



DM Signatures generated by anomalies in hidden sectors



Yann Mambrini

Laboratoire de Physique Théorique Orsay, Université Paris XI

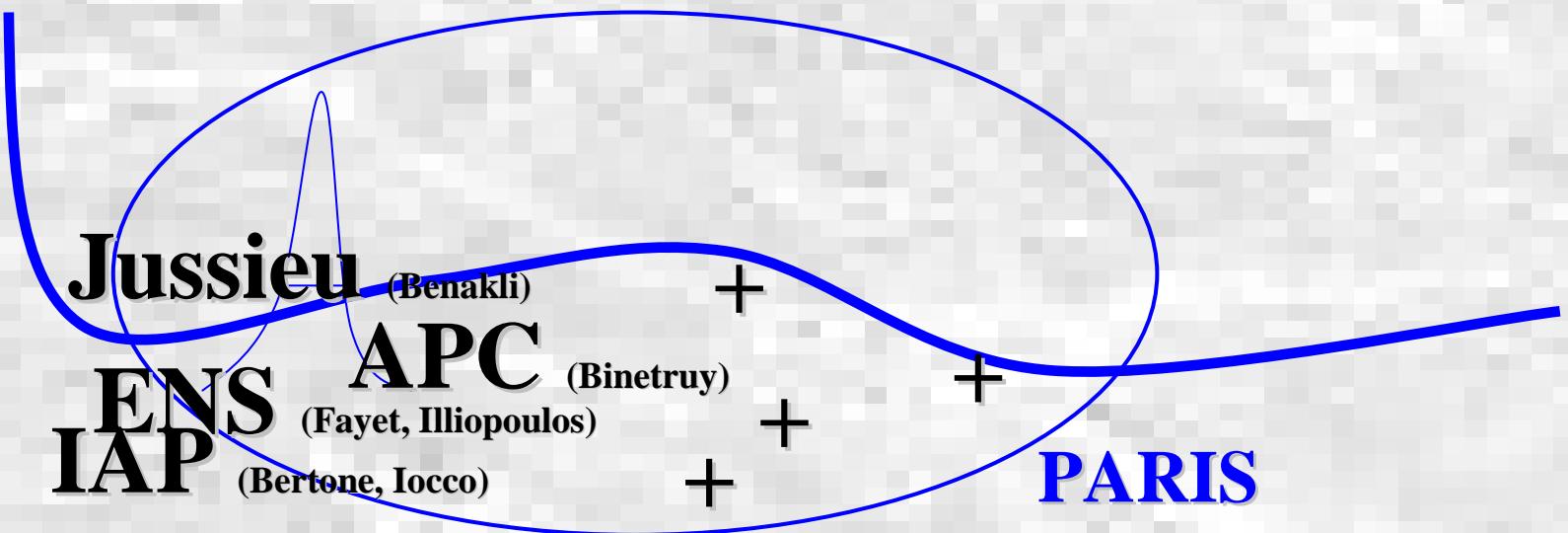
E. Dudas, S. Pokorski, A. Romagnoni



GGI, Florence,, May 20th 2010



Where is Orsay?



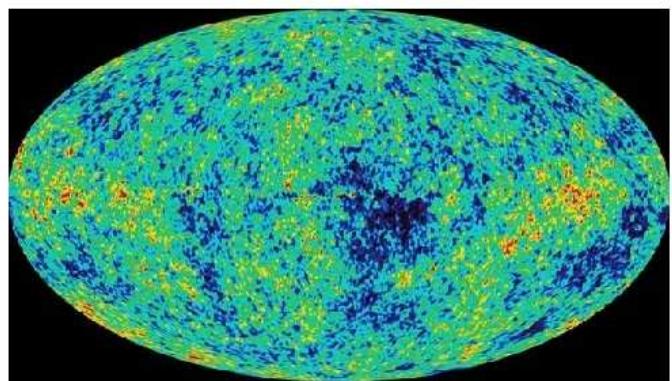
Polytechnique +
Saclay^(Antoniadis,) (Cirelli, Zaharijas) + +
Orsay

Overview

- I) The DM puzzle
- II) Gamma ray lines signals
- III) Importance of Anomalies in Particle Physics
- IV) Anomalies cancelation mechanisms
- V) Gamma-ray lines generated by anomalies
- VI) FERMI analysis
- VII) Conclusions

Dark Matter Evidences

CMB (WMAP)



Galactic Scale

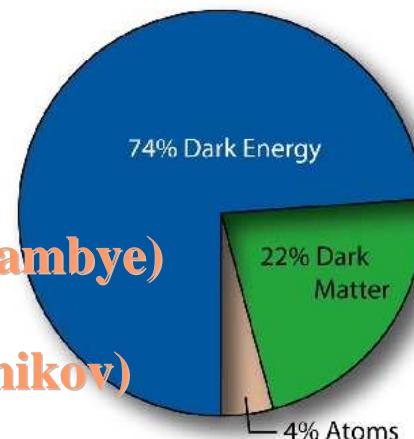


SUSY : neutralino, gravitino.. (N.
Fornengo, P. Ullio, L. Covi)
KK modes (Extra Dim.) (Servant, Tait)

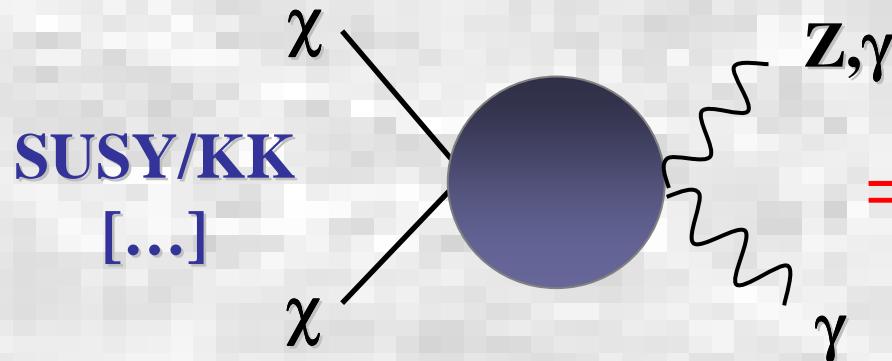
Extra U(1) boson (Arkani-Ahmed, T. Hambye)

Sterile right handed neutrino (Shaposhnikov)

Weak scale scalar (Gustafsson, Tytgat)

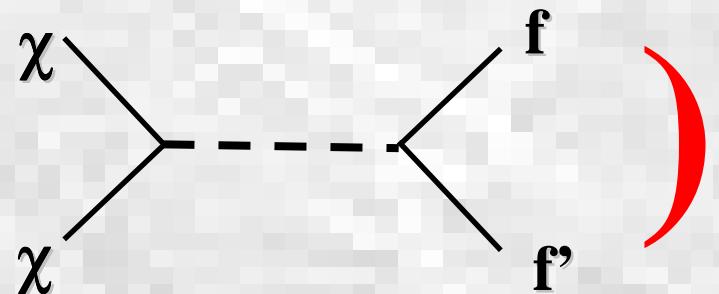


γ ray lines (annihilation)



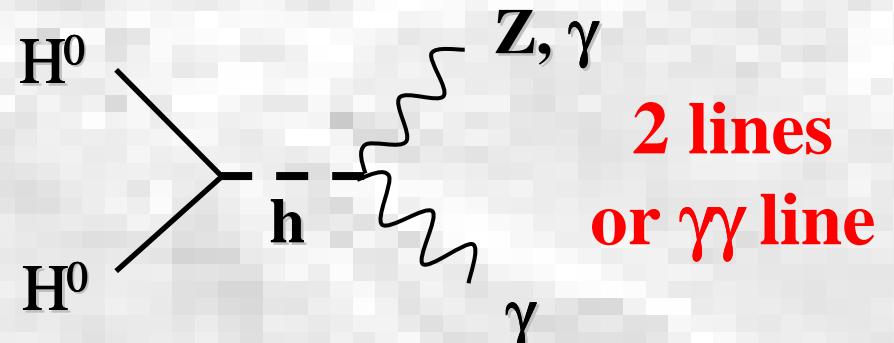
SUSY/KK
[...]

