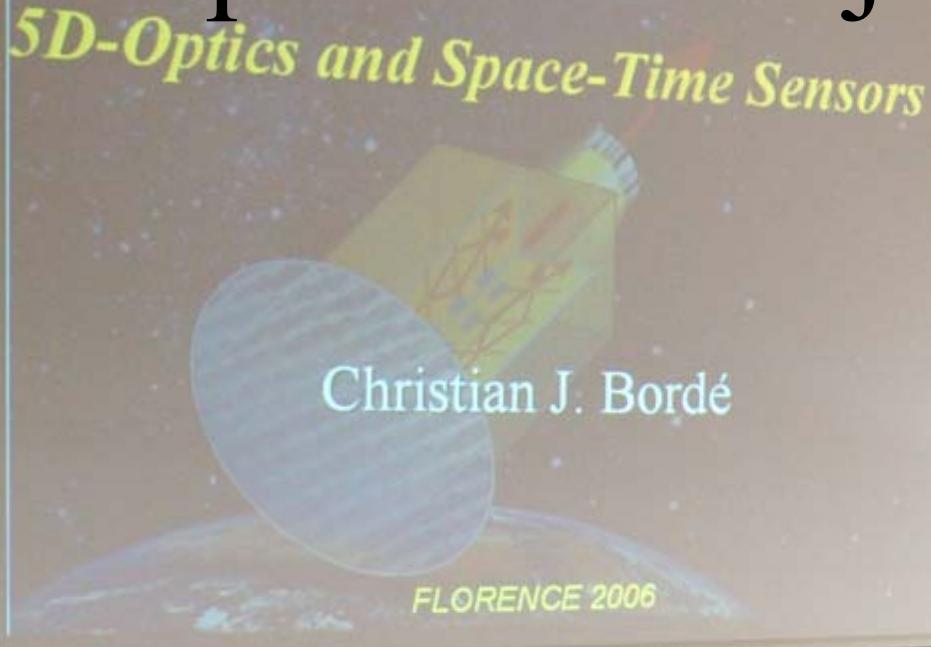


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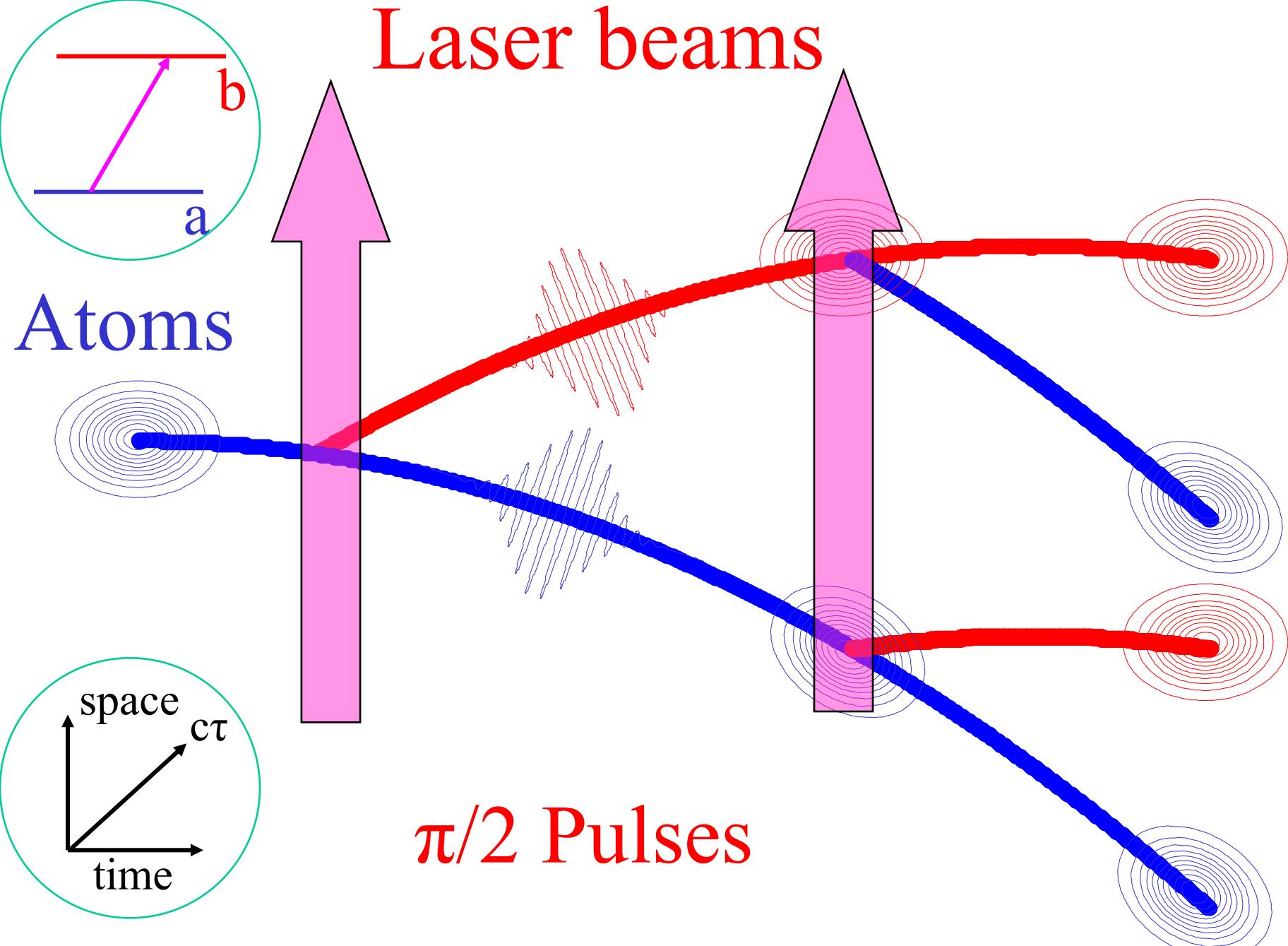
Comparison of atom and photon interferometers using 5D optics

Christian J. Bordé

Académie des Sciences

*FLORENCE
2009*

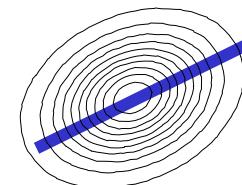
Laser beams



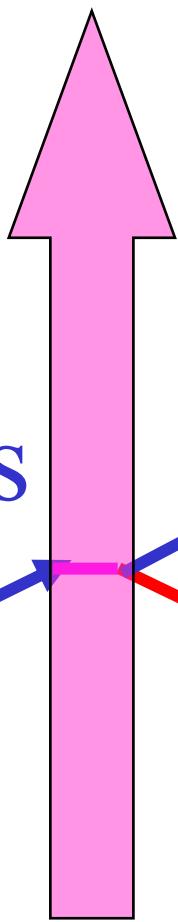
Total phase=Action integral+End splitting+Beam splitters

