

# The four basic ways of creating dark matter through a portal

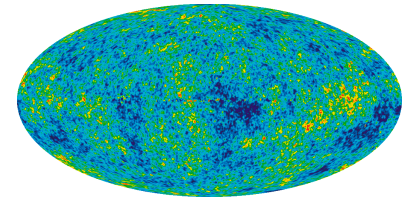
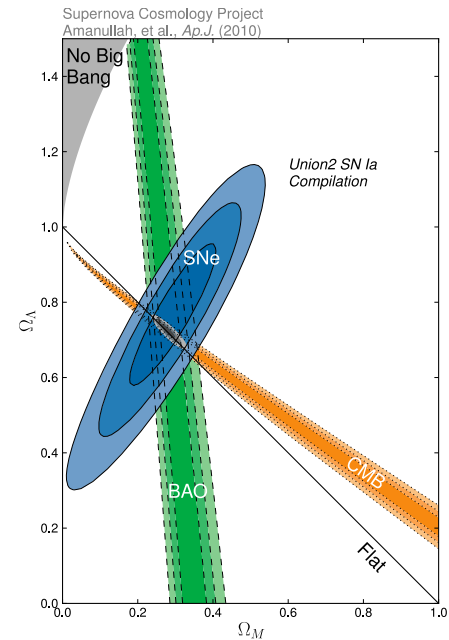
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Univ. Autonoma Madrid & IFT, Spain

# Gravitational evidences of dark matter

- At galactic scale: velocity distribution of stars
- At galaxy cluster scale: -velocity distribution of galaxies  
-bullet cluster
- At cosmological scales: CMB data (WMAP),  
supernovae,....

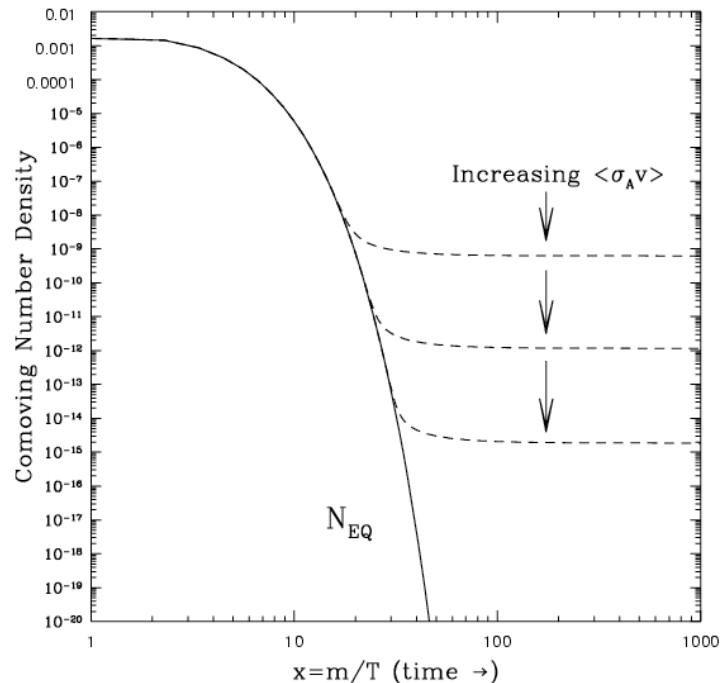
lead consistently to:  $\Omega_{DM} \simeq 0.229 \pm 0.015\%$



- DM is neutral, stable ( $\tau_{DM} > 10^{26}$  sec), cold,  $\Omega_{DM} \simeq 23\%$ , has constrained cross section on Nucleon, produces constrained fluxes of cosmic rays, BBN, ....
- but this still leaves an enormous freedom for the DM particle (mass, spin, interactions, stabilization mechanism, ...)

# The WIMP freeze-out mechanism

Relic density from annihilation freeze out:



down to  $T \sim m_{DM}$  DM is thermal equilibrium  
 for  $T \lesssim m_{DM}$  Boltzmann suppression of  $n_{DM}^{eq}$

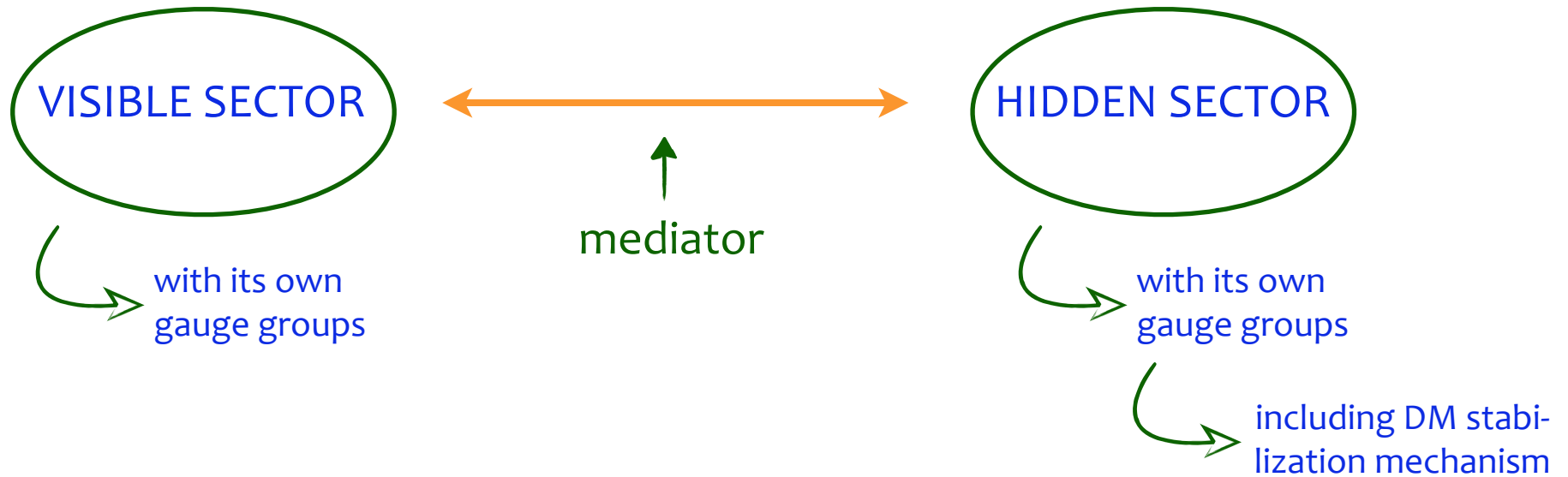
freeze-out of annihilation

$$\Omega_{DM} \propto 1 / \langle \sigma_{annih} v \rangle$$

$\Rightarrow$  if  $m_{DM} \sim 1 \text{ GeV} - 10 \text{ TeV}$  and  $\lambda \sim 1 \Rightarrow \Omega_{DM} \sim 23\%$   $\leftarrow \sigma_{annih} v \simeq 10^{-26} \text{ cm}^3/\text{sec}$

$\Rightarrow$  most straightforward/natural mechanism but not at all the only possible/simple one

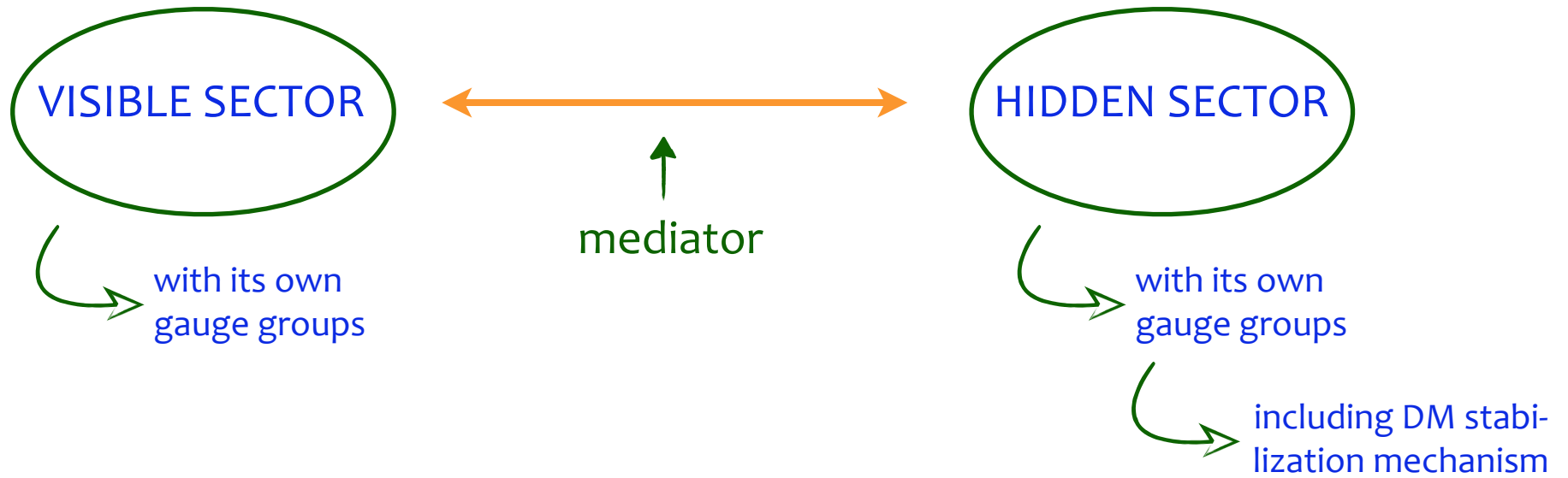
# A general visible sector/hidden sector/ mediator DM setup



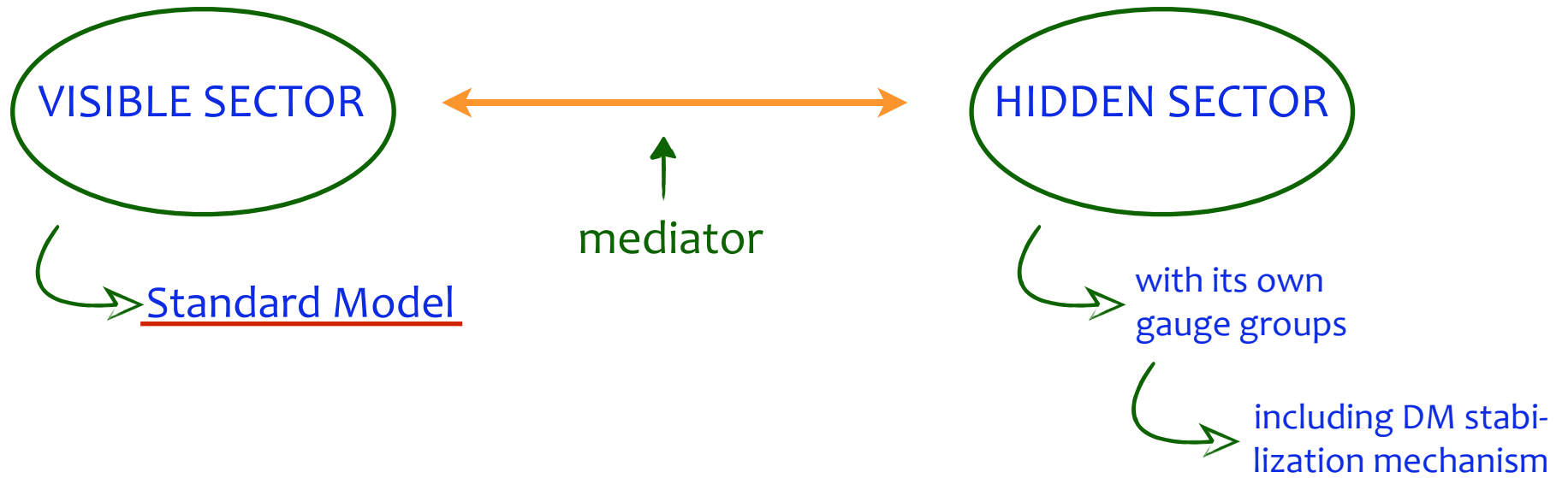
such a structure gives 4 regimes to get the observed relic density