$$= 10^{-2} \left($$

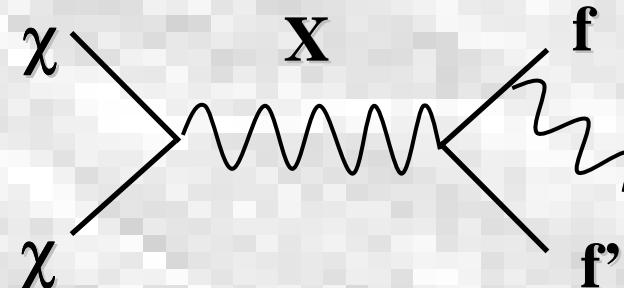


Inert Higgs Doublet
[Gustafsson, Lundstrom,
Bergstrom, Edsjo 07]

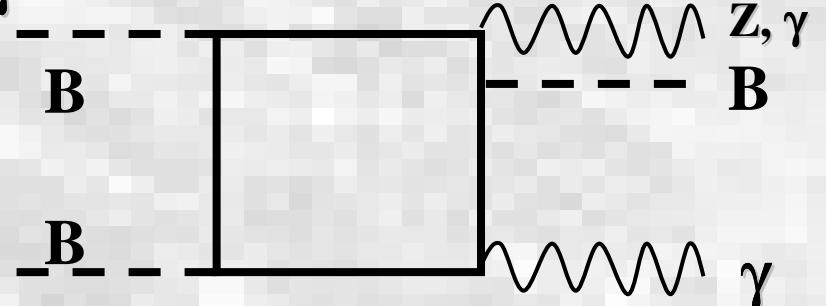
If $m_{H^0} < M_W$



Visible X
[Baek & Ko, 08]
[Bergstrom,
Bringmann, Edsjo 08]

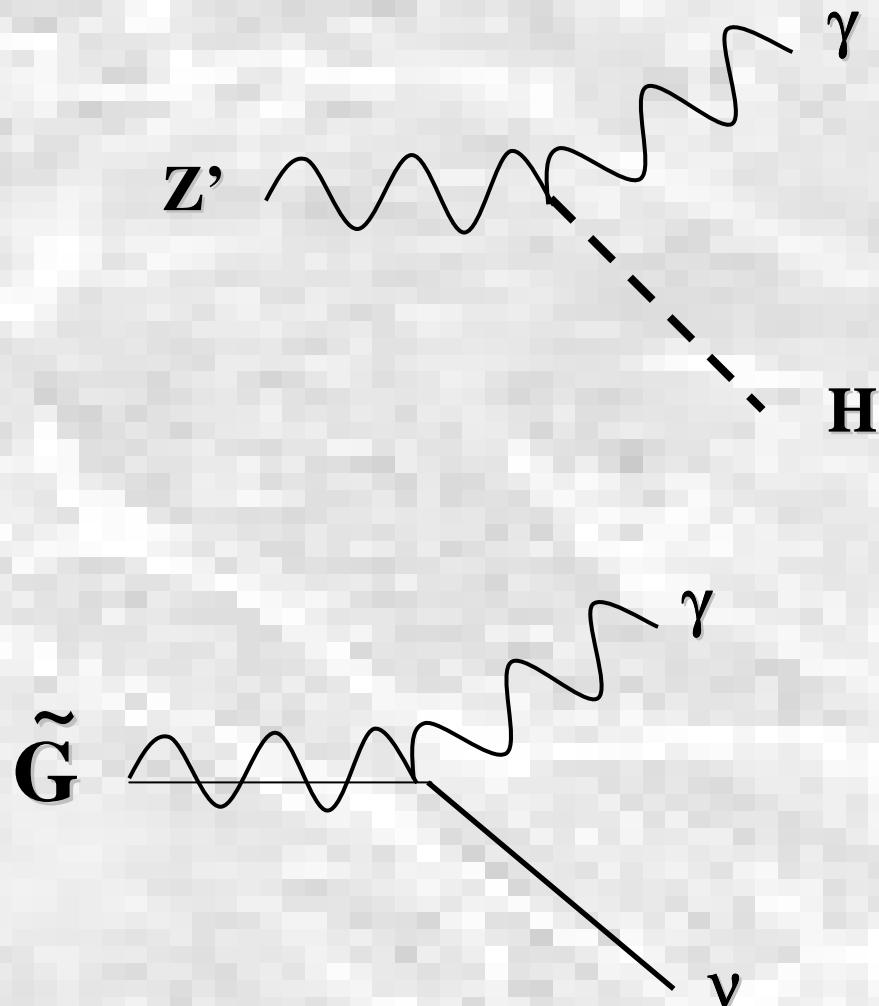


3 visible lines



Chiral Square
[Bertone et al. 09]

γ ray lines (Decaying)



T. Hambye 09

K.Y. Choi, C. Munoz 08
G. Bertone, W. Buchmuller,
L. Covi, A. Ibarra 07

Anomalies in particle physics

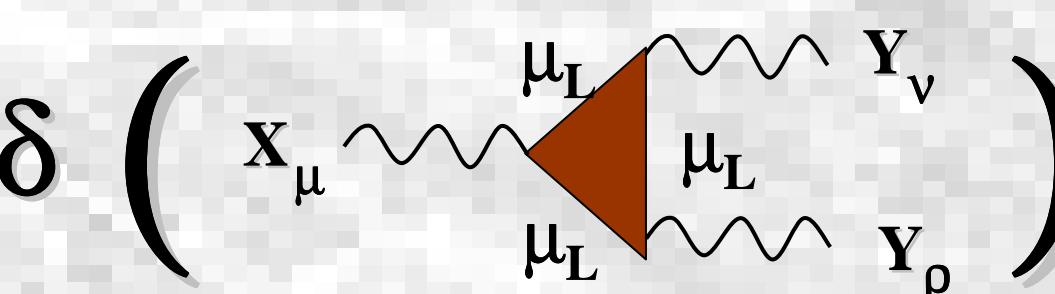
$$\cancel{L_{Sym} + \delta L_{quant}}$$

Anomalies : quantum corrections destroy the symmetries of the classical theory

If the symmetry is local, the gauge anomaly causes the **loss of unitarity**, which is to say the **consistency** of the theory

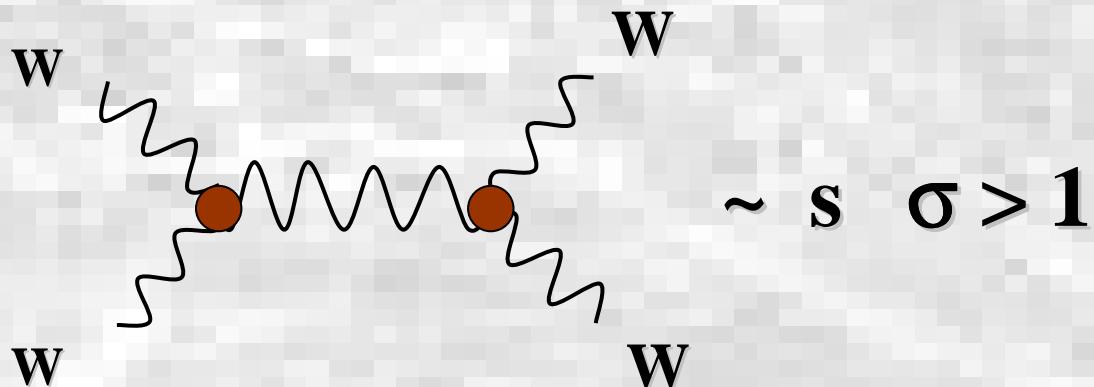
A mechanism of anomaly cancellation should be imposed to the model