Laser beams

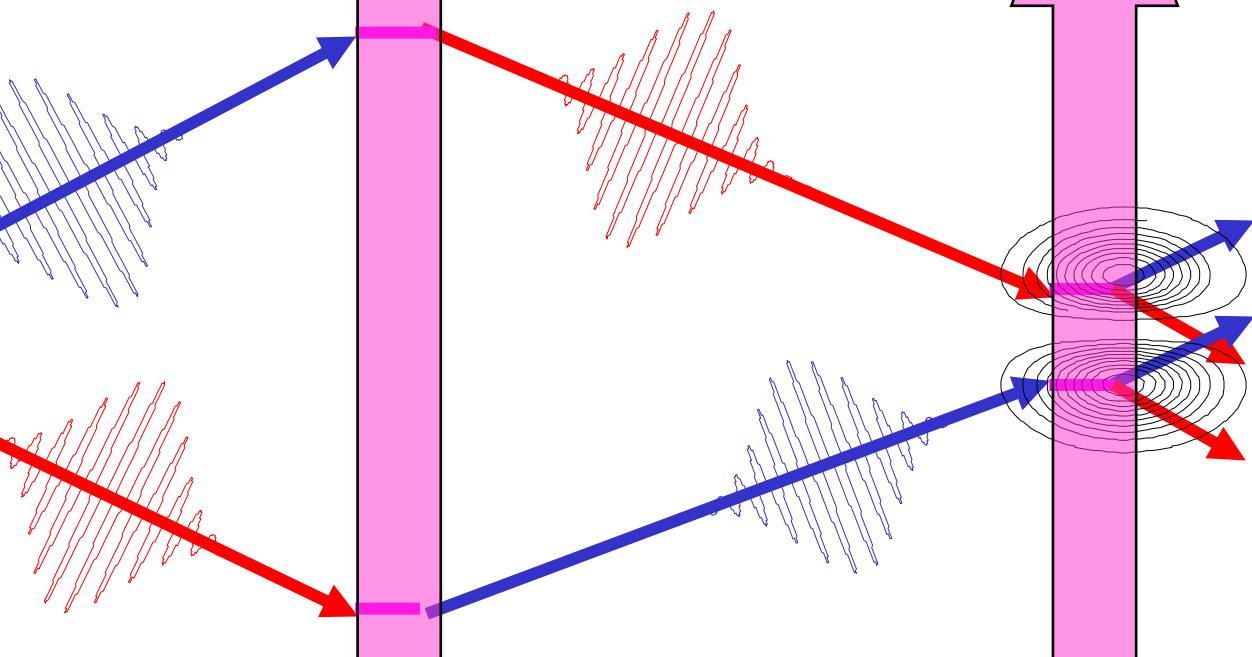
Atoms



$\pi/2$



Laser beams



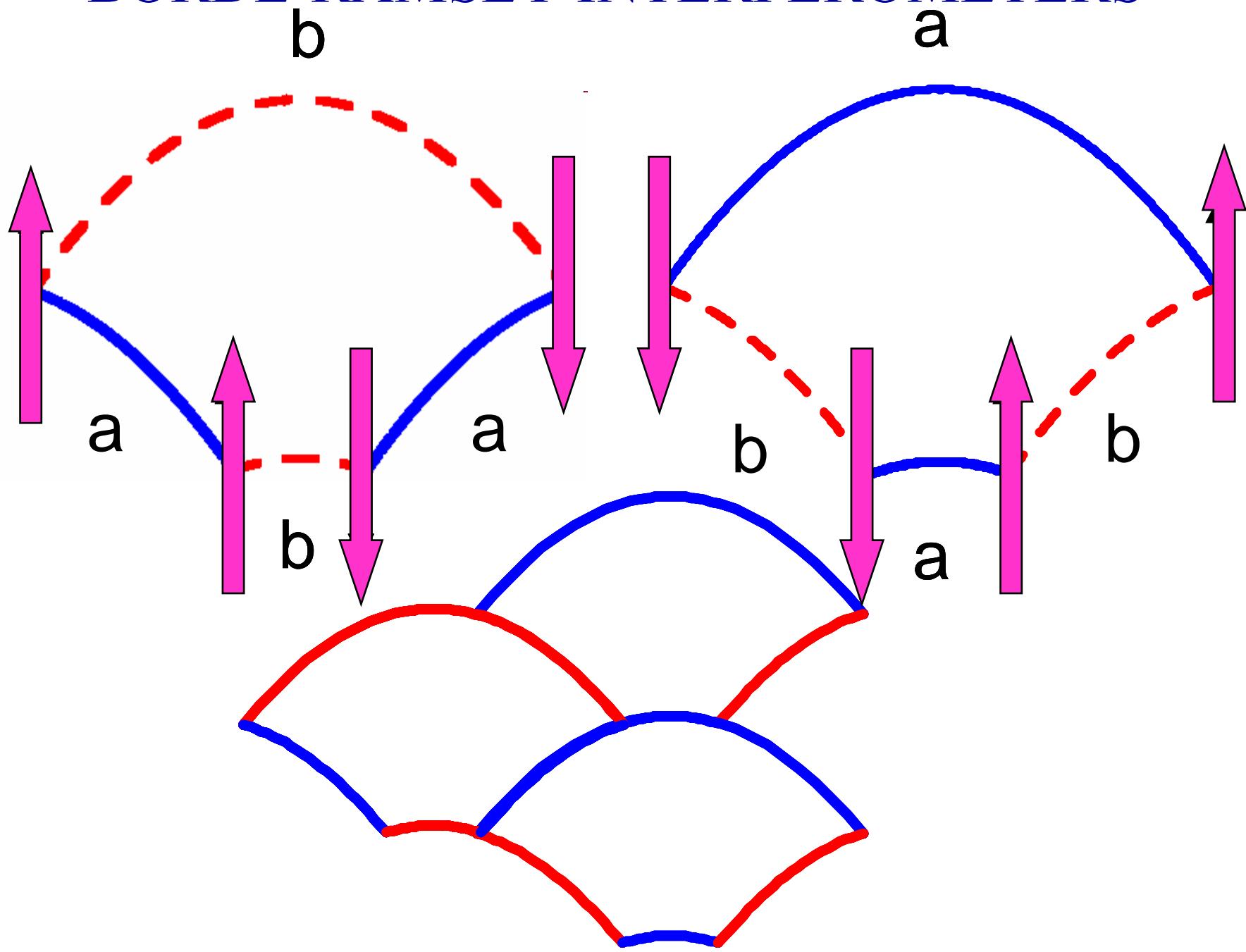
π

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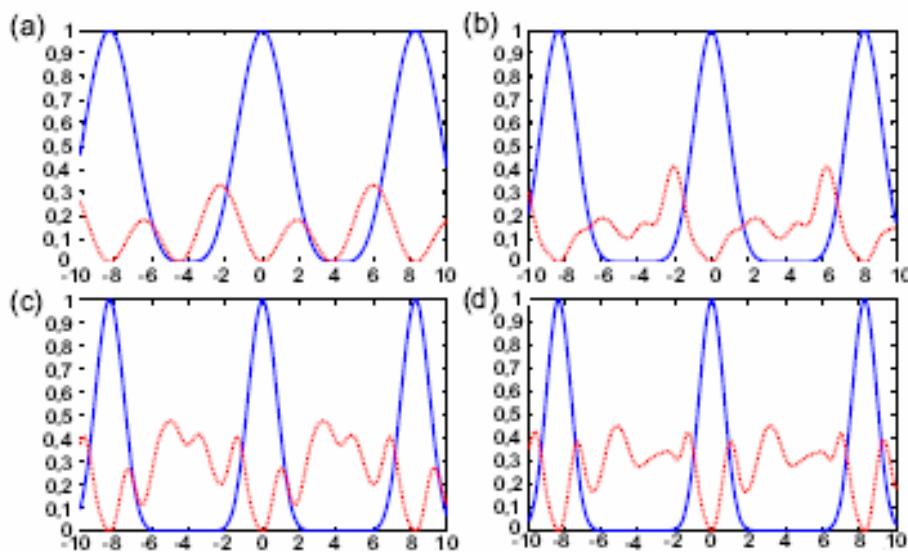
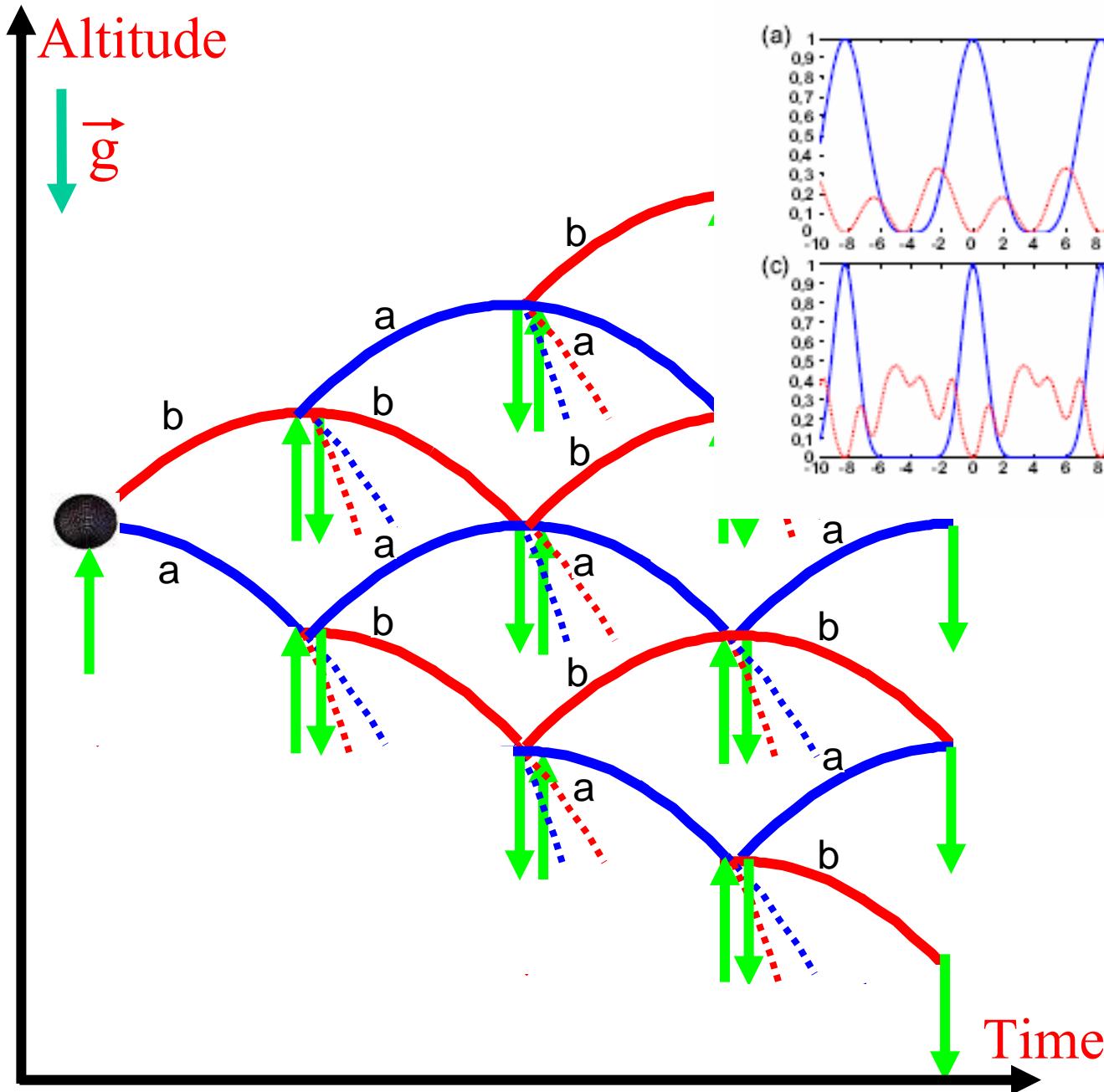
BORDÉ-CHU INTERFEROMETER