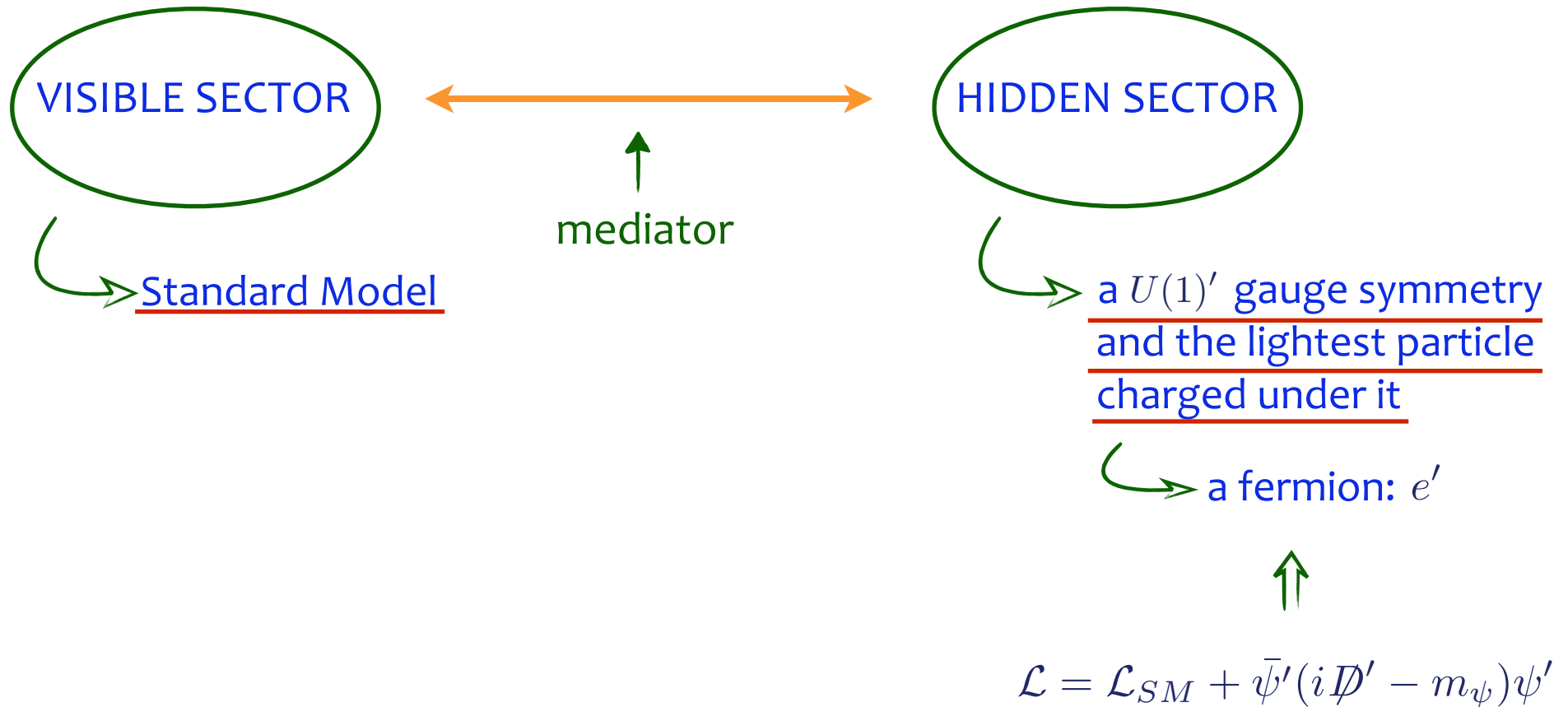
# A very simple light mediator model



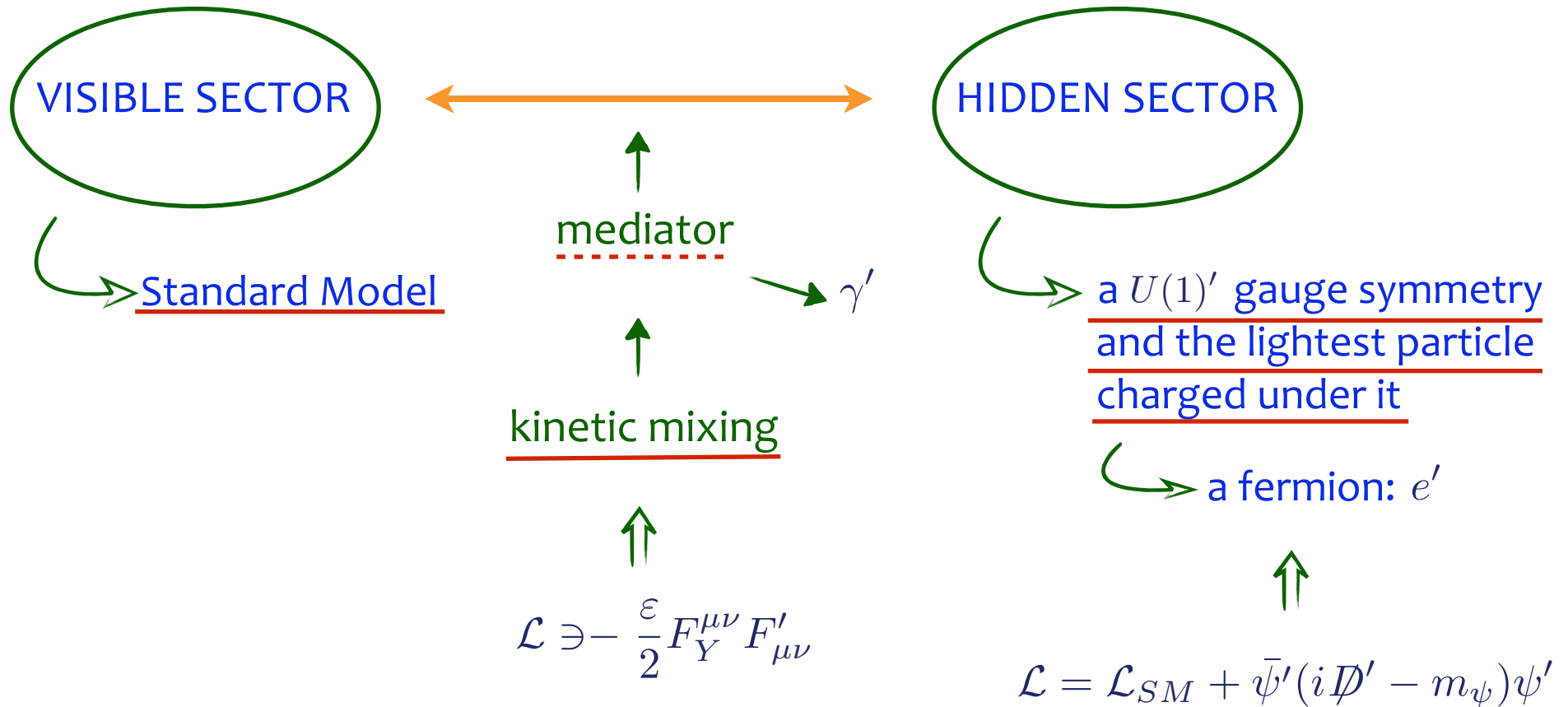
# A very simple light mediator model



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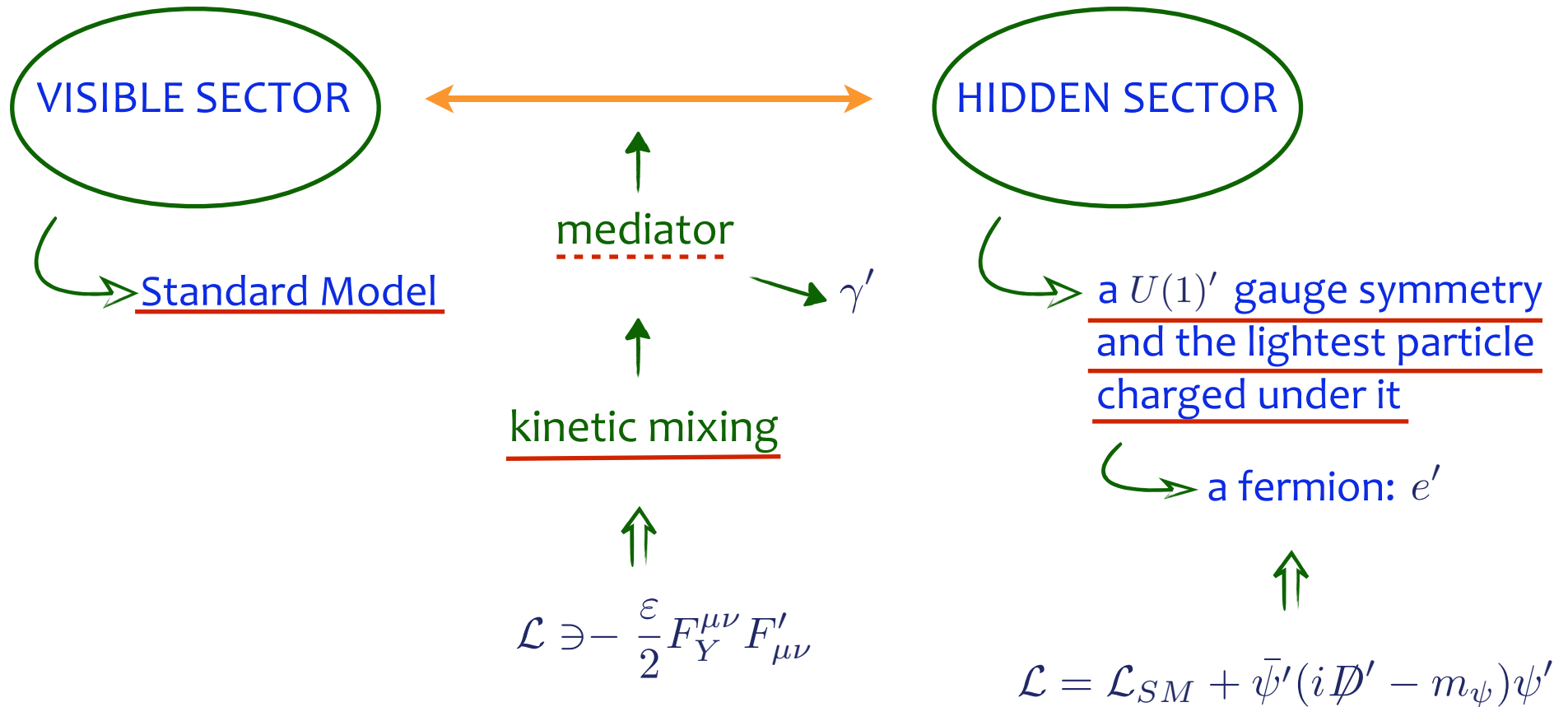
# A very simple light mediator model



Feldman, Kors, Nath 06'  
 Pospelov, Ritz, Voloshin 08'

↪ for a massive  $Z'$

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↪ for a massive  $Z'$

⇒ a good DM candidate based on 3 parameters:  $m_{DM}, \alpha', \epsilon$   
 ↪ =  $m_{e'}$

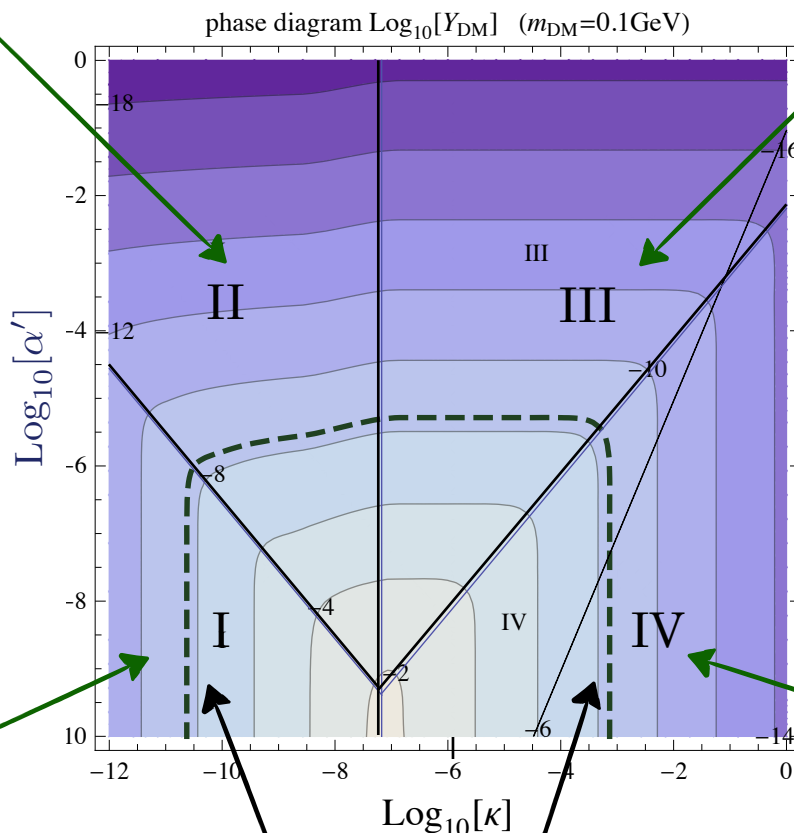
# Motivations for such a Hidden sector gauge structure

- UV.....
- simplicity
- the stability of the DM particle is a fundamental issue!  $\leftarrow \tau_{DM} > 10^{26} \text{ sec}$ 
  - $\hookrightarrow$  not that many stabilization mechanisms
  - $\hookrightarrow$  one of the simplest: the lightest charged particle under a new gauge group
- visible sector = Standard Model  $\Rightarrow$  mass and interactions of source particles are known
- new DM long range force from  $\gamma'$   $\Rightarrow$  rich cosmological phenomenology
  - $\hookrightarrow$  studied in details in: Ackerman, Buckley, Carroll, Kamionkowski 08', Feng Kaplinghat, Tu, Yu 09', Feng, Tu, Yu 08 see also Foot et al. 06'-10'
- prototype of visible sector/hidden sector/mediator structure
  - $\hookrightarrow$  with such a structure: not only freeze-out and freeze-in but 4 DM production regimes

# Relic density phase diagram

Reannihilation regime

Hidden sector freeze-out regime



Freeze-in regime

Connector freeze-out regime

$$\Omega_{DM} \sim 23\%$$

$$\kappa \equiv \epsilon \sqrt{\alpha'/\alpha}$$

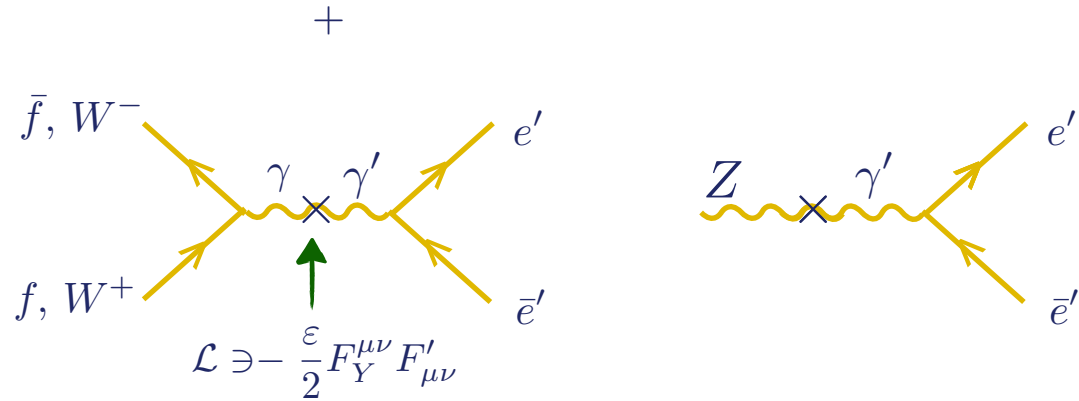
Observed relic density: "square" or "mesa" shape

in each regime  $\Omega_{DM}$  depends essentially on one coupling

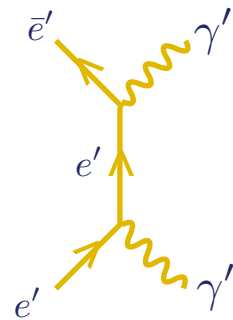
characteristic of the visible sector/hidden sector/mediator structure

# Relevant processes

Connector processes:



Hidden sector process:



↪ convenient to go in a  $\gamma - \gamma'$  basis where kinetic terms are canonical (i.e. no  $\gamma - \gamma'$  mixing)

↪ basis where  $\gamma'$  couples to  $f$   
 $\gamma$  couples to  $e'$  and  $f$

Holdom 86'

⇒  $\sigma(SMSM \rightarrow e'\bar{e}') \propto \alpha^2 \kappa^2$  ←  $\kappa \equiv \epsilon \sqrt{\alpha'/\alpha}$   
 $\sigma(e'\bar{e}' \rightarrow \gamma'\gamma') \propto \alpha'^2$



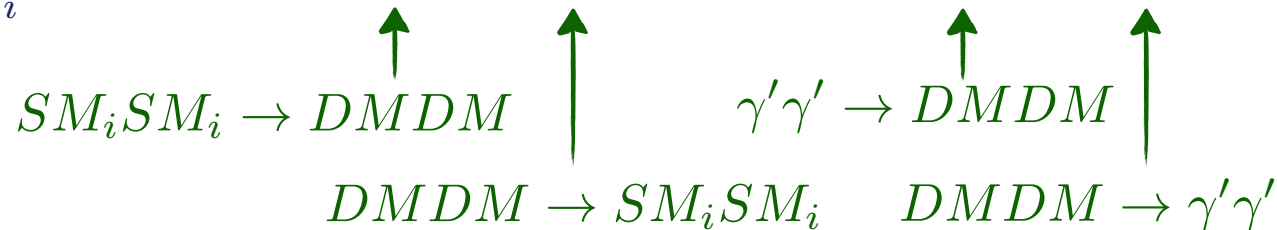
# Boltzmann equation

in terms of the usual  $\langle \sigma v \rangle$ :

$$z \equiv m_{DM}/T$$

$$Y \equiv \frac{n_{e'}}{s} \quad (= Y_{DM}/2)$$

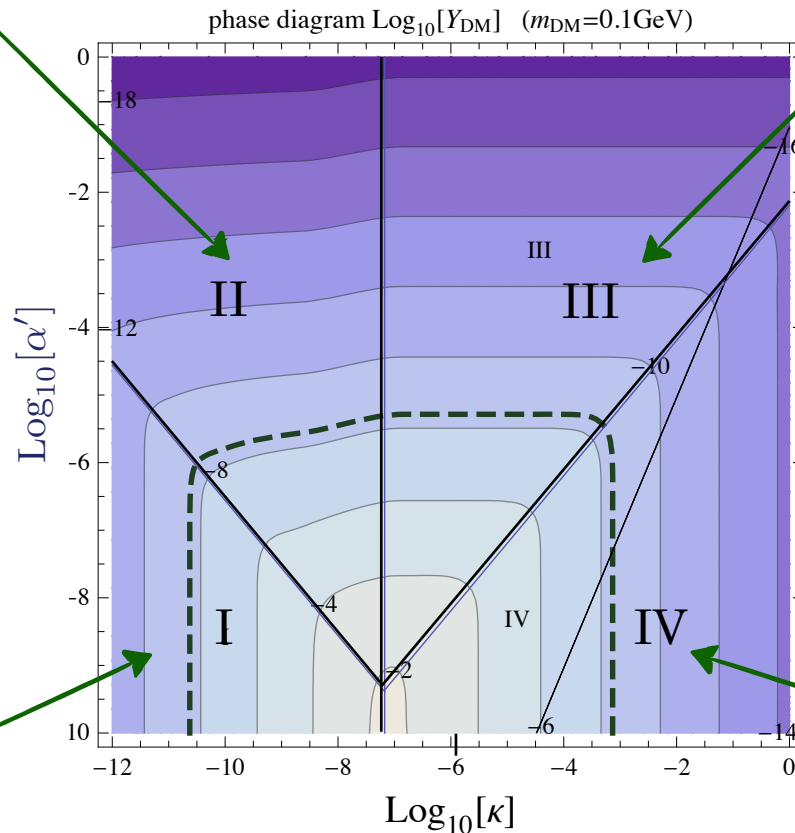
$$z \frac{H}{s} \frac{dY}{dz} = \sum_i \langle \sigma_{connect} v \rangle_i (Y_{eq}^2(T) - Y^2) + \langle \sigma_{HS} v \rangle (Y_{eq}^2(T') - Y^2)$$



