

The out of equilibrium birth of a superfluid

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arXiv:1407.1862 Accepted PRX



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Fundação para a Ciência e a Tecnologia
MINISTÉRIO DA CIÊNCIA, TECNOLOGIA E ENSINO SUPERIOR



MARIE CURIE ACTIONS

EPSRC

Engineering and Physical Sciences
Research Council

Dynamical phase transitions

$$\epsilon(t) = 1 - \frac{T(t)}{T_c} = t/\tau_Q$$

$$\tau_{eq} = \tau_0 |\epsilon|^{-\nu z}$$

$$\xi_{eq} = \xi_0 |\epsilon|^{-\nu}$$

T_c

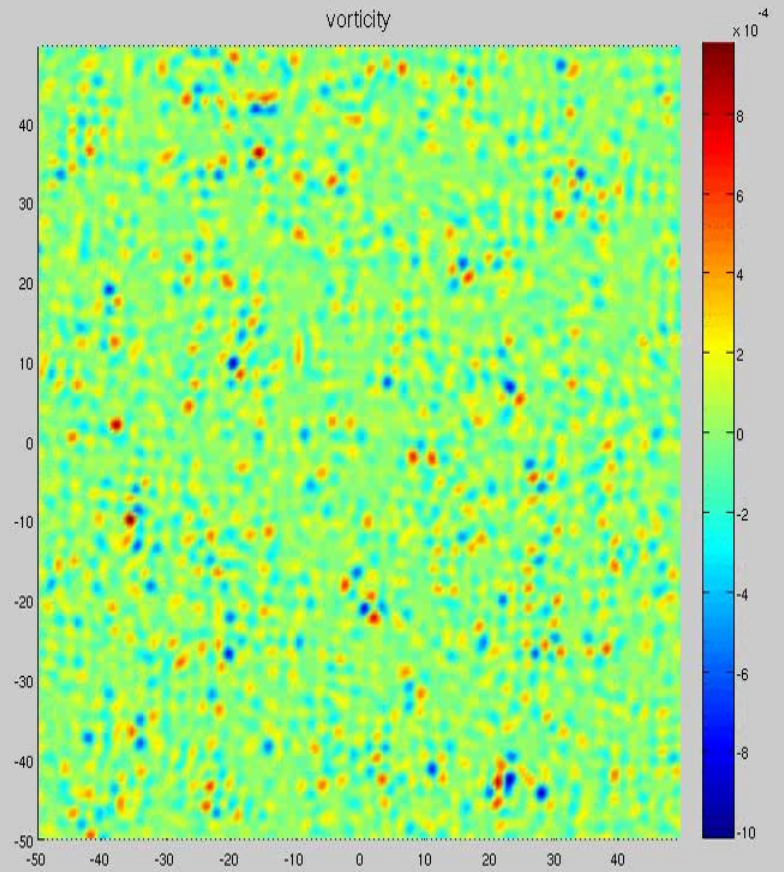
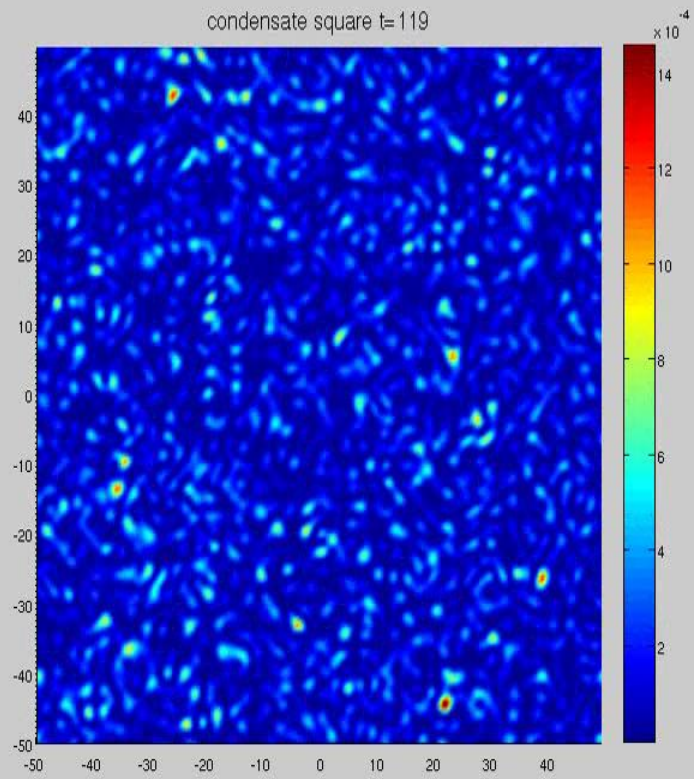
Unbroken Phase

$$\langle \psi \rangle = 0$$

Broken phase

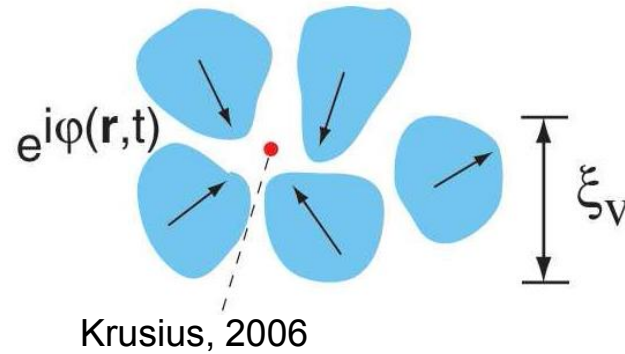
$$\langle \psi \rangle \neq 0$$

$$\langle \psi \rangle = \Delta(x, t) e^{i\theta(x, t)} ? \quad t$$

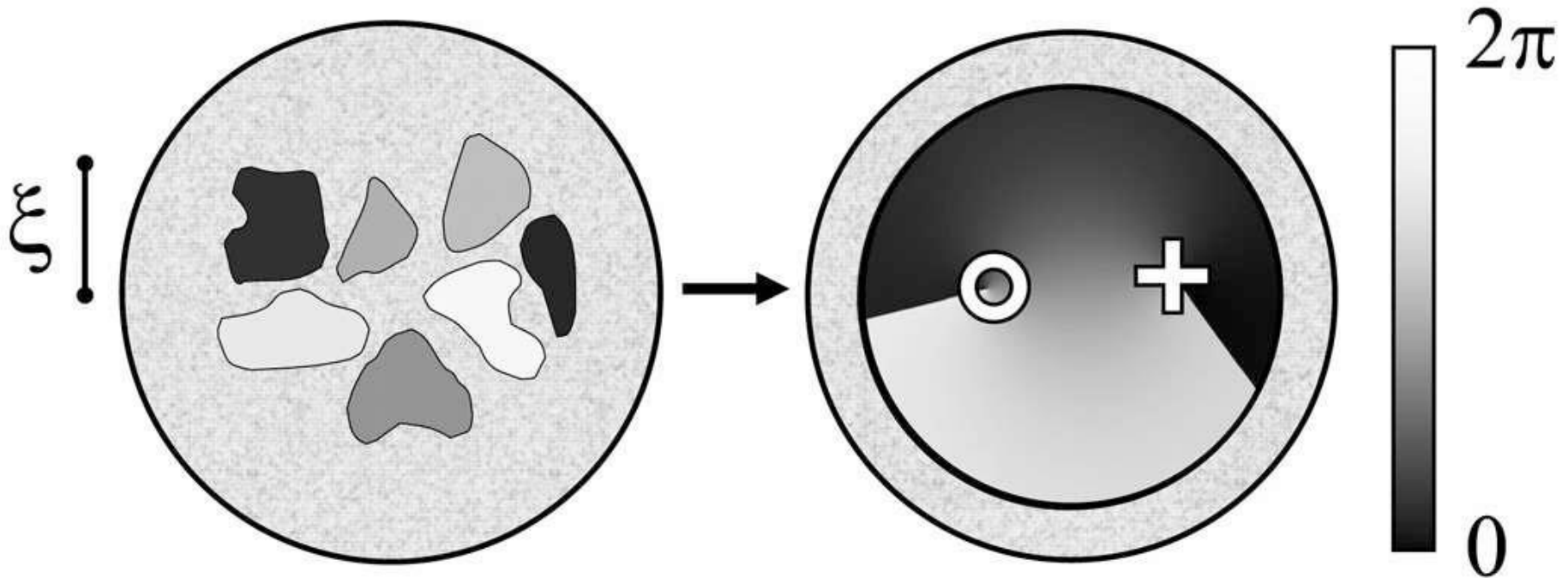


Vortices in the sky

Cosmic strings



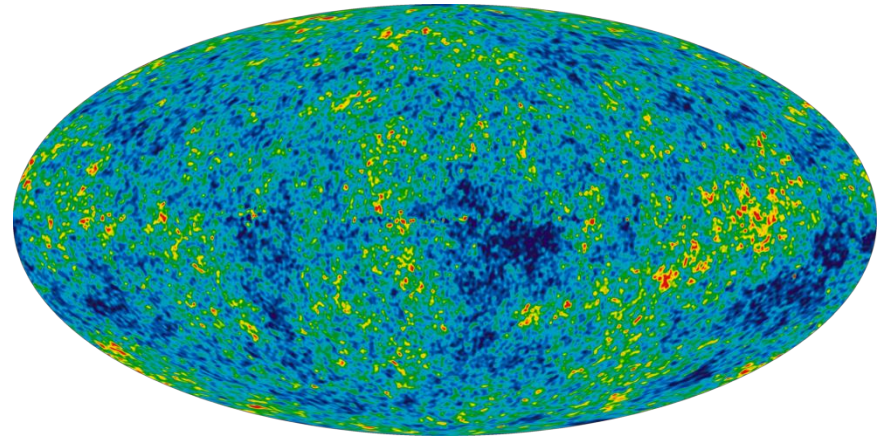
Generation of Structure



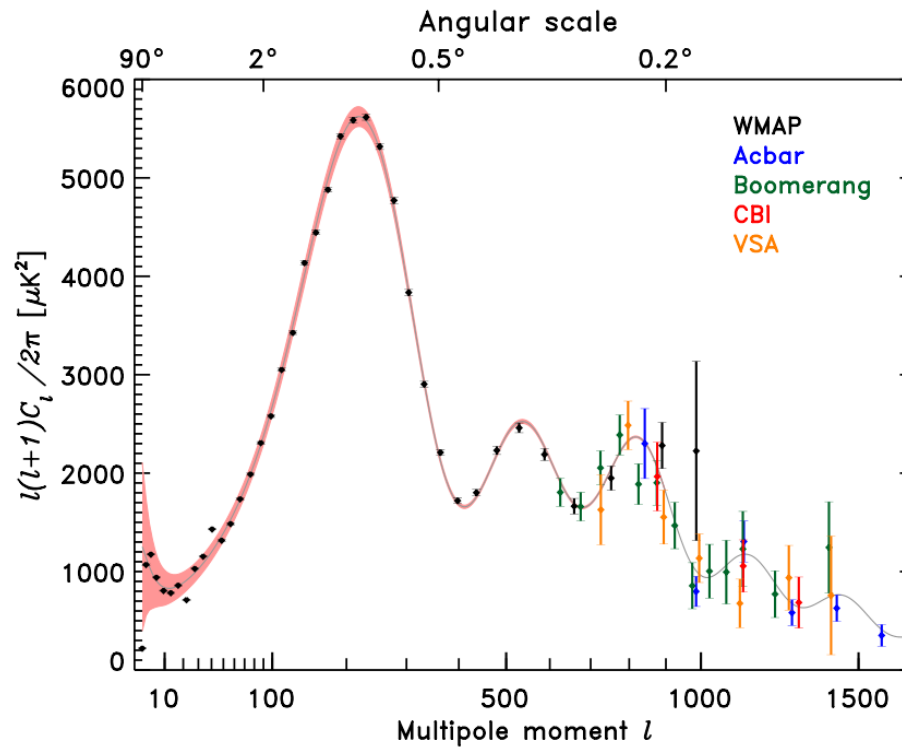
Weyler, Nature 2008

No evidence so far !

