

Bimetric theory, partial masslessness and conformal gravity

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Based on

- ▶ SFH, Angris Schmidt-May, Mikael von Strauss
arXiv:1203.5283, 1204.5202, 1208:1515, 1208:1797, 1212:4525, 1303.6940
- ▶ SFH, Rachel A. Rosen,
arXiv:1103.6055, 1106.3344, 1109.3515, 1109.3230, 1111.2070

Outline of the talk

Review: Linear and Nonlinear massive spin-2 fields

Ghost-free bimetric theory

Mass spectrum of bimetric theory

Partially Massless bimetric theory

Higher derivative gravity & Conformal gravity

Equivalence between CG and PM bimetric theory

Higher derivative gravity from bimetric theory

Discussion

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Linear massive spin-2 fields

The Fierz-Pauli equation:

Linear massive spin-2 field, $h_{\mu\nu}$, in background $\bar{g}_{\mu\nu}$

$$\bar{\mathcal{E}}_{\mu\nu}^{\rho\sigma} h_{\rho\sigma} - \Lambda \left(h_{\mu\nu} - \frac{1}{2} \bar{g}_{\mu\nu} h_{\rho}^{\rho} \right) + \frac{m_{\text{FP}}^2}{2} \left(h_{\mu\nu} - \bar{g}_{\mu\nu} h_{\rho}^{\rho} \right) = 0$$

[Fierz-Pauli, 1939]

- ▶ 5 propagating modes (massive spin-2)
- ▶ Massive gravity (?)
- ▶ What determines $\bar{g}_{\mu\nu}$? (flat, dS, AdS, ...)
- ▶ Nonlinear generalizations?

[The Boulware-Deser ghost (1972)]

Nonlinear massive spin-2 fields

- ▶ “Massive gravity” (fixed $f_{\mu\nu}$):

$$\mathcal{L} = m_p^2 \sqrt{-g} \left[R - m^2 V(g^{-1} f) \right]$$

- ▶ Interacting spin-2 fields (dynamical g and f):

$$\mathcal{L} = m_p^2 \sqrt{-g} \left[R - m^2 V(g^{-1} f) \right] + \mathcal{L}(\nabla f)(?)$$

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Bimetric: $\mathcal{L}(\nabla f) = m_f^2 \sqrt{-f} R_f(?)$

[Isham-Salam-Strathdee, 1971, 1977]

Generically, both contain a *GHOST* at the nonlinear level

[Boulware-Deser, 1972]

Counting modes:

Generic massive gravity:

- ▶ Linear modes: 5 (massive spin-2)
- ▶ Non-linear modes: 5 + 1 (**ghost**)

Generic bimetric theory:

- ▶ Linear modes: 5 (massive, $\delta g - \delta f$)
+ 2 (massless, $\delta g + \delta f$)
- ▶ Non-linear modes: 7 + 1 (**ghost**)

Complication: Since the ghost shows up nonlinearly, its absence needs to be established nonlinearly

Construction of ghost-free nonlinear theories

Based on “Decoupling limit” analysis:

A specific $V_{dRGT}(\sqrt{g^{-1}}\eta)$ was obtained and shown to be ghost-free in a “decoupling limit”, also perturbatively in $h = g - \eta$

[de Rham, Gabadadze, 2010; de Rham, Gabadadze, Tolley, 2010]

Non-linear Hamiltonian methods (non-perturbative):

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Non-linear Hamiltonian methods (non-perturbative):

Questions not answerable by “decoupling limit”:

- ▶ Is massive gravity with $V(\sqrt{g^{-1}\eta})$ ghost-free nonlinearly?

[SFH, Rosen (1106.3344, 1111.2070)]

- ▶ Is it ghost-free for generic fixed $f_{\mu\nu}$?

[SFH, Rosen, Schmidt-May (1109.3230)]

- ▶ Can $f_{\mu\nu}$ be given ghost-free dynamics?