# Relic density phase diagram

Reannihilation regime

Hidden sector freeze-out regime



Freeze-in regime

Connector freeze-out regime

$$\Omega_{DM} \sim 23\%$$

$$\kappa \equiv \epsilon \sqrt{\alpha' / \alpha}$$

- we consider a HS negligible at high temperature  $\begin{cases} Y_{DM} \sim 0 \\ \rho' \sim 0 \end{cases}$  HS energy density
- likely that DM production from inflaton decay negligible
- if reheating occurs mostly in one of the feebly coupled sectors
- can be tested if the DM mass and coupling measured are the ones which give the right relic density

# $\kappa$ and $\alpha'$ are small $\Rightarrow$ freeze-in regime

$\hookrightarrow$  if  $\kappa$  and  $\alpha'$  small:  $SMSM \leftrightarrow DMDM$  does not thermalize  
 $DMDM \leftrightarrow \gamma'\gamma'$  does not thermalize

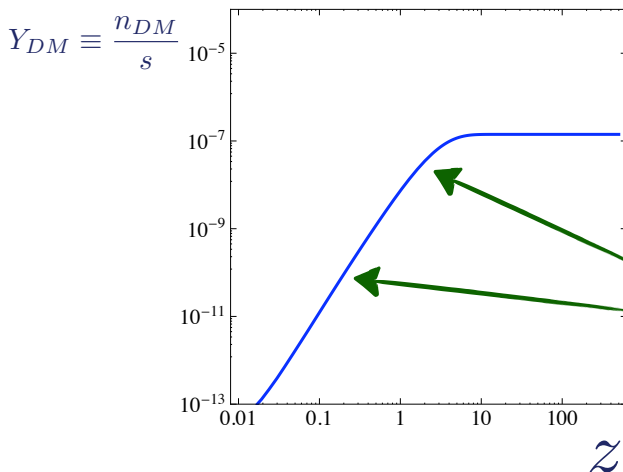
$$z \frac{H}{s} \frac{dY}{dz} = \sum_i \langle \sigma_{connect} v \rangle_i (Y_{eq}^2(T) - \cancel{Y^2}) + \langle \sigma_{HS} v \rangle (Y_{eq}^2(T') - \cancel{Y^2})$$

$\begin{matrix} \uparrow & & \uparrow & & \uparrow & & \uparrow \\ SM_i SM_i \rightarrow DMDM & & & & \gamma' \gamma' \rightarrow DMDM & & \\ \uparrow & & \uparrow & & \uparrow & & \uparrow \\ DMDM \rightarrow SM_i SM_i & & & & DMDM \rightarrow \gamma' \gamma' & & \end{matrix}$

only  $SMSM \rightarrow DMDM$  is relevant because only  $Y_{SM}^{eq}$  is large

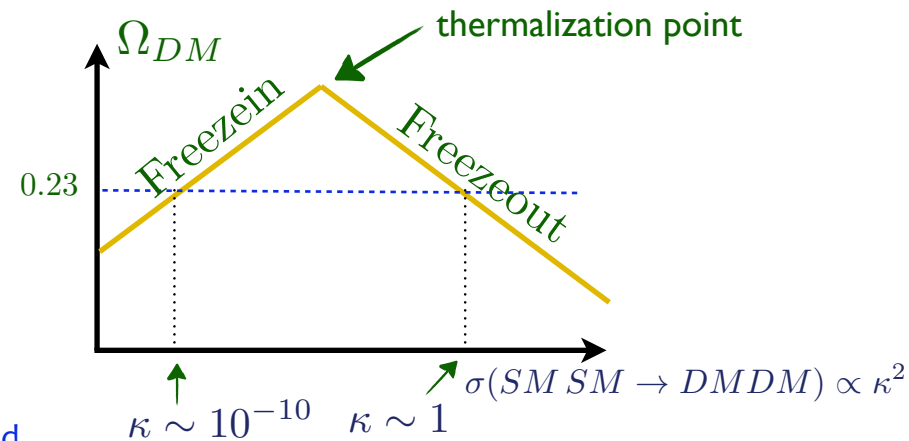
$\Rightarrow$  freeze-in regime:

Mc Donald 02'  
Hall, Jedamzik,  
March-Russell, West 09'



$Y_{DM} \propto 1/T$  down to  
 $T \sim m_{DM}$  where  $n_A^{eq}$   
 becomes Boltz. suppressed

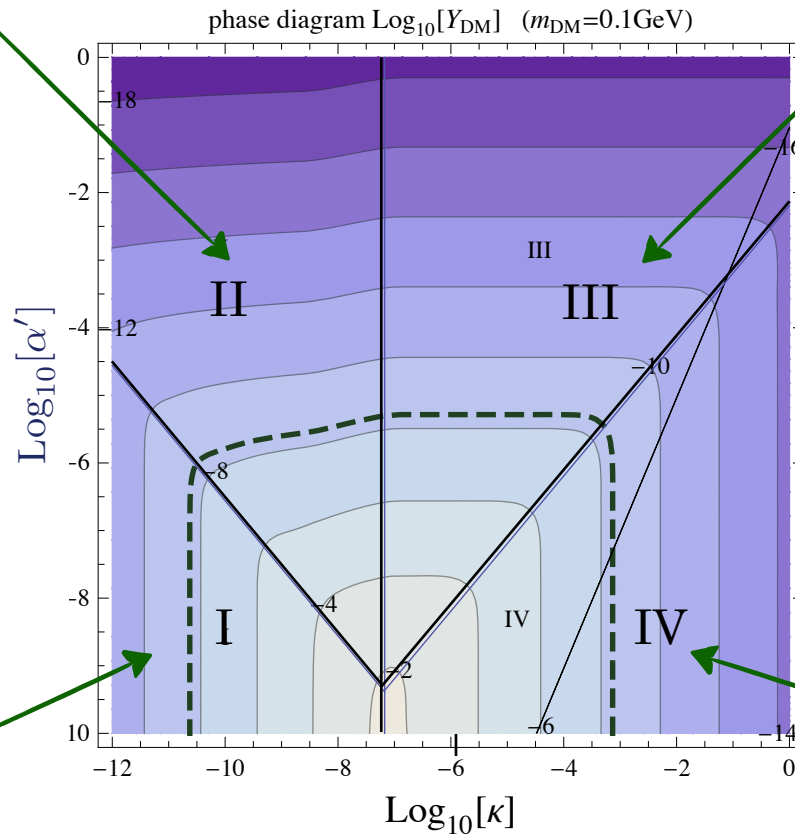
$\hookrightarrow$  DM production IR dominated



# Relic density phase diagram

Reannihilation  
regime

Hidden sector  
freeze-out regime



$$\Omega_{\text{DM}} \sim 23\%$$

Freeze-in  
regime

Connector  
freeze-out regime

# Reannihilation regime

 if one increases  $\kappa \Rightarrow$  more DM created  
 if one increases  $\alpha' \Rightarrow \langle \sigma_{HSv} \rangle$  increases

} both favor thermalization of  $\langle \sigma_{HSv} \rangle$



at some point the  $\gamma'$  thermalize with the  $e'$ :

$$\Gamma_{annih} > H$$

$$\Gamma_{annih} = \underline{\underline{n_{e'}}} \langle \sigma_{HSv} \rangle$$



$$n_{eq}(T) \langle \underline{\underline{\sigma_{effv}}} \rangle > H$$

$$n_{eq}^2(T) \langle \sigma_{connectv} \rangle / H$$

$$\langle \sigma_{effv} \rangle \equiv \sqrt{\langle \sigma_{HSv} \rangle \langle \sigma_{connectv} \rangle}$$

 we can define a HS temperature  $T'$

$$n_{\gamma'} = n_{eq}(T') \sim g_{\gamma'} T'^3$$

$$\rho' \sim g_*^{HS} T'^4$$



necessary to know the number of  $\gamma'$ , i.e. the  $\underline{\underline{\gamma' \gamma' \rightarrow DMDM}}$  rate

$$z \frac{H}{s} \frac{dY}{dz} = \sum_i \langle \sigma_{connectv} \rangle_i (Y_{eq}^2(T) - Y^2) + \langle \sigma_{HSv} \rangle (\underline{\underline{Y_{eq}^2(T')}} - Y^2)$$

 we need to calculate the HS energy density  $\rho'$  in order to determine  $T'$

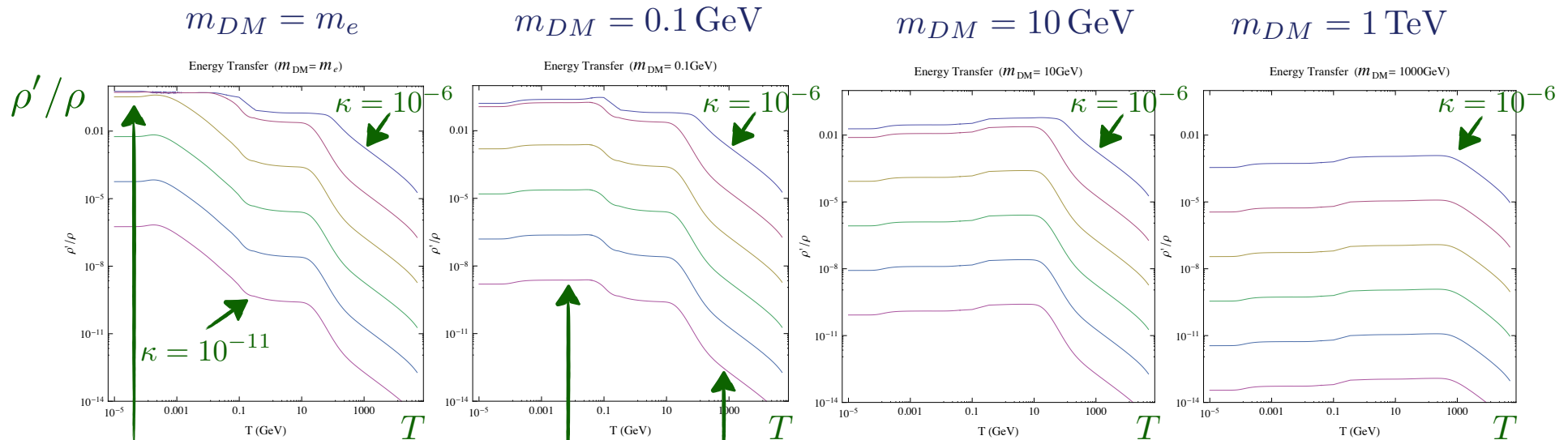
# Calculation of the energy transfer from the SM to the HS

energy transfer Boltzmann equation

$$\frac{d\rho'}{dt} + 3H(\rho' + p') = \int \prod_{i=1}^4 d^3\vec{p}_i \cdot g_i f_1(\vec{p}_1) f_2(\vec{p}_2) |i\mathcal{M}_{12\leftrightarrow 34}|^2 (2\pi)^4 \delta^{(4)}(p_1 + p_2 - p_3 - p_4) \Delta E_{tr}$$

$$\frac{d(\rho'/\rho)}{dT} = -\frac{1}{H(T)T\rho} \frac{g_1 g_2}{32\pi^4} \int ds \cdot \sigma_{connect}(s) (s - 4m^2) s T K_2\left(\frac{\sqrt{s}}{T}\right)$$

$SM_i SM_i \rightarrow DMDM$



$\rho'/\rho$  saturates when  $T$  reaches  $T$   
 $\rho'/\rho \sim 1/T$  for  $T > m_{DM}$   
 $\rho'/\rho \sim \text{const}$  for  $T < m_{DM}$

# Reannihilation regime: Boltzmann equation

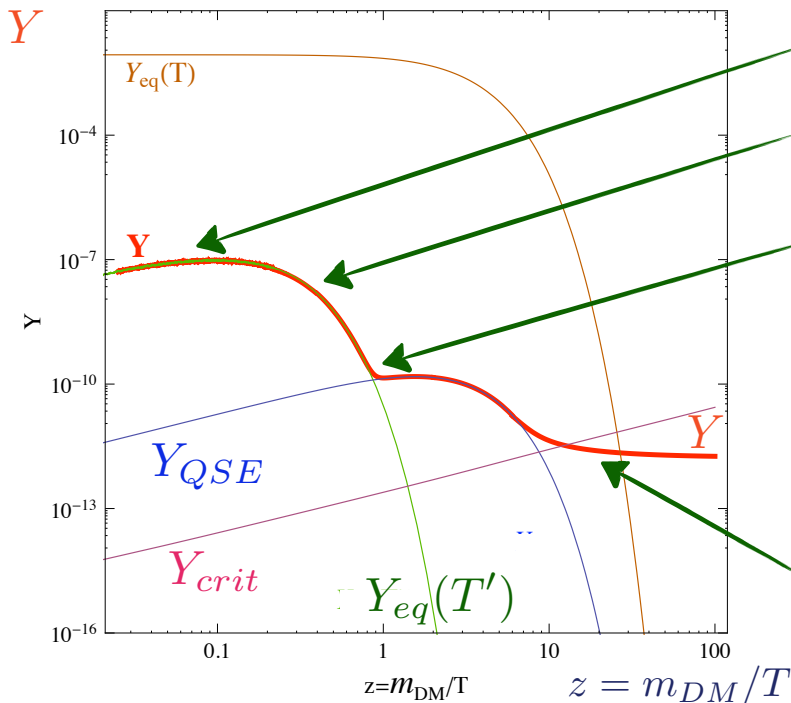
HS process in thermal equilibrium but not the connector:  $T'/T \ll 1$

$$z \frac{H}{s} \frac{dY}{dz} = \sum_i \langle \sigma_{connect} v \rangle_i (Y_{eq}^2(T) - \cancel{Y^2}) + \langle \sigma_{HS} v \rangle (Y_{eq}^2(T') - Y^2)$$

$SM_i SM_i \rightarrow DMDM$        $DMDM \rightarrow SM_i SM_i$        $\gamma' \gamma' \rightarrow DMDM$        $DMDM \rightarrow \gamma' \gamma'$

Cheung, Elor, Hall, Kumar 10'  
 ↪ in  $A \rightarrow DMB$  decay context

Region II ( $m_{DM} = 200 \text{ GeV}, \kappa = 10^{-8}, \alpha' = 10^{-2.4}$ )



after thermalization  $Y$  follows  $Y_{eq}(T')$

at  $T' \lesssim m_{DM}$ :  $Y_{eq}(T')$  becomes Boltzmann suppressed

$SMSM \rightarrow DMDM$  rate (which decouples only at  $T \lesssim m_{DM}$ ) gets larger than the  $\gamma' \gamma' \rightarrow DMDM$  rate (which decouples already at  $T' \lesssim m_{DM}$ )  $\Rightarrow$  reannihilation





at  $T \lesssim m_{DM}$  the  $SMSM \rightarrow DMDM$  source term gets Boltzmann suppressed  $\Rightarrow$  freezes

$$\Gamma_{annih} = H \leftrightarrow Y = Y_{crit} \equiv H / \langle \sigma_{HS} v \rangle$$

# Reannihilation regime: Boltzmann equation

 when the  $\gamma'\gamma' \rightarrow DMDM \propto Y_{eq}^2(T')$  rate goes below the  $SM_i SM_i \rightarrow DMDM \propto Y_{eq}^2(T)$  rate:
   
 in thermal equilibrium
  out of thermal equilibrium

$$z \frac{H}{s} \frac{dY}{dz} = \sum_i \langle \sigma_{connect} v \rangle_i (Y_{eq}^2(T) - \cancel{Y^2}) + \langle \sigma_{HS} v \rangle (Y_{eq}^2(T') - \cancel{Y^2})$$

