

Multi-jet production at NLO with NJet

Valery Yundin

Max-Planck-Institut für Physik

in collaboration with S. Badger, B. Biedermann, A. Guffanti and P. Uwer

HP2, 3–5 August 2014, GGI Firenze

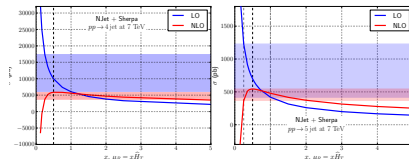
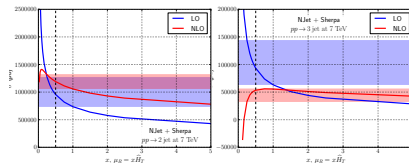
NLO results provide more accurate predictions and theoretical uncertainties for multi-jet backgrounds in new physics searches.

NLO vs LO

- ▶ Reduced theoretical uncertainty

NLO automation

- ▶ Great advances in the recent years
- ▶ High-multiplicity still remains a challenge



— LO
— NLO

Hard process ingredients

$$\sigma^{\text{NLO}} = \int_n \left(\boxed{d\sigma_n^{\text{B}}} + \boxed{d\sigma_n^{\text{V}}} + \int_1 \boxed{d\sigma_{n+1}^{\text{S}}} \right) + \int_{n+1} \left(\boxed{d\sigma_{n+1}^{\text{R}}} - \boxed{d\sigma_{n+1}^{\text{S}}} \right)$$

↑ bottleneck
 ↑ bottleneck

Complicated pieces

1. Virtual matrix elements [NJet, QCDLoop]
 - ▶ Integration over loop momentum
 - ▶ A number of new competing advanced methods

2. Real + subtraction [Sherpa, Comix]
 - ▶ Tree-like
 - ▶ Difficult phase-space integration

3. Linked with BLHA interface

NJet version 2.0¹

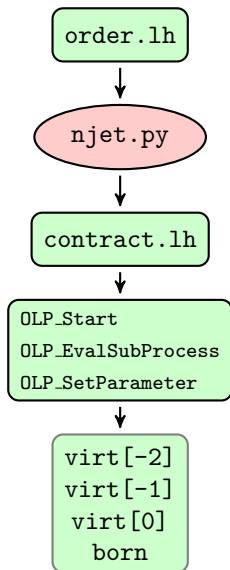
Multi-parton **matrix elements** in massless QCD [arXiv:1209.0100]

- ▶ Full colour-summed amplitudes for up to 5 outgoing partons
- ▶ Reliable accuracy estimate and rescue system
- ▶ BLHA interface for MC generators

New in version 2.0

- ▶ $W^\pm/Z/\gamma$ with up to **5 jets** and $\gamma\gamma$ with up to **4 jets**.
- ▶ Leading/Subleading colour splitting.
- ▶ Hardware vectorization for scaling test.
- ▶ BLHA2 support.
- ▶ Fast analytic amplitudes for 2 and 3 jets.

¹available from project homepage <https://bitbucket.org/njet/njet>



BLHA

- ▶ Simple uniform interface between Monte-Carlo (MC) and One Loop Providers (OLP)
[\[arXiv:1001.1307, arXiv:1308.3462\]](#)

BLHA in NJet 2.0

- ▶ Support BLHA1 and BLHA2
- ▶ Provide colour/spin-correlated trees
- ▶ Provide leading/subleading colour and desymmetrized amplitudes

BLHA extensions

- ▶ Control all settings via order file
- ▶ Single point of interaction with OLP

Advanced methods for computing amplitudes

- ▶ On-shell methods to avoid unphysical degrees of freedom (amplitudes from trees, rational terms from massive loop cuts)
- ▶ Efficient recursive construction of building blocks
- ▶ Relations between primitive amplitudes

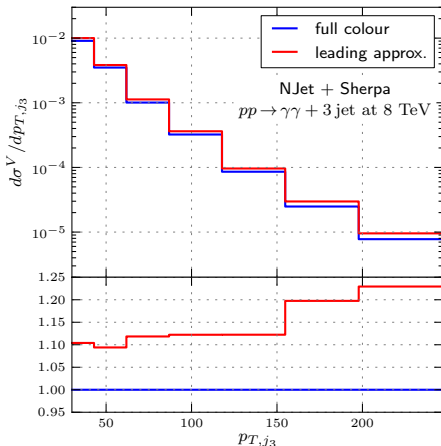
Time per phase-space point for dominating channels

$$T(n) \sim 2^n n^6 \boxed{n!}, \quad n - \text{number of final states}$$

Getting rid of the factorial (offload to MC)

