

Finite Baryon-Number · Induced
Chern-Simons Couplings
in Holographic QCD

(or: Smectic Phases in Neutron Stars?)

- I. Introduction and review of hard wall model
- II. Adding CS term and finite baryon density
- III. New couplings and dispersion relations
- IV. Outlook and Conclusions

I. HARD WALL MODEL

(Erlich, Katz, Son, Stephanov, Phys. Rev. Lett.; hep-ph/0501129)
 (da Rold, Pomarol, Nucl. Phys. B; hep-ph/0501218)

"bottom-up approach" to AdS/CFT/QCD

- Ingredients

- (i) gravitational background: AdS_5 (forget the S^5)
 with IR cutoff to induce confinement

$$ds^2 = \frac{1}{z^2} (dx^\mu dx_\mu - dz^2) \quad \text{with } 0 \leq z \leq z_m$$

(set AdS_5 radius $R=1$)

- (ii) field content:

(only include fields relevant to chiral dynamics)

5d masses given by $(\Delta - p)(\Delta + p - 4) = m_5^2$

\uparrow
operator scaling dimension \uparrow
form degree

- ① $SU(N_f)_L \times SU(N_f)_R$ gauge fields

$$A_{L,R\mu}^a \xleftrightarrow{\text{dual to}} \bar{q}_{L,R}^\alpha \gamma^\mu t^a q_{L,R}^\beta$$

$$(\Delta = 3, p = 1, m_5^2 = 0)$$

- ② Bifundamental tachyonic field (Higgs-like rôle)

$$\frac{2}{z} X^{\alpha\beta} \longleftrightarrow \bar{q}_{BR}^\alpha q_{BL}^\beta$$

$$(\Delta = 3, p = 0, m_5^2 = -3)$$

(' $\frac{2}{z}$ factor dictated by dimension of $\bar{q}_R q_L$)

(iii) Action

$$S = \int d^4x dz \sqrt{g} \text{Tr} \left[|DX|^2 + 3|X|^2 - \frac{1}{q g_s^2} (F_L^2 + F_R^2) \right] + S_{CS}$$

with $D_M X = Q_M X - i A_{L_M} X + i X A_{R_M}$

and generators normalized as $\text{Tr}(t_{LR}^a t_{LR}^b) = \frac{1}{2} \delta^{ab}$

↑
will
discuss
this
later

- Induce X_{SB} : by giving $X^{\alpha\beta}$ a z -dependent ver.

- (i) Solve linearized e.o.m. for X at zero 4d momentum:

$$X_0^{\alpha\beta} = \frac{1}{2} (m_q z + \sigma z^3) \mathbb{1}^{\alpha\beta} = \frac{v(z)}{2} \mathbb{1}^{\alpha\beta}$$

- z is non-normalizable mode
 \Rightarrow its coefficient is dual to source of operator $\bar{q}_R q_L$
- z^3 is normalizeable mode
 \Rightarrow its coefficient is dual to $\langle \bar{q}_R q_L \rangle$ (chiral condensate)

- (ii) Solve for hadron spectrum generated by normalizeable modes of vector $V_M^\alpha = \frac{1}{2} (A_{LM}^\alpha + A_{RM}^\alpha) = V_n(\hat{x}) \psi_n(z)$

and axial-vector $A_M^\alpha = \frac{1}{2} (A_{LM}^\alpha - A_{RM}^\alpha)$ gauge fields.

- confining hard wall at $z=z_m$ induces finite spectrum
- ver for $X^{\alpha\beta}$ induces mass splitting between vector and axial vector mesons

(iv) Fix undetermined parameters

g_5 z_m m_f σ
 by fitting m_p^2, f_π, \dots to data
 by matching
 to QCD result
 for $\langle J_{V_\mu}^a(x) J_{V_\nu}^b(0) \rangle$
 $(g_5^2 = \frac{12\pi^2}{N_c})$

RESULTS : Amazingly good agreement with data!