Example : leptophylic DM

$$x_\mu \sim \begin{array}{c} \mu^+ \\ \diagdown \\ \mu^- \end{array} \quad \delta \left(x_\mu \sim \begin{array}{c} \mu_L \\ \diagdown \\ \mu_L \end{array} \right) = 0$$


INDEPENDENT OF THE MASSES OF THE FERMIONS
RUNNING IN THE LOOPS (decoupling theorem)

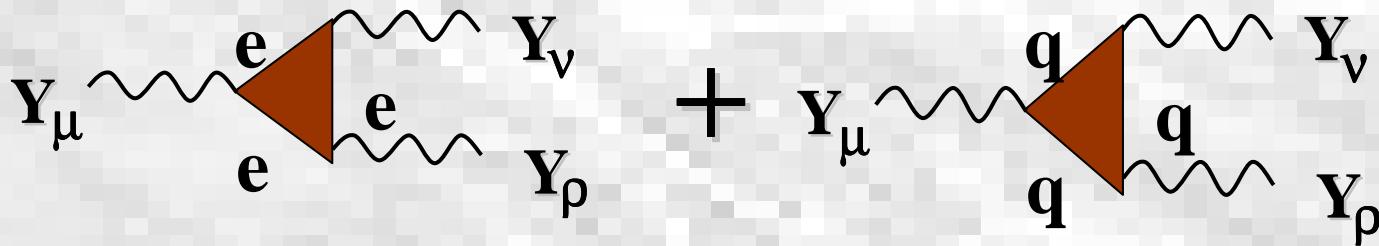
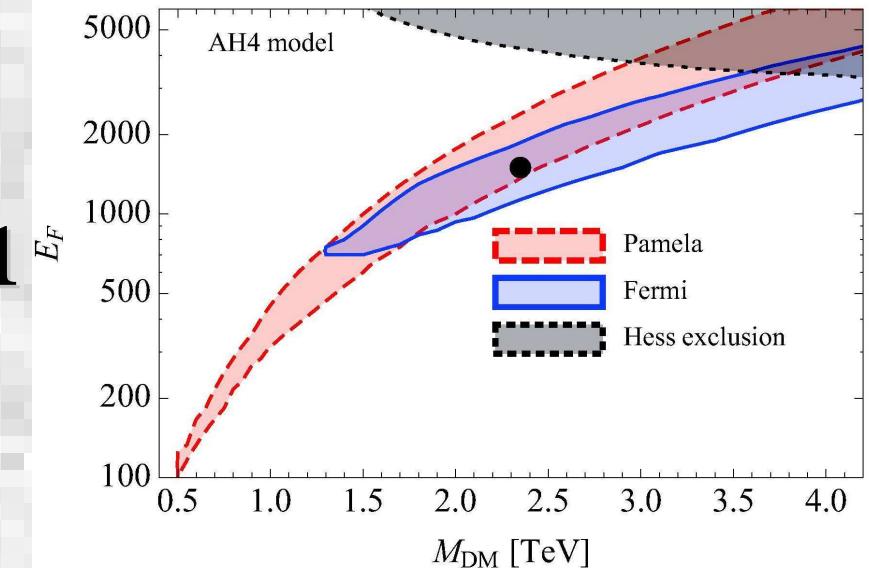
Solving the problem..



$\sim S$

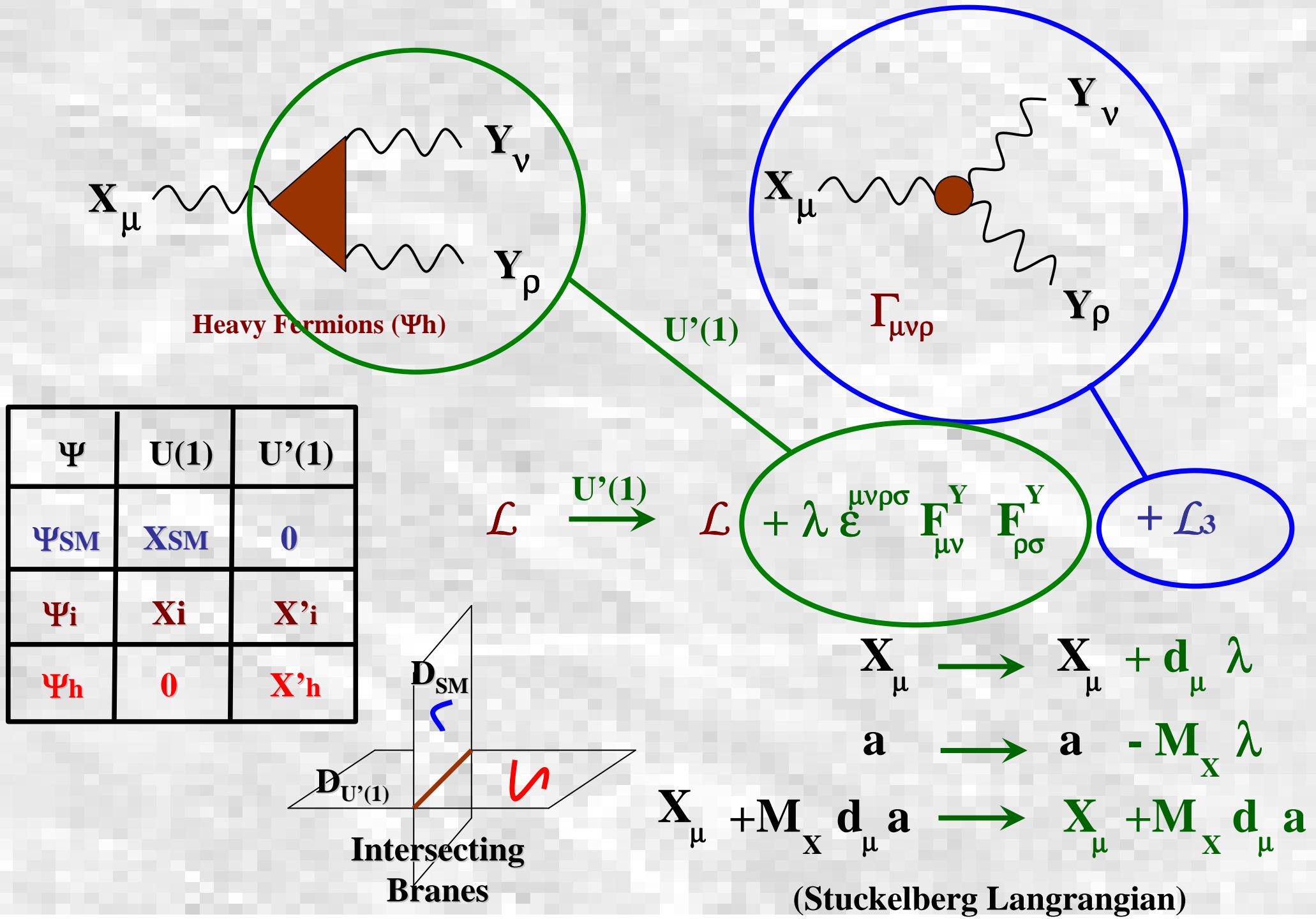
$\sigma > 1$

Lost of unitarity



$\alpha(Q_e - 3 Q_q)$
Miracle !!