- ▶ Desymmetrizing final states (available in NJet)
- ▶ Separate integration of leading/subleading colour (available in NJet)
- ▶ Colour-dressed approach (available in Sherpa/COMIX)

Why split into leading/subleading colour (at high multiplicity)



Subleading colour

- ▶ Order of magnitude **slower**
- ▶ Order of magnitude **smaller**
- ▶ Often cannot be ignored

Separate integration

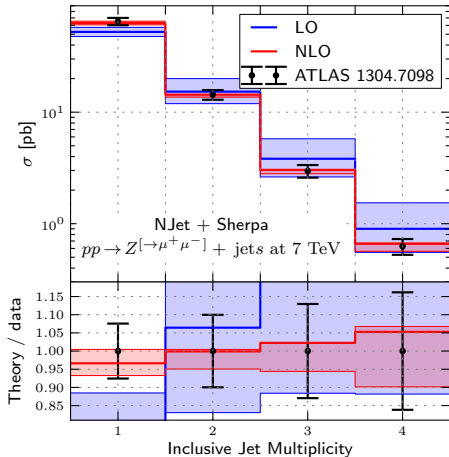
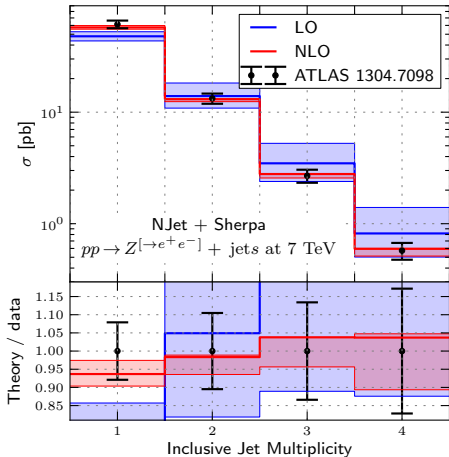
- ▶ Full colour 5–10 times **faster**

Disadvantages

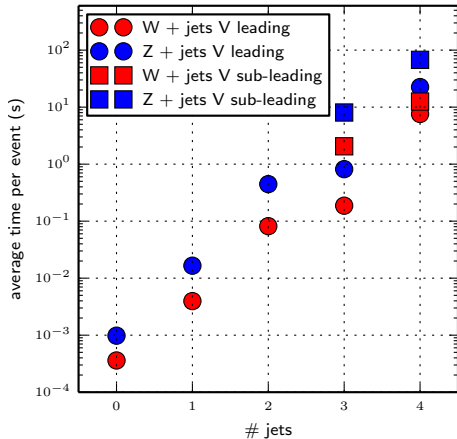
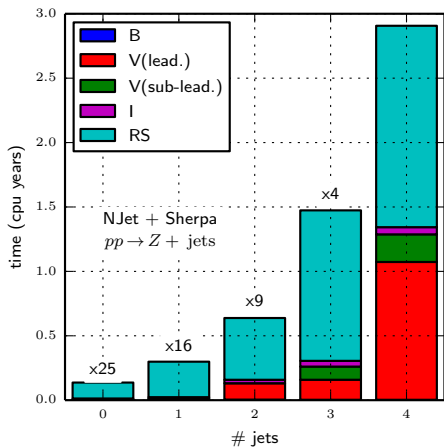
- ▶ Manual (no MC support)
- ▶ μ_R dep. has to be corrected
- ▶ Not standardized in BLHA

Example: Z + jets production with NJet+Sherpa

ATLAS cuts. Agreement with [1304.1253] and [1108.2229]



Time spent in different parts of NLO calculation



High multiplicity calculations are **expensive**

- ▶ Typical 5 final state calculations take $\sim 10^5$ CPU · hours.
- ▶ Do not want to run it more than once.