  $SM_i SM_i \rightarrow DMDM$ 
  $DMDM \rightarrow SM_i SM_i$ 
  $\gamma'\gamma' \rightarrow DMDM$ 
  $DMDM \rightarrow \gamma'\gamma'$

$$z \frac{dY}{dz} = \frac{\langle \sigma_{connect} v \rangle_s}{H} Y_{eq}^2(T) - \frac{Y^2}{Y_{crit}}$$


---

 period of Quasi Static Equilibrium where both terms compensates each other

$$\Rightarrow Y = Y_{QSE} \equiv \sqrt{Y_{crit} \frac{\langle \sigma_{connect} v \rangle_s}{H} Y_{eq}^2(T)}$$

 until  $Y = Y_{QSE} = Y_{crit}$  where Y freeze  at  $T \equiv T_f$

 simultaneous freezing of both connector source term and HS reannihilation term



# Reannihilation regime: analytical result

in practice: if the HS thermalize but the connector does not  
one has always a period of reannihilation

← (except through Z decay  
for  $m_{DM} \lesssim m_Z/2$ )

Freeze-out equation to determine  $T_f$ :

$$n_{eq}(T_f) \langle \sigma_{eff} v \rangle = H(T_f) \quad \leftarrow \langle \sigma_{eff} v \rangle \equiv \sqrt{\langle \sigma_{connect} v \rangle \langle \sigma_{HS} v \rangle}$$

↪ ordinary freeze-out equation but with another cross section:  $\langle \sigma_{eff} v \rangle$

$$\Rightarrow x_f = \log\left[0.038 \frac{g_{e'}}{\sqrt{g_*^{eff}}} m_{Pl} m_{DM} \langle \sigma_{eff} v \rangle c(c+2)\right] \quad x_f \equiv m_{DM}/T_f$$

$$- \frac{1}{2} \log\left[\log\left[0.038 \frac{g_{e'}}{\sqrt{g_*^{eff}}} m_{Pl} m_{DM} \langle \sigma_{eff} v \rangle c(c+2)\right]\right]$$

$$\Rightarrow Y(T_f) \equiv Y_{QSE}(T_f) = \frac{1}{\langle \sigma_{HS} v \rangle} \frac{3.79 x_f}{(g_{*s}/\sqrt{g_*^{eff}})} m_{Pl} m_{DM}$$

$$\Rightarrow \Omega_{DM} \propto \frac{x_f}{\langle \sigma_{HS} v \rangle} \text{ as in usual freeze-out but with a } x_f \text{ determined by another cross section}$$

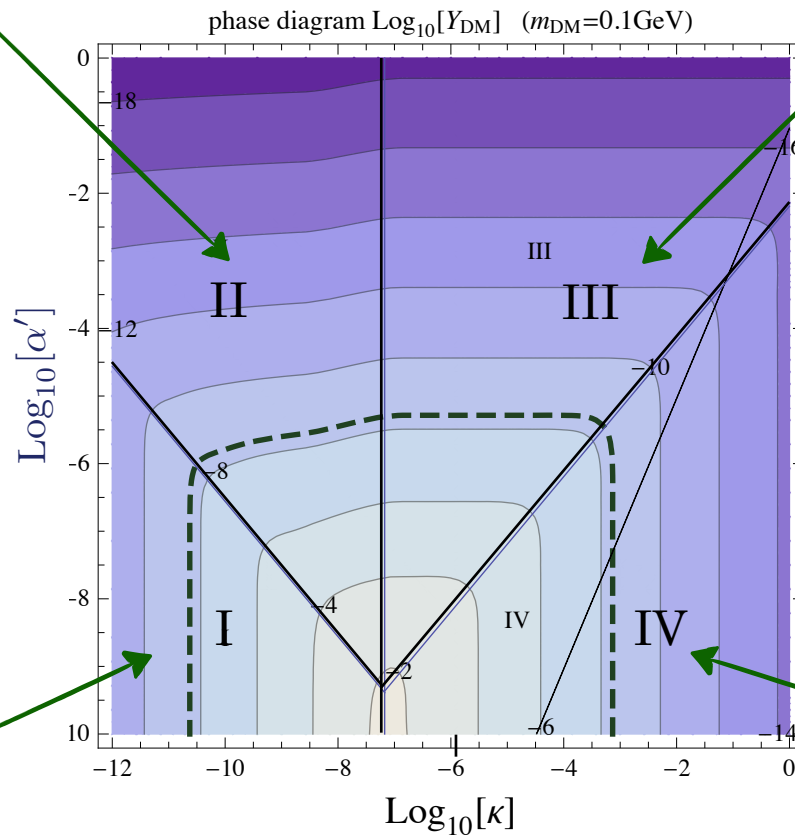
$$\hookrightarrow \propto \log(\alpha' \kappa) / \alpha'^2$$

$$\hookrightarrow \langle \sigma_{eff} v \rangle$$

# Relic density phase diagram

Reannihilation  
regime

Hidden sector  
freeze-out regime



$$\Omega_{DM} \sim 23\%$$

Freeze-in  
regime

Connector  
freeze-out regime

# Hidden sector freeze-out regime

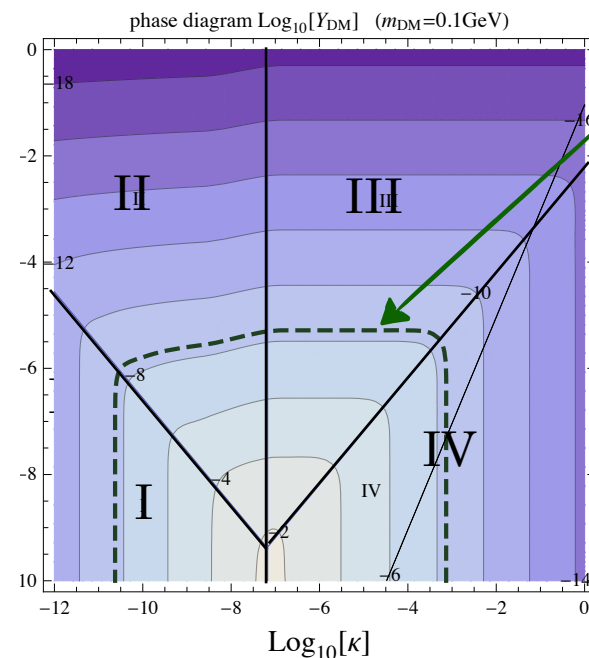
starting from a reannihilation situation  
 if one increases  $\kappa$  further  $\Rightarrow$  the connector interaction thermalizes:  $T' = T$

$SMSM \leftrightarrow DMDM$

one enters a regime where even if the connector thermalizes  
 the HS interaction thermalizes much more  $\langle \sigma_{HS\nu} \rangle > \langle \sigma_{connect\nu} \rangle$

standard freeze-out (only one temperature) dominated by the HS interaction

$$\Rightarrow \Omega_{DM} \propto \frac{1}{\langle \sigma_{HS\nu} \rangle} \propto \frac{1}{\alpha'^2}$$

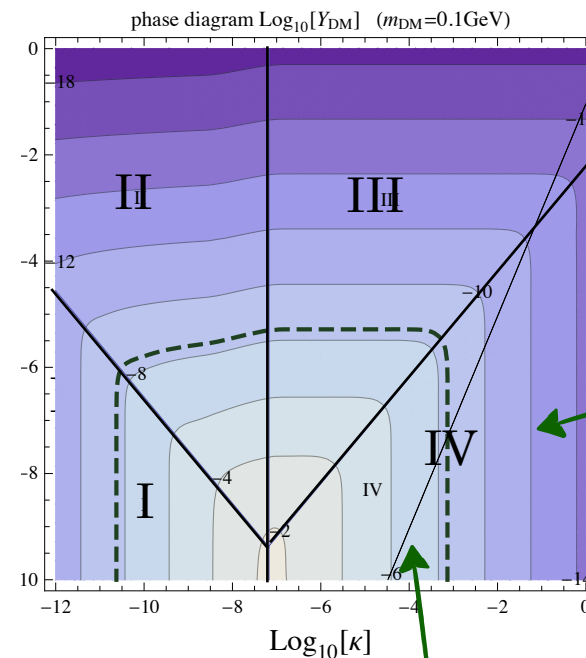


Hidden sector  
 freeze-out regime

# Connector freeze-out regime

if one increases  $\kappa$  further  $\Rightarrow$  the connector not only thermalizes but dominates the freeze-out process  $\langle \sigma_{connect} v \rangle > \langle \sigma_{HS} v \rangle$

$$\Omega_{DM} \propto \frac{1}{\langle \sigma_{connect} v \rangle} \propto \frac{1}{\kappa^2}$$

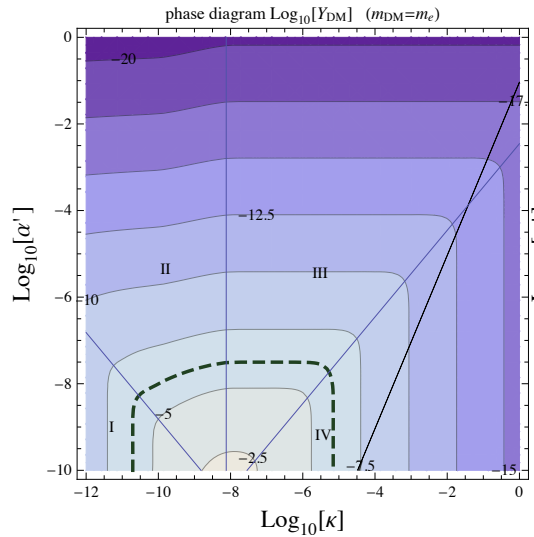


Connector freeze-out regime

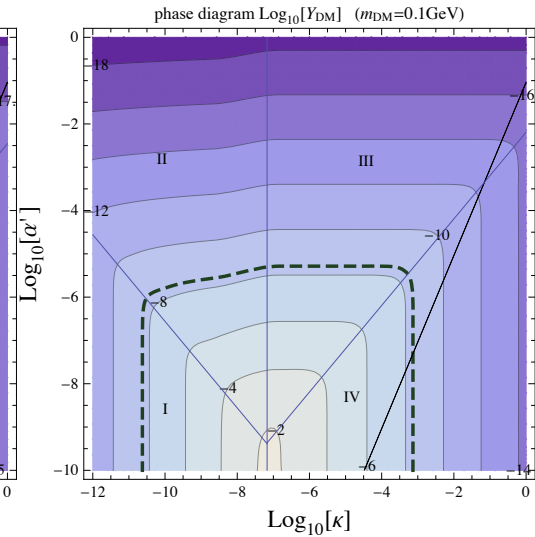
$\kappa \equiv \epsilon \sqrt{\alpha'/\alpha} \Rightarrow$  if  $\kappa$  big  $\alpha'$  small  $\Rightarrow \epsilon$  non perturbative