[SFH, Rosen (1109.3515)]

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Ghost-free bimetric theory

Digression: Elementary symmetric polynomials of \mathbb{X} with eigenvalues $\lambda_1, \dots, \lambda_4$:

$$e_0(\mathbb{X}) = 1, \quad e_1(\mathbb{X}) = \lambda_1 + \lambda_2 + \lambda_3 + \lambda_4,$$

$$e_2(\mathbb{X}) = \lambda_1\lambda_2 + \lambda_1\lambda_3 + \lambda_1\lambda_4 + \lambda_2\lambda_3 + \lambda_2\lambda_4 + \lambda_3\lambda_4,$$

$$e_3(\mathbb{X}) = \lambda_1\lambda_2\lambda_3 + \lambda_1\lambda_2\lambda_4 + \lambda_1\lambda_3\lambda_4 + \lambda_2\lambda_3\lambda_4,$$

$$e_4(\mathbb{X}) = \lambda_1\lambda_2\lambda_3\lambda_4 = \det \mathbb{X}.$$

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$$\begin{aligned}e_0(\mathbb{X}) &= 1, & e_1(\mathbb{X}) &= \lambda_1 + \lambda_2 + \lambda_3 + \lambda_4, \\e_2(\mathbb{X}) &= \lambda_1\lambda_2 + \lambda_1\lambda_3 + \lambda_1\lambda_4 + \lambda_2\lambda_3 + \lambda_2\lambda_4 + \lambda_3\lambda_4, \\e_3(\mathbb{X}) &= \lambda_1\lambda_2\lambda_3 + \lambda_1\lambda_2\lambda_4 + \lambda_1\lambda_3\lambda_4 + \lambda_2\lambda_3\lambda_4, \\e_4(\mathbb{X}) &= \lambda_1\lambda_2\lambda_3\lambda_4 = \det \mathbb{X}.\end{aligned}$$

$$\begin{aligned}e_0(\mathbb{X}) &= 1, & e_1(\mathbb{X}) &= [\mathbb{X}], \\e_2(\mathbb{X}) &= \frac{1}{2}([\mathbb{X}]^2 - [\mathbb{X}^2]), \\e_3(\mathbb{X}) &= \frac{1}{6}([\mathbb{X}]^3 - 3[\mathbb{X}][\mathbb{X}^2] + 2[\mathbb{X}^3]), \\e_4(\mathbb{X}) &= \frac{1}{24}([\mathbb{X}]^4 - 6[\mathbb{X}]^2[\mathbb{X}^2] + 3[\mathbb{X}^2]^2 + 8[\mathbb{X}][\mathbb{X}^3] - 6[\mathbb{X}^4]), \\e_k(\mathbb{X}) &= 0 \quad \text{for } k > 4,\end{aligned}$$

$$[\mathbb{X}] = \text{Tr}(\mathbb{X}), \quad e_n(\mathbb{X}) \sim (\mathbb{X})^n$$

- ▶ The $e_n(\mathbb{X})$'s and $\det(\mathbb{1} + \mathbb{X})$:

$$\det(\mathbb{1} + \mathbb{X}) = \sum_{n=0}^4 e_n(\mathbb{X})$$

- ▶ Introduce “deformed determinant” :

$$\widehat{\det}(\mathbb{1} + \mathbb{X}) = \sum_{n=0}^4 \beta_n e_n(\mathbb{X})$$

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- ▶ Observation:

$$V(\sqrt{g^{-1}f}) = \sum_{n=0}^4 \beta_n e_n(\sqrt{g^{-1}f})$$

[SFH & R. A. Rosen (1103.6055)]

Ghost-free bi-metric theory

Ghost-free combination of kinetic and potential terms for g & f :

$$\mathcal{L} = m_g^2 \sqrt{-g} R_g - 2m^4 \sqrt{-g} \sum_{n=0}^4 \beta_n e_n(\sqrt{g^{-1}f}) + m_f^2 \sqrt{-f} R_f$$

[SFH, Rosen (1109.3515, 1111.2070)]

Note,

$$\sqrt{-g} \sum_{n=0}^4 \beta_n e_n(\sqrt{g^{-1}f}) = \sqrt{-f} \sum_{n=0}^4 \beta_{4-n} e_n(\sqrt{f^{-1}g})$$

Hamiltonian analysis: 7 nolinear propagating modes, **no ghost!**

$$C(\gamma, \pi) = 0, \quad C_2(\gamma, \pi) = \frac{d}{dt} C(x) = \{H, C\} = 0$$

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Mass spectrum of bimetric theory

[SFH, A. Schmidt-May, M. von Strauss 1208:1515, 1212:4525]

$$S_{gf} = - \int d^d x \left[m_g^{d-2} \sqrt{g} R_g - 2m^d \sqrt{g} \sum_{n=0}^d \beta_n e_n(\mathcal{S}) + m_f^{d-2} \sqrt{f} R_f \right]$$

