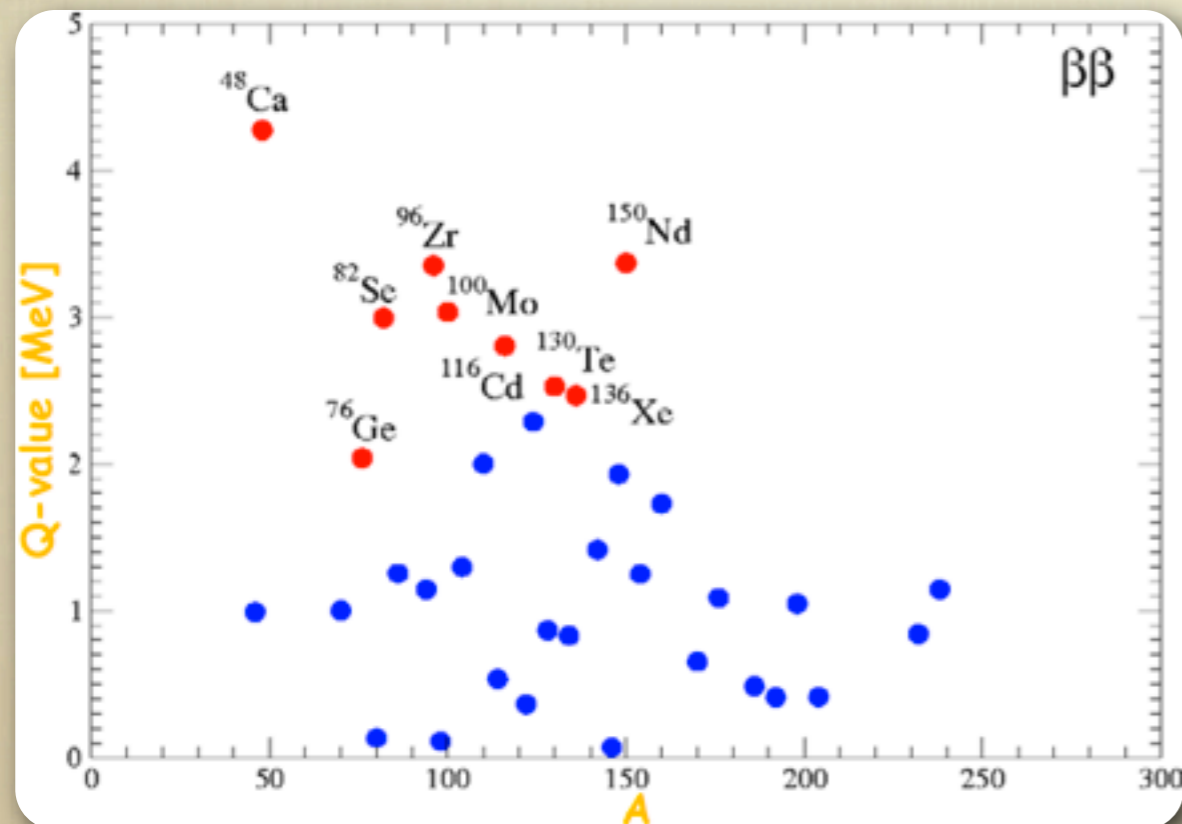
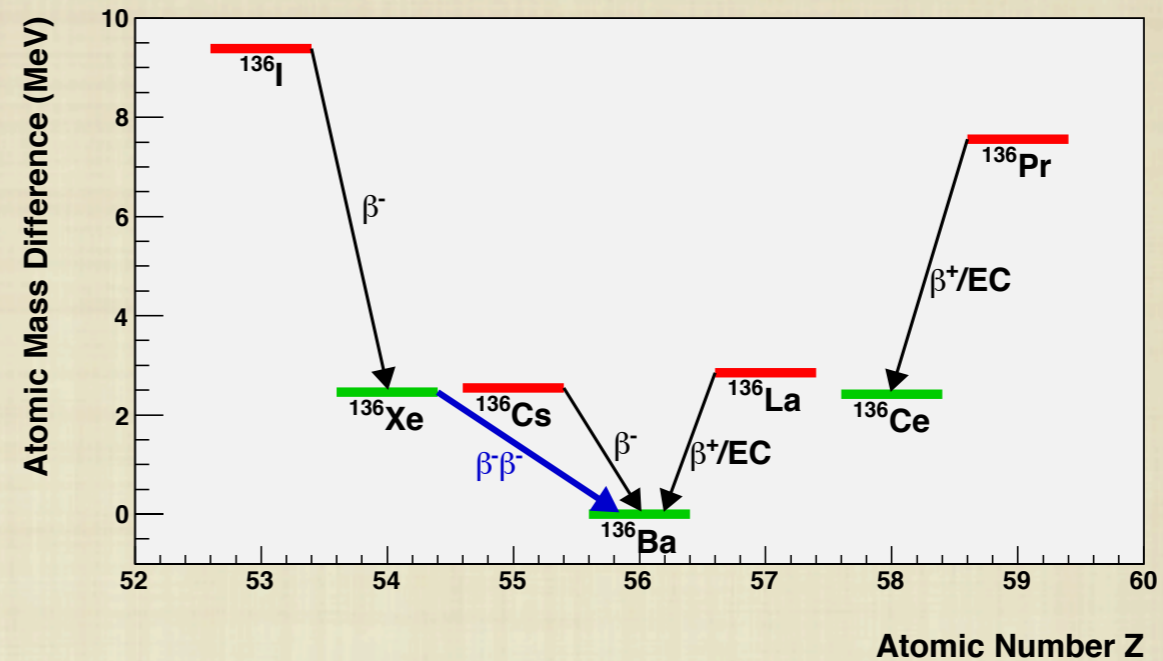
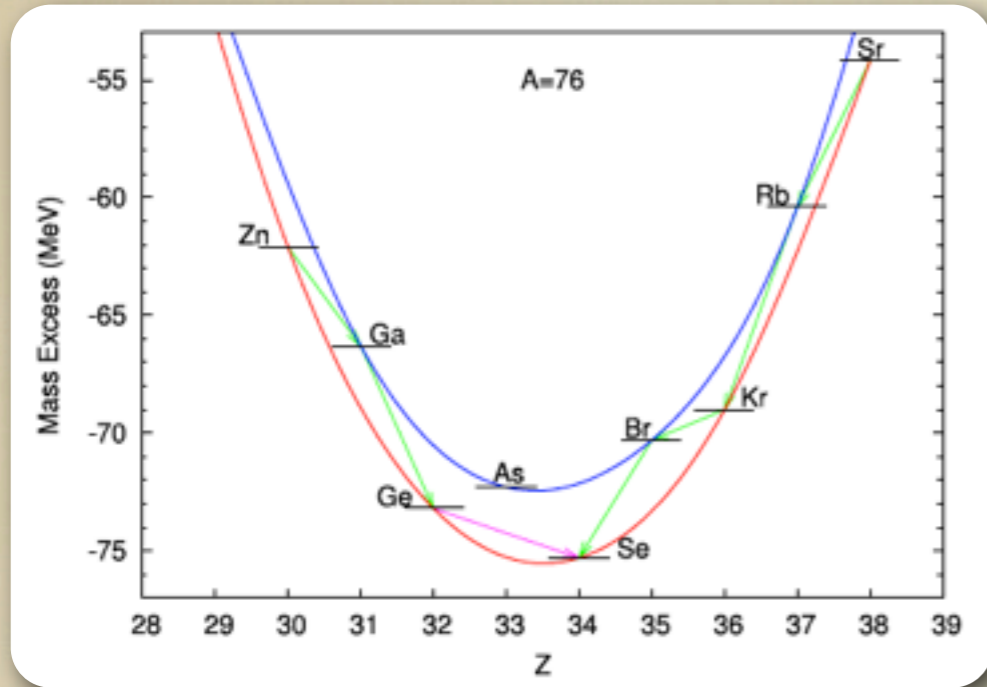


Exploring Majorana landscape

J.J. Gómez-Cadenas
Instituto de Física Corpuscular (CSIC & UVEG)

Florence, July, 2012

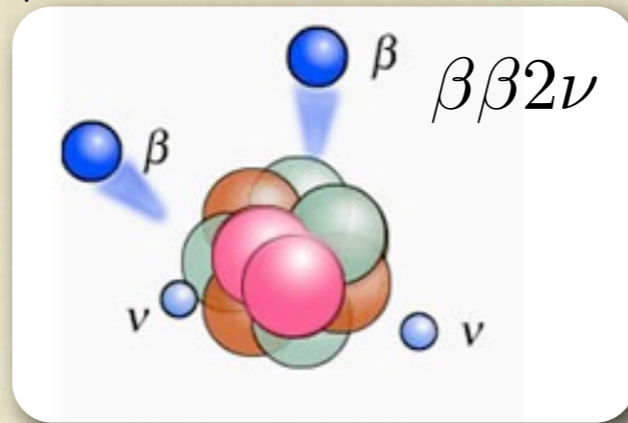
Double beta decay



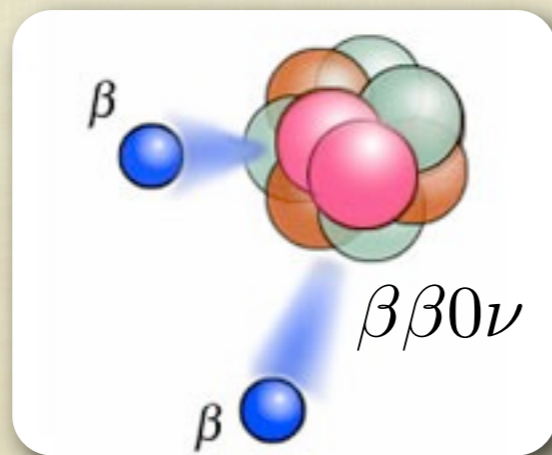
- About 10 isotopes, A ~70-150
- $Q_{bb} \sim 2-3.5$ MeV
- $T_{1/2} \sim 10^{18}-10^{20}$ yr.

Neutrinoless Double Beta Decay

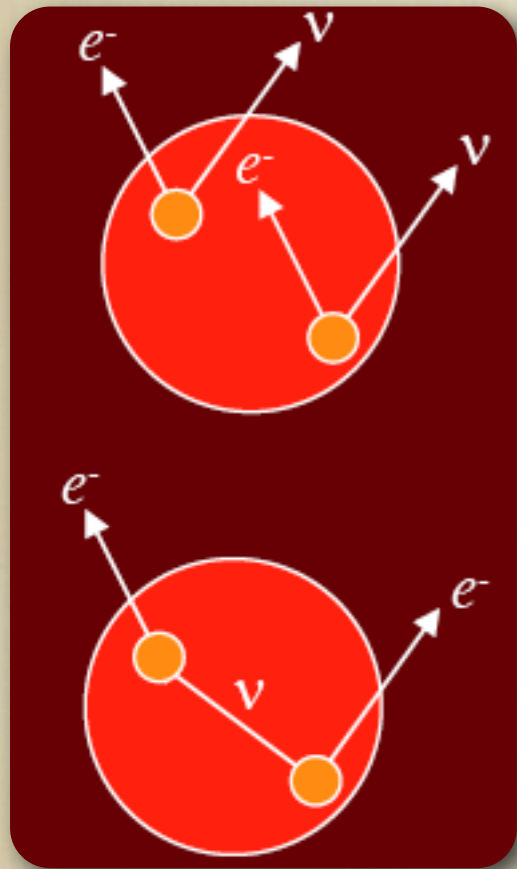
$$T_{1/2} \sim 10^{18} - 10^{20} \text{ y}$$



$$T_{1/2} > 10^{25} \text{ y}$$

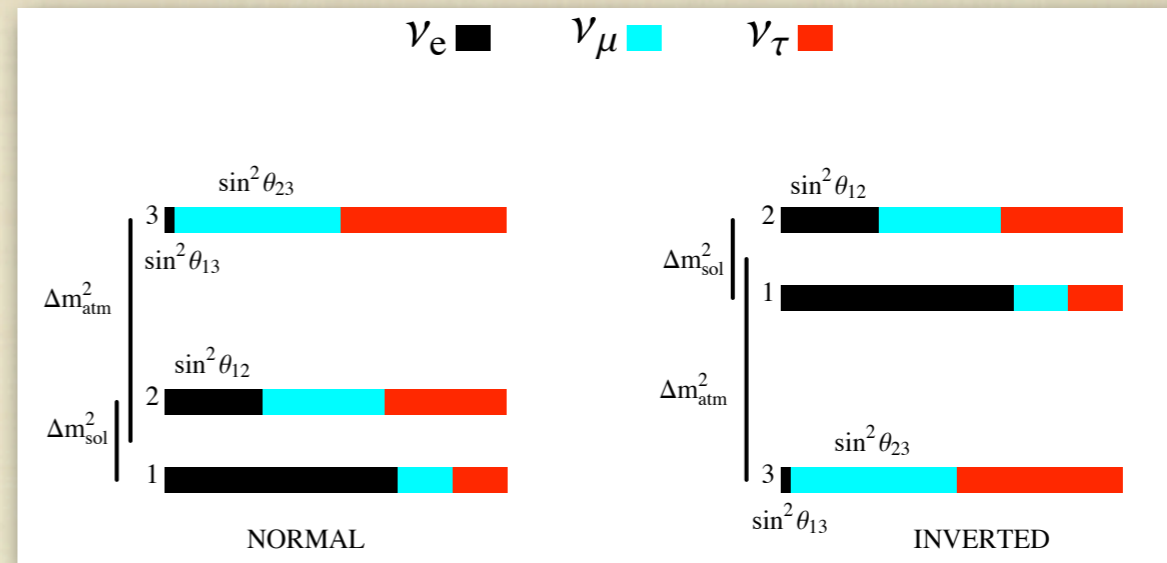
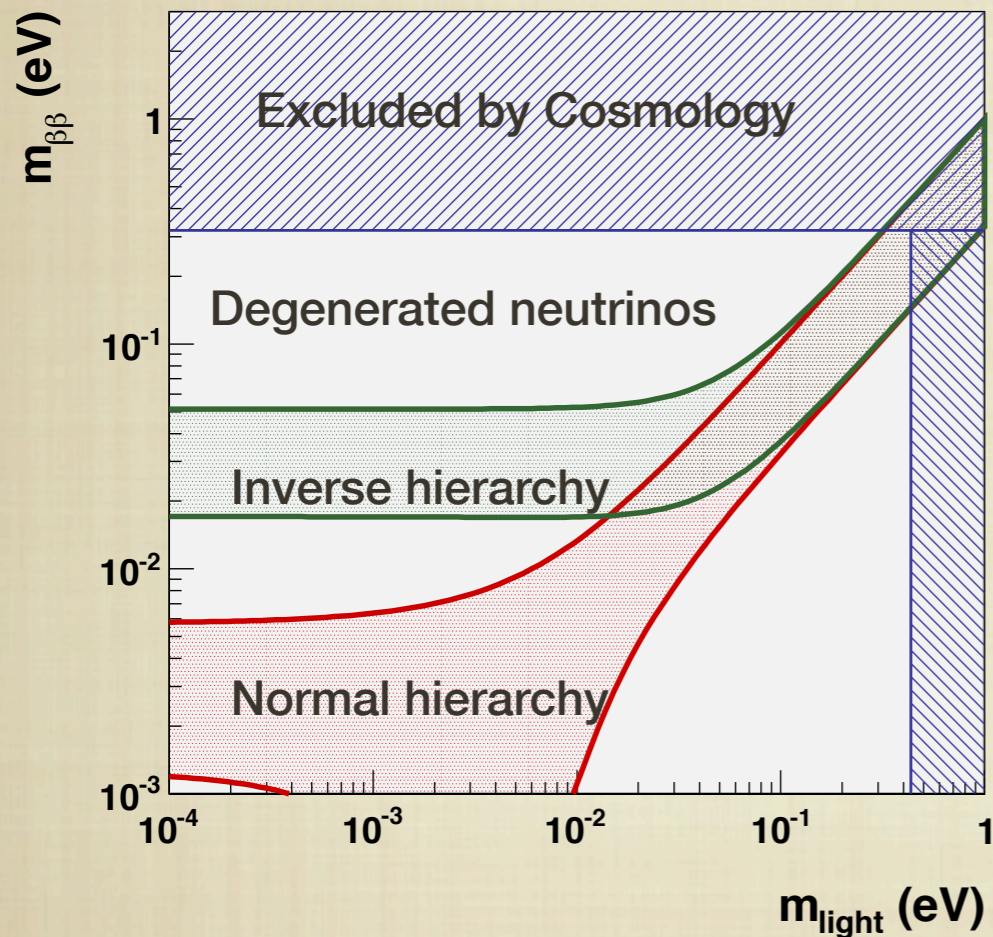


- If the neutrino is a Majorana particle the process called Neutrinoless Double Beta Decay may exist
- In $\beta\beta 0\nu$, no neutrinos are emitted. The sum of the energies of the two electrons equals the mass difference between mother and daughter nuclei ($Q_{\beta\beta}$).
- The process requires an helicity flip, and therefore it becomes more likely as the neutrino mass increases.