# Relic density phase diagram

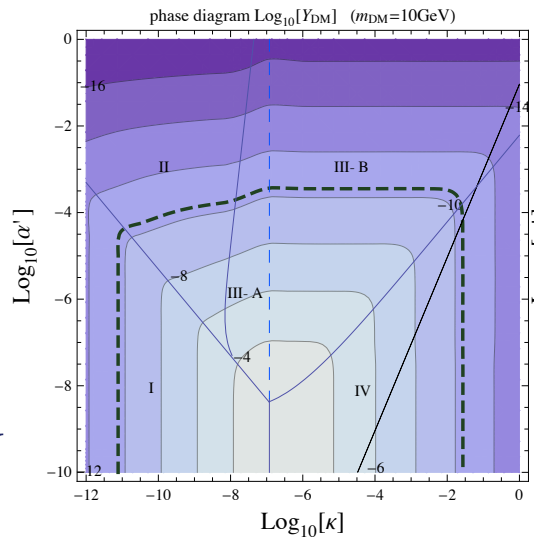
$$m_{DM} = m_e$$



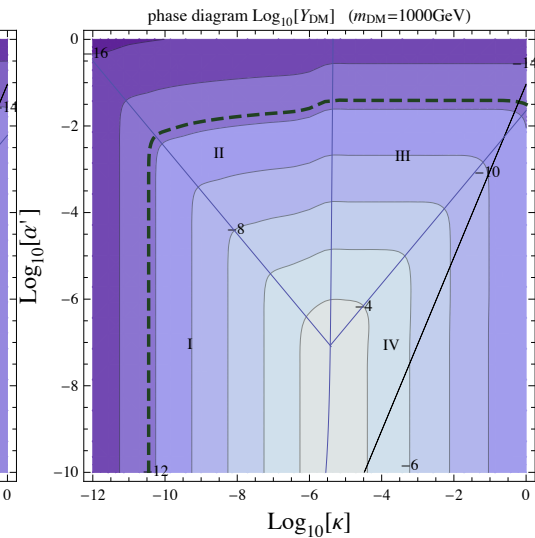
$$m_{DM} = 0.1 \text{ GeV}$$



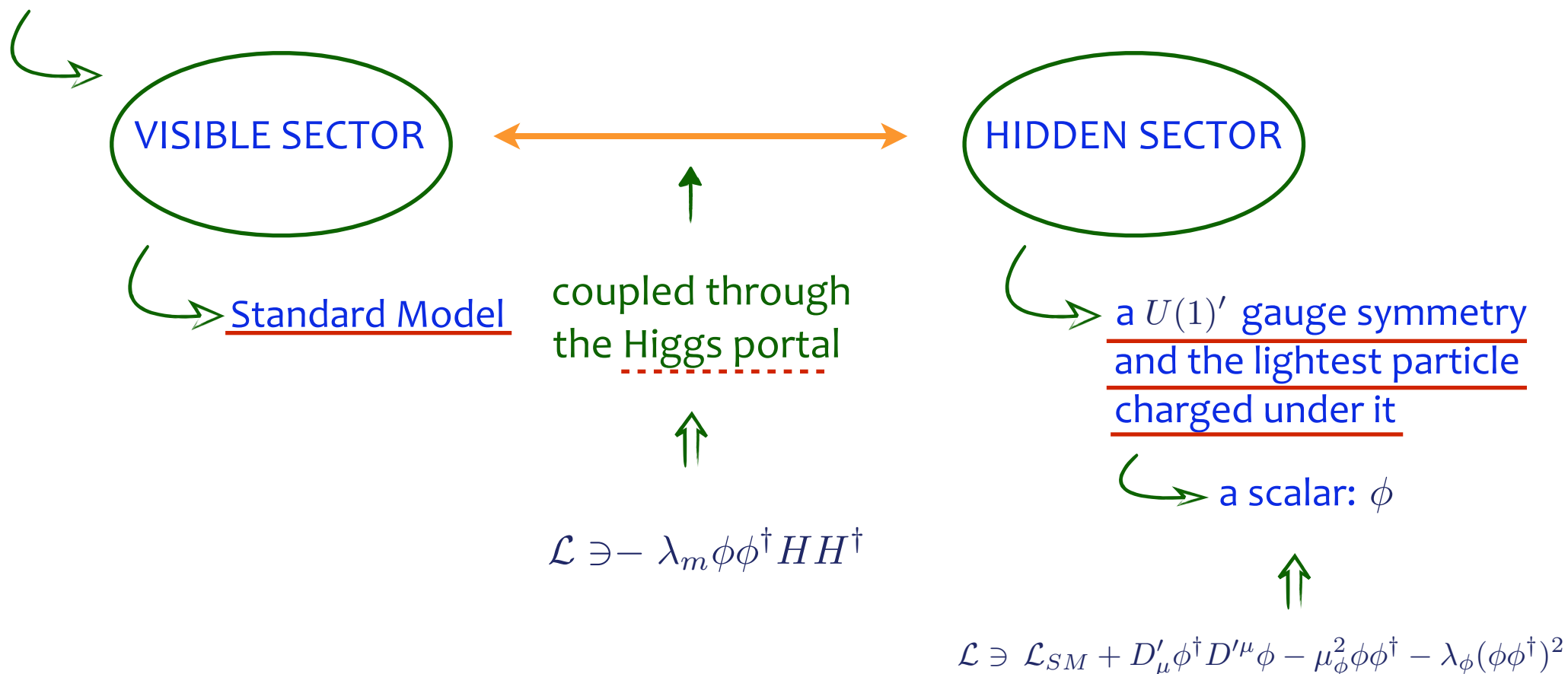
$$m_{DM} = 10 \text{ GeV}$$



$$m_{DM} = 1 \text{ TeV}$$

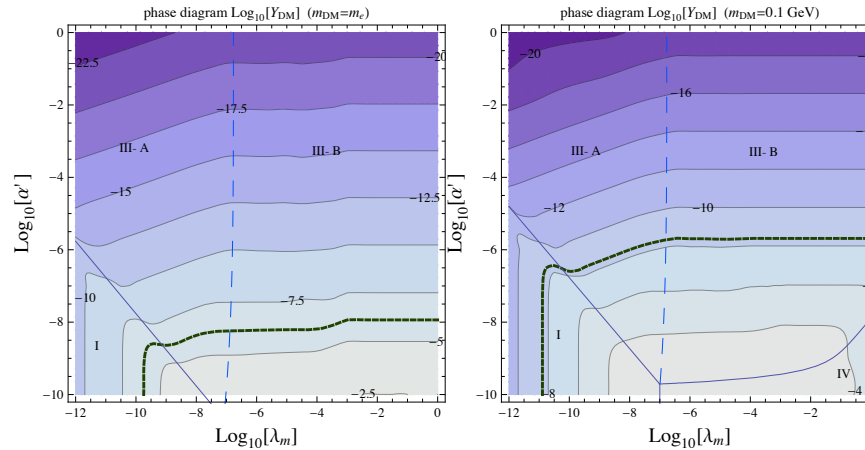


# Generality of the “mesa” phase diagram: the Higgs portal example



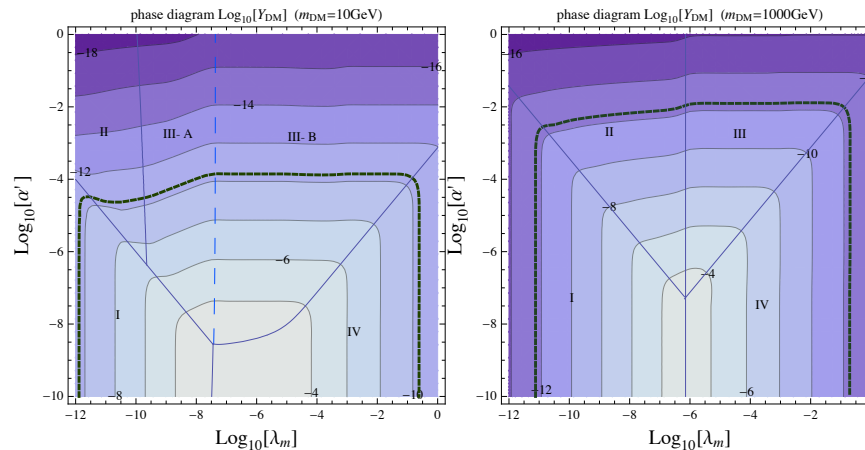
# Generality of the “mesa” phase diagram: the Higgs portal example

$$m_{DM} = m_e$$



$$m_{DM} = 0.1 \text{ GeV}$$

$$m_{DM} = 10 \text{ GeV}$$



$$m_{DM} = 1 \text{ TeV}$$

⇒ same general structure despite of important differences: the mediator is massive

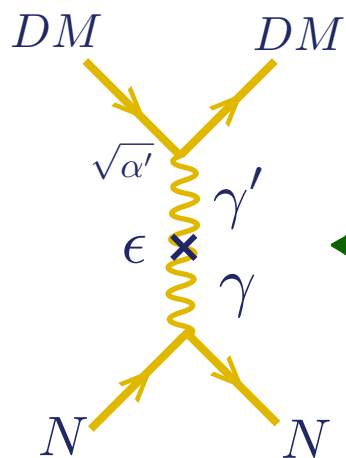
↪ production through  $m_h \rightarrow \phi\phi^\dagger$  decay if  $m_{DM} < m_h/2$   $m_h \simeq 120 \text{ GeV}$

visible/hidden sector communication cut-off at  $T \sim \text{Max}[m_h, m_{DM}]$

⇒ if the HS thermalizes but not the connector: both reannihilation and HS freeze-out possible

# Test of mesa phase diagrams for kinetic mixing: direct detection

DM elastic cross section on Nucleon



$$\frac{1}{q^2} \rightarrow \frac{1}{E_r^2}$$



$$\frac{d\sigma}{dE_r} = \frac{1}{E_r^2} \frac{1}{v^2} \frac{2\pi\kappa^2 Z^2 \alpha^2}{m_A} F_A^2(qr_A)$$

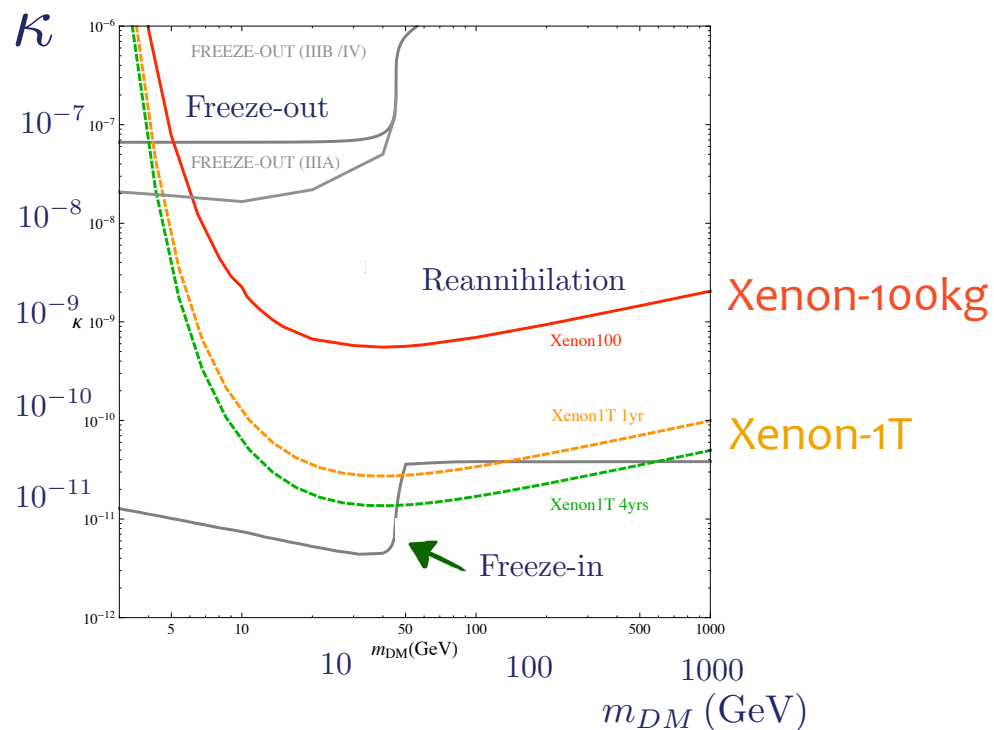
$E_r \sim \text{few KeV}$

huge enhancement

direct detection sensitive to very small  $\kappa$  values



# Test of mesa phase diagrams for kinetic mixing: direct detection



Xenon-100kg: excludes all regimes for  $m_{DM} > \text{few GeV}$  except freeze-in and part of reannihilation

Xenon-1T: will test freeze-in for

$50 \text{ GeV} < m_{DM} < 140 \text{ GeV}$  ← 1 T/year

$40 \text{ GeV} < m_{DM} < 600 \text{ GeV}$  ← 4 T/year

↪ characteristic  $\sim \frac{1}{E_r^2}$  spectrum to be observed!!

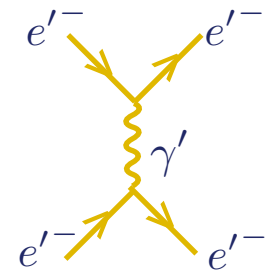
# Cosmological constraints

↪ associated to new long range  $U(1)'$  force

- BBN
- bullet cluster
- galactic dynamics
- ...

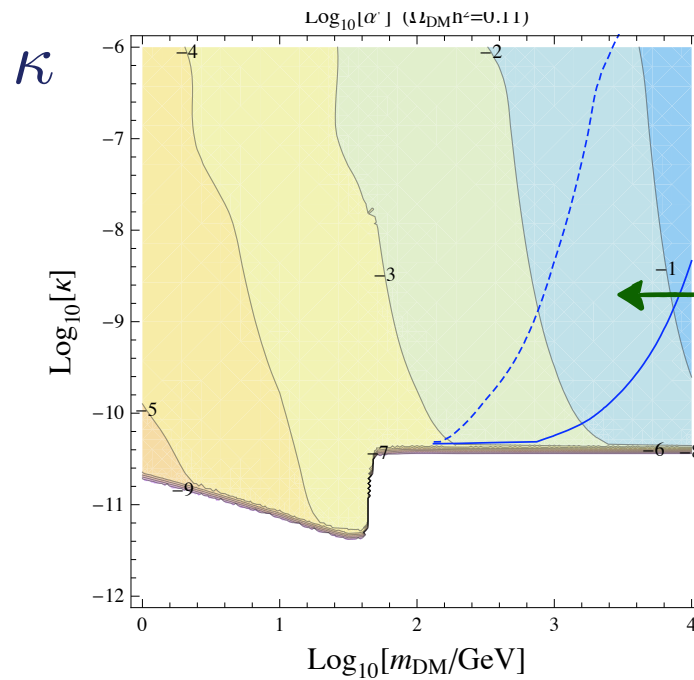
↪ DM Rutherford scattering may affect formation of DM halo

↪ ellipticity of galaxies:  $\alpha' \lesssim 10^{-7} (m_{DM}/\text{GeV})^{3/2}$



Ackerman, Buckley,  
Carroll, Kamionkowski 08'  
Feng Kaplinghat, Tu, Yu 09'  
Feng, Tu, Yu 08

# Ellipticity and relic density constraints combined



reannihilation allowed for:

$$m_{DM} > \sim \text{few } 100 \text{ GeV}$$

freeze-in always allowed

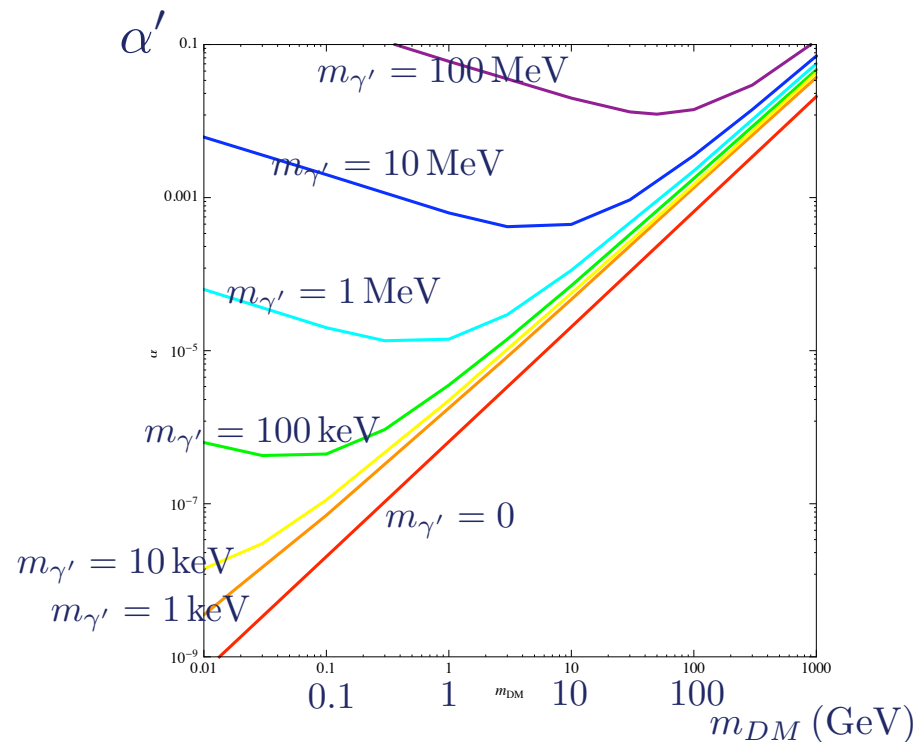
# Ellipticity constraint for the case of a slightly massive $\gamma'$

if we break the  $U(1)'$  slightly with  $m_{\gamma'} \ll m_{DM}$

the relic density plot doesn't change

the lightest charged fermion remains stable

but the cosmological constraints change a lot



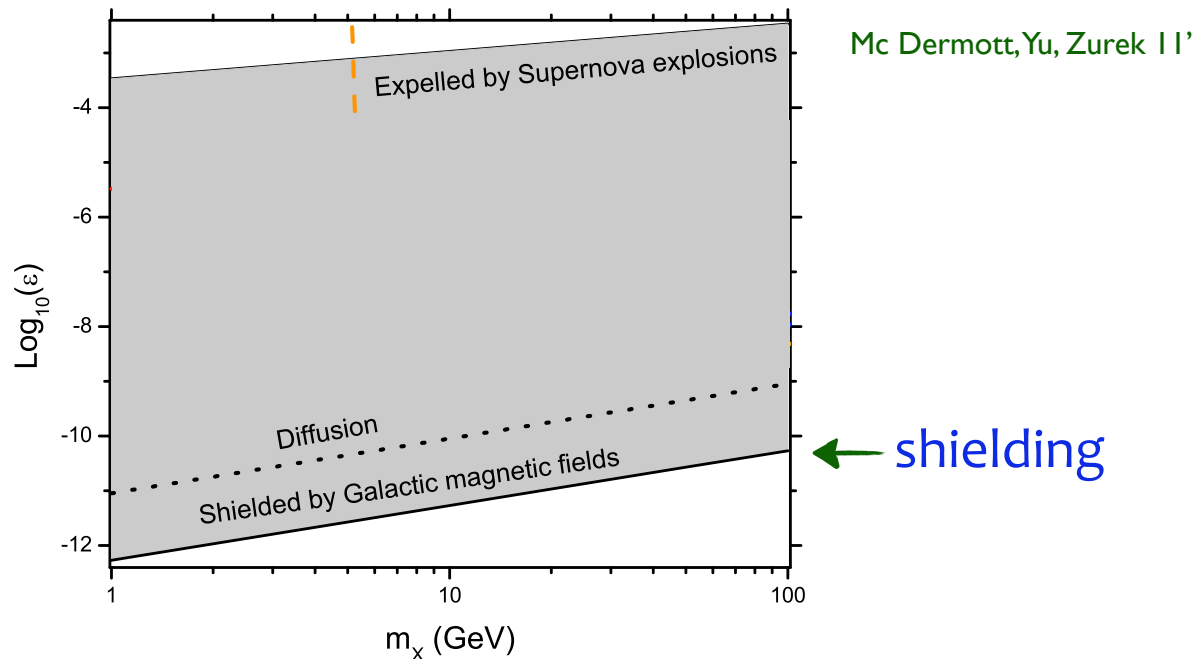
# Depletion of DM in galactic disk

→ the DM feels the galactic magnetic field via  $\kappa$

→ sufficient for a DM coming off the disk not to enter in the disk

Chuzhoy, Kolb 09'


Mc Dermott, Yu, Zurek 11'



⇒ only freeze-in regime is allowed for  $m_{DM} \lesssim 100 \text{ GeV}$  but the constraint vanishes as soon as the  $\gamma'$  becomes slightly massive

# Summary

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- Visible/hidden sectors/mediator structure:  
the observed relic density can be produced through characteristic 4 regimes  
“mesa” phase diagram...  
 natural “analytic prolongation” of the usual freeze-out regime towards small coupling values
- Kinetic mixing portal:
  - all 4 regimes can be tested from direct detection, even the freeze-in one
  - rich cosmological phenomenology (which strongly depends on the mass of the  $\gamma'$ )

