

# Current Status of LSND and Reactor Anomalies and Future Prospects for Sterile Neutrino Searches

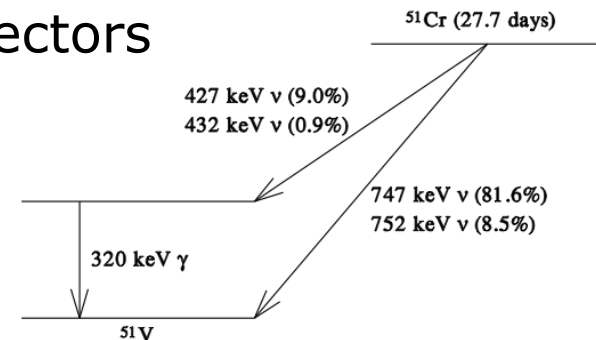
Steve Brice, Fermilab  
25 June 2012

## Current Indications of Tension in the Standard 3 Neutrino Mixing Scheme

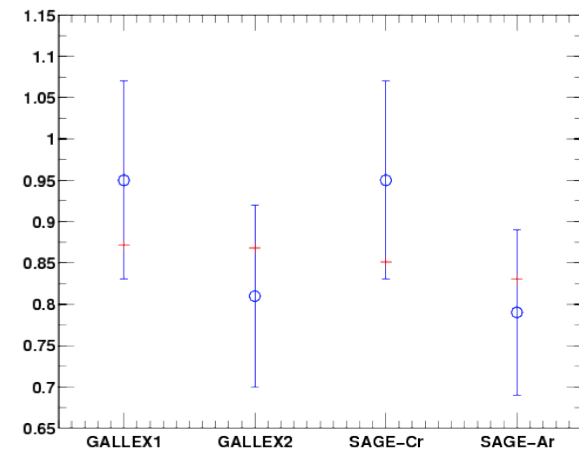
- Gallium:  $2.7\sigma$  evidence for  $\nu_e$  disappearance
- LSND:  $3.8\sigma$  evidence for anti- $\nu_e$  appearance
- MiniBooNE:  $3.8\sigma$  evidence for  $\nu_e$  and anti- $\nu_e$  appearance
- Reactor:  $3.0\sigma$  evidence for anti- $\nu_e$  disappearance
  
- Can be interpreted as a 4<sup>th</sup> neutrino state at eV scale mass
  
- Only 3 light, Weakly interacting neutrinos (LEP Z width)
- Oscillations with  $\Delta m^2_{\text{solar}}$  and  $\Delta m^2_{\text{atm}}$  are well established
- Therefore a 4<sup>th</sup> light state must be sterile
  
- Many thanks to K.Heeger, T. Lasserre, L.Huillier, C.Polly, M.Shaevitz for material

# Gallium Anomaly

- Calibration of the Gallium Solar  $\nu$  Detectors
- e-capture sources
  - $^{51}\text{Cr}$  (750 keV) &  $^{37}\text{Ar}$  (810keV)

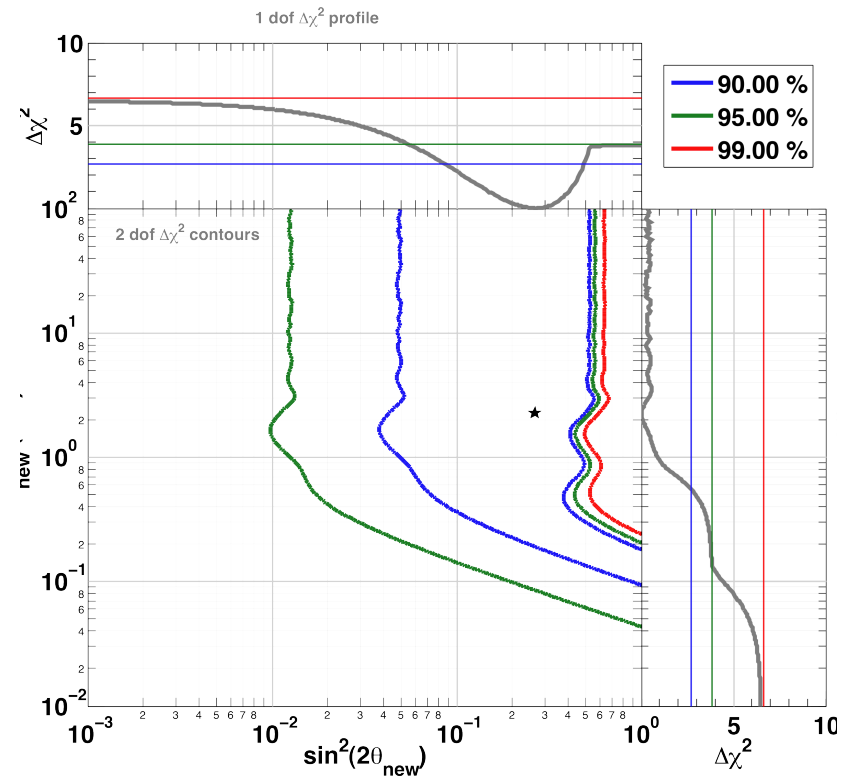


- The goal was to calibrate the production and extraction efficiency of the SAGE and GALLEX experiments
- Deficit observed
- $R_{\text{obs}/\text{pred}} = 0.86 \pm 0.05$  ( $\sigma_{\text{Bahcall}}$ )
- $R_{\text{obs}/\text{pred}} = 0.76 \pm 0.085$  ( $\sigma_{\text{Haxton}}$ )



# Gallium Anomaly

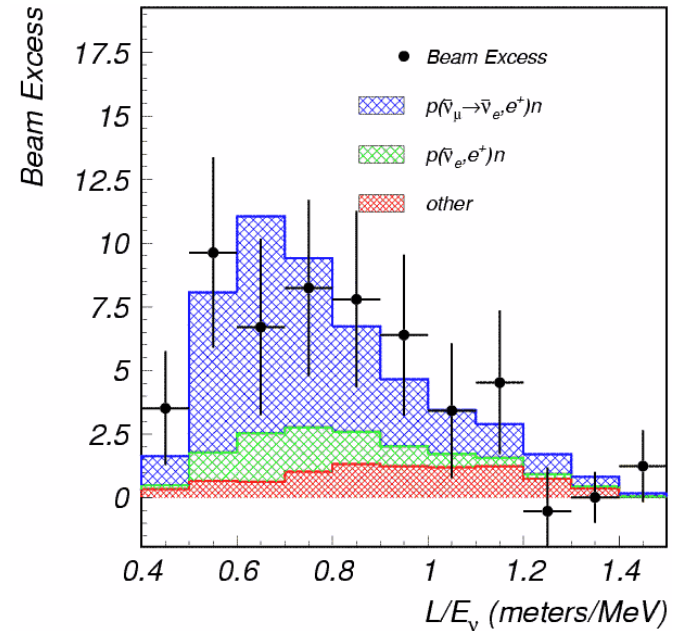
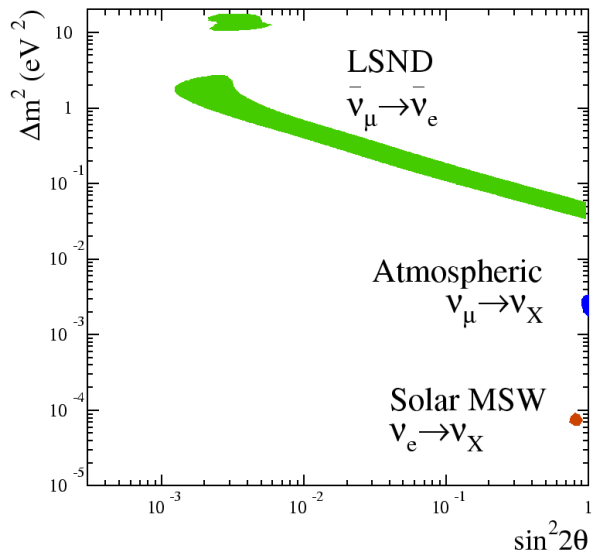
- No-oscillation hypothesis disfavored at  $2.7\sigma$  (PRC 83 065504,2011)
- Was not treated as evidence for new physics until the other anomalies appeared



Plot from Th.Lasserre

# LSND

- Used 800 MeV protons from LAMPF at Los Alamos in the 1990's
- Searched for anti- $\nu_e$  appearance in neutrino beam from pion decay at rest.

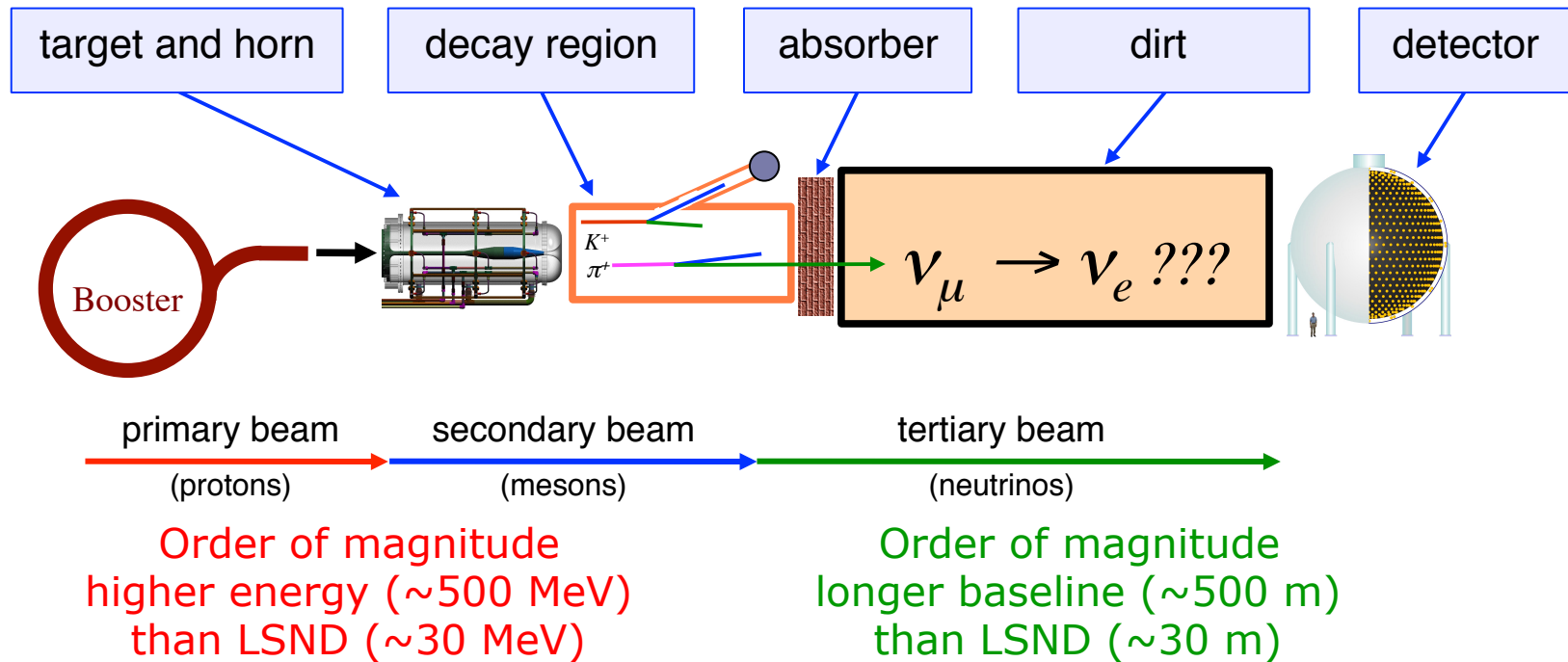


- Found an excess of anti- $\nu_e$  over background prediction
  - $87.9 \pm 22.4 \pm 6.0$  ( $3.8\sigma$ )

# MiniBooNE

Keep L/E same as LSND  
while changing systematics, energy & event signature

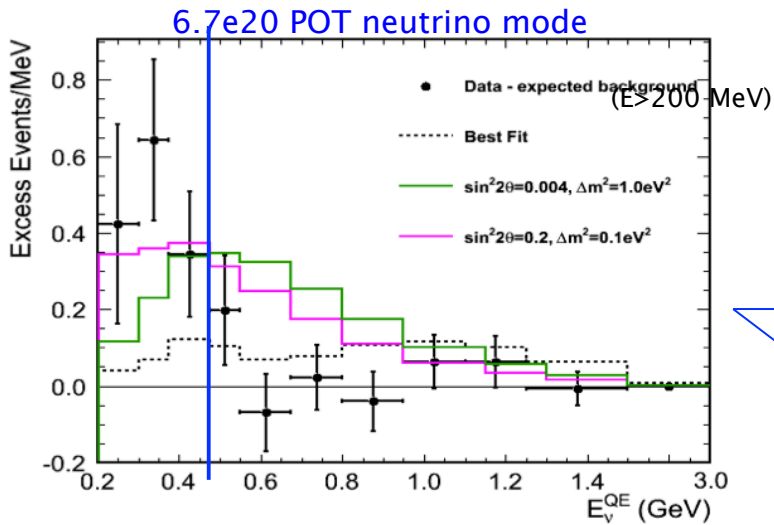
$$P(\nu_{\mu} \rightarrow \nu_e) = \sin^2 2\theta \sin^2(1.27 \Delta m^2 L/E)$$



# MiniBooNE Update

- $6.6 \times 10^{20}$  POT in neutrino mode
  - No neutrino mode data added
- $11.3 \times 10^{20}$  POT in anti-neutrino mode
  - Results updated in Kyoto with double the POT previously published
- Modest Improvements to the (anti-)neutrino analysis
  - *In situ* measurement of WS contamination in anti- $\nu$  beam  
[Phys.Rev.D84,072005 \(2011\)](#)
  - New SciBooNE constraint on intrinsic  $\nu_e$  from K+
  - Added error matrix for intrinsic  $\nu_e$  from K-
  - Improved smoothing algorithm used to assess systematics due to discriminator thresholds and PMT response
  - CC $\pi^+$  events (bkg for  $\nu_\mu$  CCQE when  $\pi^+$  is absorbed)  $Q^2$  reweighting applied based on internal MB measurement  
[Phys.Rev.D83,052007 \(2011\)](#)

# Neutrino mode 3+1 fits with all updates

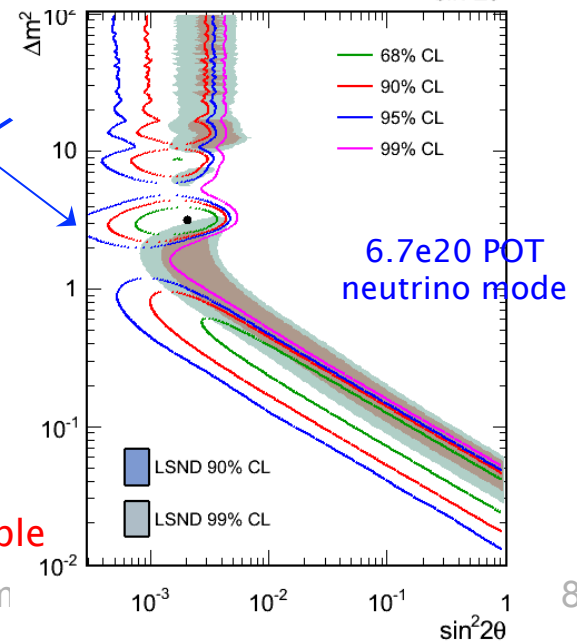
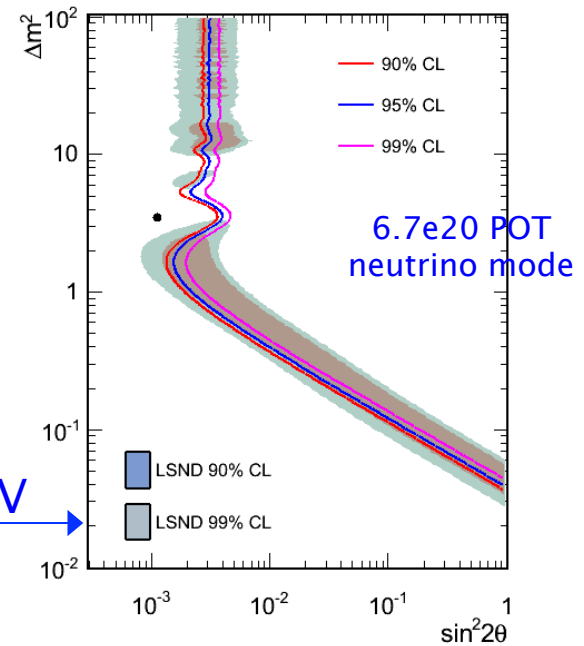