Explicit construction



Anomalie cancellation : Green-Schwarz mechanism

Anastopoulos, Bianchi, Dudas, Kiritsis 06

Anastopoulos, Fucito, Lionetto, Pradisi, Racioppi, Stanev 08

$$\mathcal{L}_{inv} = F^{Y\mu\nu} F^Y_{\mu\nu} - (d_\mu a - M_X X_\mu)^2 - i \bar{\Psi}_h \gamma^\mu D_\mu \Psi_h$$

$$\mathcal{L}_{var} = \textcolor{blue}{B a \epsilon^{\mu\nu\rho\sigma} F^Y_{\mu\nu} F^Y_{\rho\sigma}} + \textcolor{red}{C \epsilon^{\mu\nu\rho\sigma} X_\mu Y_\nu F^Y_{\rho\sigma}}$$

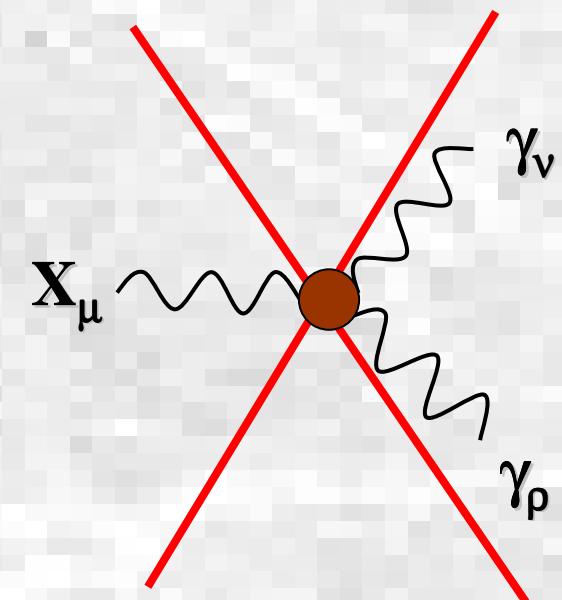
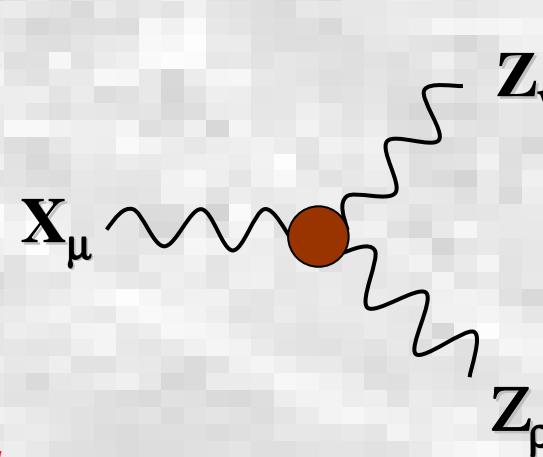
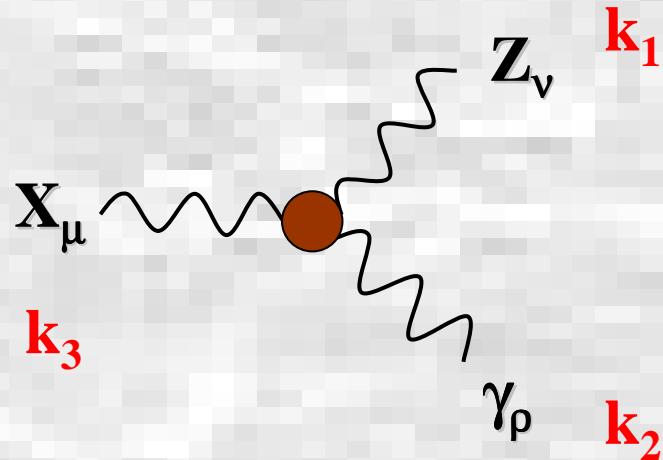
Peccei-Quinn terms

Chern-Simons terms

$$\delta \mathcal{L}_{var} = - \delta \left(X_\mu \begin{array}{c} \diagdown \\ \diagup \end{array} Y_\nu \right) Y_\rho$$

Heavy Fermions (Ψ_h)

Effective couplings : $\mathcal{L}_{eff} = \mathcal{L}_{loops} + \mathcal{L}_{var}$



$$\Gamma_{\mu\nu\rho}^{\alpha} = t^{\alpha} \{ A_1 \epsilon^{\mu\nu\rho\sigma} k_{2\sigma} - A_2 \epsilon^{\mu\nu\rho\sigma} k_{1\sigma} \\ + B_1 k_{2\nu} \epsilon^{\mu\rho\sigma\tau} k_{2\sigma} k_{1\tau} + B_2 k_{1\nu} \epsilon^{\mu\rho\sigma\tau} k_{2\sigma} k_{1\tau} + B_3 k_{2\rho} \epsilon^{\mu\nu\sigma\tau} k_{2\sigma} k_{1\tau} + B_4 \epsilon^{\mu\nu\sigma\tau} k_{2\sigma} k_{1\tau} \\ + C k_{3\mu} / k_3^2 \epsilon^{\nu\rho\sigma\tau} k_{2\sigma} k_{1\tau} + D \epsilon^{\mu\nu\rho\sigma} (k_{2\sigma} - k_{1\sigma}) \}$$

Peccei-Quinn

Chern-Simons

[Kumar, Rajaraman,
Wells 08]
[YM 09]

$\delta \mathcal{L}_{eff} = 0$ 3 Ward identities + $(k_1; k_2)$ symmetries

-> the vertex can be express as function of $|B_2 - B_1| = 1/\Lambda_X^2$

With $B_1, B_2 = \underline{\text{computable loops integrals}}$

Cc : only 3 parameters : $\Lambda_X [S]$; $M_X [g_X]$; $M_\chi [Y_{heavy}]$

Interpretation as higher dimensional operators

Antoniadis, Boyarski, Ruchavki, Wells 09
Dudas, YM, Pokorski, Romagnoni 09

$$\mathcal{L}_1 = \frac{1/M^2}{\text{blue circle}} * \{ b \text{ Tr}[F' F^Y F^Y] + c \epsilon^{\mu\nu\rho\sigma} (\mathcal{D}_\mu a) (D_\nu H)^+ F^Y_{\rho\sigma} H \},$$

with

$$\mathcal{D}_\mu a = d_\mu a - g' A'_\mu ; \quad D_\nu = d_\nu - i g Y_\nu - i g' A'_\nu$$