Total phase=Action integral+End splitting+Beam splitters

BORDÉ-RAMSEY INTERFEROMETERS



Multiple wave interferometer



NEW OPTICAL ATOMIC INTERFEROMETERS FOR PRECISE MEASUREMENTS OF RECOIL SHIFTS. APPLICATION TO ATOMIC HYDROGEN

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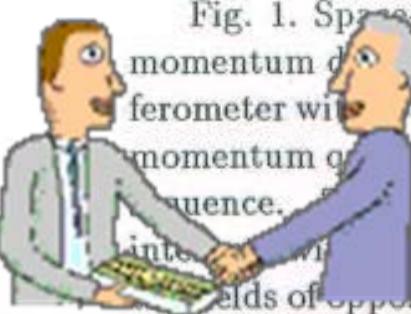
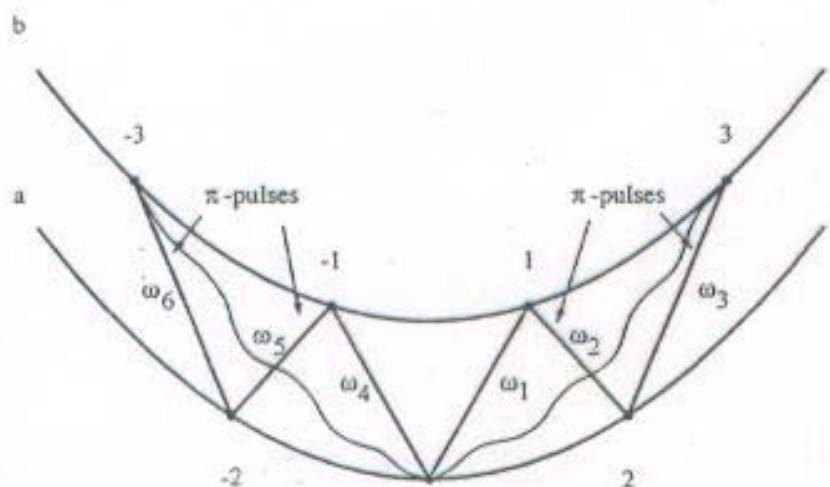
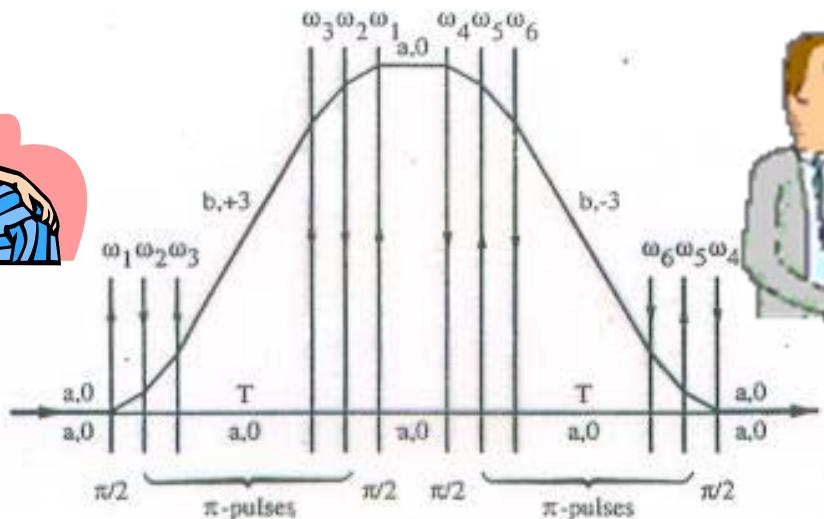
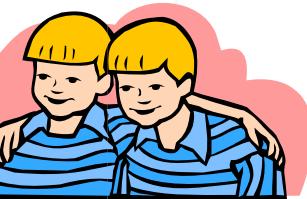
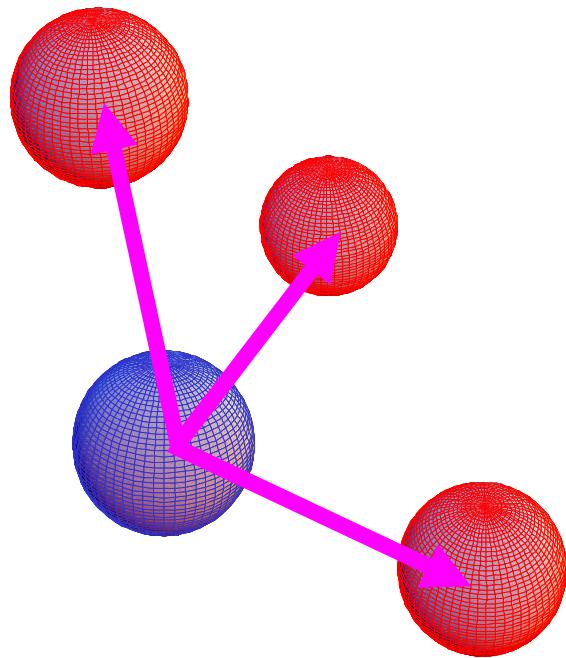


Fig. 1. Space-time and energy-momentum diagrams for an interferometer with $|m_1| = 3$ exchanged momentum quanta per interaction sequence. A two-level system interacts with selective multiphoton fields of opposite directions either perpendicular or collinear to the atomic motion. The space-time diagram displays the deflection along the optical axis versus the proper time in the "atomic frame" at the velocity \vec{p}_0/M . A coherent superposition of the two states $|a, 0\rangle$ and $|b, m_1\rangle$ is created (wiggly line) and travels freely during the time T leading to a phase shift $\varphi = (\omega_1 - \omega_2 + \omega_3 + \omega_4 - \omega_5 + \omega_6 - 2\omega_0 - 18\delta)T$. A second interferometer with opposite recoil shift is obtained by exchanging the roles of states a and b .



$$\square \equiv g^{\mu\nu} \nabla_\mu \nabla_\nu$$

$$\square \varphi = 0$$

$$p^\mu p_\mu = M^2 c^2 \rightarrow \square \varphi + \frac{M^2 c^2}{\hbar^2} \varphi = 0$$