**Backup**

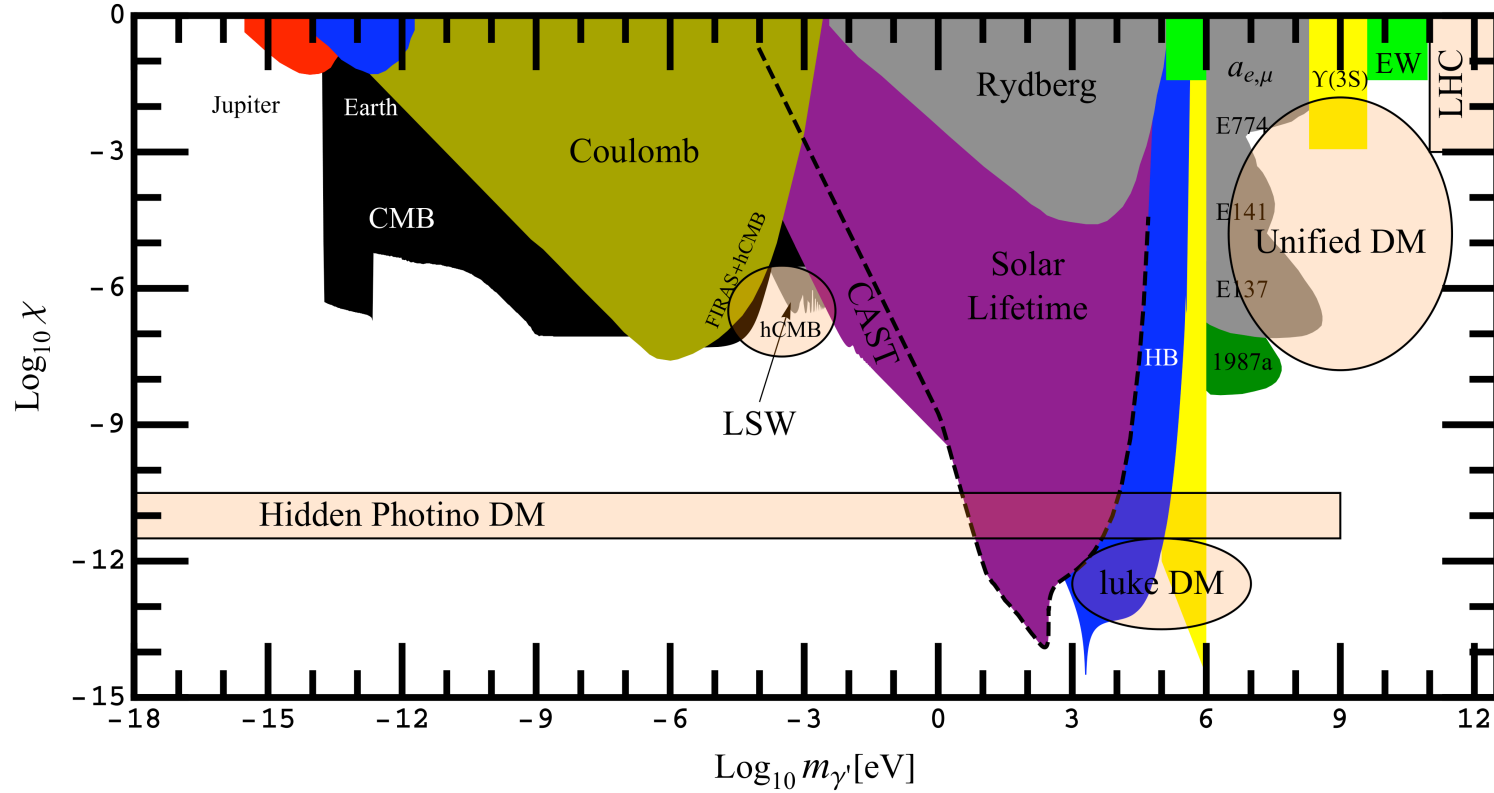


Figure 10. Summary of astrophysical, cosmological and laboratory constraints for hidden photons (kinetic mixing  $\chi$  vs. mass  $m_{\gamma'}$ ). At higher mass we have electroweak precision measurements (EW), bounds from upsilon decays ( $\Upsilon_{3S}$ ) and fixed target experiments (EXXX). Areas that are especially interesting are marked in light orange. Compilation from Ref. [93].



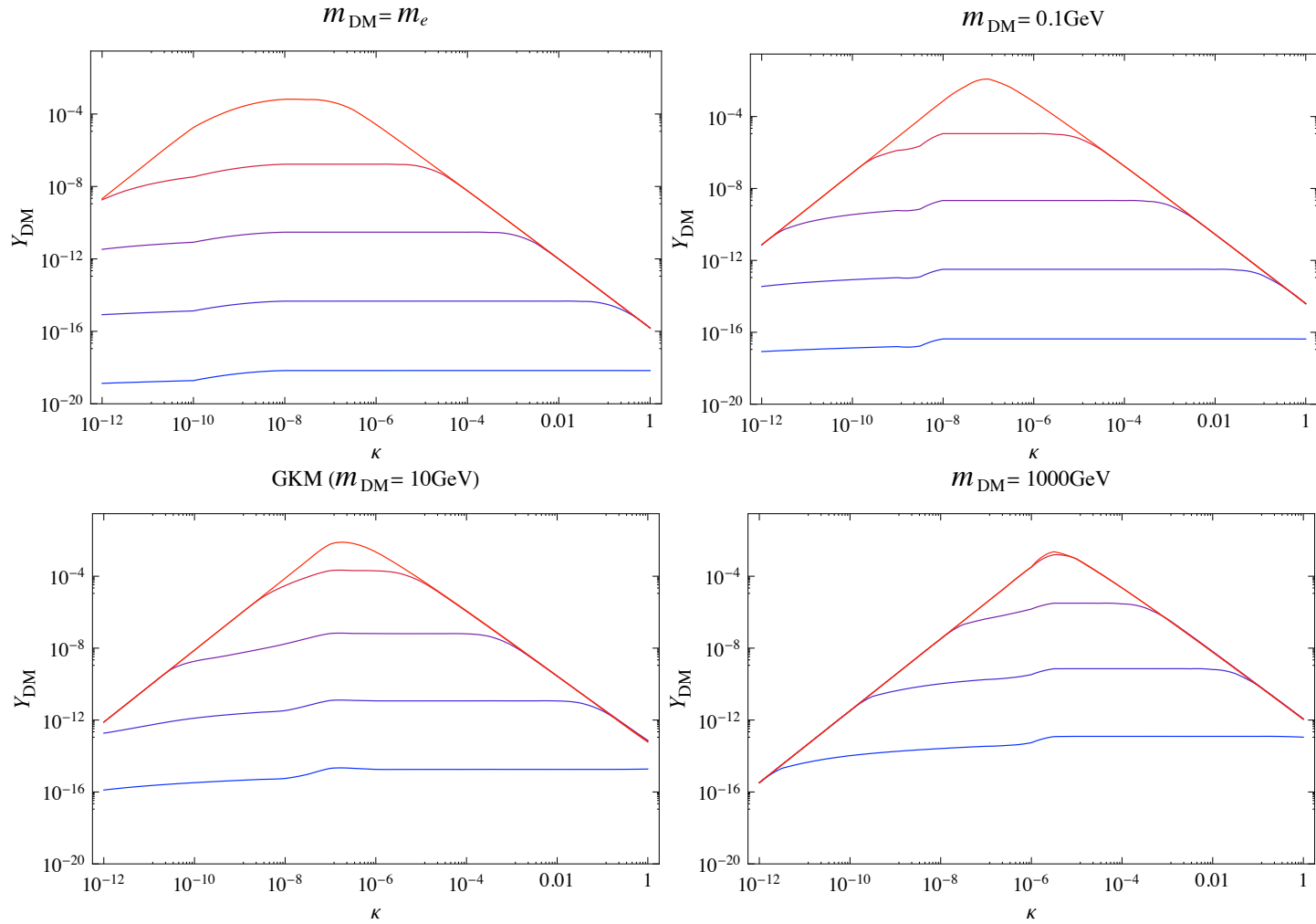
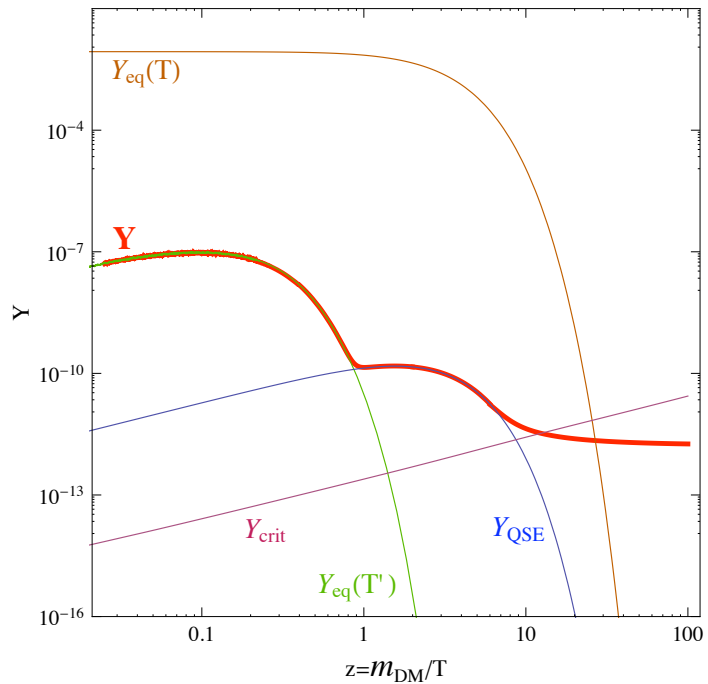
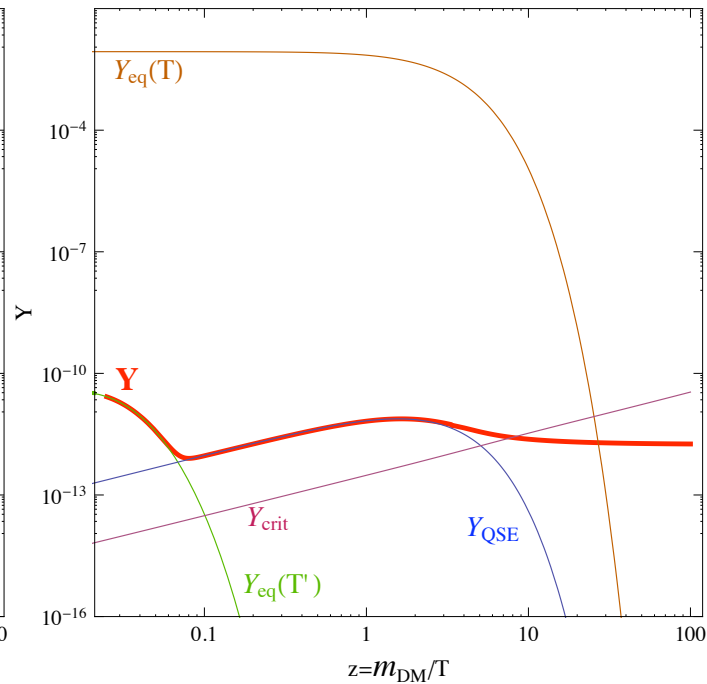


Figure 5: DM relic abundance  $Y_{DM}$  as a function of the connector parameter  $\kappa$  for different DM masses  $m_{DM}$  and values of the hidden sector interaction,  $\log_{10}(\alpha'/\alpha) = 1, -1, -3, -5, -7$  (bottom-up).

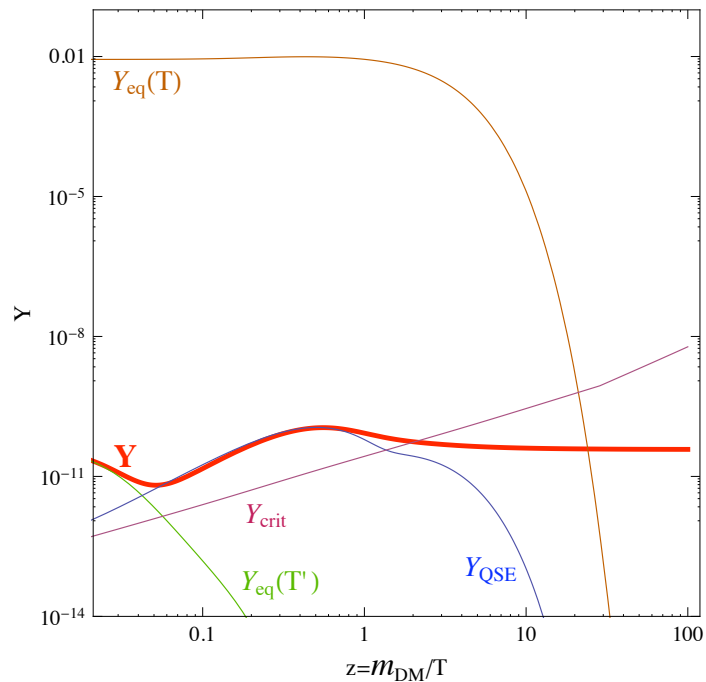
Region.II ( $m_{\text{DM}}=200\text{GeV}$ ,  $\kappa=10^{-8}$ ,  $\alpha'=10^{-2.4}$ )



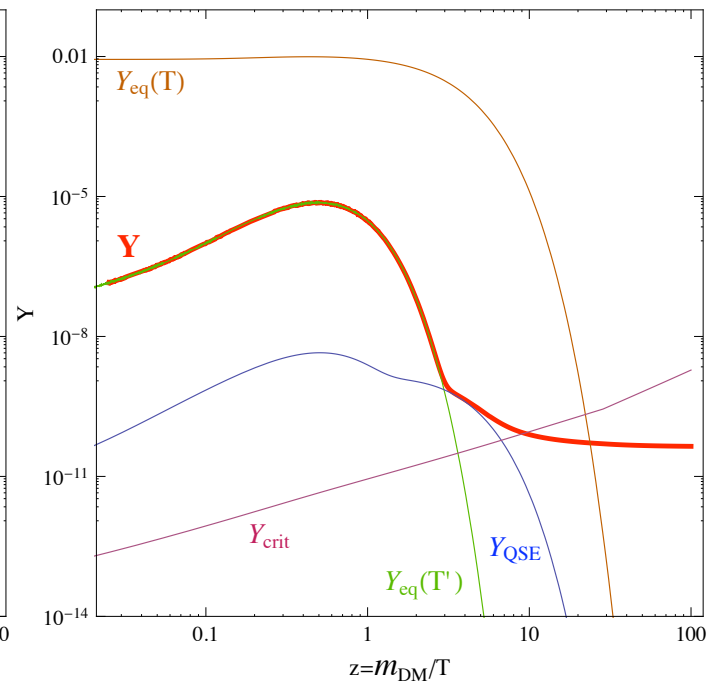
Region.II ( $m_{\text{DM}}=200\text{GeV}$ ,  $\kappa=10^{-9.5}$ ,  $\alpha'=10^{-2.6}$ )



Region.II-A ( $m_{\text{DM}}=10\text{GeV}$ ,  $\kappa=10^{-10.5}$ ,  $\alpha'=10^{-4.2}$ )



Region.II-B ( $m_{\text{DM}}=10\text{GeV}$ ,  $\kappa=10^{-8.5}$ ,  $\alpha'=10^{-3.8}$ )



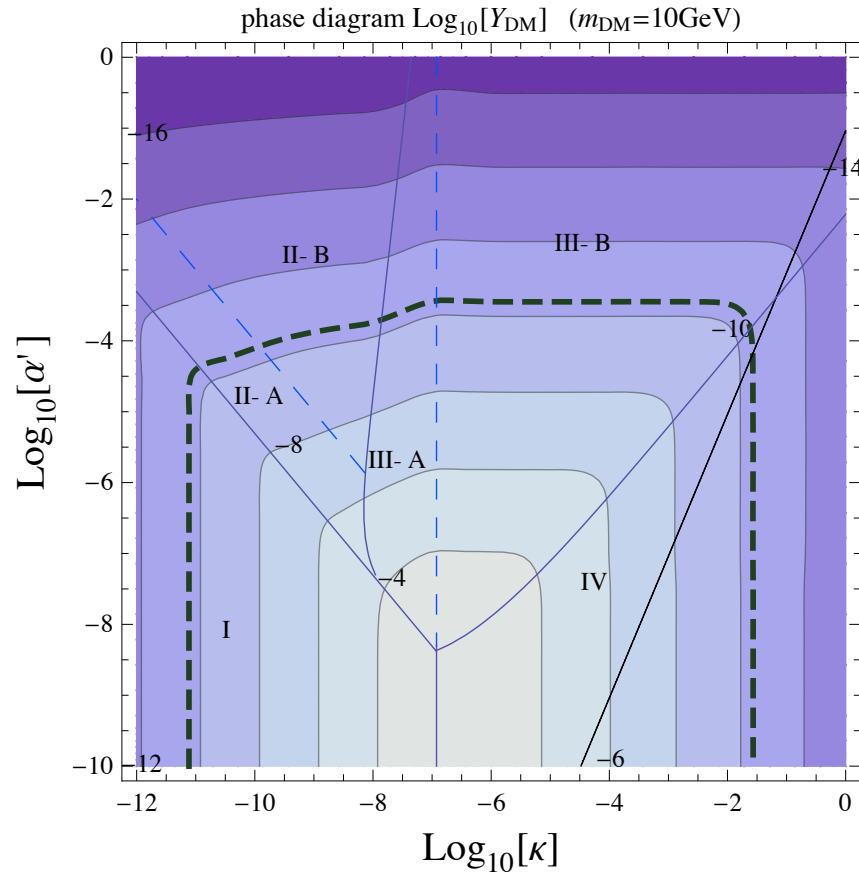


Figure 18: Phase diagram for the kinetic mixing portal and  $m_{DM} = 10$  GeV separating explicitly the reannihilation regimes dominated by the decay (IIA) and by the  $\gamma$  mediated scattering (IIB).