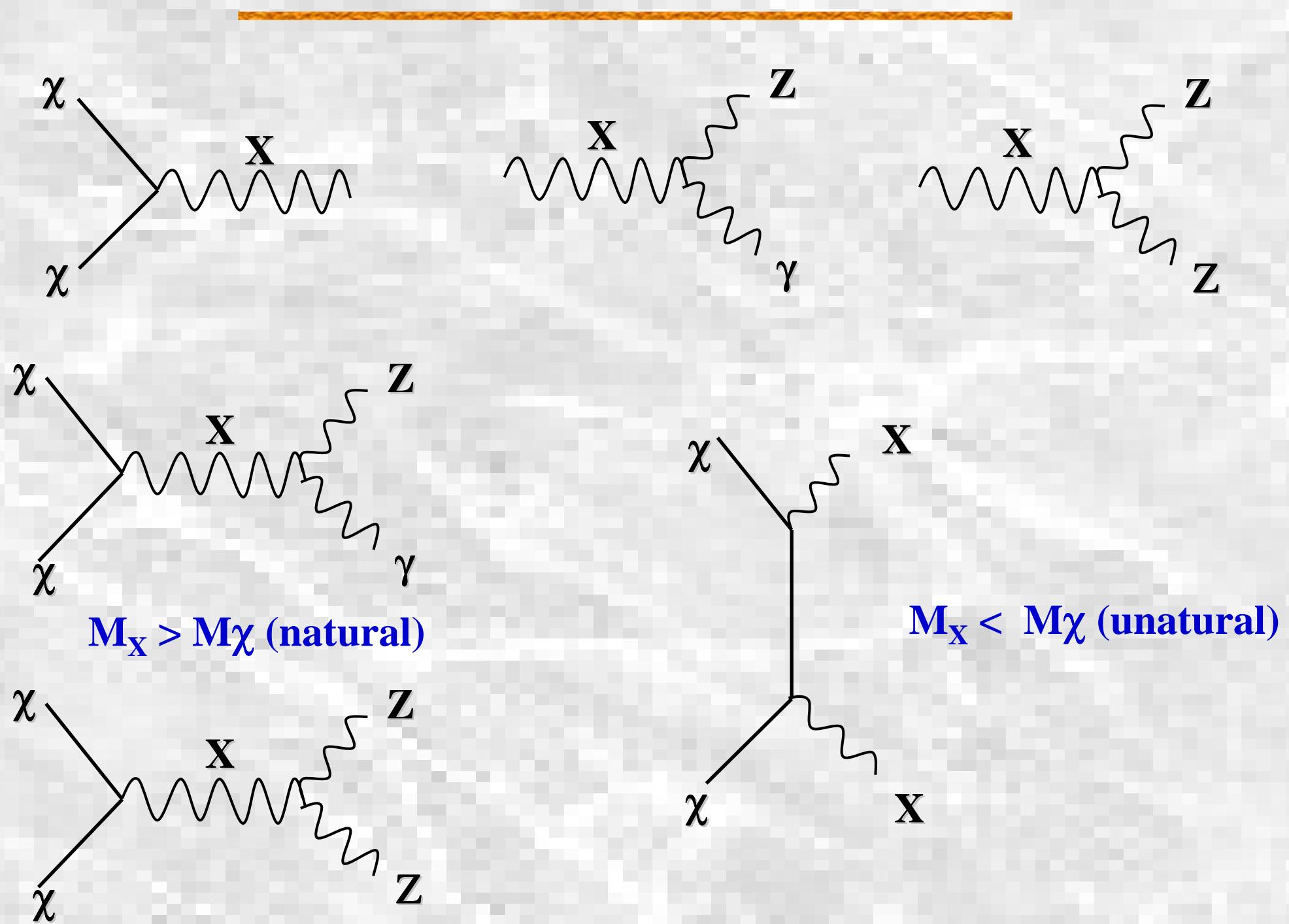
→ Masses suppression coming from the fermions
which decouple after **U'(1)** breaking

Equivalent to the D'Hoker-Farhi term

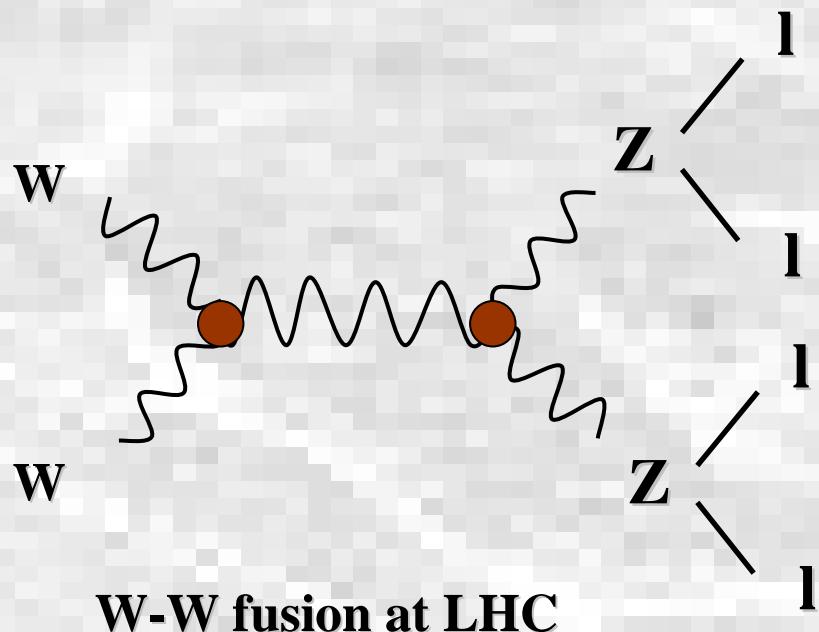
$$\{ 1/(H^+ H) \epsilon^{\mu\nu\rho\sigma} (\mathcal{D}_\mu a) (D_\nu H)^+ F^Y_{\rho\sigma} H \} \text{ for SM}$$

Remark :if two Z' are present, we can build an unsuppressed operator,
 $\epsilon^{\mu\nu\rho\sigma} (\mathcal{D}_\mu a_1) (\mathcal{D}_\nu a_2) F^Y_{\rho\sigma}$