ν mode	E > 200 MeV	E > 475 MeV
χ <sup>2</sup> (null)	22.81	6.35
Prob(null)	0.5%	36.6%
χ <sup>2</sup> (bf)	13.24	3.73
Prob(bf)	6.12%	42.0%

Overall probability of bf 6%...not great, but also not terrible

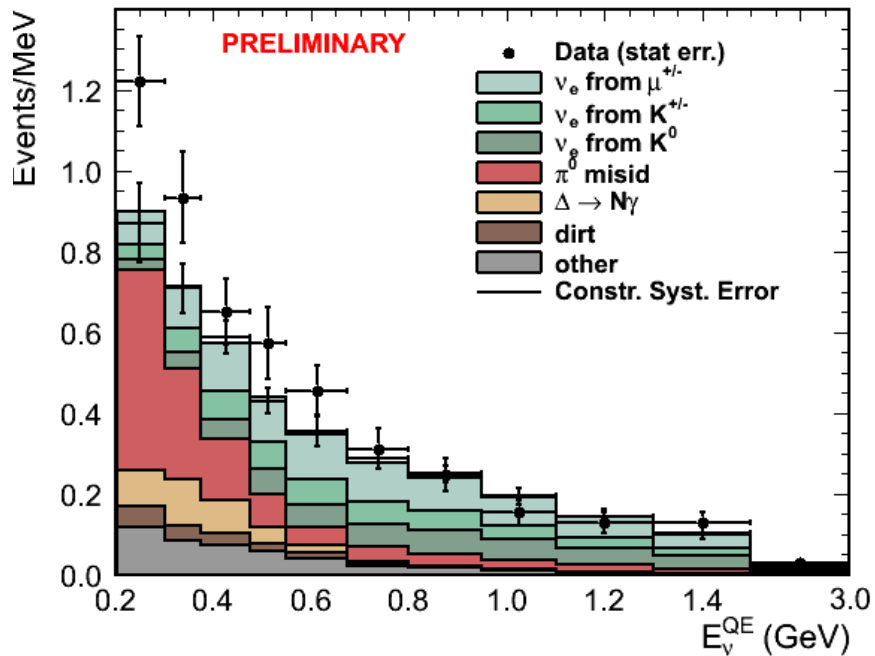
Fit E > 475 MeV

Fit E > 200 MeV





# Full Anti-neutrino dataset



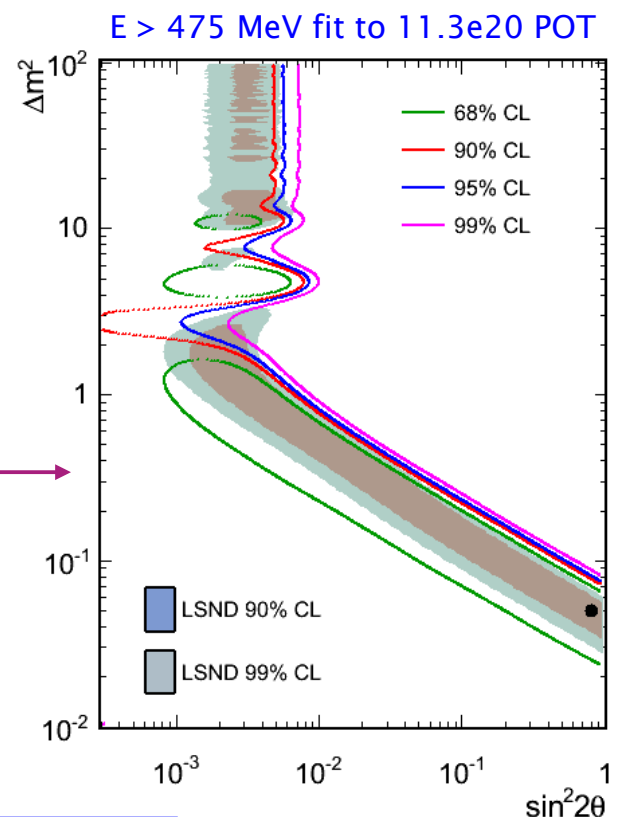
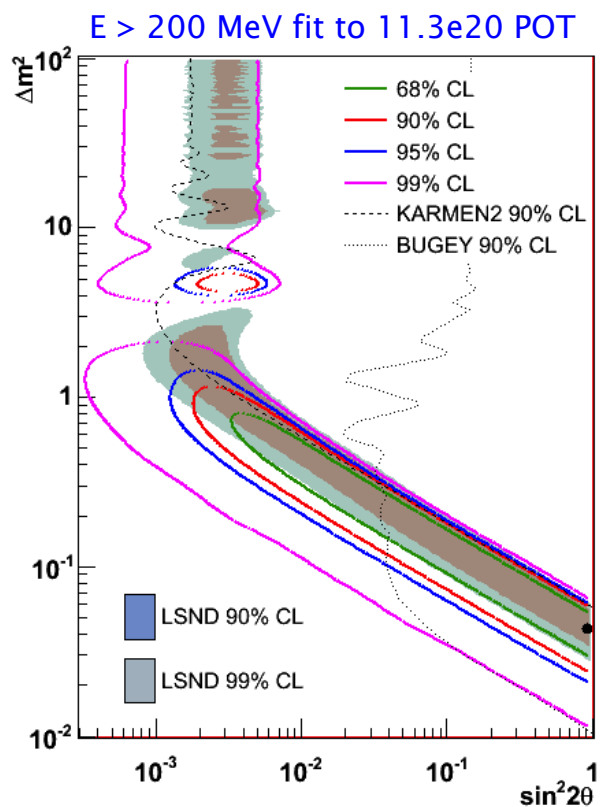
anti- $\nu_e$  CCQE signal candidates  
with 1.3e20 POT

Higher stat anti-neutrino data is now much more consistent with what was observed in the data taken with a neutrino beam

\* Systematic error after all other data constraints applied, e.g.  $\nu_\mu$  CCQE, NC  $\pi^0$ , dirt events, SciBooNE  $K^+$

$E_\nu(QE)$ range	Data	Bkg $\pm$ stat $\pm$ syst*	Excess
200-475 MeV	257	$199.1 \pm 14.1 \pm 16.3$	$57.9 \pm 21.6 (2.7\sigma)$
475-1250 MeV	221	$201.1 \pm 14.2 \pm 17.9$	$19.9 \pm 22.8 (0.9\sigma)$
200-1250 MeV	478	$400.2 \pm 20.0 \pm 23.4$	$77.8 \pm 30.8 (2.5\sigma)$

# Fitting anti-neutrino data to 2ν model

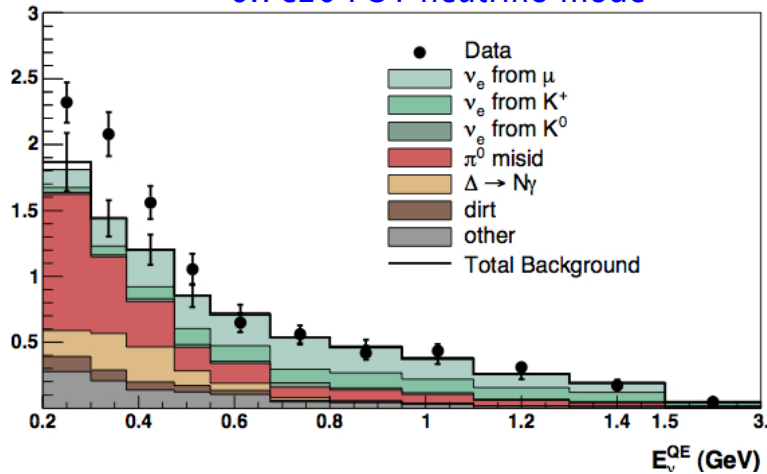


No tension

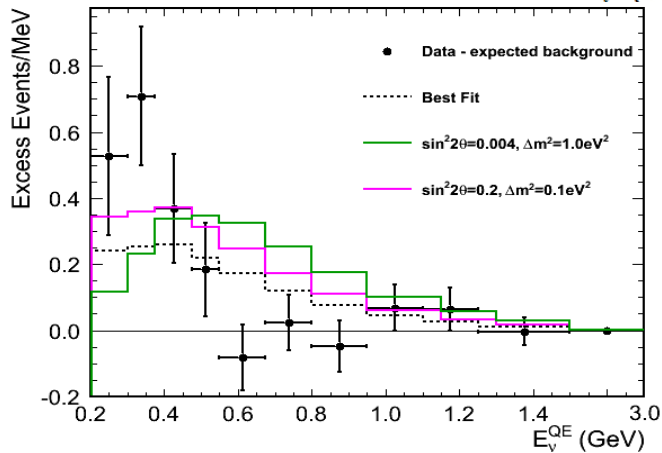
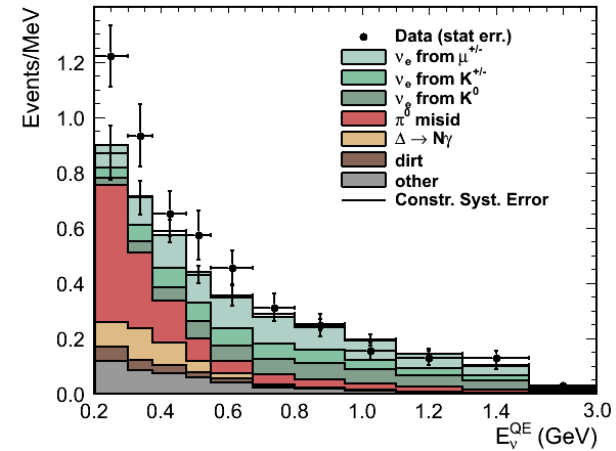
anti-ν mode	E > 200 MeV	E > 475 MeV
$\chi^2(\text{null})$	16.3	7.59
Prob(null)	5.8%	26.4%
$\chi^2(\text{bf})$	4.76	3.23
Prob(bf)	67.5%	50.2%

# Comparing neutrino to anti-neutrino mode

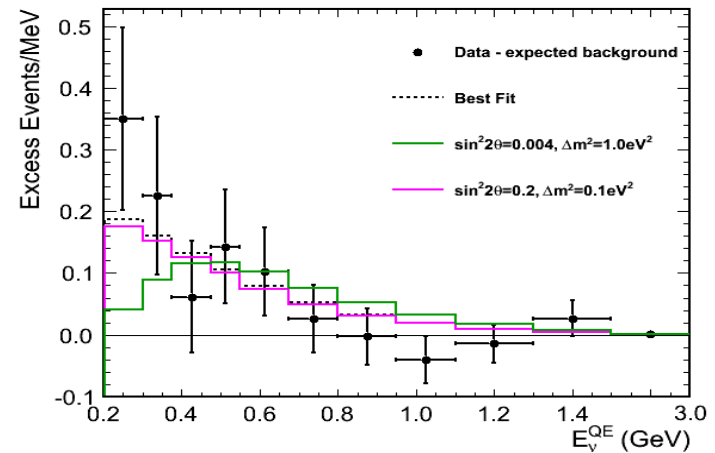
6.7e20 POT neutrino mode



11.3e20 POT anti-neutrino mode

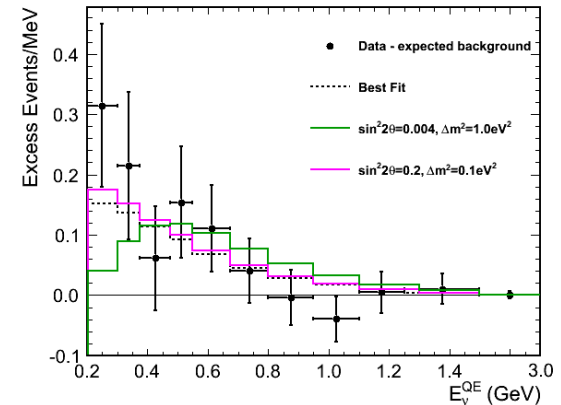
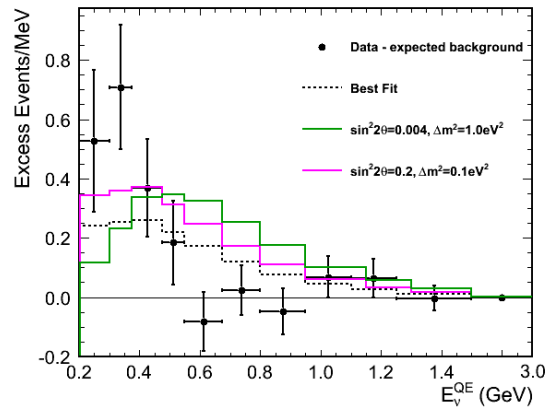
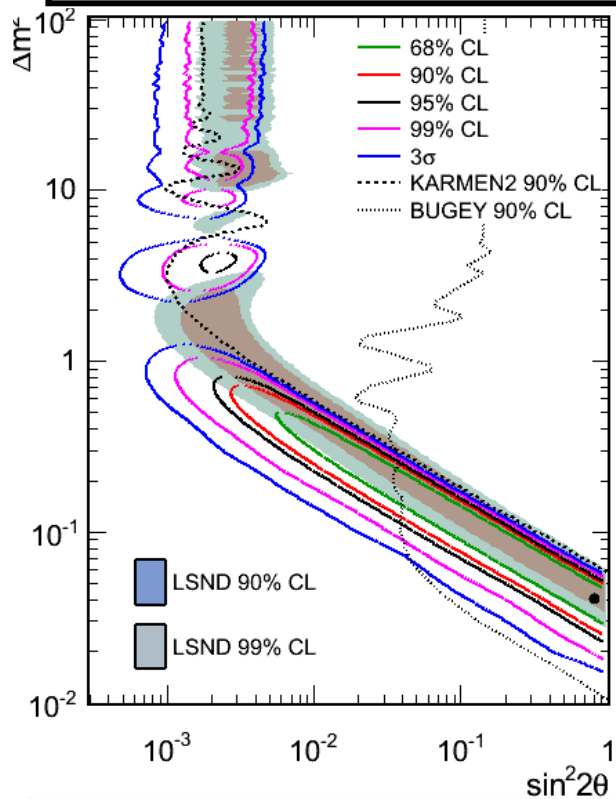


Excess:  $146.3 \pm 28.4 \pm 40.2$



Excess:  $77.8 \pm 20.0 \pm 23.4$

# Simultaneous 3+1 fit to $\nu$ and anti- $\nu$ data



Total Excess:  $240.3 \pm 34.5 \pm 52.6$

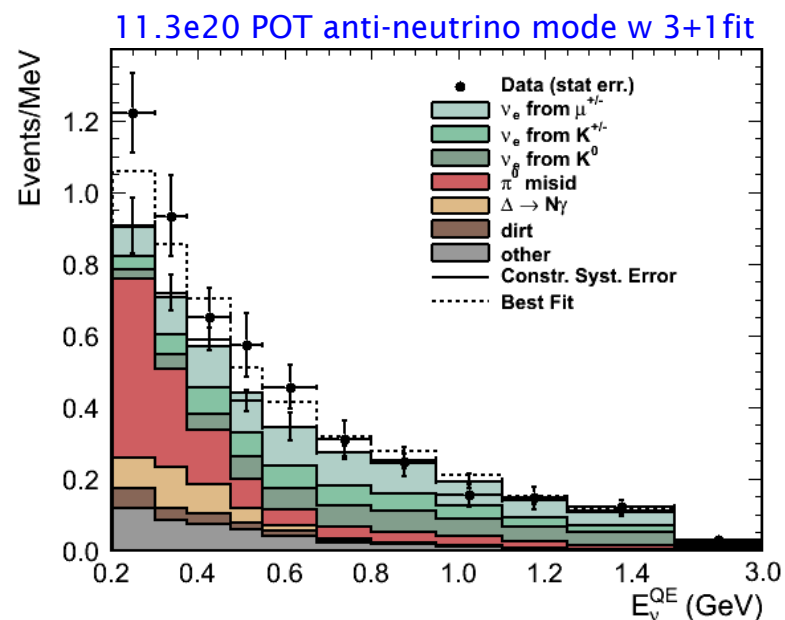
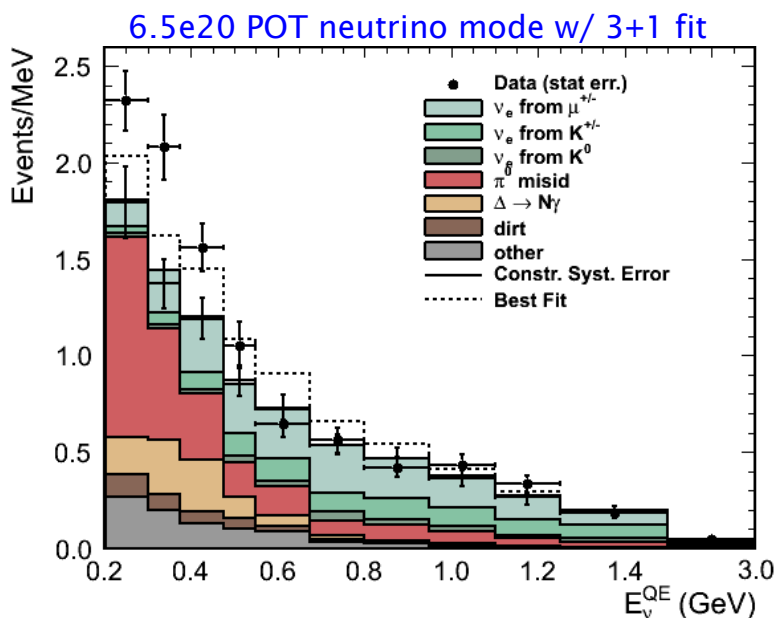
- WS accounted for properly
- Construction of correlated systematic error matrix
- $E > 200$  MeV BF preferred at  $3.6\sigma$  over null

combined	$E > 200$ MeV	$E > 475$ MeV
$\chi^2(\text{null})$	42.53	12.87
Prob(null)	0.1%	35.8%
$\chi^2(\text{bf})$	24.72	10.67
Prob(bf)	6.7%	35.8%

