New ideas in SM predictions for the LHC

Christian Bauer The Search for New States and Forces of Nature GGI 10/27/2009

Work discussed here in collaboration with Jesse Thaler and Frank Tackmann



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Complicated interactions



Different distance scales



- Distance scale set by d~1/s
- Distance scale set by $d\sim 1/m_{\rm J}$
- Distance scale set by $d{\sim}1/\Lambda$
- Various scales involved

How can anything be calculated?

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Factorization

Separating physics from different distance scales crucial to allow theoretical predictions for any process

This has been major effort over past 30 years

Most trivial factorization separates partonic scattering from the parton distribution functions

Rigorous proofs exist only for the simplest processes (Drell-Yan ...)

Much work still invested in this direction

Seffective field theory (SCET) has allowed new angle at this problem



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Obtain the required ingredients

Data	Calculations	Models
Extract expressions from data	Calculate using expansion of QCD	Use more or less inspired models of QCD
PDF's Hadron decays	hard process QCD radiation	Hadronization Underlying event Multi scattering
Best	Good	Worst

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Using Data

Measured BR's used for unstable particles decay

PDF's extracted using simple well understood processes

 \otimes

10⁻¹

1

Χ

 $\sigma(p+e^{-} \rightarrow X+e^{-})$ f_a **MSTW 2008** xf(x,Q²) 1 $\tilde{\sigma}(x,Q^2)_{HERA} (F_2(x,Q^2)_{FixedTarget}) + c$ 2 ♦ H1 ♦ ZEUS NMC $Q^2 = 10 \text{ GeV}^2$ BCDMS x=0.0005 (c=0.35) SLAC x=0.00016 (c=0.4) x=0.0013 (c=0.3) q/10 1.5 0.8 =0.0032 (c=0.25) 0.6 0.4 =0.05 (c=0.1) C,C 0.2 0.5 10⁻³ 10⁻² **10⁻⁴** NNLO ---- NLO LO $10^2 \text{ Q}^2(\text{GeV}^2) 10^2$ LBNL, 9/24/09

Partonic cross section calculated to NNLO in perturbation theory

 $\sigma(q+e^-\rightarrow q+e^-)$



Models

Hadronization, underlying event and multiple proton scatterings are all non-perturbative effects that can not be calculated

Not universal enough to extract from data

Need more or less motivated models that capture some basic properties of QCD and adjust free parameters in the model to measured distributions

Very important to describe the physics, but need to remember that large uncertainties present



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Tuesday, October 27, 2009

Calculations



To obtain cross sections, need to integrate $|M|^2$ over phase space of final state particles. Calculation of the square of the amplitude. At higher orders need to perform loop integrals.

This is what I want to focus on in this talk



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Phase space integrals require high dimensional integrals Only for very low multiplicity can perform by hand Need numerical tools ⇒ Monte Carlo integration

Main bottleneck in calculations is the efficiency of the Monte Carlo integration

Singularities in $|M|^2$ make MC integration inefficient



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Very inefficient way of doing the integration

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Need to distribute points according to singularities

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Importance sampling (VEGAS) Start from uniform distribution of points Rapid changes in weights ⇒ divide into smaller areas



Better results, still inefficient for high multiplicity

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Multi-channel (MadEvent)



Singularities as propagators go on-shell



Integrate each term using singularity structure of $|M_i|^2$



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Use Parton Shower (GenEvA)

Parton showers generate events with given kinematics
 Kinematics correspond to point in phase space Φ_n
 If one knows probability P(Φ_n) to generate phase space, can assign weight

 $w(\Phi_n) = |M|^2(\Phi_n) / P(\Phi_n)$

Since parton shower reproduces singularity structure of QCD

$w(\Phi_n) = O(1)$



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What are the difficulties?

Some issues with a PS implementation

- Momentum is not conserved at leading order in parton shower
- Need to make somewhat ad-hoc corrections
- Can change the shape of distribution functions



Start with single branch
Add branching of left daughter
will change values of EL and ER
Add branching of right daughter
will change values of EL and ER
might not be allowed given tL



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Ringberg, 01/09/07

What are the difficulties?

Some issues with PS implementation

Momentum is not conserved at leading order in parton shower

Need to make somewhat ad-hoc corrections

Can change the shape of distribution functions

Changes distribution





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What are the difficulties?