Dark matter: Annihilation channels

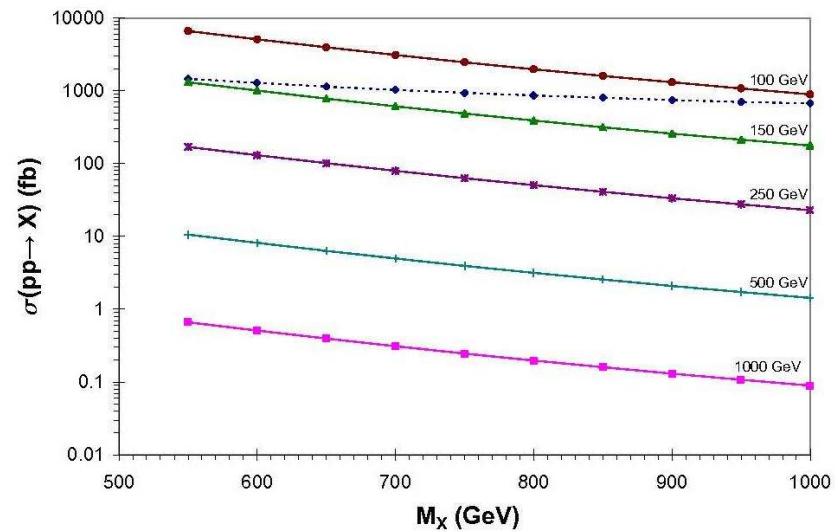


LHC production

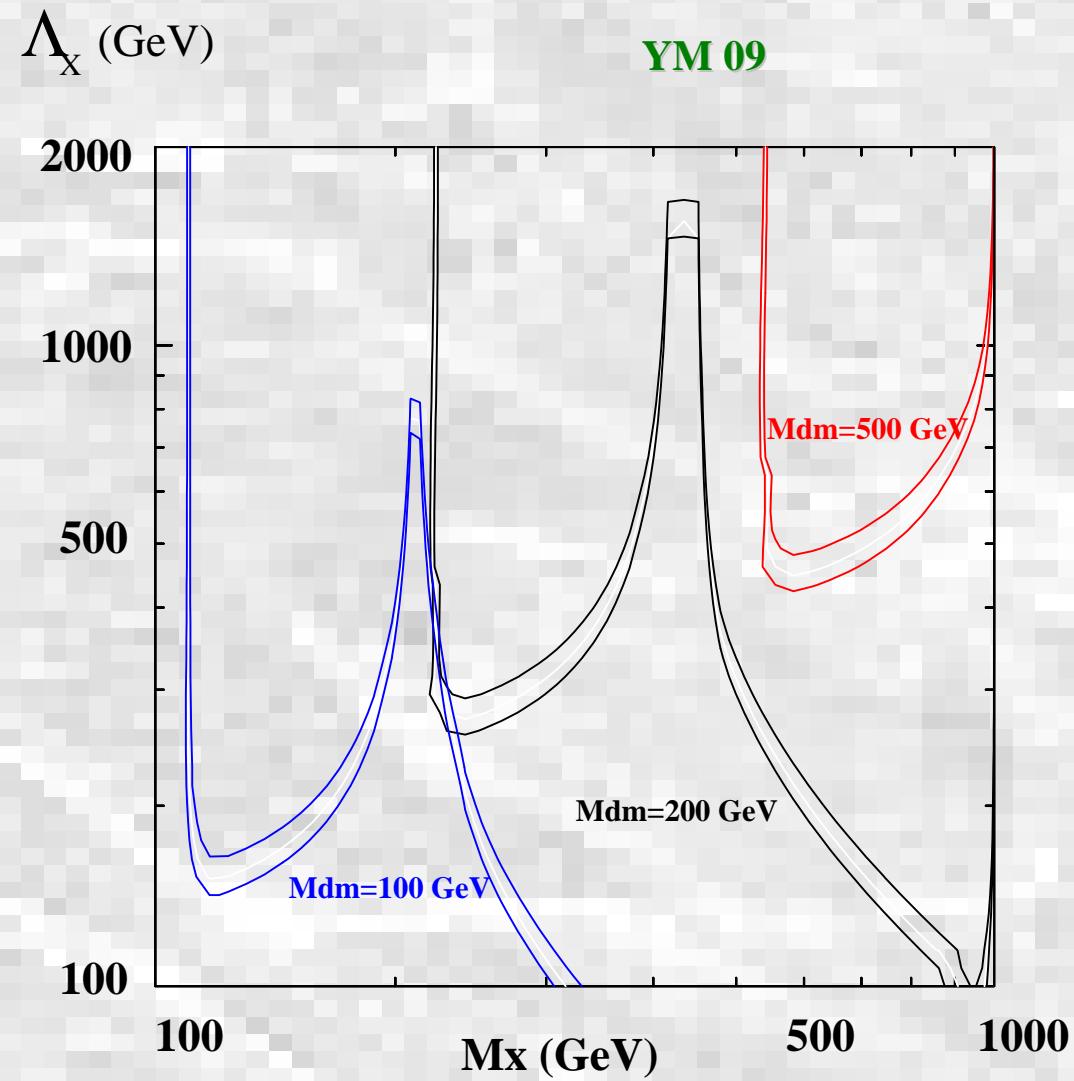


W-W fusion at LHC

Kumar, Wells 08
Antoniadis Boyarski Ruchavski, Wells 09



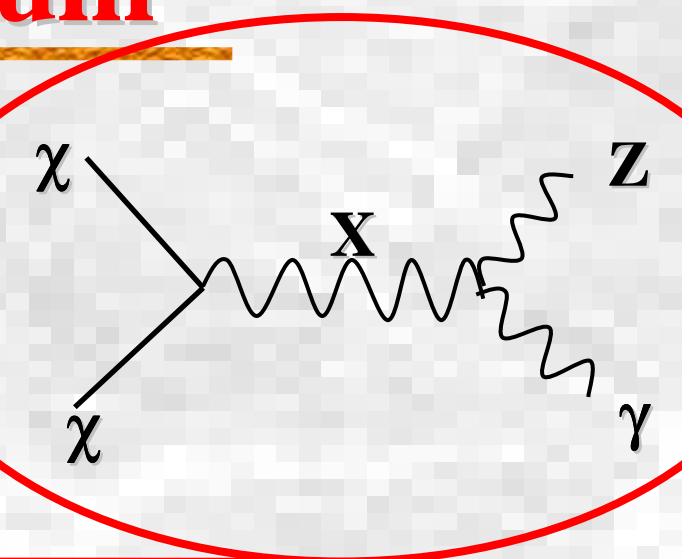
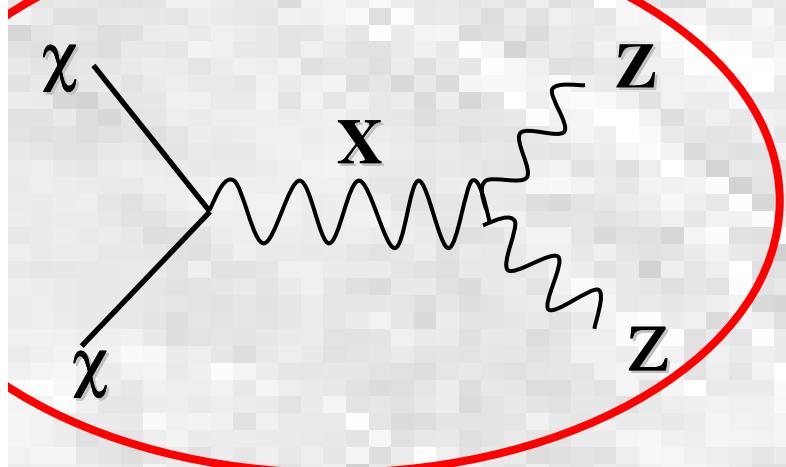
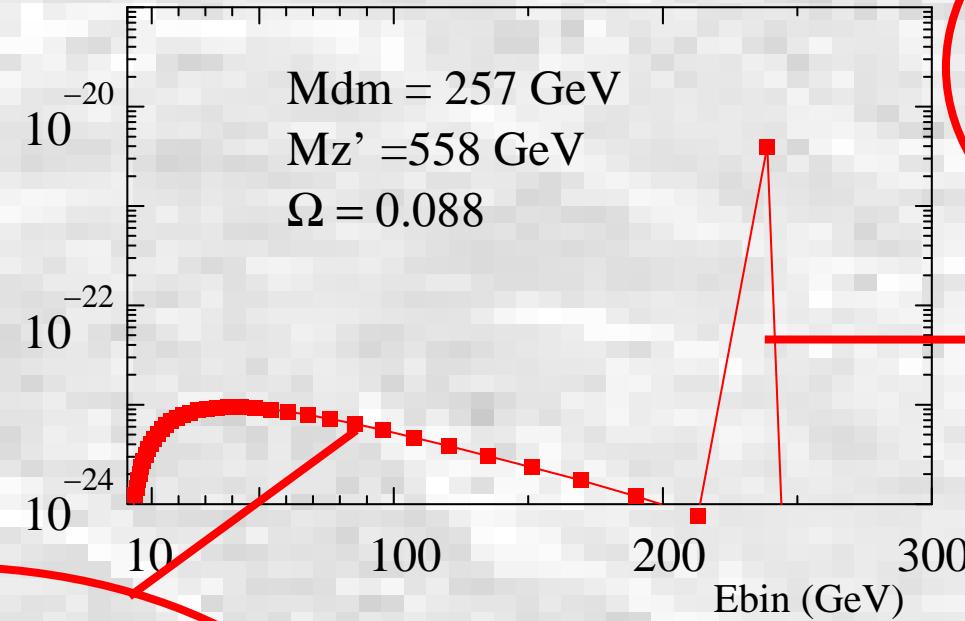
The relic density



$$1/\Lambda_X^2 = |B_2 - B_1| = g_h * g^2 / (8\pi^2) * \text{Tr}[X' X^2 / M_{heavy}^2] * \text{Integral}$$