Layered computation set-up

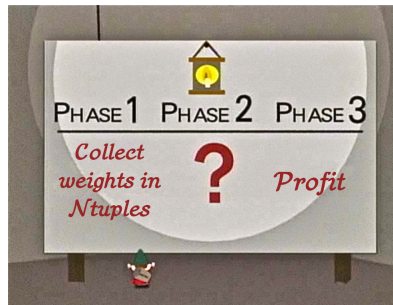
- ▶ Save generated events in ROOT NTuples. [\[arXiv:1003.1241\]](#)
- ▶ Analyze later (still several days per analysis).
- ▶ Interpolation grids with **APPLgrid** for fast PDF convolution and scale variations. [\[arXiv:1312.4460\]](#)

NLO calculations are expensive and we need to use the results as efficiently as possible.

ROOT NTuples output

[arXiv:1003.1241]

- ▶ Save weights, PDFs, scheme dependence
- ▶ Compact compressed storage
- ▶ Can change scales/PDFs during analysis
- ▶ Several jet algorithms at low cost

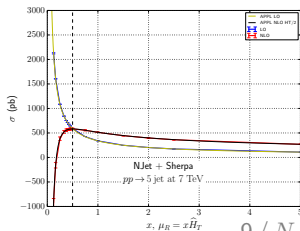
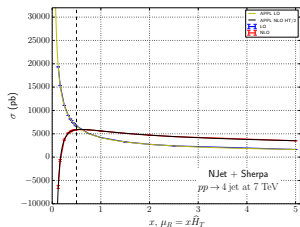
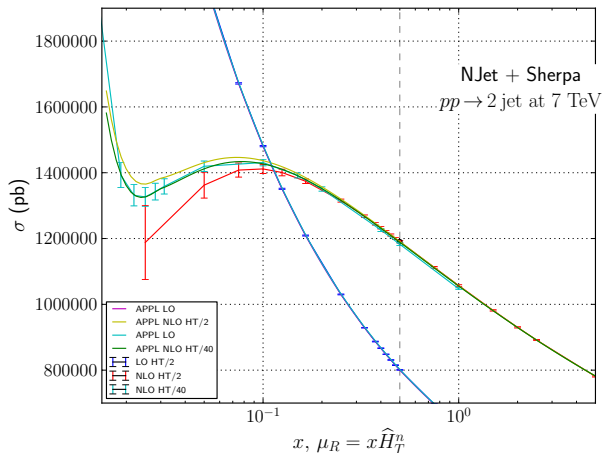


Interpolation grids to speed-up PDF convolution

APPLgrid – interpolation grids in Q , x_1 , x_2 for each bin in histogram.

NTuples: ~ 1000 GB space, ~ 30 hours to analyze, completely generic

APPLgrid: ~ 1 GB space, ~ 0.1 hour to analyze, specific observable/binning



On-the-fly:

- + **Zero** disk space cost
- Scale/PDF vars.
Extremely high CPU cost
- Flexibility **low**
- + Can use standard tools: **Rivet**

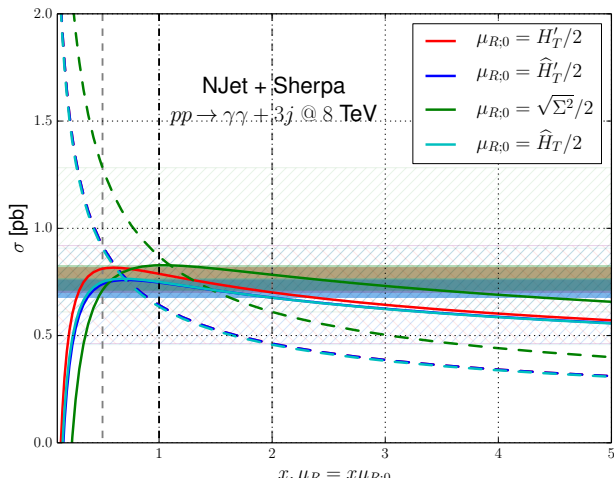
NTuples:

- **High** disk space cost
- ± Scale/PDF vars.
moderate CPU cost
- ++ Flexibility **high**
- Needs custom software

APPLgrid:

- + **Low** disk space cost
- ++ Scale/PDF vars.
low CPU cost
- Flexibility **low**
- Needs custom software

Most important for high-multiplicity fixed order NLO are flexibility and cheap scale/PDF variations: NTuples+APPLgrid