CMB, galaxy distributions...



NASA/WMAP



Cosmological experiments in superfluid helium?

Doable for ^4He !!



Zurek

W. H. Zurek

Theoretical Astrophysics, Los Alamos National Laboratory,
Los Alamos, New Mexico 87545, USA Nature 317, 505 (1985)

$$T \approx T_c$$

2nd order



Scaling

$$\tau(T_c) = \infty$$

$$\epsilon(t) = 1 - \frac{T(t)}{T_c} = t/\tau_Q$$

$$t = -\hat{t} \equiv -t_{freeze}$$

$$t = \hat{t} \equiv t_{freeze}$$

Non adiabatic evolution

Defect generation!

$$\epsilon(t) = t/\tau_Q$$

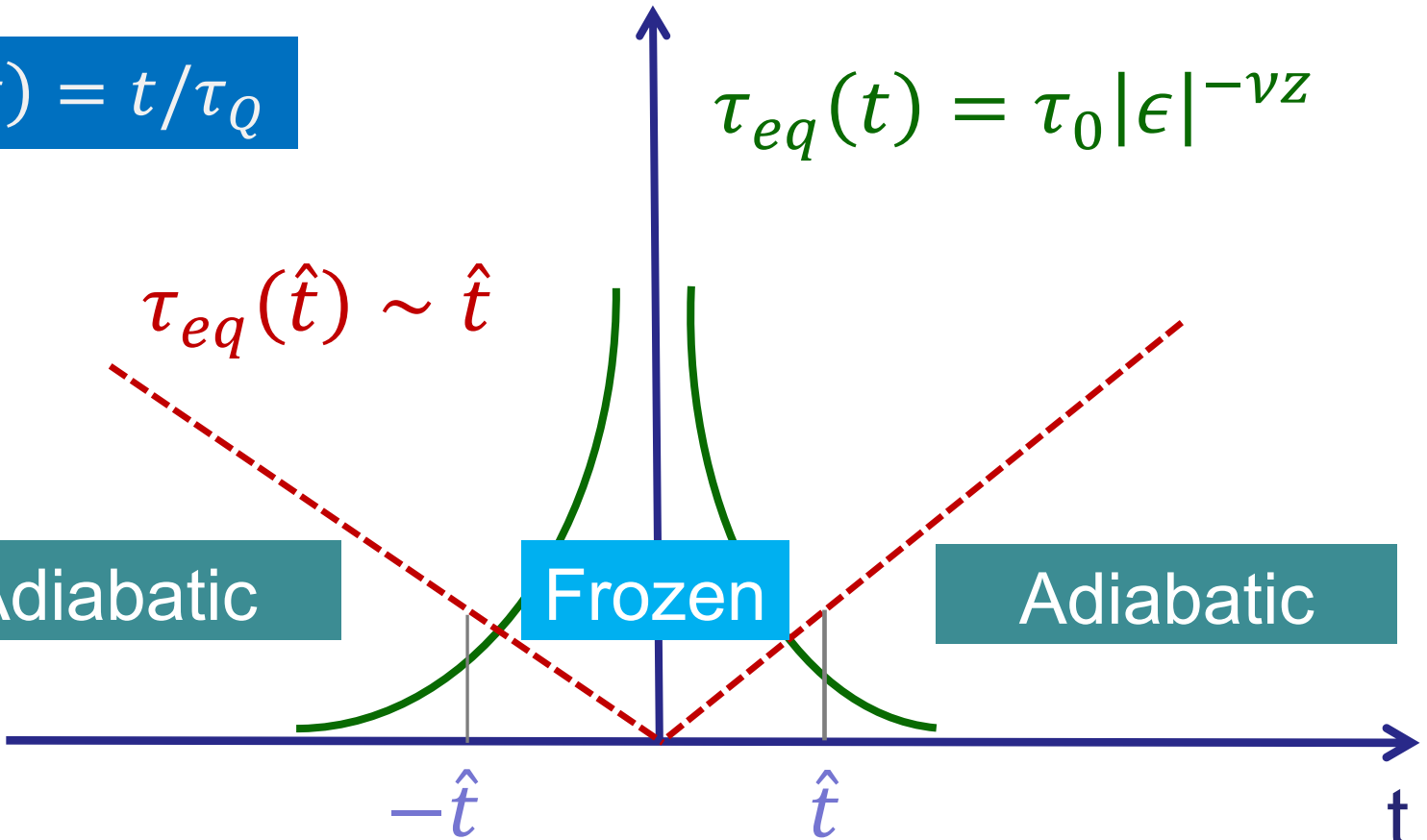
$$\tau_{eq}(t) = \tau_0 |\epsilon|^{-\nu z}$$

$$\tau_{eq}(\hat{t}) \sim \hat{t}$$

Adiabatic

Frozen

Adiabatic



$$\hat{\xi} = \xi_0 |\hat{\epsilon}|^{-\nu} = \xi_0 (\tau_Q/\tau_0)^{\nu/(1+\nu z)}$$

Kibble-Zurek mechanism

$$\rho \sim \hat{\xi}^{-d} \sim \tau_Q^{-d\nu/(1+\nu z)}$$

Generation of defects in superfluid ^4He as an analogue of the formation of cosmic strings

P. C. Hendry*, N. S. Lawson*, R. A. M. Lee*, P. V. E. McClintock* & C. D. H. Williams†

* School of Physics and Materials, Lancaster University, Lancaster LA1 4YB, UK

† Department of Physics, University of Exeter, Exeter EX4 4QL, UK

NATURE · VOL 368 · 24 MARCH 1994

Transient
attenuation of
second sound
amplitude

But vortices
induced by stirring
up!

VOLUME 81, NUMBER 17

PHYSICAL REVIEW LETTERS

26 OCTOBER 1998

Nonappearance of Vortices in Fast Mechanical Expansions of Liquid ^4He through the Lambda Transition

M. E. Dodd,¹ P. C. Hendry,¹ N. S. Lawson,¹ P. V. E. McClintock,¹ and C. D. H. Williams²

¹*Department of Physics, Lancaster University, Lancaster, LA1 4YB, United Kingdom*

²*Department of Physics, University of Exeter, Stocker Road, Exeter EX4 4QL, United Kingdom*

No vortices in ^4He !!

G. Karra, R. J. Rivers, PRL. 81, 3707 (1998)

Vortex formation in neutron-irradiated superfluid ^3He as an analogue of cosmological defect formation

Ruutu, Nature 382, 334-336 (1996)

Laboratory simulation of cosmic string formation in the early Universe using superfluid ^3He

C. Bäuerle et al. Nature 382, 332 (1996)

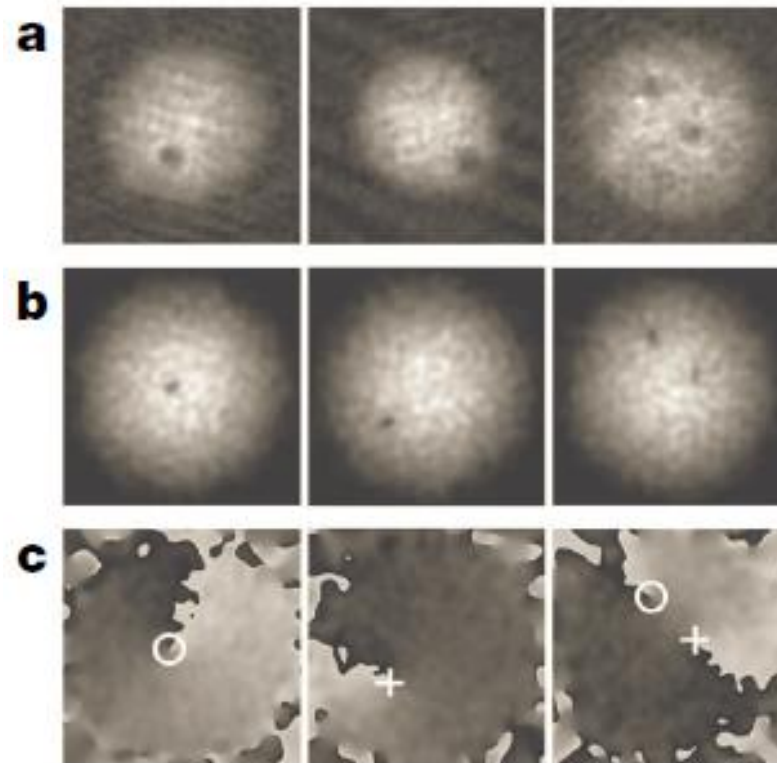
Thin SC films, nematic liquid crystal..