FROM 3 TO 4 SPATIAL DIMENSIONS

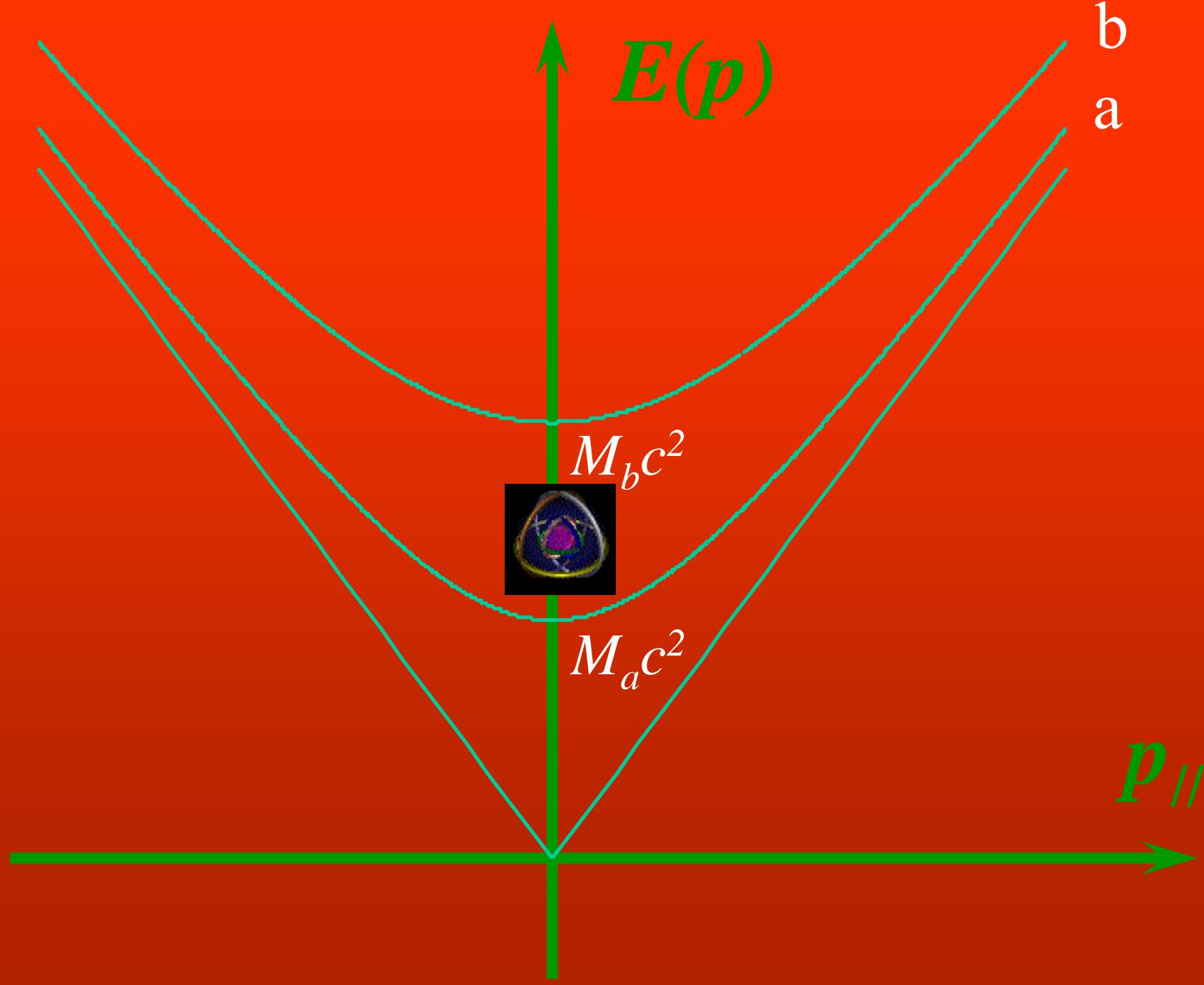
$$\varphi(x, c\tau) = \exp\left[i \frac{Mc^2}{\hbar}(\tau - \tau_0)\right] \varphi(x, Mc)$$

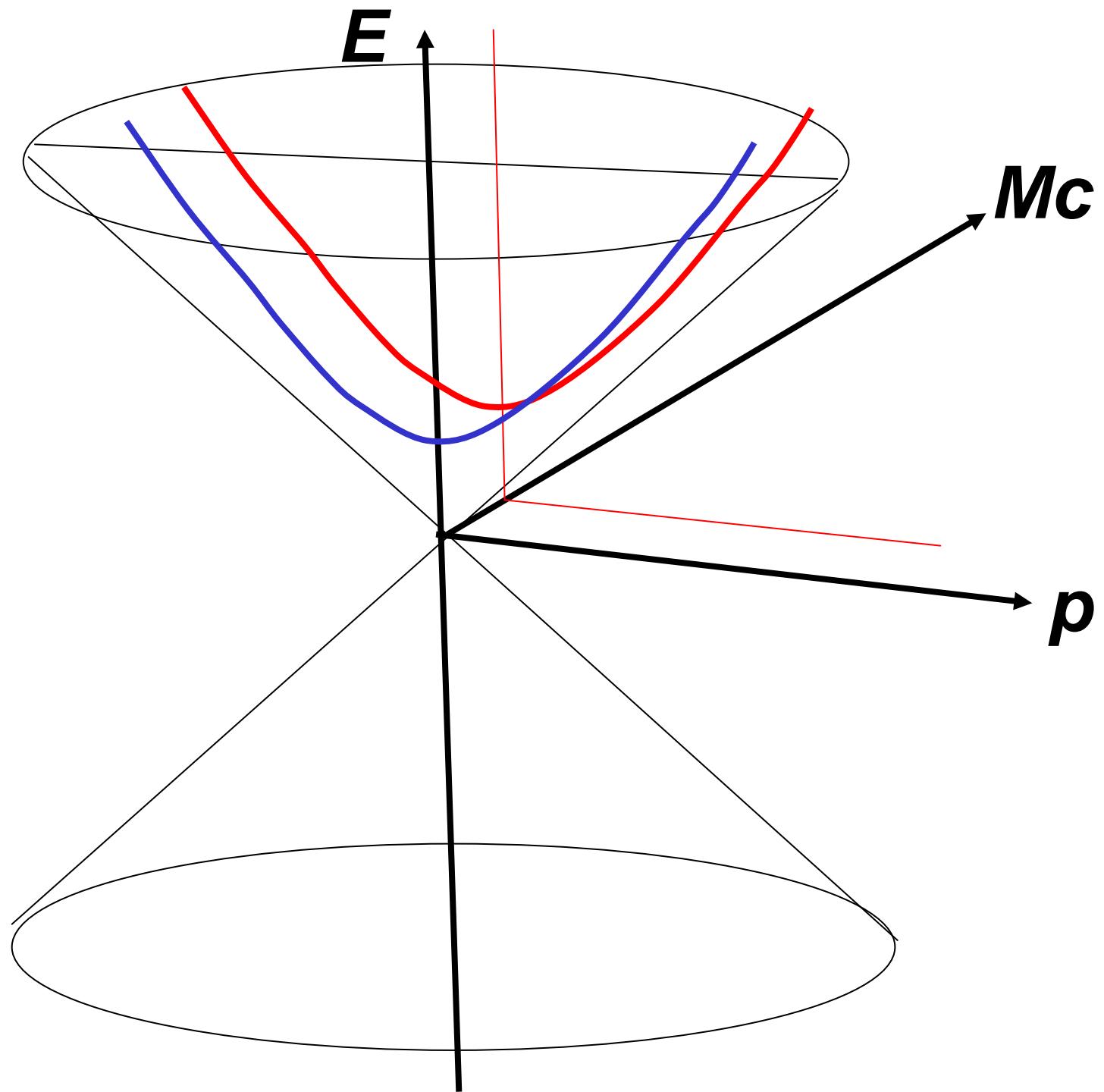
$$i\hbar \frac{\partial \varphi(x, c\tau)}{\partial \tau} = -Mc^2 \varphi(x, \tau)$$