DBD and neutrino mass

$$(T_{1/2}^{0\nu})^{-1} = G^{0\nu}(Q, Z) |M^{0\nu}|^2 m_{\beta\beta}^2$$

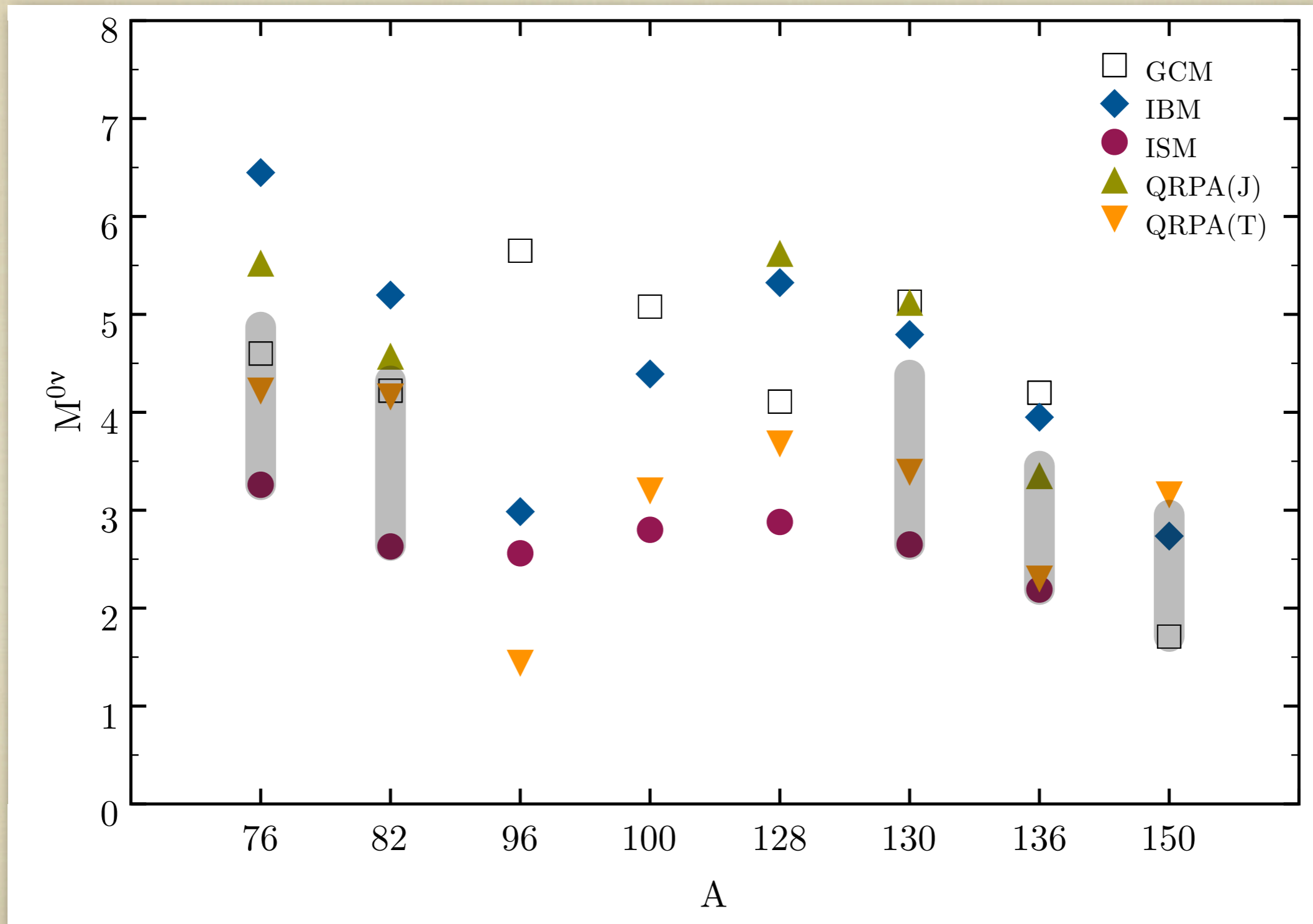


- EXO sets a limit of $T_{1/2}(Xe^{136}) = 1.6 \times 10^{25}$ yr (90% CL)

- Corresponding to $m_{\beta\beta} \sim 140 - 380$ meV

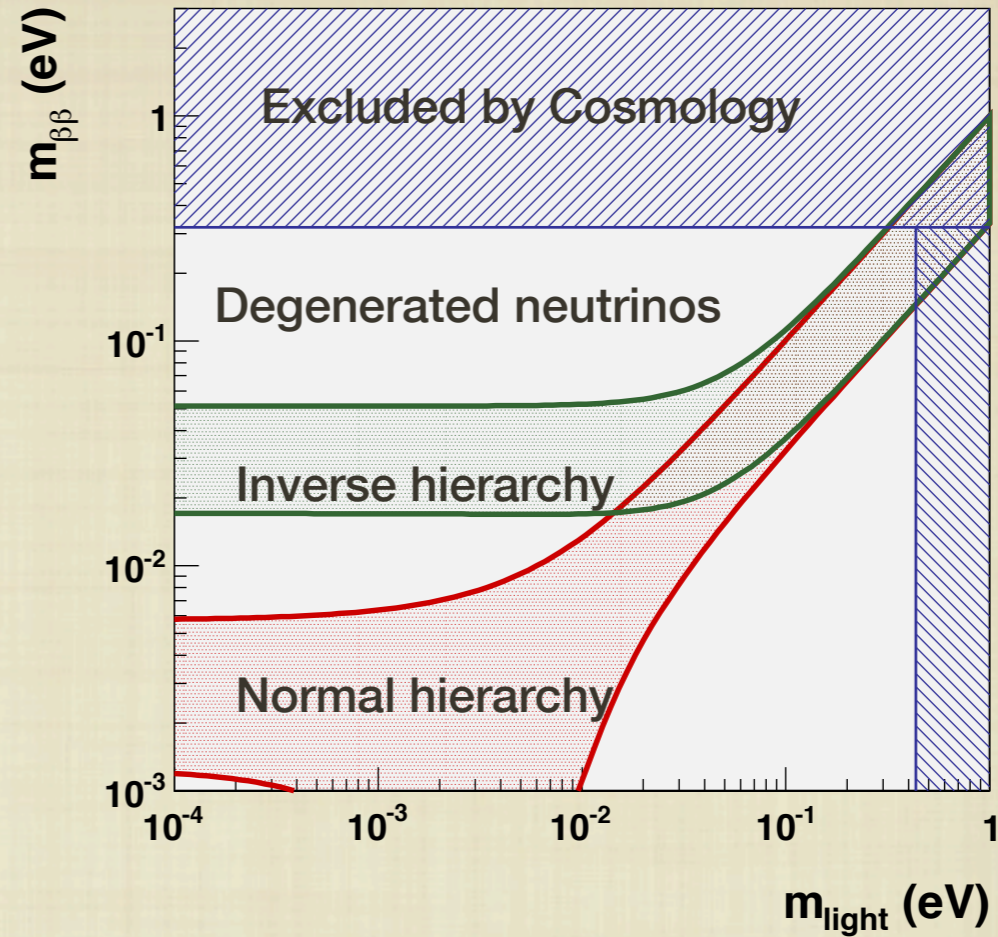
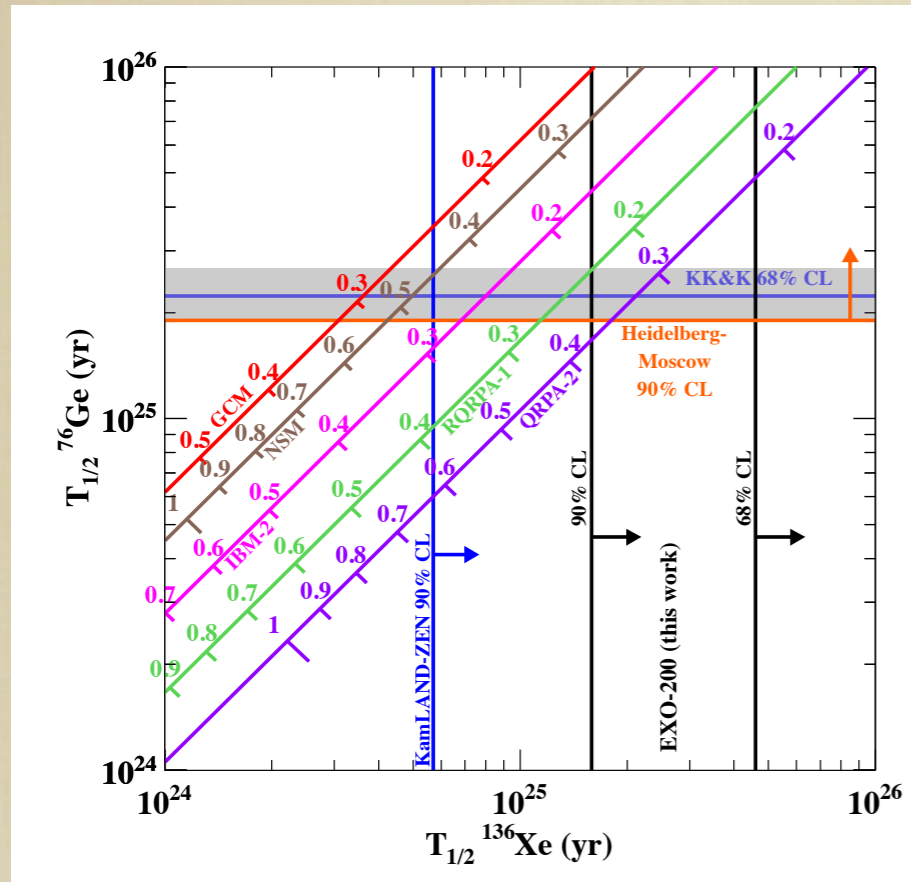
$$m_{\beta\beta} = \left| \sum_i m_i U_{ei}^2 \right|$$

NME Industry



Gomez-Cadenas *et al.*, *JCAP* **1106** (2011) 007

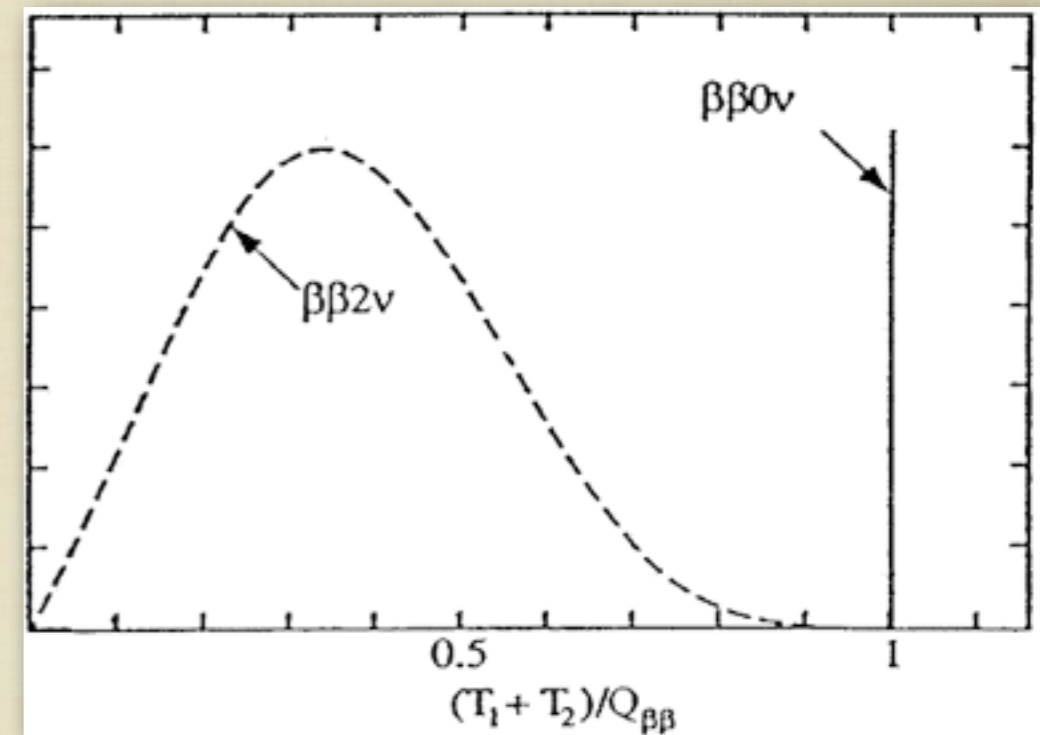
Current experimental situation



- EXO result has excluded the claim of KK for all but one of the NME's sets. It appears that KK claim will not hold water. GERDA should settle the matter soon.
- The region of $m_{\beta\beta}$ between 50-300 meV corresponds to the so-called degenerated hierarchy.
- EXO: $T_{1/2} \sim 10^{25}$ y ($m_{\beta\beta} \sim 150$ meV). To reach $m_{\beta\beta} \sim 50$ meV needs $T_{1/2} \sim 10^{26}$ y

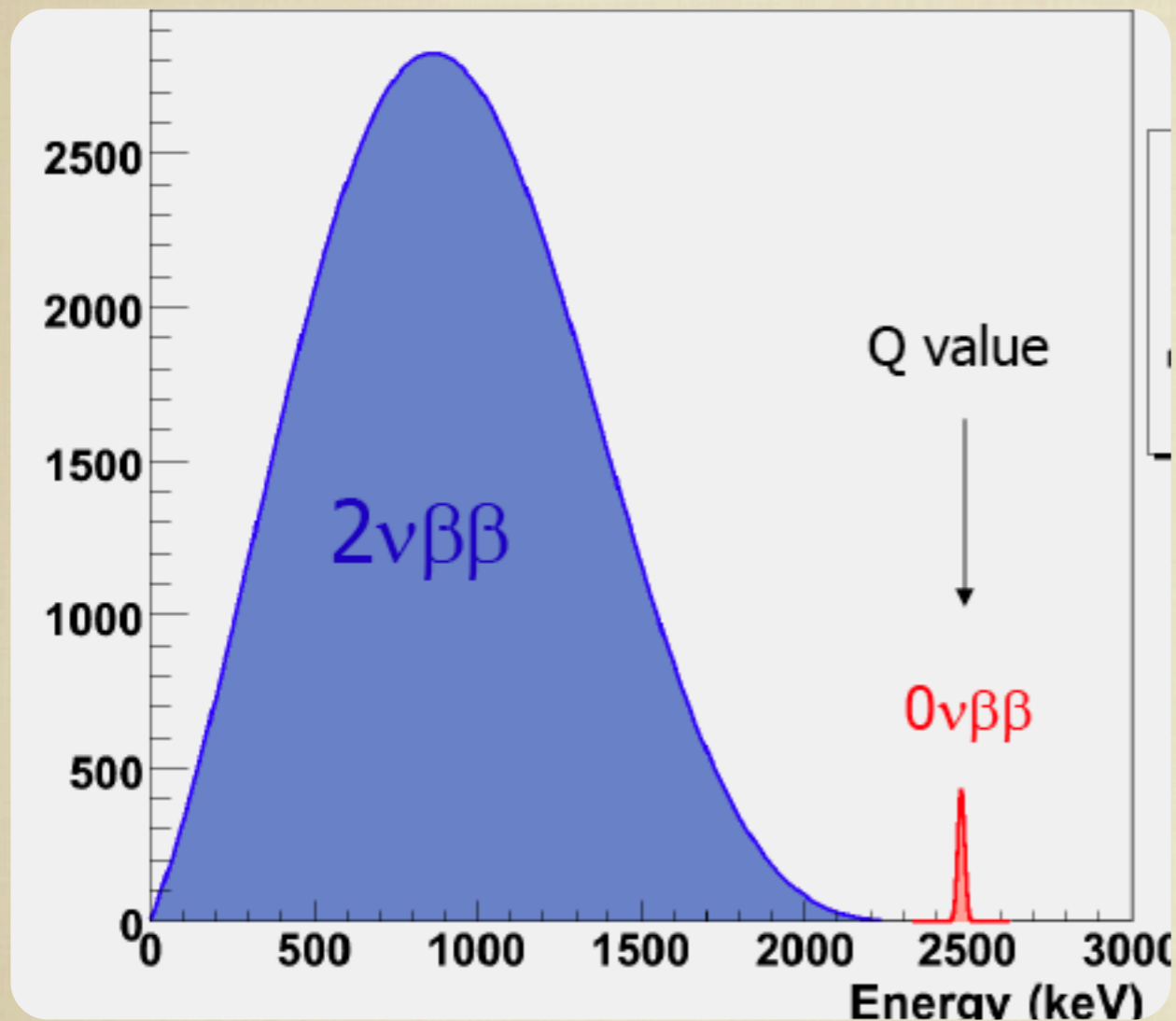
Ideal $bb0\nu$ experiment

- Get a large mass of double beta decay source ($N = MtN_A/A$).
- Measure the energy of the emitted electrons.
- Select those with $(T_1+T_2)/Q_{bb} = 1$
- Count the number of events and calculate the corresponding half-life.