LETTERS

Spontaneous vortices in the formation of Bose–Einstein condensates

Chad N. Weiler¹, Tyler W. Neely¹, David R. Scherer¹, Ashton S. Bradley²†, Matthew J. Davis² & Brian P. Anderson¹



ARTICLE

Received 25 Mar 2013 | Accepted 11 Jul 2013 | Published 7 Aug 2013

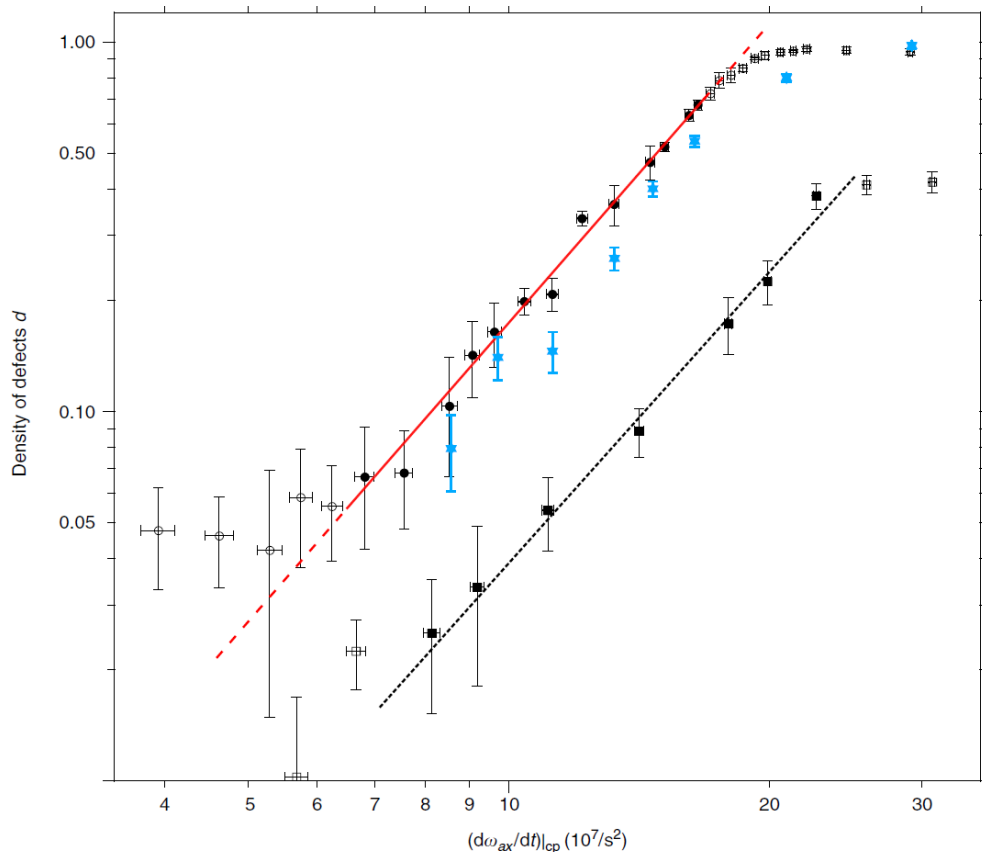
DOI: 10.1038/ncomms3290

Observation of the Kibble–Zurek scaling law for defect formation in ion crystals

S. Ulm¹, J. Roßnagel¹, G. Jacob¹, C. Degünther¹, S.T. Dawkins¹, U.G. Poschinger¹, R. Nigmatullin^{2,3}, A. Retzker⁴, M.B. Plenio^{2,3}, F. Schmidt-Kaler¹ & K. Singer¹

KZ scaling with the
quench speed

Too few defects



Extension to quantum phase transitions

Zurek, Zoller, et al, "Dynamics of a quantum phase transition." , PRL 95.10 (2005): 105701.

Demonstration of KZ scaling in 1d Ising chain in transverse field

Dziarmaga "Dynamics of a quantum phase transition: Exact solution of the quantum Ising model." PRL 95.24 (2005): 245701.

Calculation of correlation functions

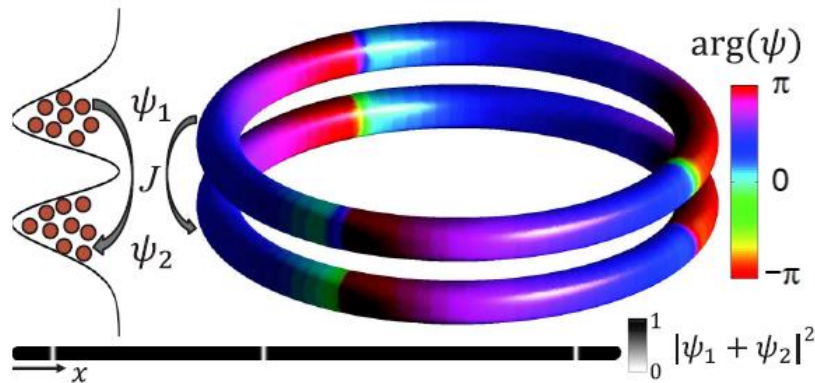
Kibble-Zurek problem: Universality and the scaling limit
PRB 86, 064304, (2012), Gubser, Sondhi et al.

Fast quenches

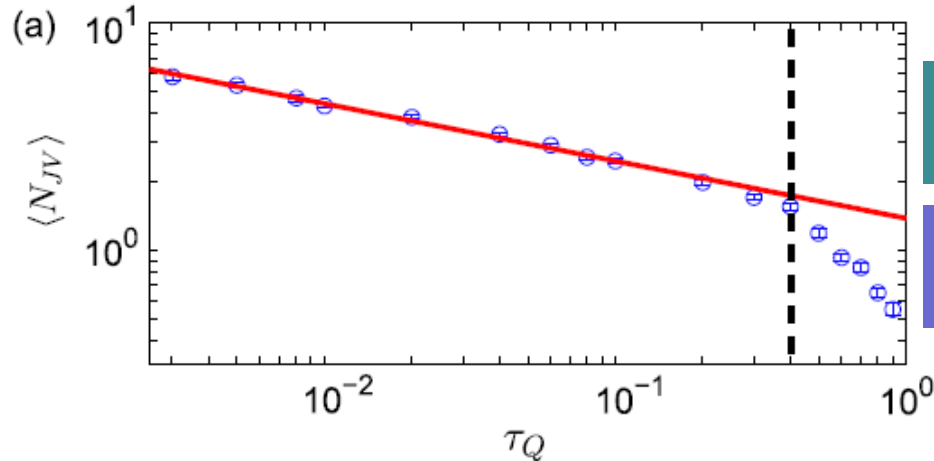
Universal Scaling in Fast Quantum Quenches in Conformal Field Theories, S. R. Das, D. Galante, R. C. Myers PRL 112, 171601 (2015)

Kibble-Zurek Scaling and its Breakdown for Spontaneous Generation of Josephson Vortices in Bose-Einstein Condensates

Shih-Wei Su,¹ Shih-Chuan Gou,² Ashton Bradley,³ Oleksandr Fialko,⁴ and Joachim Brand⁴



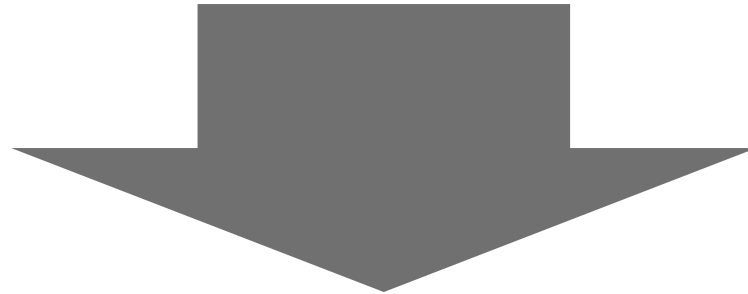
Stochastic
Gross-Pitaevskii



Breaking of KZ scaling

Too few vortices !

Adiabatic at t_{freeze} ?
Defects without a
condensate?



$t_{eq} > t > t_{freeze}$ is relevant

arXiv:1407.1862 PRX Accepted

Slow Quenches

Linear response

$$t > t_{\text{freeze}}$$

Scaling



t_{freeze} t_{eq}

$$\frac{t_{\text{eq}}}{t_{\text{freeze}}} \sim (\log R)^{\frac{1}{1+\nu z}}$$

$$|\psi|^2(t) \propto e^{a_2 \bar{t}^{1+z\nu}}$$

$$\Lambda = (d - z)\nu - 2\beta$$

$$R \sim \xi^{-1} \tau_Q^{\Lambda/1+\nu z}$$

$$\gamma = \frac{1 + (z - 2)\nu}{2(1 + z\nu)}$$

$$\rho(t_{\text{eq}}) \sim [\log R]^\gamma \rho_{\text{KZ}}$$

Non adiabatic growth after t_{freeze}

$$C(t, \mathbf{r}) \equiv \langle \psi^*(t, \mathbf{x} + \mathbf{r}) \psi(t, \mathbf{x}) \rangle$$

$$\psi(t, \mathbf{q}) = \int dt' G_R(t, t', q) \varphi(t, \mathbf{q})$$

$$\langle \varphi^*(t, \mathbf{x}) \varphi(t', \mathbf{x}') \rangle = \zeta \delta(t - t') \delta(\mathbf{x} - \mathbf{x}')$$

$$G_R(t, t', q) = \theta(t - t') H(q) e^{-i \int_{t'}^t dt'' \mathbf{w}_0(\epsilon(t''), q)}$$

$$C(t, q) = \int dt' \zeta |G_R(t, t', q)|^2$$

$$C(t, q) = \int_{t_{freeze}}^t dt' \zeta |H(q)|^2 e^{2 \int_{t'}^t dt'' \text{Im } \mathbf{w}_0(\epsilon(t''), q)} + \dots$$

$$\mathbf{w}_0(\epsilon, q) = \epsilon^{z\nu} h(q\epsilon^{-\nu})$$

$$\text{Im } \mathbf{w}_0 = -a\epsilon^{(z-2)\nu} q^2 + b\epsilon^{z\nu} + \dots, \quad q_{max} \sim \epsilon(t)^\nu$$

$$\text{Im } \mathbf{w}_0 > 0$$

Unstable Modes



Growth

$\langle \psi(t) \rangle$ $t > t_{freeze}$

Protocol

$$\epsilon(t) = t/\tau_Q$$

$$t_i = (1 - T_i/T_c)\tau_Q < 0$$

$$t \in (t_i, t_f)$$

$$t_f = (1 - T_f/T_c)\tau_Q > 0$$

Slow quenches

$$t_f \geq t_{eq}$$

$$t > t_{freeze}$$

Correlation length increases

$$C(t, r) \sim |\psi|^2(t) e^{-\frac{r^2}{2\ell_{co}^2(t)}}, \quad \bar{t} \equiv \frac{t}{t_{freeze}}$$
$$\ell_{co}(\bar{t}) = a_3 \zeta_{freeze} \bar{t}^{\frac{1+(z-2)\nu}{2}}$$

Condensate growth

$$|\psi|^2(t) \sim \tilde{\varepsilon}(t) e^{a_2 \bar{t}^{1+z\nu}}$$
$$\tilde{\varepsilon}(t) \equiv \zeta t_{freeze} \ell_{co}^{-d}(t)$$

Adiabatic evolution
 $t = t_{eq} \gg t_{freeze}$

$$|\psi|^2(t = t_{eq}) \sim |\psi|_{eq}^2(\epsilon(t_{eq}))$$

Defects

$$\rho(t_{eq}) \sim 1/\ell_{co}^{d-D}(t_{eq}) \sim [\log(\zeta^{-1} \tau_Q^\Lambda)]^{-\frac{(d-D)(1+(z-2)\nu)}{2(1+z\nu)}} \rho_{KZ}$$

Fast quenches

$$t_f \ll t_{eq}$$

$$q_{max}(T_f) = \epsilon(t_f)^{d\nu}$$

Breaking of τ_Q scaling

$$KZ \quad t_f < t_{freeze}$$

$$US \quad t_{freeze} \ll t_f \ll t_{eq}$$

Exponential
growth

$$|\psi|^2(t) \sim \epsilon_f^{(d-z)\nu} \zeta \exp [2b(t - t_{freeze})\epsilon_f^{\nu z}]$$

Number of
defects

Independent
of τ_Q

$$\rho \sim \begin{cases} \epsilon_f^{(d-D)\nu} & R_f \lesssim O(1) \\ \epsilon_f^{(d-D)\nu} \log^{-\frac{d-D}{2}} R_f & R_f \gg 1 \end{cases}$$

$$R_f \equiv \frac{\epsilon_f^{2\beta}}{\zeta \epsilon_f^{(d-z)\nu}} \quad \epsilon_f \equiv \frac{T_c - T_f}{T_c}$$

Holography?