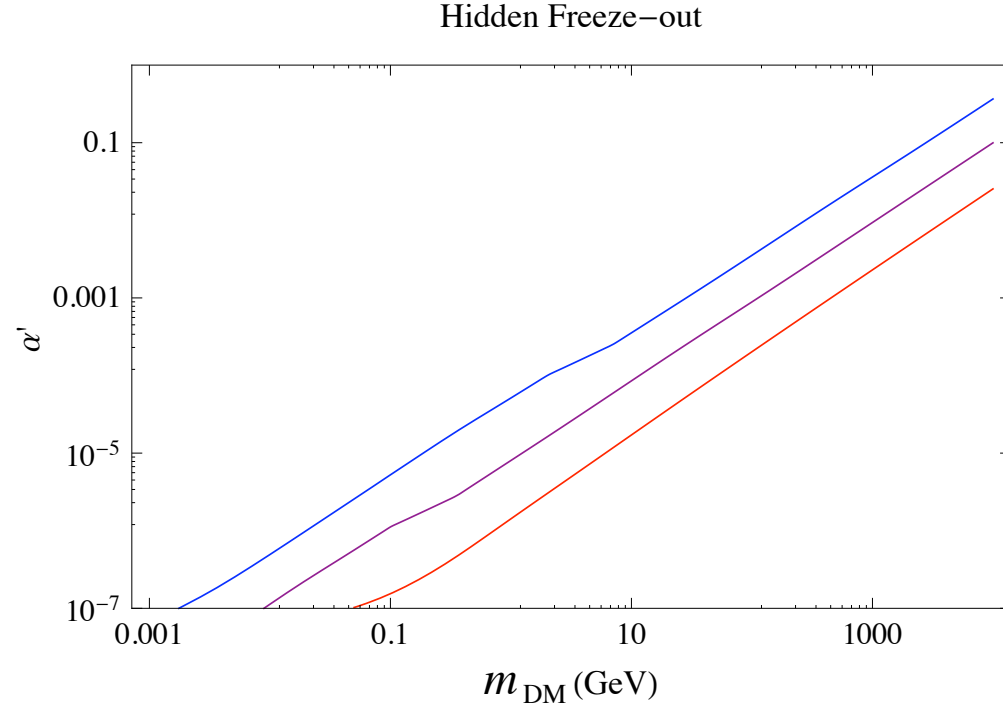


Figure 6: Values of  $\alpha'$  required to get the WMAP relic density as a function of  $m_{DM}$  assuming no connector between the hidden sector and the SM sector, for different values of the temperature ratio  $\xi \equiv T'/T = 0.01, 0.1, 1$  (bottom-up).

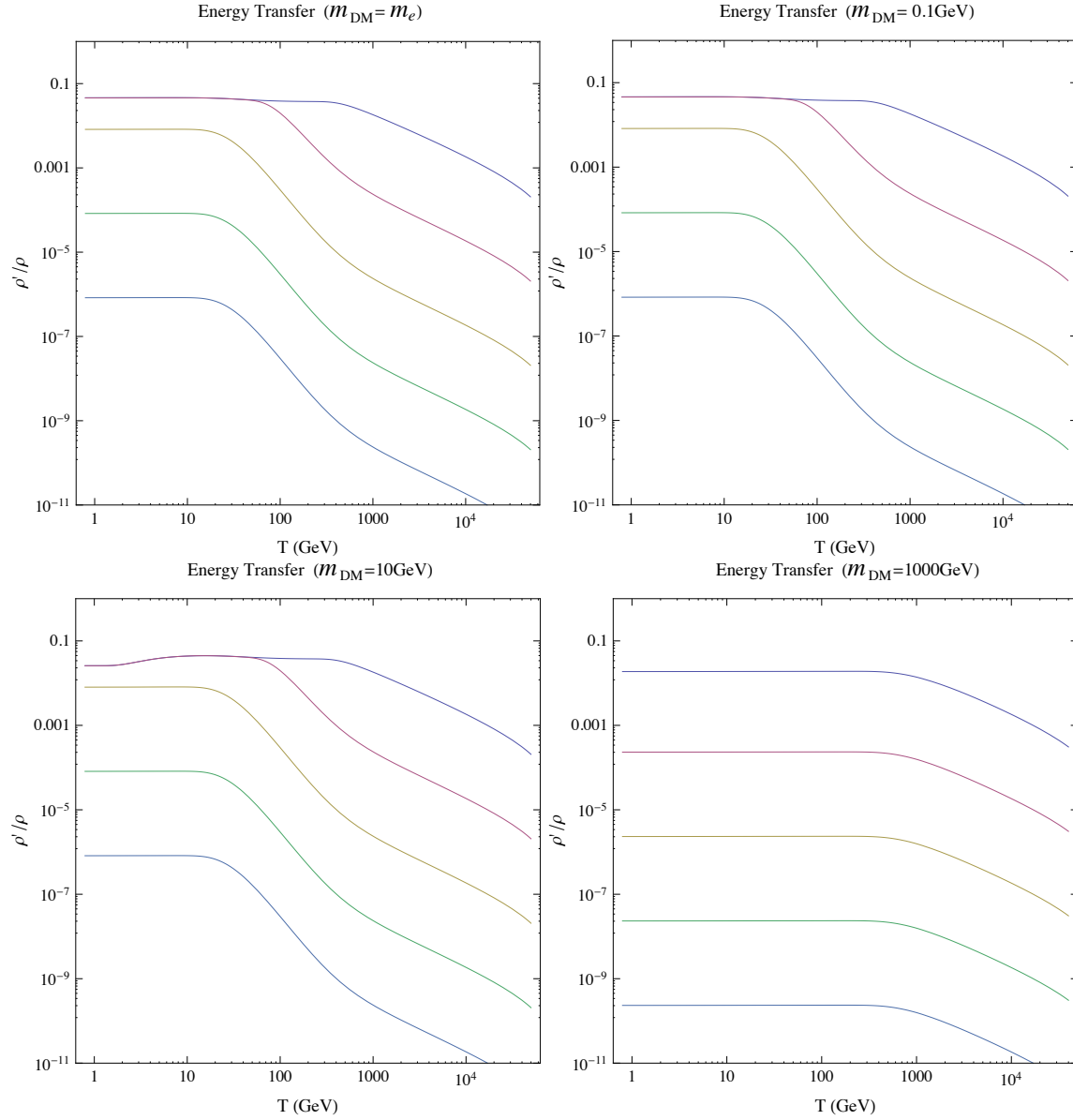


Figure 12: Evolution of the ratio of the visible and hidden sectors energy densities, for a range of connector parameter,  $\lambda_m = 10^{-6, -7, -8, -9, -10}$  (from up to down), and for various DM masses.

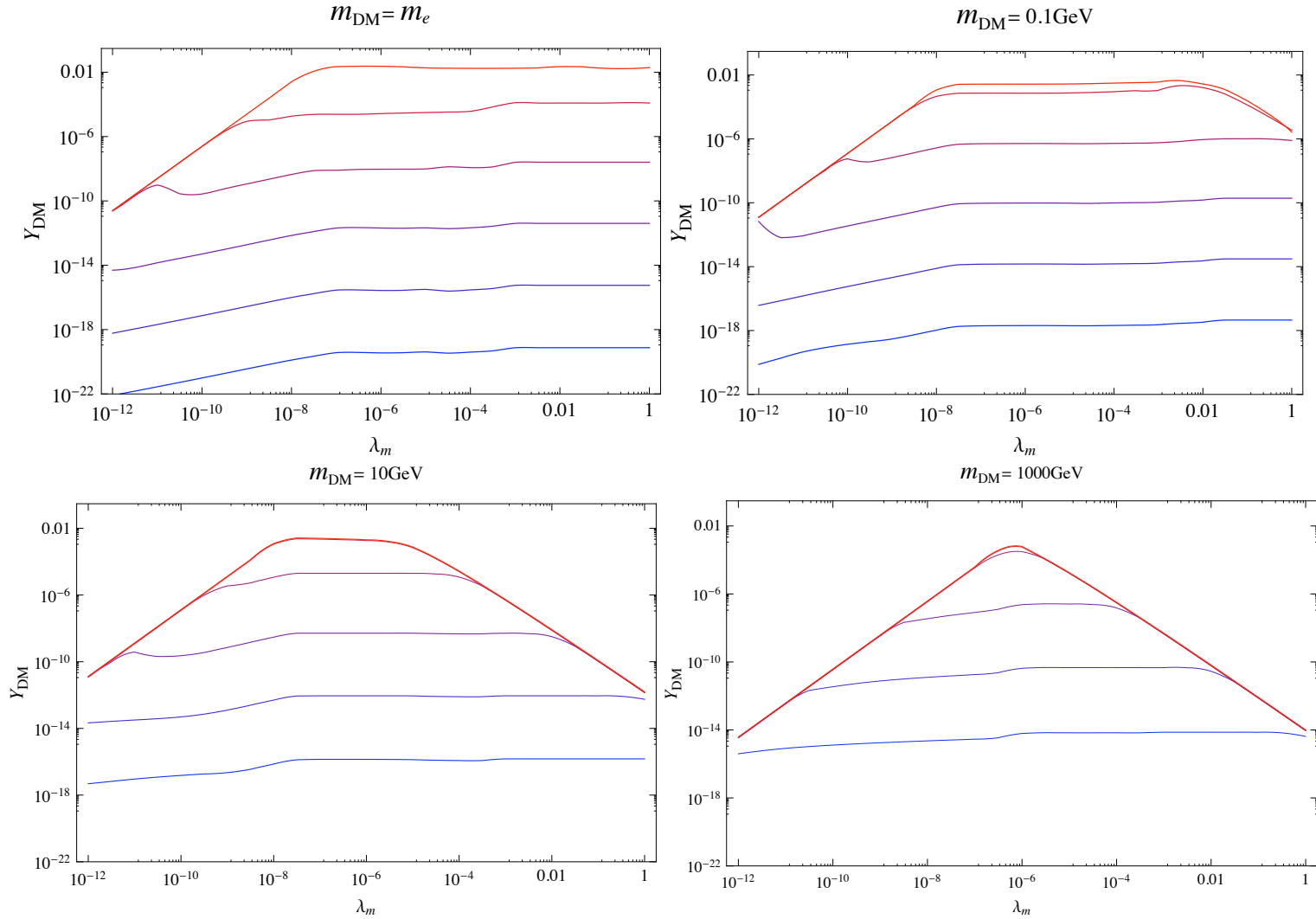


Figure 14: DM relic abundance  $Y_{DM}$  as a function of the connector parameter  $\lambda_m$  for different DM masses  $m_{DM}$  and values of the hidden sector interaction,  $\log_{10}(\alpha'/\alpha) = 1, -1, -3, -5, -7, -9$ , bottom-up (the last two lines are the same for  $m_{DM} = 10 \text{ GeV}$ , as well as for  $m_{DM} = 1 \text{ TeV}$ ).