Different parton shower histories can result in same point in phase space

Need to give different parton shower histories weights w_i , such that $\Sigma w_i = 1$

and

Have to make sure that these weights don't spoil efficiency



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Verification of Phase Space

Use LO tree level matrix elements and compare output to other event generator (MadEvent)

process	MadEvent	GenEvA	process	MadEvent	GenEvA
LO 3 (fb)	216.71 ± 0.21	216.77 ± 0.22	LO 5 (ab)	2542 ± 3	2543 ± 3
$u \overline{u} g$	86.62 ± 0.13	86.60 ± 0.18	$u\bar{u}ggg$	912 ± 2	912 ± 2
$d ar{d} g$	21.75 ± 0.07	21.55 ± 0.10	$d\bar{d}ggg$	227.5 ± 0.9	228.3 ± 0.8
$s \overline{s} g$	21.63 ± 0.06	21.73 ± 0.10	$u\bar{u}d\bar{d}g$	33.8 ± 0.2	34.3 ± 0.4
$c\bar{c}g$	86.71 ± 0.13	86.70 ± 0.18	$u\bar{u}u\bar{u}g$	25.6 ± 0.2	25.7 ± 0.3
LO 4 (fb)	36.44 ± 0.04	36.49 ± 0.04	LO 6 (ab)	67.9 ± 0.3	68.0 ± 0.2
$u \bar{u} g g$	14.00 ± 0.03	14.00 ± 0.02	$u\bar{u}gggg$	22.41 ± 0.09	22.29 ± 0.12
d ar d g g	3.504 ± 0.013	3.511 ± 0.011	$u \bar{u} u \bar{u} g g$	1.117 ± 0.006	1.14 ± 0.03
$u \bar{u} d \bar{d}$	0.175 ± 0.001	0.180 ± 0.003	$ u \bar{u} u \bar{u} u \bar{u} u \bar{u}$	0.005 ± 0.001^{-1}	0.005 ± 0.001
$u \bar{u} u \bar{u}$	0.132 ± 0.001	0.132 ± 0.002	$u\bar{u}d\bar{d}s\bar{s}$	0.019 ± 0.001^{-1}	0.020 ± 0.005

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Capri, 06/17/08

Verification of Phase Space Use LO tree level matrix elements and compare output to other event generator (MadEvent)





UAM, 06/24/08

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Comparison to MadEvent

	$\eta_{\rm eff}$	$T_{\rm eff} \ ({\rm msec})$	$T_{0.9} (\text{msec})$
GenEvA LO 3	0.789	0.57	0.62
GenEvA LO/LL inc. 3	0.965	0.47	< 0.47
MadEvent 3	0.982	2.6	< 2.6
MadEvent $u \bar{u} g$	0.994	3.0	< 3.0
GenEvA LO 4	0.525	1.7	2.2
GenEvA LO/LL inc. 4	0.713	1.3	1.5
MadEvent 4	0.809	11.1	11.4
MadEvent $u \bar{u} g g$	0.752	5.4	5.7
GenEvA LO 5	0.390	10.0	15
GenEvA LO/LL inc. 5	0.557	8.6	10.8
MadEvent 5	0.843	62	64
MadEvent $u \bar{u} g g g$	0.833	27	27
GenEvA LO 6	0.298	160	250
GenEvA LO/LL inc. 6	0.396	150	230
MadEvent 6	0.809	1900	2300
MadEvent $u ar{u} g g g g$	0.784	330	350

GenEvA very competitive with MadEvent, even without tuning shower for efficiency

MadEvent is worse for LL improved results, while GenEvA is better

GenEva beats MadEvent dramatically for LL improved results

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Calculation of $|M|^2$

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Peturbative calculations

Calculations proceed by calculating Feynman diagrams $\frac{\text{Example: pp} \rightarrow \text{W j}}{\text{Example: pp}} \rightarrow \text{W j}$





real

At NLO need:

virtual





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Logarithmic resummation

Perturbative expressions at higher order always contain logarithms of ratios of scales in the problem

Take example of $pp \rightarrow W+jets$ at small p_T

 $d\sigma/dp_T = N \times \left[1 + \alpha_s L^2 + \alpha_s L + \alpha_s + \alpha_s^2 L^2 + \alpha_s^2 L^2 + \alpha_s^2 L^2 + \alpha_s^2 L + \alpha_s^2 + \alpha_s^2 L$

Presence of logarithms can clearly spoil perturbative expansion for large ratio p_T/Q

Need to sum all large logarithmic terms



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Exclusive event generators

Most useful theoretical result for experimentalists is in the form of completely exclusive events

Need to generate events with many particles in the final state

Need to merge calculations with parton showers

Requiring such general results make calculations much more difficult

How precise can we make calculations in this form?



