



Precision Measurement of Gravity at Micrometer Scale using Ultracold Atoms

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Outline

- *Motivation*
- *Previous experiments*
- *Measuring gravity with atoms*
- *Precision gravity measurement at μm scale with laser-cooled Sr atoms in an optical lattice*
- *Prospects*

Motivation

- **Physics beyond the standard model**

Extra space-time dimensions

Deviations from $1/r^2$ law

Hierarchy problem: why is gravity so weak?

New boson-exchange forces

Radion – low-mass spin-0 fields with gravitational-strength couplings

Moduli – massive scalar particles producing gravitylike forces

Dilaton – Light scalar in string theory, coupling to nucleons

Axion – pseudoscalar particles explaining smallness of CP violation in QCD for strong nuclear force

Multi-particle exchange forces

- **Small observed size of Einstein cosmological constant**

- **Experimental challenge**

N. Arkani-Hamed, S. Dimopoulos, G. Dvali, Phys. Lett. B 429, 263 (1998)
N. Arkani-Hamed, S. Dimopoulos, G. Dvali, Phys. Rev. D 59, 086004 (1999)

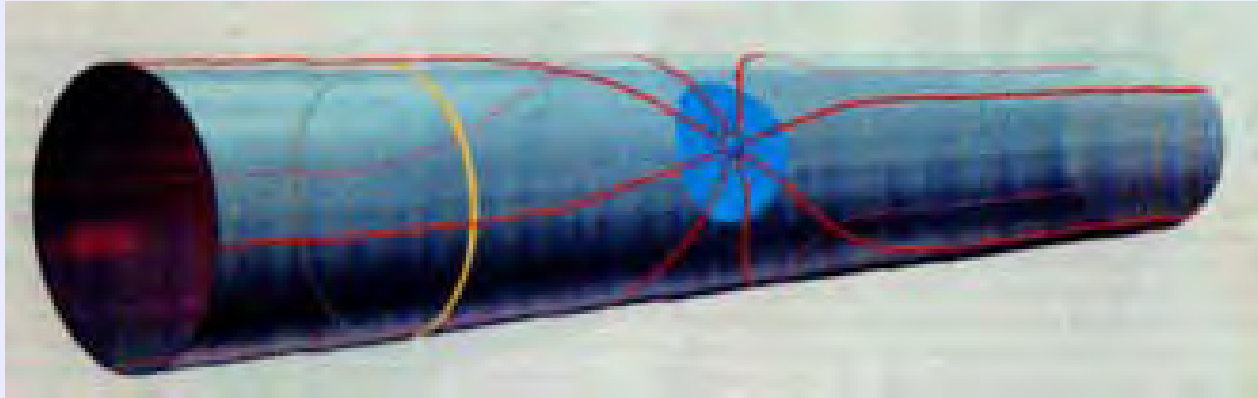
S. Dimopoulos and G. F. Giudice, Phys. Lett. B 379, 105 (1996)
I. Antoniadis, S. Dimopoulos, and G. Dvali, Nuc. Phys. B 516,70 (1998)

T.R. Taylor, G. Veneziano, Phys. Lett. B 213, 450 (1988)
D. B. Kaplan, M. B. Wise, J. High Energy Phys. 8, 37 (2000)

Moody and Wilczek, Phys Rev. D 30, 130 (1984)
R. Barbieri, A. Romanino, A. Strumia, Phys. Lett. B 387, 310 (1996)
L.J. Rosenberg, K.A. van Bibber, Phys. Rep. 325, 1 (2000))

S.R. Beane, Gen. Rel. Grav. 29, 945 (1997)
R. Sundrum, Phys. Rev. D 69, 044014 (2004)

Extra Dimensions



| | | | |
|--------------------|---------------|-------------------|--------------------------|
| 1 extra dimension | \Rightarrow | $F \propto 1/r^3$ | size $\approx 10^{11}$ m |
| 2 extra dimensions | \Rightarrow | $F \propto 1/r^4$ | size $\approx 10^{-4}$ m |
| 3 extra dimensions | \Rightarrow | $F \propto 1/r^5$ | size $\approx 10^{-9}$ m |
| ... | | | |

N. Arkani-Hamed, S. Dimopoulos, G. Dvali, *The hierarchy problem and new dimensions at a millimeter*, Phys. Lett. B 429, 263 (1998)

N. Arkani-Hamed, S. Dimopoulos, G. Dvali, *Phenomenology, astrophysics, and cosmology of theories with submillimeter dimensions and TeV scale quantum gravity*, Phys. Rev. D 59, 086004 (1999)

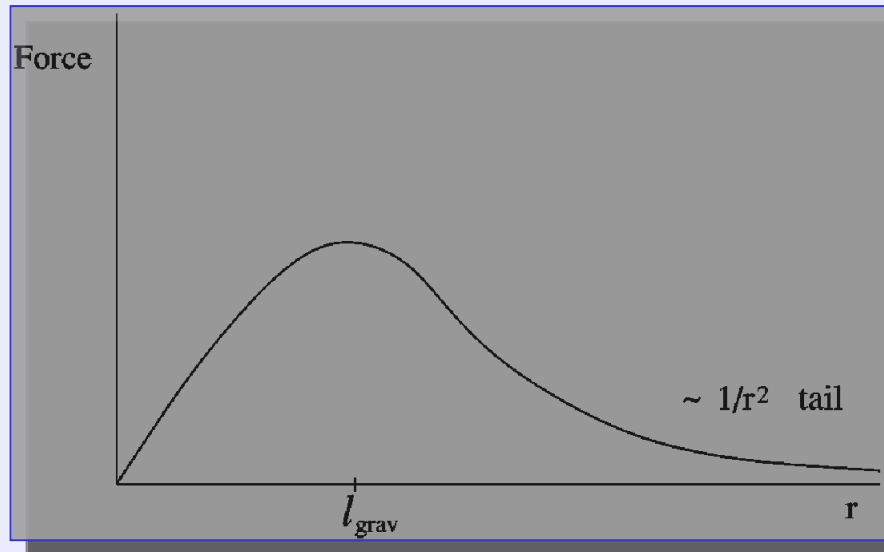
Cosmological constant problem and “fat” gravitons

Cosmological data \Rightarrow vacuum-energy density $\rho_{\text{vac}} \sim 3 \cdot 10^3 \text{ eV/cm}^3$



$$l \sim (\hbar c / \rho_{\text{vac}})^{1/4} \sim 0.1 \text{ mm}$$

The graviton is “fat” at distances $\leq 1 \text{ mm}$?
Solution to the CC Problem?



R. Sundrum, *Fat gravitons, the cosmological constant and submillimeter tests*, Phys. Rev. D 69, 044014 (2004)

$$20 \mu\text{m} \leq l_{\text{grav}} \leq 0.2 \text{ mm}$$

Parametrizations for deviations from Newtonian gravity

- **Modification of power law in Newton-type force**

$$F(r) = G \frac{M_1 M_2}{r^{2+\delta}}$$

- **Newton+Yukawa potential**

$$V(r) = -G \frac{M_1 M_2}{r} \left[1 + \alpha e^{-\frac{r}{\lambda}} \right]$$

- Exchange of a boson with $m = \hbar/\lambda c$
- Extra dimensions

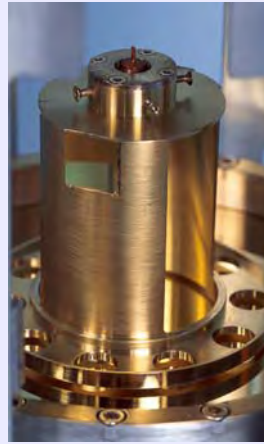
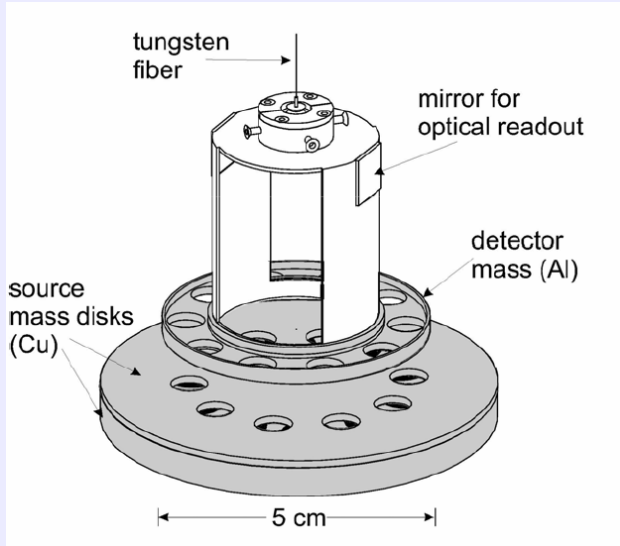
- **Modified power-law potential**

$$V(r) = -G \frac{M_1 M_2}{r} \left[1 + \alpha_N \left(\frac{r_0}{r} \right)^{N-1} \right]$$

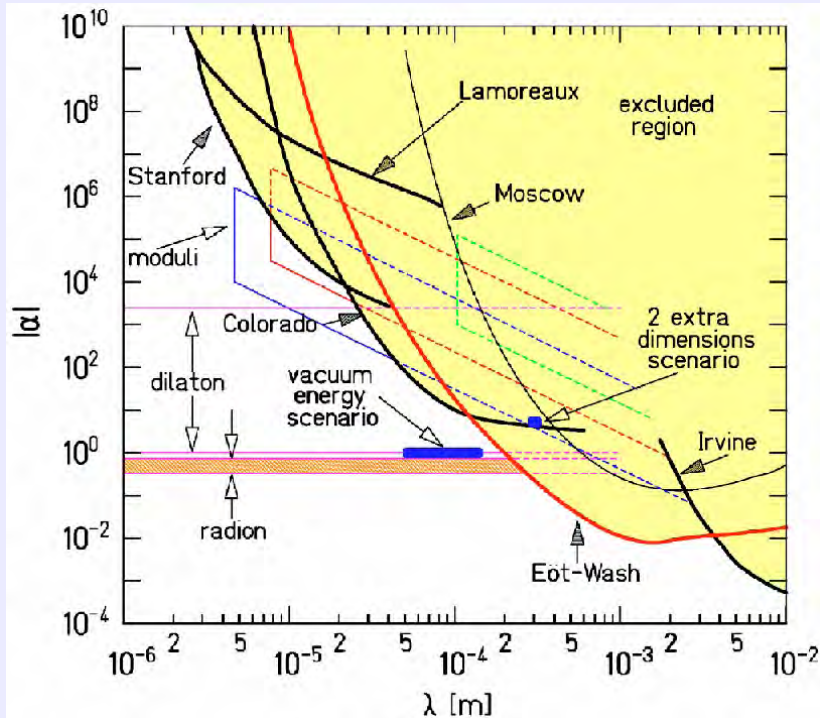
Exchange of 2 massless particles

Experiments on gravity at small spatial scale

Torsion balance - Washington experiment



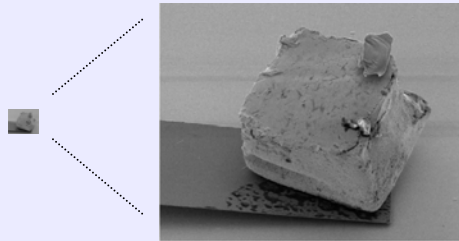
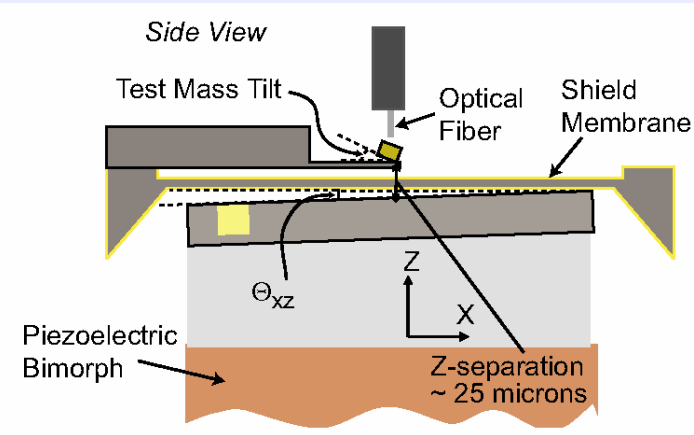
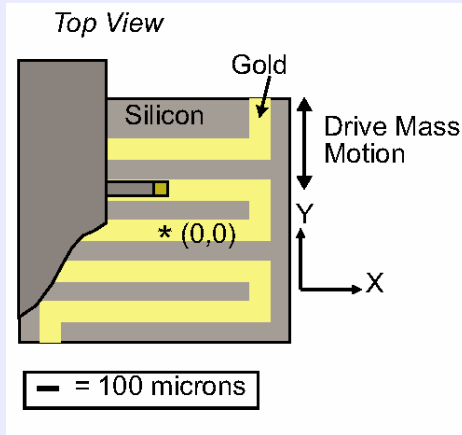
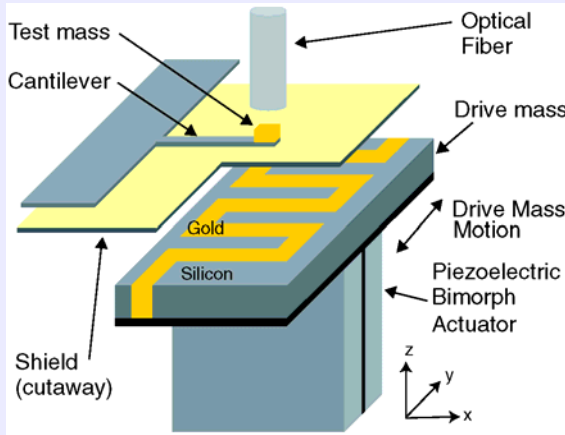
- **Test bodies: “missing masses” of holes bored into plates**
- **Torsion pendulum**
7075 aluminum, gold coated
disk height = 2 mm
10 cylindrical holes evenly spaced about the azimuth
- **Attractor**
high-purity copper disk
top surface coated with gold
10 cylindrical holes evenly spaced about the azimuth
uniformly rotating
- **Electrostatic shield**
tightly stretched 20- μ m-thick BeCu foil



- **Distance from top of attractor to bottom of pendulum**
from 10.77 mm to 137 μ m

C. D. Hoyle, D. J. Kapner, B. R. Heckel, E. G. Adelberger, J. H. Gundlach, U. Schmidt, H. E. Swanson, *Submillimeter tests of the gravitational inverse-square law*, PRD 70, 042004 (2004)

Microcantilever - Stanford experiment

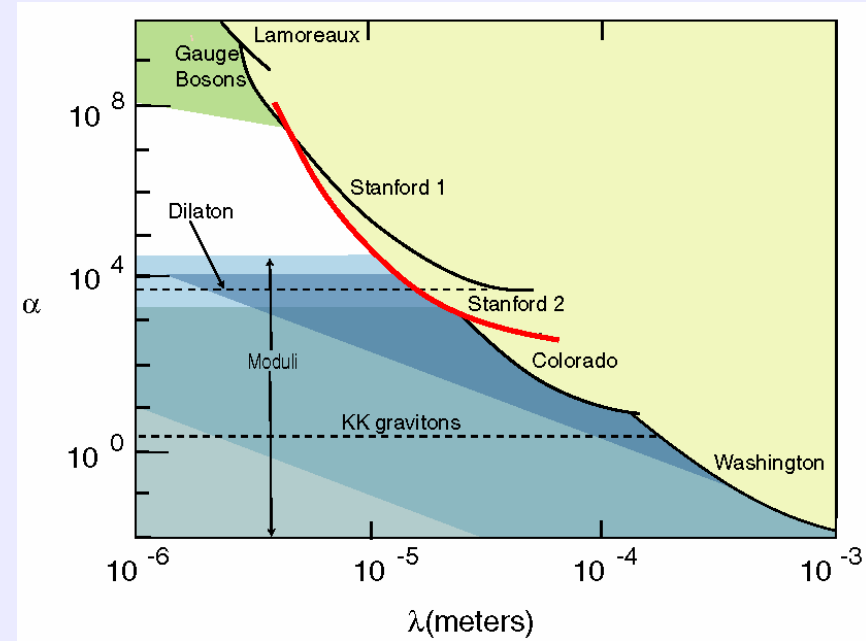
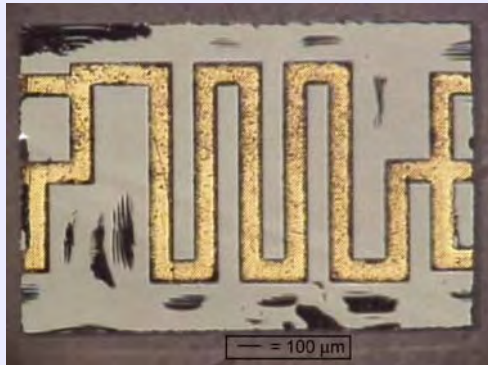


Probe mass (gold)
 $50 \mu\text{m} \times 50 \mu\text{m} \times 30 \mu\text{m}$
 $m_t \sim 1.6 \mu\text{g}$

Cantilever (<100> Si)
 $50 \mu\text{m} \times 250 \mu\text{m} \times 0.33 \mu\text{m}$
 $Q \sim 80\,000$
 $\omega_0 \sim (k/m_t)^{1/2} \sim 300 \text{ Hz}$

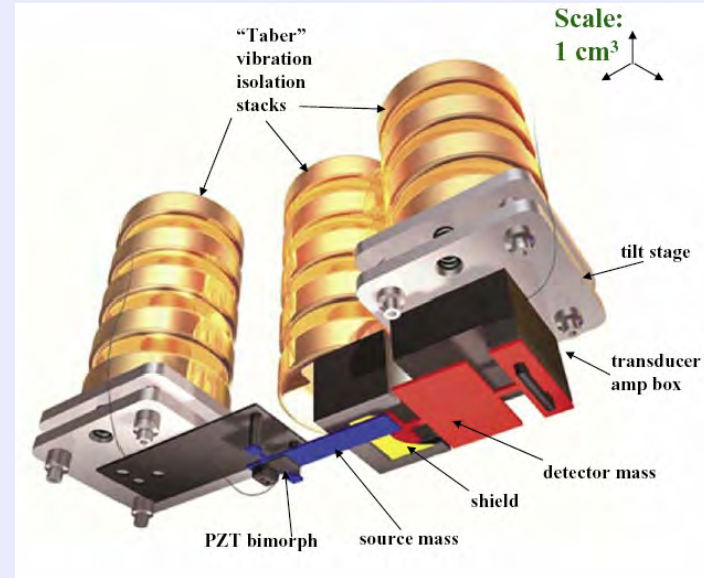
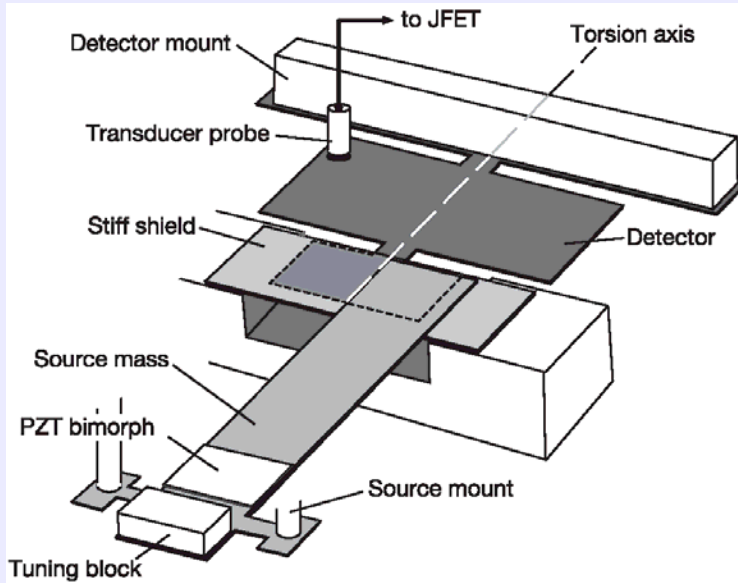
Source mass
 5 sets of gold and silicon bars
 $100 \mu\text{m} \times 1 \text{ mm} \times 100 \mu\text{m}$