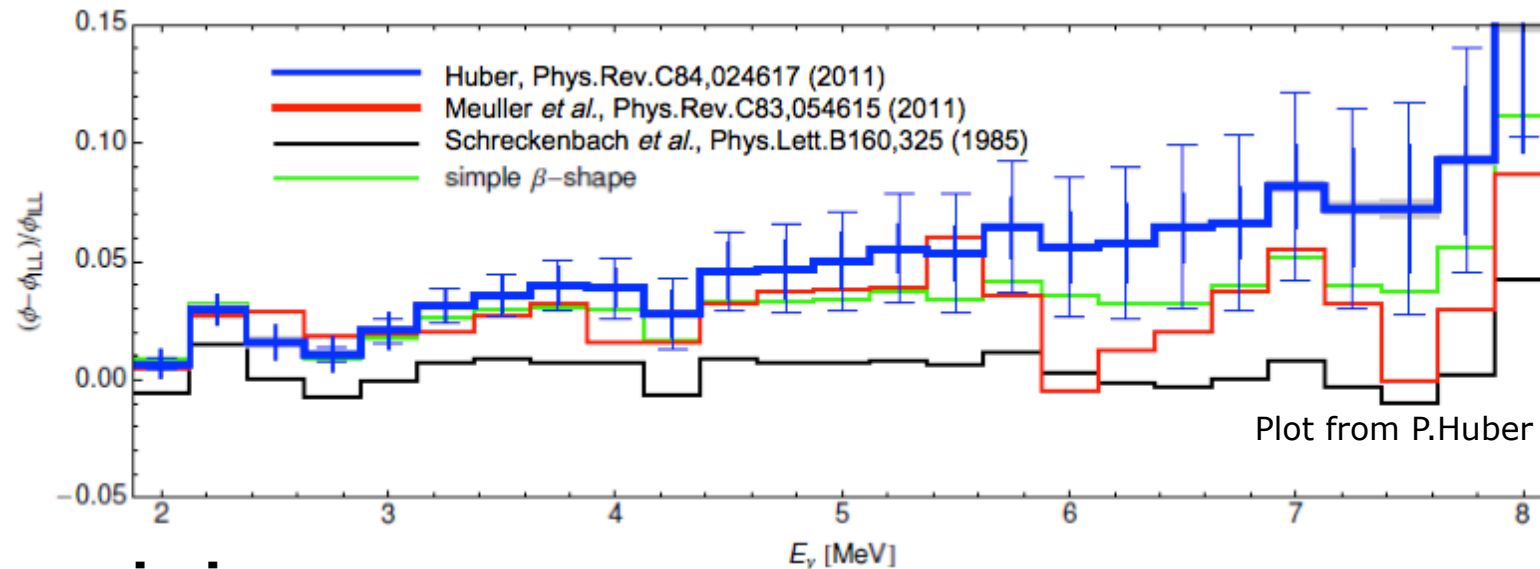
\* Simultaneous fit ( $E > 200$  MeV) with fully-correlated systematic to entire MB neutrino and anti-neutrino data

# MiniBooNE Conclusions

- MiniBooNE observes an excess of  $\nu_e$  candidates in the 200-1250 MeV energy range in neutrino mode ( $3.0\sigma$ ) and in anti-neutrino mode ( $2.5\sigma$ ). The combined excess is  $240.3 \pm 34.5 \pm 52.6$  events ( $3.8\sigma$ )
- The event excess is concentrated in the 200-475 MeV region where NC  $\pi^0$  and other processes leading to a single  $\gamma$  dominate
- Higher statistics anti- $\nu$  data is now similar to the neutrino mode data
- It is not yet known whether the MiniBooNE excesses are due to oscillations, some unrecognized NC  $\gamma$  background, or something else



# Reactor Neutrino Anomaly



## $\nu$ emission:

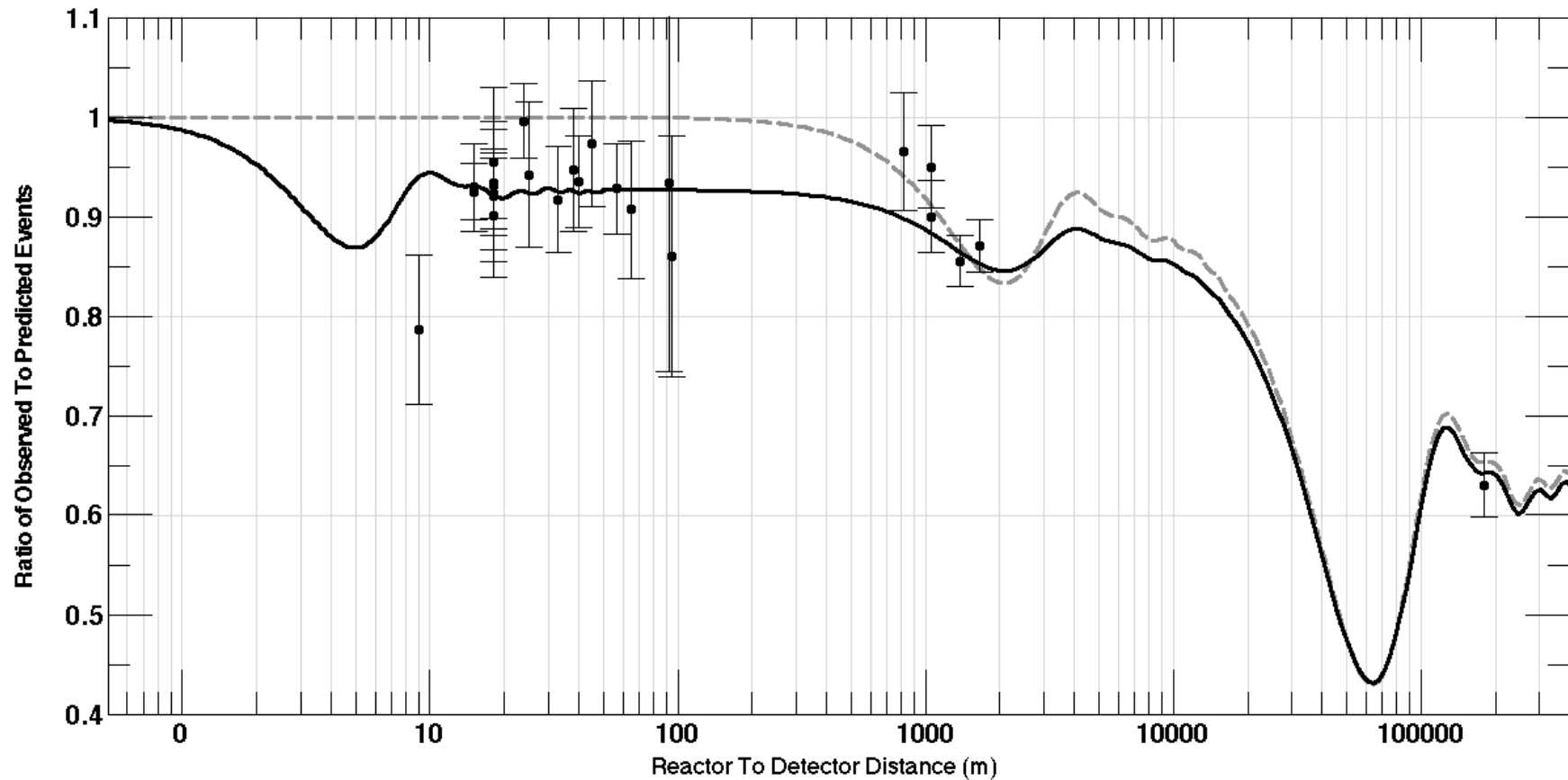
- Improved reactor neutrino spectra produces +3.5%
- Accounting for long-lived isotopes accumulating in reactors produces +1%
- PRC83, 054615 (2011)
- PRC84, 024617 (2011)

## $\nu$ detection:

- Reevaluation of  $\sigma_{\text{IBD}}$  Improved neutron life time measurements produces +1%

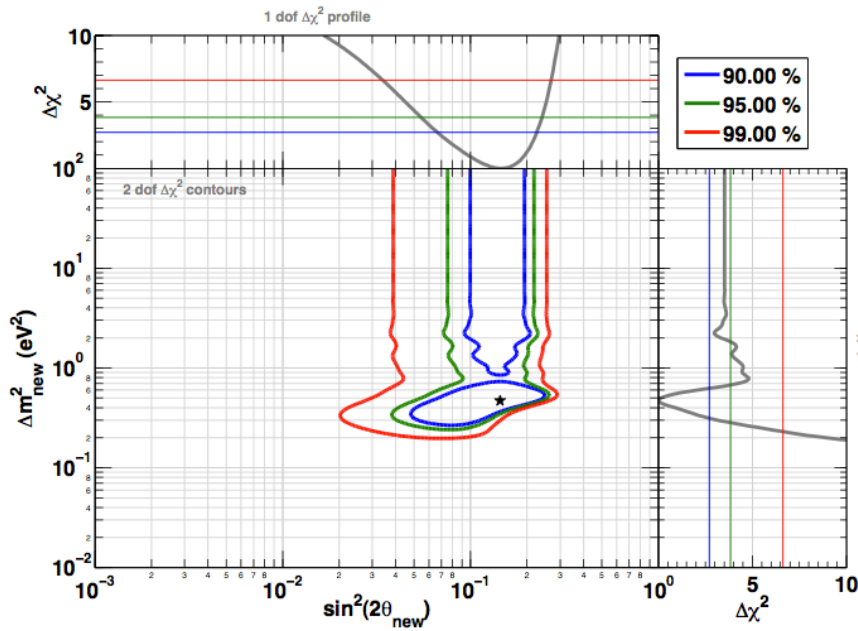
Observed/predicted averaged event ratio:  $R=0.927\pm0.023$  ( $3.0 \sigma$ )

# Interpreted as Oscillation with 4th State

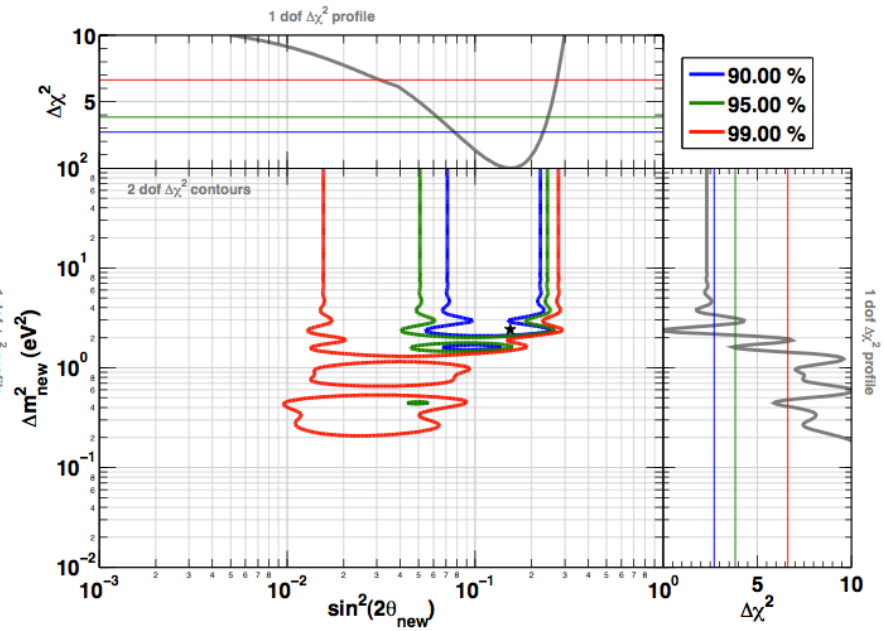


# Interpreted as Oscillation with 4th State

## Rate Analysis



## Rate + Shape Analysis



Plot from G.Mention et al arXiv:1101.2755



## Counter Evidence for 4<sup>th</sup> State

There are a number of results that are sensitive, but see no evidence for a 4<sup>th</sup> neutrino state with  $\sim eV$  mass:-

- CDHS and MiniBooNE searches for  $\nu_\mu$  disappearance
- MiniBooNE search for  $\bar{\nu}_\mu$  disappearance
- MINOS search for  $\nu_\mu \rightarrow \nu_s$
- Karmen search for  $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$

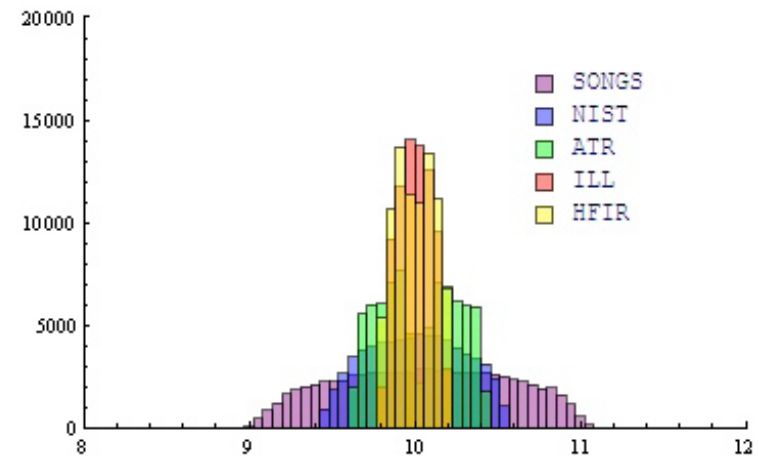
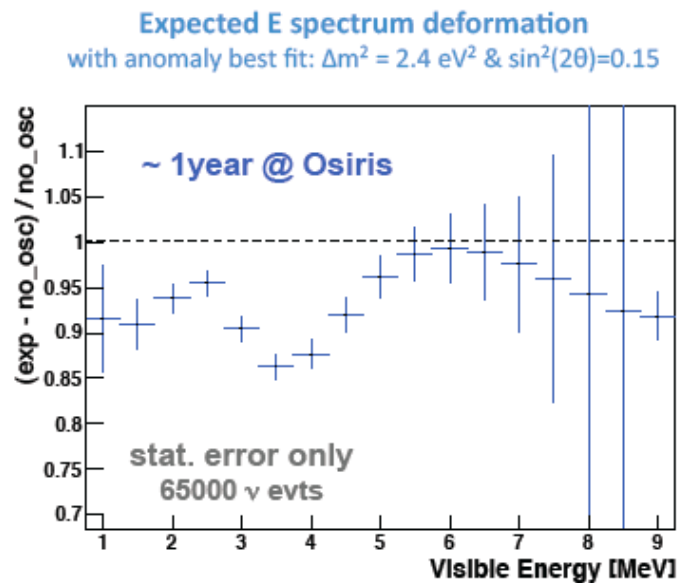
It is hard (impossible?) to fit all data with a single oscillation hypothesis