Defects survive
large N limit

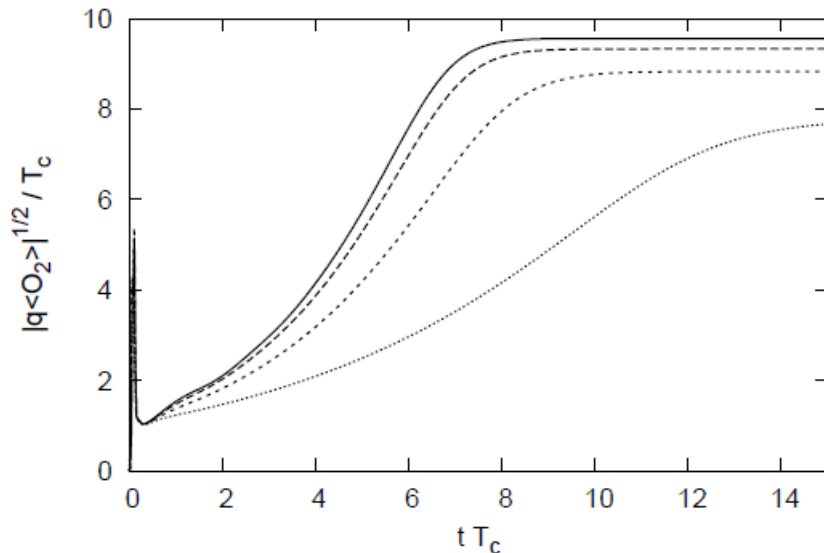
Universality

Real time

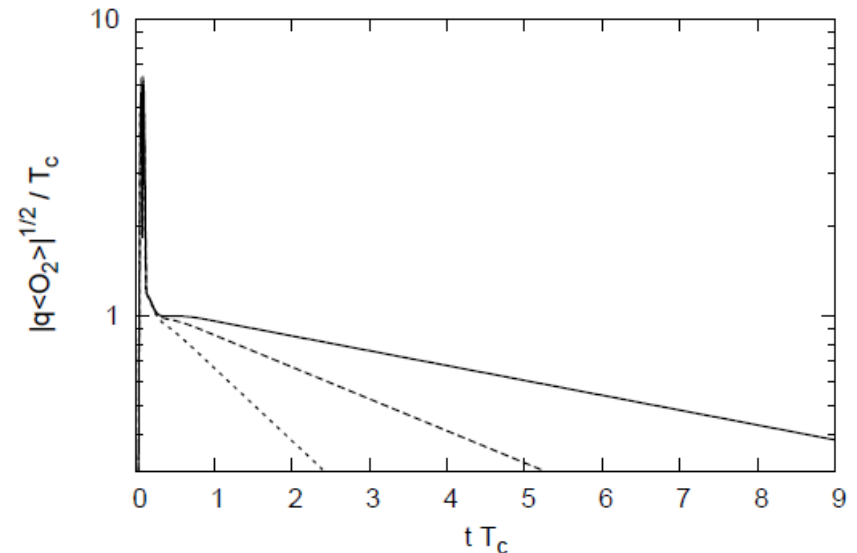
Holographic SC out of equilibrium

$$S_{\text{grav}} = \frac{1}{16\pi G_{\text{Newton}}} \int d^4x \sqrt{-G} \left[R + \Lambda + \frac{1}{e^2} \left(-\frac{1}{4} F_{MN} F^{MN} - |D\Phi|^2 - m^2 |\Phi|^2 \right) \right]$$

Gubser, Herzog, Horowitz, Hartnoll



(a) $T < T_c$



(b) $T > T_c$

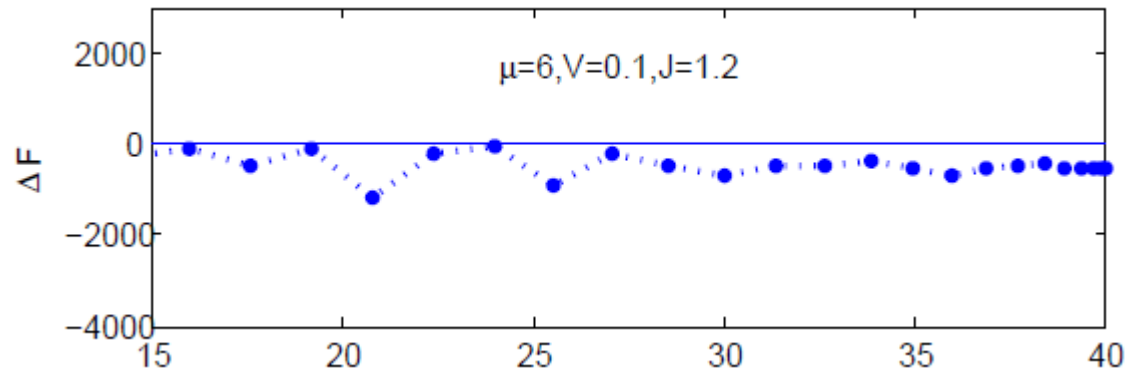
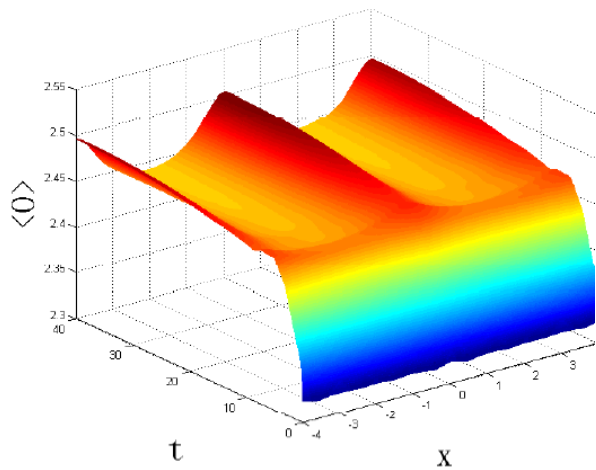
$$|\langle \mathcal{O}_2(t) \rangle| = C_1 \exp(-t/t_{\text{relax}}) + C_2$$

$$|\langle \mathcal{O}_2(t) \rangle| = C \exp(-t/t_{\text{relax}})$$

Exponential growth/decay

Murata, et al.,
arXiv:1005.0633

Oscillations in space

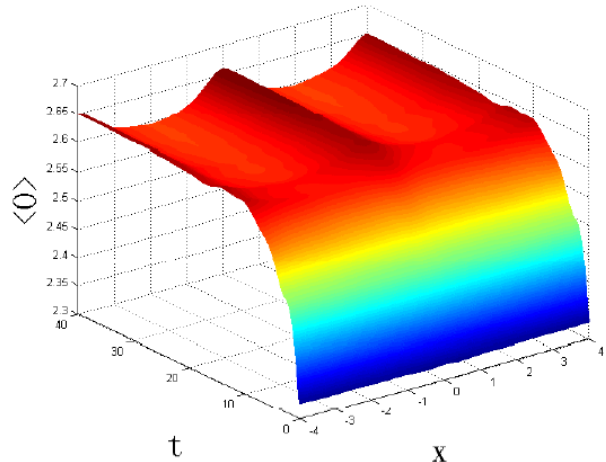


$$\Psi \approx z\psi_1 + \psi(x, t)z^2$$
$$\psi_1(t) = J(1 - \tanh vt)$$

$$\langle O \rangle \sim \psi(x, t)$$

Probe limit

Conservation laws!



Basu et al., arXiv:1308.4061

AGG, Zhang, Bi, arXiv:1308.5398

Oscillations in space: BdG

$$\hat{\xi} = -\vec{\nabla}^2/2m - \mu$$

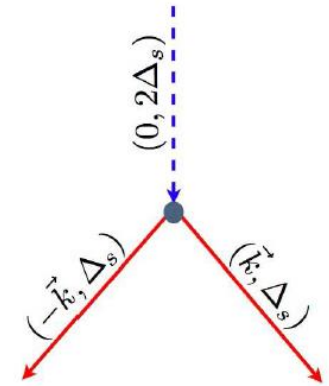
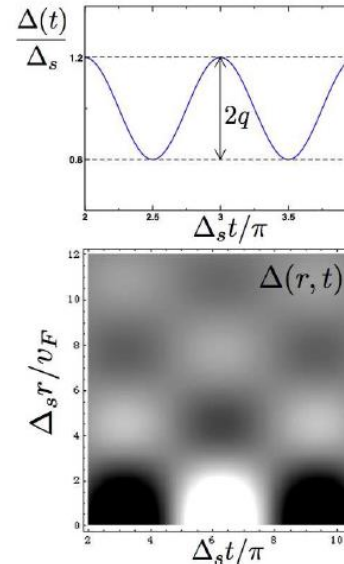
$$i\dot{u}_{\mathbf{p}}(\mathbf{r}, t) = \hat{\xi}u_{\mathbf{p}}(\mathbf{r}, t) + \Delta(\mathbf{r}, t)v_{\mathbf{p}}(\mathbf{r}, t),$$

$$i\dot{v}_{\mathbf{p}}(\mathbf{r}, t) = -\hat{\xi}v_{\mathbf{p}}(\mathbf{r}, t) + \bar{\Delta}(\mathbf{r}, t)u_{\mathbf{p}}(\mathbf{r}, t)$$

$$\Delta(\mathbf{r}, t) = \bar{\Delta}(t) + \delta\Delta(\mathbf{r}, t)$$

$$\delta\Delta(\vec{r}, t) \approx \frac{C e^{\nu_m t} \cos[\Delta_s(t - \tau)] \sin(k_m R) e^{-R^2/l^2(t)}}{\sqrt{\Delta_s t} k_m R}$$

$$l(t) \approx \xi \sqrt{\Delta_s t} \quad \nu_m \approx 2q\Delta_s$$



Conservation laws

Instability to spatial inhomogeneity

KZ scaling in holographic superconductors

Quantum Quench Across a Zero Temperature Holographic Superfluid Transition,

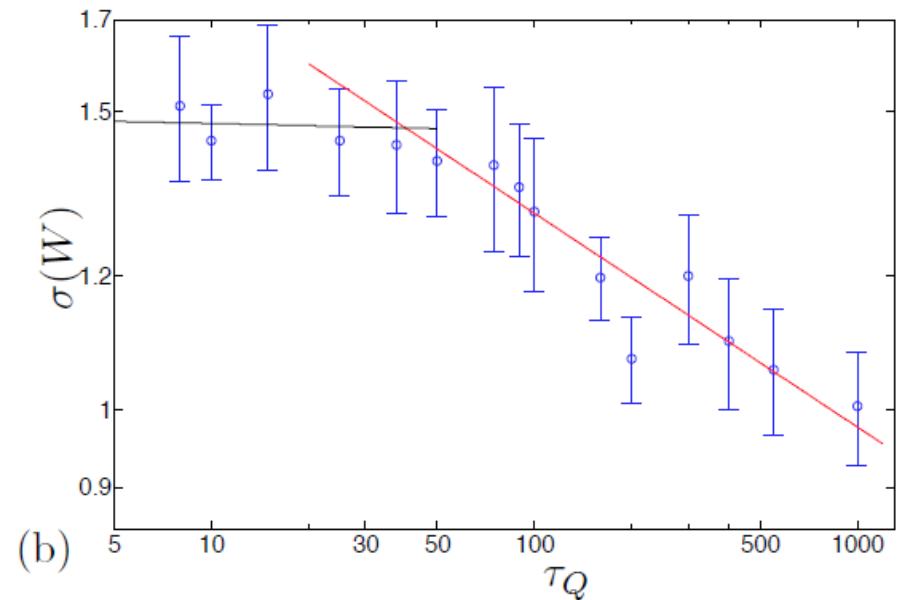
P. Basu, D. Das, S. R. Das, T. Nishioka
arXiv:1211.7076