Separation 25 μm



S. J. Smullin, A. A. Geraci, D. M. Weld, J. Chiaverini, S. Holmes, A. Kapitulnik, *Constraints on Yukawa-type deviations from Newtonian gravity at 20 microns*, Phys. Rev. D 72, 122001 (2005)

Microcantilever - Colorado experiment

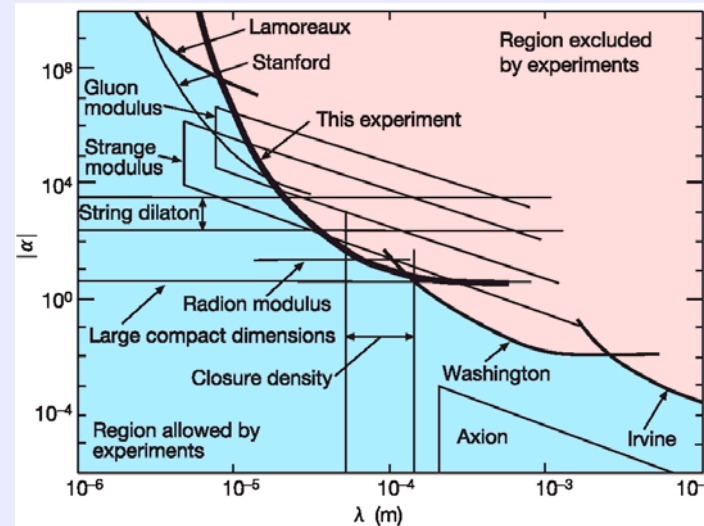


Detector (tungsten)
11.455 mm x 5.080 mm x 195 μm

$Q \sim 25\,000$
 $\omega_0 \sim 1173\text{ Hz}$

Source mass (tungsten)
35 mm x 7 mm x 305 μm

Separation 108 μm



J.C. Long, H.W. Chan, A.B. Churnside, E.A. Gulbis, M.C.M. Varney, J.C. Price, *Upper limits to submillimetre-range forces from extra space-time dimensions*, Nature 421, 922 (2003)

Experiments on gravity at small spatial scale

Experiments based on torsion balances ($\lambda \leq 1$ mm)

J. Gundlach and E. Adelberger (Washington) – torsion balance

R. Newman and P. Boynton (Irvine, Washington) – cryogenic torsion balance

Experiments based on high-frequency oscillators ($\lambda \leq 0.1$ mm)

J. Long and J. Price group (Colorado) – torsional oscillator

A. Kapitulnik group (Stanford) - microcantilever

R. Decca and E. Fischbach group (Purdue, Indiana) – torsional oscillator

New experiments based on atomic probes ($\lambda \leq 0.01$ mm)

E.A. Cornell group (Colorado) – Oscillations of a Bose-Einstein condensate

G.M. Tino group (Firenze) – Atom interferometry

Also experiments on Casimir effect ($\lambda \leq 0.001$ mm)

Departures observed at $\sim 70 \mu\text{m}$?
(Washington, Stanford)

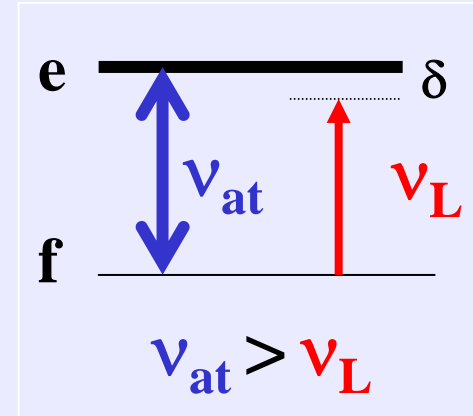
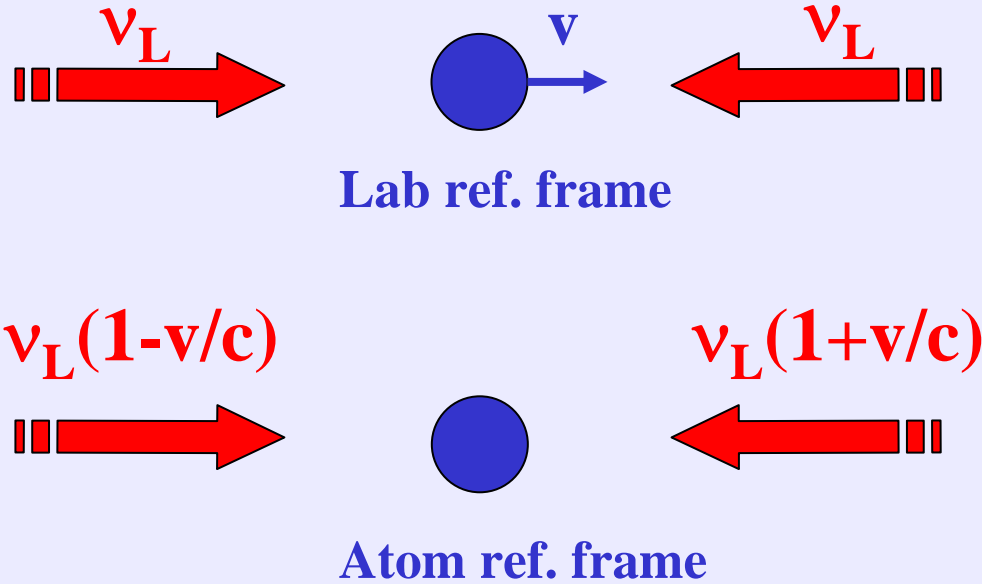
Experimental artifact?

Why atoms?

- Extremely small size
- Well known and reproducible properties
- Quantum systems
- Precision gravity measurement by atom interferometry
- Potential immunity from stray fields effects

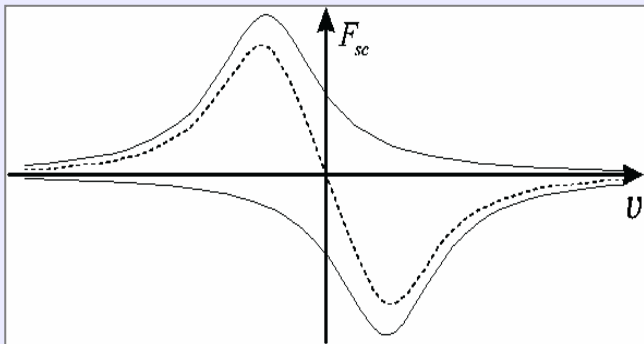
Laser cooling and manipulation of atoms

Optical molasses



$(I/I_0 \ll 1) \rightarrow$

$$F(\nu) \approx \frac{h\nu_L}{c} \cdot \frac{1}{2\tau} \cdot \left[\frac{I/I_0}{1 + I/I_0 + \frac{4}{\Gamma^2} (\delta - \frac{\nu_L}{c} v)^2} - \frac{I/I_0}{1 + I/I_0 + \frac{4}{\Gamma^2} (\delta + \frac{\nu_L}{c} v)^2} \right]$$



$$F(v) \approx \frac{h}{4\pi^2} \frac{\omega_L^2 8\delta}{c^2 \Gamma} \frac{I/I_0}{[1 + (\frac{2\delta}{\Gamma})^2]^2} v = -\alpha v$$

Idea: T.W. Hänsch, A. Schawlow, 1975
Exp. demonstration: S. Chu et al., 1985

Laser cooling: atomic temperatures

Atomic Temperature : $k_B T = M v_{\text{rms}}^2$

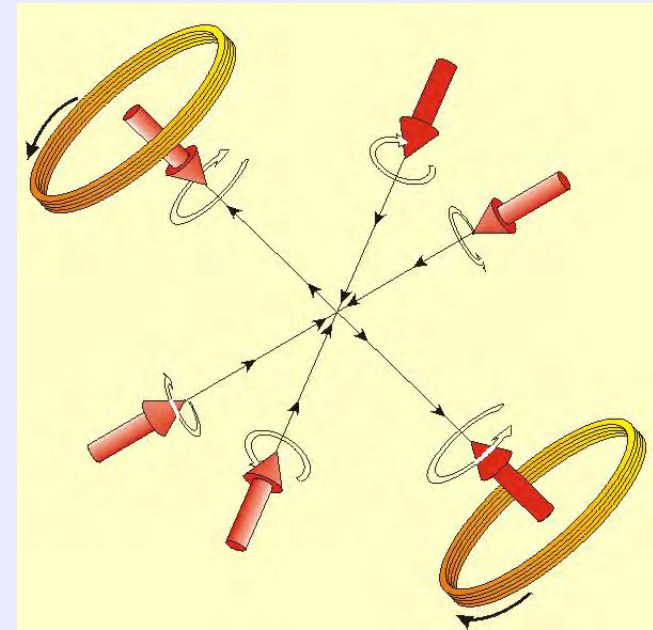
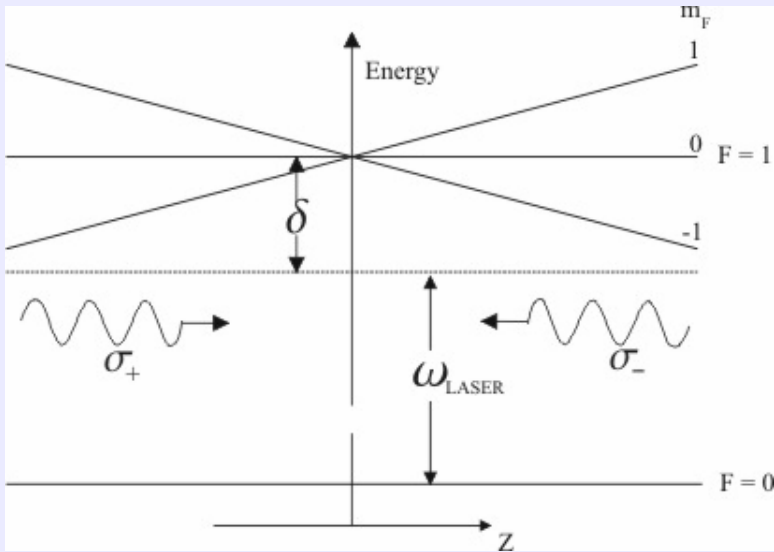
Minimum temperature for Doppler cooling: $k_B T_D = \frac{h\Gamma}{2}$

Single photon recoil temperature: $k_B T_r = \frac{1}{M} \left(\frac{h\nu_L}{c} \right)^2$

Examples:

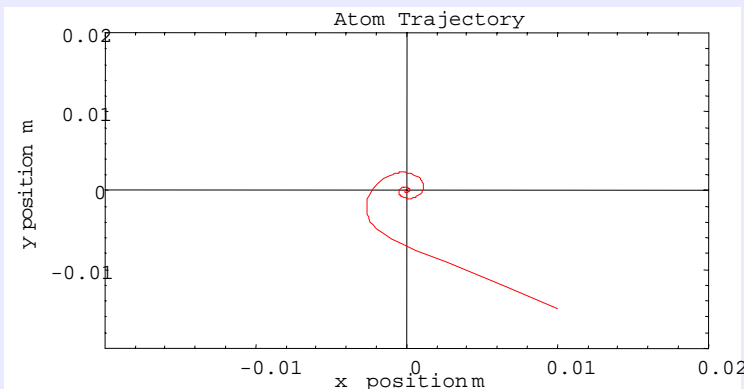
| | T_D | T_r |
|----|-------------------|-------------------|
| Na | 240 μK | 2.4 μK |
| Rb | 120 μK | 360 nK |
| Cs | 120 μK | 200 nK |
| Sr | 180 nK | 460 nK |

Magneto-Optical Trap (MOT)



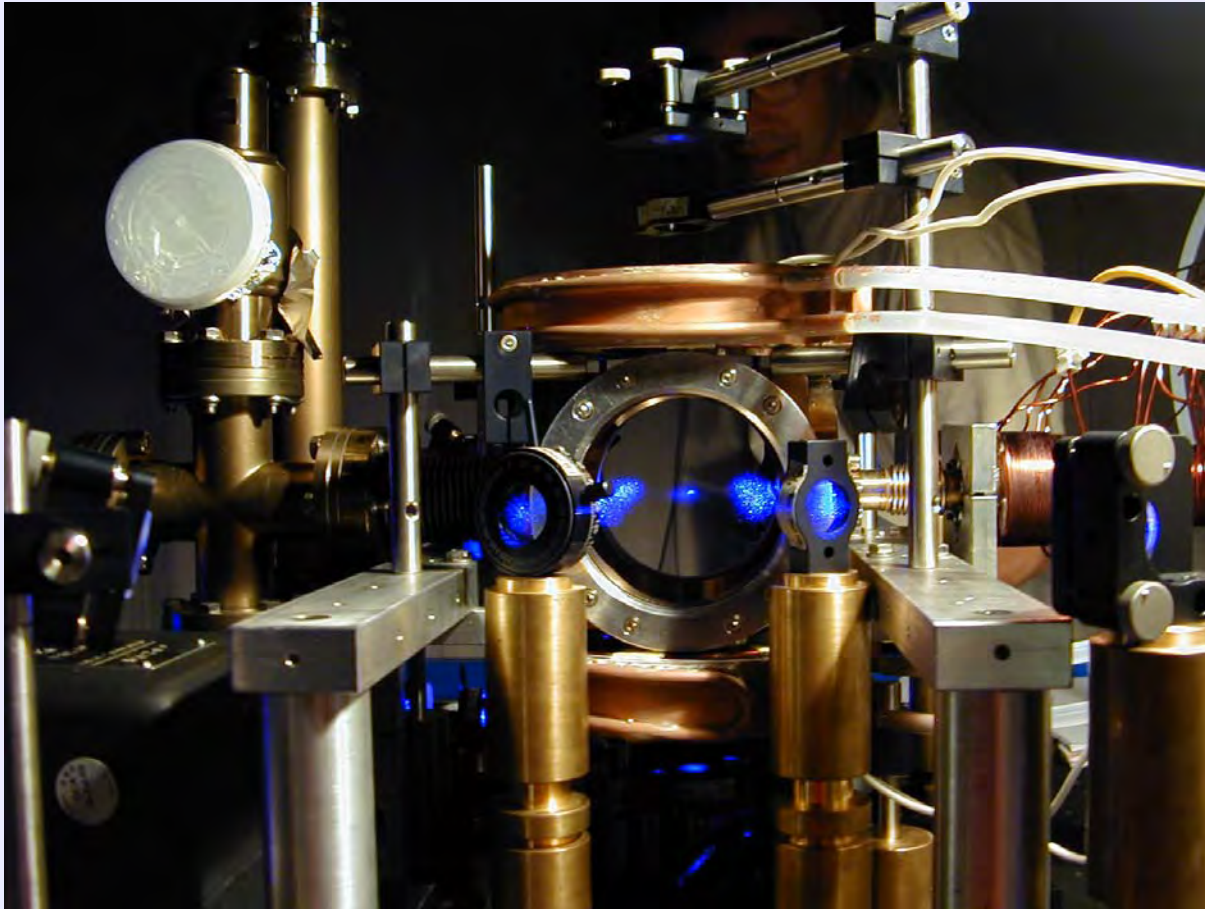
$$F(z, v) \approx \frac{4hk I \delta}{\pi I_0 \Gamma} \frac{kv + \beta z}{[1 + (\frac{2\delta}{\Gamma})^2]^2}$$

density $n \approx 10^{11} \text{ cm}^{-3}$
 temperature $T \approx 100 \text{ } \mu\text{K}$
 size $\Delta x \approx 1 \text{ mm}$



E. Raab *et al.*, Phys. Rev. Lett. **59**, 2631 (1987)

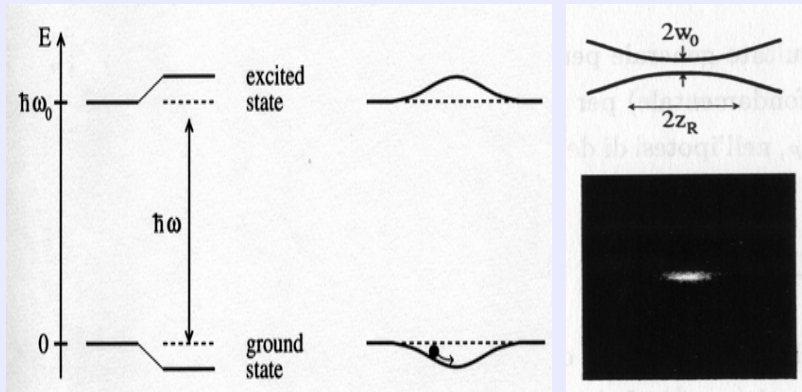
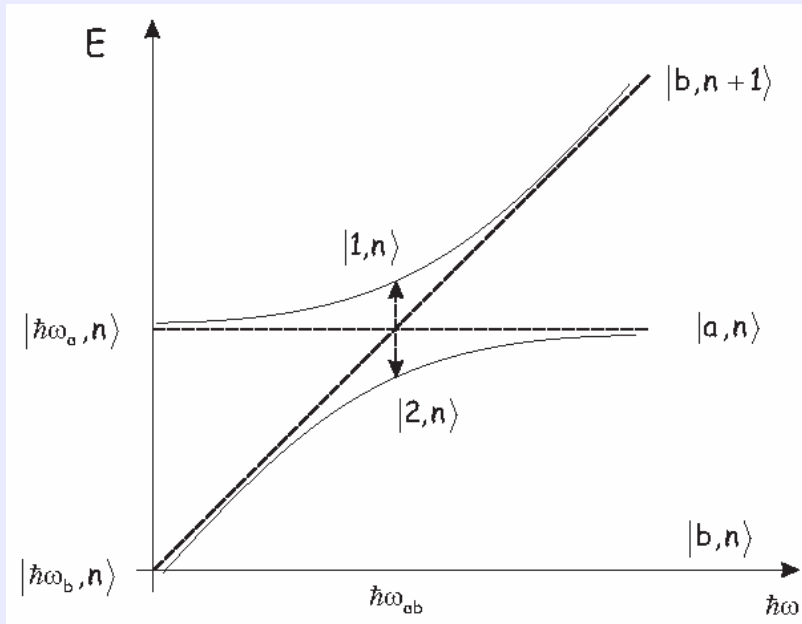
Sr MOT picture



LENS, Firenze

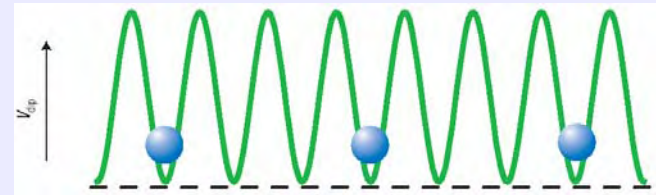


Light shifts and optical traps

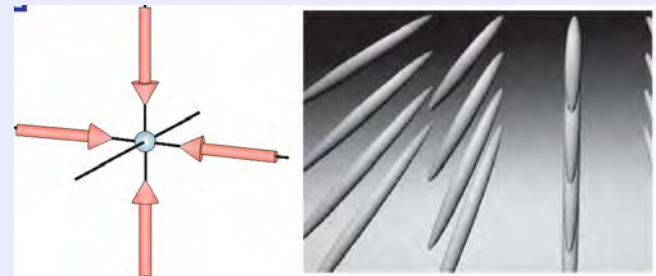


$$V_{\text{dip}}(\mathbf{r}) = -\mathbf{d} \cdot \mathbf{E}(\mathbf{r}) \propto \alpha(\omega_L) |\mathbf{E}(\mathbf{r})|^2$$