# Future Tests

- Need a definitive test(s) of the 4<sup>th</sup> neutrino hypothesis hinted at by the current anomalies
- Many tests proposed. They fall into three types:-
  - 1) Detector <15 m from compact nuclear reactor
  - 2) Accelerator based short baseline
  - 3) Intense sources close to or in detector
- For definitive test would like oscillation evidence in E and L and redundant cross-checks
- See Sterile Neutrino White Paper
  - arXiv:1204.5379
- Upcoming report from Fermilab Short Baseline Working Group

# Reactor Searches for 4<sup>th</sup> State

- Get a ton scale detector 2-10m from a compact core



- Look for depleted rate and altered shape

# Proposed Reactor Short Baseline Experiments

Proposal	Reactor	Fuel (#fissions)	Core Size (m)	<L> (m)	Depth (mwe)	Status	Comment
Nucifer Saclay	Osiris 70 MW	<sup>235</sup> U ON-Off cycle	<1	7	5	Data Taking	Non proliferation 1 m <sup>3</sup> Gd-LS Mostly Rate + Shape?
Stereo Genoble	ILL 50 MW	<sup>235</sup> U ON-Off cycle	<1	10	10	Proposal	2 m <sup>3</sup> Gd-LS Rate + Mostly shape
SCRAMM (Ca)	San-Onofre 3 GW PWR	<sup>235,238</sup> U <sup>239,241</sup> Pu	3x3.8	24	30	Proposal	2 m <sup>3</sup> Gd-LS Mostly Rate + Shape
SCRAMM (Idaho)	ATR 150 MW	<sup>235</sup> U ON-Off cycle	<1	12	15	Proposal	2 m <sup>3</sup> Gd-LS Rate + Mostly shape
DANSS (Russia)	KNPP 3 GW PWR	<sup>235,238</sup> U <sup>239,241</sup> Pu	few	14	70	Being Built	Segmented detector 1 m <sup>3</sup> Rate + Shape?
NIST (US)	NCNR 20 MW	<sup>235</sup> U ON-Off cycle	≈1	4-11	0	Proposal	Rate + Mostly shape
Nu4 (Russia)	SM-3 100 MW	<sup>235</sup> U ON-Off cycle	0.35x0.42 x0.42	6-12	10	Being Built	14 m <sup>3</sup> Gd-LS Rate + shape

Table from T.Lasserre

# Accelerator Based Short Baseline Search

- Wish (Requirements) list:-
  - Need significance at  $>5\sigma$  level
  - Would like to see effect in L and E
  - Would like to have redundant crosschecks within an experiment
  - Would like to see a consistent picture with appearance and disappearance
- Four experiment types:-
  - Accelerator isotope production with large detector close by
  - Pion/Kaon decay at rest (C.F. LSND, Karmen)
  - Pion decay in flight (C.F. MiniBooNE)
  - Low energy neutrino factory

# Accelerator Short Baseline Experiments

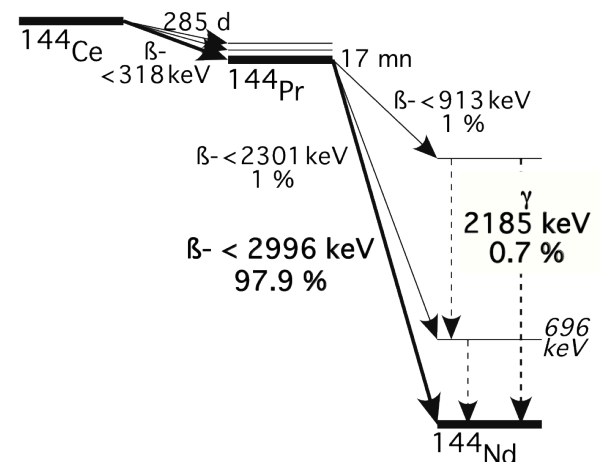
Isotope Source	Disapp	$\bar{\nu}_e \rightarrow \bar{\nu}_e$	IsoDAR
Pion / Kaon Decay-at-Rest Source	Appearance & Disapp	$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_e \rightarrow \nu_e$	OscSNS, CLEAR, DAEδALUS, KDAR
Accelerator $\bar{\nu}$ using Pion Decay-in-Flight	Appearance & Disapp	$\nu_\mu \rightarrow \nu_e, \bar{\nu}_\mu \rightarrow \bar{\nu}_e$ $\nu_\mu \rightarrow \nu_\mu, \nu_e \rightarrow \nu_e$	MINOS+, MicroBooNE, LAr1kton+MicroBooNE, CERN SPS
Low-Energy $\nu$ -Factory	Appearance & Disapp	$\nu_e \rightarrow \nu_\mu, \bar{\nu}_e \rightarrow \bar{\nu}_\mu$ $\nu_\mu \rightarrow \nu_\mu, \nu_e \rightarrow \nu_e$	$\nu$ STORM at Fermilab

Table from M.Shaevitz

# Intense Radioactive Sources

- Test 4<sup>th</sup> state hypothesis with  $\sim$ MeV (anti-)neutrino sources placed a few meters from large low background detector
  - Similar to the  $^{51}\text{Cr}$  calibrations of the SAGE and GALLEX solar  $\nu$  detectors
  - Can use existing reactor/solar neutrino detectors
  - Can place source inside or just outside detector
- Can search for effect on energy spectrum and rate as a function of distance from source.

- Typically need compact MCi source
  - Technically non-trivial
  - e.g.  $^{144}\text{Ce}$



# Proposals with Intense Radioactive Sources

Type	channel	Background	Source	Production	Activity (Mci)		Proposal
$\nu_e$	$\nu_e e \rightarrow \nu_e e$  Compton edge	radioactivity (managable)	$^{51}\text{Cr}$ 0.75 MeV $t_{1/2}=26\text{d}$	$n_{\text{th}}$ irradiation in Reactor	in	>3	Sage LENS
		Solar $\nu$ (irreducible)			out	5-10	SOX SNO+
	5% $E_{\text{res}}$ 15cm $R_{\text{res}}$	$\nu$ -Source (out ok but in ?)	$^{37}\text{Ar}$ 0.8 MeV $t_{1/2}=35\text{d}$	$n_{\text{fast}}$ irradiation in Reactor (breeder)	in	>1	-
					out	5	Ricochet (NC)
$\bar{\nu}_e$	$\bar{\nu}_e p \rightarrow e^+ n$  $E_{\text{th}}=1.8 \text{ MeV}$  ( $e^+, n$ ) Coincidence	reactor $\nu$ & $\nu$ -Source	$^{144}\text{Ce}$ $E < 3\text{MeV}$ $t_{1/2}=285\text{d}$	spent nuclear fuel reprocessing	in	0.005-0.05	CeLAND SOX
		$\rightarrow$			out	0.5	Daya-Bay
	5% $E_{\text{res}}$ 15cm $R_{\text{res}}$	Background free!	$^{90}\text{Sr}$ $^{106}\text{Rh}$	-	-	-	
			$^{42}\text{Ar}$	?	-	-	-

Table from T.Lasserre

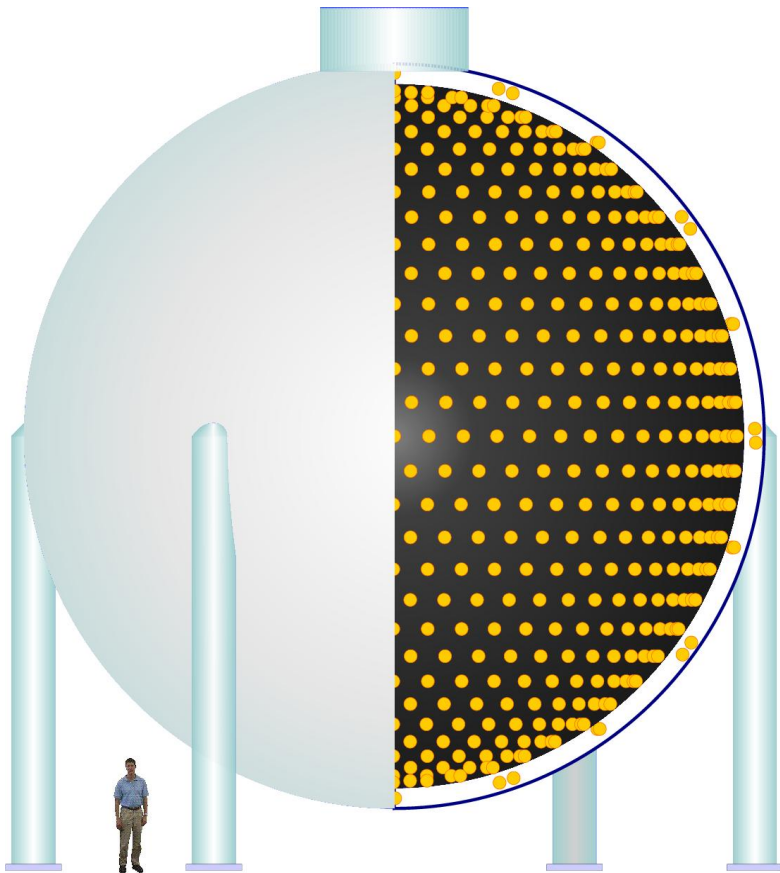


# Conclusions

- A number of intriguing hints at oscillations involving a 4<sup>th</sup>  $\nu$  state
- No single hint is compelling
- Much experimental evidence is in tension with such a 4<sup>th</sup> state
- Nonetheless the situation cannot be ignored
- Definitive experiments are needed in more than one experimental domain
- Many proposals are on the table and some would be definitive

# Backup slides

# MiniBooNE Detector



541 meters downstream of target

3 meter overburden of dirt

12 meter diameter sphere

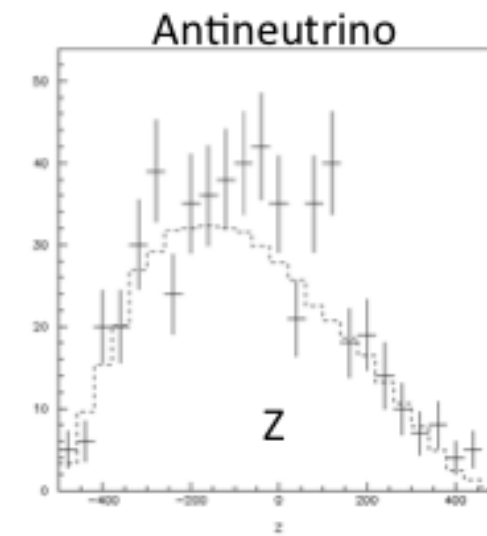
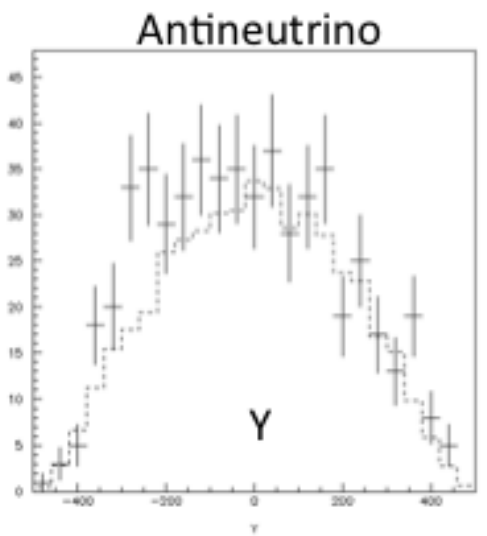
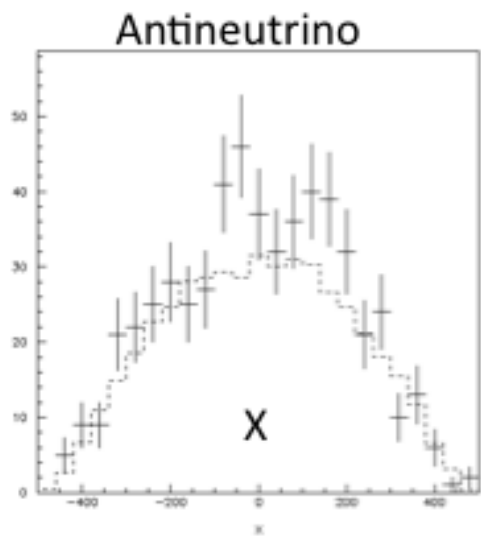
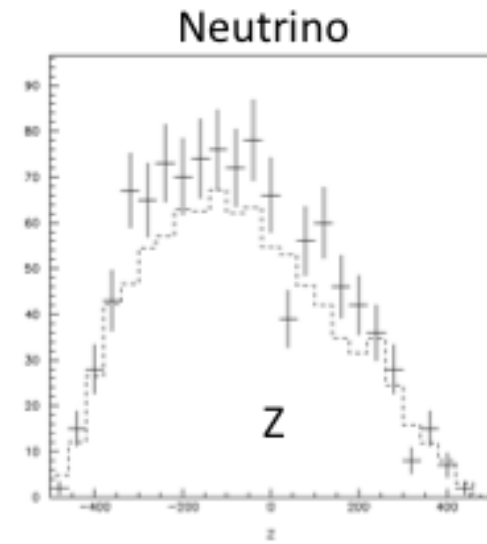
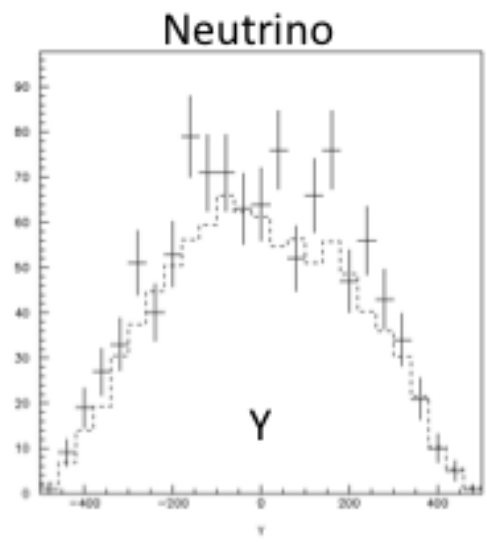
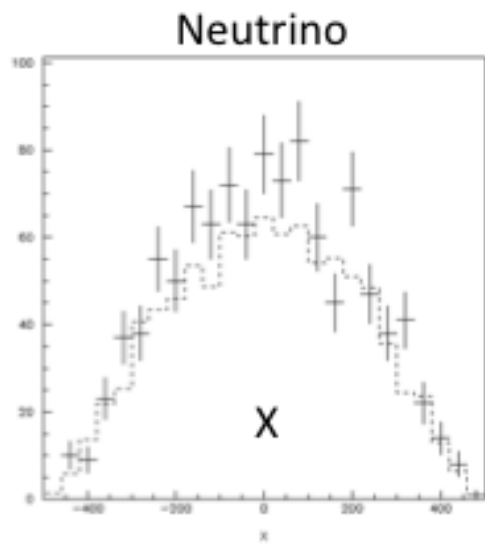
Filled with 800 t of pure mineral oil

( $\text{CH}_2$ --density 0.86,  $n=1.47$ )