Kibble-Zurek Scaling in Holographic Quantum Quench : Backreaction

S. R. Das, T. Morita
arXiv:1409.7361

Universal far-from equilibrium dynamics of a holographic superconductor

Sonner, Campo, and Zurek
arXiv:1406.2329



Dual gravity theory

$$S_{\text{grav}} = \frac{1}{16\pi G_{\text{Newton}}} \int d^4x \sqrt{-G} \left[R + \Lambda + \frac{1}{e^2} \left(-\frac{1}{4} F_{MN} F^{MN} - |D\Phi|^2 - m^2 |\Phi|^2 \right) \right]$$

$$\Lambda = -3 \quad m^2 = -2$$

Herzog, Horowitz, Hartnoll, Gubser

*AdS*₄

$$ds^2 = r^2 g_{\mu\nu}(t, \mathbf{x}, r) dx^\mu dx^\nu + 2dr dt$$

Eddington-Finkelstein
coordinates

$$0 = \nabla_M F^{NM} - J^M,$$

$$0 = (-D^2 + m^2)\Phi.$$

Probe limit

EOM's:

PDE's in x, y, r, t

Boundary conditions:

$$r \rightarrow \infty$$

$$\Psi = \frac{\psi^{(1)}(x, y, t)}{r} + \frac{\psi^{(2)}(x, y, t)}{r^2} + \dots$$

$$A_t = \mu - \rho/r$$

hep-th/9905104v2

arXiv:1309.1439

Science 2013

Drive:

$$\epsilon(t) = t/\tau_Q \quad t_i = (1 - T_i/T_c)\tau_Q$$

$$t \in (t_i, t_f) \quad t_f = (1 - T_f/T_c)\tau_Q$$

Dictionary:

$$\langle O_2 \rangle \sim \psi_2$$

No solution of Einstein equations but do not worry, Hubeny 2008

Stochastic driving

$$\psi^{(1)} = \varphi(t, x)$$

$$\langle \varphi^*(t, x) \varphi(t', x') \rangle = \xi \delta(t - t') \delta(x - x')$$

Field theory:

$$\xi(T, \nu)$$

Quantum/thermal fluctuations

Gravity:

$$\xi \propto 1/N^2$$

Hawking radiation

Predictions

Mean field critical exponents

Slow quenches:

$$C(t, r) \sim |\psi|^2(t) e^{-\frac{r^2}{2\ell_{\text{co}}^2(t)}}, \quad |\psi|^2(t) \sim \tilde{\epsilon} t_{\text{freeze}} \bar{t} e^{a_2 \bar{t}^2}, \quad \ell_{\text{co}}(t) \sim \xi_{\text{freeze}} \sqrt{\bar{t}}$$

$$\frac{t_{\text{eq}}}{t_{\text{freeze}}} \sim \sqrt{\log \frac{N^2}{\sqrt{\tau_Q}}}$$

$$\rho \sim \frac{1}{\sqrt{\log \frac{N^2}{\sqrt{\tau_Q}}}} \rho_{\text{KZ}}$$

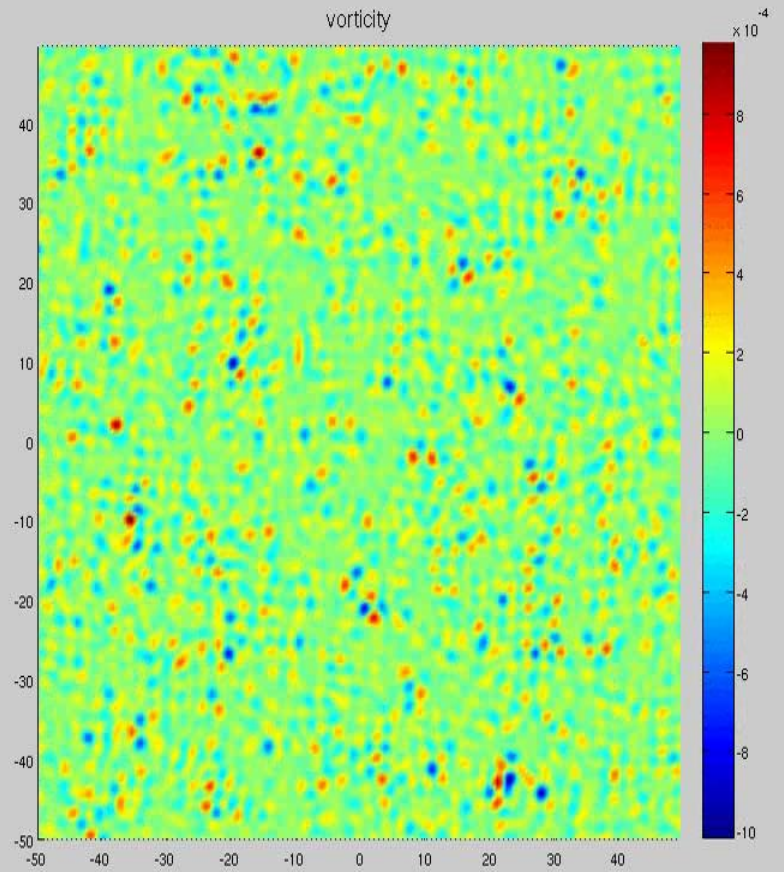
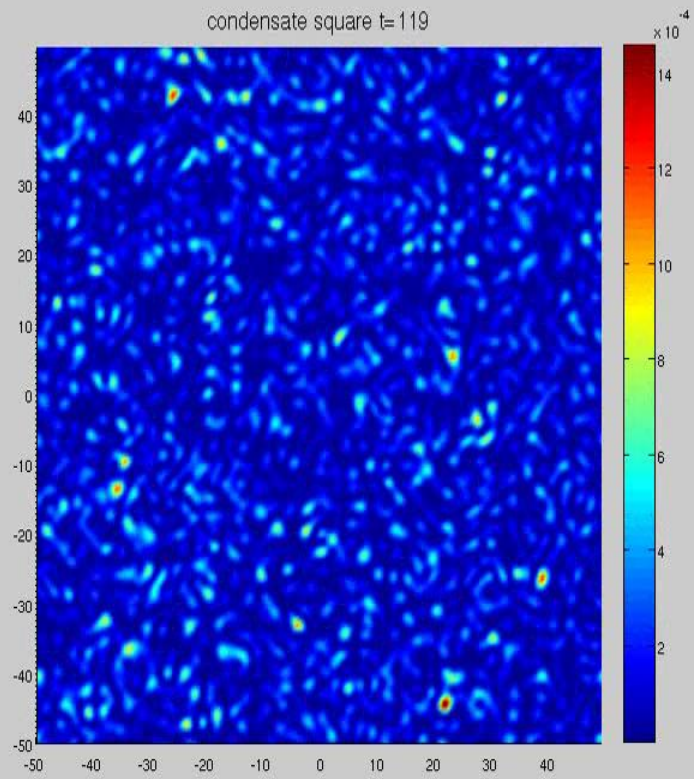
Fast quenches:

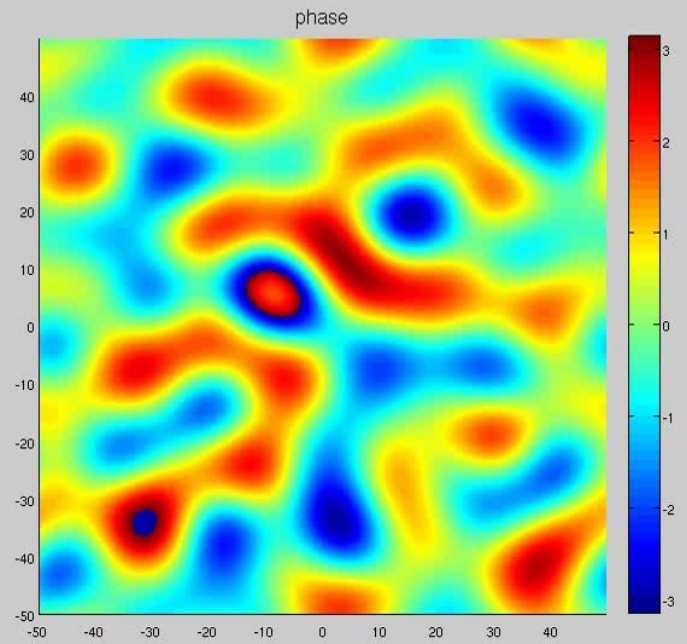
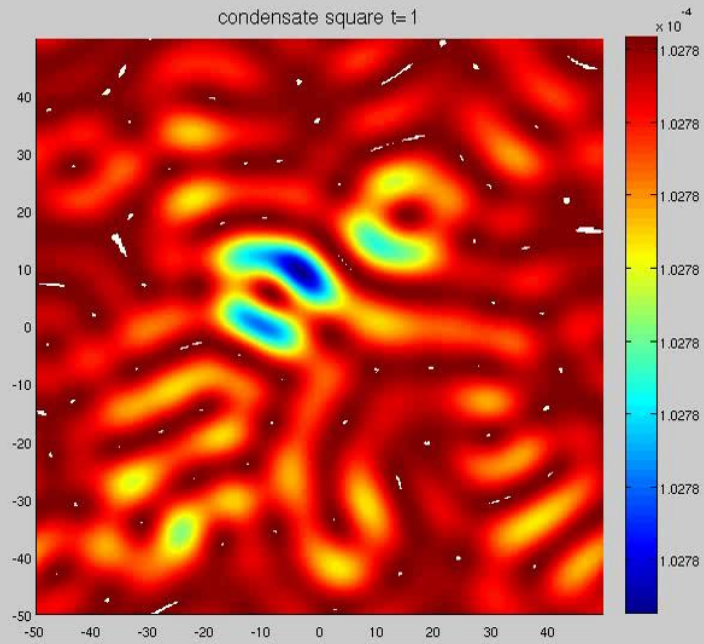
$$C(t, r) = |\psi|^2(t) e^{-\frac{r^2}{2\ell_{\text{co}}^2(t)}}, \quad |\psi|^2(t) \sim \zeta \exp [2b(t - t_{\text{freeze}})\epsilon_f]$$