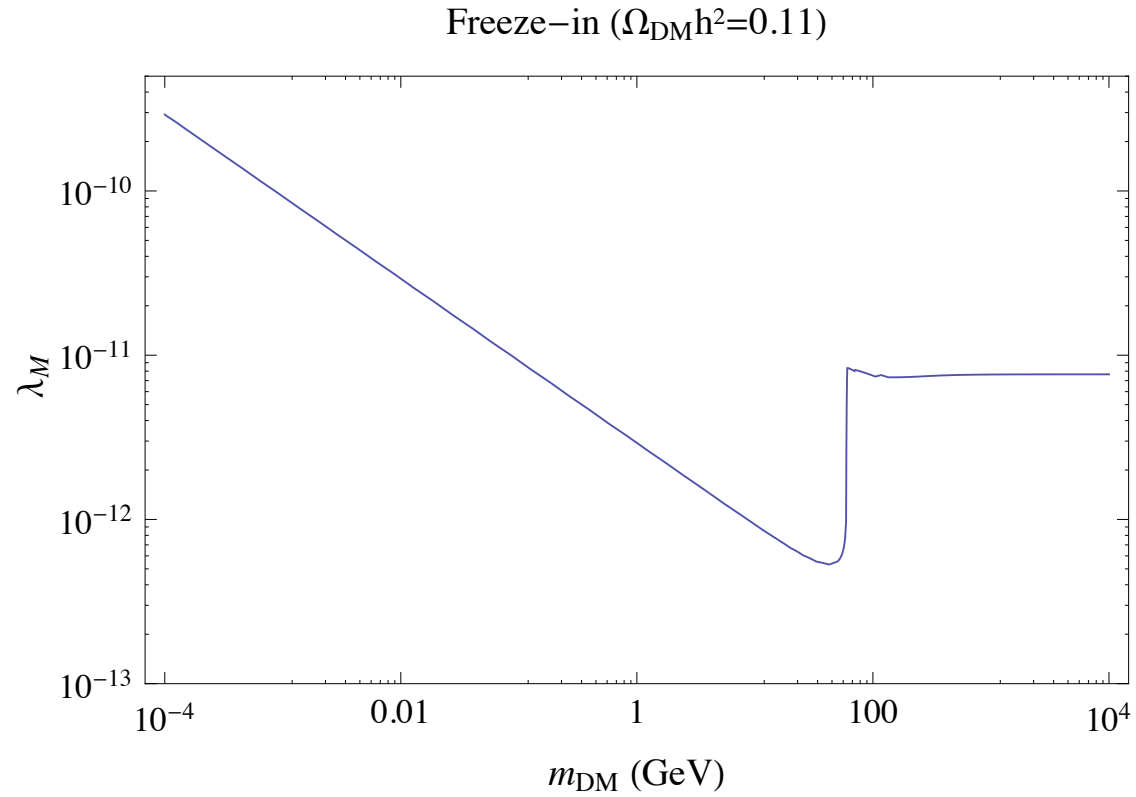
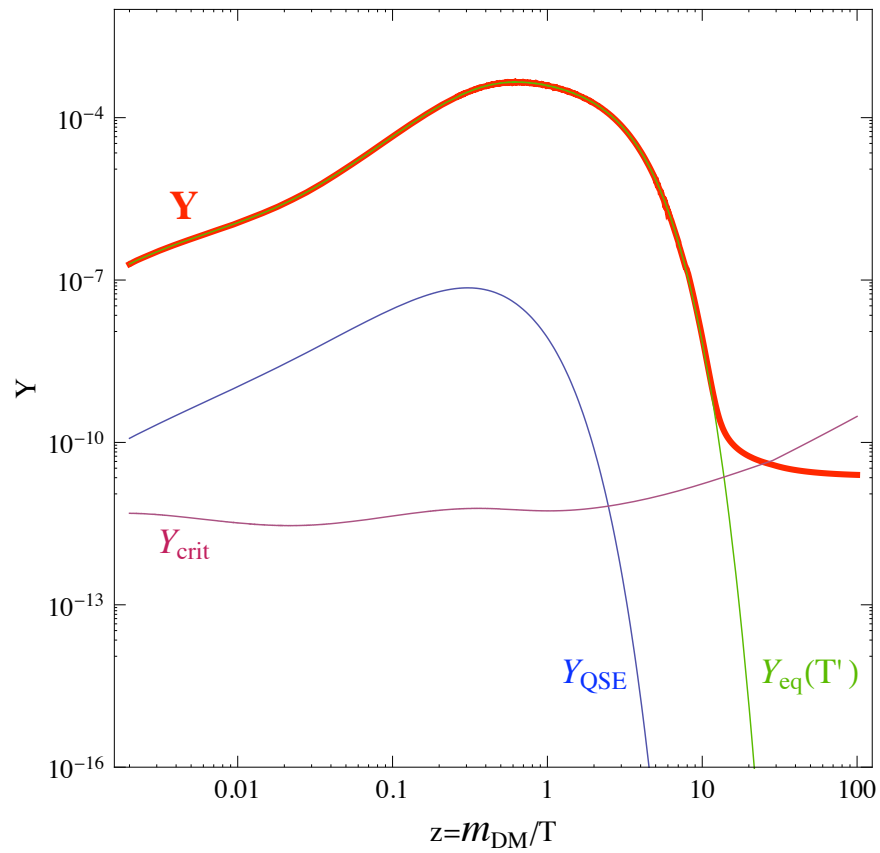
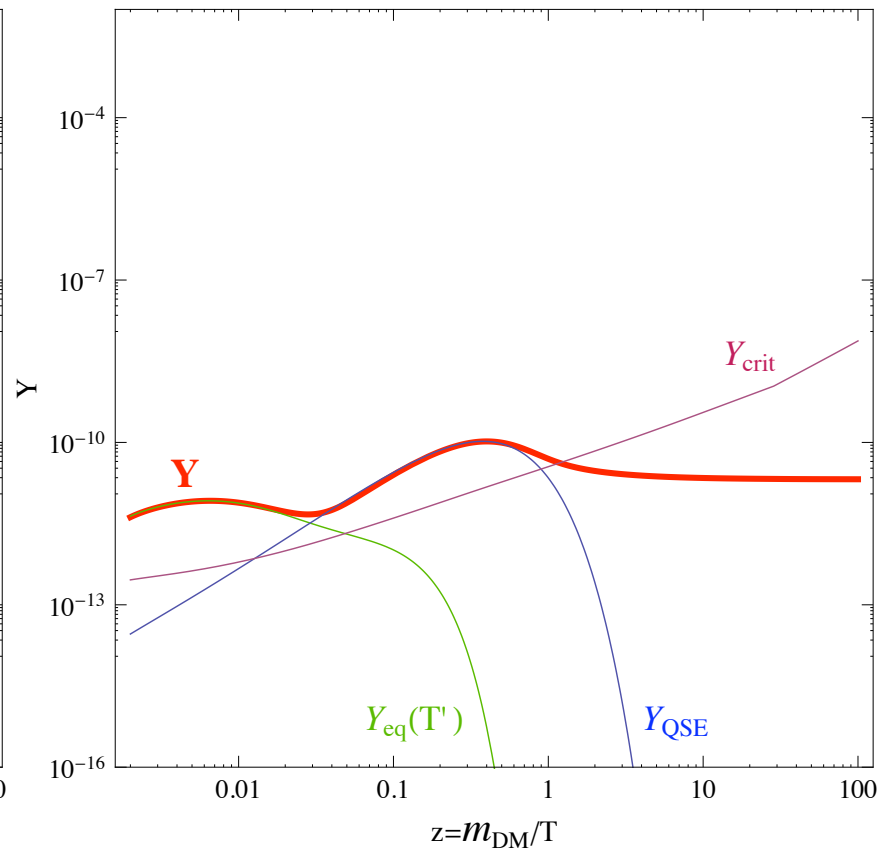


Figure 14: Higgs portal parameter required to get the observed DM relic density through freeze-in ( $\alpha' = 0$ ).

Region.II ( $m_{\text{DM}}=10\text{GeV}, \lambda_m=10^{-8}, \alpha'=10^{-4}$ )



Region.II ( $m_{\text{DM}}=10\text{GeV}, \lambda_m=10^{-11}, \alpha'=10^{-4.7}$ )





# The freeze-in mechanism



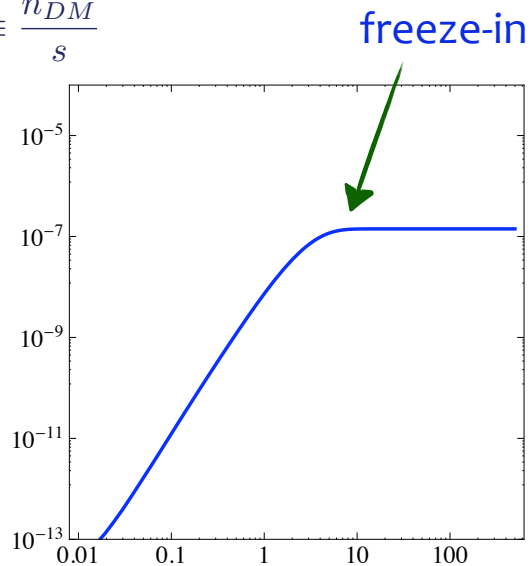
DM couples only feebly to the SM particles  
 production through out-of-equilibrium  $AA \rightarrow DMDM$   
 or  $A \rightarrow DMB$  processes

Mc Donald 02'  
 Hall, Jedamzik,  
 March-Russell, West 09'

example:  $A \rightarrow DMB$

at  $T > m_A$  :  $\frac{dn_{DM}}{dt} = n_A^{eq}(T)\Gamma_{A \rightarrow DMB}(T)$

$Y_{DM} \equiv \frac{n_{DM}}{s}$



$z = m_A/T$



$\frac{dY_{DM}}{dT} = \frac{n_A^{eq}(T)\Gamma_{A \rightarrow DMB}(T)}{TH(T)s(T)} \propto 1/T^2$



$Y_{DM} \propto 1/T$  down to  $T \sim m_A$

in a comoving volume



$Y_{DM} \equiv \frac{n_{DM}}{s}$

where  $n_A^{eq}$  becomes Boltzmann suppressed



DM production IR dominated

# The freeze-in mechanism

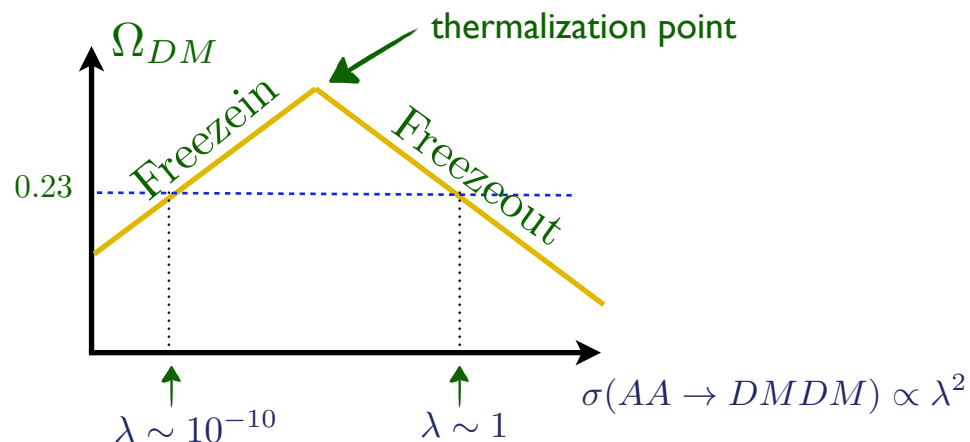
$$\Rightarrow Y_{DM}(T \ll m_A) \simeq \frac{n_A^{eq} \Gamma_{A \rightarrow DMB}}{s} \frac{1}{H} \Big|_{T=m_A}$$

freeze-in is “thermal” in the sense that DM is produced by a A particle in thermal equilibrium

$Y_{DM}$  produced depends only on mass and interactions of particles at freezing

$\Omega_{DM} \sim 23\%$  requires tiny coupling  $\sim 10^{-10}$

$\Rightarrow$  for a  $AA \rightarrow DMB$  scattering production process:



Mc Donald 02'  
Hall, Jedamzik,  
March-Russell, West 09'

# Freeze-in issues

- what about a primordial DM density?

↪ not washed-out by (out-of-equil.) DM production process  
negligible if reheating occurs mostly in one of the feebly coupled sectors

↪ can be tested if the DM mass and coupling measured are the ones which give the right relic density

- what about testing the freeze-in mechanism? ←  $\lambda \sim 10^{-10}$  !!

↪ one possibility:  $A \rightarrow DM B$  decay very slow ⇒ displaced vertex at colliders

⇓  
requires:

Cheung, Elor, Hall, Kumar 10'

- $A$  and  $B$  couple sizably to the SM
- sym. which stabilizes the DM particle also shared by visible sector

↪ e.g.  $A$  and  $DM$  odd under a  $Z_2$

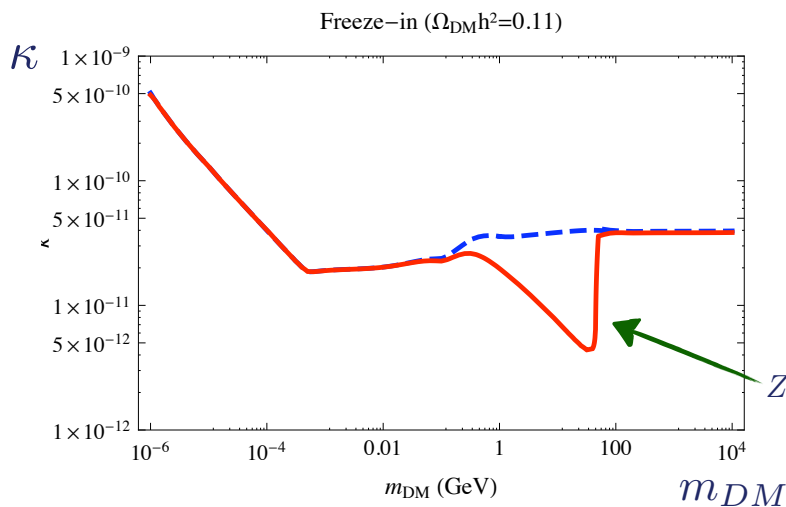
⇒ rich phenomenology at LHC

# $\kappa$ and $\alpha'$ are small $\Rightarrow$ freeze-in regime

we consider a HS negligible at high temperature  $\begin{cases} Y_{DM} \sim 0 \\ \rho' \sim 0 \end{cases}$   
 HS energy density

if  $\kappa$  and  $\alpha'$  small:  $SMSM \leftrightarrow DMDM$  does not thermalize  
 $DMDM \leftrightarrow \gamma' \gamma'$  does not thermalize

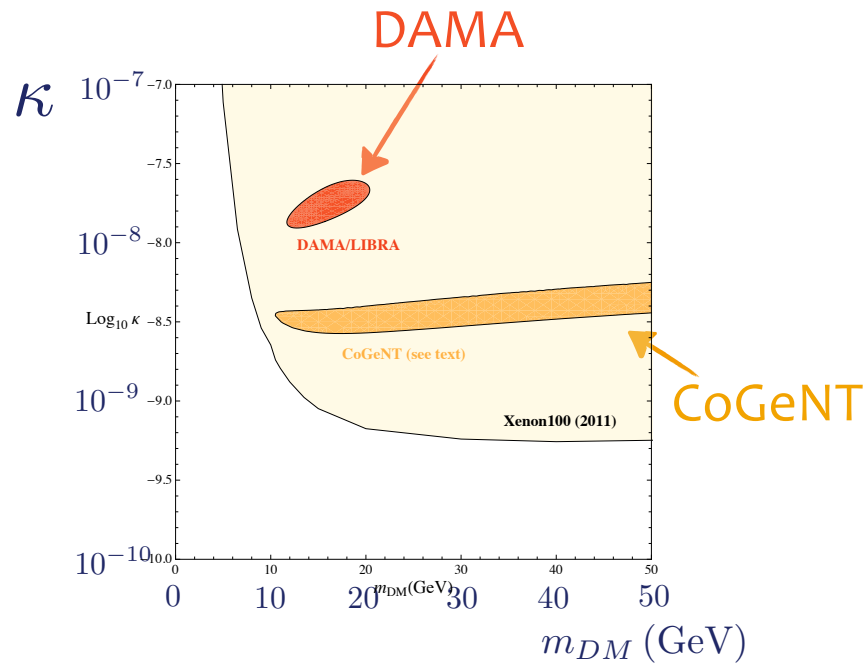
only  $SMSM \rightarrow DMDM$  is relevant because only  $Y_{SM}^{eq}$  is large



freeze-in:  $Y \sim \frac{\langle \sigma_{connect} v \rangle n_{SM}^2}{sH} \Big|_{T=m_{DM}} \propto \kappa^2$   
 $\kappa \sim 3 \cdot 10^{-11}$

dominated by  $m_{SM_i} < m_{DM}$  channels

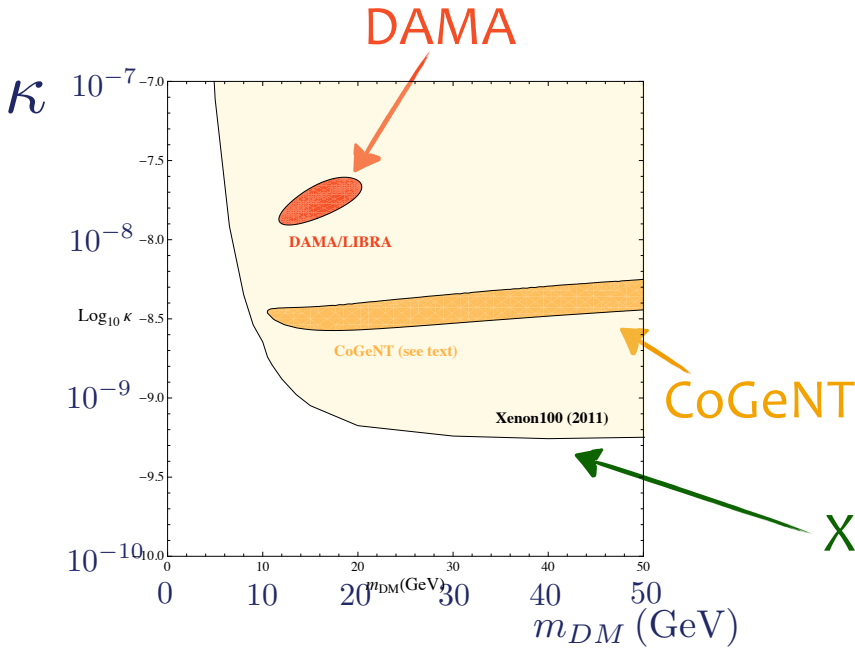
# Test of mesa phase diagrams for kinetic mixing: direct detection



DAMA-CoGeNT: small threshold  
Xenon: higher threshold

$\sim \frac{1}{E_r^2}$  gives much better agreement

# Test of mesa phase diagrams for kinetic mixing: direct detection



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but not compatible anymore

Xenon-100kg