Cuts

$$\begin{aligned}
 p_{T,j} &> 30 \text{ GeV} \\
 |\eta_j| &\leq 4.7 \\
 p_{T,\gamma_1} &> 40 \text{ GeV} \\
 p_{T,\gamma_2} &> 25 \text{ GeV} \\
 |\eta_\gamma| &\leq 2.5 \\
 R_{\gamma,j} &= 0.5 \\
 R_{\gamma,\gamma} &= 0.45
 \end{aligned}$$

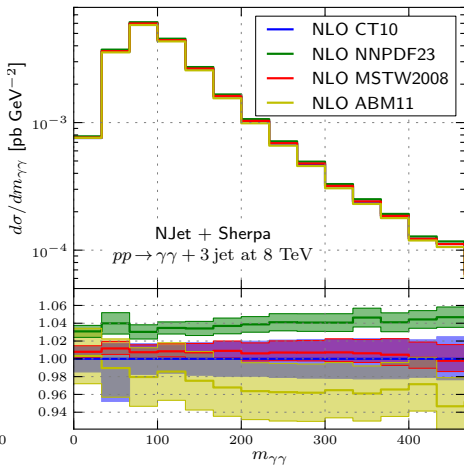
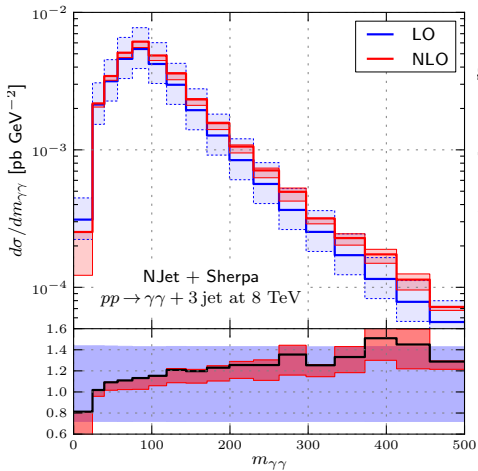
Scales

$$\begin{aligned}
 \hat{H}_T &= \sum_{i \in \{\gamma, \text{partons}\}} p_{T,i} \\
 \hat{H}'_T &= m_{\gamma\gamma} + \sum_{i \in \text{partons}} p_{T,i} \\
 H'_T &= m_{\gamma\gamma} + \sum_{i \in \text{jets}} p_{T,i} \\
 \Sigma^2 &= m_{\gamma\gamma}^2 + \sum_{i \in \text{jets}} p_{T,i}^2
 \end{aligned}$$

$$\sigma_{\gamma\gamma+3j}^{LO}(\hat{H}'_T/2) = 0.643(0.003)_{-0.180}^{+0.278} \text{ pb}$$

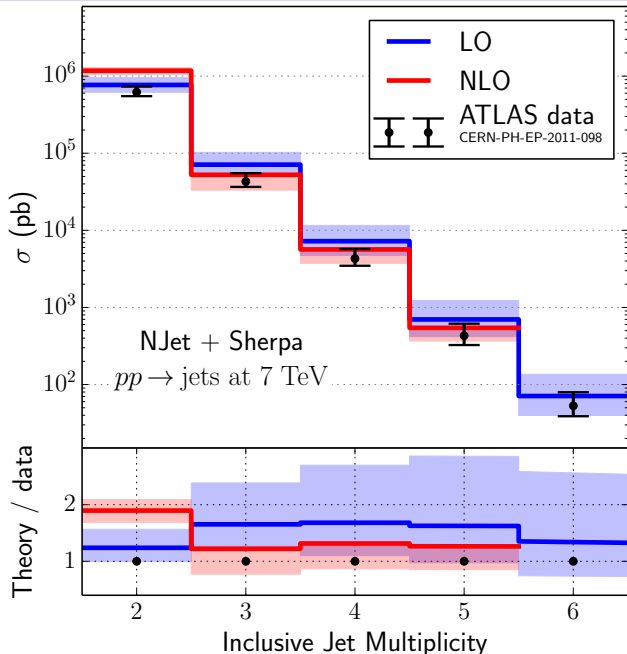
$$\sigma_{\gamma\gamma+3j}^{NLO}(\hat{H}'_T/2) = 0.785(0.010)_{-0.085}^{+0.027} \text{ pb}$$