Fiducial volume: 450 t

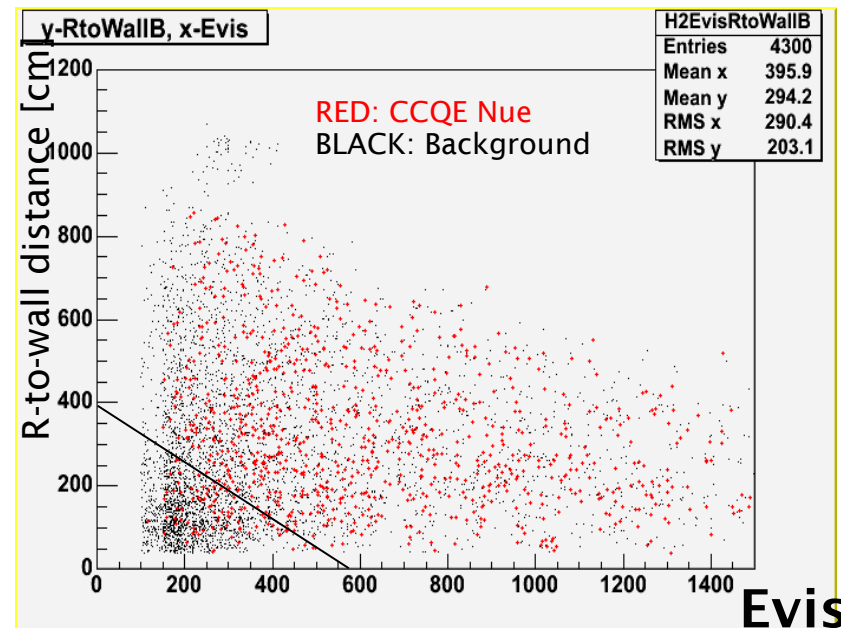
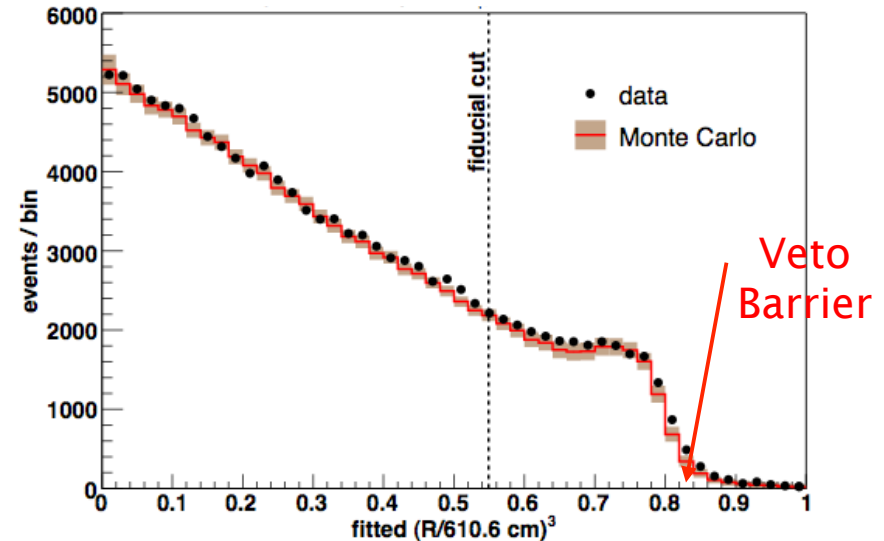
1280 inner 8" phototubes 10% coverage,

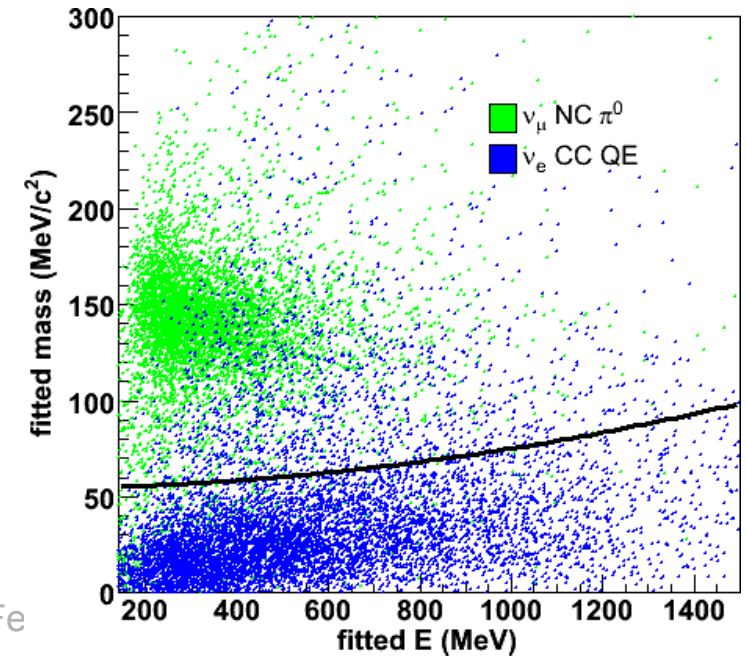
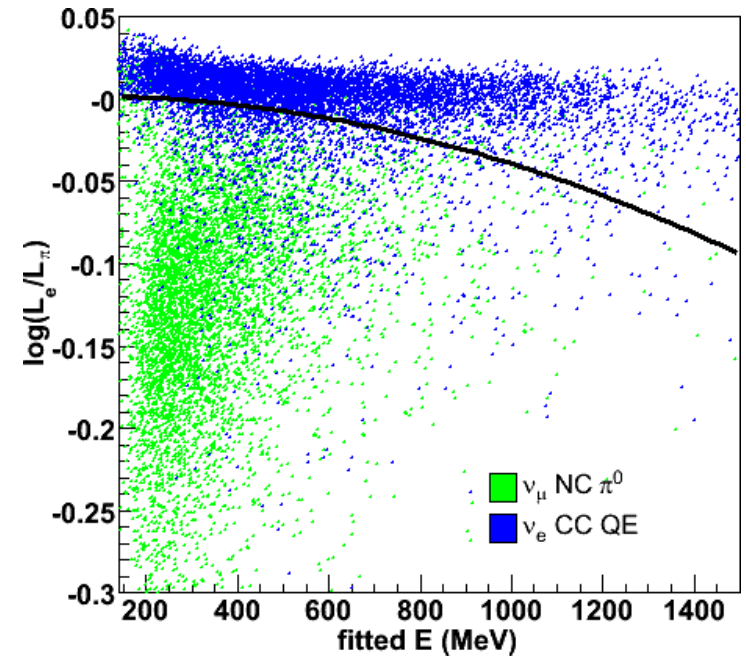
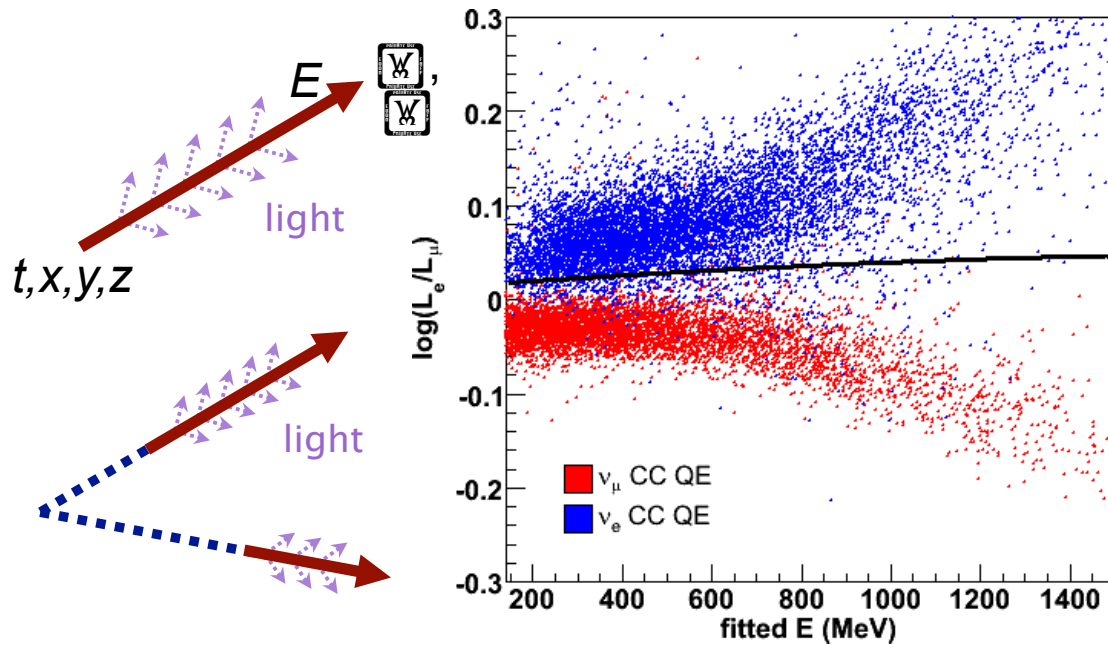
240 veto phototubes



# Signal selection in MiniBooNE

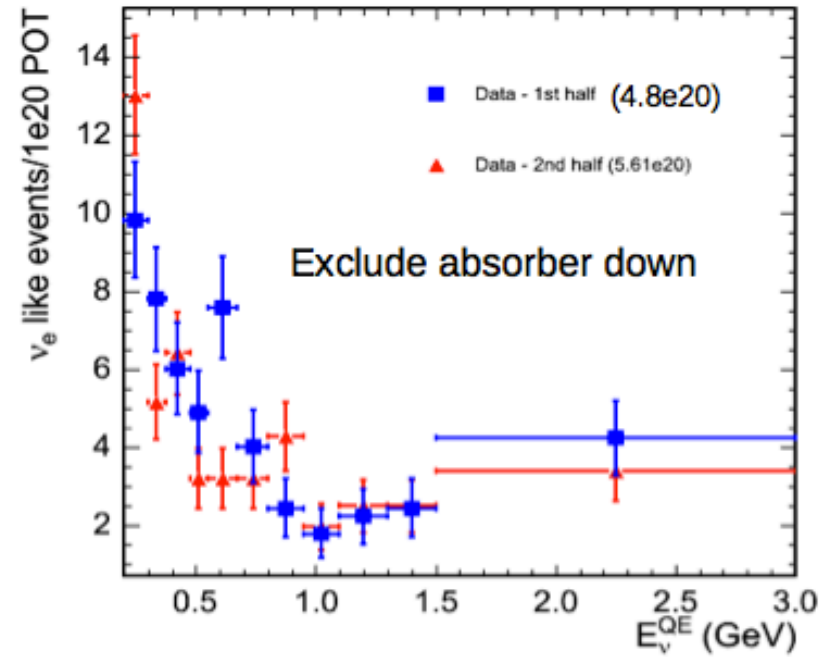
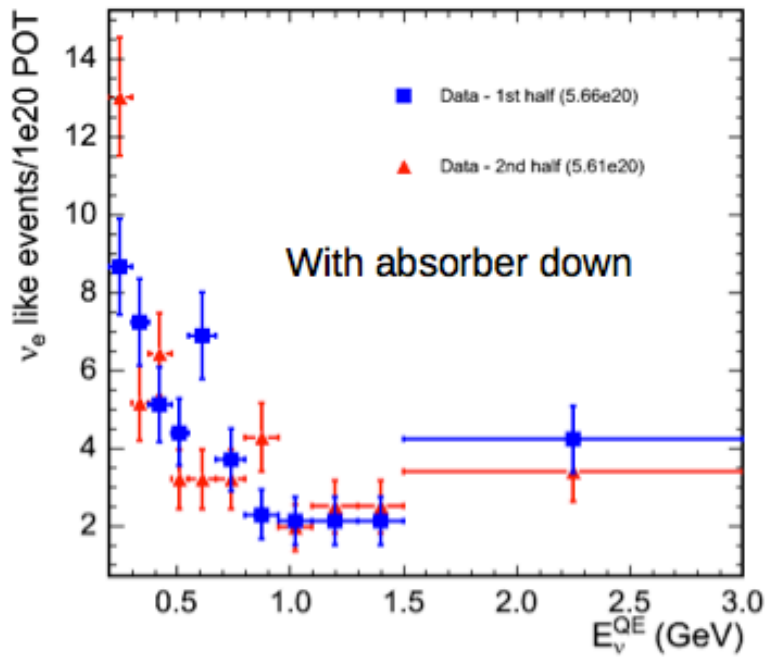
- Neutrino and anti-neutrino analyses are identical
- Start with pre-cuts
  - No late time activity, removes Michel electrons, cuts  $\sim 80\%$  of  $\nu_\mu$  CCQE events
  - Veto hits  $< 6$ , contained & not a cosmic
  - Tank hits  $> 200$  & visible  $E > 140$  MeV, removes NC elastic bkg
  - Radius  $< 500$  cm, far enough from PMTs to avoid area where light modeling becomes less certain
  - R-to-wall backward cut, removes bkg (mainly  $\gamma$ 's) from beam  $\nu$  that interact in dirt outside the detector



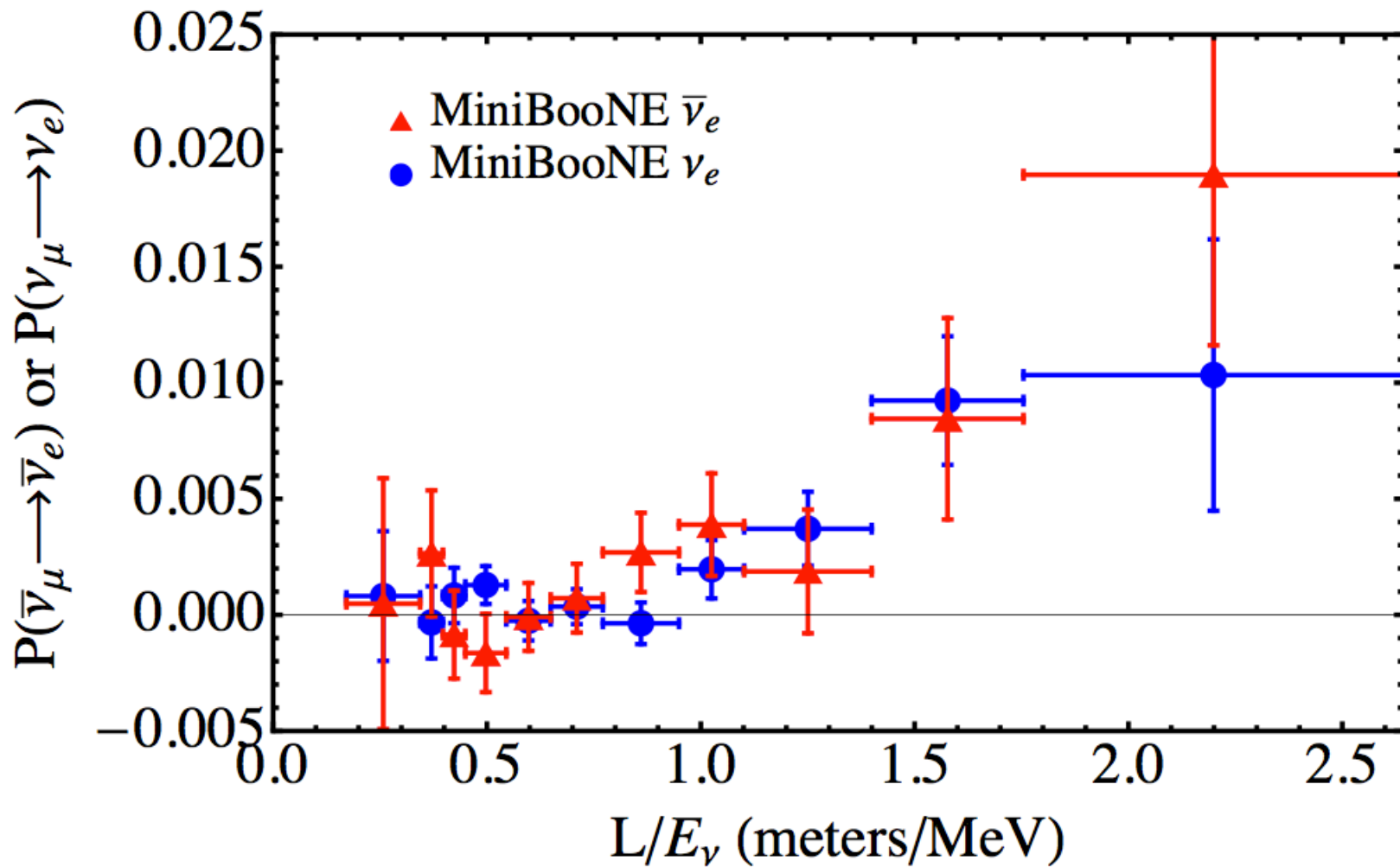


- Form sophisticated Q and T pdfs, and fit for track parameters under 3 hypotheses
  - The track is due to an electron
  - The track is coming from a muon
  - The “track” is a two-track(ring)  $\pi^0$  event
- Apply energy-dependent cuts on  $L(e/\mu)$ ,  $L(e/\pi)$ , and the  $\pi^0$  mass
- Plot remaining events versus  $E_\nu(QE)$  and fit

- KS test 17.8% (29.5% if exclude absorber down period)