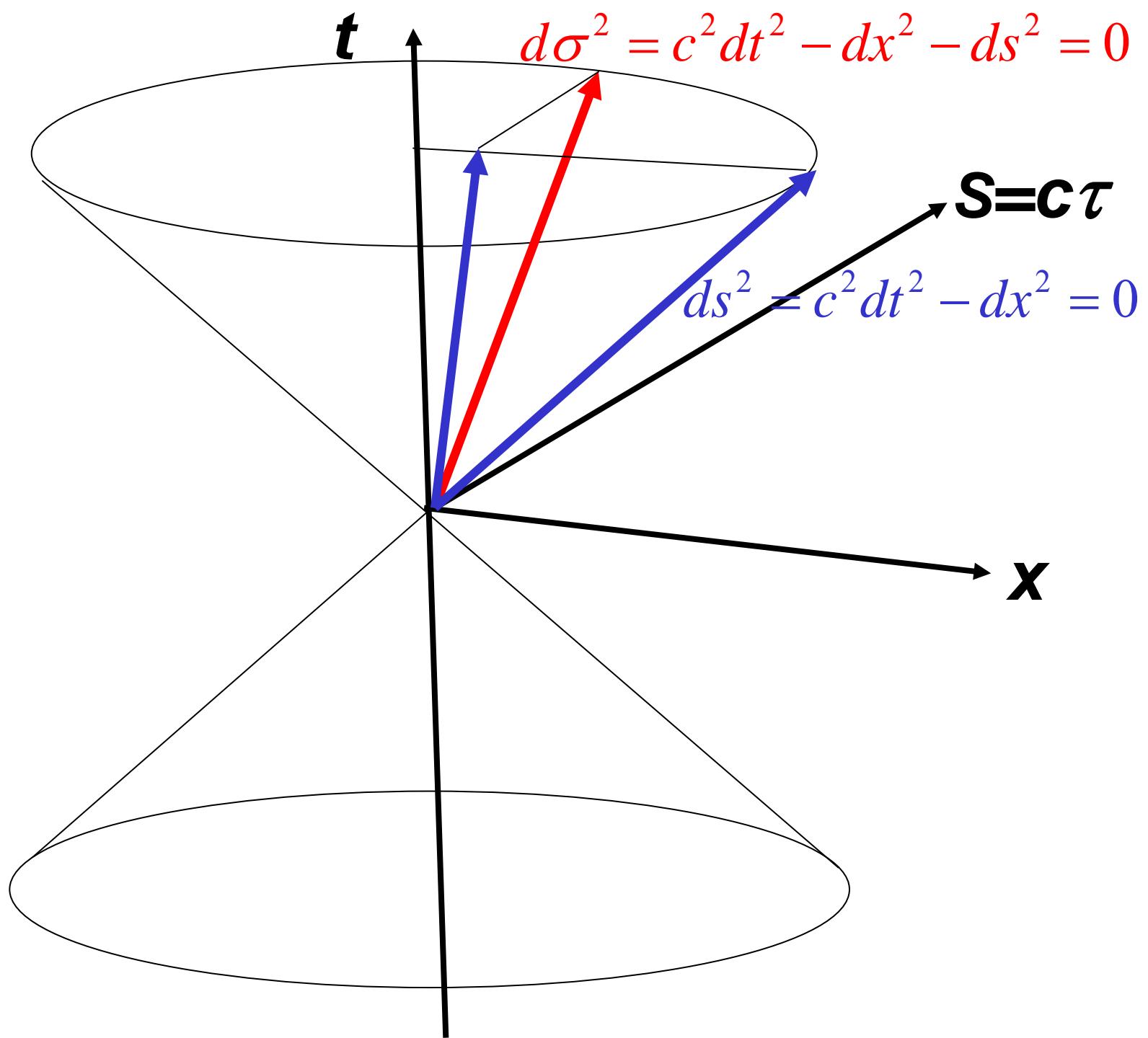
$$\frac{1}{c^2} \frac{\partial^2 \varphi}{\partial t^2} - \Delta \varphi + \frac{M^2 c^2}{\hbar^2} \varphi = 0$$

$$\hat{\square} \varphi \equiv \frac{1}{c^2} \frac{\partial^2 \varphi}{\partial t^2} - \Delta \varphi - \frac{1}{c^2} \frac{\partial^2 \varphi}{\partial \tau^2} = 0$$

$$\varphi(x, c\tau) = \int \frac{d(Mc)}{\sqrt{2\pi\hbar}} \exp\left[i \frac{Mc^2}{\hbar}(\tau - \tau_0)\right] \varphi(x, Mc)$$







OPTICAL PATH & FERMAT'S PRINCIPLE IN (4+1)D

$$\hat{\square}\varphi \equiv \square\varphi - \frac{1}{c^2} \frac{\partial^2 \varphi}{\partial \tau^2} = 0$$

eikonal equation in 5D ($\hat{\mu}, \hat{\nu} = 0, 1, 2, 3, 4$):

$$g^{\hat{\mu}\hat{\nu}} \partial_{\hat{\mu}}\phi \partial_{\hat{\nu}}\phi = 0$$

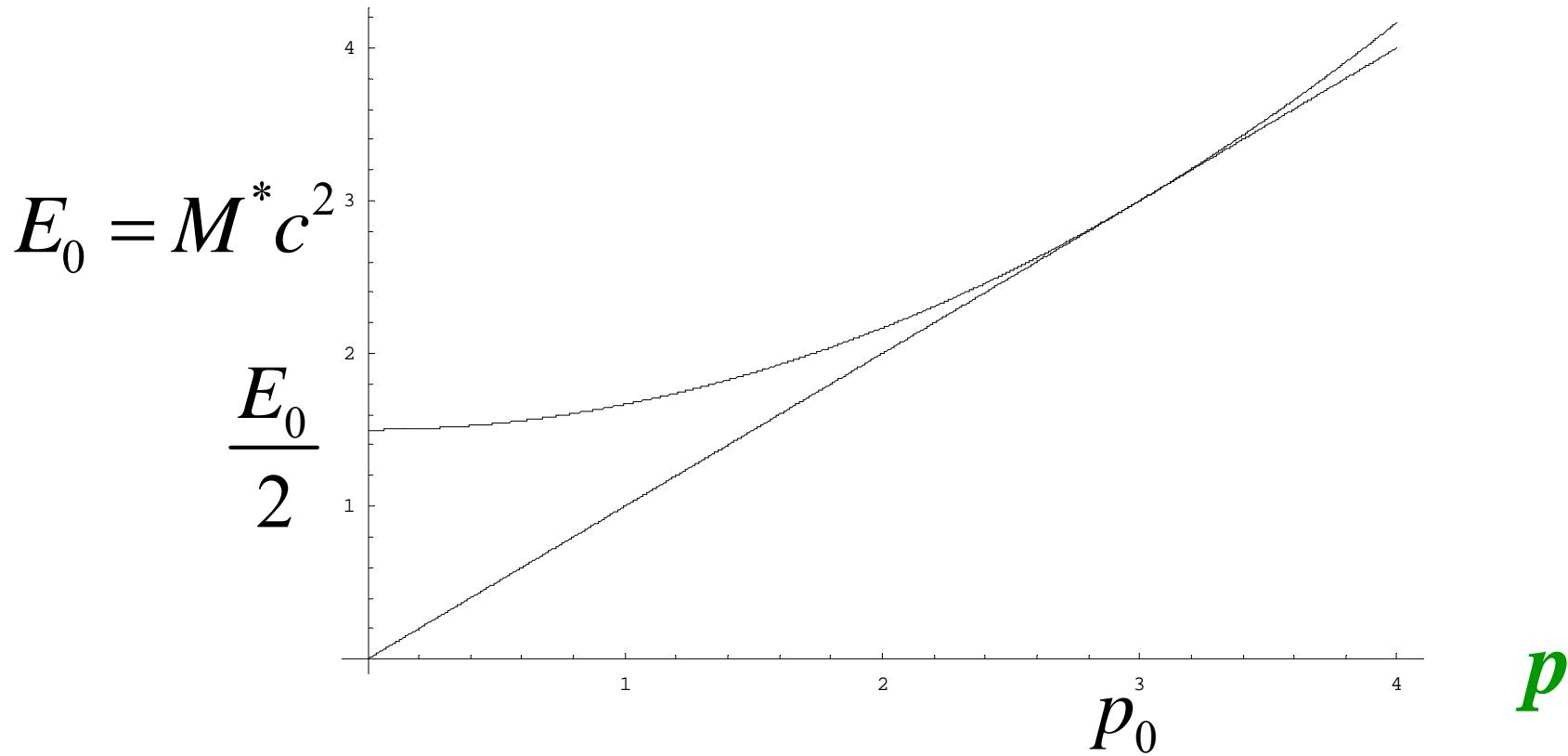
$$\phi = -\frac{E}{h} \left((t - t_0) - \int \frac{dl^{(4)}}{c\sqrt{g_{00}}} + \int \frac{g_{j0}}{cg_{00}} dx^j \right)$$

$$dl^{(4)} = \sqrt{-f_{ij}dx^i dx^j + c^2 d\tau^2} \quad \lambda^{(4)} = \frac{h}{E} c \sqrt{g_{00}}$$

$$f_{ij} = g_{ij} - \frac{g_{0i}g_{0j}}{g_{00}}$$

BASICS OF ATOM /PHOTON OPTICS

Parabolic approximation
of slowly varying phase and amplitude
 $E(p)$



BASICS OF ATOM /PHOTON OPTICS

Schroedinger-like equation for the atom /photon field:

$$i\hbar \frac{\partial \varphi}{\partial t} = \frac{M^* c^2}{2} \varphi - \frac{1}{2M^*} \left[p^j p_j + p^4 p_4 \right] \varphi + \frac{1}{2M^*} p_\mu h^{\mu\nu} p_\nu \varphi$$

$$p_j = i\hbar \partial_j; \quad p_4 = i\hbar \partial_{c\tau}; \quad p_0 = M^* c \quad (\hbar\omega/c \text{ for photons})$$