PDF uncertainty $\approx 3-6\%$



Di-photon invariant mass distribution

NJet+Sherpa: total XS for 2, 3, 4, 5 jets at 7 TeV vs ATLAS measurements



Cuts

anti-kt $R = 0.4$

$p_T^{1st} > 80$ GeV

$p_T^{other} > 60$ GeV

$|\eta| < 2.8$

NLO

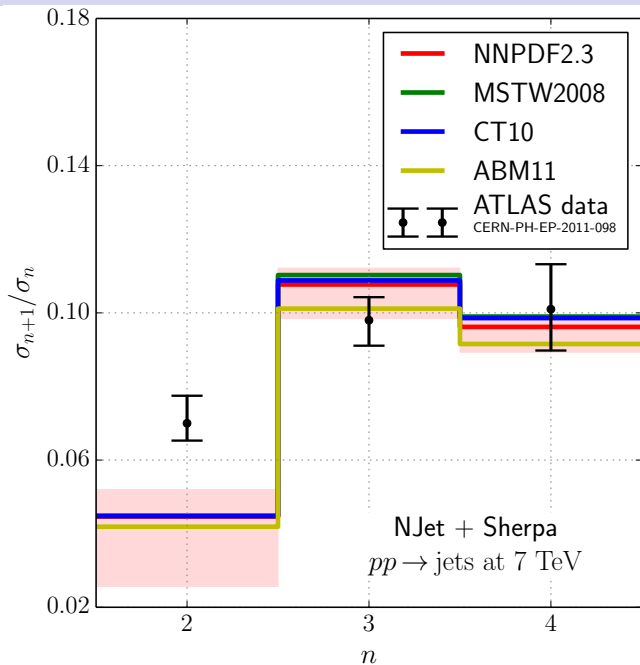
$\mu_R = \mu_F = \hat{H}_T/2$

vars. $\hat{H}_T/4$ and \hat{H}_T

$\alpha_s(M_Z) = 0.118$

NNPDF23 PDF set

NJet+Sherpa: jets ratios at 7 TeV with different PDFs vs ATLAS data



Cuts

anti-kt $R = 0.4$

$p_T^{\text{1st}} > 80 \text{ GeV}$

$p_T^{\text{other}} > 60 \text{ GeV}$

$|\eta| < 2.8$

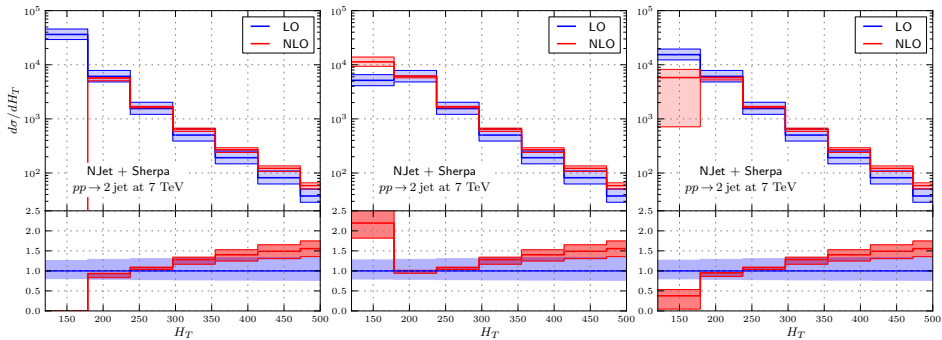
NLO

$\mu_R = \mu_F = \hat{H}_T/2$

vars. $\hat{H}_T/4$ and \hat{H}_T
 (shown for NNPDF)

