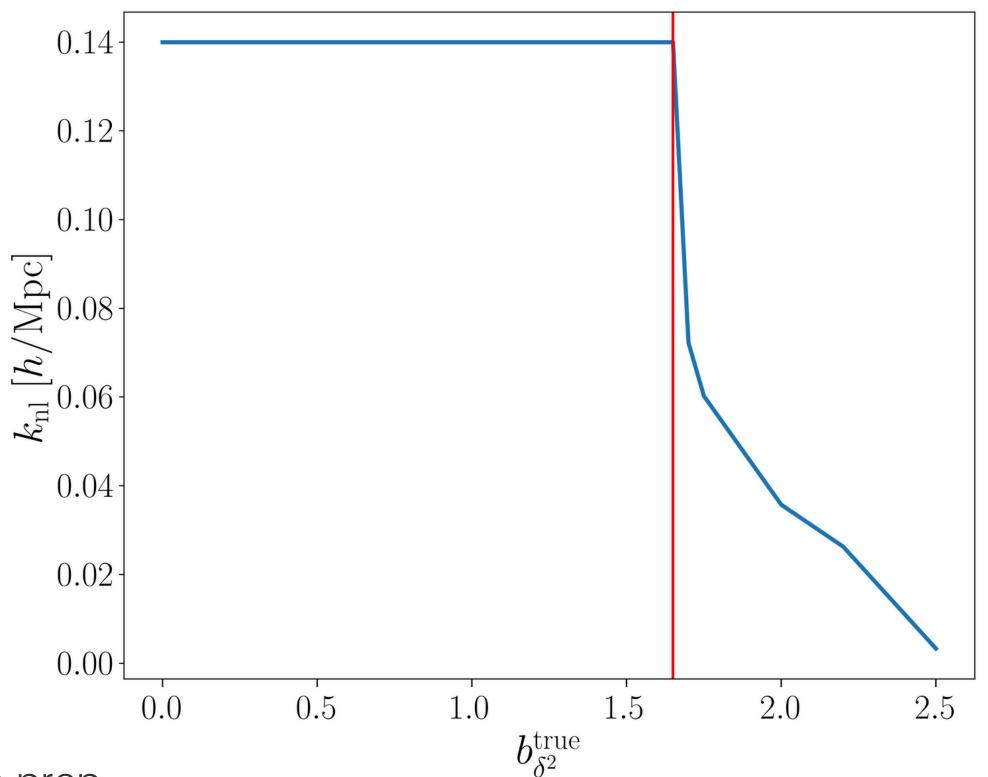


KEY TAKEAWAYS

- Field-level inference with LEFTfield is a powerful tool for galaxy clustering analysis
- Apples-to-apples comparison of FBI and SBI summaries shows that there is a lot of reliable information beyond 2+3-point functions
- Operator correlators (OC) offer an efficient way of including higher order n-point functions in our analysis
- Going higher order in OCs extracts more information in the perturbative regime