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Minimum accuracy required

NLO accuracy

1.Required to have any idea of scale dependence 2.NLO corrections can be factor of 2 over LO

LL resummation

1.Required to be able to merge with parton shower 2.Unresummed Logarithms can completely destroy convergence of perturbation theory

Combination of these two is the goal



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Implementing NLO Calculations



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Note that real emission has different final state multiplicity than virtual graph However, difference not detectable for kg°pq→0 due to finite resolution of any detector

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Analytical results Virtual contribution

$$\left(\frac{s\,\mathrm{d}\sigma}{\mathrm{d}t}\right)_{q\bar{q}} = e_{\mathrm{f}}^{2} \frac{K_{qq}}{s} \alpha_{\mathrm{s}}(M^{2}) T_{0}(Q^{2}, u, t) \left\{1 + \frac{\alpha_{\mathrm{s}}}{2\pi} \frac{\Gamma(1-\varepsilon)}{\Gamma(1-2\varepsilon)} \left(\frac{4\pi\mu^{2}}{Q^{2}}\right)^{\varepsilon} \times \left[-\frac{(2C_{\mathrm{F}}+N_{\mathrm{C}})}{\varepsilon^{2}} - \frac{1}{\varepsilon} \left(3C_{\mathrm{F}}-2C_{\mathrm{F}}\ln\left|\frac{s}{Q^{2}}\right| + \frac{11}{6}N_{\mathrm{C}} + N_{\mathrm{C}}\ln\frac{sQ^{2}}{ut} - \frac{2}{3}T_{\mathrm{R}}\right)\right] + \dots\right\}$$

Real contribution

$$\frac{s\,\mathrm{d}\sigma}{\mathrm{d}t\,\mathrm{d}u} = e_{\mathrm{f}}^{2} \frac{K_{\mathrm{qq}}}{s} \alpha_{\mathrm{s}} \left\{ T_{0}(Q^{2}, u, t)\,\delta(s_{2}) \left[\frac{\alpha_{\mathrm{s}}}{2\pi} \frac{\Gamma(1-\varepsilon)}{\Gamma(1-2\varepsilon)} \left(\frac{4\pi\mu^{2}}{Q^{2}} \right)^{\varepsilon} \right. \\ \left. \times \left[\frac{(2C_{\mathrm{F}}+N_{\mathrm{C}})}{\varepsilon^{2}} + \frac{1}{\varepsilon} \left(3C_{\mathrm{F}}+2C_{\mathrm{F}} \ln \frac{Q^{2}}{s} + N_{\mathrm{C}} \ln \frac{Q^{2}s}{ut} + \frac{11}{6}N_{\mathrm{C}} - \frac{2}{3}T_{\mathrm{R}} \right) \right] + \dots \right\}$$

Divergences cancel, finite pieces left over



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Infrared divergences

Problem is that IR divergences only cancel after phase space integration over real emission

Therefore, can not have fully exclusive events

Need to define exclusive cross section precisely



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Definition of exclusive x-section

$$\frac{\mathrm{d}\sigma_n^{\mathrm{excl}}}{\mathrm{d}\Phi_n} = \left[\frac{\mathrm{d}\sigma_n^{\mathrm{parton}}}{\mathrm{d}\Phi_n} + \sum_{m \ge n+1} \int \mathrm{d}\Phi_m \frac{\mathrm{d}\sigma_m^{\mathrm{parton}}}{\mathrm{d}\Phi_m} J_{\mathrm{excl}}(\Phi_m, \Phi_n; \mu)\right] R_n(\Phi_n)$$

"Jet" definition, ensuring that we have no more than n final state "particles"

Restriction, making sure that we don't have less than n "particles"

In the end, have exactly n final states



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Calculate @ NLO

Include both Born and Virtual term

$$\frac{\mathrm{d}\sigma_n^{\mathrm{parton}}}{\mathrm{d}\Phi_n} = B_n(\Phi_n) + V_n(\Phi_n) + \cdots$$

This gives

$$\frac{\mathrm{d}\sigma_n^{\mathrm{excl}}}{\mathrm{d}\Phi_n}\Big|_{\mathrm{NLO}} = \left[B_n(\Phi_n) + V_n(\Phi_n) + \int \mathrm{d}\Phi_{n+1} B_{n+1}(\Phi_{n+1}) J_{\mathrm{excl}}(\Phi_{n+1}, \Phi_n, \mu_n)\right] R_n(\Phi_n)$$
$$\pm \int \mathrm{d}\Phi_{n+1} S_{n+1}(\Phi_{n+1}) J_{\mathrm{excl}}(\Phi_{n+1}, \Phi_n, \mu_n) R_n(\Phi_n)$$

Define $V^{S} = V + \int S \qquad B^{S} = B - S$



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Calculate @ NLO

Include both Born and Virtual term

$$\frac{\mathrm{d}\sigma_n^{\mathrm{parton}}}{\mathrm{d}\Phi_n} = B_n(\Phi_n) + V_n(\Phi_n) + \cdots$$

This gives

$$\frac{\mathrm{d}\sigma_n^{\mathrm{excl}}}{\mathrm{d}\Phi_n}\Big|_{\mathrm{NLO}} = \left[B_n(\Phi_n) + V_n^S(\Phi_n) + \int \mathrm{d}\Phi_{n+1} B_{n+1}^S(\Phi_{n+1}) J_{\mathrm{excl}}(\Phi_{n+1}, \Phi_n, \mu_n)\right] R_n(\Phi_n)$$