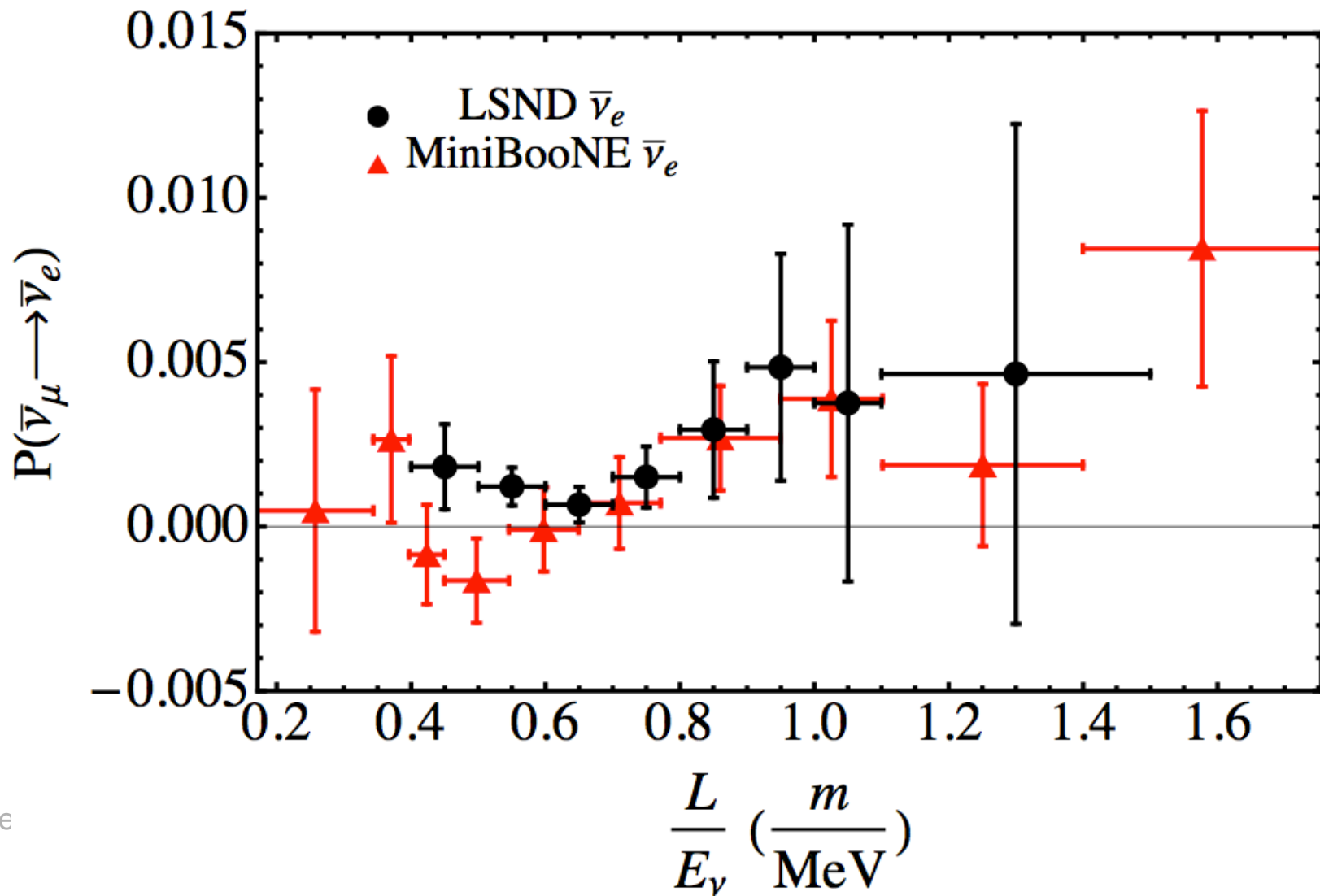
		1st half			2nd half		
	data	mc	excess	data	mc	excess	
200-475	119	100.5±14.3	18.5 (1.3s)	138	100.0±14.1	38 (2.7s)	
475-1250	120	99.1±14.0	20.9 (1.5s)	101	103.1±14.4	-2.2 (-0.2s)	



\* Note this plot assumes that the excess events in anti-neutrino mode come only from the anti-neutrino beam content,  $P(\text{osc})$  at highest 3  $L/E$  bins would be reduced by 25% WS contamination were also included



Can also compare to LSND P(osc)



- Maximum likelihood fit:

$$-2 \ln(L) = (x_1 - \mu_1, \dots, x_n - \mu_n) M^{-1} (x_1 - \mu_1, \dots, x_n - \mu_n)^T + \ln(|M|)$$

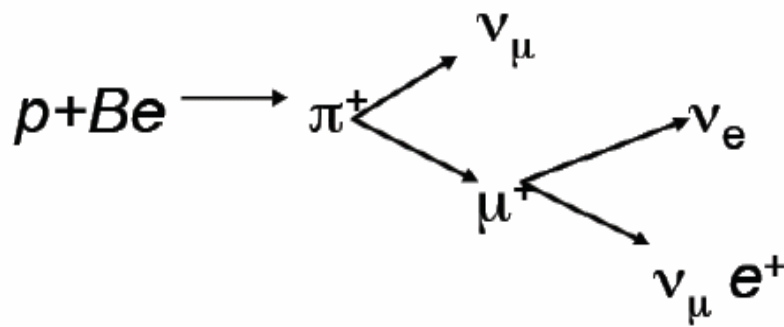
$$M = M_{om} + M_{xsec} + M_{flux} + M_{\pi^0} + M_{dirt} + M_K^0 + \dots$$

- Simultaneously fit (FC-corrected)

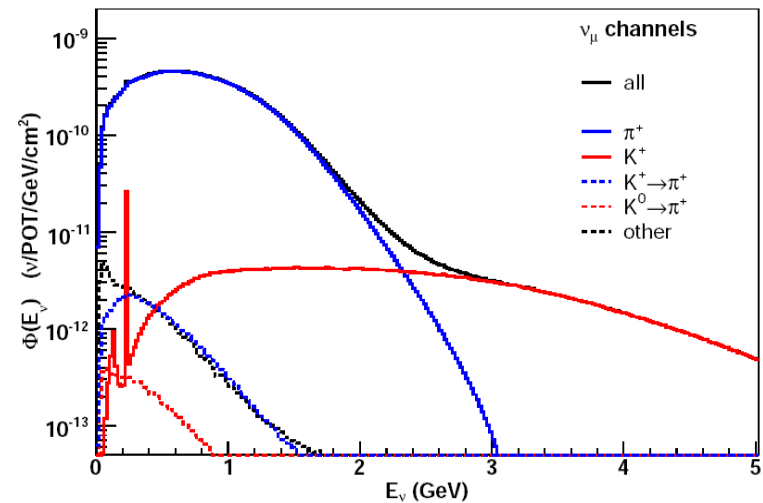
1000's of MC universes go into forming M

- $\nu_e$  CCQE signal + high E  $\nu_e$  sample
- High statistics  $\nu_\mu$  CCQE sample

- $\nu_\mu$  CCQE sample acts like a near detector, i.e. same flux as oscillation  $\nu_e$  by definition, lepton universality + muon mass corrections fix relative cross-section

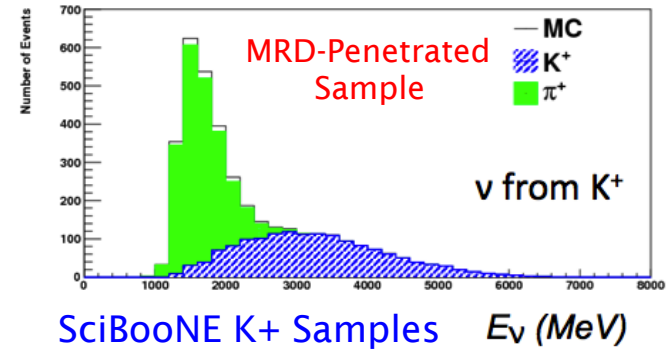
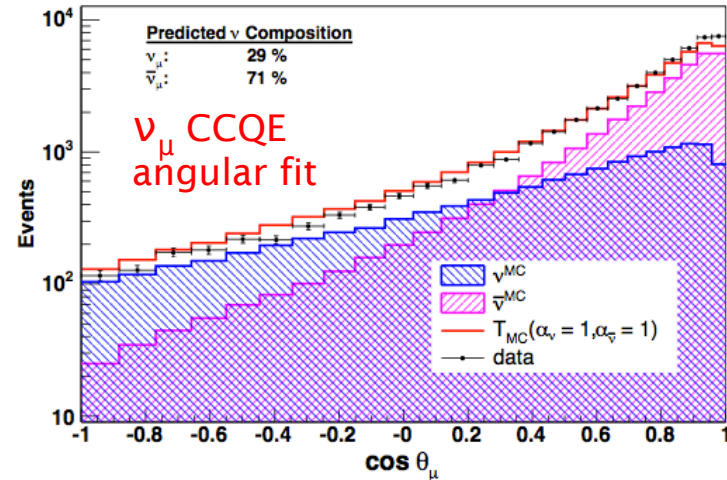
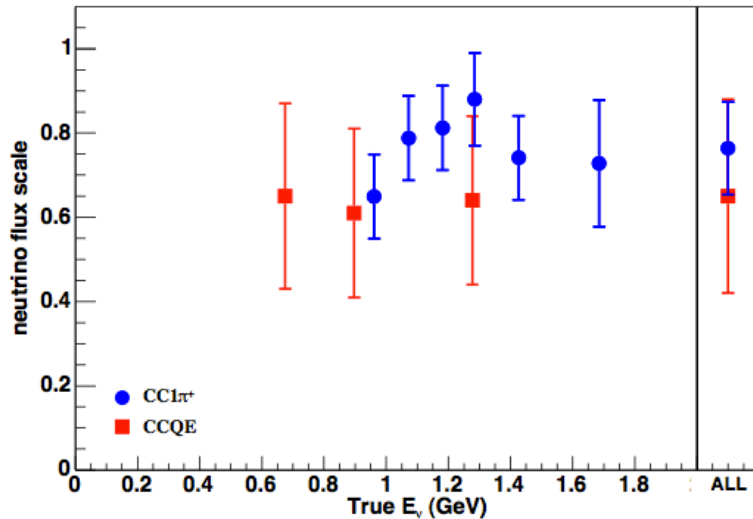


$\nu_\mu$  flux through detector ( $\nu$  mode)

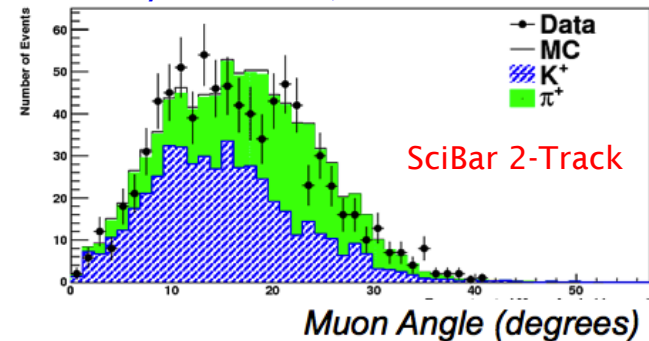


● *In situ* measurement of WS contamination in anti- $\nu$  beam [Phys.Rev.D84,072005 \(2011\)](#)

→  $\nu_\mu$  CCQE angular fit, and new constraint from CC $\pi^+$  rate...good agreement with expectation



SciBooNE K+ Samples  $E_\nu$  (MeV)  
[Phys.Rev.D84,012009 \(2011\)](#)



● New SciBooNE constraint on intrinsic  $\nu_e$  from K+

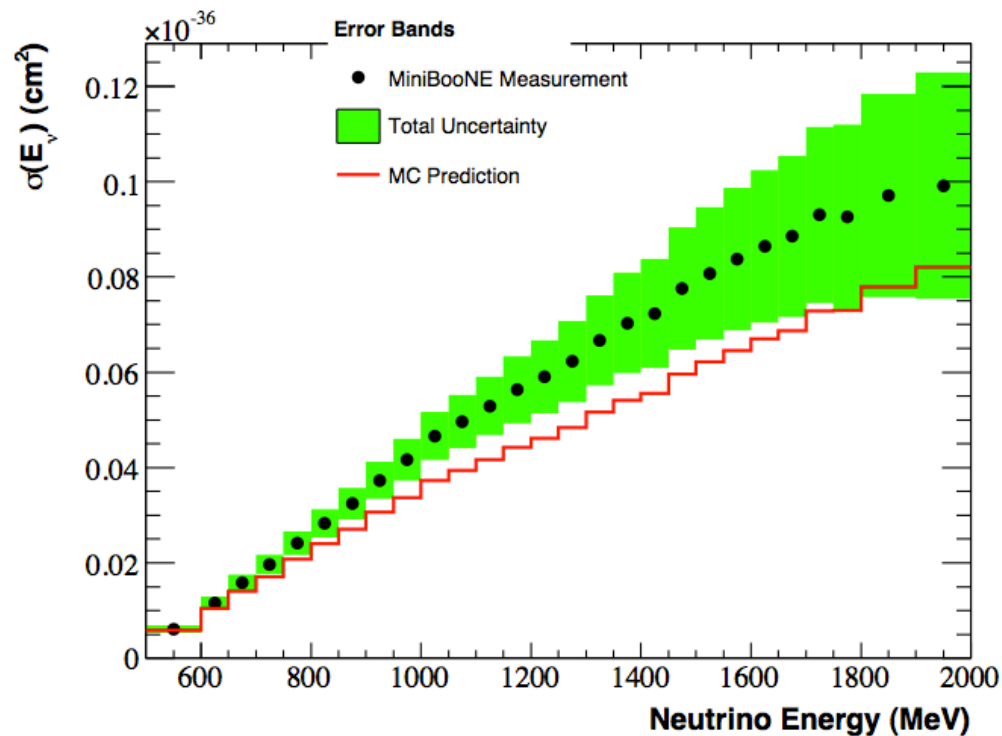
- Found K+ production to be  $0.85 \pm 0.12$  relative to prediction, consistent with prior MiniBooNE assessment of  $1.00 \pm 0.30$
- Combined with world K+ production data, reduces error on K+ flux to 9% in MB Ev range
- Leading error on K+ bkg becomes  $\sim 20\%$  error from cross section

June 20 2012

Steve Brice

● Few other minor updates...