$$\ell_{\text{co}}^2(t) = 4a(t - t_{\text{freeze}})$$

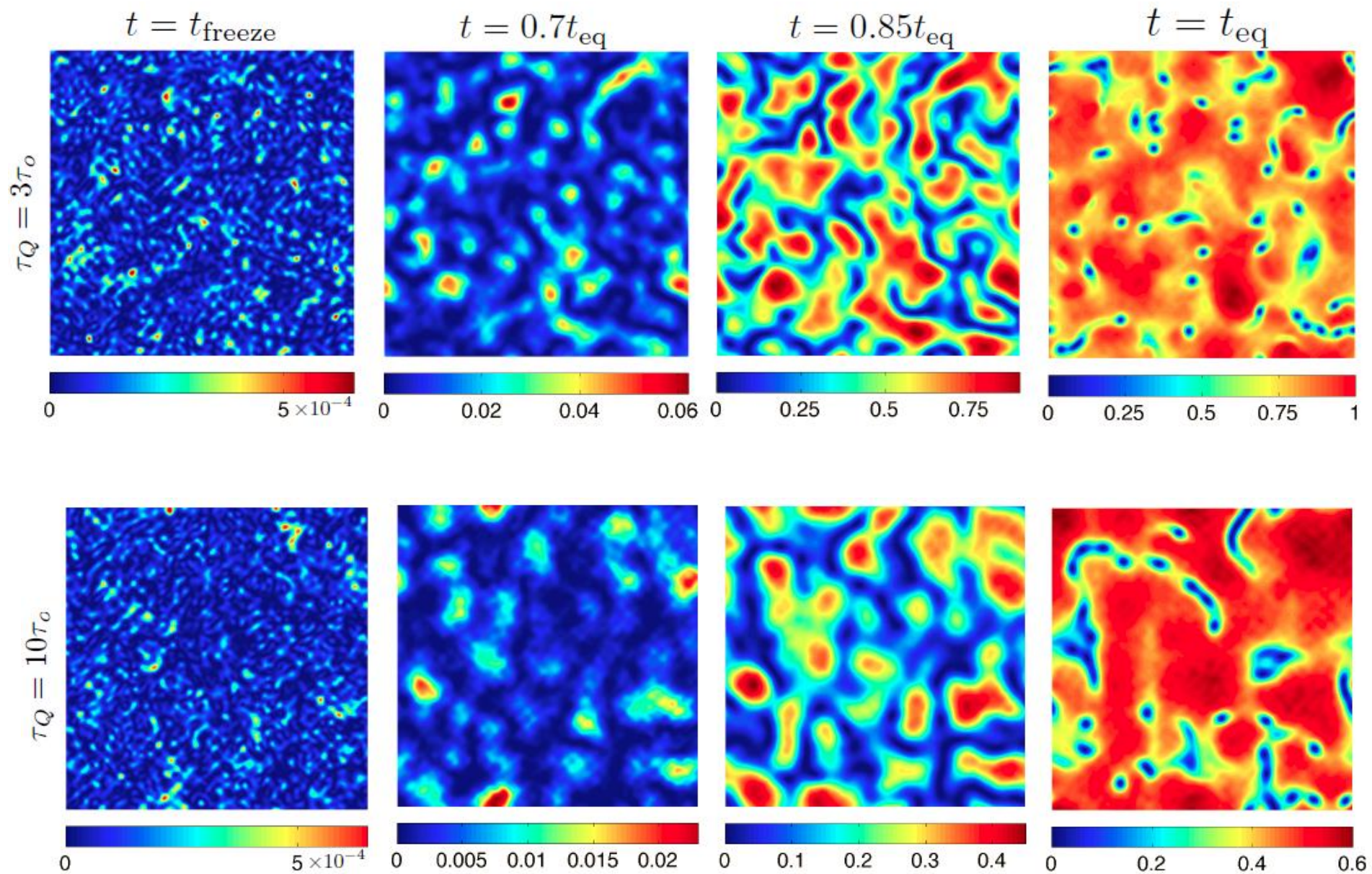
$$\rho \sim \frac{\epsilon_f}{\log \frac{N^2}{\epsilon_f}}$$

Movies!!





Slow

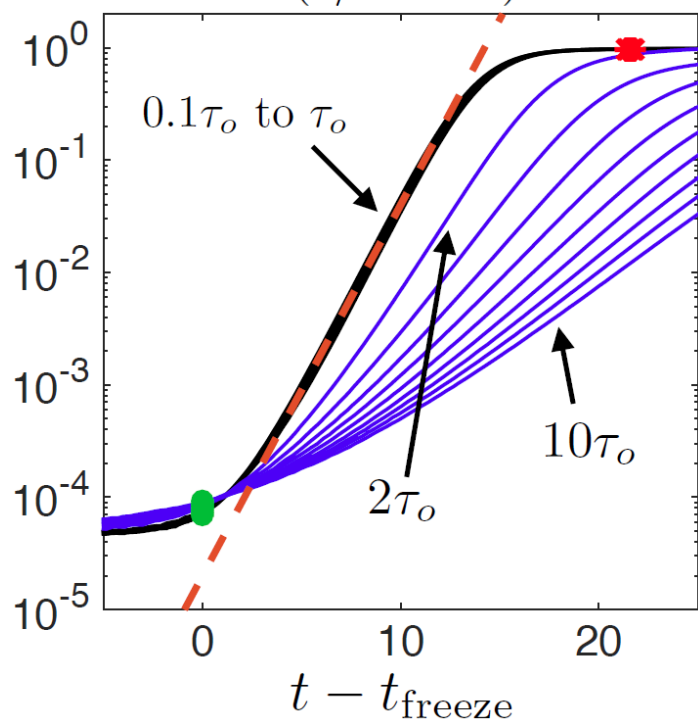
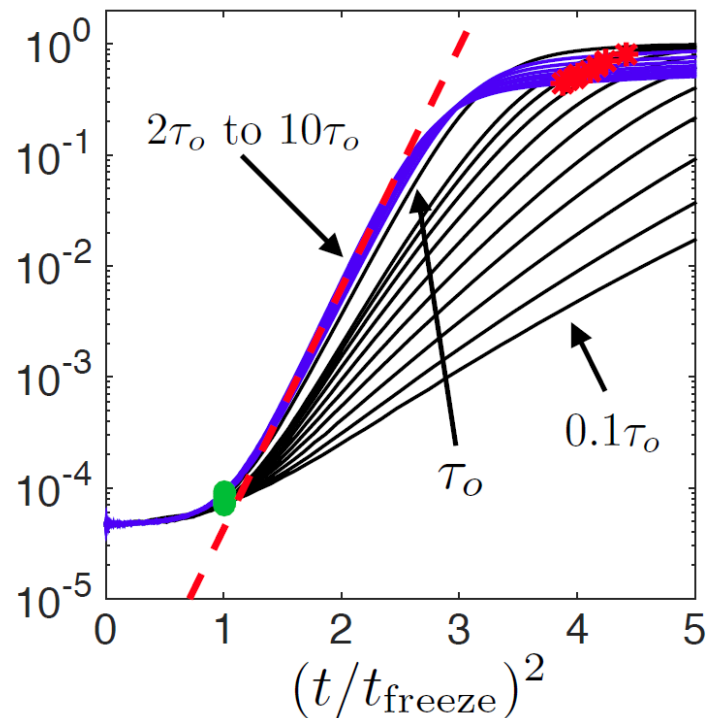
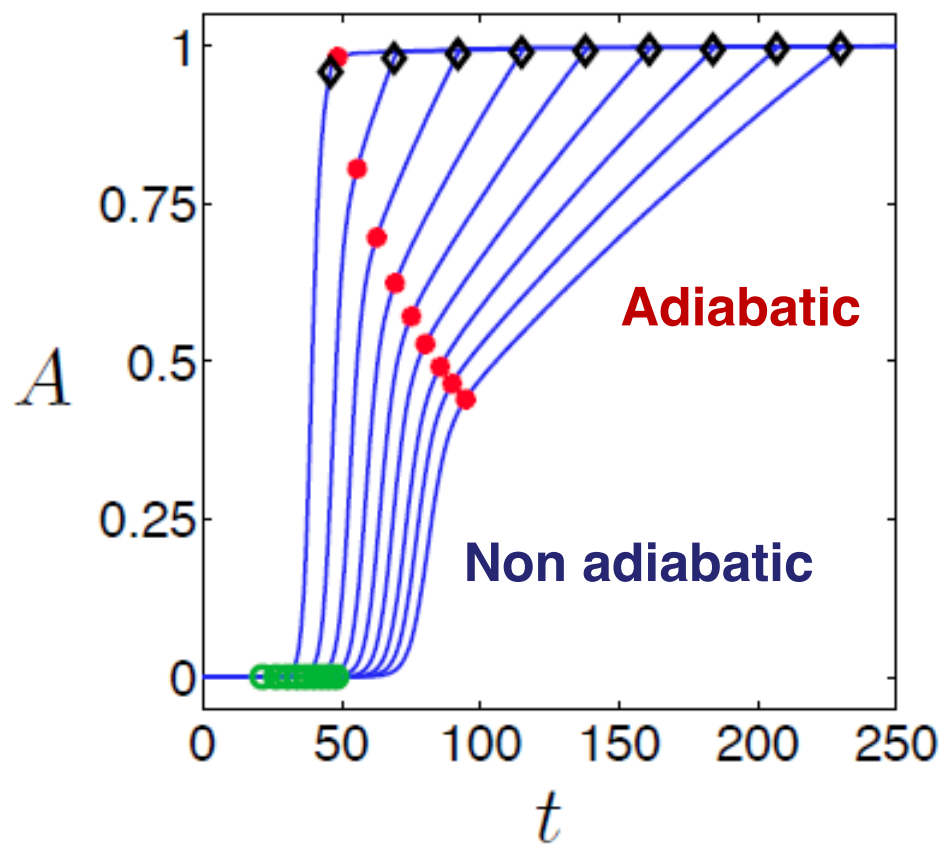