- gravitation field: $h^{00} = -2\vec{g} \cdot \vec{q} / c^2 - \vec{q} \cdot \overset{\Rightarrow}{\vec{\gamma}} \cdot \vec{q} / c^2$

- rotation field: $\vec{h} = -\overset{\Rightarrow}{\vec{\alpha}} \cdot \vec{q} / c$

- gravitational wave: $\overset{\Rightarrow}{\vec{h}} = \overset{\Rightarrow}{\vec{\beta}} - \overset{\Rightarrow}{\vec{\delta}}$

$$H_{ext} = \overset{\Rightarrow}{\vec{p}} \cdot \overset{\Rightarrow}{\vec{\alpha}}(t) \cdot \vec{q} + \overset{\Rightarrow}{\vec{p}} \cdot \overset{\Rightarrow}{\vec{\beta}}(t) \cdot \overset{\Rightarrow}{\vec{p}} / 2M^* - M^* \overset{\Rightarrow}{\vec{q}} \cdot \overset{\Rightarrow}{\vec{\gamma}}(t) \cdot \vec{q} / 2 - M^* \overset{\Rightarrow}{\vec{g}} \cdot \vec{q} + \overset{\Rightarrow}{\vec{f}} \cdot \overset{\Rightarrow}{\vec{p}}$$

ABCD $\xi\phi$ LAW OF ATOM/PHOTON OPTICS

wavepacket(q, t) =

$$\exp \left[i p_c(t) (q - q_c(t)) / \hbar \right] F(q - q_c(t), X(t), Y(t))$$

$$p = (p_x, p_y, p_z, Mc); \quad q = (x, y, z, c\tau)$$

$$q_c(t) = A q_c(t_0) + B p_c(t_0) / M^* + \xi(t, t_0)$$

$$p_c(t) / M^* = C q_c(t_0) + D p_c(t_0) / M^* + \phi(t, t_0)$$

$$X(t) = A X(t_0) + B Y(t_0)$$

$$Y(t) = C X(t_0) + D Y(t_0)$$

Ehrenfest theorem

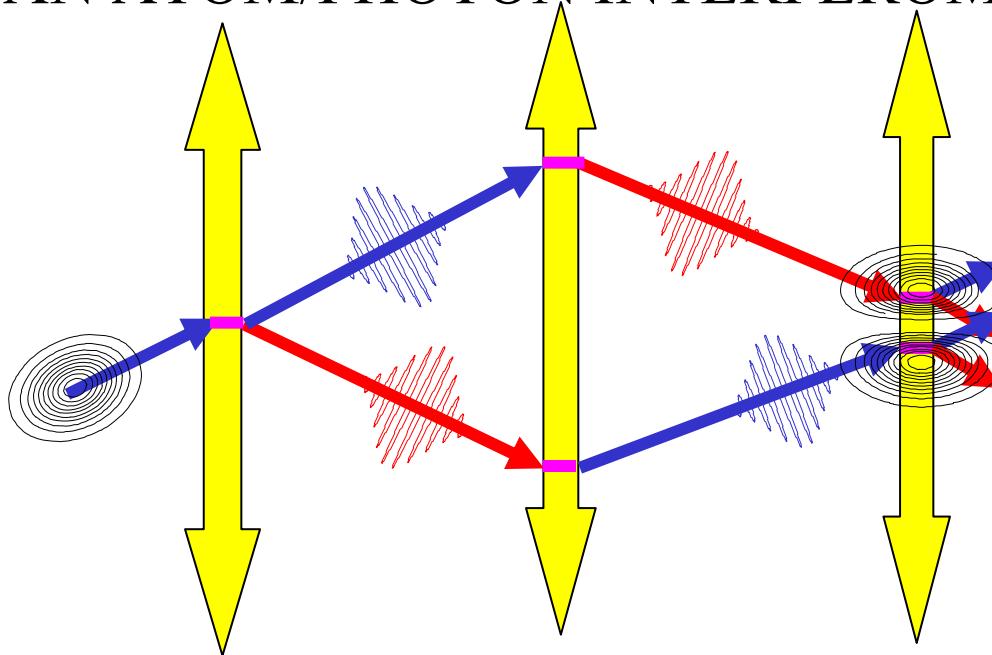
+

Hamilton equations

$$H_{ext} = \vec{p} \cdot \overset{\Rightarrow}{\alpha}(t) \cdot \vec{q} + \vec{p} \cdot \overset{\Rightarrow}{\beta}(t) \cdot \vec{p} / 2M^* - M^* \vec{q} \cdot \overset{\Rightarrow}{\gamma}(t) \cdot \vec{q} / 2 - M^* \vec{g} \cdot \vec{q} + \vec{f} \cdot \vec{p}$$

$$\begin{pmatrix} A(t, t_0) & B(t, t_0) \\ C(t, t_0) & D(t, t_0) \end{pmatrix} = T \exp \left[\int_{t_0}^t \begin{pmatrix} \alpha(t') & \beta(t') \\ \gamma(t') & \alpha(t') \end{pmatrix} dt' \right]$$

GENERAL FORMULA FOR THE PHASE SHIFT OF AN ATOM/PHOTON INTERFEROMETER

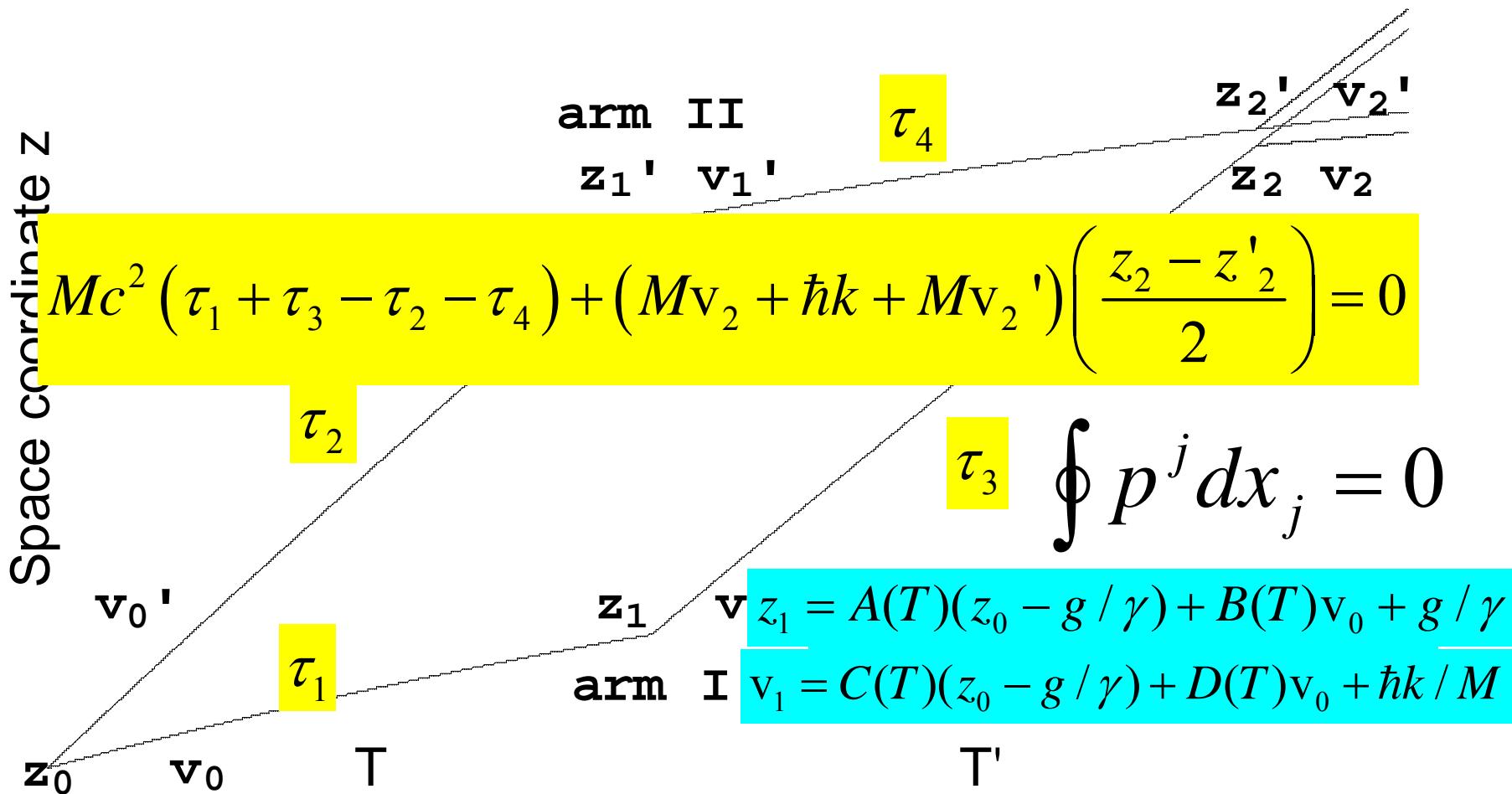


$$\delta\varphi = \sum_{j=1}^N (\delta\tilde{k}_j^{(5)} \cdot q_j^{(5)} + \delta\varphi_j)$$

$$\tilde{k}^{(5)} = \left[(k_x, k_y, k_z, \frac{\omega^{(0)}}{c}), \frac{\omega}{c} \right]; \quad \tilde{q}^{(5)} = [(x, y, z, c\tau), ct]$$

$$\delta\tilde{k}^{(5)} = \tilde{k}_{\beta j}^{(5)} - \tilde{k}_{\alpha j}^{(5)}; \quad q_j^{(5)} = (q_{\beta j}^{(5)} + q_{\alpha j}^{(5)})/2$$