- ➔ Added error matrix for intrinsic  $\nu_e$  from K-
- ➔ Improved smoothing algorithm that was being used to assess systematics due to discriminator thresholds and PMT response
- ➔ CC $\pi^+$  events (bkg for  $\nu_\mu$  CCQE when  $\pi^+$  is absorbed)  $Q^2$  reweighting applied based on internal MB measurement... [Phys.Rev.D83,052007 \(2011\)](#)



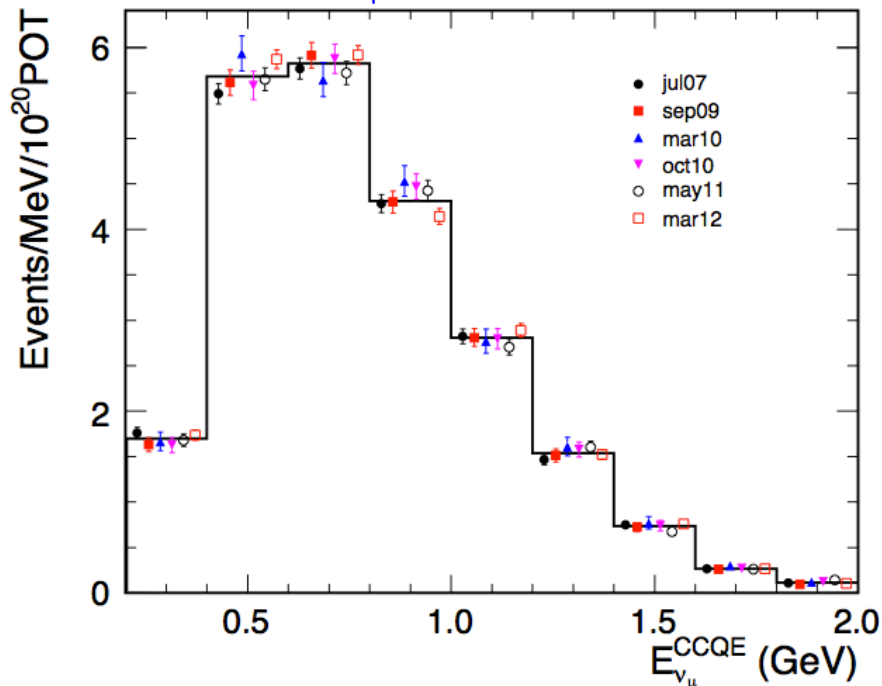
● Statistics of anti-neutrino running has doubled since [Phys.Rev.Lett.105 181801 \(2010\)](#)

➡ 5.66e20 POT --> 11.3e20 POT

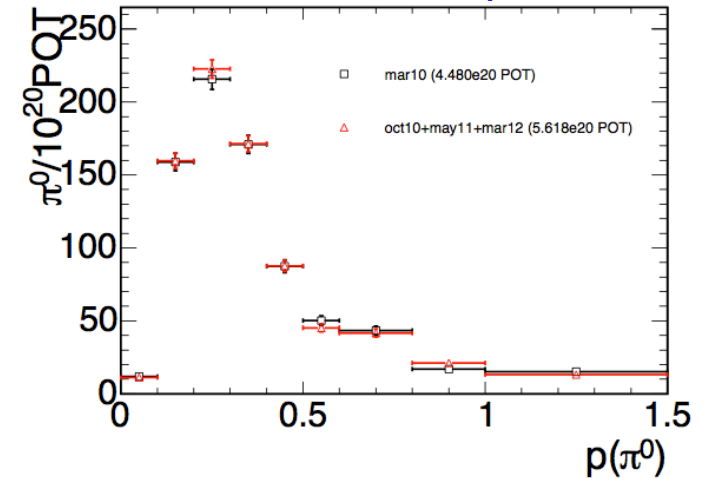
➡ higher statistics in anti- $\nu_e$  appearance

➡ ...and samples used for constraints

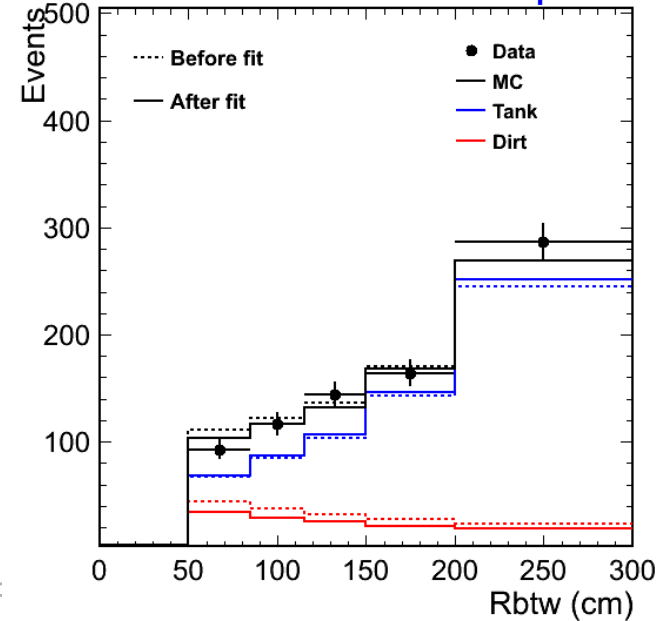
anti- $\nu_\mu$  CCQE Sample



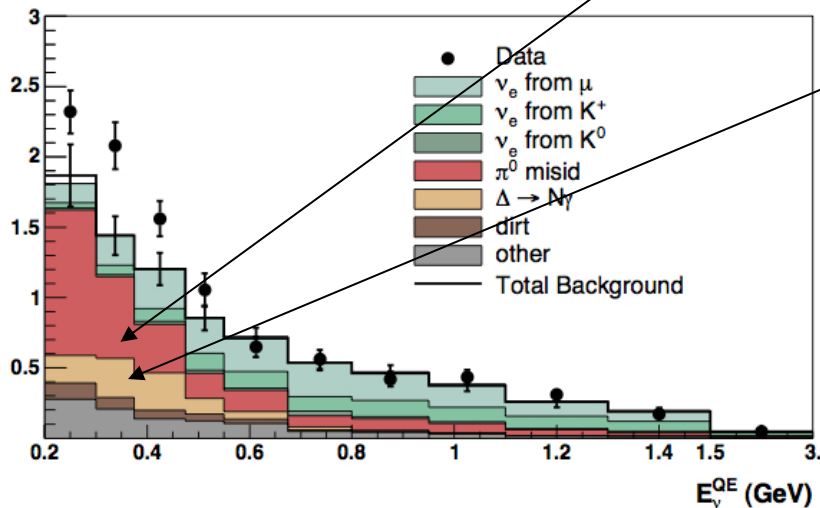
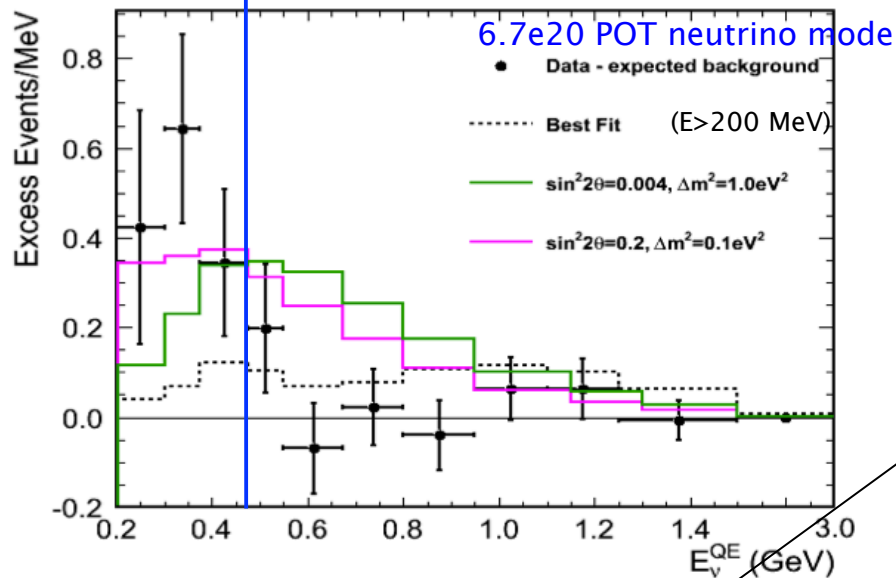
NC  $\pi^0$  Sample



Dirt-enhanced Sample



# What can we say about low-E excess...



- Not a stat fluctuation, statistically  $6\sigma$
- Unlikely to be intrinsic  $\nu_e$ , small bkg at low E
- NC  $\pi^0$  background dominates
  - reduce significance to  $3\sigma$
  - heavily constrained by NC  $\pi^0$  *in situ* rate
- Region where single  $\gamma$  backgrounds can contribute
- MB ties  $\Delta \rightarrow N\gamma$  expected rate to be  $\sim 1\%$  of measured NC  $\pi^0$  rate
  - Number of theory calculation for various single  $\gamma$  processes
  - All find total cross-section within 20% of MiniBooNE's  $\sim 5 \times 10^{-42} \text{ cm}^2/\text{N}$
  - Would need nearly 300% change...

R. Hill, [arXiv:0905.0291](https://arxiv.org/abs/0905.0291)

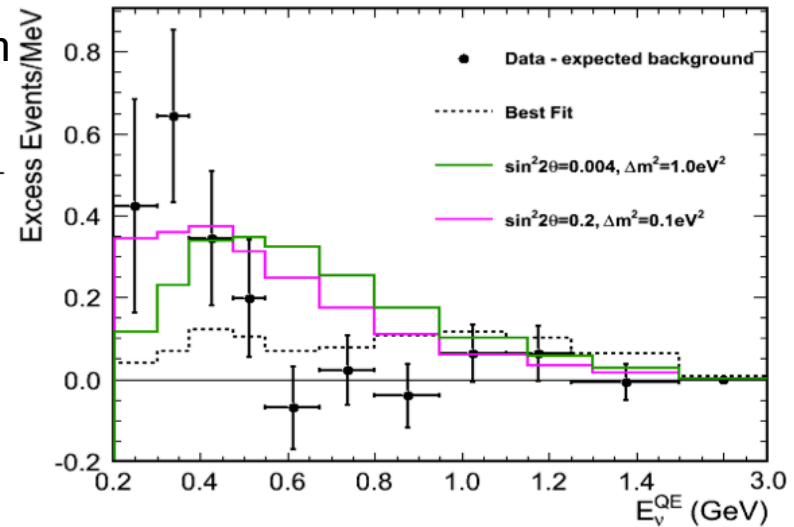
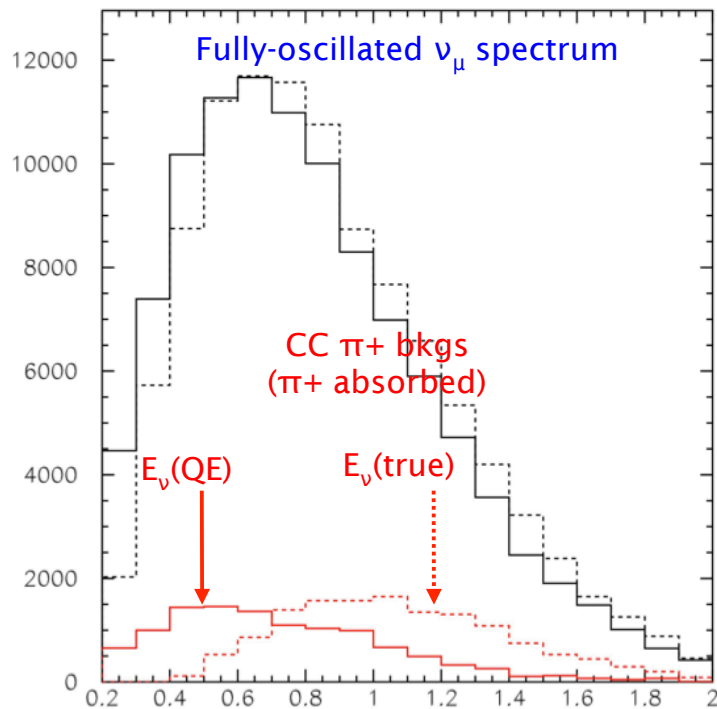
Jenkins & Goldman, [arXiv:0906.0984](https://arxiv.org/abs/0906.0984)

Serot & Zhang, [arXiv:1011.5913](https://arxiv.org/abs/1011.5913)

# Something to consider...

- This plot assumes CCQE-like reconstruction

$$E_\nu = \frac{2(M_n - E_B)E_\mu - (E_B^2 - 2M_nE_B + m_\mu^2 + \Delta M^2)}{2[(M_n - E_B) - E_\mu + p_\mu \cos \theta_\mu]}$$



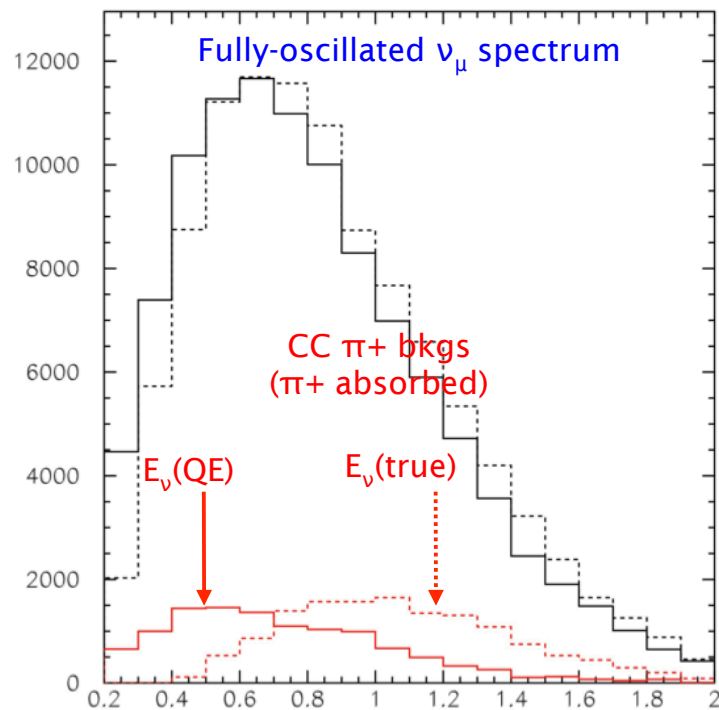
- Additional participants other than the outgoing lepton and struck nucleon will cause events to reconstruct at lower  $E_\nu(\text{QE})$

# Something to consider...

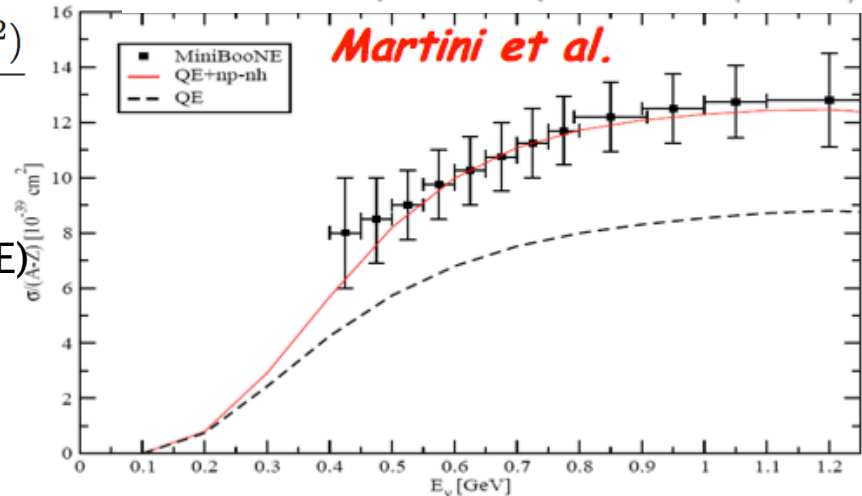
- This plot assumes CCQE-like reconstruction

$$E_\nu = \frac{2(M_n - E_B)E_\mu - (E_B^2 - 2M_n E_B + m_\mu^2 + \Delta M^2)}{2[(M_n - E_B) - E_\mu + p_\mu \cos \theta_\mu]}$$

- Additional participants other than the outgoing lepton and struck nucleon will cause events to reconstruct at lower  $E_\nu$  (QE)