DETERMINING k_nl

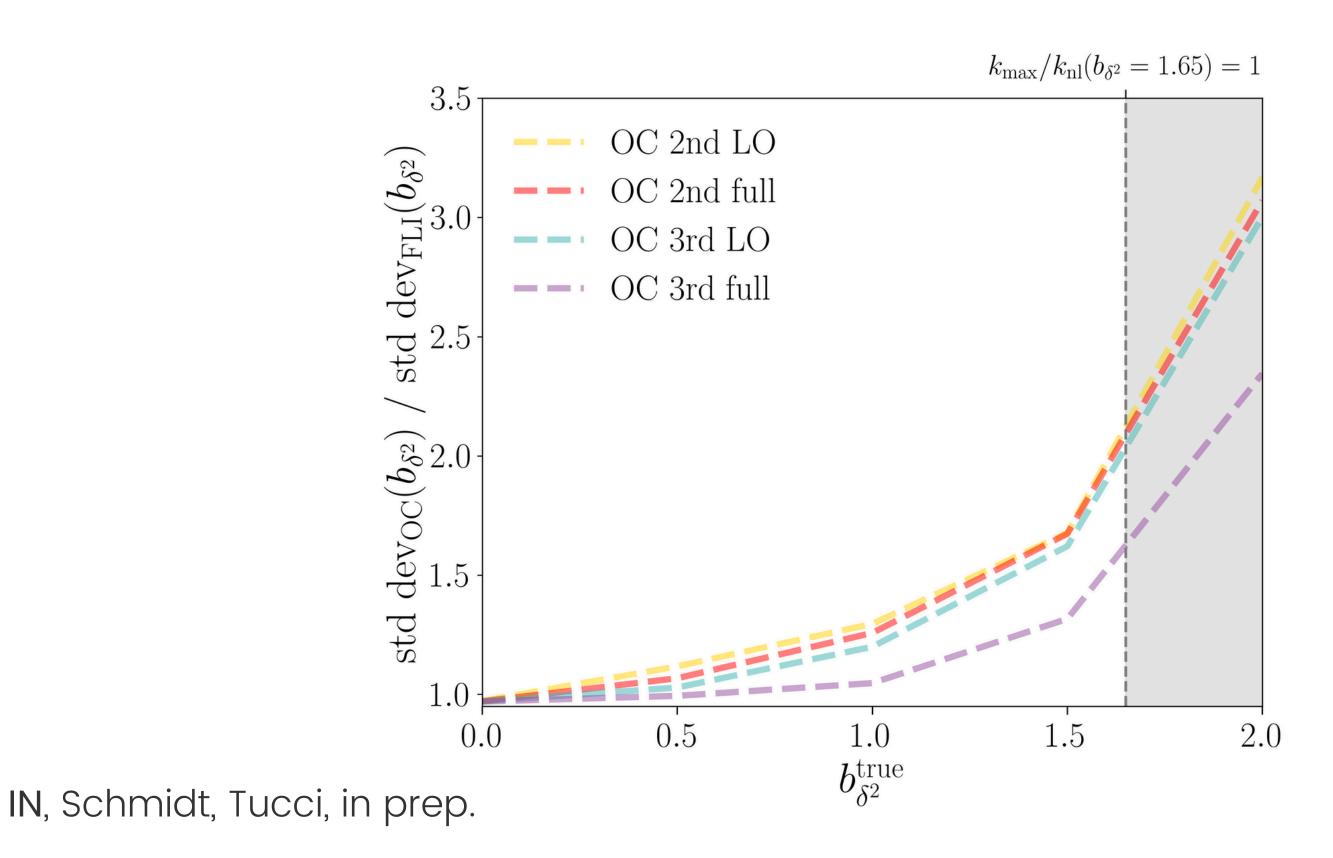
(backup slides)



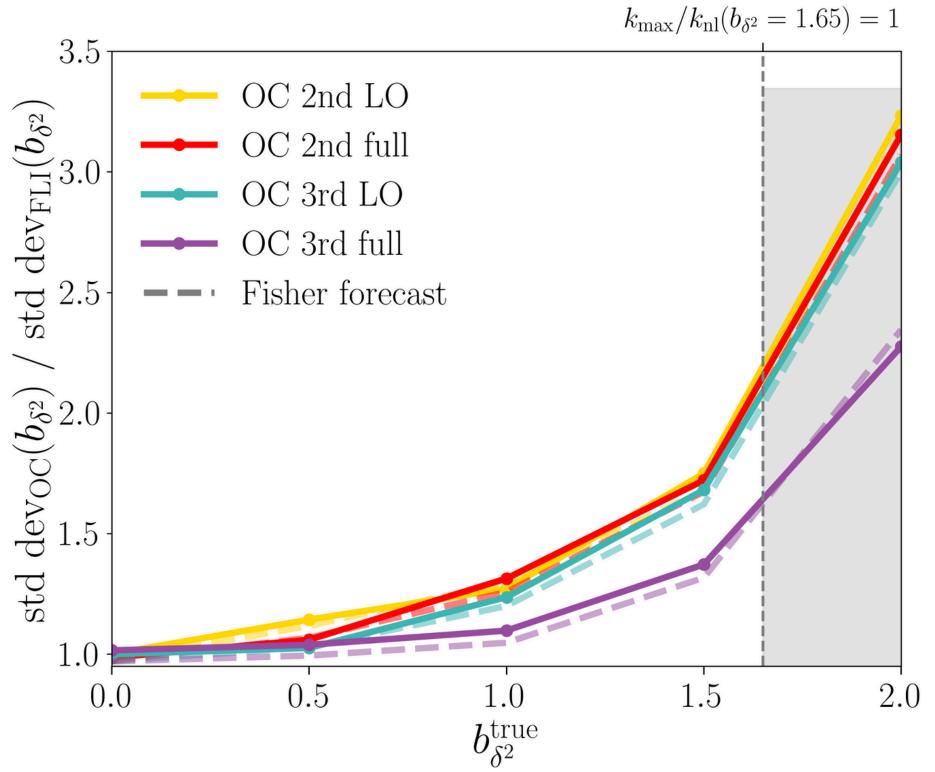
$$\max_{k \leq k_{ ext{nl}}} \ rac{P_{ ext{1-loop}}(k)}{P_{ ext{tree}}(k)} = 1$$

FBI to SBI comparison - OC Fisher forecast

(backup slides)



FBI to SBI comparison - OC cases with Fisher (backup slides)



FBI to SBI comparison - OC cases with Fisher (backup slides)