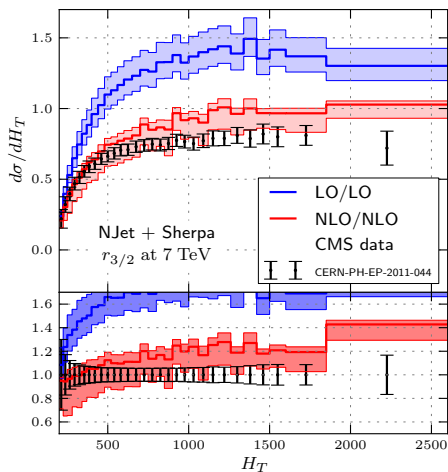
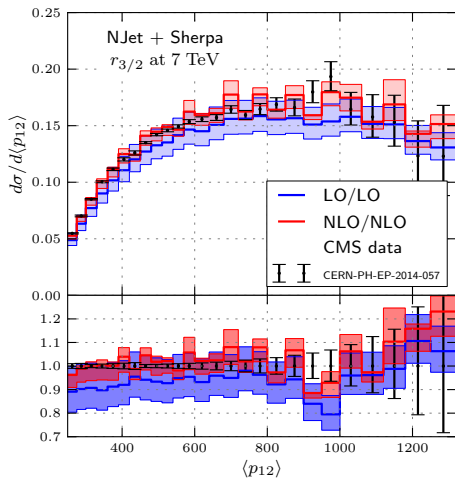
$\alpha_s(M_Z) = 0.118$

Dijet phase-space cuts choices

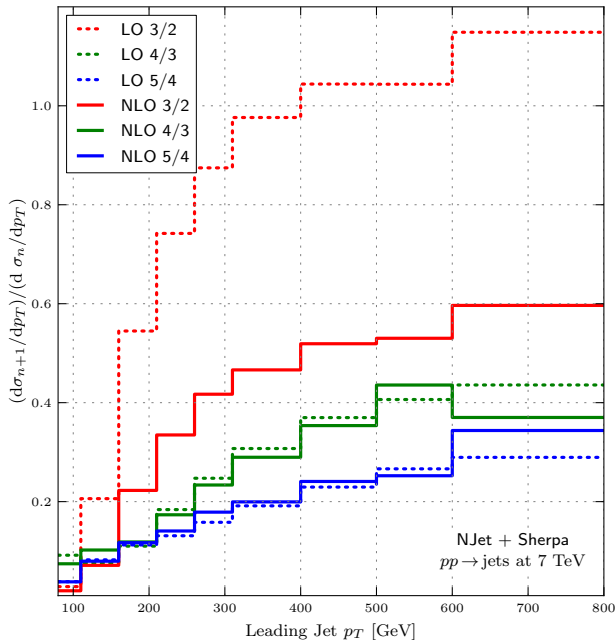


- ▶ **Left** symmetric p_T cuts: $p_{T,1} > 60$ GeV, $p_{T,2} > 60$ GeV
doesn't work for NLO (negative XS)
- ▶ **Middle** asymmetric p_T cuts: $p_{T,1} > 80$ GeV, $p_{T,2} > 60$ GeV
doesn't give full NLO accuracy (LO is symmetric)
- ▶ **Right** symmetric $H_{T,2}$ cuts: $p_{T,1} + p_{T,2} > 140$ GeV
seems to work best

3/2 jet ratios compared with CMS data



$\langle p_{T,12} \rangle = H_{T,2}/2$ results in smaller K -factor than H_T



Cuts

anti-kt $R = 0.4$

$p_T^{\text{1st}} > 80$ GeV

$p_T^{\text{other}} > 60$ GeV

$|\eta| < 2.8$

NLO

$\mu_R = \mu_F = \hat{H}_T/2$

vars. 60%, 12%, 9%
 (not shown)

$\alpha_s(M_Z) = 0.118$

NNPDF23 PDF set

Summary

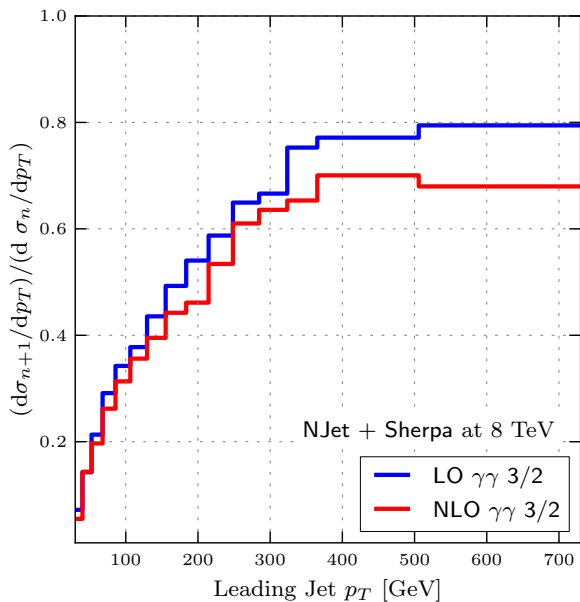
- ▶ NJet library version 2 with improved speed and new processes
- ▶ Multi-leg NLO workflow with NTuples and APPLgrid
- ▶ NJet+Sherpa: $pp \rightarrow 3, 4, 5$ jets at NLO at 7 and 8 TeV
 $pp \rightarrow \gamma\gamma + 2, 3$ jets at NLO at 8 TeV
- ▶ 3/2 ratio is sensitive to cuts and observable definition,
4/3 and 5/4 ratios are more stable