Atomic Gravimeter



$$\delta\varphi = -k(z_2 - z_1 - z_1' + z_0) + k(z_2 - z_2')/2$$

Exact phase shift for the atom gravimeter

$$\begin{aligned}\delta\varphi = & -k(z_2 - z_1 - z'_1 + z_0) + k(z_2 - z'_2)/2 \\ = & \frac{k}{\sqrt{\gamma}} \left\{ \left[\sinh(\sqrt{\gamma}(T + T')) - 2 \sinh(\sqrt{\gamma}T) \right] \left(v_0 + \frac{\hbar k}{2M} \right) \right. \\ & \left. + \sqrt{\gamma} \left[1 + \cosh(\sqrt{\gamma}(T + T')) - 2 \cosh(\sqrt{\gamma}T) \right] \left(z_0 - \frac{g}{\gamma} \right) \right\}\end{aligned}$$

which can be written to first-order in γ , with $T=T'$:

$$\delta\varphi = kgT^2 + k\gamma T^2 \left[\frac{7}{12} gT^2 - \left(v_0 + \frac{\hbar k}{2M} \right) T - z_0 \right]$$

Reference: Ch. J. B., Theoretical tools for atom optics and interferometry,
C.R. Acad. Sci. Paris, 2, Série IV, p. 509-530, 2001

ARBITRARY 3D TIME-DEPENDENT GRAVITO-INERTIAL FIELDS

Hamiltonian: $H = \vec{p} \cdot \overset{\Rightarrow}{\alpha}(t) \cdot \vec{q} + \vec{p} \cdot \overset{\Rightarrow}{\beta}(t) \cdot \vec{p} / 2M^* - M^* \vec{q} \cdot \overset{\Rightarrow}{\gamma}(t) \cdot \vec{q} / 2$

Hamilton's eqns: $\begin{pmatrix} A & B \\ C & D \end{pmatrix} = T \exp \int dt \begin{pmatrix} \alpha & \beta \\ \gamma & \alpha \end{pmatrix}$

Example: Phase shift induced by a gravitational wave

Einstein coord.: $\overset{\Rightarrow}{\beta} = 1 + h \cos(\xi t + \phi)$, $\overset{\Rightarrow}{\gamma} = 0$, with $\overset{\Rightarrow}{h} = \{h^{ij}\}$

Fermi coord.: $\overset{\Rightarrow}{\beta} = 1$, $\overset{\Rightarrow}{\gamma} = (\xi^2 / 2) \overset{\Rightarrow}{h} \cos(\xi t + \phi)$

Einstein coord.: $\begin{cases} A = 1 \\ B = t + \frac{h}{\xi} [\sin(\xi t + \phi) - \sin \phi] \end{cases}$

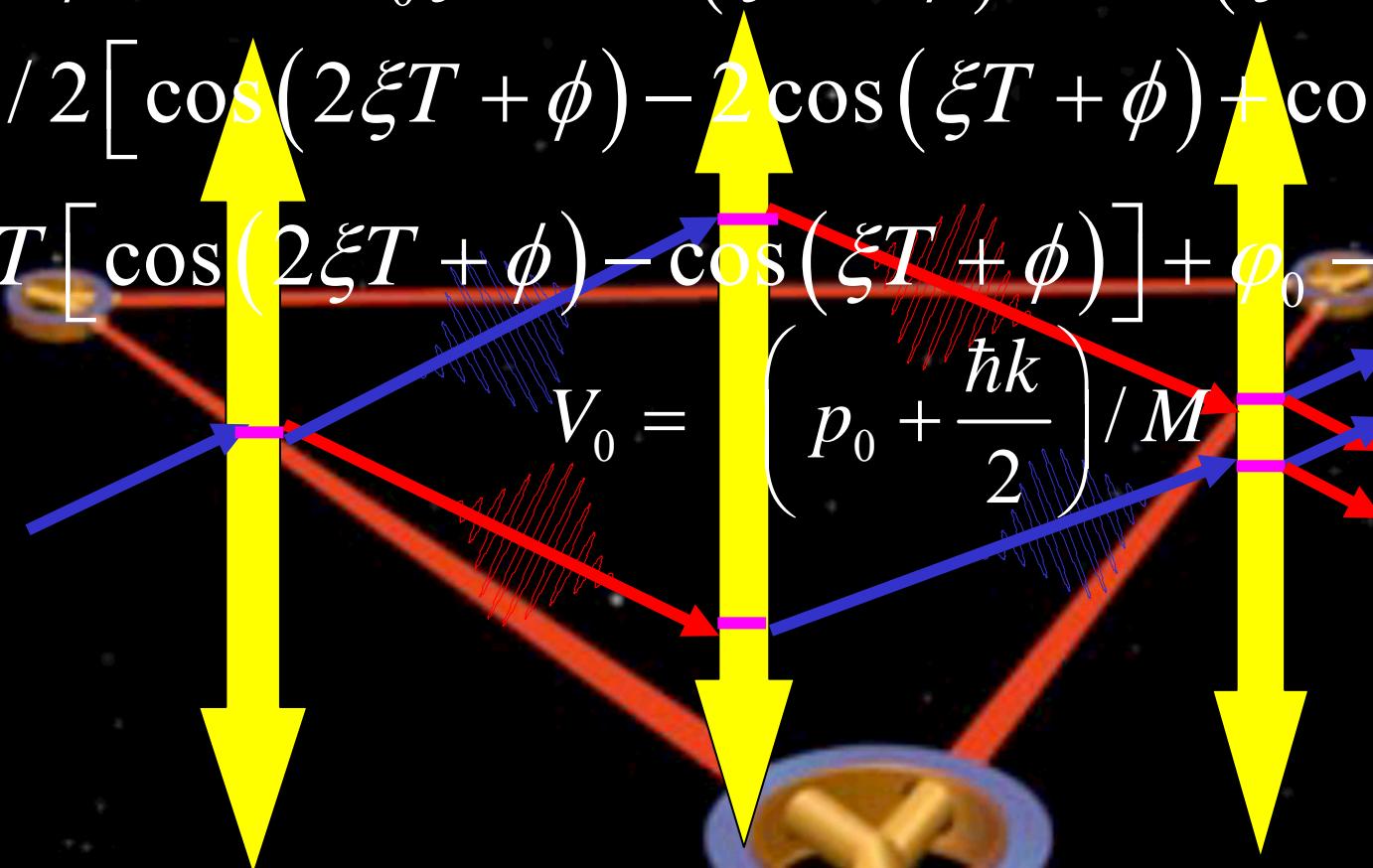
Fermi coord.: $\begin{cases} A = 1 - \frac{h}{2} [\cos(\xi t + \phi) - \cos \phi] - \frac{h\xi t}{2} \sin \phi \\ B = t + \frac{h}{\xi} [\sin(\xi t + \phi) - \sin \phi] - \frac{ht}{2} [\cos(\xi t + \phi) + \cos \phi] \end{cases}$

Atomic phase shift induced by a gravitational wave

$$\delta\phi = -khV_0\xi T^2 \sin(\xi T + \phi) \text{sinc}^2(\xi T / 2)$$

$$-khq_0/2 [\cos(2\xi T + \phi) - 2\cos(\xi T + \phi) + \cos \phi]$$

$$-khV_0T [\cos(2\xi T + \phi) - \cos(\xi T + \phi)] + \varphi_0 - 2\varphi_1 + \varphi_2$$



Ch.J. Bordé, Gen. Rel. Grav. 36 (March 2004)