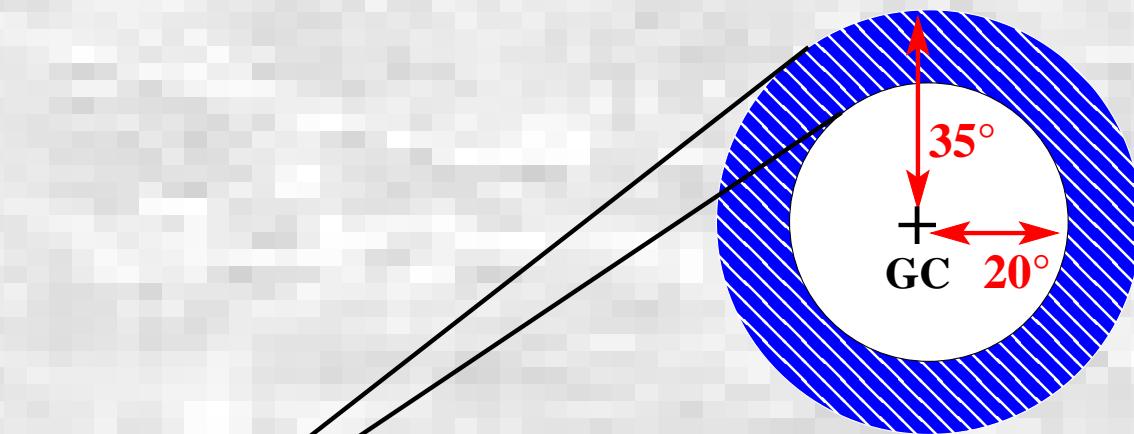
Indirect detection: examples of spectrum

$E^2 * dN/dE$



$$E_\gamma = M_\chi [1 - (M_Z/2M_\chi)^2]$$

Indirect detection astro-parameters



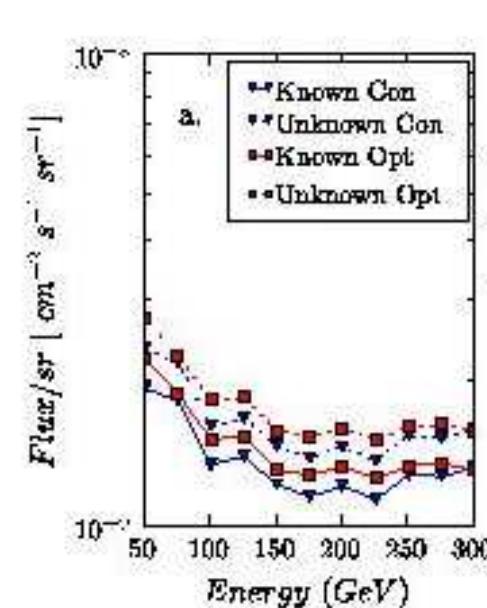
Galactic Centered Annulus
(Stoehr et al 2003, GLAST col. 2008)

Signal to noise ratio
12 time greater than
GC

Independant of the Galactic profile

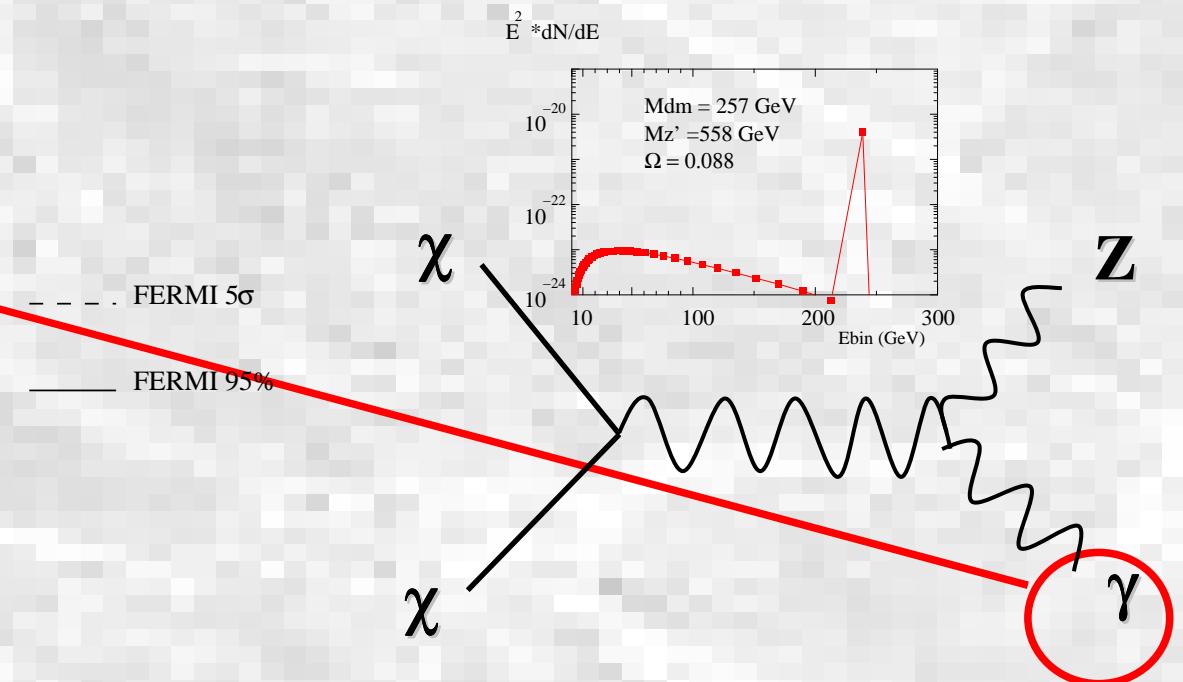
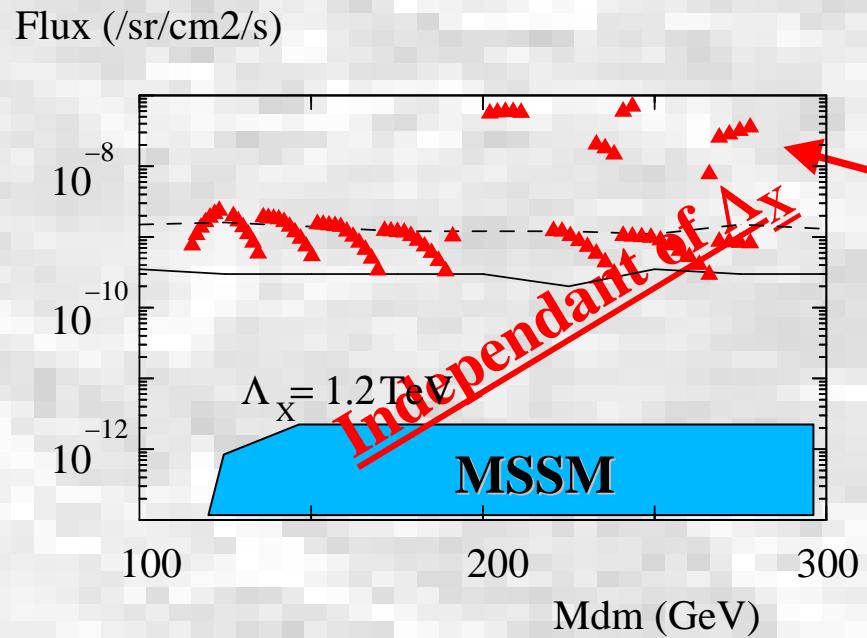
$$J \Delta\Omega \sim 10$$