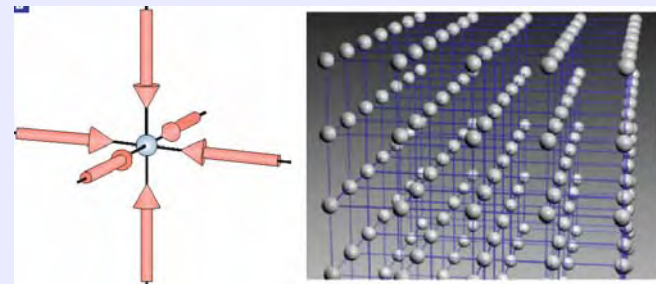
optical lattices



1D optical lattice \Rightarrow array of 2D disk-like trapping potentials



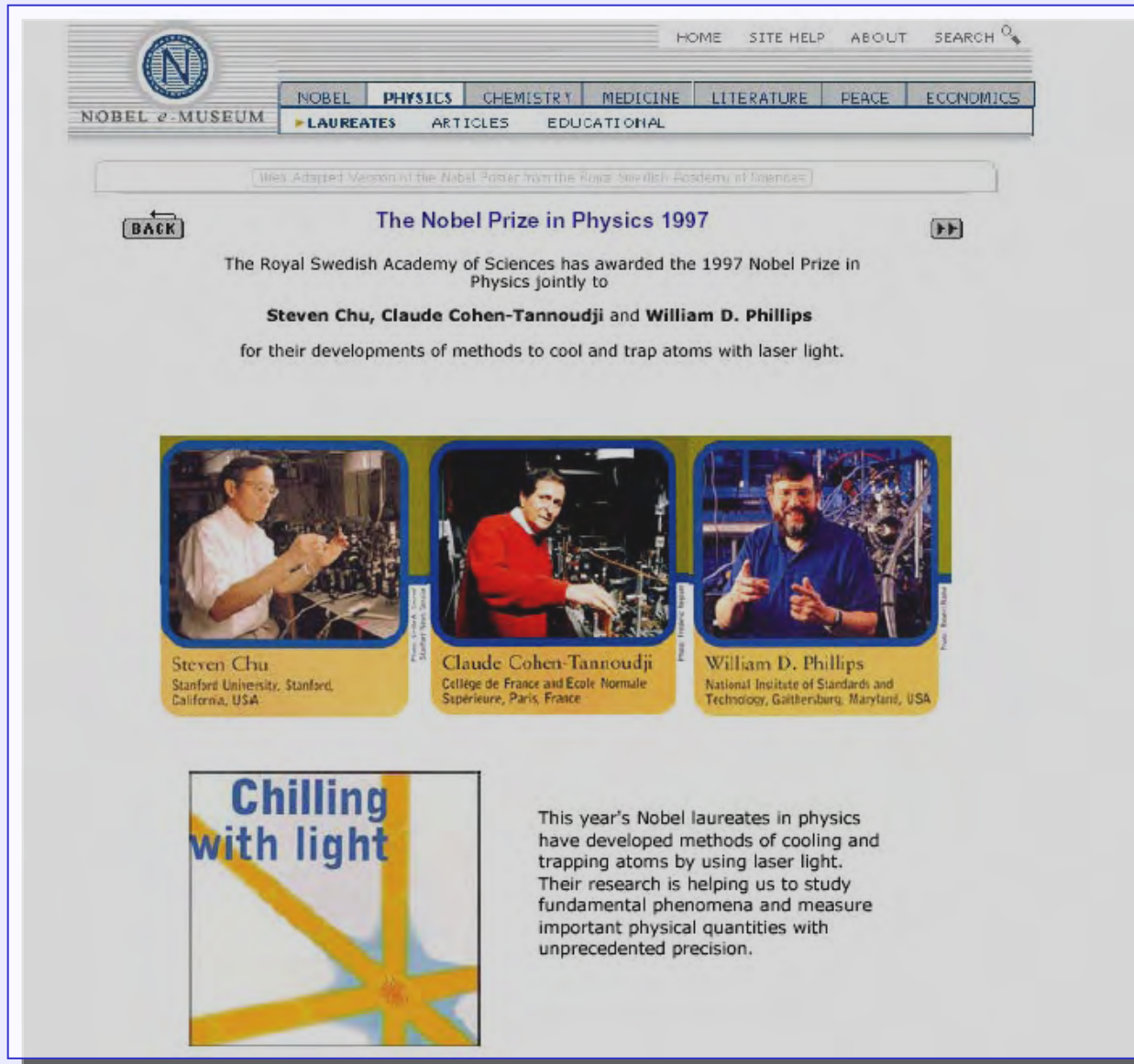
2D optical lattice \Rightarrow array of 1D potential tubes



3D optical lattice \Rightarrow 3D simple cubic array of h.o. potentials

Cooling and trapping atoms with laser light

The Nobel Prize in Physics 1997



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Web Adapted Version of the Nobel Poster from the Royal Swedish Academy of Sciences

[BACK](#) **The Nobel Prize in Physics 1997** [FORWARD](#)


The Royal Swedish Academy of Sciences has awarded the 1997 Nobel Prize in Physics jointly to

Steven Chu, Claude Cohen-Tannoudji and William D. Phillips


for their developments of methods to cool and trap atoms with laser light.



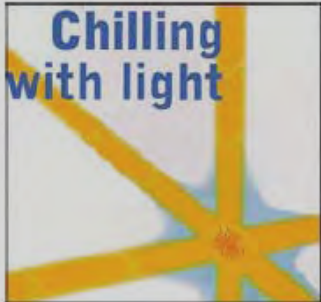
Steven Chu
Stanford University, Stanford, California, USA



Claude Cohen-Tannoudji
Collège de France and École Normale Supérieure, Paris, France



William D. Phillips
National Institute of Standards and Technology, Gaithersburg, Maryland, USA

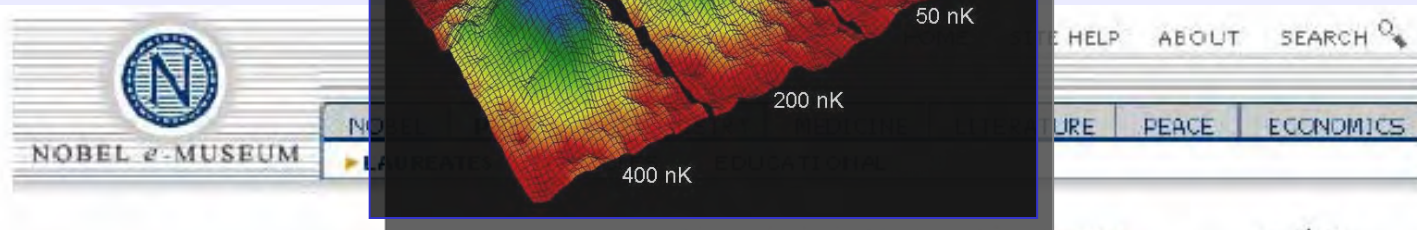


Chilling with light

This year's Nobel laureates in physics have developed methods of cooling and trapping atoms by using laser light. Their research is helping us to study fundamental phenomena and measure important physical quantities with unprecedented precision.

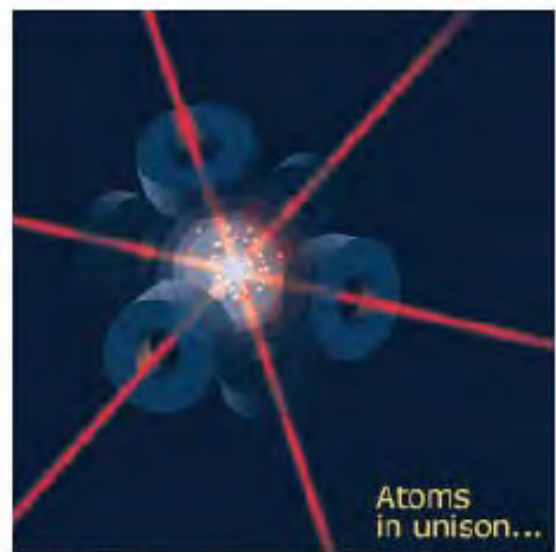
Bose-Einstein condensation in dilute gases of atoms

The Nobel Prize in Physics 2001



The Nobel Prize in Physics 2001

The Royal Swedish Academy of Sciences has awarded the Nobel Prize in Physics for 2001 jointly to Eric A. Cornell, Wolfgang Ketterle and Carl E. Wieman "for the achievement of Bose-Einstein condensation in dilute gases of alkali atoms, and for early fundamental studies of the properties of the condensates".



Eric A. Cornell
JILA and National Institute of Standards and Technology (NIST), Boulder, Colorado, USA.

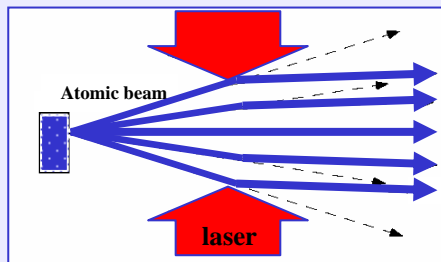
Carl E. Wieman
JILA and University of Colorado, Boulder, Colorado, USA.

Wolfgang Ketterle
Massachusetts Institute of Technology (MIT), Cambridge, Massachusetts, USA.

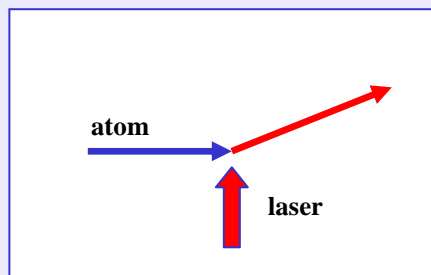
Contents:

Atom optics

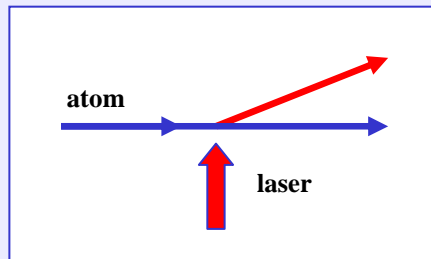
lenses



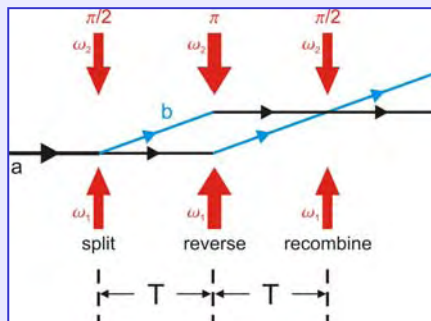
mirrors



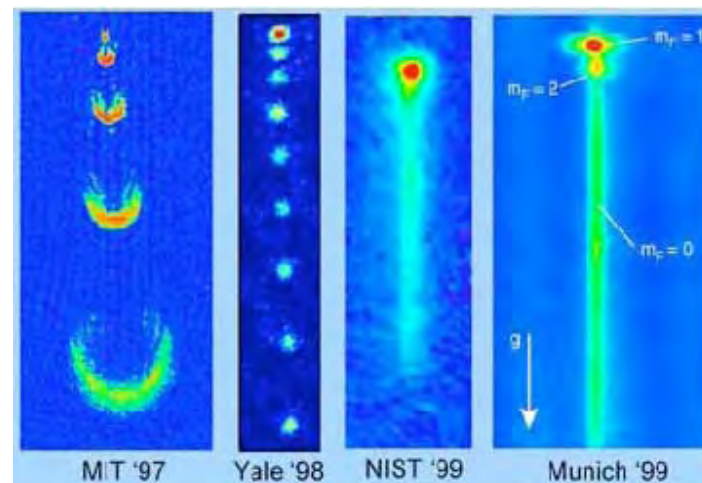
beam-splitters



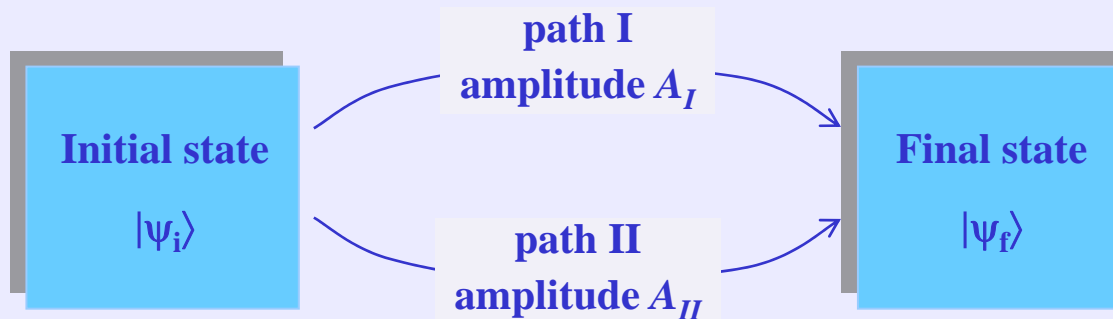
interferometers



atom laser



Quantum interference

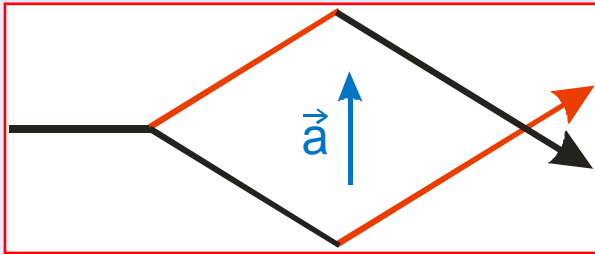


Interference of transition amplitudes

$$P(|\psi_i\rangle \Rightarrow |\psi_f\rangle) = |A_I + A_{II}|^2 = |A_I|^2 + |A_{II}|^2 + 2 \operatorname{Re}(A_I A_{II}^*)$$

Matter wave sensors

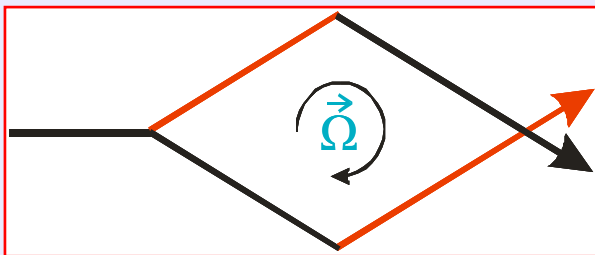
accelerations:



$$\Delta\Phi_{\text{acc}} = k T_{\text{drift}}^2 \cdot a$$

$$\frac{\Delta\varphi_{\text{mat}}}{\Delta\varphi_{\text{ph}}} \sim \left(\frac{c}{v_{\text{at}}} \right)^2 \approx 10^{11} - 10^{17}$$

rotations:



$$\Delta\Phi_{\text{rot}} = 2\pi \frac{2 m_{\text{at}}}{h} A \cdot \Omega$$

$$\frac{\Delta\varphi_{\text{mat}}}{\Delta\varphi_{\text{ph}}} \sim \frac{m_{\text{at}} \cdot \lambda \cdot c}{h} \approx 5 \cdot 10^{10}$$

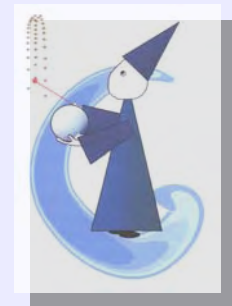


MAGIA

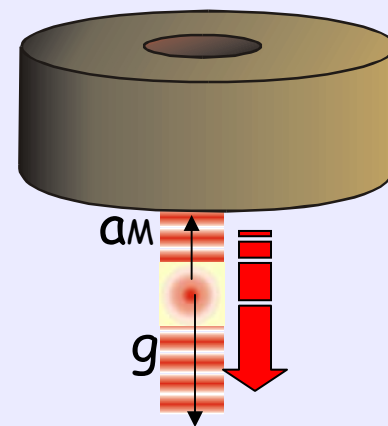
Misura Accurata di G mediante Interferometria Atomica

<http://www.fi.infn.it/sezione/esperimenti/MAGIA/home.html>

MAGIA

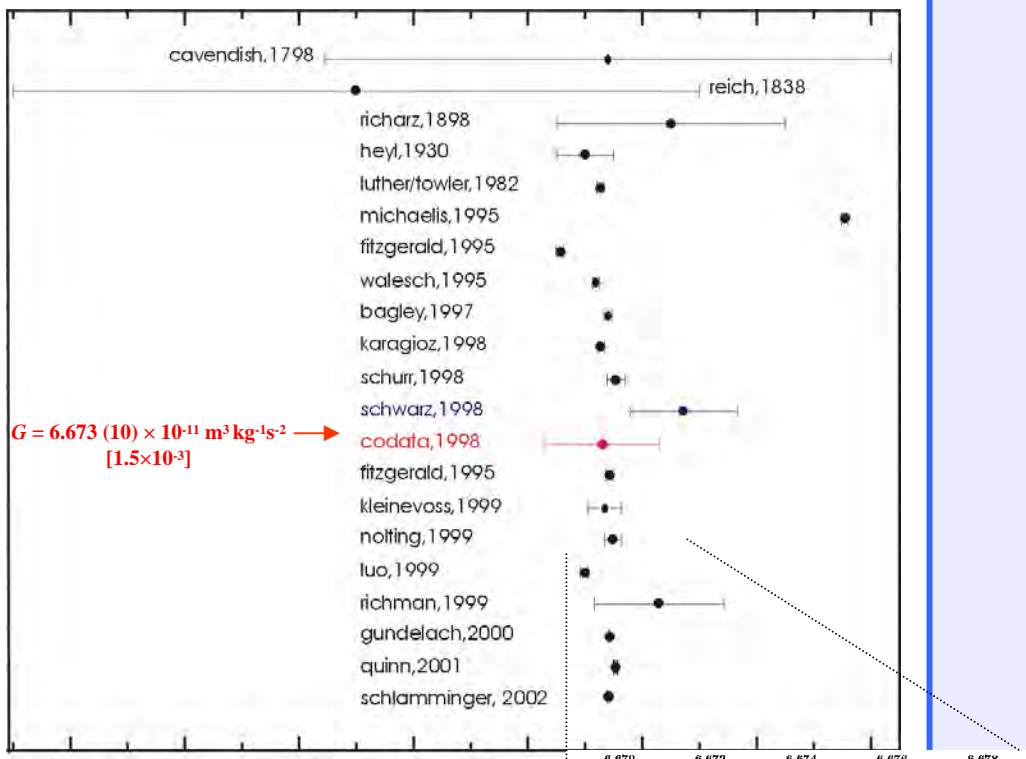


- Measure g by atom interferometry
- Add source masses
- Measure change of g

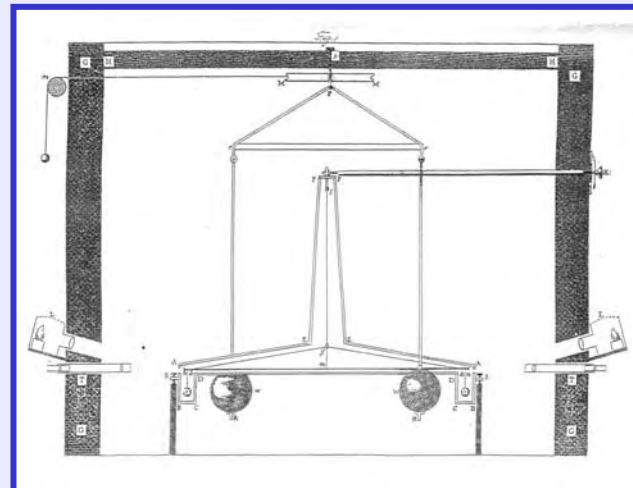


- *Precision measurement of G*
- *Measurement of G at sub-mm distances*

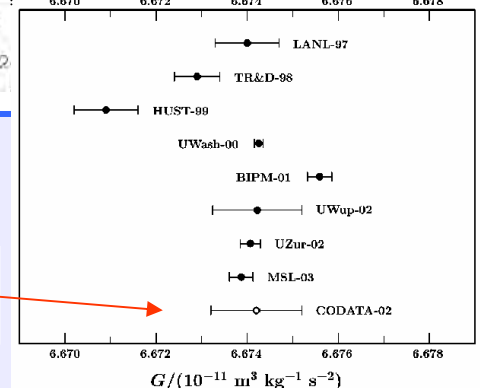
Why measuring G with atom interferometry?