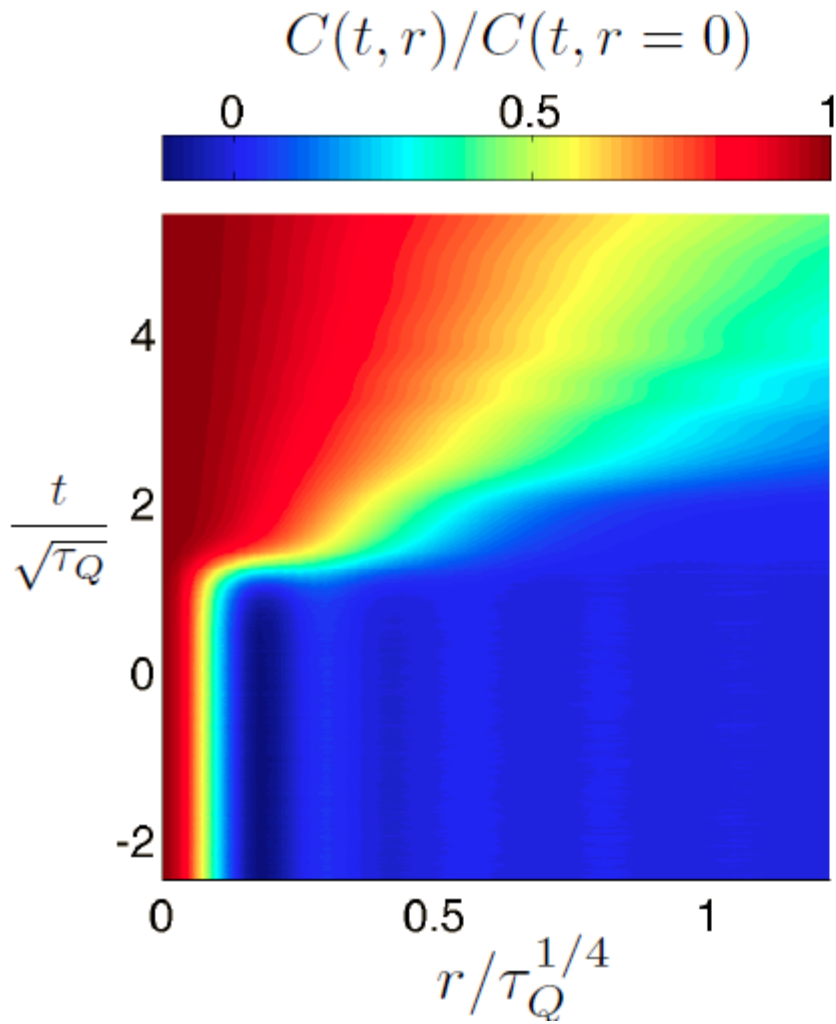
t_{eq} is the relevant scale

$$A(t) = \frac{1}{M} \sum_{i=1}^M \frac{a_i(t)}{a_i(\infty)}$$

$$a_i(t) \equiv \int d^2x |\psi_i(t, \mathbf{x})|^2$$

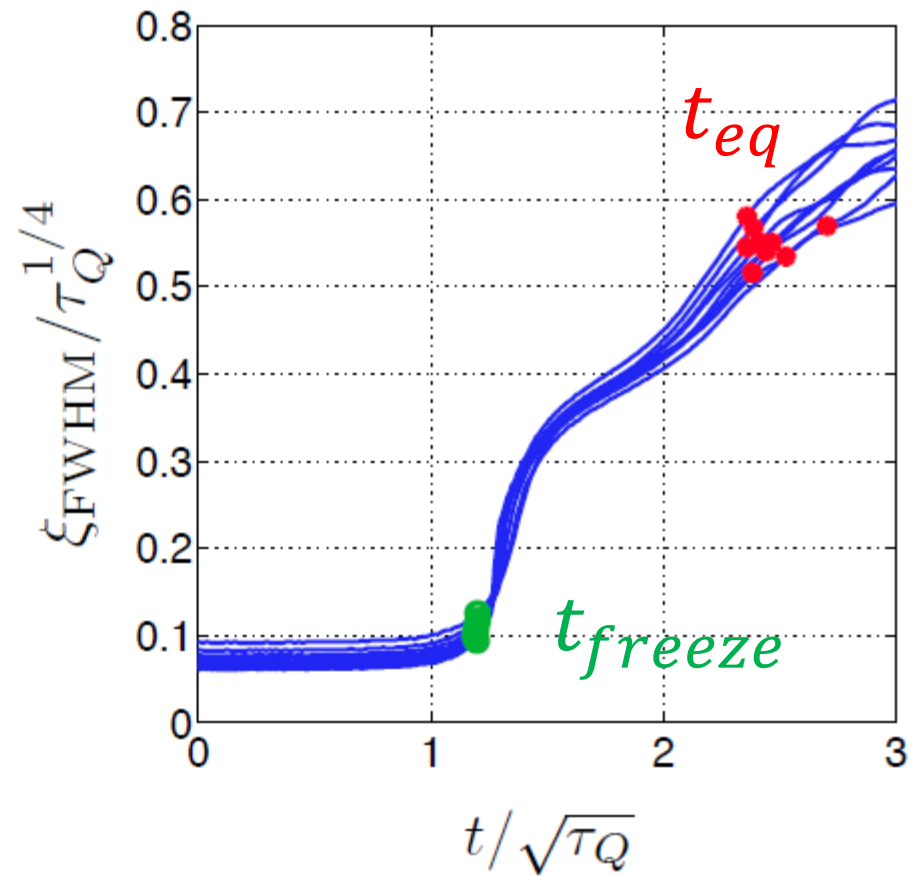
$$\tau_Q = 2\tau_o, \dots, 10\tau_o$$



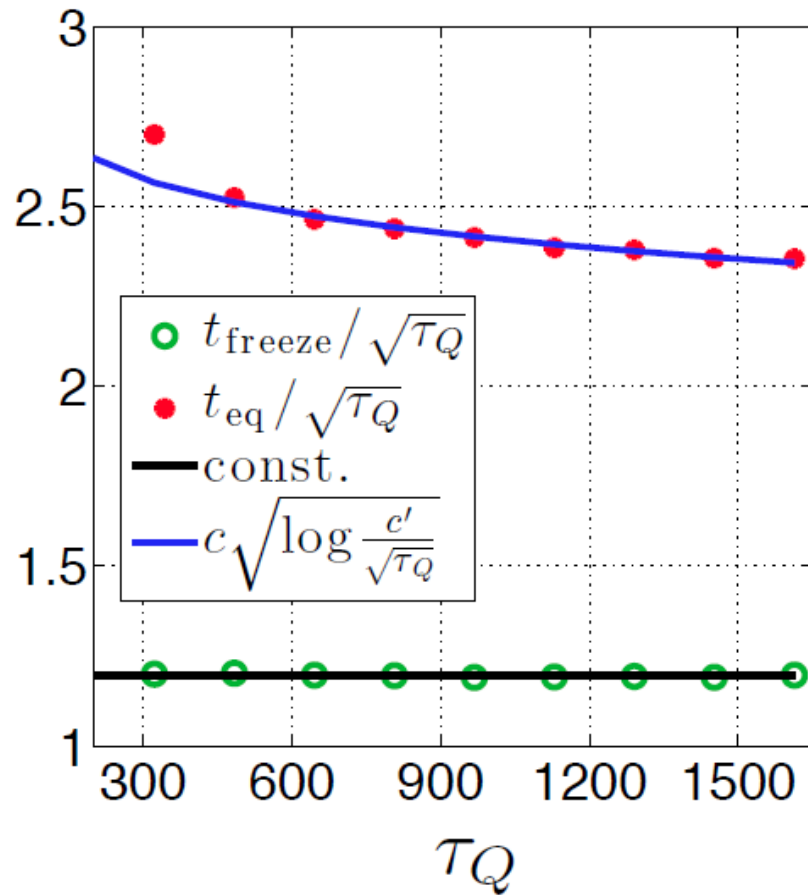
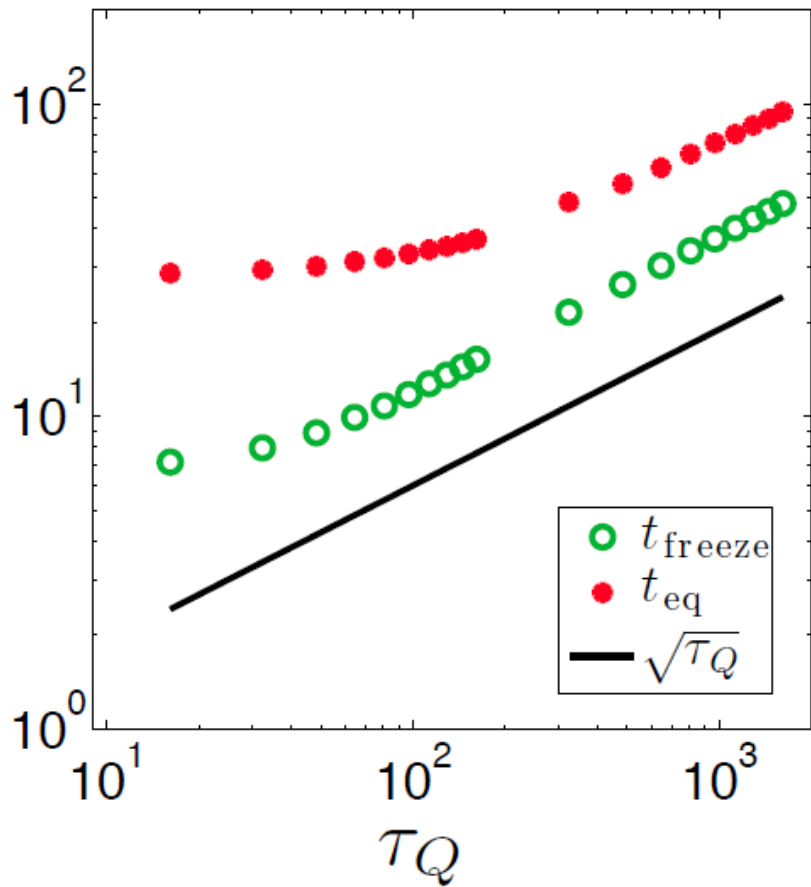