Three Questions:

- ▶ **Q1:** When are the 7 fluctuations in $\delta g_{\mu\nu}$, $\delta f_{\mu\nu}$ good mass eigenstates? (FP mass)
- ▶ **Q2:** In what sense is this Massive spin-2 field + gravity ?
- ▶ **Q3:** How to characterize deviations from General Relativity?

$$R_{\mu\nu}(g) - \frac{1}{2} g_{\mu\nu} R(g) + V_{\mu\nu}^g = T_{\mu\nu}^g$$

$$R_{\mu\nu}(f) - \frac{1}{2} f_{\mu\nu} R(f) + V_{\mu\nu}^f = T_{\mu\nu}^f$$

Proportional backgrounds

A1: FP masses exist only around,

$$\bar{f}_{\mu\nu} = c^2 \bar{g}_{\mu\nu}$$

g and f equations:

$$R_{\mu\nu}(\bar{g}) - \frac{1}{2} \bar{g}_{\mu\nu} R(\bar{g}) + \begin{pmatrix} \Lambda_g \\ \Lambda_f \end{pmatrix} \bar{g}_{\mu\nu} = 0 \text{ or } \begin{pmatrix} T_{\mu\nu}^g \\ T_{\mu\nu}^f \end{pmatrix}$$

$$\Lambda_g = \frac{m^d}{m_g^{d-2}} \sum_{k=0}^{d-1} \binom{d-1}{k} c^k \beta_k, \quad \Lambda_f = \frac{m^d}{m_f^{d-2}} \sum_{k=1}^d \binom{d-1}{k-1} c^{k+2-d} \beta_k$$

Implication:

$$\Lambda_g = \Lambda_f \quad \Rightarrow \quad c = c(\beta_n, \alpha \equiv m_f/m_g)$$

(Exception: Partially massless (PM) theory)

Mass spectrum around proportional backgrounds

Linear modes:

$$\delta M_{\mu\nu} = \frac{1}{2c} \left(\delta f_{\mu\nu} - c^2 \delta g_{\mu\nu} \right), \quad \delta G_{\mu\nu} = \left(\delta g_{\mu\nu} + \alpha^{d-2} c^{d-4} \delta f_{\mu\nu} \right)$$

$$\bar{\mathcal{E}}_{\mu\nu}^{\rho\sigma} \delta G_{\rho\sigma} - \Lambda_g \left(\delta G_{\mu\nu} - \frac{1}{2} \bar{g}_{\mu\nu} \bar{g}^{\rho\sigma} \delta G_{\rho\sigma} \right) = 0,$$

$$\begin{aligned} \bar{\mathcal{E}}_{\mu\nu}^{\rho\sigma} \delta M_{\rho\sigma} - \Lambda_g \left(\delta M_{\mu\nu} - \frac{1}{2} \bar{g}_{\mu\nu} \bar{g}^{\rho\sigma} \delta M_{\rho\sigma} \right) \\ + \frac{1}{2} m_{\text{FP}}^2 \left(\delta M_{\mu\nu} - \bar{g}_{\mu\nu} \bar{g}^{\rho\sigma} \delta M_{\rho\sigma} \right) = 0 \end{aligned}$$

The FP mass of δM :

$$m_{\text{FP}}^2 = \frac{m^d}{m_g^{d-2}} \left(1 + (\alpha c)^{2-d} \right) \sum_{k=1}^{d-1} \binom{d-2}{k-1} c^k \beta_k$$

Bimetric as massive spin-2 field + gravity

A2: The massless mode is not gravity!

$$G_{\mu\nu} = g_{\mu\nu} + c^{d-4} \alpha^{d-2} f_{\mu\nu}, \quad M_{\mu\nu}^G = G_{\mu\rho} (\sqrt{g^{-1}f})^\rho{}_\nu - c G_{\mu\nu}$$

$G_{\mu\nu}$ has no ghost-free matter coupling!

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Hence:

- ▶ Gravity: $g_{\mu\nu}$
- ▶ Massive spin-2 field: $M_{\mu\nu} = g_{\mu\rho} (\sqrt{g^{-1}f})^\rho{}_\nu - c g_{\mu\nu}$
- ▶ $m_g \gg m_f$: $g_{\mu\nu}$ mostly massless (opposite to massive gr.)

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Hence:

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A3: $M_{\mu\nu} = 0 \Rightarrow$ GR.