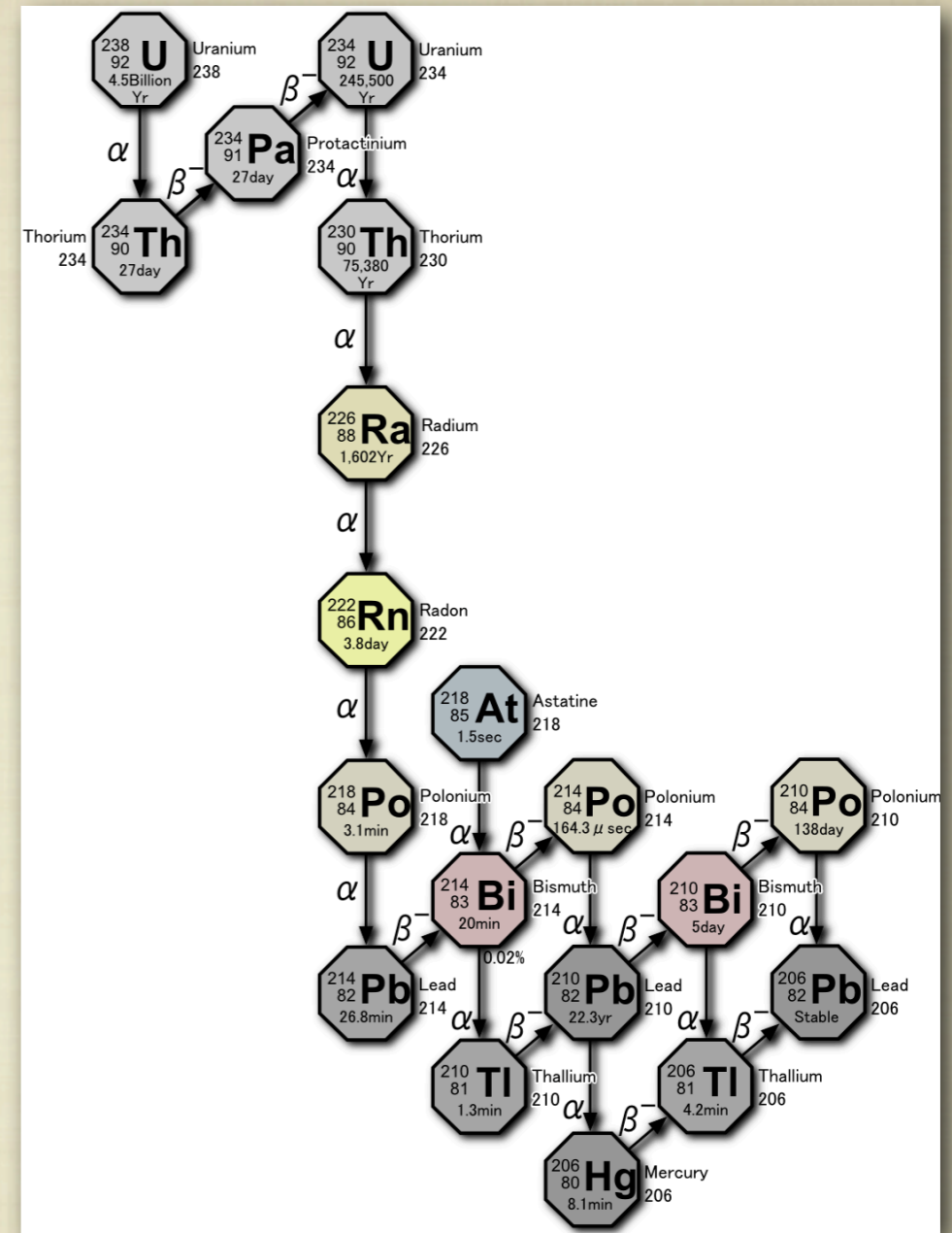
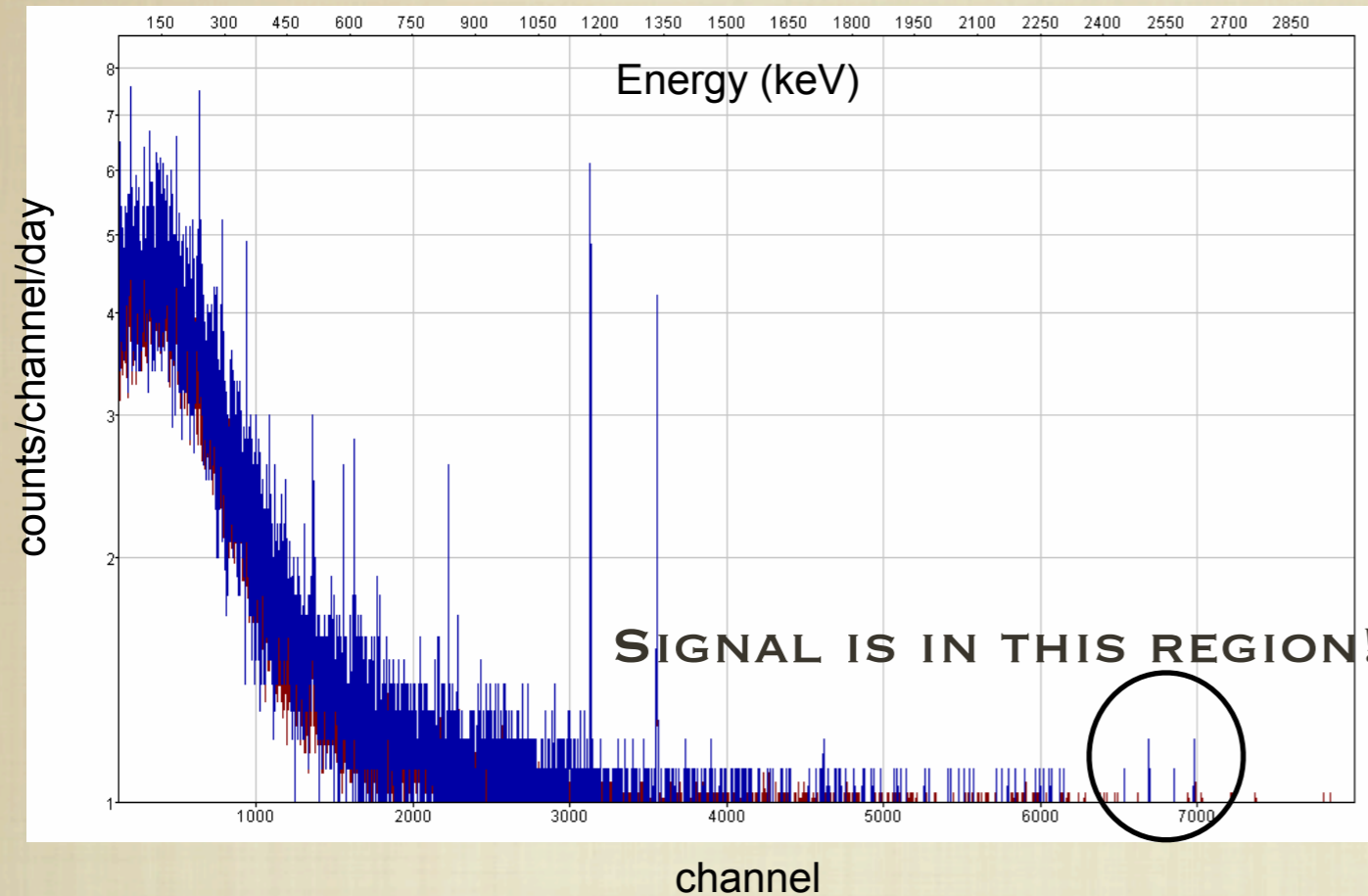
$$T_{1/2} = \log 2 \frac{N_A Mt}{A N_{\beta\beta}}$$

Energy resolution



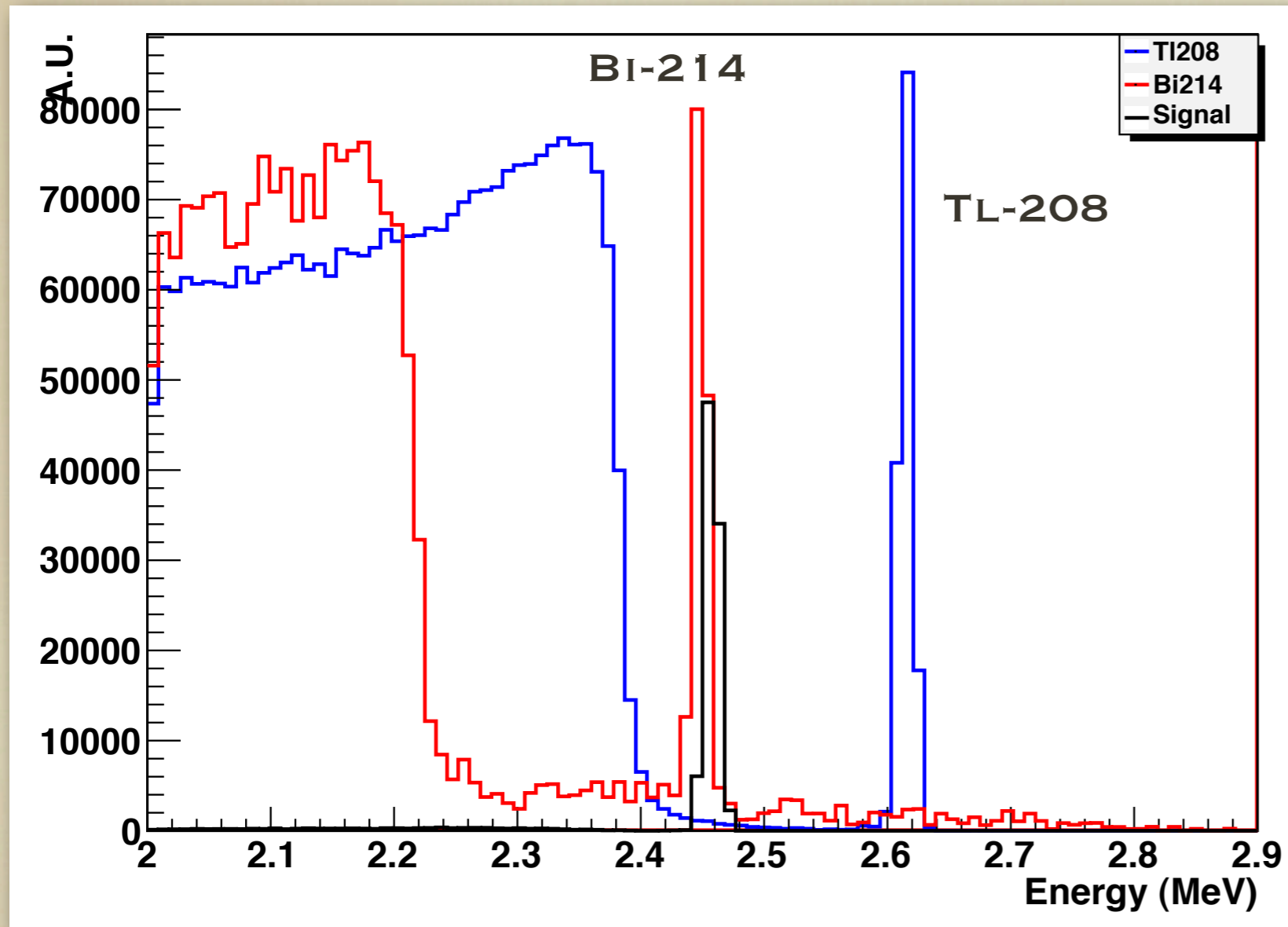
- If detector resolution is not perfect, energy spike around Q becomes a Gaussian.
- Background events, both from $bb2n$ process and from U & Th radioactive chains will leak into the Gaussian region (the ROI)
- Everything else (radiopurity, extra handles) being the same, experiments with superb energy resolution are preferred, to minimize the impact of background events.

Radioactive background



- Due to the radioactive chains of U and Th.
- Earth is a very radioactive planet. About 1 gr of U and 3 gr of Th per ton of rock.
- Lifetime of U-238 is 4.5×10^9 yr to be compared with a signal lifetime of $\sim 10^{26}$ (10^{27}) yr

Radioactive background in Xenon

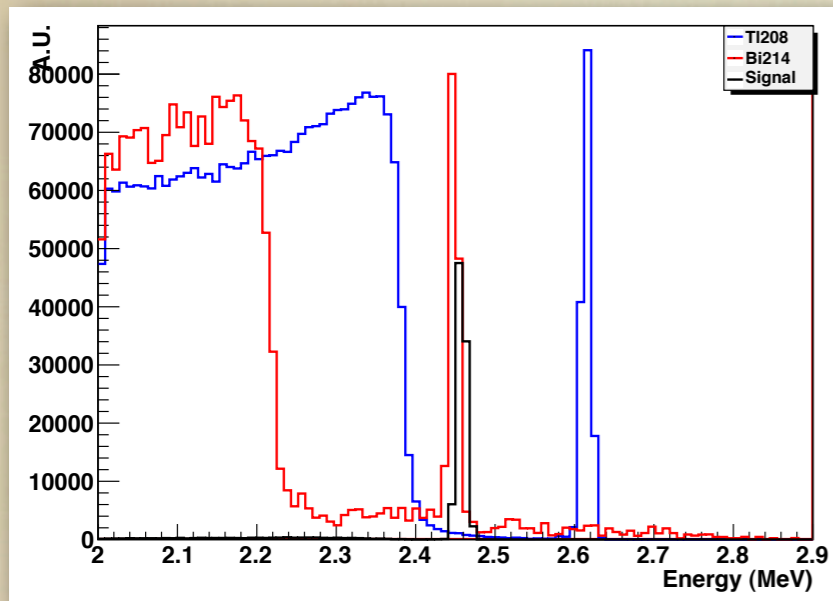


Arbitrary normalization

Assumed resolution: 1%
Q_{bb}

Main backgrounds: high-energy gammas from TI-208 and Bi-214
(natural radioactivity in detector materials).

Signal & Background

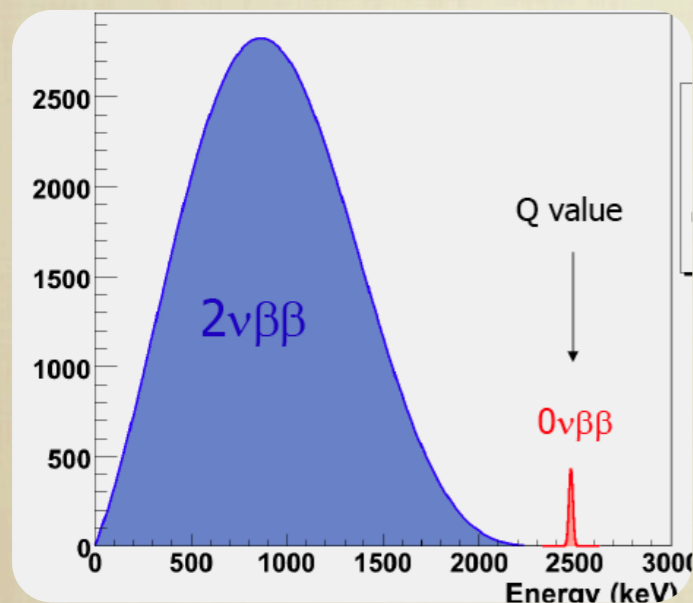


Imagine that signal is at a level of $T_{1/2} \sim 10^{25}$ y. How many events per year and kg?

$$T_{1/2} = \log 2 \frac{N_A M t}{A N_{\beta\beta}}$$

$$N_{\beta\beta}(Xe^{136}) = \log 2 \frac{6 \cdot 10^{23} \cdot 10^3(g) \cdot 1(y)}{10^{25} \cdot 136} \sim 0.1$$

Need a large mass to see the signal (100 kg yr for $T_{1/2} = 10^{25}$ y)



Imagine that natural background for Bi-214 line is of the order of 1 μ Bq. This is one count of background every 10^6 seconds. One year has $3 \cdot 10^7$ seconds. Thus 30 counts of Bi-214 go into the Bi-214 peak which overlaps at 50 % with signal peak. B/S $\sim 1!!!$

Improving T is difficult

$$T_{1/2}^{-1} \propto a \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{\Delta E \cdot B}}$$

- K → isotope yield
- ε → detection efficiency
- M → isotope mass
- t → running time
- b → background rate
- ΔE → energy resolution

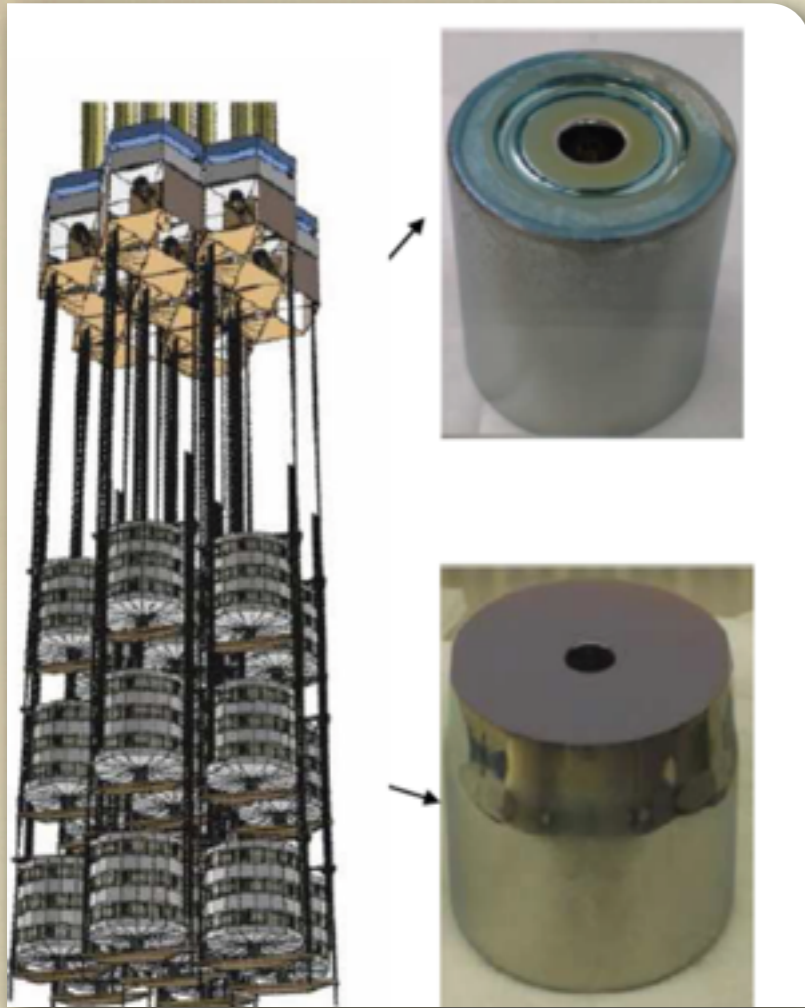
- Assume a , ε and ΔE constant.
- To improve $T_{1/2}$ by 10 one needs to: a) Increase Mt by 10^2 or decrease B by 10^2 or increase Mt by 10 and decrease B by 10.
- CHALLENGE: Build a detector with 10 times the mass and 10 times less background.