(For global fit to seven observables rms error $\sim 9\%$)

Some shortcomings :

- don't get correct Regge behavior ($m_n \sim n^2$)
- coupling doesn't run

... fixed in more complicated modifications,
 e.g. soft wall model

(Karch, Katz, Son, Stephanov,
 Phys. Rev D74; hep-ph/0602229)

II. CS TERM AND FINITE BARYON DENSITY

(SKD : J. Harvey, Phys. Rev. Lett. 99, arXiv: 0704.1604 [hep-ph])

Add to hard-wall model (with gauge group $U(N_f)_L \times U(N_f)_R$)

(i) Chern-Simons term to reproduce QCD chiral anomaly

$$S_{CS} = \frac{N_c}{24\pi^2} \int_M [\omega_5(A_L) - \omega_5(A_R)]$$

where for hermitian generators

$$\omega_5(A) = \text{Tr} \left[A F^2 + \frac{i}{2} A^3 F - \frac{1}{10} A^5 \right].$$

The variation of this $\omega_5 = d(F^2)$ gives a boundary variation precisely dual to the chiral anomaly:

$$\delta_\lambda \overset{(0)}{\omega_5} = d \overset{(1)}{\omega_4}$$

$$\text{where } \overset{(1)}{\omega_4} = \text{Tr} (\Lambda d(A dA - \frac{i}{2} A^3))$$

The coefficient is determined by anomaly - no extra undetermined parameters added.

(ii) Finite baryon density

i.e. turn on normalizable ($q^2 \approx 0$) mode of dual AdS_5 field

① Recall: we upgraded the gauge group to

$$U(N_f)_L \times U(N_f)_R$$

So we have a $U(1)_V$ gauge field dual to the baryon number current:

$$\hat{V}_B = \hat{V}_B^{\text{total}} \hat{t} + V_B^a \hat{t}^a$$

↓
 $U(1)_V$ generator ↑
 $SU(N_f)_V$ generator

$\Rightarrow \hat{V}_B$ is dual to the baryon number operator

② How to normalize \hat{V}_B ?

(i) Solve linearized equations of motion from action above

$$\hat{V}_B = A + B \frac{z^2}{2}$$

$\Rightarrow A$ is proportional to μ_q (or μ_B)

$\Rightarrow B$ is proportional to n_B

(ii) Evaluate S for Euclidean time (i.e. finite temp.)

(should give grand canonical potential)

$$\begin{aligned}
 S &= -\frac{1}{2g_5^2} \int d^4x dz \left(-\frac{1}{2} (\hat{F}_{0z})^2 \right) = \frac{1}{2g_5^2} \int d^4x \hat{V}_B \partial_z \hat{V}_B \Big|_{z=0} = \\
 &= \frac{1}{2g_5^2} AB \int d^4x = \frac{1}{2g_5^2} AB (\text{vol}_3 \cdot \beta)
 \end{aligned}$$



Compare to grand canonical ensemble

$$S_{\text{Euclid.}} = \mu_B \underbrace{(n_B \cdot \text{vol}_3)}_{N_B} \beta$$

Setting these equal we have

$$\frac{1}{2g_5^2} AB (\text{vol}_3 \cdot \beta) = \mu_B n_B (\text{vol}_3 \cdot \beta)$$

(relates A and B)

(iii) Fix A:

- Now use the Chern-Simons term.

$$S_{CS} = \frac{N_c}{64\pi^2} \int d^4x dz \epsilon^{MNPQ} \left[\hat{A}_{(L)} \text{Tr}(F_{MN} F_{UP}) - \hat{A}_{(R)} \text{Tr}(F_{(U)MN} F_{(P)UP}) \right] + \text{trace over SU(N_f)}$$

... + other terms ...

$$(M, N, P, Q = 1, 2, 3, \bar{z})$$

- Define $SU(N_f)$ instanton number in the usual way

$$N_{LR} = \frac{1}{32\pi^2} \int d^3x dz \epsilon^{MNPQ} \text{Tr}(F_{(L,R)MN} F_{(L,R)PQ})$$

but using z instead of t as fourth dimension

(Assume F_{MN} is independent of t : static configuration)

This gives a coupling in S_{CS} proportional to N_{LR} :

$$S_{CS} = \frac{N_c}{2} \int dx^0 (\hat{A}_{(L)} N_L - \hat{A}_{(R)} N_R)$$

Relying on the connection between static 3d "instantons" and Skyrme configurations (i.e. pion field configurations with nontrivial winding number)* we identify states with instanton number

$$N_L = -N_R = N_b$$

as baryons of baryon number N_b .