$G = 6.673 (10) \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$
 $[1.5 \times 10^{-3}]$



Cavendish 1798

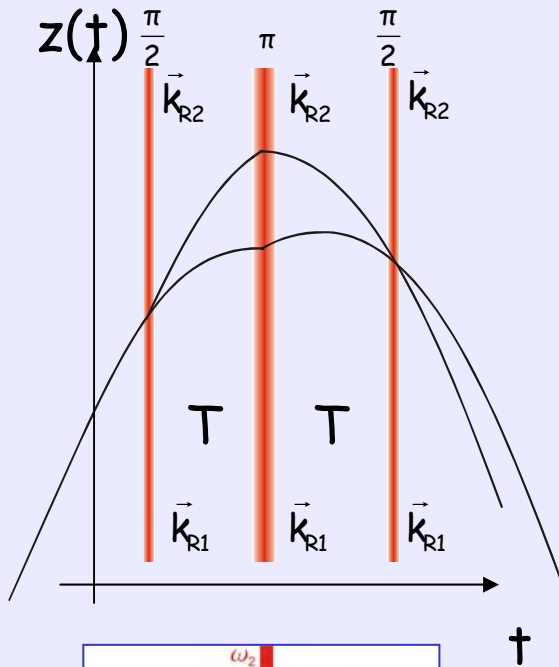


$G = 6.6742 (10) \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}$
 $[1.5 \times 10^{-4}]$



Quinn 2001

Raman interferometry in an atomic fountain



Phase difference between the paths:

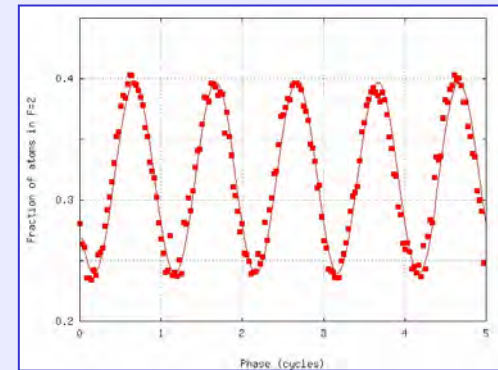
$$\Delta\Phi = k_e[z(0) - 2z(T) + z(2T)] + \Phi_e \quad \mathbf{k_e = k_1 - k_2}, \quad \omega_e = c \mathbf{k_e}$$

with $z(t) = -g t^2/2 + v_0 t + z_0$ & $\Phi_e = 0 \Rightarrow \Delta\Phi = \mathbf{k_e g T^2}$

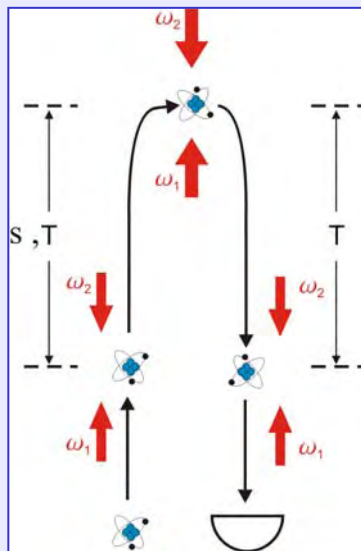
$$\mathbf{g = \Delta\Phi / k_e T^2}$$

Final population:

$$\mathbf{N_a = N/2 (1 + \cos[\Delta\Phi])}$$



Interference fringes – Firenze 2006



$$T = 150 \text{ ms} \Rightarrow 2\pi = 10^{-6} g$$

$$S/N = 1000$$

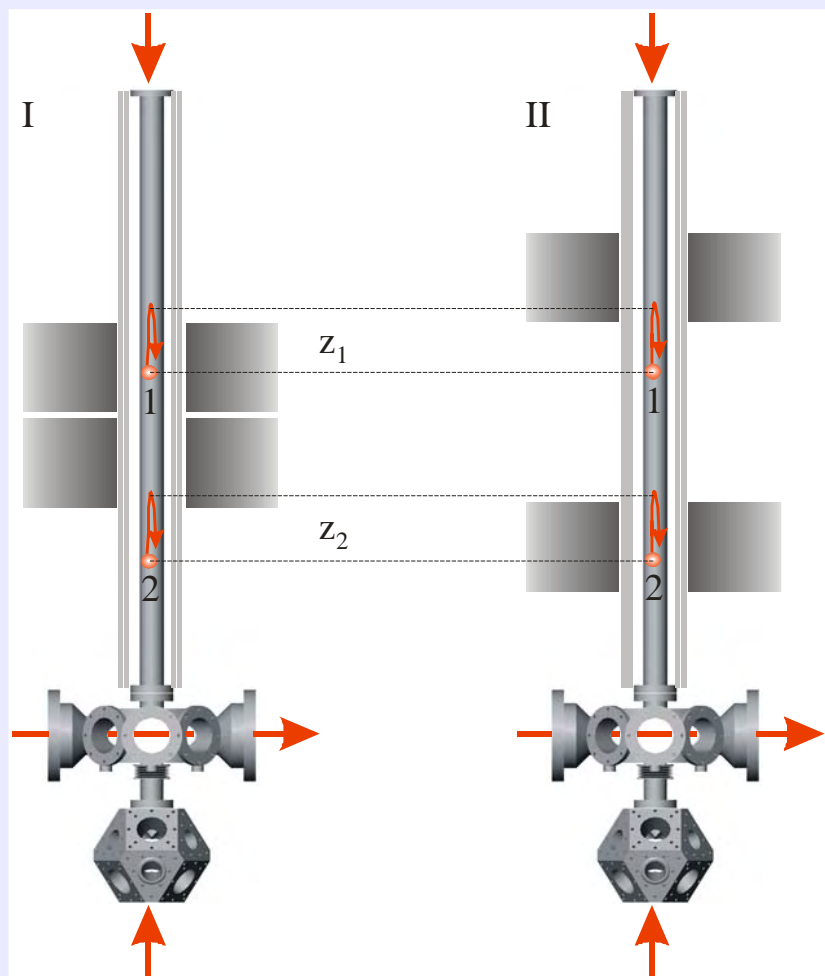
$$\Rightarrow \text{Sensitivity } 10^{-9} \text{ g/shot}$$

M. Kasevich, S. Chu, *Appl. Phys. B* **54**, 321 (1992)

A. Peters, K.Y. Chung and S. Chu, *Nature* **400**, 849 (1999)



MAGIA: Experimental procedure



- trap, cool and launch 2 clouds of Rb atoms
- apply Raman light pulses masses in position 1
- detect atoms state selectively
- repeat several times
- plot N_a/N and fit the differential phase shift $\Delta\Phi_g$ between the clouds
- move masses to position 2
- repeat all procedure
- subtract the differential phase shifts for the two mass positions

$$\phi_1^I - \phi_2^I = \phi_g(z_1) + \phi_{SM} + \phi_{Sys}(z_1, t_I) - (\phi_g(z_2) - \phi_{SM} + \phi_{Sys}(z_2, t_I))$$

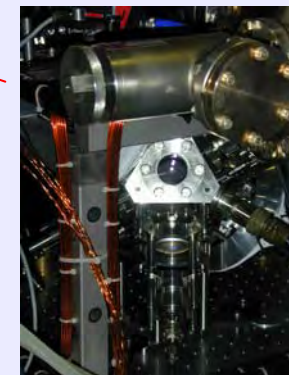
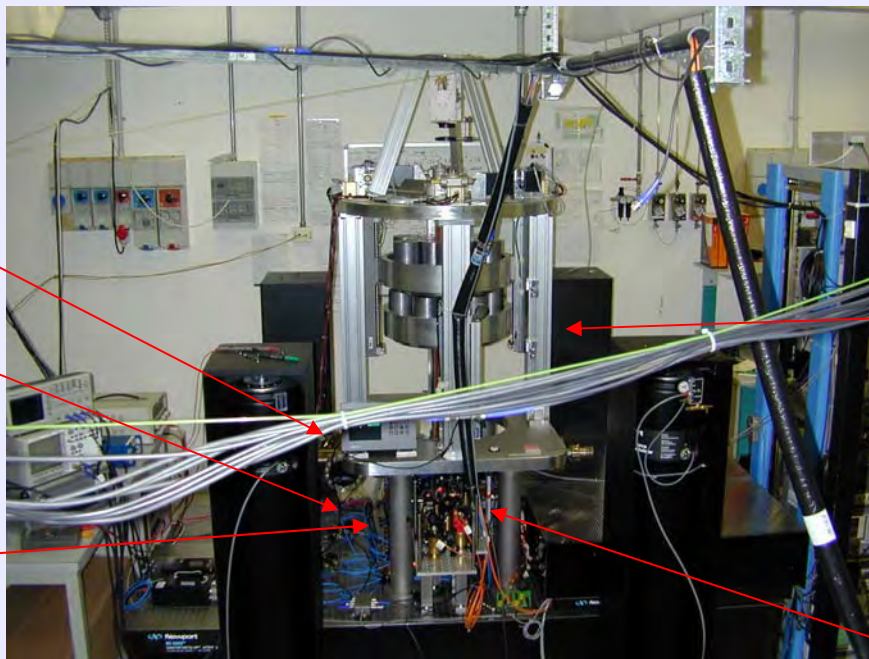
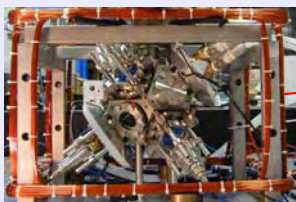
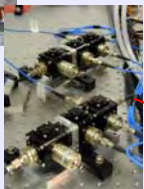
$$\phi_1^{II} - \phi_2^{II} = \phi_g(z_1) - \phi_{SM} + \phi_{Sys}(z_1, t_{II}) - (\phi_g(z_2) + \phi_{SM} + \phi_{Sys}(z_2, t_{II}))$$

$$\Rightarrow (\phi_1^I - \phi_2^I) - (\phi_1^{II} - \phi_2^{II}) = 4\phi_{SM} + \phi_{Sys}(\Delta z, \Delta t)$$

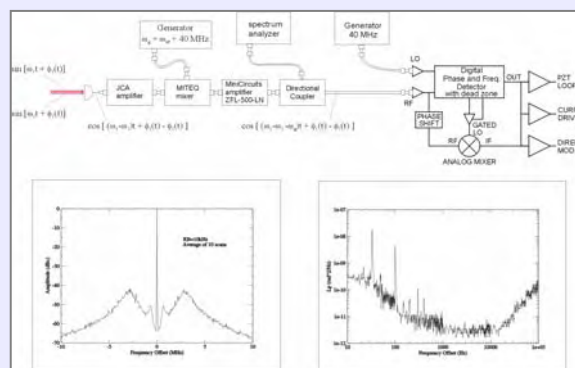


Atom gravity gradiometer apparatus

Source masses and support



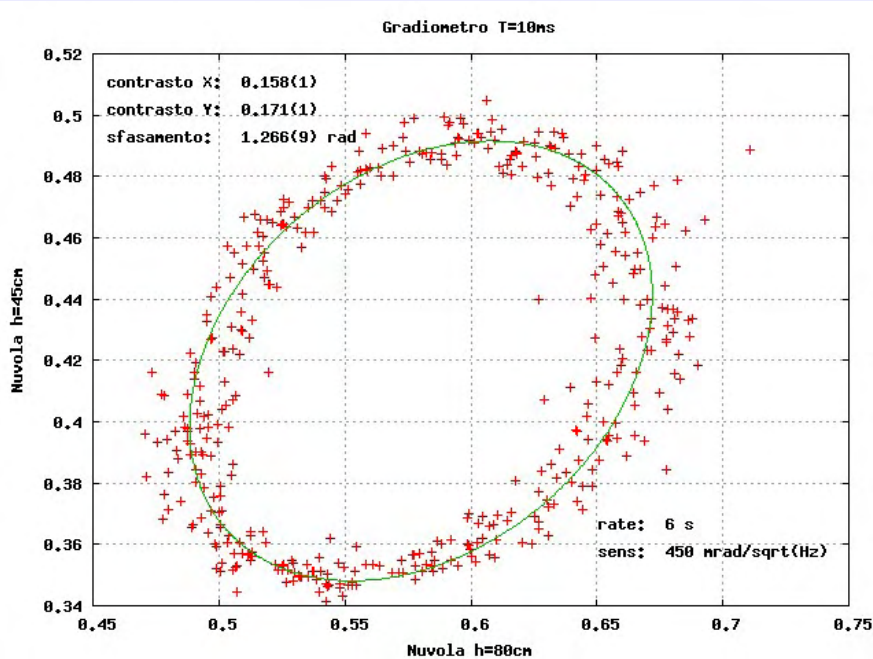
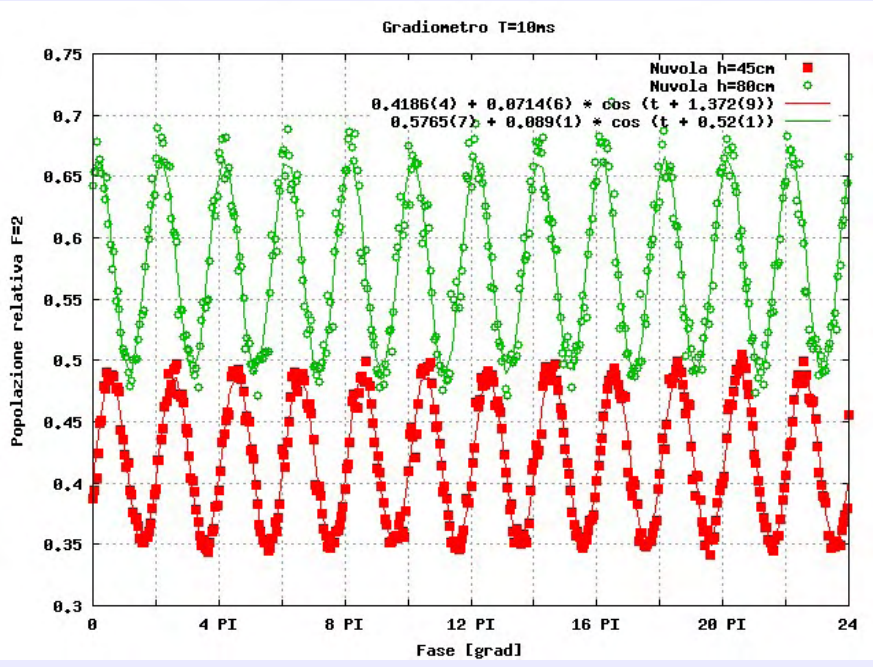
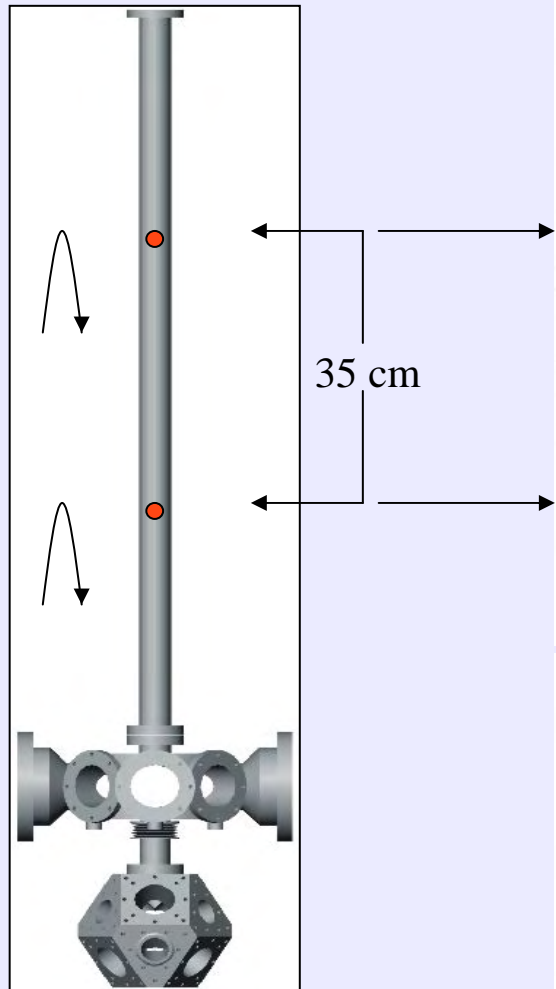
Laser and optical system