Improving m is VERY difficult

$$m_{\beta\beta} = K \sqrt{1/\varepsilon} \left(\frac{b \Delta E}{Mt} \right)^{1/4}$$

- Today: 150 meV. Degenerated: 50 meV. Inverse: 20 meV
- First jump: Improve () by $3^4 \sim 100$ Second jump: Improve by $6^4 \sim 1000$
- EXO: ~ 100 kg yr: First jump: 10,000 kg yr, second jump: 100,000 kg yr
- No go, unless one reduces B at the same time by a factor 10 (100).

Calorimeters/Bolometers



$$T_{1/2}^{-1} \propto a \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{\Delta E \cdot B}}$$

$\Delta E \rightarrow 0$ initial reach OK

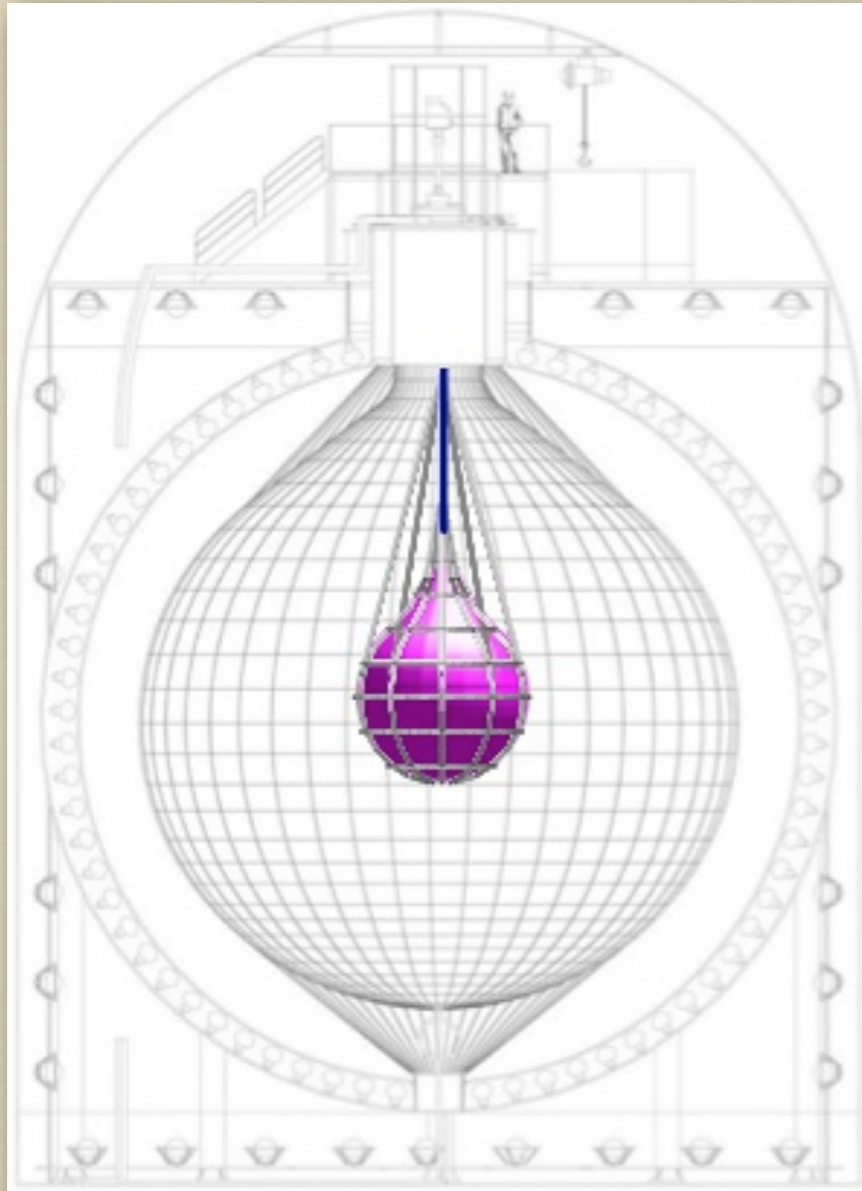
$b \rightarrow S/V$ large, alpha particles.

$M \rightarrow$ expensive and modular (no scale)

$b \rightarrow$ Can improve by a factor 10 with advanced techniques.

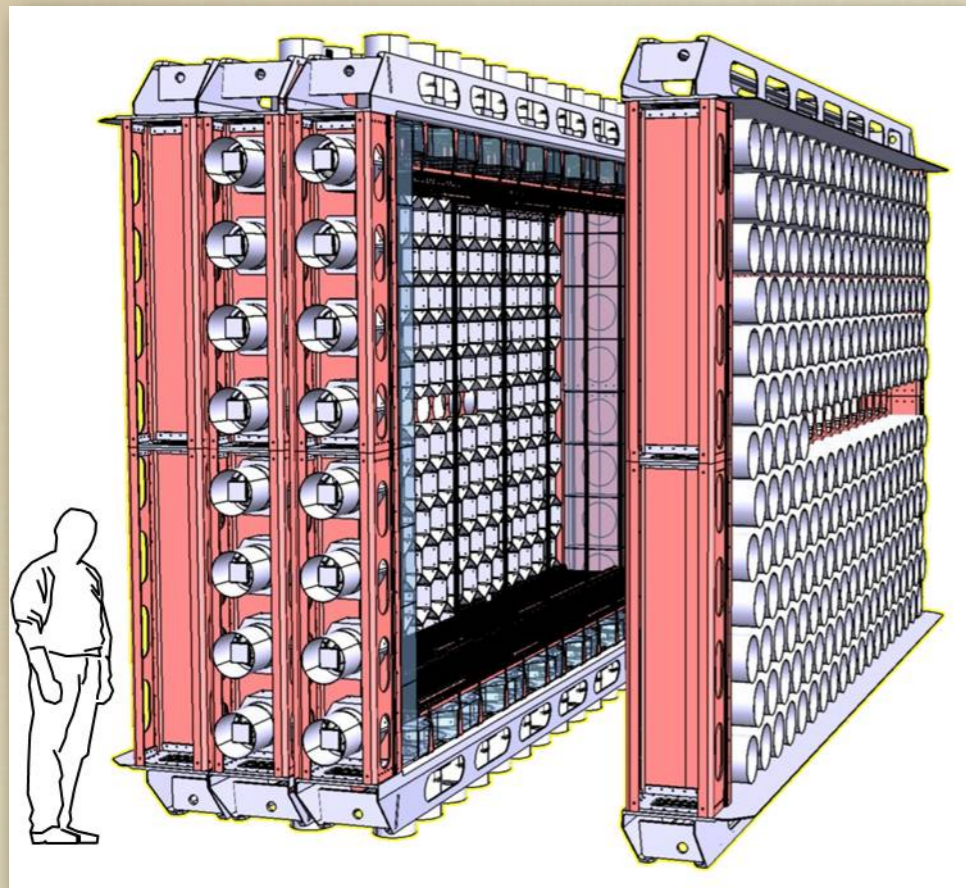
Main limitation: S/V and Mass.

Low resolution calorimeters

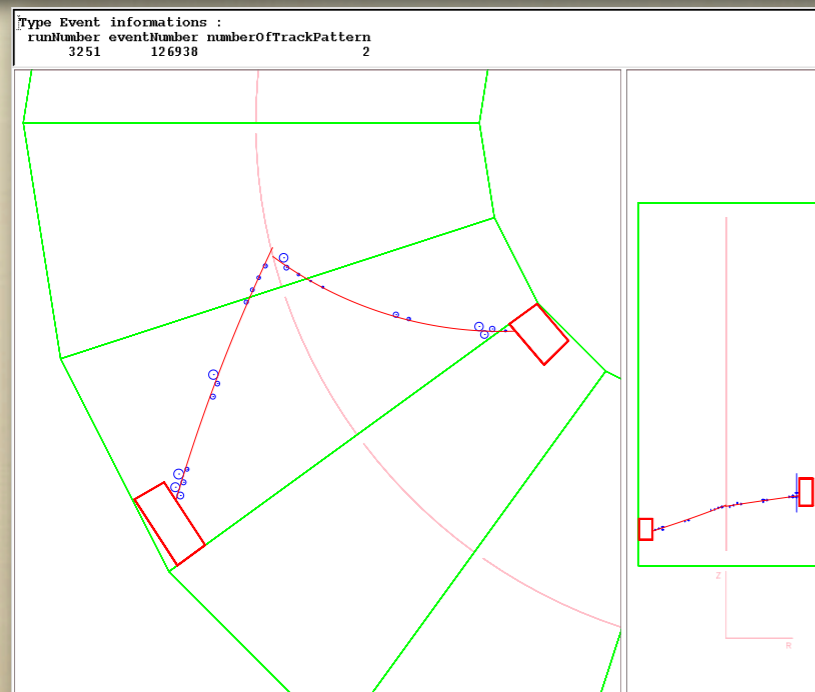


- Enriched xenon dissolved in liquid scintillator.
- Poor resolution, 10% FWHM at the Q-value.
- Easy to pile up large mass
- Difficult to control backgrounds (K-ZEN initial run 10^2 larger than expected)

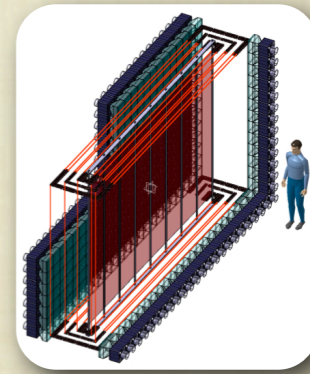
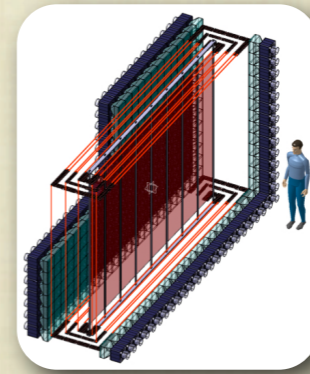
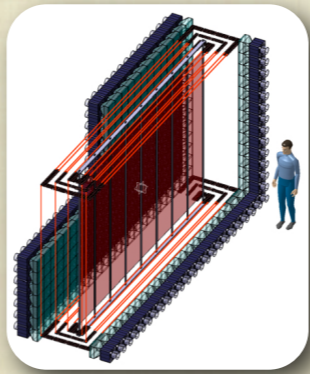
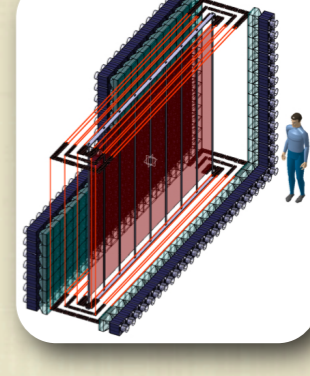
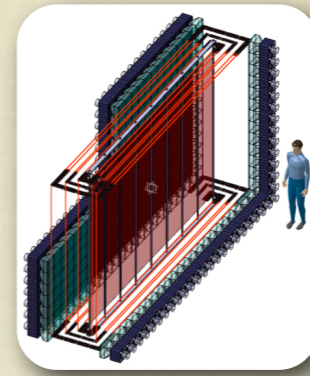
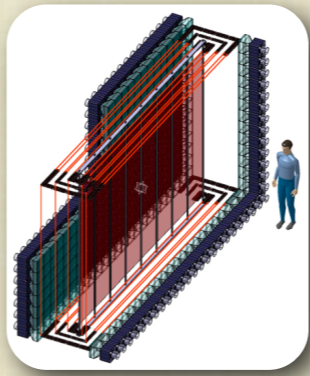
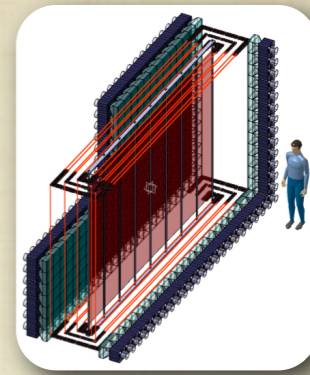
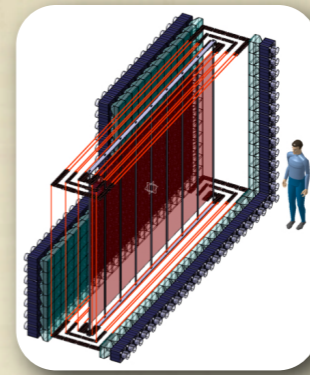
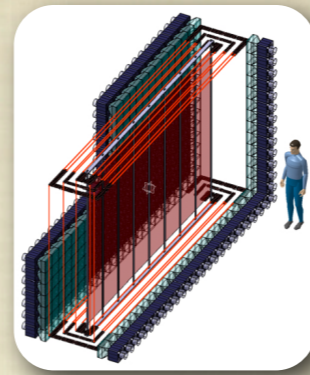
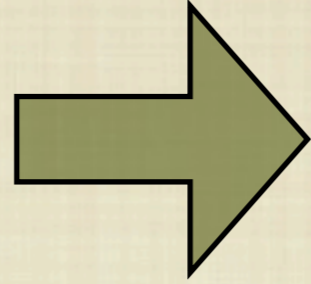
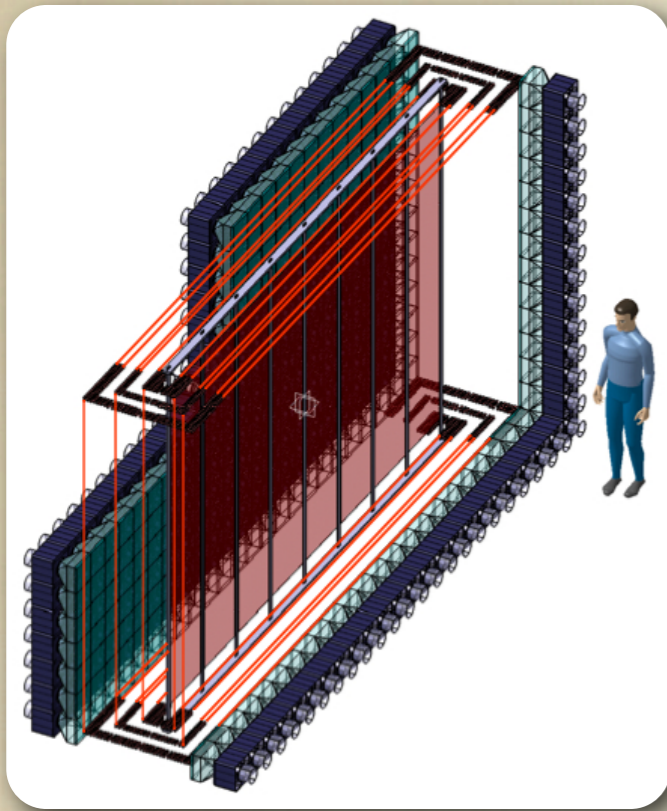
Modular/non-homogenous



- Thin source foil (Se-82) within a tracking chamber surrounded by a calorimeter.
- Mediocre resolution, 4% FWHM at Q-value.
- Low efficiency ($\sim 30\%$).
- Extra handles (tracks)



No-go technique



PRICE & EFFORT SCALES LINEARLY

BACKGROUNDS (PROPORTIONAL TO SURFACES) SCALE LINEARLY

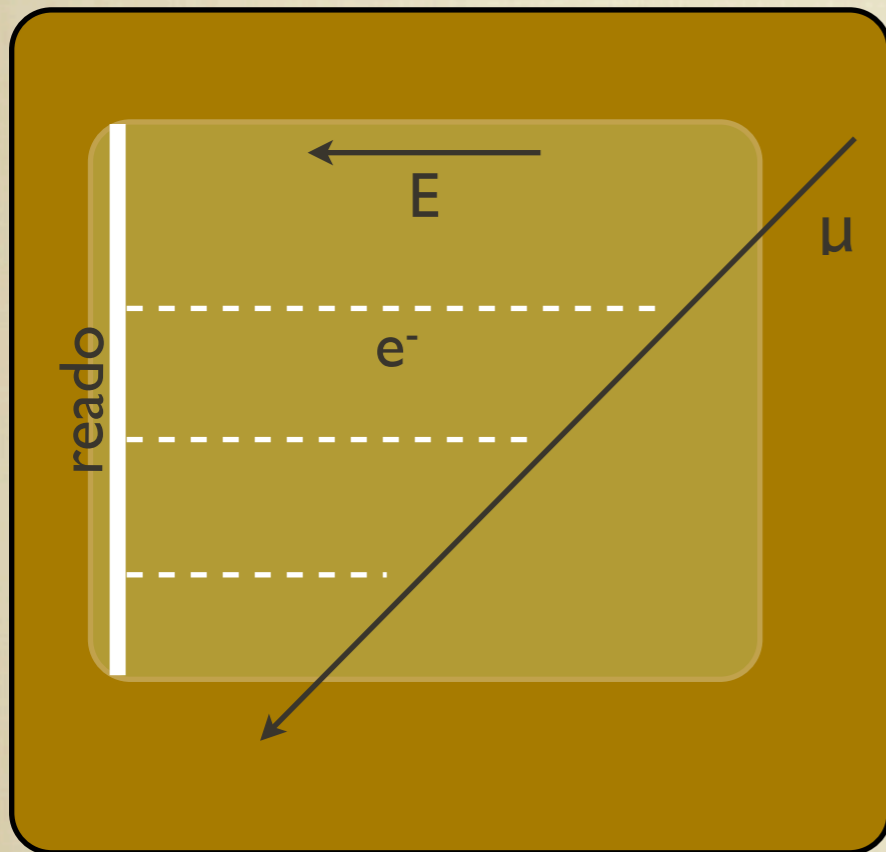
NOT HOMOGENOUS DETECTOR

NOT SUITED EVEN FOR CURRENT MODEST SCALE

BEST FEATURE OF DETECTOR: PROPAGANDA.