* [For details see Atiyah : Manton Physlett B 222
 Son : Stephanov PhysRev.D 69 hep-ph/0504182
 Hata, Sakai, Sugimoto; Yamamoto hep-th/0701276]

Now we have the coupling

$$S_{CS} = \frac{N_c}{2} N_B \int (\hat{A}_{(L)_0} + \hat{A}_{(R)_0}) dx^0 = N_c N_B \int dx^0 \hat{V}_0 = \beta N_c N_B A = \beta \mu_B N_B$$

$$\Rightarrow A = \frac{\mu_B}{N_c} = \mu_q$$

Overall, $A = \frac{\mu_B}{N_c}$ and $B = \frac{24\pi^2 n_B}{N_c} = 24\pi^2 n_B$:

$$\hat{V}_0 = \frac{\mu_B}{N_c} + 24\pi^2 n_B \frac{z^2}{2}$$

 focus on n_B from now on

III. Consequences of $n_B \neq 0$ and S_{CS}

(i) 4d effective action (to quadratic order)

Note: S_{CS} contains a term $\propto \hat{V}_R \partial_M \hat{V}_N \partial_P A_Q \epsilon^{MNPQ}$
which is quadratic order when $n_B \neq 0$!

- Expand as usual

$$\left\{ \begin{array}{l} V_\mu^a(x, z) = g_5 \rho_\mu^a(x) \Psi_p(z) \\ A_\mu^a(x, z) = g_5 a_\mu^a(x) \Psi_a(z) \end{array} \right. \quad \begin{array}{l} (\text{rho meson}) \\ (\text{a}_1 \text{ meson}) \end{array}$$

where $\Psi_{p,a}(z)$ satisfies linearized equations of motion
with $f_{4d}^2 = m_{p,a}^2$.

- Integrate out holographic (z) direction:
(and ignore pions)

$$S_{4d} = \int d^4x \left[-\frac{1}{4} (\rho_{\mu\nu}^a)^2 + \frac{1}{2} m_p^2 \rho_\mu^a \rho^\mu_a - \frac{1}{4} (a_{\mu\nu}^a)^2 + \frac{1}{2} m_{a_1}^2 a_\mu^a a^\mu_a \right. \\ \left. + \mu \epsilon^{ijk} (\rho_i^a \partial_j a_k^a + a_i^a \partial_j \rho_k) \right]$$

\uparrow
not chemical potential

where $i, j, k = 1, 2, 3$ and

$$\mu = 18 \pi^2 z_m^2 n_B I$$

(dimensions of mass)

here $I = \frac{1}{z_m^2} \int_0^{z_m} dz z \Psi_p(z) \Psi_{a_1}(z)$ gives overlap

between ρ and a_1 wavefunctions and $\partial_z \hat{V}_0$.

For densities typical of neutron stars $n_{B_0} \approx 0.16 \text{ fm}^{-3}$

$$\mu = (1.05 \text{ GeV}) \left(\frac{n_B}{n_{B_0}} \right)$$

(I is ~ 0.54)

(A) Mixing of transverse ρ and a_1 polarizations

- Consider ρ and a_1 in plane wave state

(without loss of generality, consider propagation along x^3)

$$\rho_\mu(x) = \epsilon_\mu^{(\rho)}(q) e^{-iq \cdot x} \quad a_{\mu}(x) = \epsilon_\mu^{(a)}(q) e^{-iq \cdot x}$$

(suppressing gauge indices)

$$\text{with } q = (q_0, 0, 0, q_3)$$

From the equations of motion we have

$$(\square + m_p^2) \rho^\beta = 2\mu \epsilon^{ij\beta} \partial_i a_j^a$$

$$(\square + m_a^2) a^\beta = 2\mu \epsilon^{ij\beta} \partial_i \rho_j^a$$

which gives the usual dispersion relations for
 $\beta=0$ and $\beta=3$, but for transverse polarizations
the CS term has an effect

$$\beta=1, 2 :$$

$$q_0^2 - q_3^2 = \frac{1}{2} (m_p^2 + m_a^2) \pm \sqrt{(m_a^2 - m_p^2) + 16\mu^2 q_3^2}$$

• for $q_3 \rightarrow 0$ "−" dispersion relation gives ρ :

$$\epsilon_1^{a_1} = \frac{iM^2(q_3)}{2\mu q_3} \epsilon_2^\rho \quad \epsilon_2^{a_1} = - \frac{iM^2(q_3)}{2\mu q_3}$$