Outlook

- ▶ Today: fully automated 4 final state predictions at NLO
- ▶ In a few years: routine calculations with 5/6 final states

Bonus material

NJet+Sherpa: p_T for $\gamma\gamma$ + jets 3/2 ratio at 8 TeV



Ratio

$$\mu_R = \mu_F = \hat{H}'_T/2$$

$$R_{3/2}^{LO} = 0.314(0.002)$$

$$R_{3/2}^{NLO} = 0.276(0.004)$$

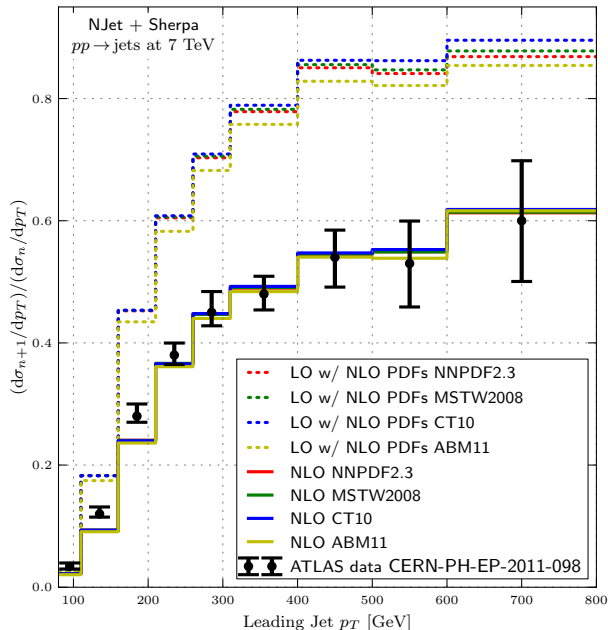
Scale

Different scales

agree within

8% for $R_{3/2}$

NJet+Sherpa: p_T for 3/2 ratio at 7 TeV with diff. PDFs vs ATLAS data



Cuts

anti-kt $R = 0.6$

$p_T^{\text{1st}} > 80$ GeV

$p_T^{\text{other}} > 60$ GeV

$|\eta| < 2.8$

NLO

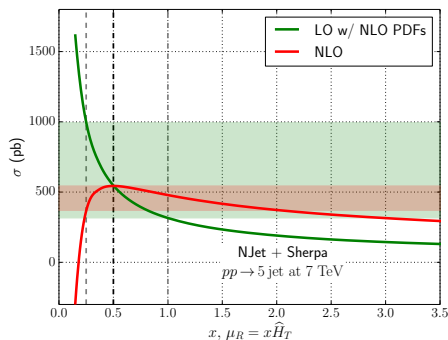
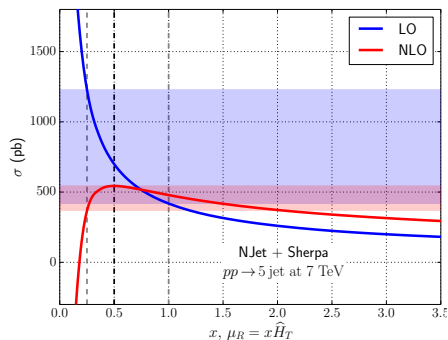
$\mu_R = \mu_F = \hat{H}_T/2$

$\alpha_s(M_Z) = 0.118$

α_s from 3/2 ratios

[ATLAS-CONF-2013-041]

[CMS-QCD-11-003]