The TPC detector

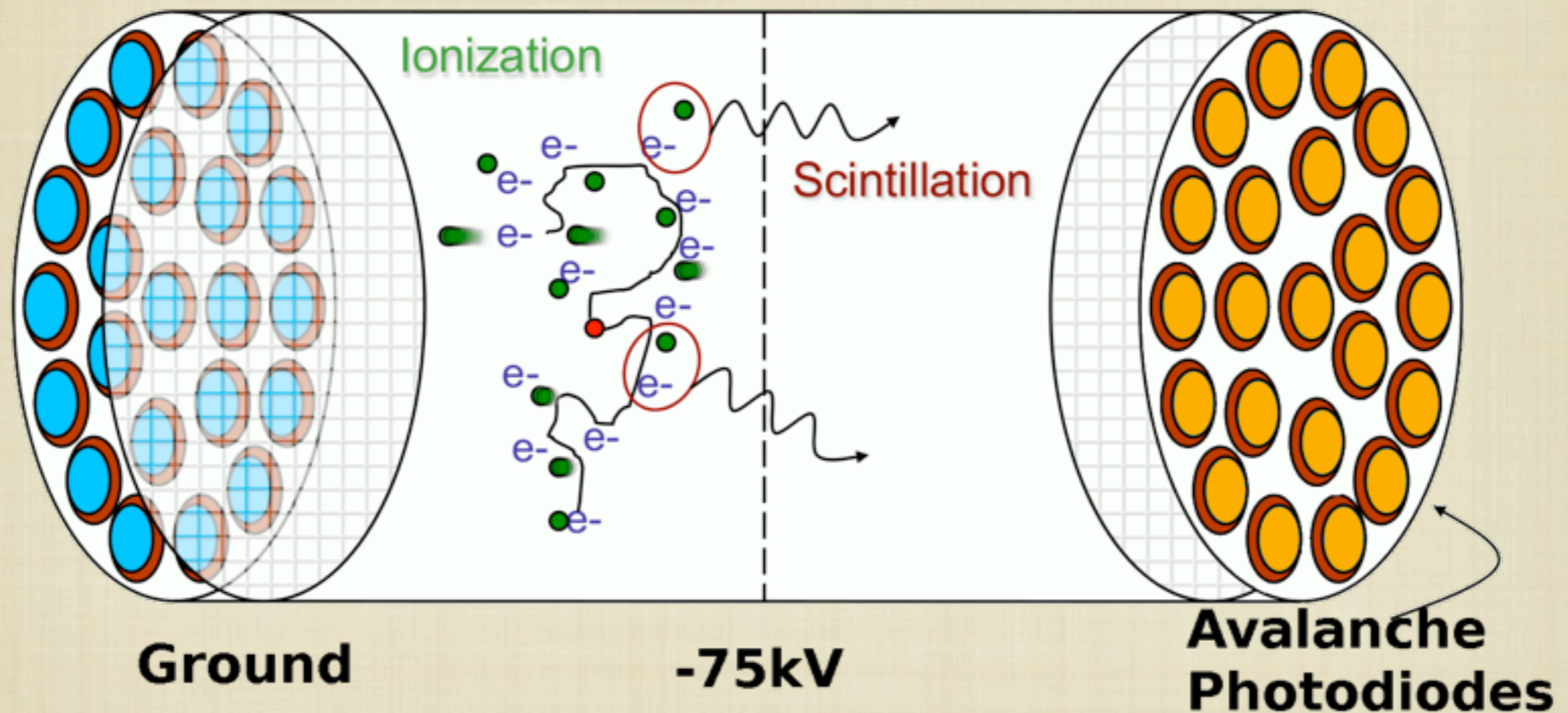
Time Projection Chamber: invented by D. Nygren in the 1970's.
Can be seen as an electronic bubble chamber.



- REQUIRES A NOBLE GAS TO OPERATE
- CHARGED PARTICLES TRAVERSING TPC IONIZE GAS LEAVING A TRACK
- IF TRACK STOPS INSIDE TPC THEN ITS ENERGY IS CALORIMETRICALLY MEASURED (WITH GOOD RESOLUTION)
- LARGE VOLUME POSSIBLE (THUS LARGE MASS)
- NO SURFACES IN FIDUCIAL VOLUME FOR BACKGROUND IONS TO ATTACH TO

Imaging Xe chambers

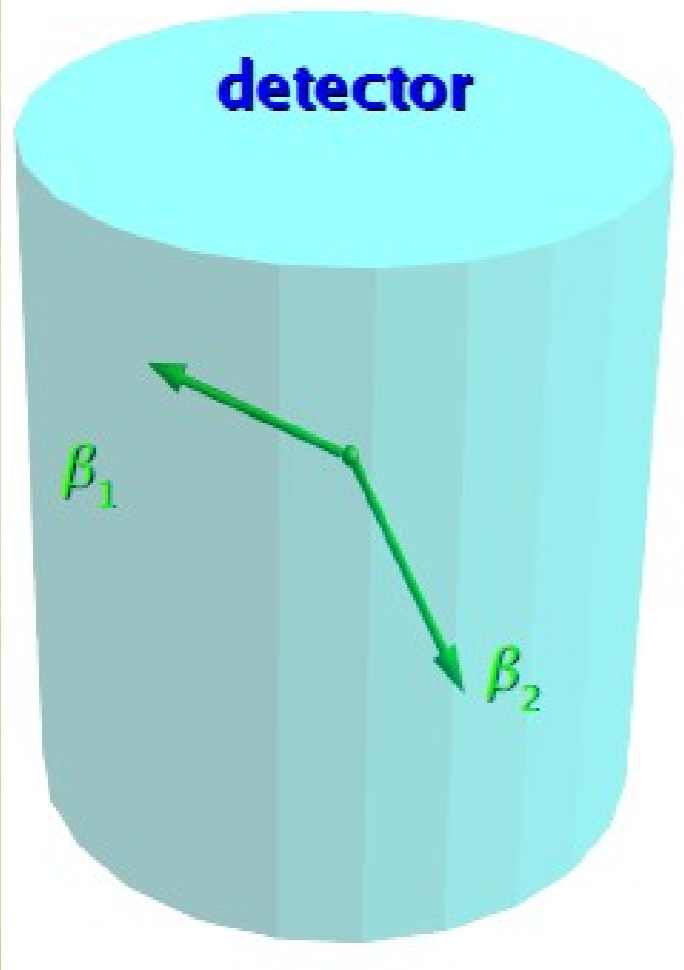
high-pressure gaseous Xenon (HPGXe) or LXe



Ionization and Scintillation in Xenon can be recorded in Xe chambers

Xe chambers are an homogenous detector

Source = Detector



$$T_{1/2}^{-1} \propto a \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{\Delta E \cdot B}}$$

Large TPC

Mass goes with L^3

Large mass and good fiduciality

Backgrounds scale with L^2 .

Improve (doing nothing) as you make it larger.

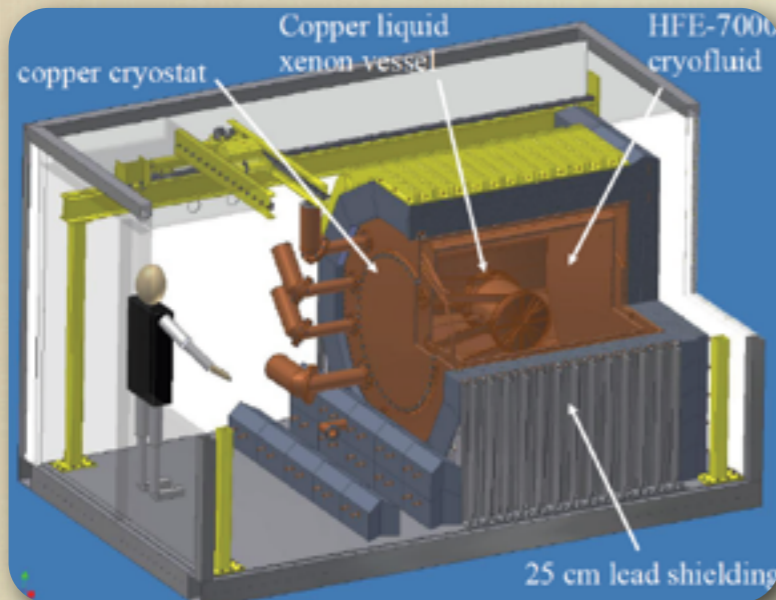
$$T_{1/2}^{-1} \propto a \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{\Delta E \cdot B}}$$

Easy to enrich to >90% and “cheap”.

Isotope	Natural Abundance (%)
^{48}Ca	0.2
^{76}Ge	7.8
^{82}Se	9.2
^{96}Zr	2.8
^{100}Mo	9.6
^{110}Pd	11.8
^{116}Cd	7.5
^{124}Sn	5.6
^{130}Te	34.5
^{136}Xe	8.9
^{150}Nd	5.6

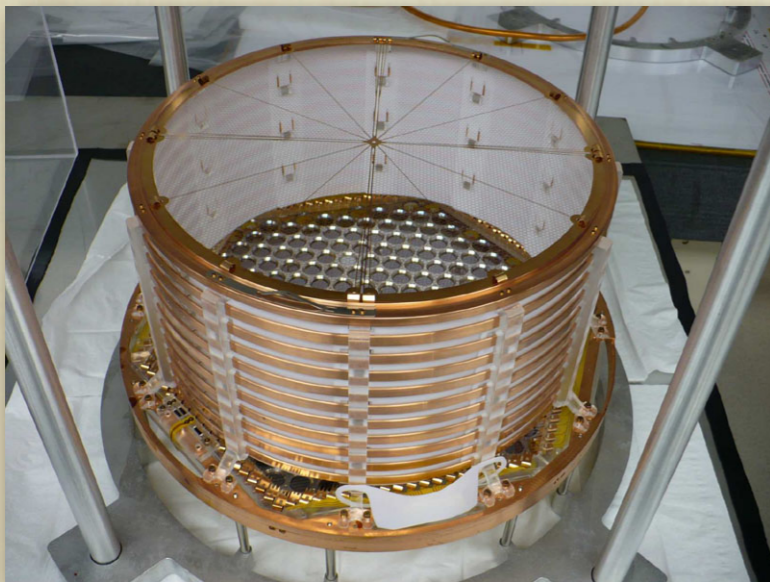


EXO-200



200 kg of liquid Xenon TPC
~4 % FWHM at Q_{bb}
70% efficiency (hard fiducial cut needed for self-shielding)
Bkgnd --> $\sim 10^{-3}$ c/(kg keV y)

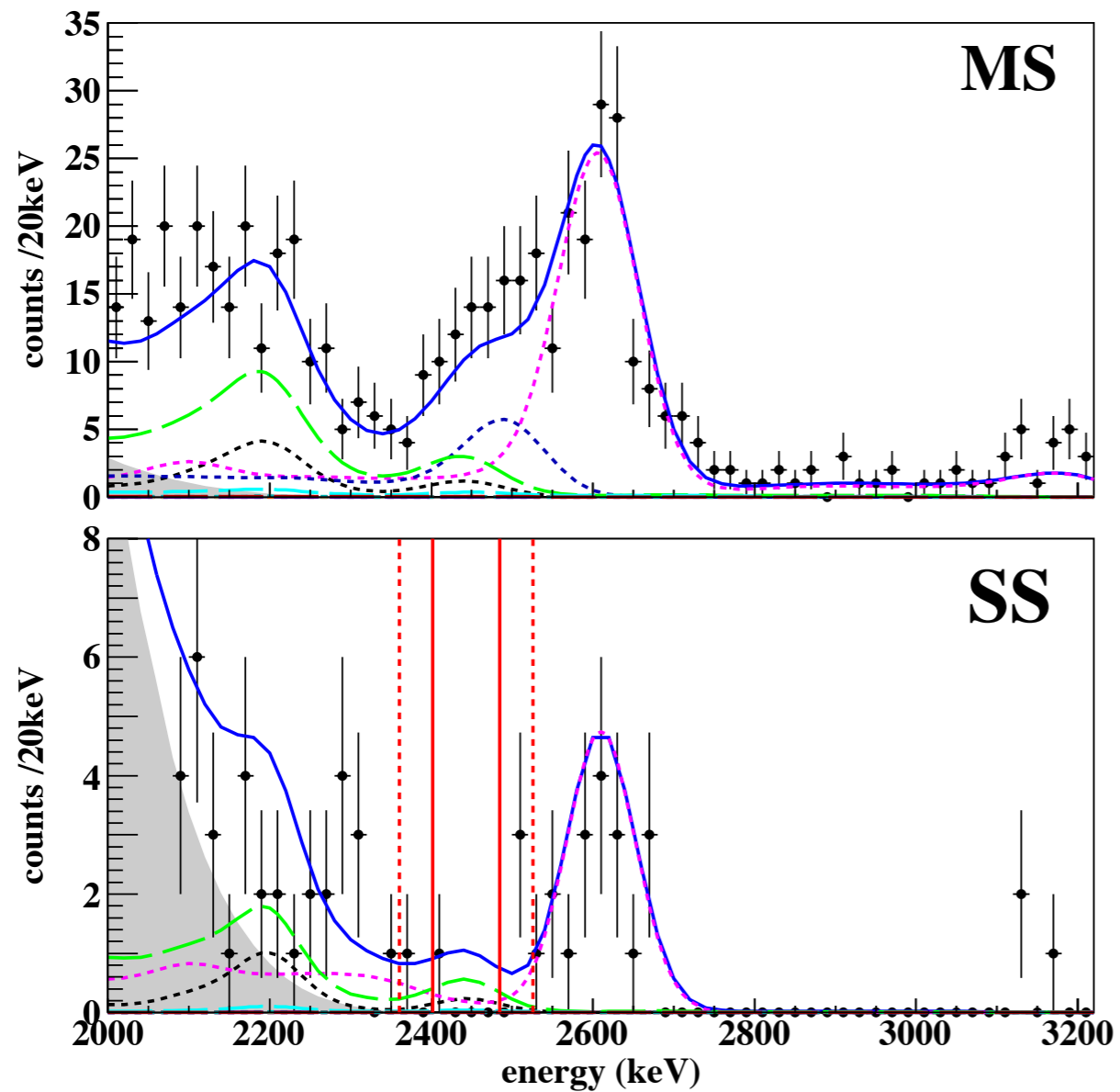
Large mass easy to achieve (liquid Xenon is very dense).



Strong points: compact, self-shielding, mass, scalability.

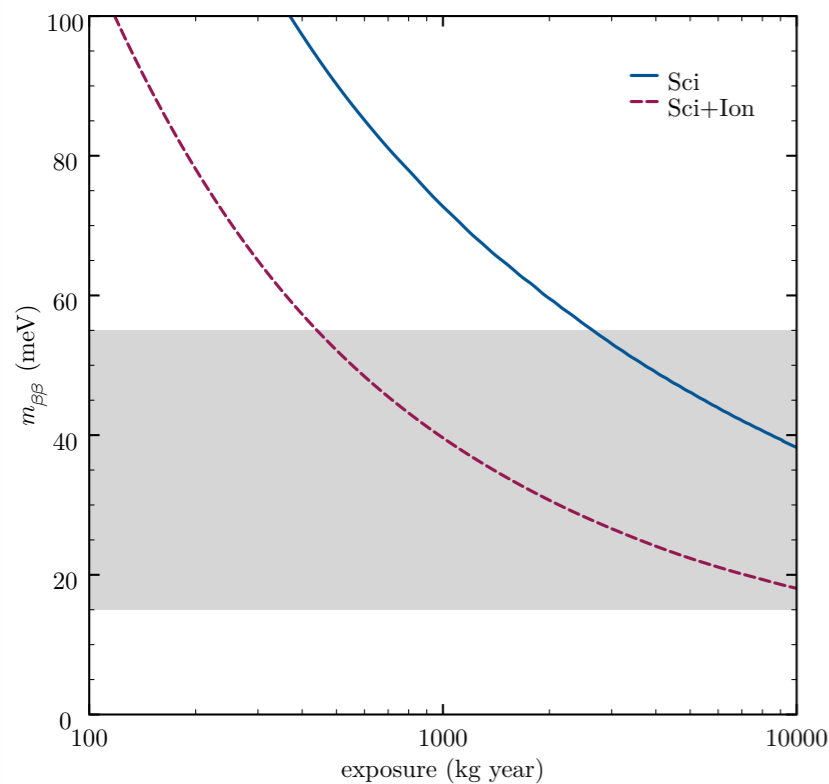
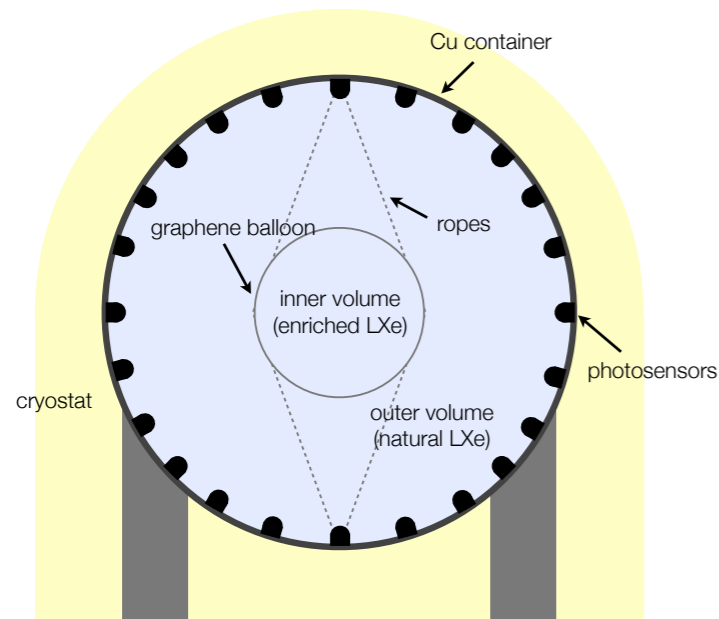
Weak points: mediocre resolution, low efficiency for effective shielding.