$M_{\mu\nu} \neq 0 \Rightarrow$ deviations from GR, driven by matter couplings

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Partial masslessness in FP theory

$$\bar{\mathcal{E}}^{\rho\sigma}_{\mu\nu} h_{\rho\sigma} - \Lambda(h_{\mu\nu} - \frac{1}{2}\bar{g}_{\mu\nu}h^\rho{}_\rho) + \frac{m_{FP}^2}{2}(h_{\mu\nu} - \bar{g}_{\mu\nu}h^\rho{}_\rho) = 0$$

dS/Einstein backgrounds:

$$\bar{g}_{\mu\nu} : \quad R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R + \Lambda g_{\mu\nu} = 0$$

Higuchi Bound:

$$m_{FP}^2 = \frac{2}{3}\Lambda$$

New gauge symmetry:

$$\Delta h_{\mu\nu} = (\nabla_\mu \nabla_\nu + \frac{\Lambda}{3})\xi(x)$$

Gives $5-1=4$ propagating modes

[Deser, Waldron, ... (1983-2012)]

Can a nonlinear extension of PM theory exist?

Partial masslessness beyond FP theory

Non-linear PM theory = Nonlinear spin-2 fields with a gauge invariance!

Does it exist? Independent of dS/Einstein backgrounds?

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Known perturbative results around dS:

- ▶ Cubic PM vertices ($\sim h^3$) in $d = 4$ *[Zinoviev (2006)]*
- ▶ Cubic PM vertices exist only in $d = 3, 4$ with 2 derivatives
For $d > 4$, higher derivative terms needed.

[Joung, Lopez, Taronna (2012)]

We will identify a specific bimetric theory as the candidate nonlinear PM theory

Partial masslessness in Bimetric theory

[SFH, Schmidt-May, von Strauss, 1208:1797, 1212:4525]

- 1) Assume a nonlinear bimetric theory with PM symmetry exists
- 2) Around $\bar{f} = c^2 \bar{g}$, $\delta M_{\mu\nu}$ satisfies the FP equation. Then the action of PM symmetry must be:

$$\delta M_{\mu\nu} \rightarrow \delta M_{\mu\nu} + \left(\nabla_\mu \nabla_\nu + \frac{\Lambda}{3} \bar{g}_{\mu\nu} \right) \xi(x), \quad \delta G_{\mu\nu} \rightarrow \delta G_{\mu\nu}$$

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- ▶ Find the transformation of $\delta g_{\mu\nu}$ & $\delta f_{\mu\nu}$.
- ▶ Shift the transf. to dynamical backgrounds $\bar{g}_{\mu\nu}$ & $\bar{f}_{\mu\nu}$
- ▶ For the dS-preserving subset $\xi = \xi_0$ (const), this gives,

$$\bar{g}'_{\mu\nu} = (1 + a\xi_0) \bar{g}_{\mu\nu}, \quad \bar{f}'_{\mu\nu} = (1 + b\xi_0) \bar{f}_{\mu\nu}$$

$$\bar{f}' = c'^2(\xi_0) \bar{g}' \quad c' \neq c$$

A symmetry can exist only if $\Lambda_g = \Lambda_f$ does not determine c

Candidate PM bimetric theory in d=4

The necessary condition for the existence of PM symmetry is that c is **not** determined by $\Lambda_g = \Lambda_f$, or

$$\beta_1 + \left(3\beta_2 - \alpha^2\beta_0\right) c + \left(3\beta_3 - 3\alpha^2\beta_1\right) c^2 + \left(\beta_4 - 3\alpha^2\beta_2\right) c^3 + \alpha^2\beta_3c^4 = 0$$

This gives the candidate nonlinear PM theory (d=4)

$$\alpha^2\beta_0 = 3\beta_2, \quad 3\alpha^2\beta_2 = \beta_4, \quad \beta_1 = \beta_3 = 0$$

Nonlinear PM bimetric theory

Checks:

- ▶ $m_{\text{FP}}^2 = 2 \frac{m^4}{m_g^2} (\alpha^{-2} + c^2) \beta_2 = \frac{2}{3} \Lambda_g$
- ▶ For $d > 4$, all $\beta_n = 0$. Nonlinear PM bimetric exists only for $d = 3, 4$.
- ▶ In $d > 4$ PM is restored by Lanczos-Lovelock terms
- ▶ Realization of the ξ_0 gauge transformation in the nonlinear theory on dS.