Strong coarsening
 $t > t_{freeze}$

Full width half max of $C(t, r)$



$$\ell_{co}(t) \sim \xi_{freeze} \sqrt{t/t_{freeze}}$$



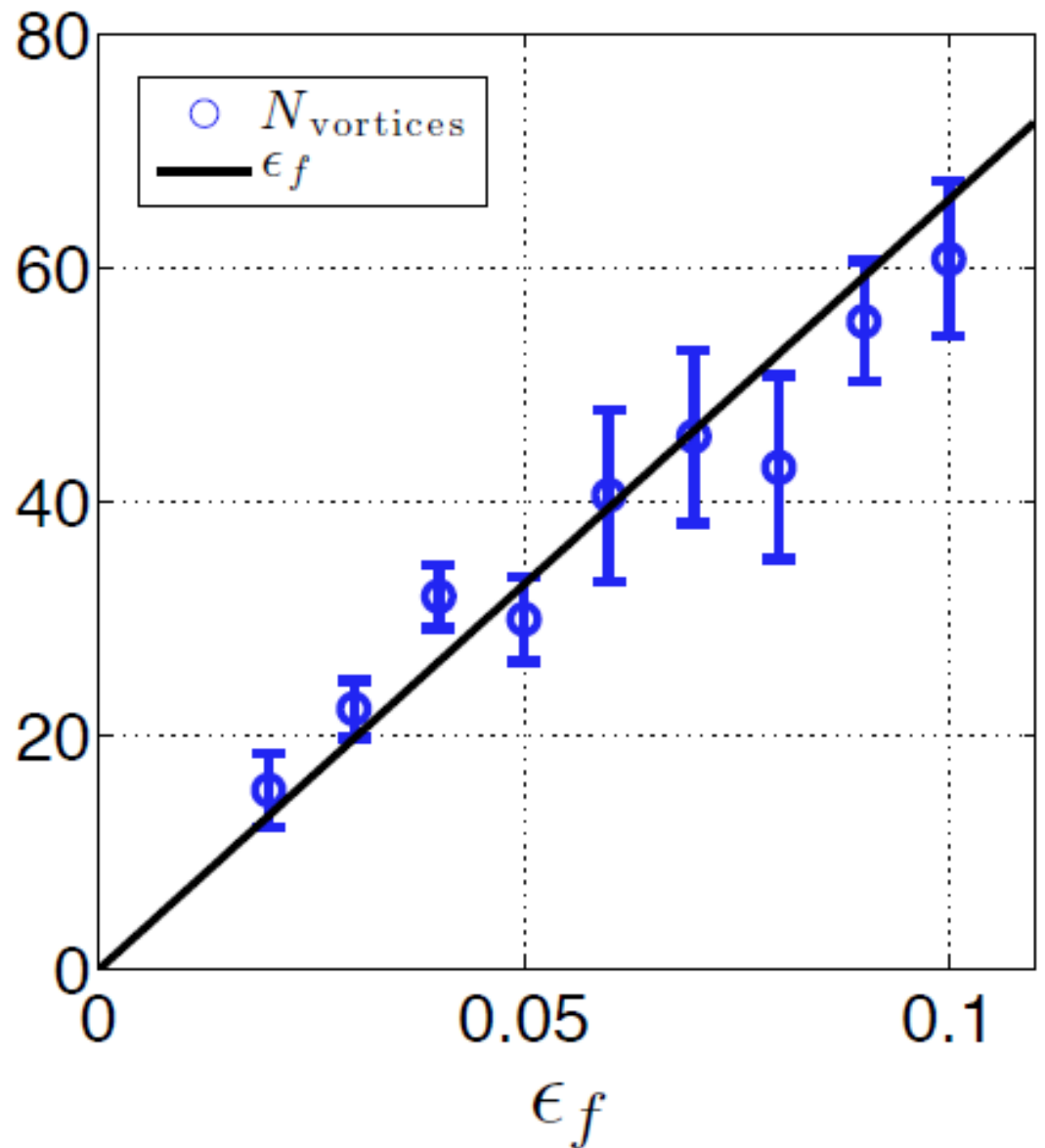
$$t_{freeze} \ll t_{eq}$$

Correct scaling

Fast
quenches

$$\rho \sim \frac{\epsilon_f}{\log\left(\frac{N^2}{\epsilon_f}\right)}$$

$T < T_c$
dynamic
irrelevant



Slow

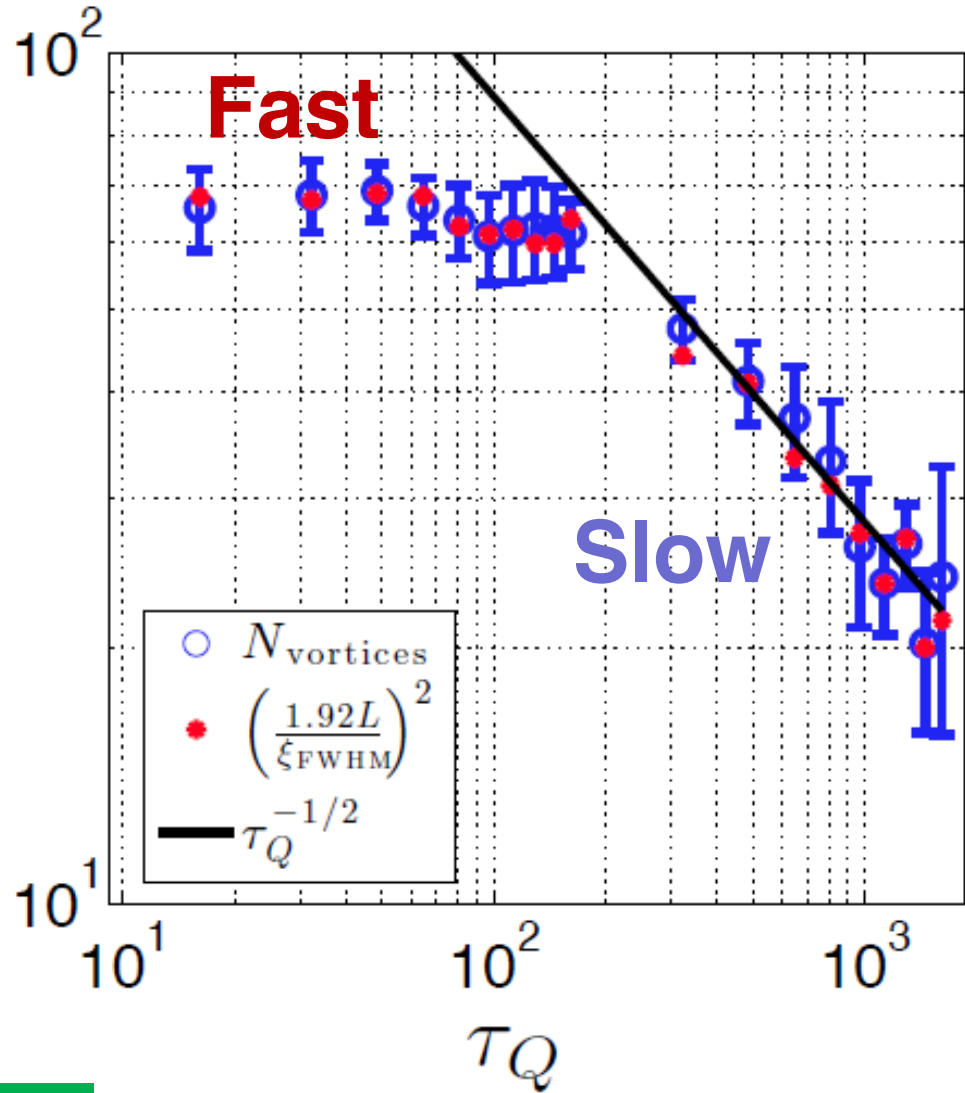
$$\rho \sim \frac{\rho_{KZ}}{(\log(N^2/\tau_Q^{1/2}))^{1/2}}$$

Fast

$$\rho \sim \frac{\epsilon_f}{\log\left(\frac{N^2}{\epsilon_f}\right)}$$

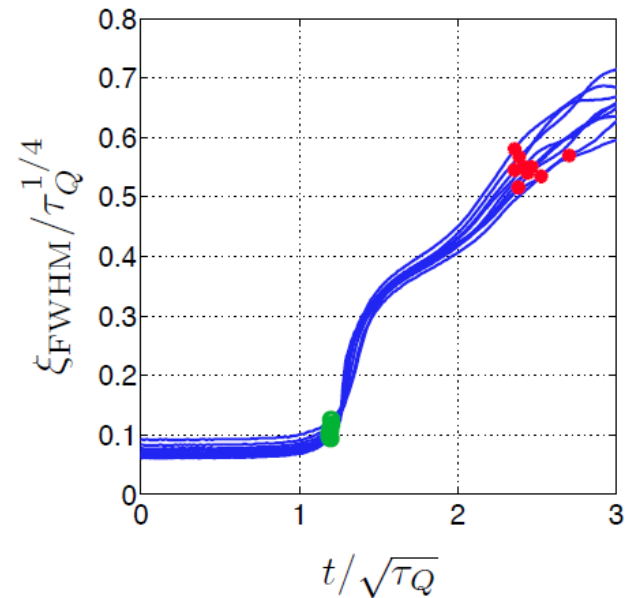
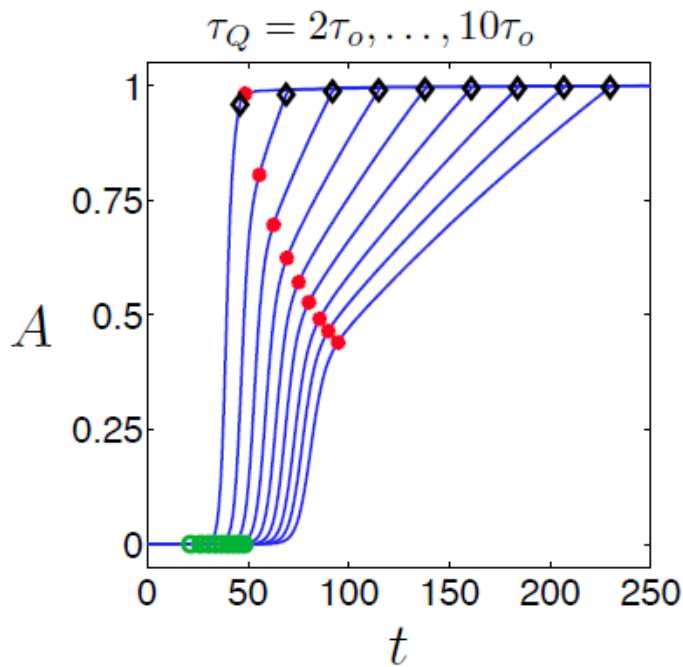
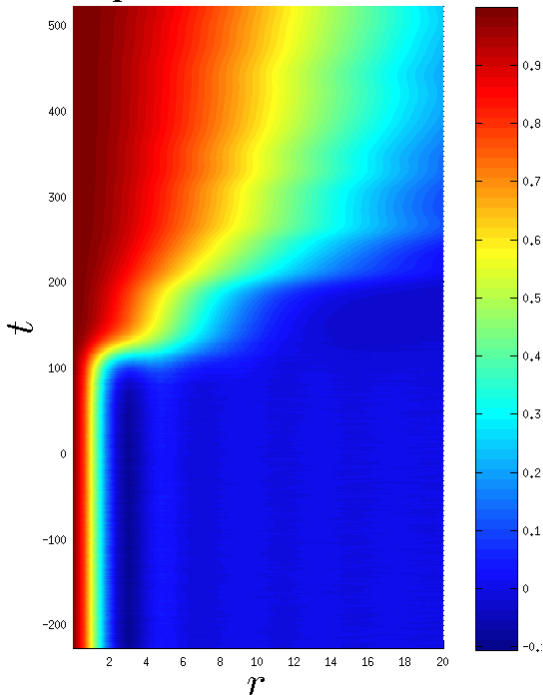
Relevant for ^4He ?

$$t_{\text{eq}} \sim [\log \tau_Q]^{1/(1+\nu z)} t_{\text{freeze}}$$



~25 times less defects than KZ prediction!!

phase correlator



time



Freezing

Condensate formation

Defect generation

Phase coherence ?

Physics beyond
Kibble-Zurek

Novel dynamical region

$$t_{\text{eq}} > t > t_{\text{freeze}}$$

Holography duality
helpful to discover and
model this region

More efficient than
SGPE?

NEXT

Cracking
thermalization?

${}^4\text{He?}$

Cosmology

Vortex physics

BKT transition

Thanks!