EXO-Limitations



- Expects 4 events in 1 sigma.
- Observes 4.
- Got lucky.
- Hard to improve with exposure (could get worse).
- Limitations: energy resolution lack of extra handles, expensive self-shielding.

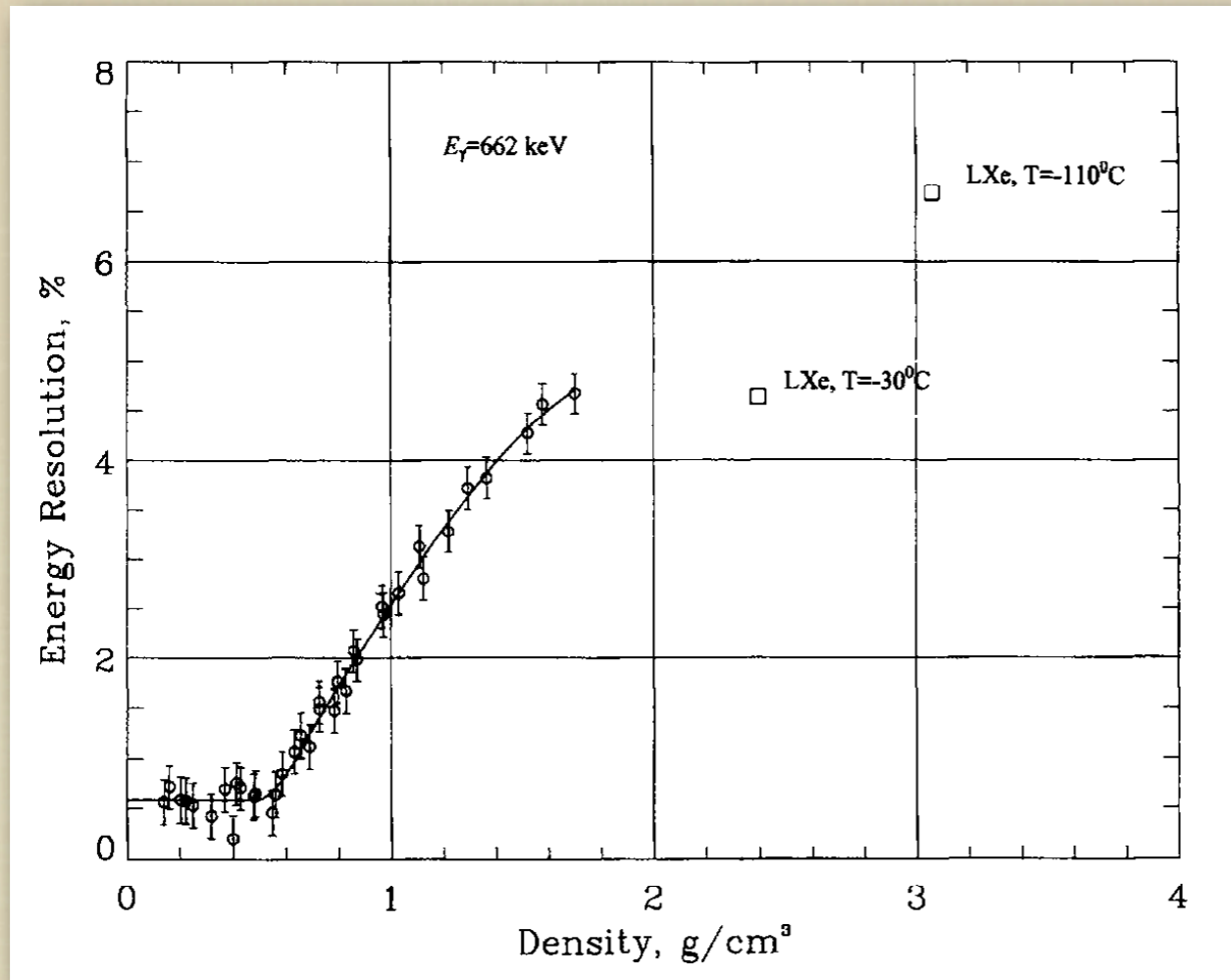
GRAXE: A concept to improve LXe reach



- GraXe is an spherical TPC. Conceptually identical to EXO.
- But EXE isolated from background by a buffer of pure NXE (no radioactive background)
- EXE enclosed in a graphene balloon that lets UV light through (also perfect metallic conductor, for spherical TPC).
- 20 tons of NXE will kill PMT radioactive background (and make up for a nice DM experiment)
- 1 ton extremely isolated EXE.
- Sci only (mediocre due to poor resolution: KZEN is a no-go in the long run)

Energy resolution in HPXe

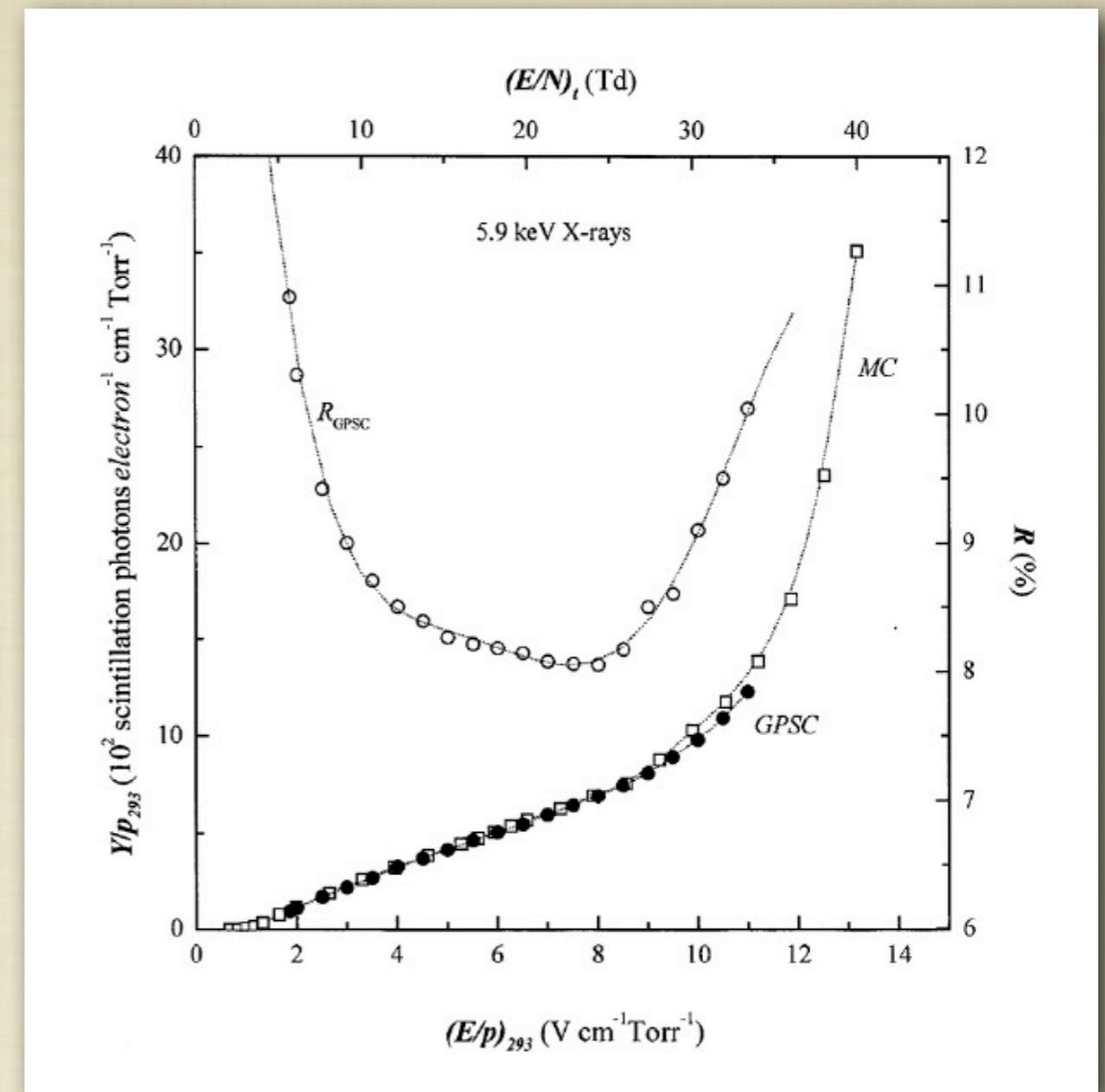
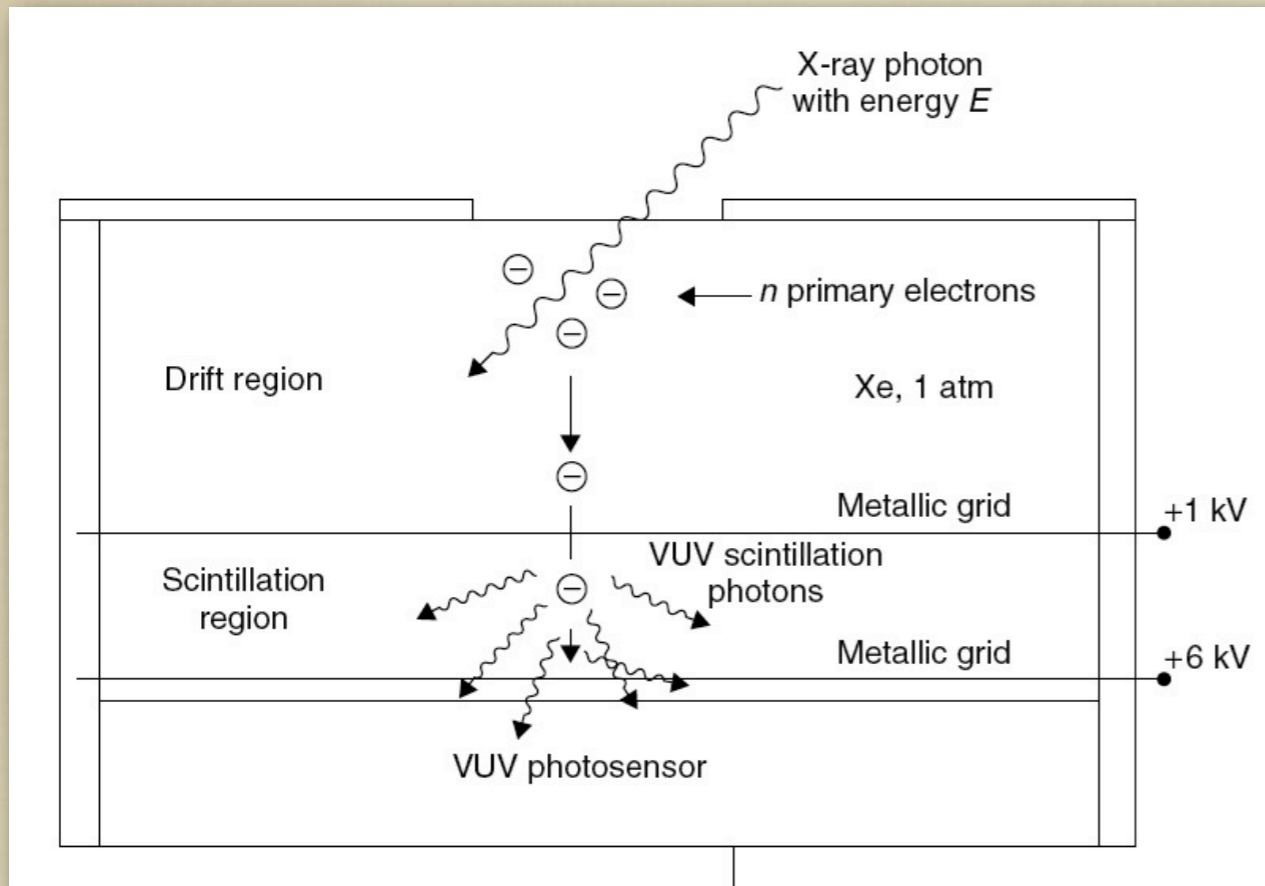
Bolotnikov and Ramsey, NIM A 396 (1997)



$$T_{1/2}^{-1} \propto a \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{\Delta E \cdot B}}$$

- Intrinsic resolution (Fano factor) at $Q_{\beta\beta}$ (2458 keV): 3×10^{-3} FWHM.
- Best experimental result: 5×10^{-3} FWHM.

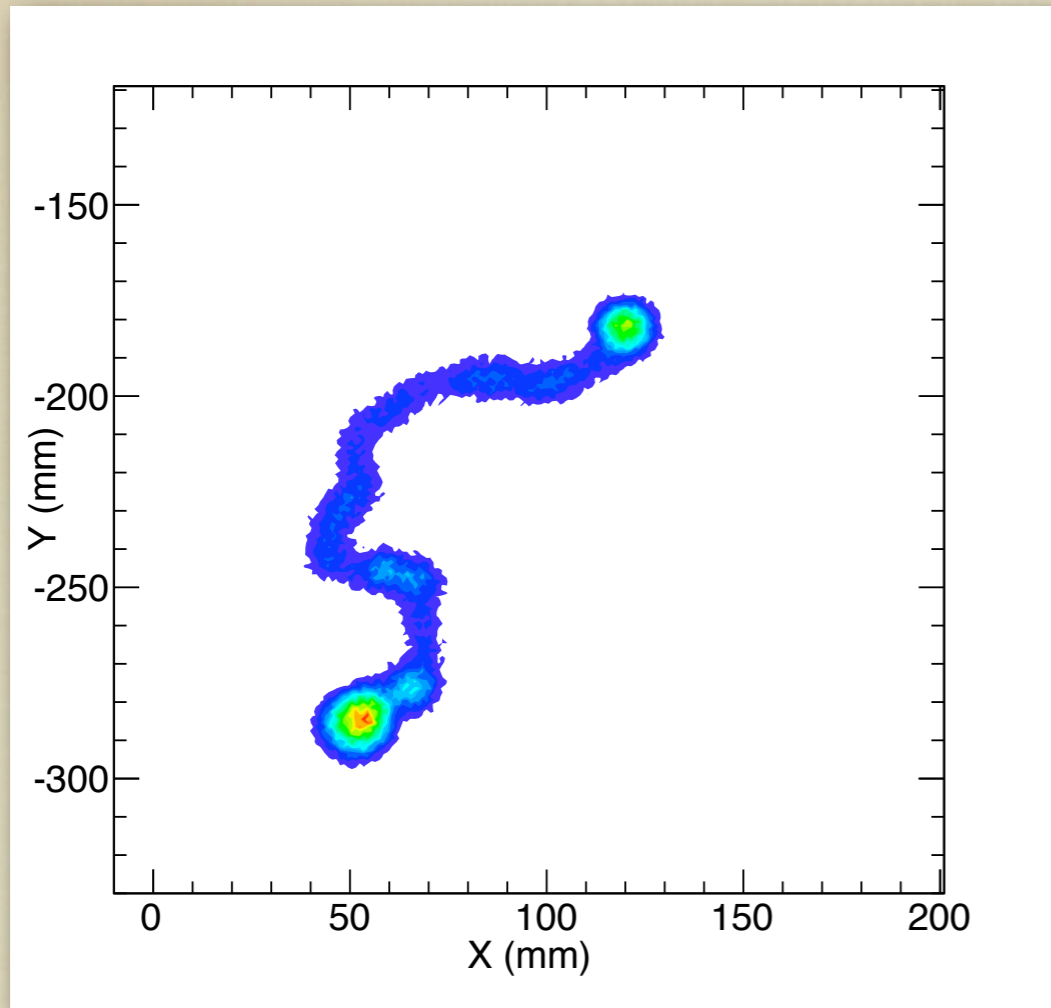
Electroluminescence



Emission of scintillation light by atoms excited by a charge accelerated by a moderate electric field.

Linear process, sub-poissonian fluctuations, huge gain at $3 < E/p < 6$ kV/cm/bar.