B^s has no more singularities, and J_{excl} only integrates over very small region

$$\Rightarrow \int B^{S} J_{excl} \rightarrow C$$

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Calculate @ NLO

Include both Born and Virtual term

$$\frac{\mathrm{d}\sigma_n^{\mathrm{parton}}}{\mathrm{d}\Phi_n} = B_n(\Phi_n) + V_n(\Phi_n) + \cdots$$

This gives

$$\frac{\mathrm{d}\sigma_n^{\mathrm{excl}}}{\mathrm{d}\Phi_n}\Big|_{\mathrm{NLO}} = \left[B_n(\Phi_n) + V_n^S(\Phi_n,\mu)\right]R_n(\Phi_n) + \mathcal{O}(\mu)$$

Can extract this from programs such as Blackhat etc Note that σ_n^{excl} depends on scale μ



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Resummation of Logs



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Logarithmic resummation

Logarithms related to IR divergences in theory



Divergences from loop integrations $\int dk_1 (k_1-Q)^{-1-\epsilon} = 1/\epsilon + ...$



Divergences from phase space integrations $\int_{0}^{1} dk_{g} (k_{g})^{-1-\epsilon} = -1/\epsilon + ...$

Restrictions on phase space give rise to logarithmic remainders $0 < k_g < \mu \implies -1/\epsilon + \log(\mu)$

 $\sigma_{V} + \sigma_{R} = \log(\mu) + \dots$



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Relating IR and UV divergences Effective theories are defined to reproduce the IR physics of underlying theory If EFT only contains no intrinsic mass scales, it only depends on Λ_{IR} and Λ_{UV} In pure dim-reg $\Lambda_{IR} \rightarrow 0$ and $\Lambda_{UV} \rightarrow \infty$ No scale in problem \Rightarrow result in EFT is 0 $\Rightarrow Log(\Lambda_{IR}) = - Log(\Lambda_{UV})$ IR dependence of full theory can be extracted from UV dependence of EFT



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Sum logs using RG Equations

Derive RG Equation, by taking $\mu d/d\mu$

$$\mu \frac{\mathrm{d}}{\mathrm{d}\mu} \frac{\mathrm{d}\sigma_n^{\mathrm{LL}}(\mu)}{\mathrm{d}\Phi_n} = \gamma_n(\mu) \frac{\mathrm{d}\sigma_n^{\mathrm{LL}}(\mu)}{\mathrm{d}\Phi_n}$$

Solve to find

$$\frac{\mathrm{d}\sigma_n^{\mathrm{LL}}(\mu)}{\mathrm{d}\Phi_n} = \frac{\mathrm{d}\sigma_n^{\mathrm{LL}}(\mu_0)}{\mathrm{d}\Phi_n} \,\Delta_n(\mu_0,\mu)$$

Δ_n = Sudakov factor

Resums all logarithms of μ

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Resumming kinematic logs

Many scales in problem, such as different values of p_T

For universal result, want to resum ratios of all scales

For very particular choice of J_{excl}, possible to write the result as

 $\sigma_n(\mu_1,\mu_2,...) = \sigma_n(Q) \Delta_n(Q,\mu_1) \Delta_n(Q,\mu_1) \dots$

Possible to sum all logarithms in the exclusive cross sections



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Combining NLO and LL results

Schematically, write

$\sigma_n^{excl} = \sigma_n^{LL} + M_n$

Determine matching coefficient M_n by requiring correct NLO expression when expanded

$$M_n^{i_n,(0)}(\Phi_n) = S_n^{i_n}(\Phi_n) \left(\frac{B_n(\Phi_n)}{S_n(\Phi_n)} - 1\right),$$
$$M_n^{i_n,(1)}(\Phi_n) = S_n^{i_n} \left(\frac{V_n^S(\Phi_n, \mu_n)}{S_n(\Phi_n)} - \frac{V_{n-1}^S(\Phi_{n-1}^{i_n}, t_n^{i_n})}{B_{n-1}(\Phi_{n-1}^{i_n})} - \Delta_n^{(1)}(t_n^{i_n}, \mu_n)\right)$$

Everything can be calculated analytically



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Status of the work? Have all the analytical results worked out in detail for ete-Finished implementation in C++ code Working on the implementation with Pythia parton shower Working on the extension to allow for hadron colliders

Will be interesting to see how well it works!



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Conclusions

SM predictions need to include many effects to give adequate predictions for experimental observables

In order to compare with data, need exclusive events, distributed in phase space according to SM predictions

New ideas in phase space generation can remove major bottleneck in efficiency of calculations

Combination of NLO calculations with generic LL resummation is becoming reality

Look forward to testing these theoretical ideas against real data from the LHC



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