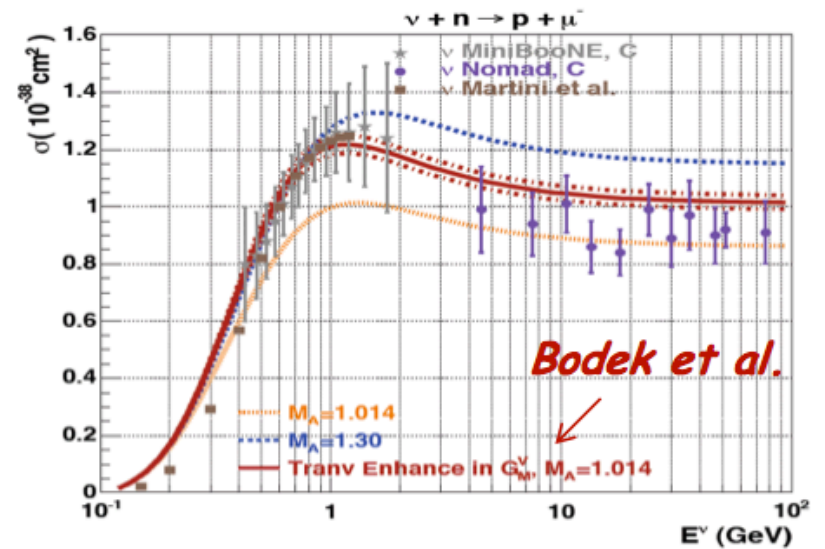
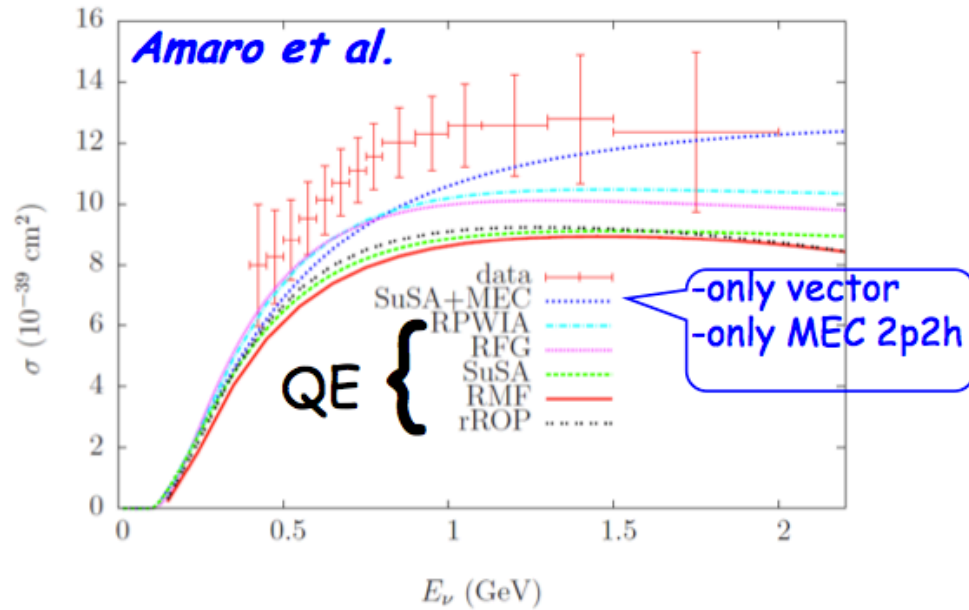
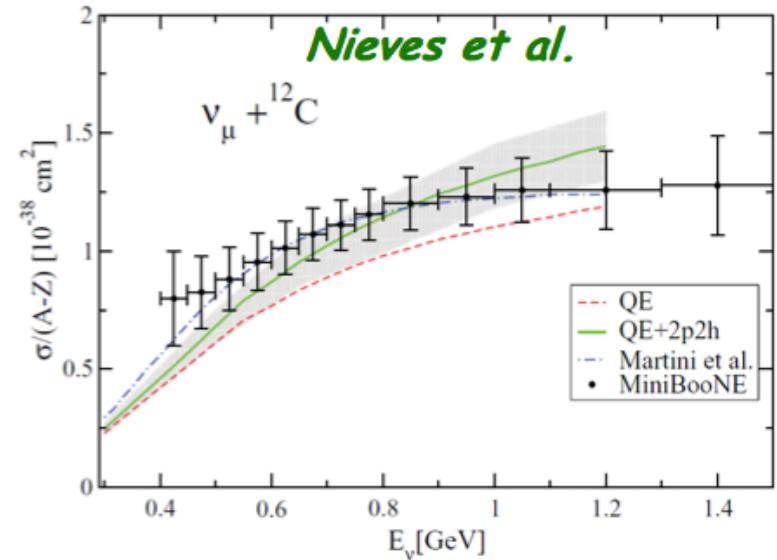
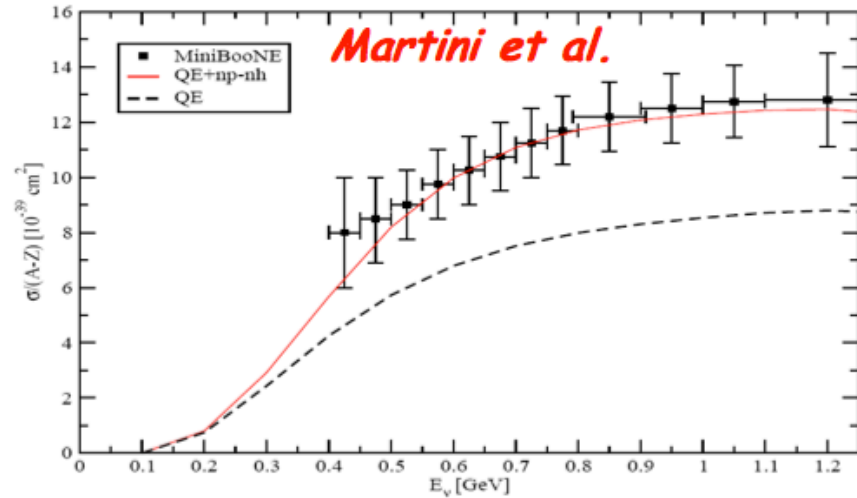
Martini et al., PRC 80, 065001 (2009)



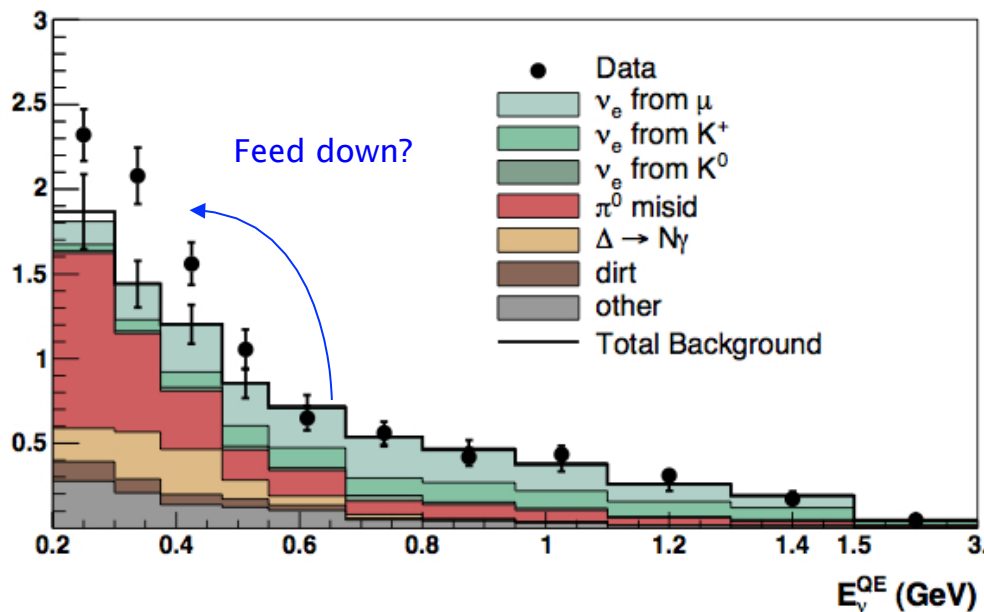
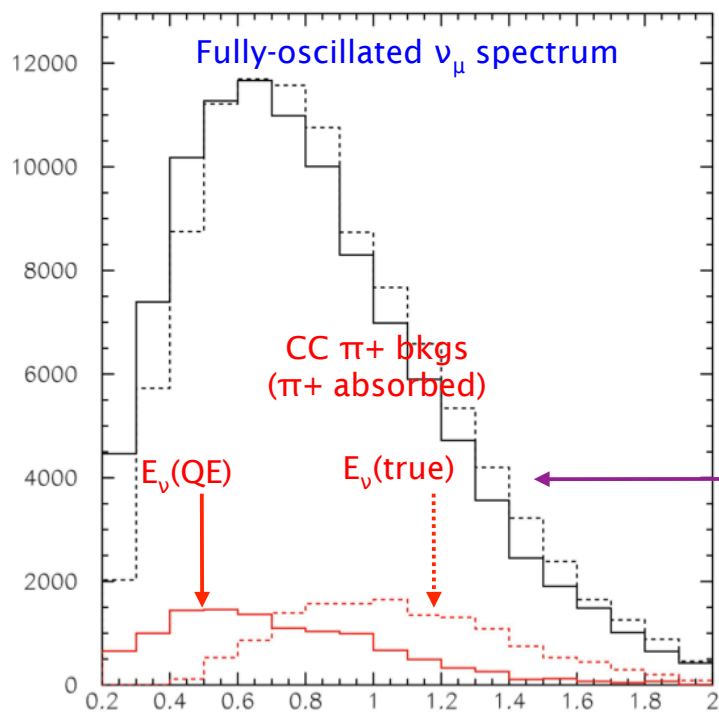
- MiniBooNE finds a cross-section for  $\nu_\mu$  CCQE that is 20-30% higher than expected
- Number of theorists suggesting this could arise from multi-nucleon correlations, observed many years ago in e-scattering
- Could help explain why MB xsec is higher than free nucleon, differences between expts where event selection can depend on final state nucleons



# Comparison between models including np-nh



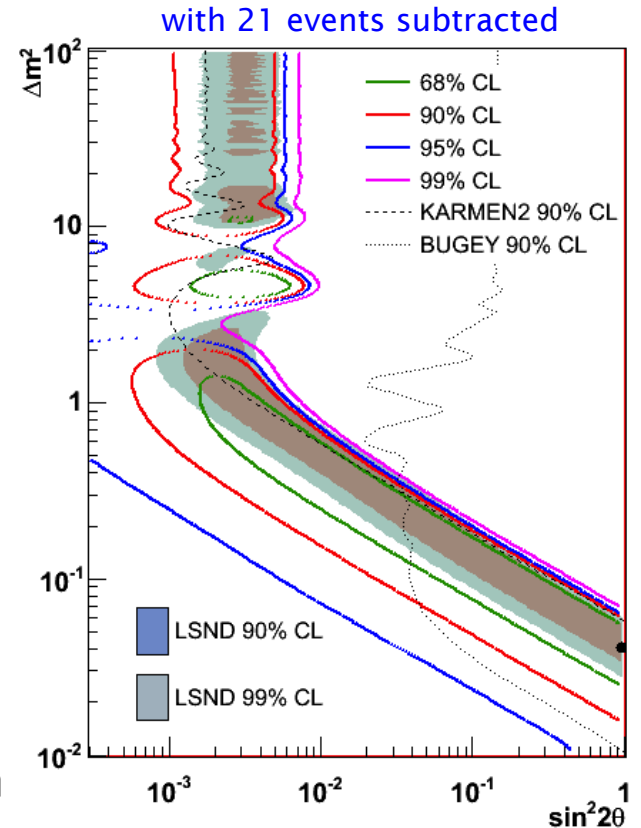
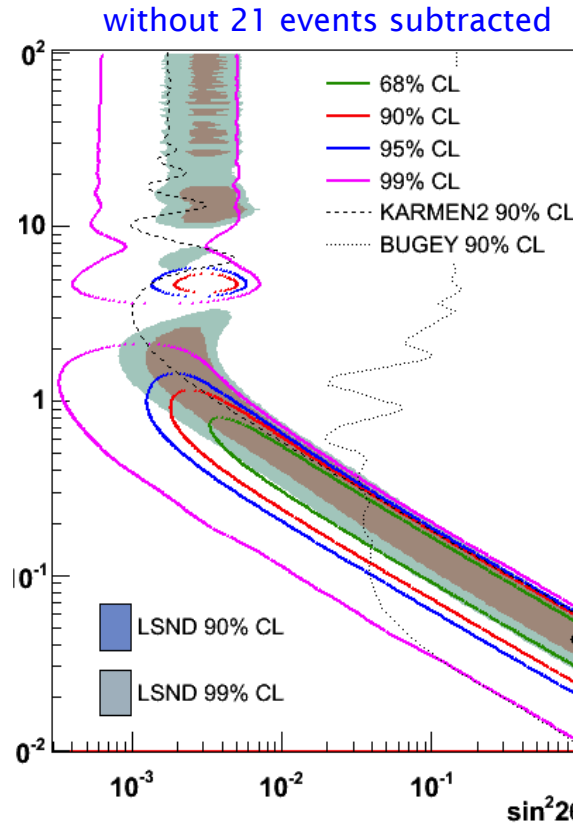
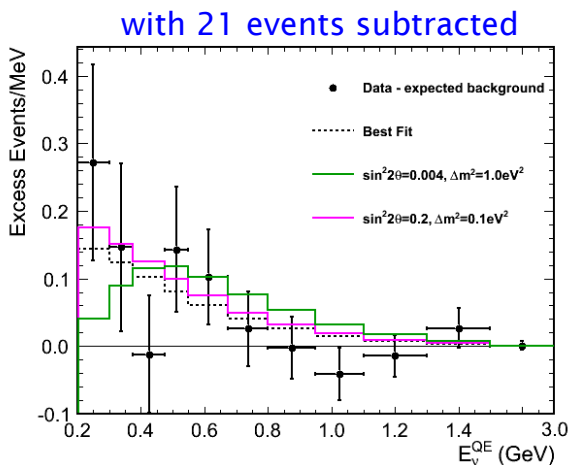
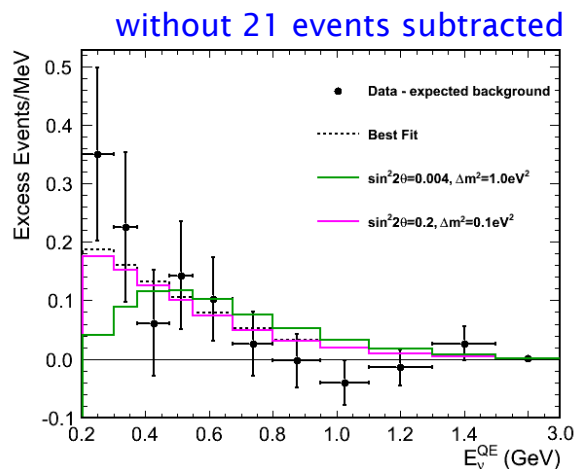
# Relevant for oscillation analysis?



- Means a fraction of oscillated  $\nu_e$  could be misreconstructed (similar to CC $\pi^+$  case)
- Could feed down help relax tension between low and mid range energies?
- Possible, but MiniBooNE corrects sig/bkg predictions based on the measured  $\nu_\mu$  spectrum
- Studies where we double the  $\pi^+$  absorption rate, and then retune sig/bkg predictions to match to CCQE...negligible impact

# Account for neutrino low-E events

- Fits on prior page assume only anti-neutrinos are oscillating, but we know there is a low E excess in  $\nu$  mode data
- Simplest scaling is to assume that there should be an excess in the low energy region proportional to the WS content (21 events)



\*  $E > 200$  MeV fits to full anti-neutrino statistics ( $11.3 \times 10^{20}$  POT)

# Reanalysis of Reactor Experiments

result	Det. type	$\tau_n$ (s)	$^{235}\text{U}$	$^{239}\text{Pu}$	$^{238}\text{U}$	$^{241}\text{Pu}$	old	new	err(%)	corr(%)	L(m)
Bugey-4	$^3\text{He}+\text{H}_2\text{O}$	888.7	0.538	0.328	0.078	0.056	0.987	0.926	3.0	3.0	15
ROVNO91	$^3\text{He}+\text{H}_2\text{O}$	888.6	0.614	0.274	0.074	0.038	0.985	0.924	3.9	3.0	18
Bugey-3-I	$^6\text{Li-LS}$	889	0.538	0.328	0.078	0.056	0.988	0.930	4.8	4.8	15
Bugey-3-II	$^6\text{Li-LS}$	889	0.538	0.328	0.078	0.056	0.994	0.936	4.9	4.8	40
Bugey-3-III	$^6\text{Li-LS}$	889	0.538	0.328	0.078	0.056	0.915	0.861	14.1	4.8	95
Goesgen-I	$^3\text{He+LS}$	897	0.620	0.274	0.074	0.042	1.018	0.949	6.5	6.0	38
Goesgen-II	$^3\text{He+LS}$	897	0.584	0.298	0.068	0.050	1.045	0.975	6.5	6.0	45
Goesgen-III	$^3\text{He+LS}$	897	0.543	0.329	0.070	0.058	0.975	0.909	7.6	6.0	65
ILL	$^3\text{He+LS}$	889	$\approx 1$	—	—	—	0.832	0.7882	9.5	6.0	9
Krasn. I	$^3\text{He+PE}$	899	$\approx 1$	—	—	—	1.013	0.920	5.8	4.9	33
Krasn. II	$^3\text{He+PE}$	899	$\approx 1$	—	—	—	1.031	0.937	20.3	4.9	92
Krasn. III	$^3\text{He+PE}$	899	$\approx 1$	—	—	—	0.989	0.931	4.9	4.9	57
SRP I	Gd-LS	887	$\approx 1$	—	—	—	0.987	0.936	3.7	3.7	18
SRP II	Gd-LS	887	$\approx 1$	—	—	—	1.055	1.001	3.8	3.7	24
ROVNO88-1I	$^3\text{He+PE}$	898.8	0.607	0.277	0.074	0.042	0.969	0.901	6.9	6.9	18
ROVNO88-2I	$^3\text{He+PE}$	898.8	0.603	0.276	0.076	0.045	1.001	0.932	6.9	6.9	18
ROVNO88-1S	Gd-LS	898.8	0.606	0.277	0.074	0.043	1.026	0.955	7.8	7.2	18
ROVNO88-2S	Gd-LS	898.8	0.557	0.313	0.076	0.054	1.013	0.943	7.8	7.2	25
ROVNO88-3S	Gd-LS	898.8	0.606	0.274	0.074	0.046	0.990	0.922	7.2	7.2	18

PRD83, 073006 (2011)