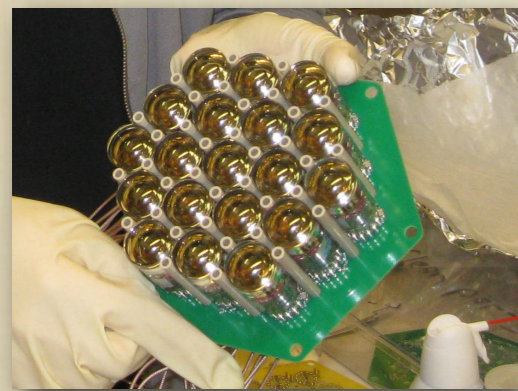
Tracking in HPXe



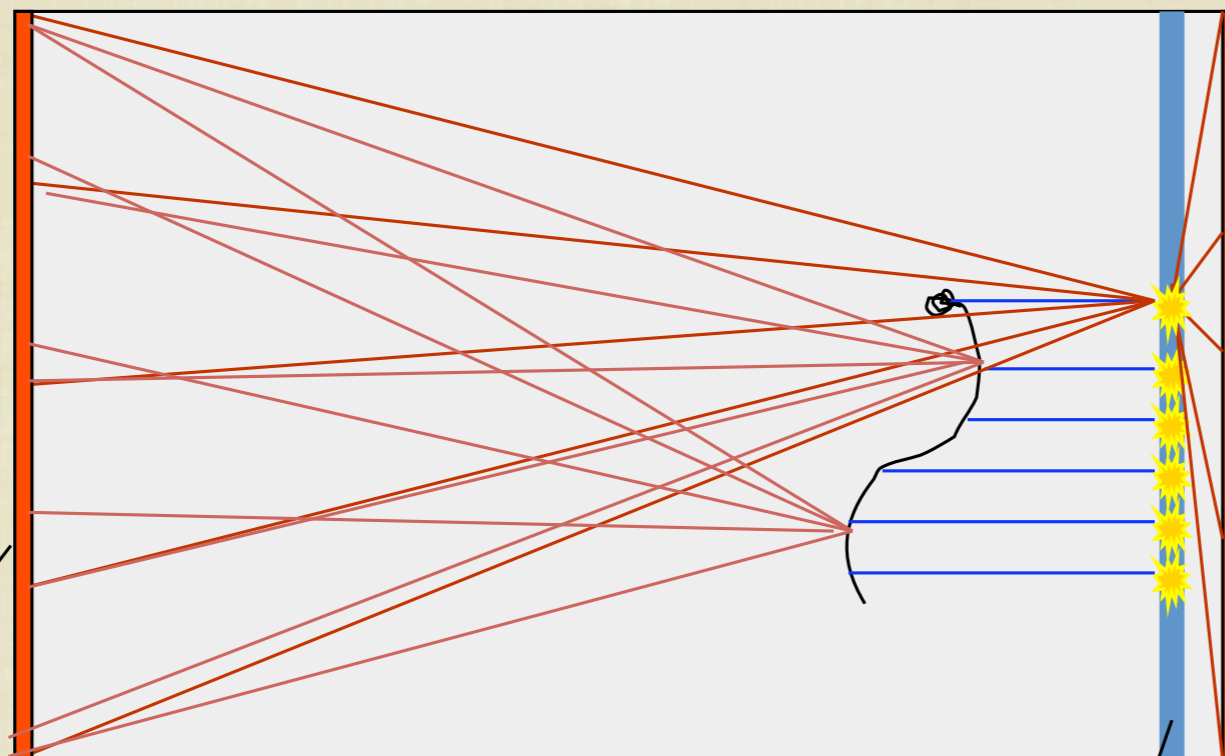
$$T_{1/2}^{-1} \propto a \cdot \varepsilon \cdot \sqrt{\frac{M \cdot t}{\Delta E \cdot B}}$$

Electrons travel on average ~15 cm each. Trajectories highly affected by multiple scattering. Electrons behave as MIPs except near the endpoints (*blobs*).

The SOFT concept



Readout
Plane B
- **energy**



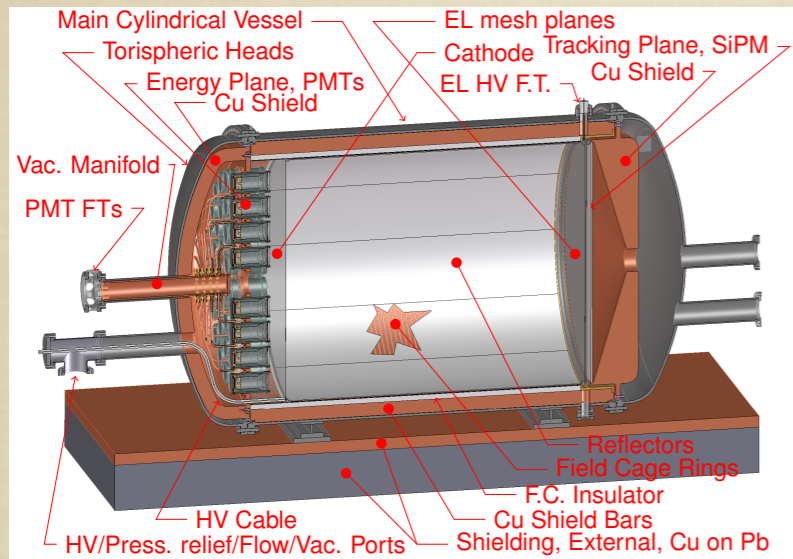
Electroluminescent
Layer

Readout
Plane A
- **position**

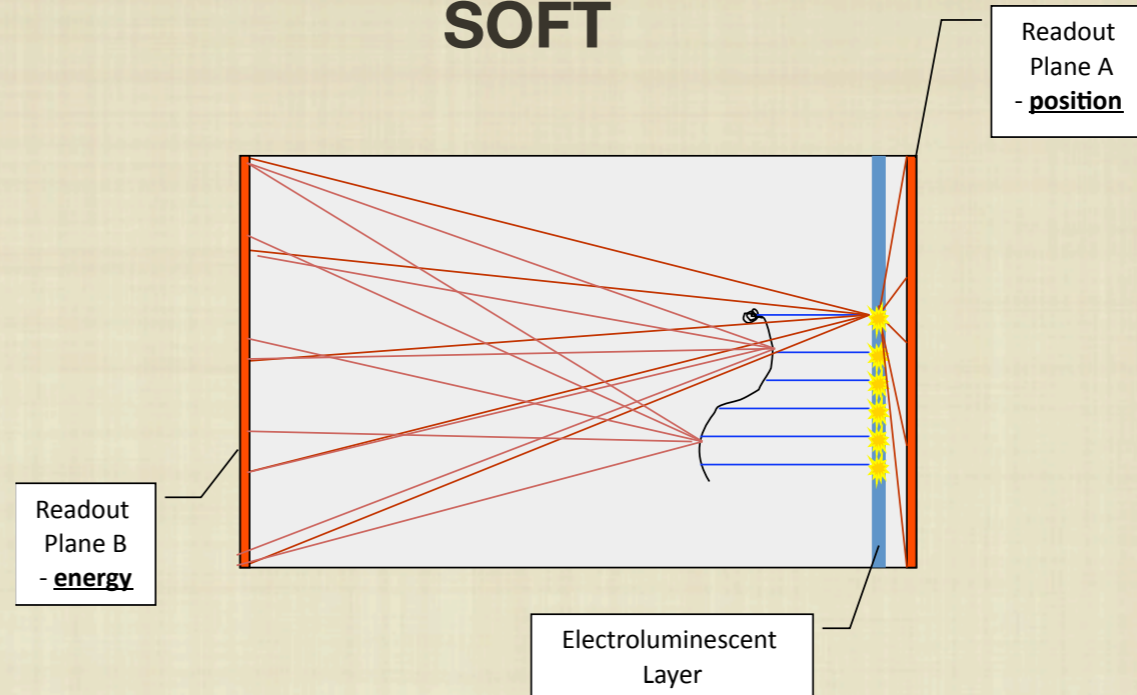


NEXT concept

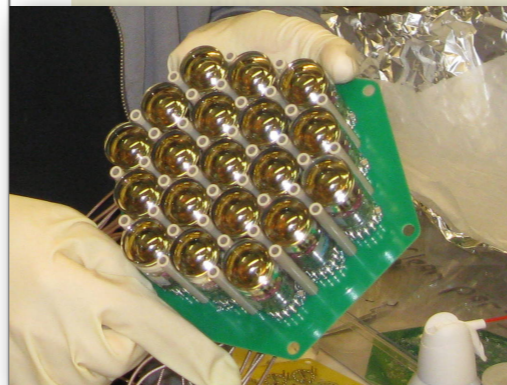
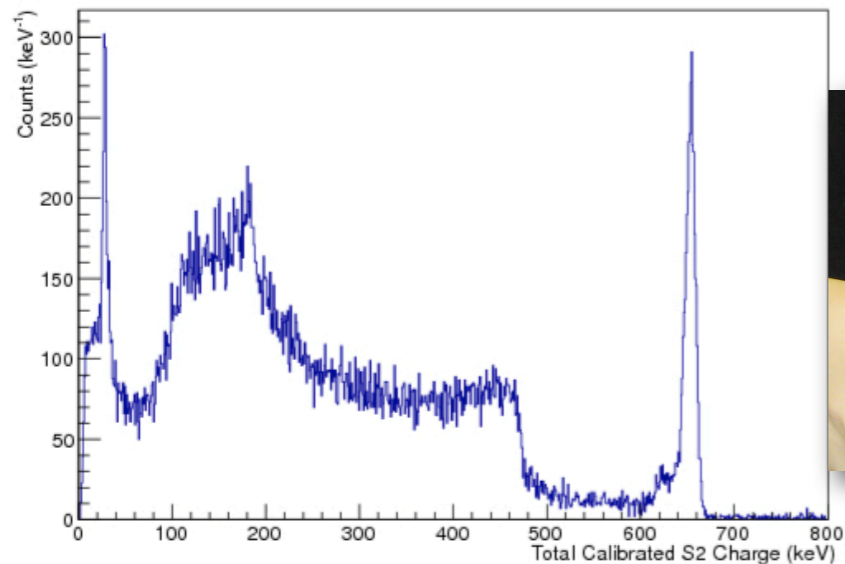
HPGXe