Galprop conventional model
for the background



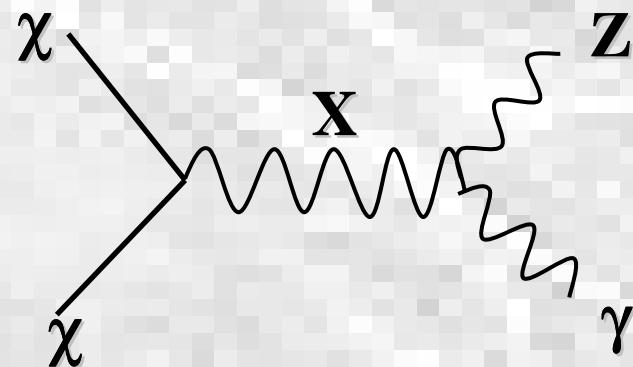
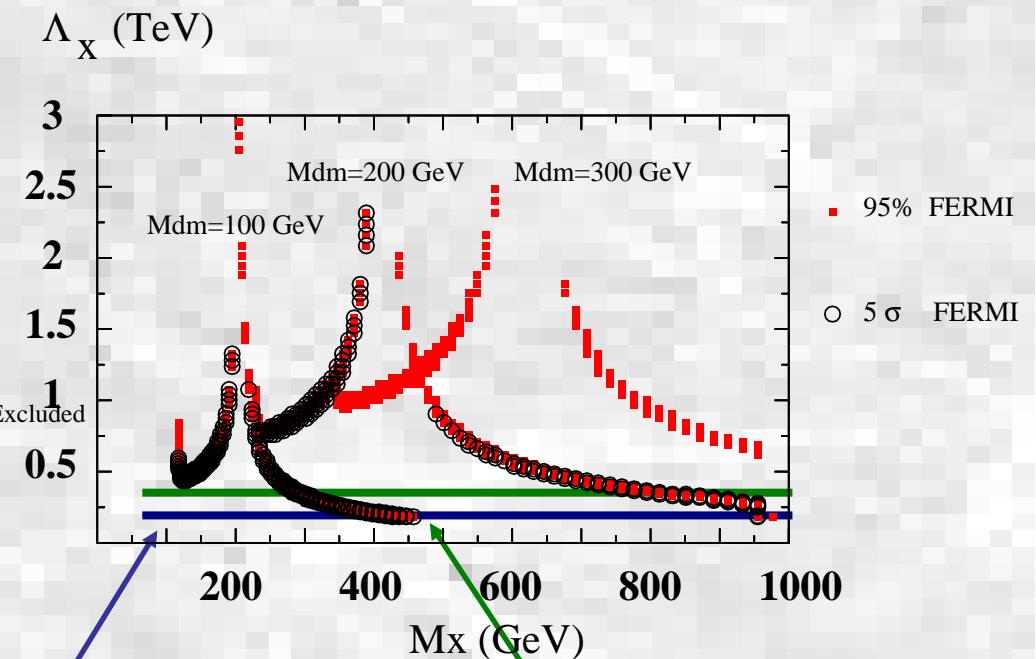
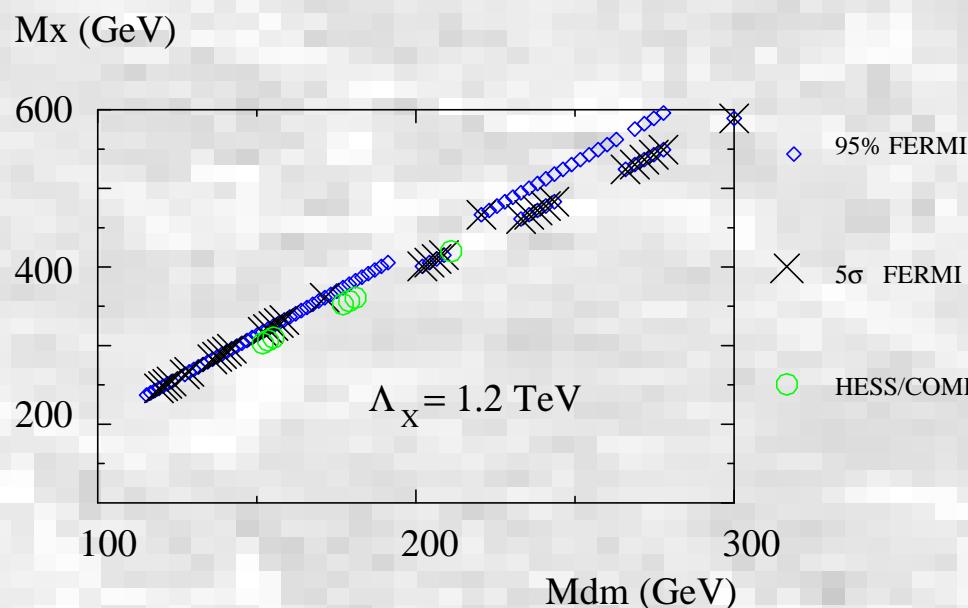
5 years of data, signals at 5σ and
95% CL
[FERMI estimates, Morselli et al. 08]

Observability



No excess with current constraints (EGRET, HESS.. [Jacques, Bell 08])

Consequences on M_x and M_{χ}

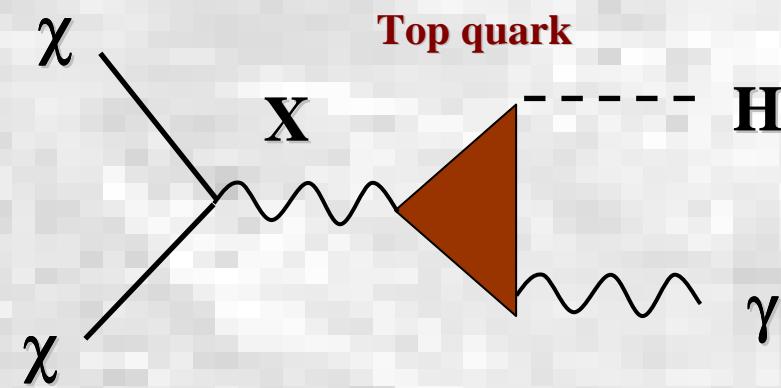


LHC (Wells 08, Antoniadis 09)

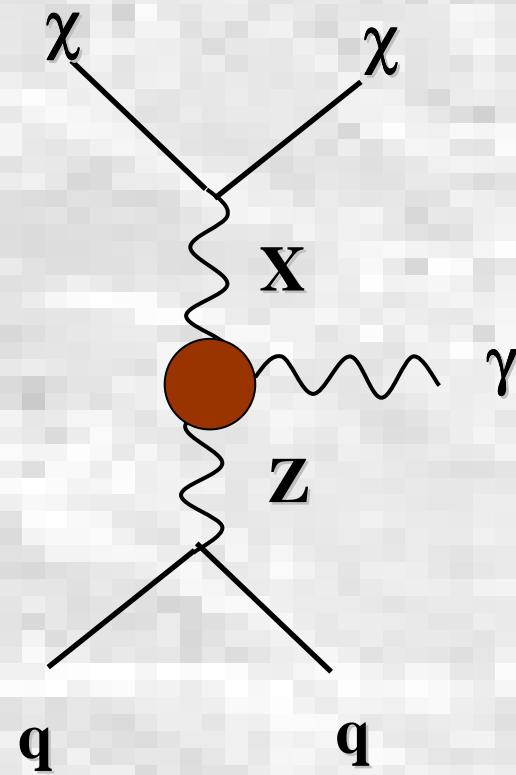
FERMI 10

An extension : Higgs in Space

Jackson, Servant, Shaughnessz, Tait, Taoso 09



Direct detection



Mini-conclusions

An (In)visible Z' can be quite visible

Indirect detection could be THE ONLY WAY to observe it

**1 γ ray line is a smoking gun signal distinguishing it clearly
from other constructions**

Possibility to test up to 5TeV BSM scale at FERMI telescope