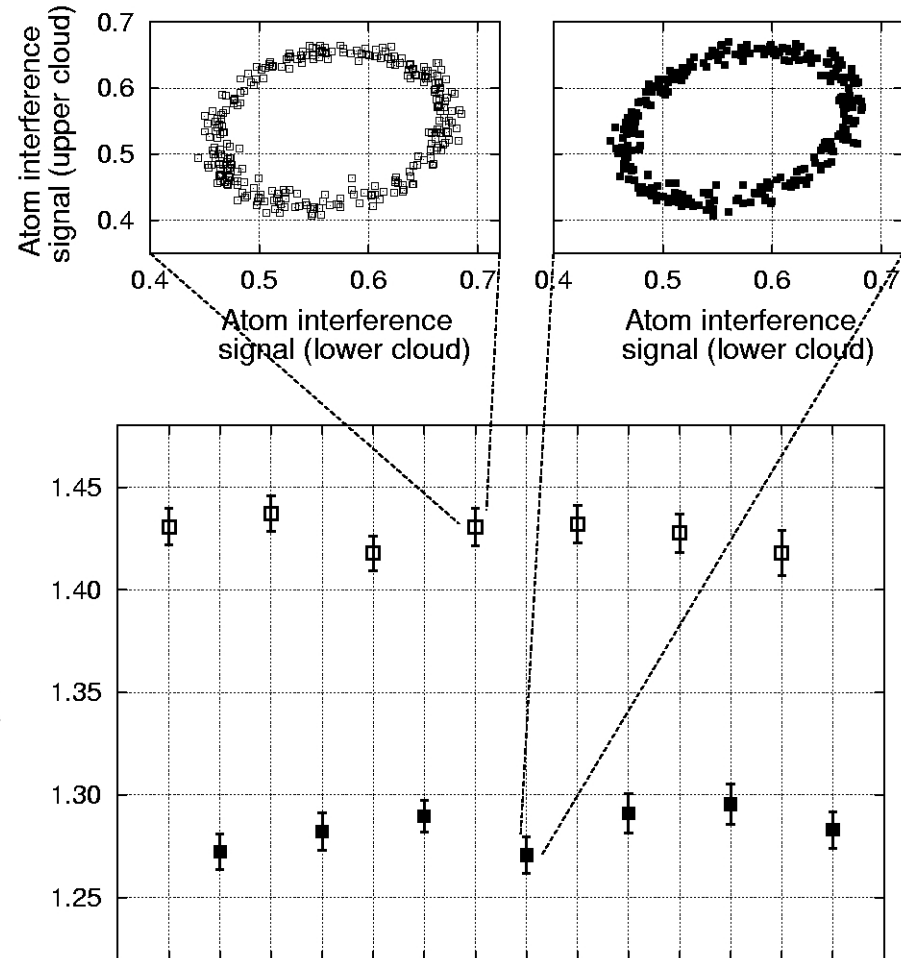
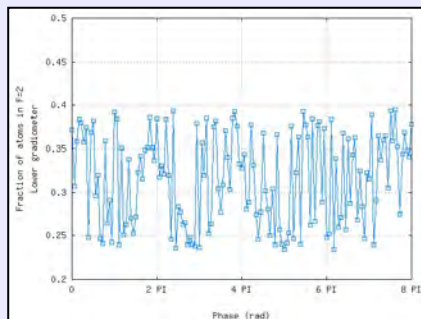
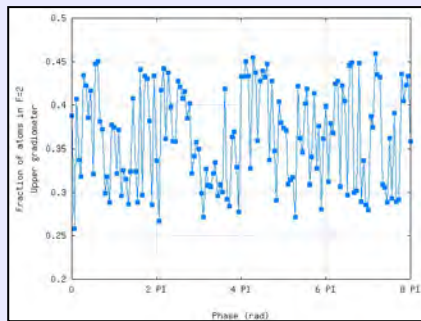
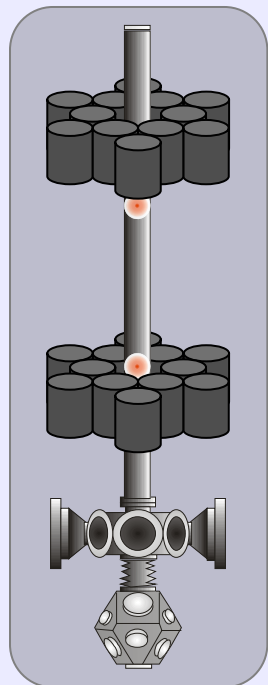
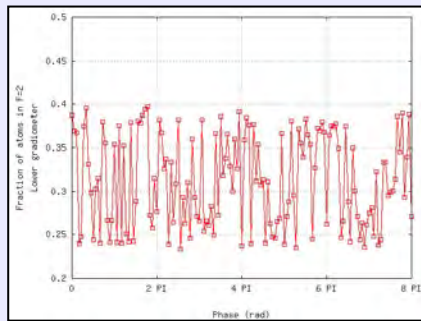
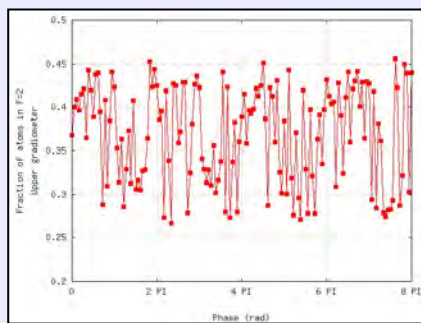
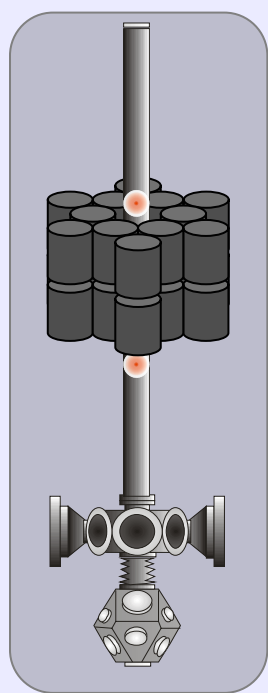
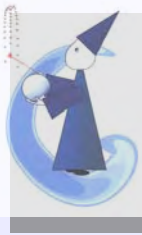
L. Cacciapuoti, M.de Angelis, M. Fattori, G. Lamporesi, T. Petelski, M.Prevedelli, J. Stuhler, G.M. Tino, *Analog+digital phase and frequency detector for phase locking of diode lasers*, Rev. Scient. Instr. 76, 053111 (2005)

GRADIOMETER

Atom interference fringes



G: first results

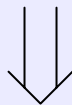


A. Bertoldi, G. Lamporesi, L. Cacciapuoti, M. de Angelis, M. Fattori, T. Petelski, A. Peters, M. Prevedelli, J. Stuhler, G. M. Tino (2006), submitted, arXiv:physics/0606126

MAGIA – Relevant numbers

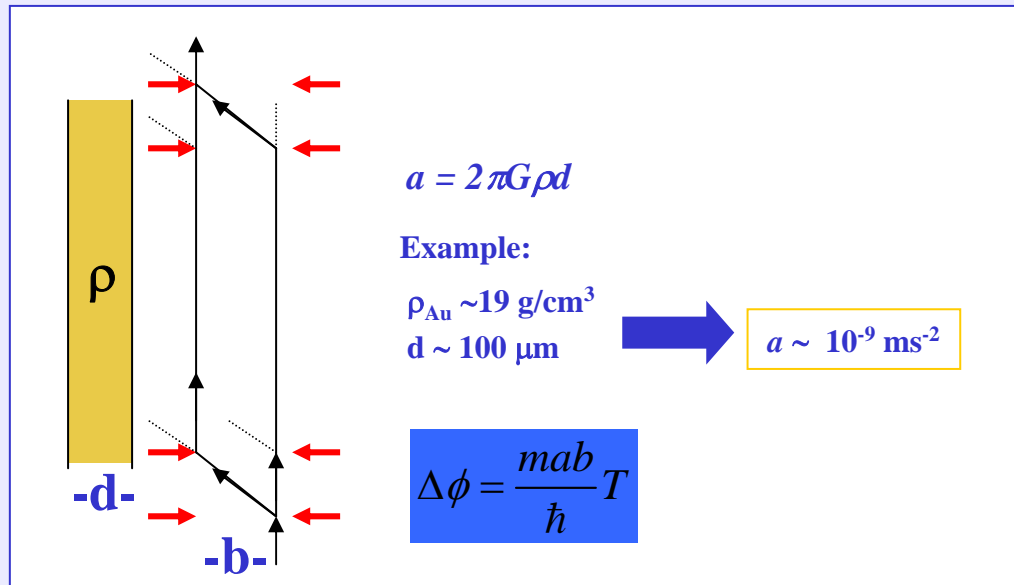


- time separation between pulses $T=150$ ms
- 10^6 atoms
- shot noise limited detection
- launch accuracy: 1 mm e $\Delta v \sim 5$ mm/s
- knowledge of the masses dimensions and relative positions: 10 μm
- 10000 measurements



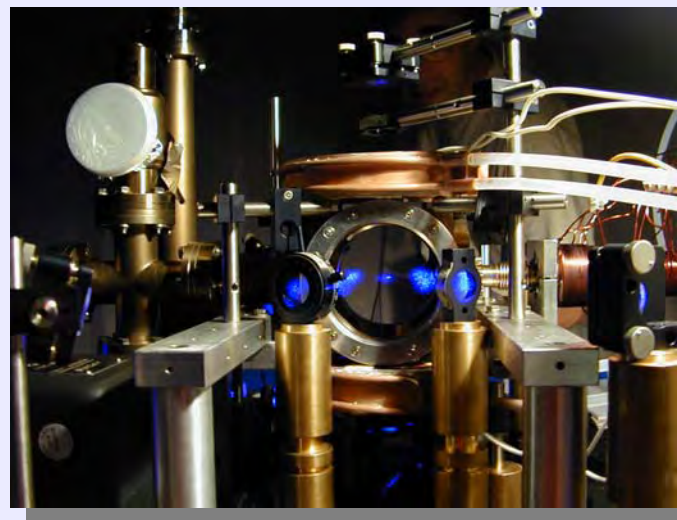
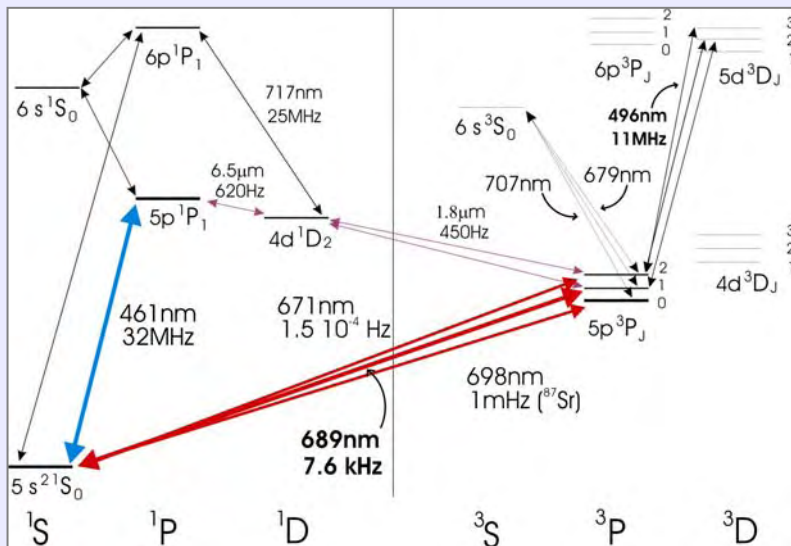
$$\Delta G/G \leq 10^{-4}$$

Possible test of the gravitational $1/r^2$ law in the sub-mm range with an atom interferometer



- G.M. Tino, in “2001: A Relativistic Spacetime Odyssey”, Firenze, 2001, World Scientific (2003)
- G.M. Tino, Nucl. Phys. B 113, 289 (2002)

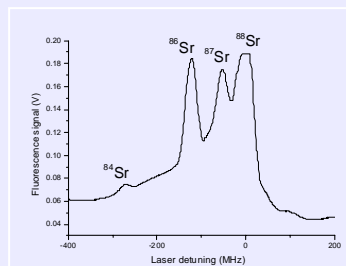
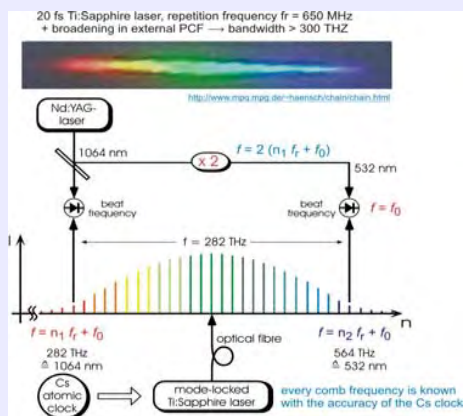
Ultracold Sr – The experiment in Firenze



Firenze 2003, Magneto-optical trapping of all Sr isotopes

- Optical clocks using visible intercombination lines

- New atomic sensors for fundamental physics tests



| | Abundance |
|------------------|-----------|
| ⁸⁸ Sr | 82.6% |
| ⁸⁶ Sr | 9.9% |
| ⁸⁷ Sr | 7.0% |
| ⁸⁴ Sr | 0.6% |

| Isotope | I | transition | lifetime | λ | t_{int} | $\sigma_y t^{-1/2}$ | abundance |
|------------------|-----|----------------|------------------|-----------|------------------|---------------------|-----------|
| ⁸⁸ Sr | 0 | $1S_0 - ^3P_1$ | 20 μs | 689 nm | 10 μs | $2 \cdot 10^{-13}$ | 83% |
| ⁸⁷ Sr | 9/2 | $1S_0 - ^3P_0$ | 200 s | 698 nm | 0.5 s | 10^{-17} | 7% |

G. Ferrari, P. Cancio, R. Drullinger, G. Giusfredi, N. Poli, M. Prevedelli, C. Toninelli, G.M. Tino, *Precision Frequency Measurement of Visible Intercombination Lines of Strontium*, Phys. Rev. Lett. 91, 243002 (2003)

G. Ferrari, R.E. Drullinger, N. Poli, F. Sorrentino, G.M. Tino, *Cooling of Sr to high phase-space density by laser and sympathetic cooling in isotopic mixtures*, Phys. Rev. A 73, 023408 (2006)

Why Sr atom?

• New quantum sensors

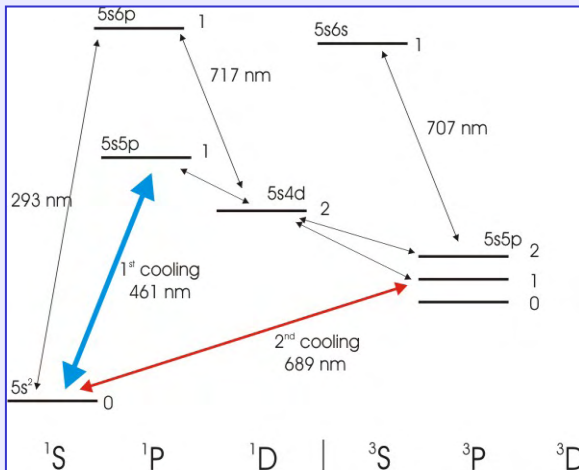
- $J = 0$ ground state
- Small collisional cross section
- Fermionic and bosonic atoms

• Optical clocks on visible intercombination lines

- $^1S_0 - ^3P_1$ (7.5 kHz) (this work)
- $^1S_0 - ^3P_0$ (1 mHz, ^{87}Sr)
- $^1S_0 - ^3P_2$ (<1 mHz)
- Optical trapping in optical lattices with negligible change of clock frequency

• Physics of ultracold atoms

- Simple 0 - 1 transitions
- $T_D \approx T_{\text{rec}}$ for $^1S_0 - ^3P_1$ transition
- Two-stage optical cooling and trapping
- All-optical cooling to quantum degeneracy
- Degenerate Bose and Fermi gases



| Sr Isotope | Nuclear Spin I | Atomic Mass (10^{-27} Kg) | Natural Abundance |
|------------|----------------|------------------------------|-------------------|
| 88 | 0 | 145,97068 | 82,6 % |
| 87 | 9/2 | 144,315568 | 7,0 % |
| 86 | 0 | 142,65567 | 9,8 % |
| 84 | 0 | 139,34150 | 0,56 % |

Optical clocks: Towards 10^{-19}

- Narrow optical transitions

$$\delta\nu_0 \sim 1 \text{ Hz}, \nu_0 \sim 10^{15} \text{ Hz}$$

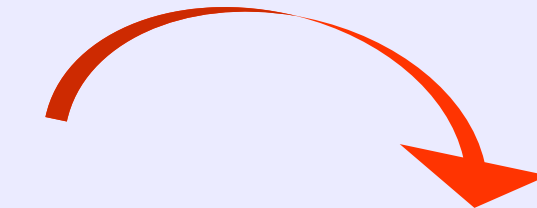
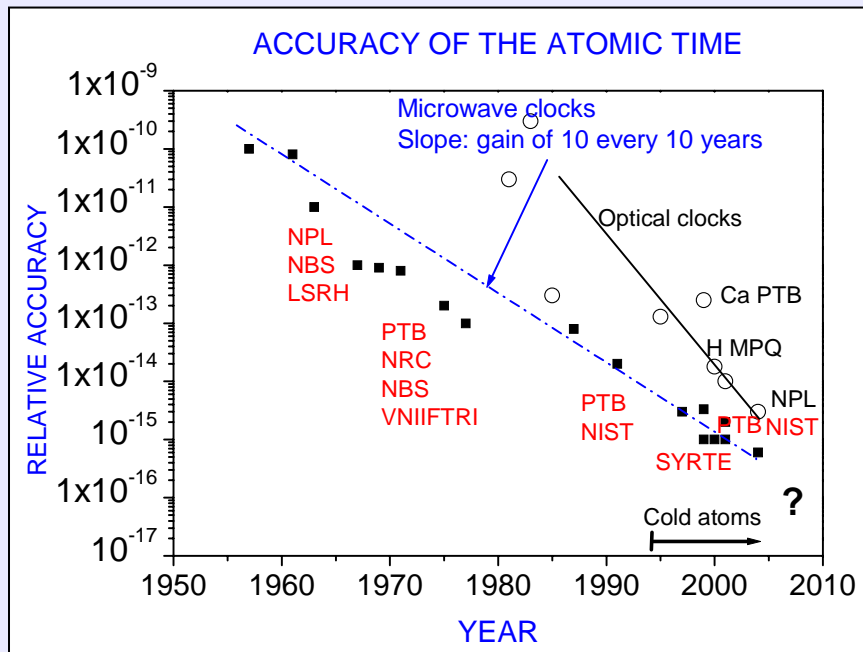
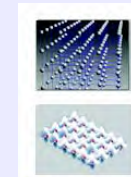
$$\sigma_y \approx \frac{\text{Noise}}{\pi Q \cdot \text{Signal}} \approx \frac{\Delta\nu}{\nu_0} \frac{1}{\sqrt{N_{\text{atom}}}} \sqrt{\frac{T_{\text{cycle}}}{2\tau_{\text{average}}}} \frac{1}{C_{\text{fringe}}}$$

- Candidate atoms

Trapped ions: $\text{Hg}^+, \text{In}^+, \text{Sr}^+, \text{Yb}^+, \dots$



Cold neutral atoms: $\text{H}, \text{Ca}, \text{Sr}, \text{Yb}, \dots$



Gravitational red shift at lab scale



The Nobel Prize in Physics 2005


"for his contribution to the quantum theory of optical coherence"

"for their contributions to the development of laser-based precision spectroscopy, including the optical frequency comb technique"



photo J. J. Ted

Roy J. Glauber

 1/2 of the prize

USA


Harvard University
Cambridge, MA, USA

b. 1925



photo C. J. L. Harwood

John L. Hall

 1/4 of the prize

USA


University of Colorado,
JILA; National Institute
of Standards and
Technology
Boulder, CO, USA

b. 1934



photo M. Urban

Theodor W. Hänsch

 1/4 of the prize

Germany

Max-Planck-Institut für
Quantenoptik
Garching, Germany;
Ludwig-Maximilians-Universität
München, Germany

b. 1941

The Nobel Prize in Physics 2005

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Roy J. Glauber

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 2004

The 2005 Prize in:

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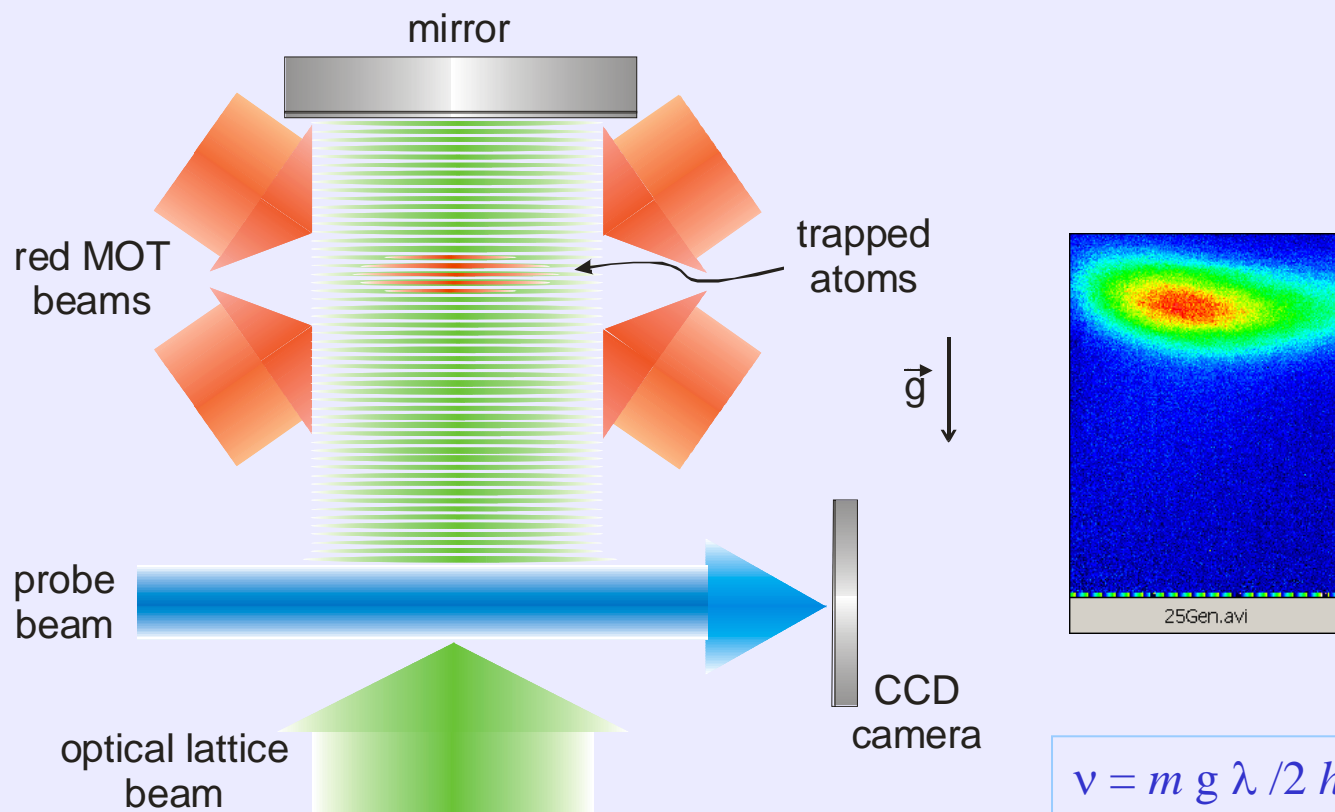
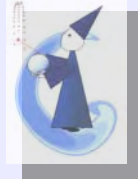
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Find a Laureate:

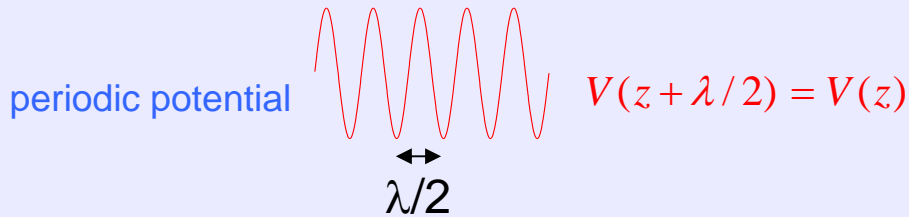


Precision gravity measurement at μm scale with Bloch oscillations of Sr atoms in an optical lattice



G. Ferrari, N. Poli, F. Sorrentino, and G. M. Tino, *Long-lived Bloch oscillations with bosonic Sr atoms and application to gravity measurement at micrometer scale* (2006) to be published, arXiv:physics/0605018

Particle in a periodic potential: Bloch oscillations

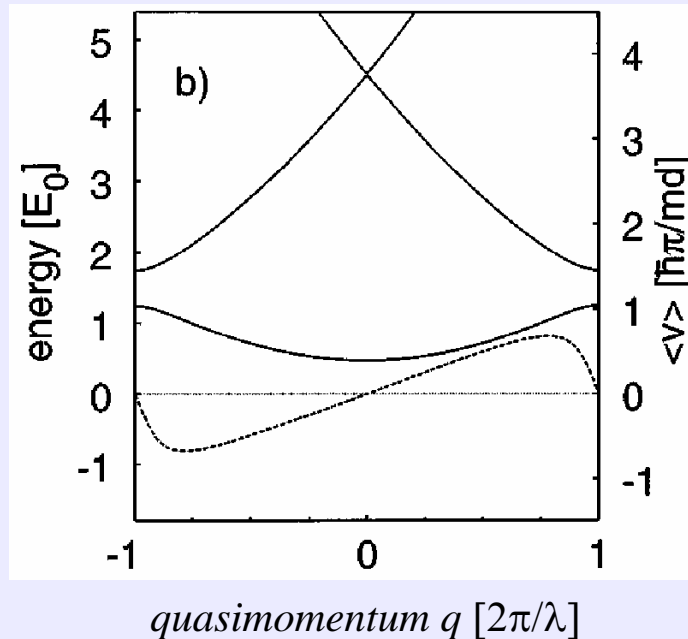


$$\Psi(z) = e^{i\frac{\mathbf{q}}{\hbar}z} u(z)$$

$$u(z + \lambda/2) = u(z)$$

Bloch's theorem

$$\Psi(z + \lambda/2) = e^{i\frac{\mathbf{q} \cdot \lambda}{\hbar}} \Psi(z)$$



$$\langle v \rangle_n(q(t)) = \frac{1}{\hbar} \frac{dE_n(R(q(t)))}{dq}$$

with a constant external force F

$$q(t) = q(0) + Ft/\hbar$$

↓
Bloch oscillations

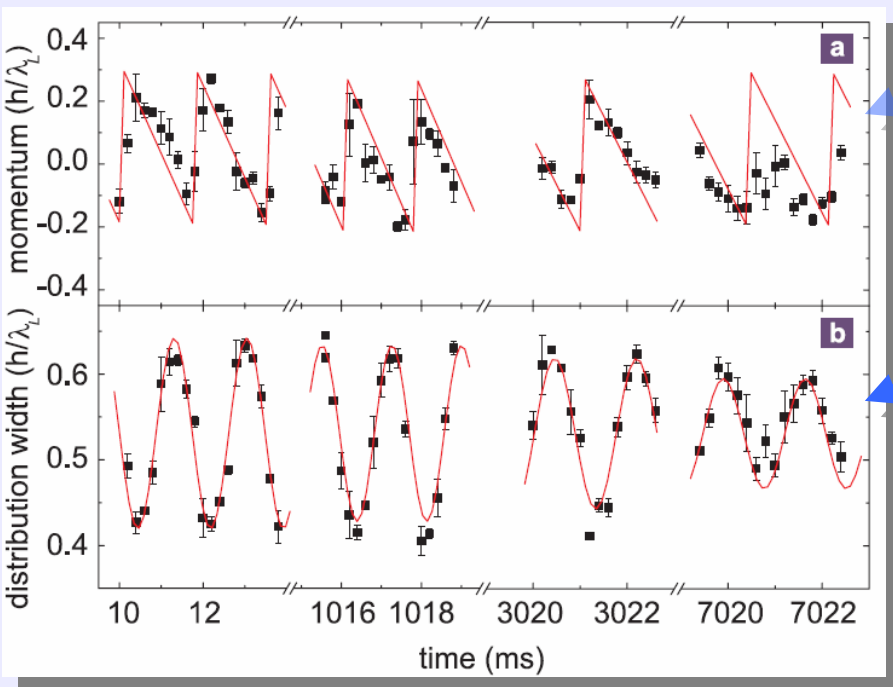
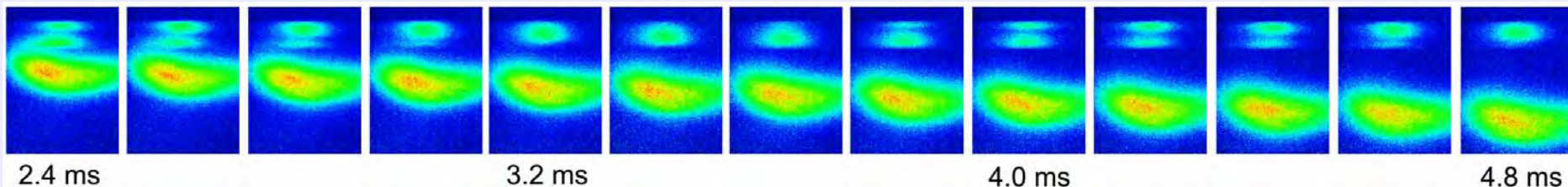
Quantum theory for electrons in crystal lattices: **F. Bloch**, *Z. Phys.* **52**, 555 (1929)

Never observed in natural crystals (evidence in artificial superlattices)

Direct observation with Cs atoms: **M. Ben Dahan, E. Peik, J. Reichel, Y. Castin, C. Salomon**, *PRL* **76**, 4508 (1996)



Persistent Bloch oscillations



average vertical momentum of the lower peak

width of the atomic momentum distribution

Bloch frequency $\nu_B = 574.568(3)$ Hz

damping time $\tau = 12$ s

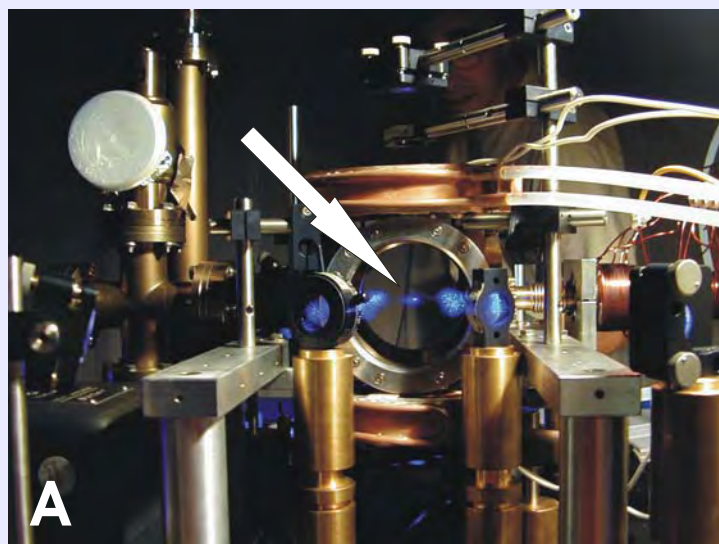
8000 photon recoils in 7s

$g_{\text{meas}} = 9.80012(5)$ ms⁻²

G. Ferrari, N. Poli, F. Sorrentino, and G. M. Tino, *Long-lived Bloch oscillations with bosonic Sr atoms and application to gravity measurement at micrometer scale* (2006) to be published, arXiv:physics/0605018

G. M. Tino, GGI Colloquium, Arcetri 22/6/2006

Decoherence

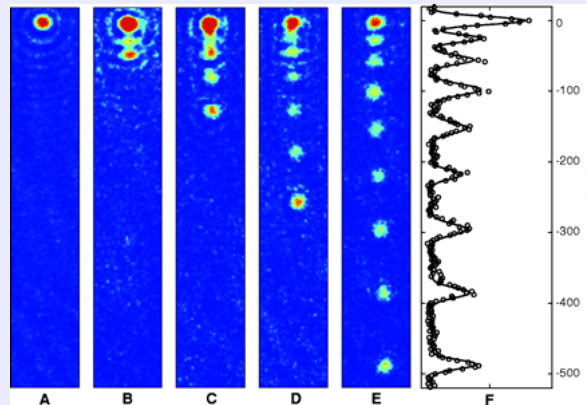


- ^{88}Sr : $\sigma_{88-88} = (3 \pm 1) \times 10^{-13} \text{ cm}^2$ \longrightarrow $\tau_{\text{coh}} = 1-10 \text{ s}$
- ^{87}Sr : suppressed collisions for fermions at low temperatures

• N. Poli, R.E. Drullinger, G. Ferrari, J. Leonard, F. Sorrentino, G.M. Tino, *Cooling and trapping of ultracold strontium isotopic mixtures*, Phys. Rev A 71, 061403 (R) (2005)

• G. Ferrari, R.E. Drullinger, N. Poli, F. Sorrentino, G.M. Tino, *Cooling of Sr to high phase-space density by laser and sympathetic cooling in isotopic mixtures*, Phys. Rev. A 73, 023408 (2006)

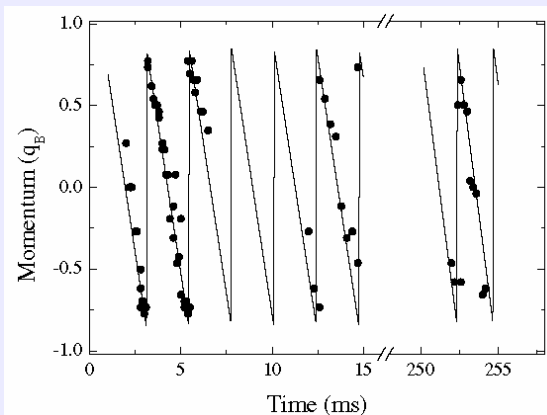
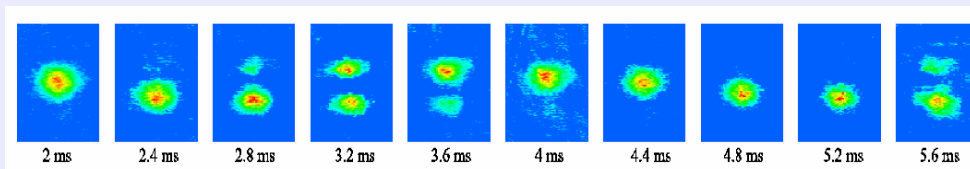
gravity measurement with quantum degenerate bosonic and fermionic atoms in optical lattices



$$g = 9.6 (4) \text{ m/s}^2$$

$$\delta g/g \sim 10^{-4} \text{ in } 10 \text{ ms}$$

B. P. Anderson, M. A. Kasevich, *Macroscopic Quantum Interference from Atomic Tunnel Arrays*, Science 282, 1686 (1998)



$$g = 9.7372 (9) \text{ m/s}^2$$

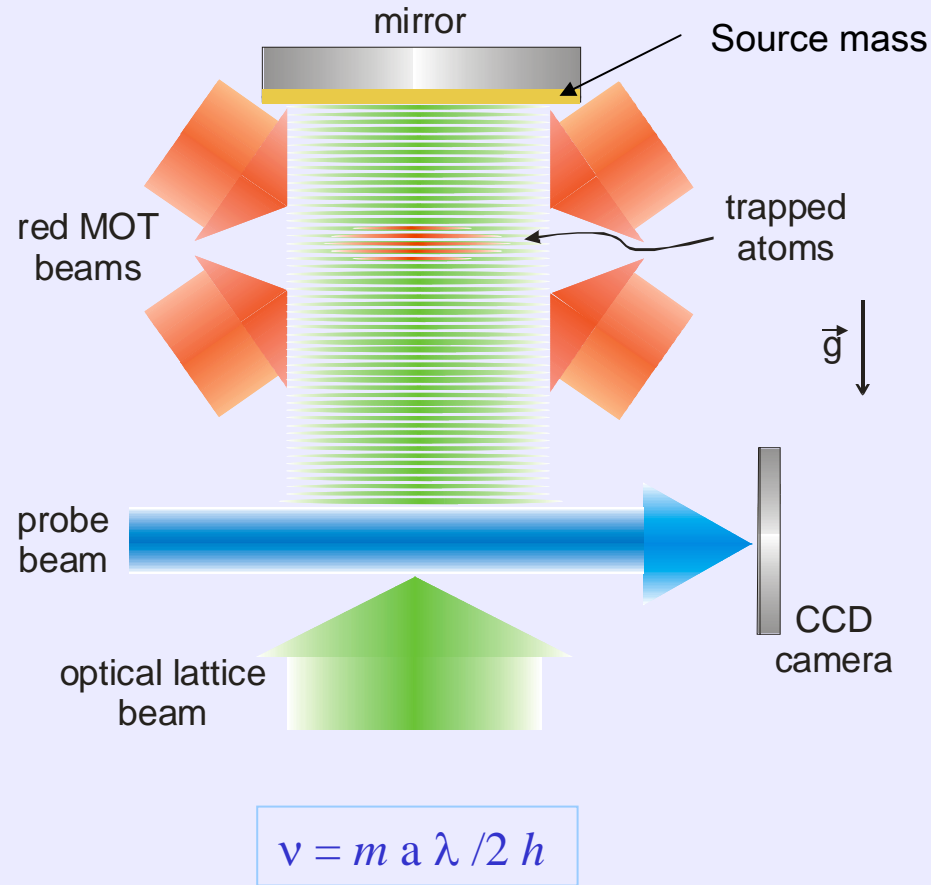
$$\delta g/g \sim 10^{-4} \text{ in } 250 \text{ ms}$$

G. Roati, E. de Mirandes, F. Ferlaino, H. Ott, G. Modugno, M. Inguscio, *Atom Interferometry with trapped Fermi Gases*, PRL 92, 230402 (2004)

Precision gravity measurement with ultracold Sr atoms in an optical lattice:

- **Measured gravity acceleration:** $g_{\text{meas}} = 9.80012 (5) \text{ ms}^{-2}$
From geophysical data: $g_{\text{ref}} = 9.805046 (9) \text{ ms}^{-2}$
- **Present sensitivity:** $5 * 10^{-6} \text{ g}$
- **Achievable sensitivity:** $\sim 10^{-7} \text{ g}$

Scheme for the measurement of small distance forces

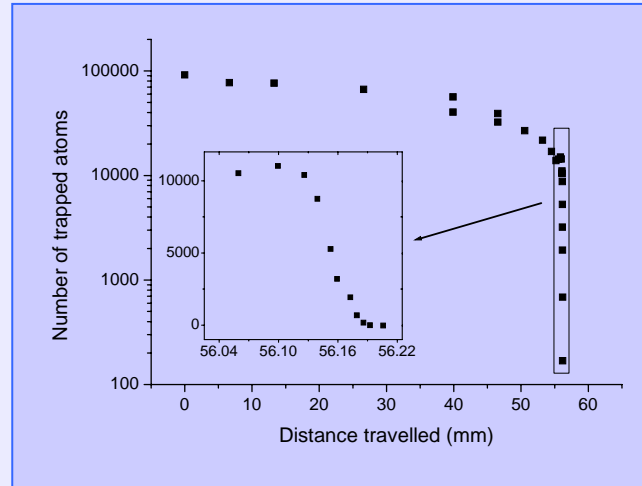
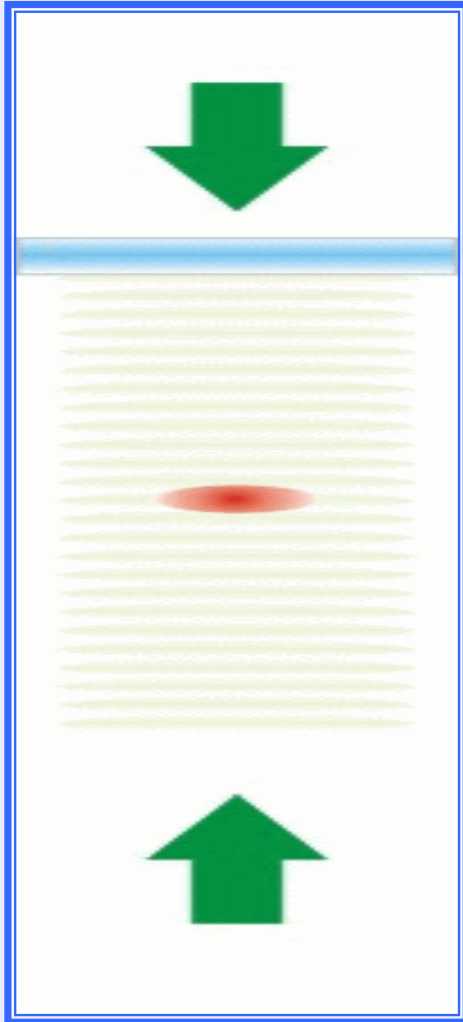