ATLAS cuts, NNPDF23 PDF set, $\alpha_s(M_Z) = 0.118$


$$\sigma_5^{7\text{TeV-LO}}(\mu = \hat{H}_T/2) = 0.699(0.004)_{-0.280}^{+0.530} \text{ nb}$$

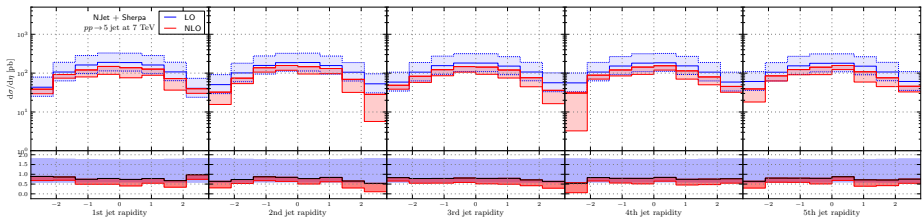
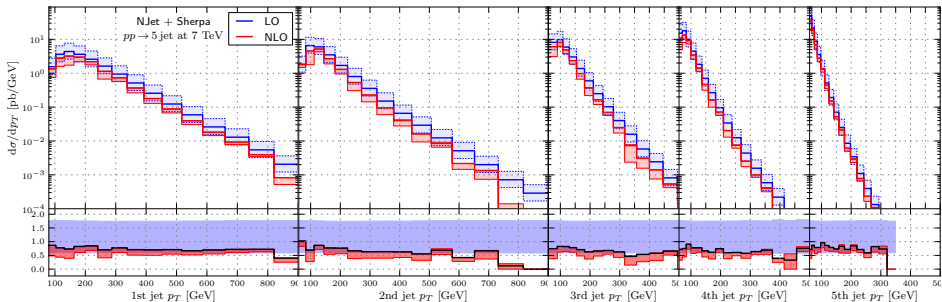
$$\sigma_5^{7\text{TeV-NLO}}(\mu = \hat{H}_T/2) = 0.544(0.016)_{-0.177}^{+0.0} \text{ nb}$$

$$\sigma_5^{8\text{TeV-LO}}(\mu = \hat{H}_T/2) = 1.044(0.006)_{-0.413}^{+0.770} \text{ nb}$$

$$\sigma_5^{8\text{TeV-NLO}}(\mu = \hat{H}_T/2) = 0.790(0.021)_{-0.313}^{+0.0} \text{ nb}$$

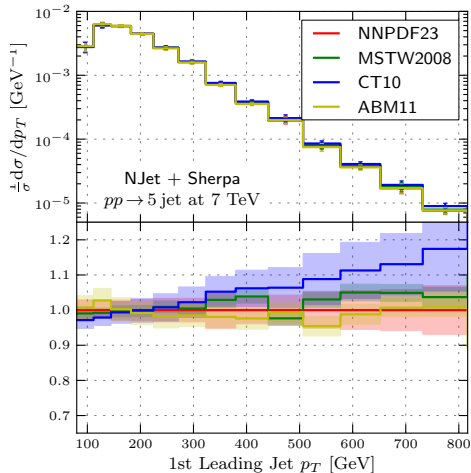
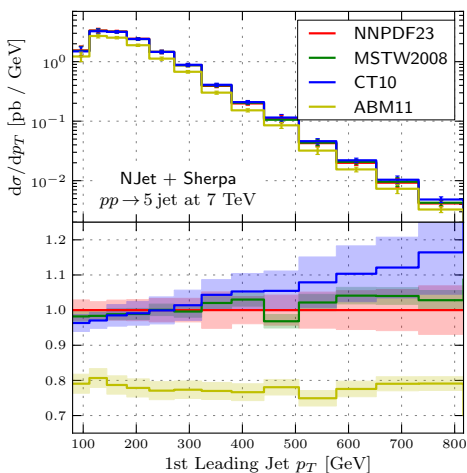
NJet+Sherpa: 5 jets at 7 TeV, p_T and η distributions

ATLAS cuts, NNPDF23 PDF set, $\alpha_s(M_Z) = 0.118$



NJet+Sherpa: 5 jets at 7 TeV, PDF uncertainties

ATLAS cuts, $\alpha_s(M_Z) = 0.118$, PDF uncertainty $\approx 3\%$



Right plot — distributions normalized to total cross-section.