Ch.J. Bordé, J. Sharma, Ph. Tourrenc and Th. Damour,

Theoretical approaches to laser spectroscopy in the presence of gravitational fields, J. Physique Lettres 44 (1983) L983-990

Classification

Physics Abstracts
32.70J — 04.80

Theoretical approaches to laser spectroscopy in the presence of gravitational fields

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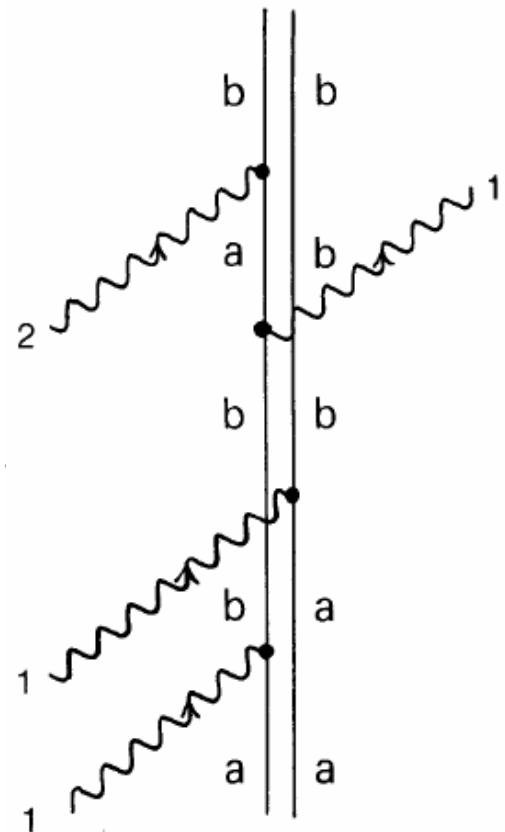
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(Reçu le 3 août 1983, accepté le 27 octobre 1983)



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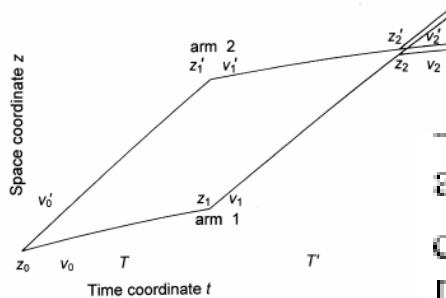


Figure 18. Space-time diagram of the atomic gravimeter.

what could be obtained with light rays, since the times T can be of the order of 1 s and gT^2 is of the order of 10 m compared with an optical wavelength. The next term is a significant correction due to the gravitational field gradient γ . One can also measure directly these field gradients, with two gravimeters using two clouds of cold atoms and sharing the same vertical laser beam splitters. It is then no longer necessary to have a very sophisticated inertial platform for the reference mirror and it is possible to measure directly the differential acceleration between these two clouds. This is the principle, illustrated in Figure 19, of the gradiometers

developed first in Stanford then in Yale by Kasevich and co-workers [27, 59]. The current sensitivity is $4 \times 10^{-9} \text{ s}^{-2}/\sqrt{\text{Hz}}$ and the uncertainty is $1 \times 10^{-9} \text{ s}^{-2}/\sqrt{\text{Hz}}$ for an extrapolated 10 m separation between accelerometers. For the future, this principle may be considered for gravitational wave detection in space

accelerometers. For the future, this principle may be considered for gravitational wave detection in space [60].

The test mass energy m_ec^2 has to be replaced by the photon energy $h\nu$ in the case of light waves. This expression can be derived by a number of approaches, the best of which is to use the rotation operator in the Schrödinger equation and the derived propagator and $ABCD$ matrices, as outlined above, which gives the Sagnac shift thanks to an exact formula. The same approach applies to the trapezoid geometry used in the previous experiment as well as to the parallelogram geometry, as suggested in [24], analogue of the Mach-Zehnder optical interferometer, which has been used in more recent experiments and which has the advantage of being insensitive to laser detuning. For the illustration, the formula calculated with the $ABCD\xi$ formalism is

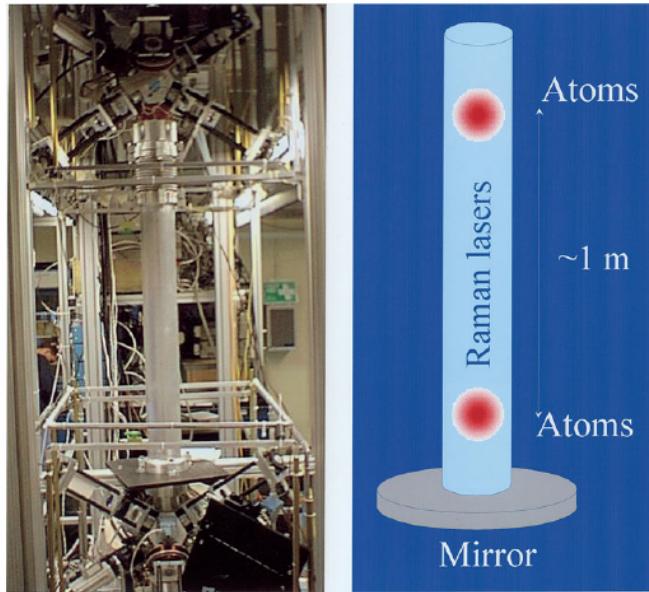
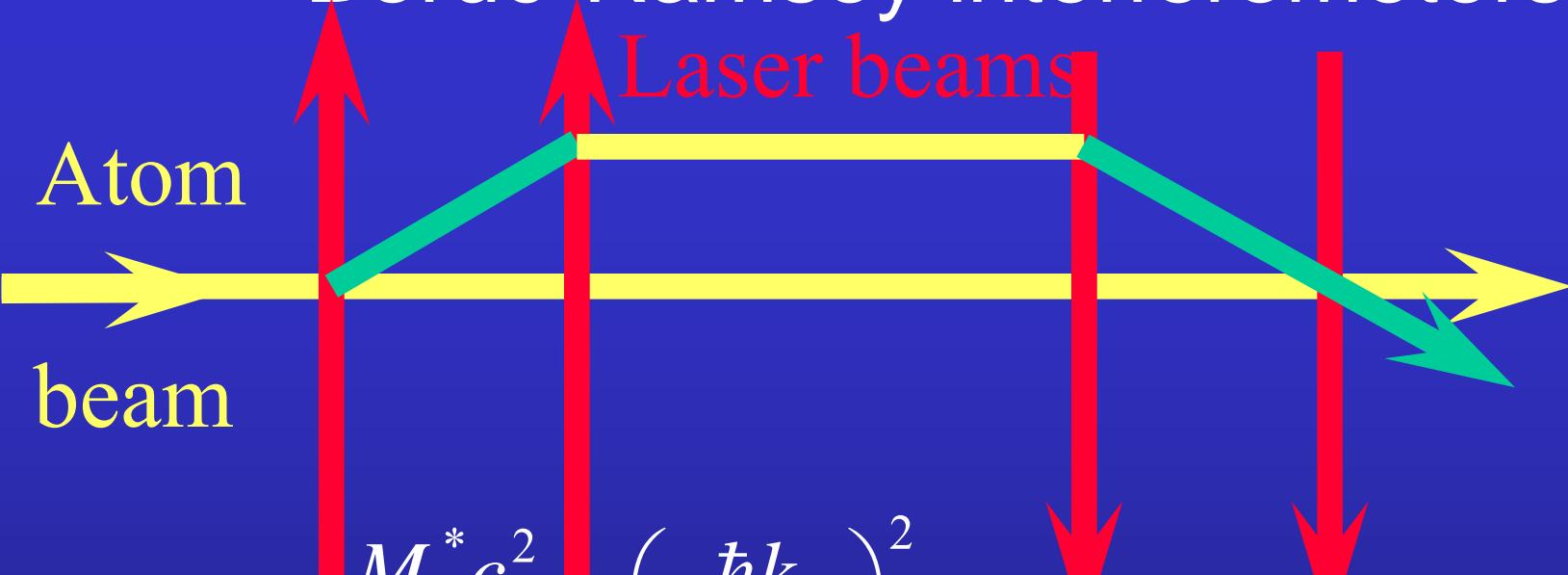


Figure 19. Atom wave gradiometer of Yale University. The two clouds of atoms share the same Raman beams, which generate two atom interferometers separated vertically by 1 m.

Bordé-Ramsey interferometers



$$\delta\phi = -\frac{M^* c^2}{\hbar} T \left(\frac{\hbar k}{M^* c} \right)^2 h \cos(\xi T + \phi) \operatorname{sinc}(\xi T)$$