Objective: $\lambda = 1- 10 \mu m, \alpha = 10^3-10^4$

G. Ferrari, G.M. Tino, INFN note, 2005

G. Ferrari, N. Poli, F. Sorrentino & G. M. Tino, 2006, to be published

Preliminary experiment

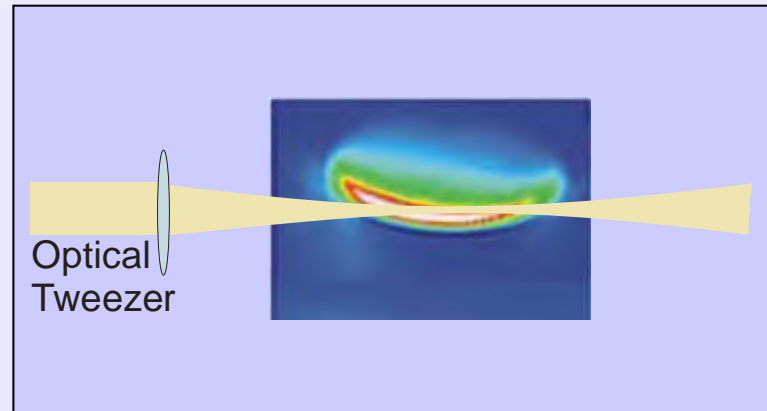


Vertical size of the atomic sample: $15 \mu\text{m}$

Atom elevator:

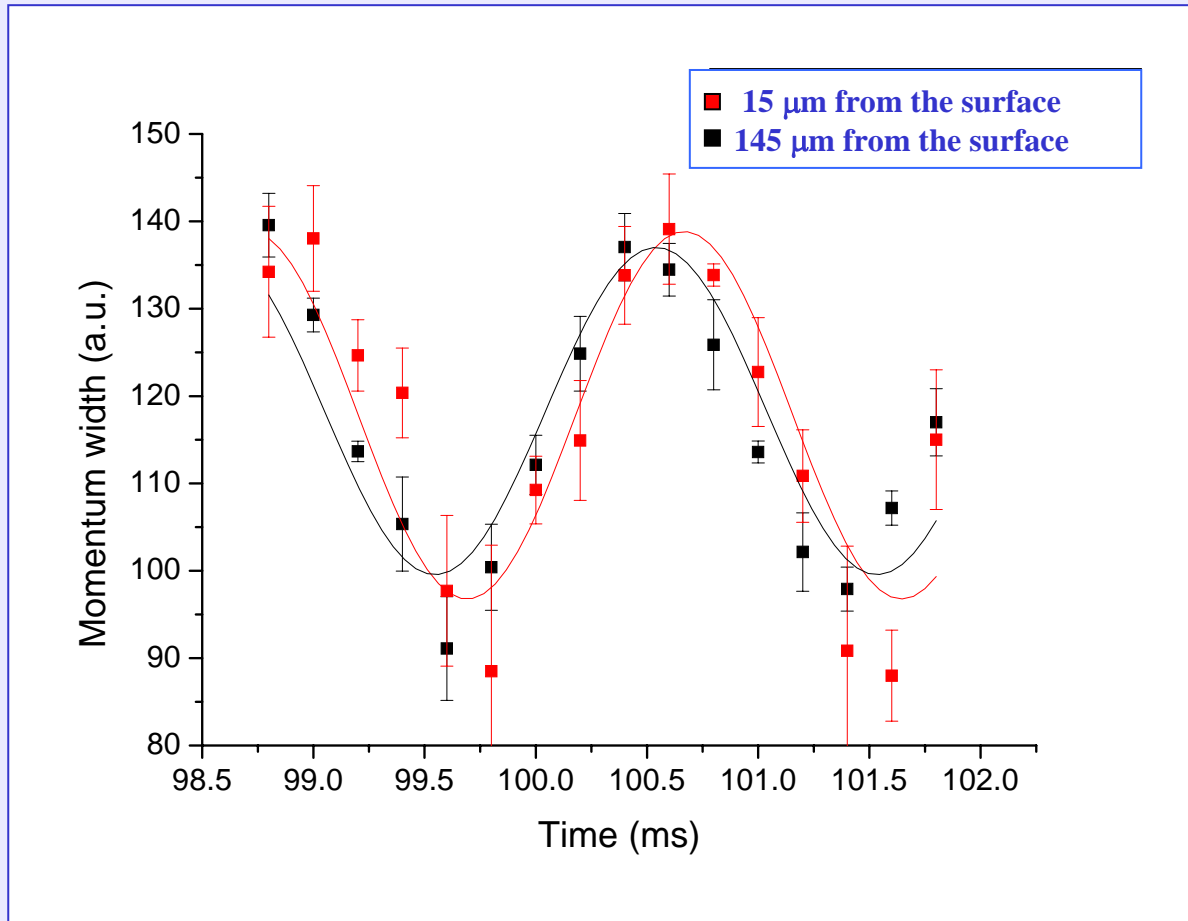
- upward acceleration (1.35 g) for 10 ms
- uniform velocity (133 mm/s) for variable time
- downward acceleration (-1.35 g) for 10 ms
- rest for 470 ms
- reverse motion back to the starting point

Vertical position fluctuations: $3 \mu\text{m rms}$



- Vertical size reduced to $4 \mu\text{m}$ with an optical tweezer

Surface effect? (Preliminary)



Atom position varied by changing the duration of uniform velocity motion
Phase shift between 145 μm and 15 μm: 0.4 rad in 235 oscillations
 $\Delta F \approx 10^{-3}$ gravity

Other experiments on atom-surface force

Hinds (1993)

Aspect-Westbrook (1997)

Vuletic (2004): effect of Casimir-Polder on atoms on chip

Shimuzu, Ketterle (2001,2005): effect of Casimir-Polder on quantum reflection

Cornell (2005): measurement of the Casimir-Polder force by oscillations of a BEC

Casimir-Polder and surface effects

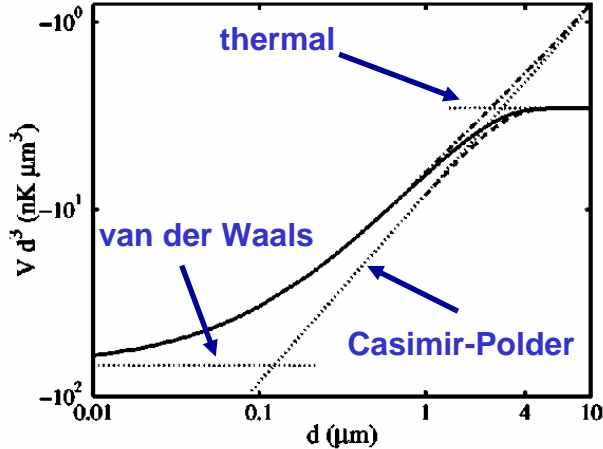


FIG. 3. The atom-surface potential is shown using the exact formula of Eq. (17) (solid line), the short-range approximation (21) (dash-dotted line), and the static approximation (26) (dashed line). The asymptotic van der Waals–London ($\sim 1/d^3$), Casimir-Polder ($\sim 1/d^4$), and high- T ($\sim 1/d^4$) potentials are also shown (dotted lines). The curves were obtained for a sapphire substrate at 300 K and for ^{87}Rb atoms in the condensate.

thermal

$$F_T = -\frac{3kT\alpha_0(\varepsilon_0 - 1)}{4z^4(\varepsilon_0 + 1)} \quad z \gg \lambda_T$$

Casimir-Polder

$$F_{CP} = -\frac{3\hbar c\alpha_0(\varepsilon_0 - 1)\phi(\varepsilon_0)}{2\pi z^5(\varepsilon_0 + 1)} \quad \lambda_{opt} \ll z \ll \lambda_T$$

^{87}Rb atoms

Sapphire substrate at 300 K

$$\lambda_T = \hbar c/k_B T \sim 7.6 \mu\text{m}$$

$$\lambda_{opt} \sim 0.1 \mu\text{m}$$

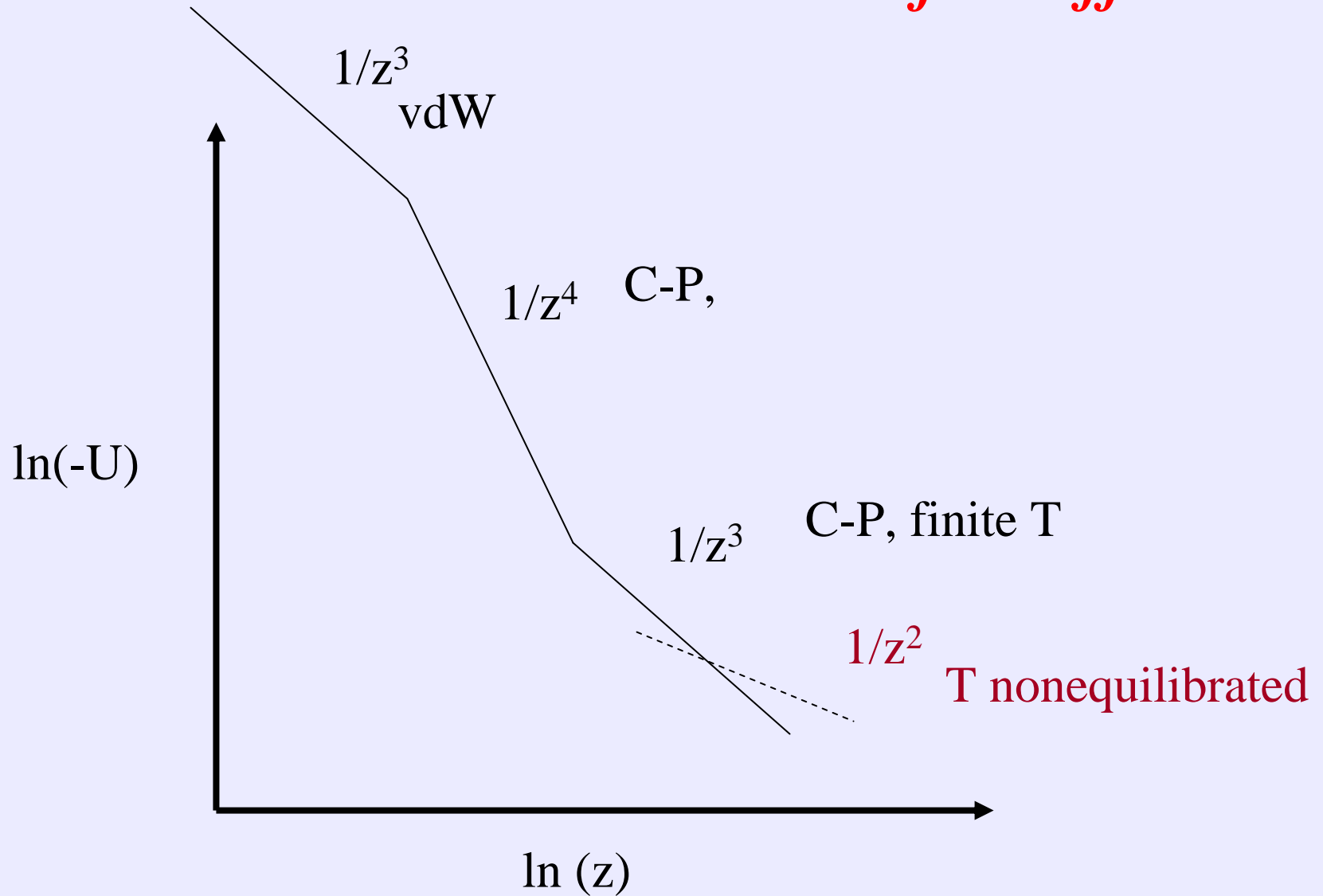
From M. Antezza et al., *Effect of the C-P force on the collective oscillations of a trapped BEC*, PRA 70, 053619 (2004)

surface-atom force
out of thermal equilibrium

$$F^{neq} = -\frac{\pi \alpha_0 k_B^2 (T_{Surface}^2 - T_{Envir}^2)}{6 z^3 c \hbar} \frac{\varepsilon_0 + 1}{\sqrt{\varepsilon_0 - 1}}$$

M. Antezza et al., *New Asymptotic Behavior of the Surface-Atom Force out of Thermal Equilibrium* PRL 95, 113202 (2005)

Casimir-Polder and surface effects



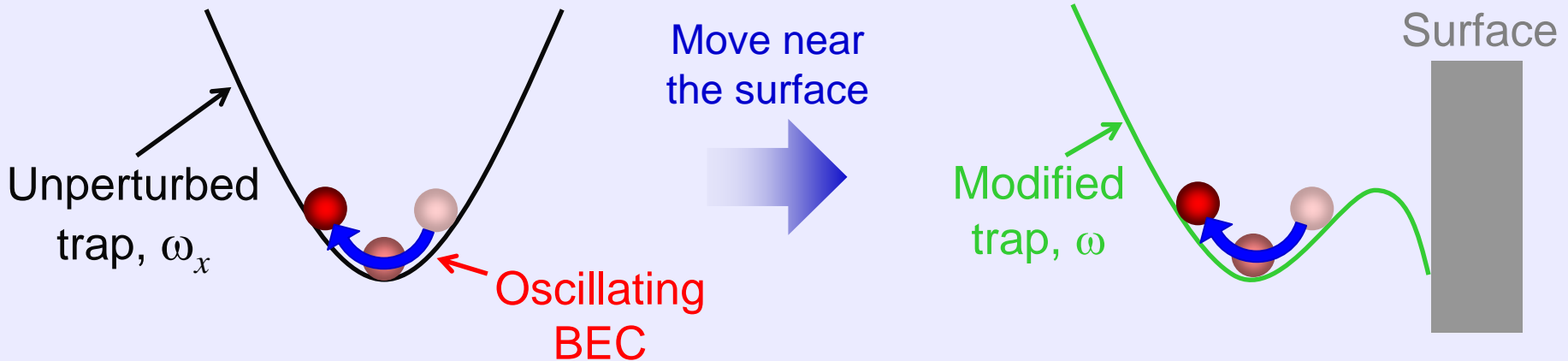
Measuring atom-surface forces

Use trapped BEC as a mechanical oscillator

Measure changes in dipole oscillation frequency

Negative curvature
attractive potential

Trap frequency decrease

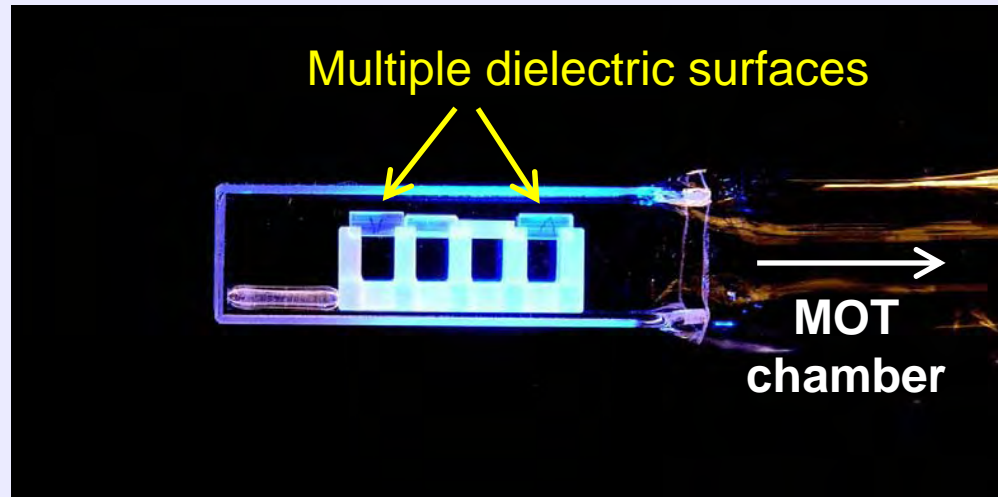


Express trap frequency changes as normalized frequency shifts:

From E.A. Cornell
San Feliu Conference, 2005

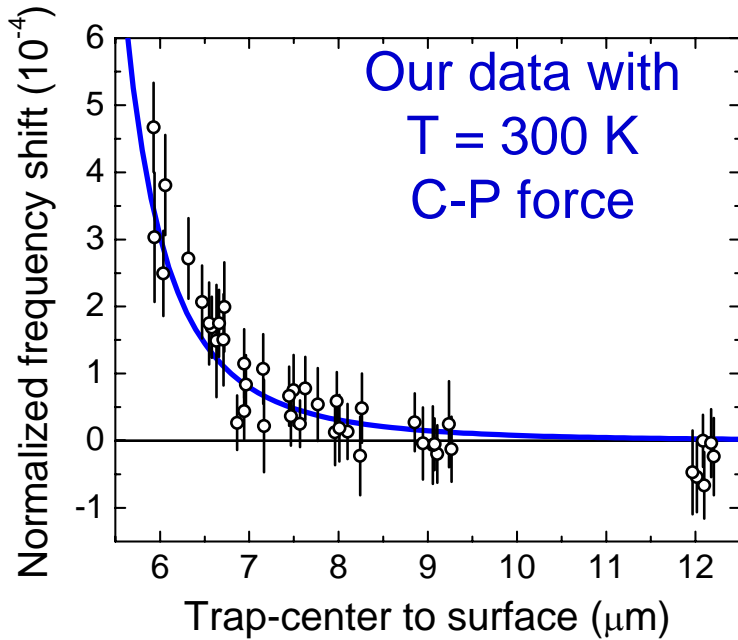
$$\frac{\omega_x - \omega}{\omega_x} \approx -\frac{1}{2\omega_x^2 m} \frac{d^2U}{dx^2}$$

Casimir-Polder force near a dielectric surface

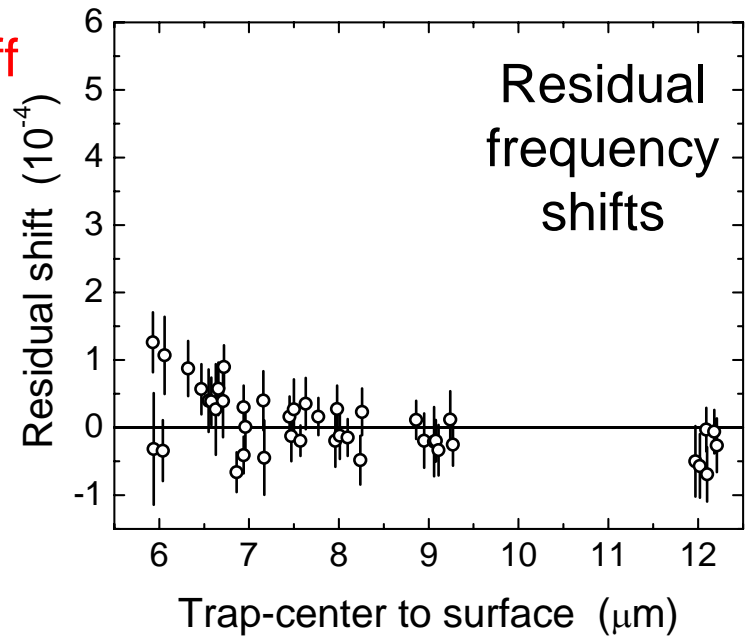


From E.A. Cornell
San Feliu Conference, 2005

The *absence* of forces in addition to C-P force allows us to obtain limits from our data:



Subtract off
C-P force



Use residuals to obtain a limit on
the presence of additional forces

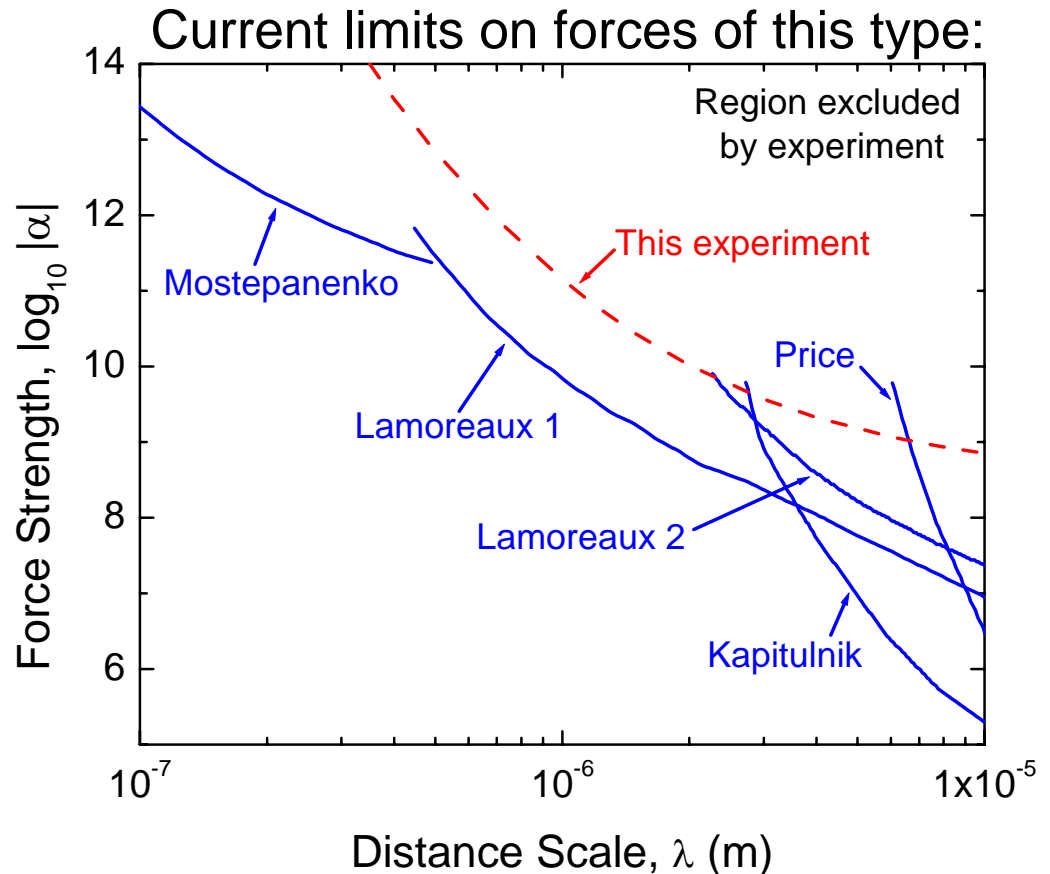
Residuals to the C-P force

From E.A. Cornell
San Feliu Conference, 2005

Limits on exotic forces

$$U = \int_V \frac{Gm\rho dV}{r} (1 + \alpha e^{-r/\lambda})$$

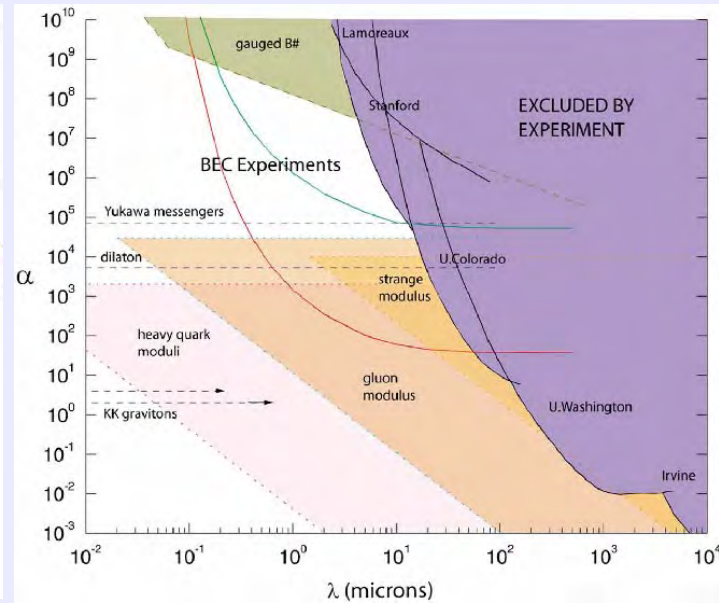
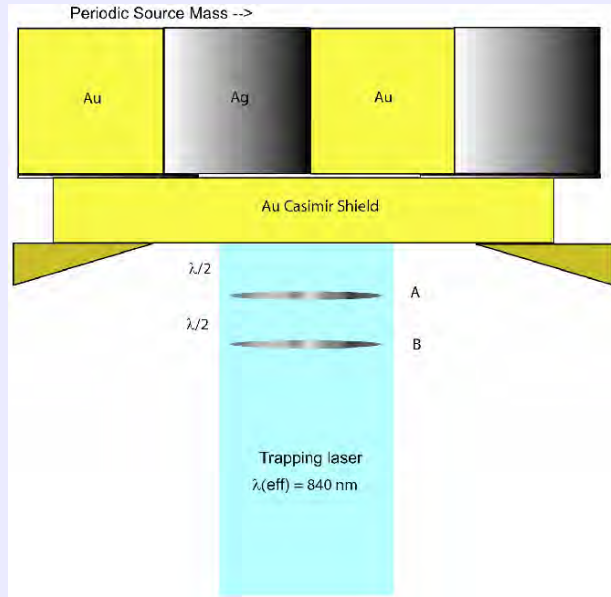
- Very different type of measurement (atom-bulk vs. bulk-bulk)
- Our experiment does not reach the current best limits in 1-10 μm range
- Experimental modifications could improve sensitivity by over an order of magnitude



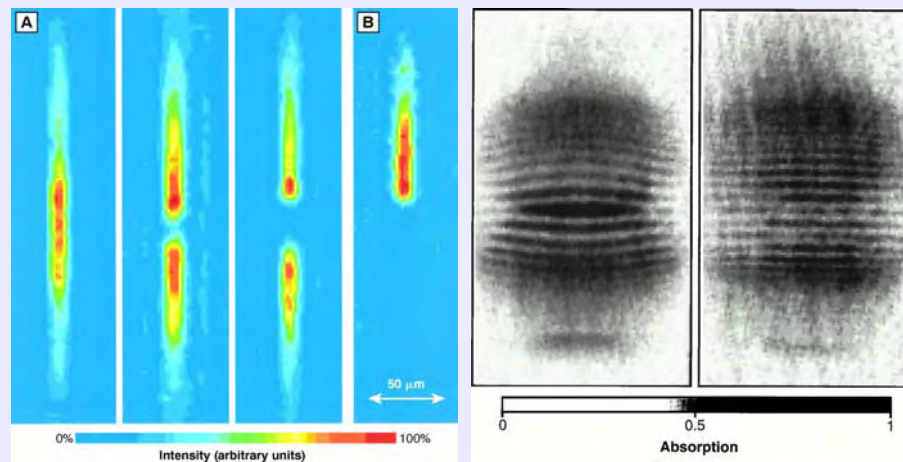
From E.A. Cornell, San Feliu Conference, 2005

D. M. Harber, J. M. Obrecht, J. M. McGuirk, E. A. Cornell, *Measurement of the Casimir-Polder force through center-of-mass oscillations of a Bose-Einstein condensate*, PRA 72, 033610 (2005)

Interferometry with BEC - Proposal



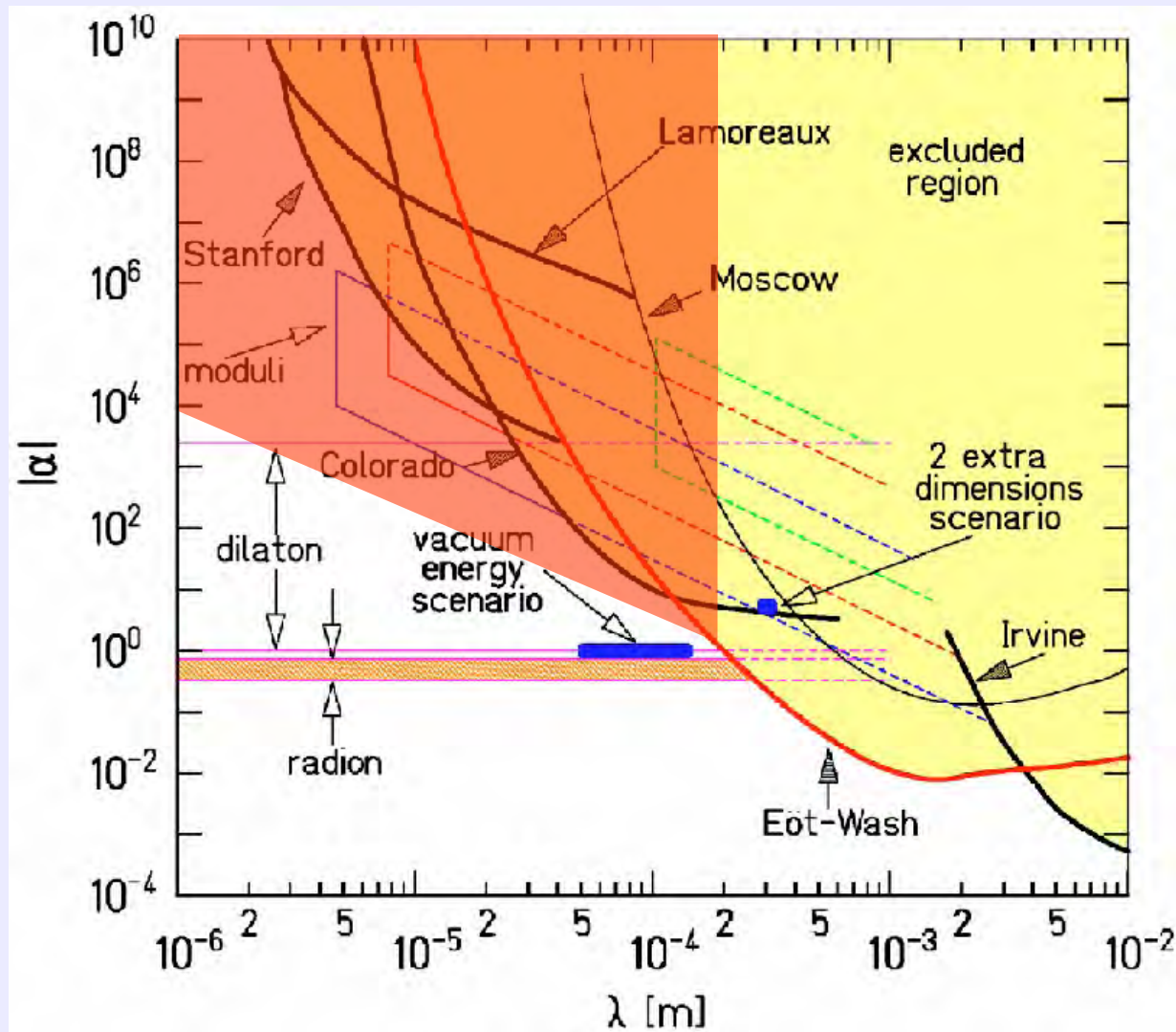
S. Dimopoulos, A. A. Geraci, *Probing submicron forces by interferometry of Bose-Einstein condensed atoms*, PRD 68, 124021 (2003)



M. R. Andrews, C. G. Townsend, H.-J. Miesner, D. S. Durfee, D. M. Kurn, W. Ketterle, *Observation of Interference Between Two Bose Condensates*, Science, 275, 637 (1997)

G. M. Tino, GGI Colloquium, Arcetri 22/6/2006

Accessible region with atomic probes



Conclusions and Prospects

The small size and high sensitivity of atomic sensors enable the investigation of small spatial-scale forces as in atom-surface interactions, surface-induced decoherence, Casimir-Polder interaction, and for the search of recently predicted deviations from the Newtonian gravitational law at micrometer scale










We observed persistent Bloch oscillations of weak-interacting bosonic Sr atoms in a vertical optical lattice for a time longer than 7 s, with more than 8000 photon momenta coherently transferred to the atoms

Longer coherence time and reduced complexity compared to alternative approaches based on degenerate Bose or Fermi gases

The new scheme can be used for a force sensor with micrometric size and sensitivity $\sim 10^{-7}$ g

Gravity can be investigated in new unexplored regions ($\lambda \sim 1-10 \mu\text{m}$, $\alpha \sim 10^3-10^4$) with no need for modeling and extrapolation as in the case of macroscopic probes

Applications of new quantum sensors based on atom interferometry

- Measurement of fundamental constants $\begin{matrix} \nearrow \alpha \\ \searrow G \end{matrix}$  
- New definition of kg 
- Test of equivalence principle 
- Short-distances forces measurement 
- Search for electron-proton charge inequality 
- New detectors for gravitational waves ? 
- Transportable sensors $\begin{matrix} \longrightarrow \text{geophysics} \\ \searrow \text{space} \end{matrix}$  



The Galileo Galilei Institute for Theoretical Physics
Arcetri, Florence
School & Workshop



**SIGRAY Graduate School on
"Experimental Gravitation In Space"**

September 25-27, 2006

Jo Galileo Galilei

The Italian Society of General Relativity and Gravitational Physics (SIGRAY) is organizing an International School on Experimental Gravitation In Space, with the goal of providing young physicists and engineers with an overview of current and future space projects in fundamental physics. The bulk of the lectures will focus on clocks, solar system tests and theoretical foundations. Due attention will be given also to the challenging engineering aspects of fundamental physics missions. The ongoing, large experimental projects (GAIA, GP-B, and LISA) will be presented in dedicated seminars.

The number of participants is limited to 60. Applications including a CV should be submitted before July 15, 2006.

The participation fee is 150 Euro including lunch and social dinner. Registration form and hotel information are available on the web page http://www.fi.infn.it/GGI-gw-space/eng_s.html.

The School will be followed by the International Workshop on ADVANCES IN PRECISION TESTS AND EXPERIMENTAL GRAVITATION IN SPACE organized by ESA, Physics Department University of Firenze, GREC, SIGRAY, INFN and CNES. Information and application forms are available at the web page: http://www.fi.infn.it/GGI-gw-space/eng_w.html.

Information and application forms are available at the web page: http://www.fi.infn.it/GGI-gw-space/eng_w.html.

Directors of the School:
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Prof. Guglielmo M. Tino (Università di Firenze, Firenze, Italy)

Local Organization and Secretariat:
Dr. Andrea Bertoldi
Dr. Francesco Thelli

Programme and Lecturers:

Christophe Salomon (ENS, Paris, France)
Cold atoms, atomic clocks and interferometers

Alberto Anselmi (Alcatel Alenia Space, Italy)
Microsat and spacecraft design

Steve Turishev (NASA/JPL, Pasadena, USA)
Optical Interferometry, LLR and VLBI

Bruno Bertoldi (Università di Pavia, Pavia, Italy)
Solar system tests

Luca Lussana (INFN, Firenze, Italy)
Relativistic effects and non inertial frames

Seminars on GAIA, GP-B and LISA

GGI: <http://www.fi.infn.it/GGI/>



The Galileo Galilei Institute for Theoretical Physics
Arcetri, Florence



**International Workshop on "Advances in Precision
Tests and Experimental Gravitation in Space"**

September 28-30, 2006

Jo Galileo Galilei

In recent years, precision tests of fundamental laws of physics have made remarkable progress, challenging our understanding of the Universe with ingenious and fascinating experiments on both macroscopic objects and quantum systems. Many of these high precision measurements would gain orders of magnitude in sensitivity in space environments, thanks to long free-fall duration, large changes in velocity and gravitational potential, large distance scales, and absence of seismic noise.

The workshop is intended to:

- Present recent results and advances in precision instruments and tests of fundamental laws of physics both on ground and in space
- Discuss how ground-based experiments can be extended into space missions to test our understanding of the Universe
- Present new ideas and proposals for the next generation of fundamental physics "experiments" in space
- Encourage international collaborations between research institutes on topics of common interest.

The number of participants is limited to 100. Applications should be submitted before July 15, 2006. The participation fee is 150 Euro including lunch and social dinner.

Workshop website: http://www.fi.infn.it/GGI-gw-space/eng_w.html.
Before the workshop, a School on "EXPERIMENTAL GRAVITATION IN SPACE" will take place at GGI. Information and application forms are available at the website: http://www.fi.infn.it/GGI-gw-space/eng_s.html.

Organizing Committee: L. Ciavarella (ESTEC, NL), W. Essler (IQ, D), C. Salomon (ENS, F), G. M. Tino (University of Firenze and INFN, I).

Scientific Committee: L. Ciavarella (ESTEC, NL), T. Damour (GHEP, F), H. Dittus (University of Braunschweig, D), W. Essler (IQ, D), P. Gill (NPL, UK), S. Lese (CNES, F), A. Nobili (University of Pisa, I), C. Nury-Main (CNRS-OCA, F), W. Phillips (NIST, USA), S. Reynaud (LKB, F), C. Salomon (LKB, F), S. Schiller (University of Darmstadt, D), G. M. Tino (University of Firenze, I), G. Vescorato (CERN, CH)

Local Organization and Secretariat: Dr. Andrea Bertoldi, Dr. Francesco Thelli

GGI: <http://www.fi.infn.it/GGI/>

People

| | |
|----------------|--|
| G. Lamporesi | PhD student, Università di Firenze |
| A. Giorgini | PhD student, Università di Napoli |
| N. Poli | Post-doc, LENS |
| F. Sorrentino | Post-doc, Università di Firenze |
| A. Bertoldi | Researcher, Università di Firenze |
| G. Ferrari | Researcher, INFN/CNR |
| L. Cacciapuoti | Long term guest, ESA-Noordwijk |
| M. de Angelis | Long term guest, CNR-Napoli |
| R. Drullinger | Long term guest, NIST-Boulder |
| M. Prevedelli | Long term guest, Università di Bologna |
| M. Fattori | ex-PhD student, now in Stuttgart |
| T. Petelski | ex-PhD student, now in Munich |
| J. Stuhler | ex-Post-doc, now in Stuttgart |

Collaborations

- **IEN, Torino**
- **IMGC, Torino**
- **Humboldt-Universitaet zu Berlin**
- **IQO, Hannover**
- **ENS and SYRTE, Paris**

Support and funding

- ✓ **Istituto Nazionale di Fisica Nucleare (INFN)**
- ✓ **European Commission (EC)**
- ✓ **Ministero dell'Istruzione, dell'Università e della Ricerca (MIUR)**
- ✓ **European Laboratory for Non-linear Spectroscopy (LENS)**
- ✓ **Ente Cassa di Risparmio di Firenze (CRF)**
- ✓ **European Space Agency (ESA)**
- ✓ **Agenzia Spaziale Italiana (ASI)**
- ✓ **Istituto Nazionale per la Fisica della Materia (INFN)**
- ✓ **Istituto Nazionale Geofisica e Vulcanologia (INGV)**