Full Gauge symmetry of the nonlinear theory not yet known, but expect $6=7-1$ propagating modes

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Higher derivative gravity and Conformal gravity

HD gravity:

$$S_{(2)}^{\text{HD}}[g] = m_g^2 \int d^4x \sqrt{g} \left[\Lambda + c_R R(g) - \frac{c_{RR}}{m^2} \left(R^{\mu\nu} R_{\mu\nu} - \frac{1}{3} R^2 \right) \right]$$

7 modes: massless spin-2 + massive spin-2 (ghost)

[Stelle (1977)]

d=3: New Massive Gravity (NMG)

[Bergshoeff, Holm, Townsend (2009)]

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Conformal Gravity:

$$S^{\text{CG}}[g] = -c \int d^4x \sqrt{g} \left[R^{\mu\nu} R_{\mu\nu} - \frac{1}{3} R^2 \right],$$

Weyl Invariance \Rightarrow 6 modes: 2 (massless spin-2) + 4 **ghosts**

[Riegert (1984), Maldacena (2011)]

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Curvature expansion of bimetric equations

Define

$$S = \sqrt{g^{-1}f}, \quad P_{\mu\nu} = R_{\mu\nu} - \frac{1}{2(d-1)}g_{\mu\nu}R$$

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$$S = \sqrt{g^{-1}f}, \quad P_{\mu\nu} = R_{\mu\nu} - \frac{1}{2(d-1)}g_{\mu\nu}R$$

Solve the bimetric $g_{\mu\nu}$ equation algebraically for $f_{\mu\nu}$, as an expansion in $R_{\mu\nu}(g)/m^2$,

$$S^\mu_\nu = a\delta^\mu_\nu + \frac{a_1}{m^2}P^\mu_\nu + \frac{a_2}{m^4} \left[\left(P^\mu_\nu{}^2 - PP^\mu_\nu \right) + \frac{1}{d-1}e_2(P)\delta^\mu_\nu \right] + \mathcal{O}(m^{-6})$$

Compute,

$$f_{\mu\nu} = a^2 g_{\mu\nu} + \frac{2aa_1}{m^2}P^\mu_\nu + \mathcal{O}(m^{-4})$$

Equivalence between CG and PM bimetric theory

[SFH, Schmidt-May, von Strauss, 1303:6940]

- ▶ CG equation of motion: The Bach equation (4-derivative),

$$B_{\mu\nu} = 0$$

Propagates 6 modes due to Weyl invariance.

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- ▶ In PM bimetric theory, solve the g -equation for $f_{\mu\nu}$.
Substitute back in f -equation to get,

$$B_{\mu\nu} + \mathcal{O}(R^3/m^2) = 0$$

In the low curvature limit, PM bimetric theory has a gauge symmetry even away from dS and definitely propagates $7 - 1 = 6$ modes! None is a ghost

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- ▶ CG eom is the low curvature limit of PM bimetric eom.
Conversely, PM bimetric is a ghost-free completion of CG

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Higher derivative gravity from Bimetric theory

- ▶ Solve the g -equation for $f = f(g)$. Then,

$$S^{\text{BM}}[g, f(g)] = S^{\text{HD}}[g]$$

- ▶ 4-derivative ($\sim R^2$) truncation:

$$S_{(2)}^{\text{BM}}[g, f(g)] = S_{(2)}^{\text{HD}}[g]$$

The spin-2 ghost in 4-derivative HD gravity is an artifact of this truncation (can be illustrated in a linear theory).

- ▶ The correspondence is not an equivalence of the truncated theories (in general). Different truncated EoM's.
- ▶ PM bimetric theory again leads to conformal gravity.
- ▶ $d=3$ reproduces NMG.

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- ▶ Recurring doubts about $7=5+2$ modes in bimetric theory. But in all cases finally confirmed.
- ▶ Superluminality (at least two light cones), but no superluminality in the matter sector.
- ▶ Accausality?: Is the Cauchy problem well posed? Yes. No generic shock waves.

[Izumi, Ong (1304.0211)]

- ▶ Stability of classical solutions: Schwarzschild with $f = c^2 g$ has a Gregory-Laflam type instability which goes away in the PM case. Not relevant for astrophysical blackholes

[Babichev, Fabbri (1304.5992), Brito, Cardoso, Pani (1304.6725)]

- ▶ Proof of PM gauge symmetry/6 modes in the candidate PM bimetric theory ?