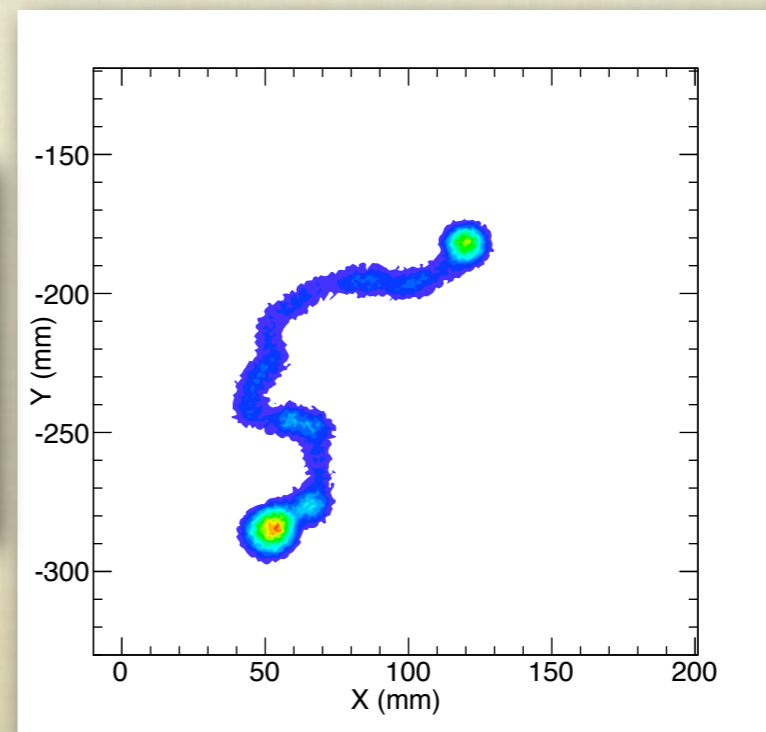
SOFT



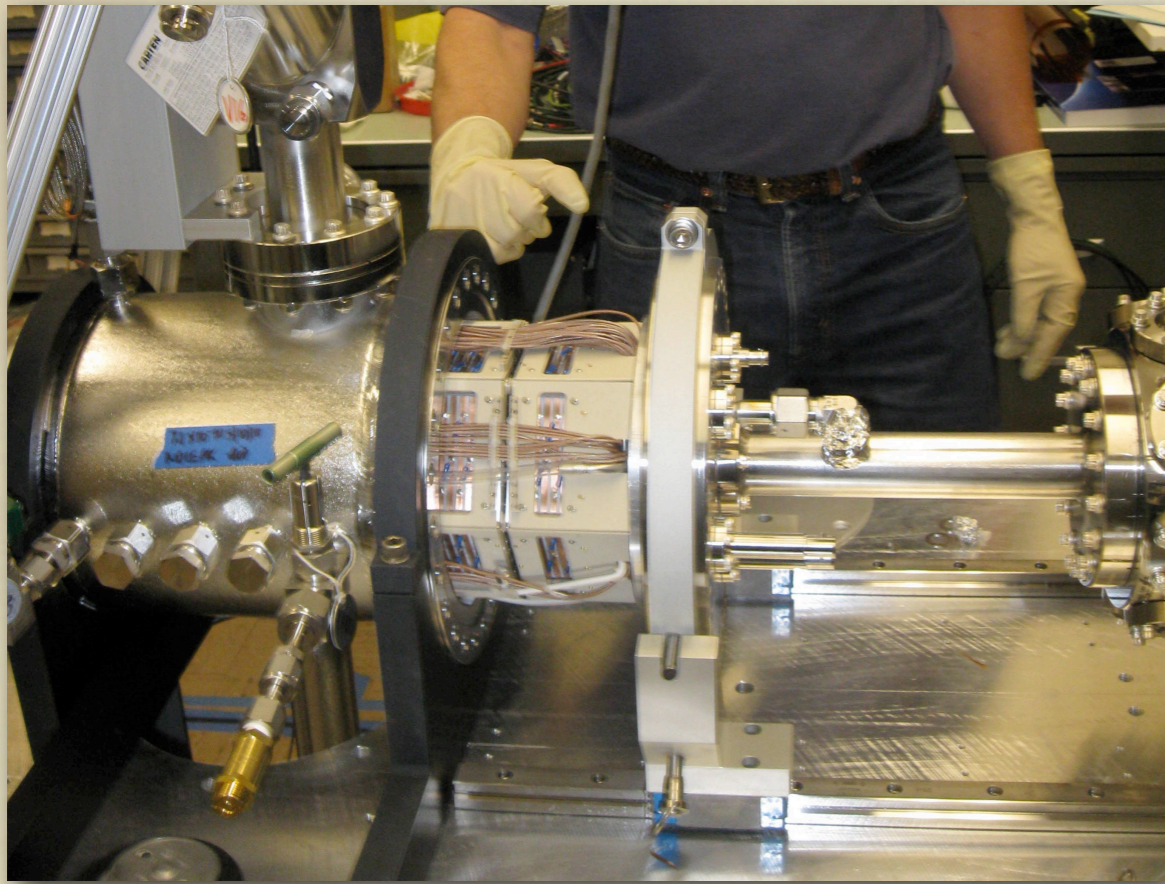
Energy resolution



Topological signature



NEXT-1 prototypes



NEXT-DBDM

Energy resolution in HPXe



NEXT-DEMO

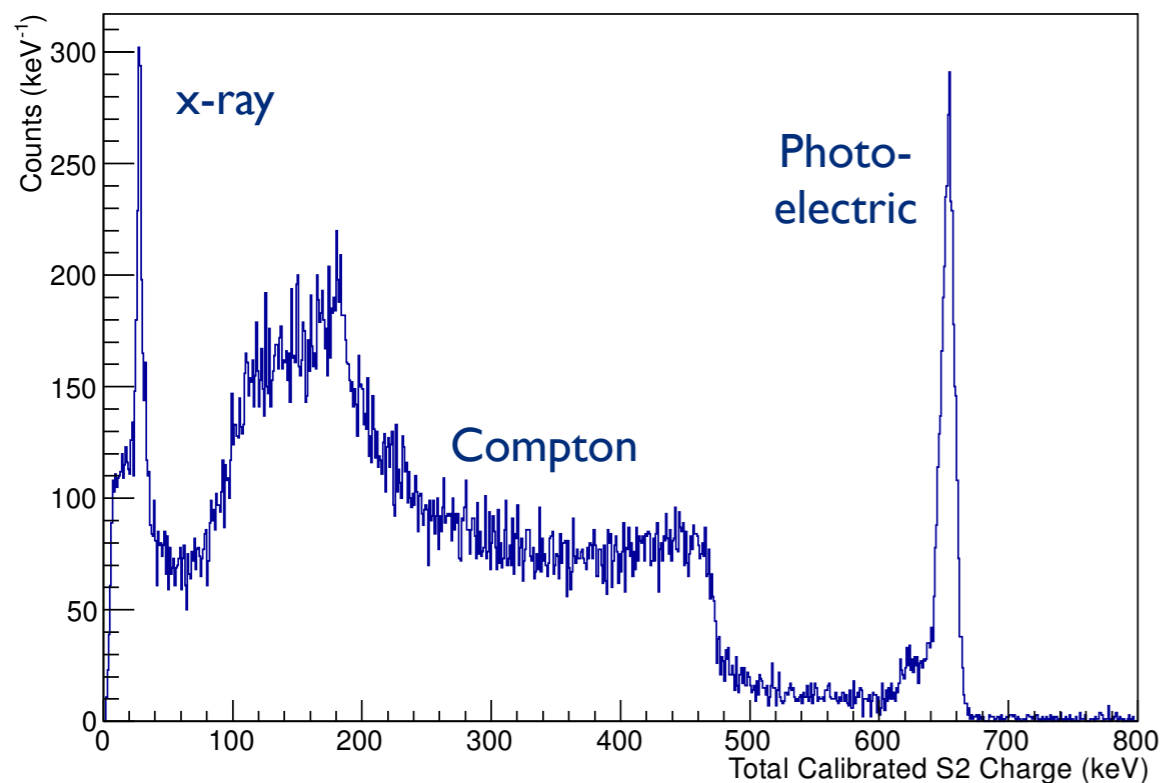
NEXT detector concept

Energy resolution demonstrated

- **Experience and results from prototypes**

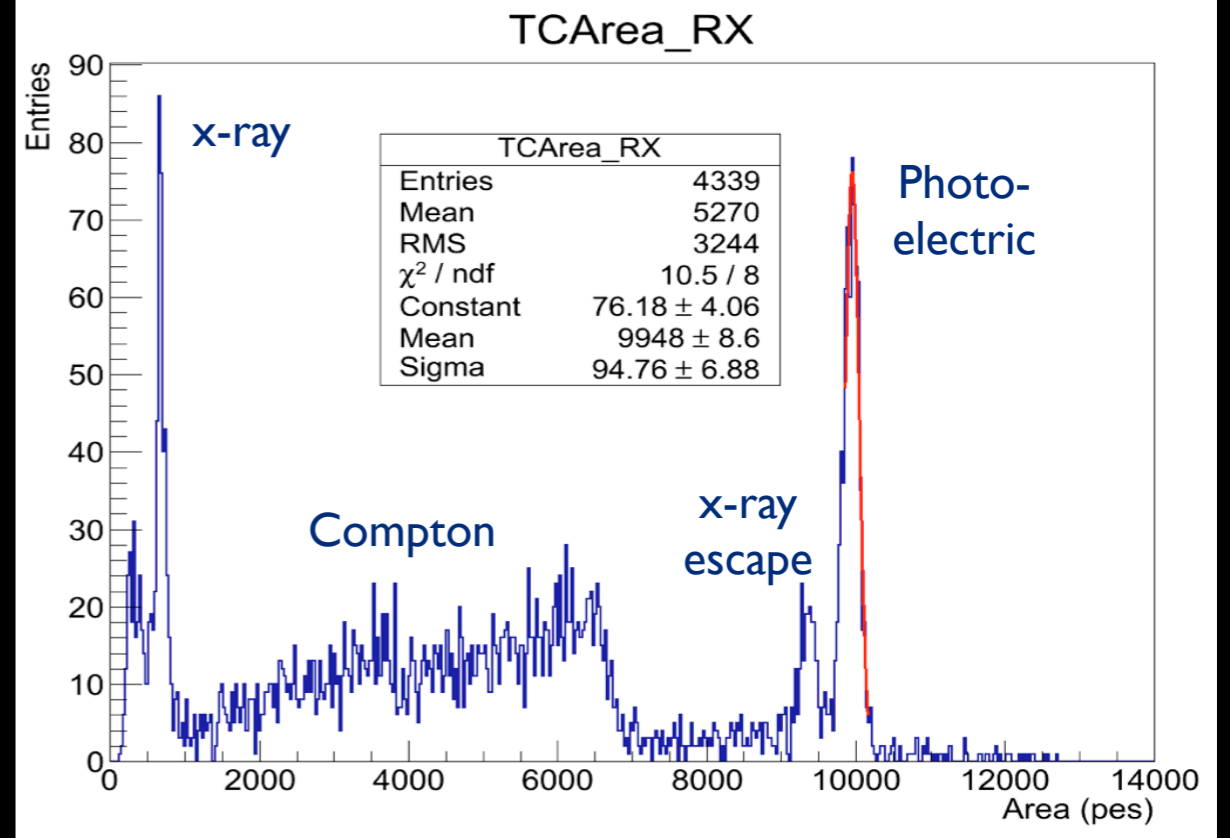
- Testing ground for all foreseeable technical hurdles in NEXT-100
- 0.5-1% FWHM energy resolution at $Q_{\beta\beta}$ demonstrated

Energy, 662 keV gammas from ^{137}Cs in NEXT-DBDM prototype



0.5% FWHM at $Q_{\beta\beta}$ in central region

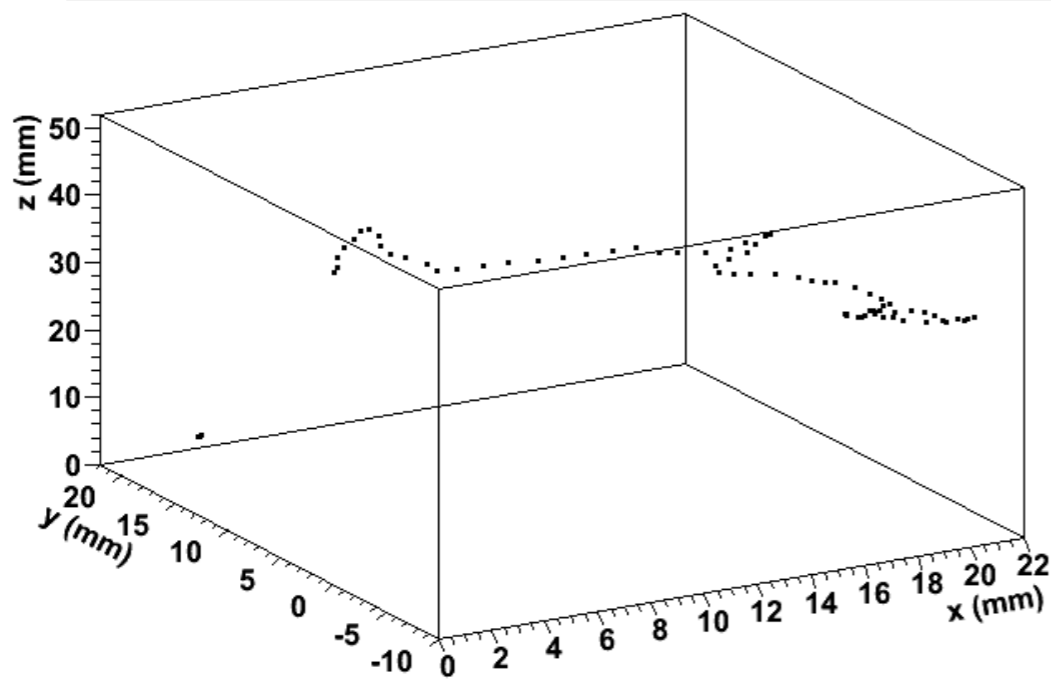
Energy, 511 keV gammas from ^{22}Na in NEXT-DEMO prototype



1% FWHM at $Q_{\beta\beta}$ in full fiducial

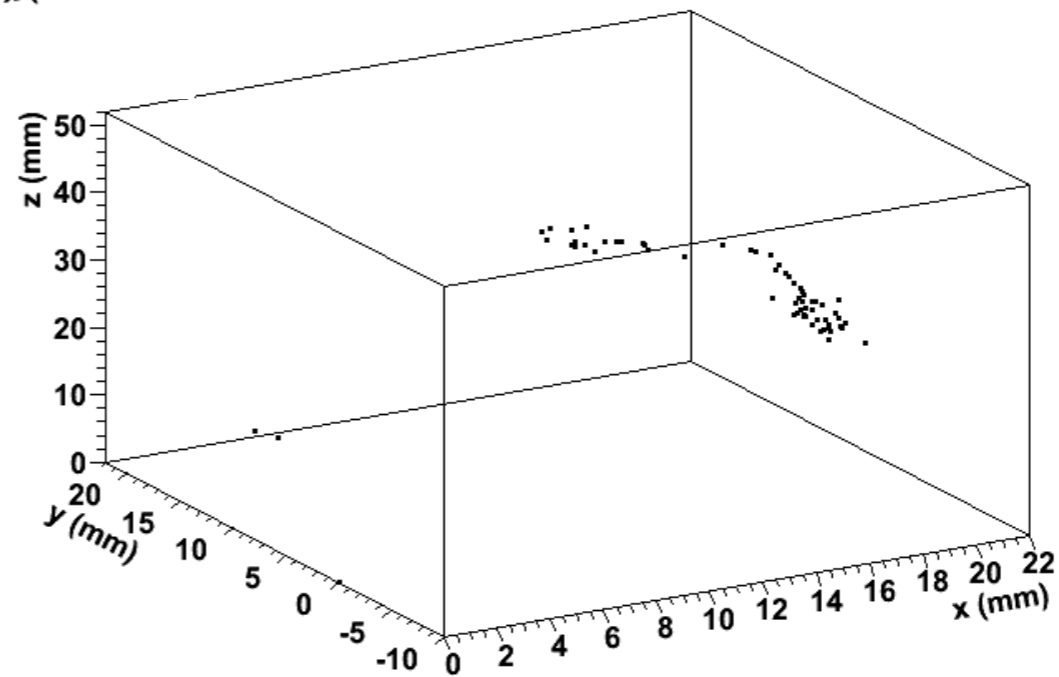
TRACK RECONSTRUCTION

Track reconstruction with SiPMs:



← MC Truth

Reconstructed →

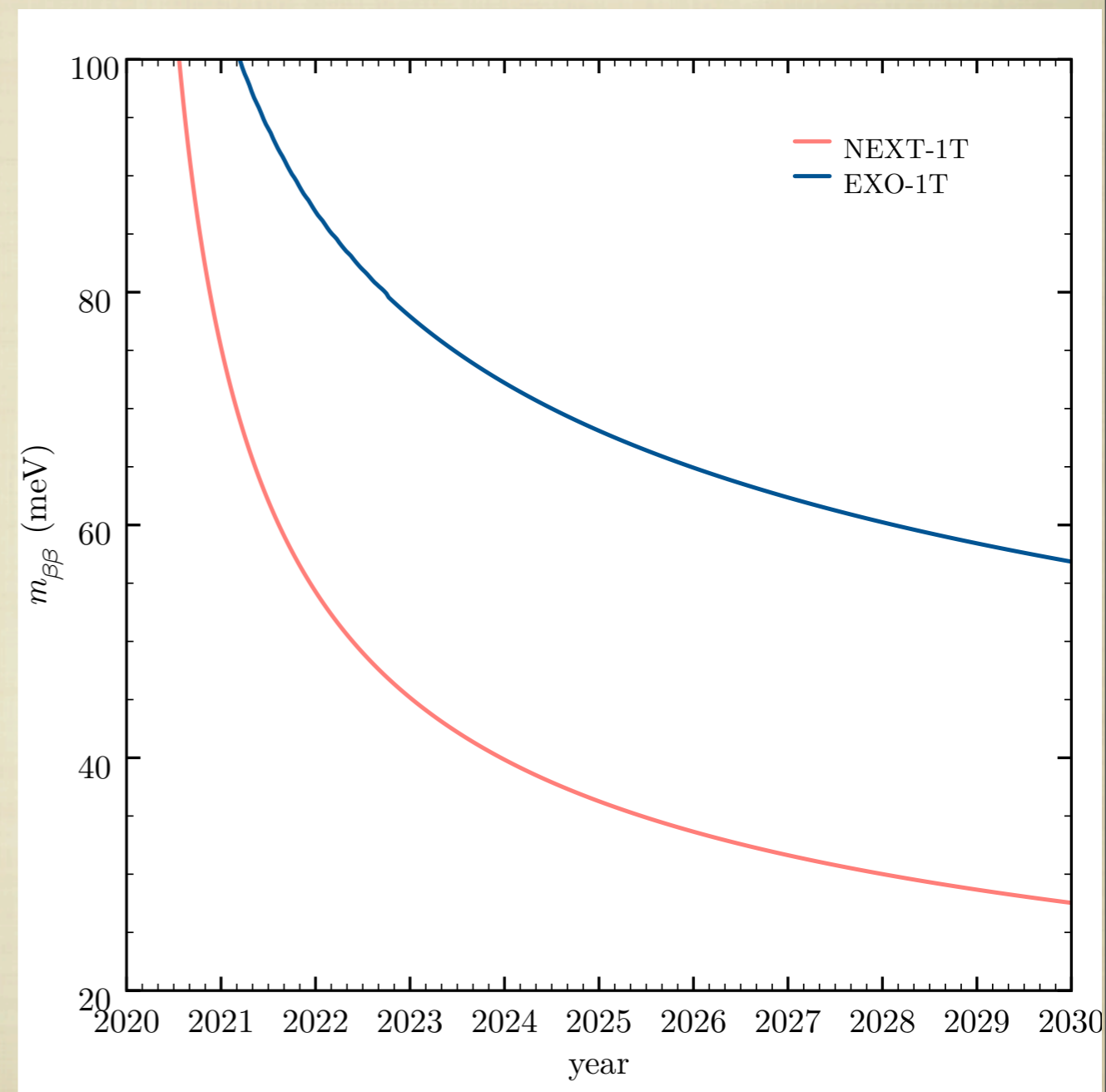
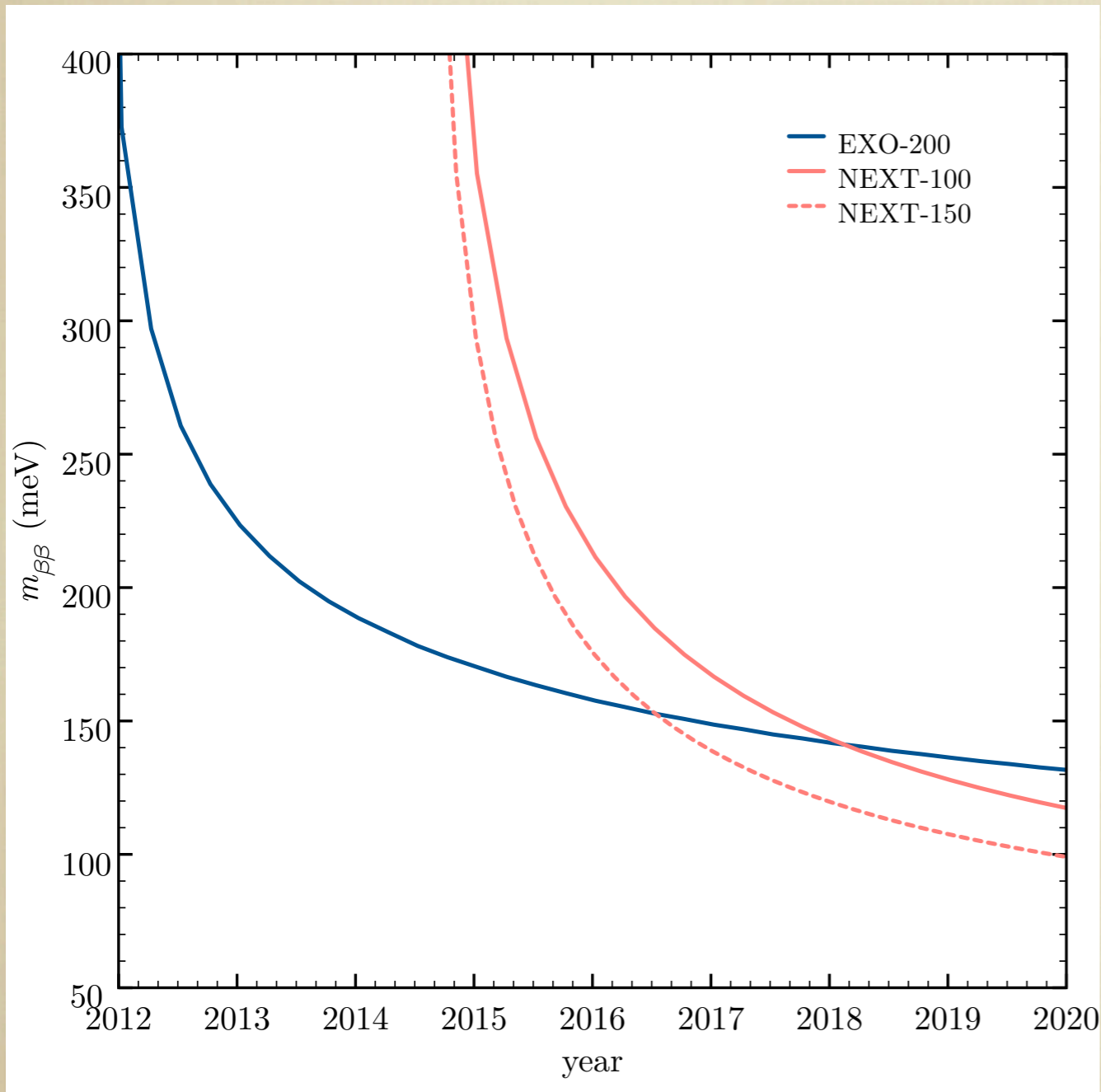


NEXT-100 performance

	Signal	^{214}Bi	^{208}Tl
1 track cut	0.48	6.0×10^{-5}	2.4×10^{-3}
ROI	0.33	2.2×10^{-6}	1.9×10^{-6}
Topological cut	0.25	1.9×10^{-7}	1.8×10^{-7}

Rejection Potential	$\sim 10^{-7}$
Background	2.0×10^{-4} counts/keV/kg/yr

NEXT vs EXO



Outlook

- A large detector (1 ton) capable of exploring periods $> 10^{25}$ y (inverse hierarchy $\sim 10^{26} - 10^{27}$ y) requires a homogenous cheap isotope. Xenon is a noble gas and the cheapest in the market, it has no radioactive isotopes, is a great calorimeter and in the gas phase has excellent resolution.
- There are two ways to reach the 1 ton mass, $T_{1/2} \sim 10^{26} - 10^{27}$ y, with manageable background (1 event per ton per year)
- LXe, if a way to kill all radioactive background is implemented (Ba⁺⁺ tagging, Graxe concept).
- HPGXe, taking advantage of excellent resolution and extra handle.

NEXT Collaboration



U. Girona • IFIC (Valencia) • U. Santiago de Compostela
• U. Polit cnica Valencia • U. Zaragoza



LBNL • Texas A&M



CEA (Saclay)



U. Coimbra • U. Aveiro



JINR (Dubna)



UAN (Bogot )

Spain provides:

- Most collaborators
- Most of secured funding
- Host laboratory (LSC)

Key contributions from international groups:

- TPC detector design
- Gaseous detectors
- Xe supply and enrichment

SPAIN: BEYOND SOCCER?

[HTTP://NEXT.IFIC.UV.ES/](http://next.ific.uv.es/)

Euro 2012

The next big thing

Jul 2nd 2012, 13:27 by P.L.

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