• for $q_3 \rightarrow 0$ "+" dispersion gives a_1 :

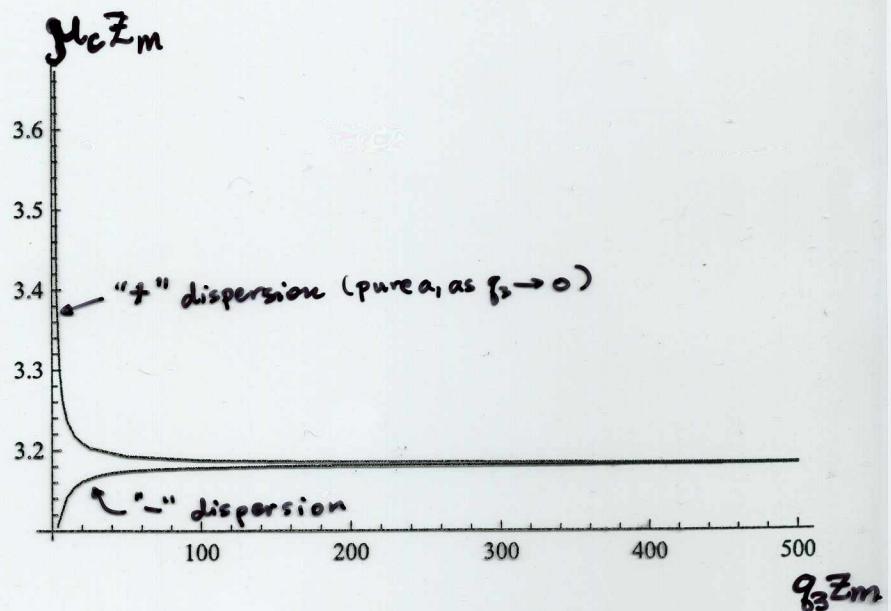
$$\epsilon_1^\rho = - \frac{iM^2(q_3)}{2\mu q_3} \epsilon_2^{a_1} \quad \epsilon_2^\rho = \frac{iM^2(q_3)}{2\mu q_3} \epsilon_1^{a_1}$$

with $M^2(q_3) = \frac{\sqrt{\Delta^4 + 16\mu^2 q_3^2} - \Delta^2}{2}$ for $\Delta^2 = m_a^2 - m_p^2$ - 10 -

- For $\mu > \mu_c$ (which depends on q_3)
 the dispersion relation exhibits
 tachyonic modes ($\frac{dq_0}{dq_3} > 1$)
 indicating a possible instability.
 (i.e. group velocity faster than c)

→ For very large q_3

$$\mu_c \approx \sqrt{\frac{m_a^2 + m_p^2}{2}} \approx 1.09 \text{ GeV}$$



μ_c as a function of q_3

($\mu > \mu_c$ gives $\frac{dq_0}{dq_3} > 1$)

(B) Meson Condensation

Instability in dispersion hints at possibility of forming a condensate!

Rough idea:

(i) Consider the energy density of the system:

$$\mathcal{L} = \frac{1}{2} (\Pi_p^2 + \Pi_a^2) + \frac{1}{4} (p_{ij}^2 + a_{ij}^2) + \frac{1}{2} m_p^2 p_\mu p^\mu + \frac{1}{2} m_a^2 a_\mu a^\mu - \mu \sum_{ijk} (p_i \partial_j a_k + a_i \partial_j p_k)$$

↑ ↑
 canonical
momenta
 (forget from
now on...)

Complete the square, with $\left\{ \begin{array}{l} \vec{B}_p = \vec{\nabla} \times \vec{p} \\ \vec{B}_a = \vec{\nabla} \times \vec{a} \end{array} \right.$

$$\mathcal{L} = \dots + \frac{1}{2} (m_a^2 - \mu^2) \vec{a} \cdot \vec{a} + \frac{1}{2} (m_p^2 - \mu^2) \vec{p} \cdot \vec{p} + \frac{1}{2} (\vec{B}_a - \mu \vec{p})^2 + \frac{1}{2} (\vec{B}_p - \mu \vec{a})^2$$

(ii) Apply the ansatz

$$\vec{a} = v \cos(\mu x_3) \hat{x}_2 \quad \vec{p} = v \sin(\mu x_3) \hat{x}_1$$

- last two terms in \mathcal{L} vanish
- for $\mu > \sqrt{\frac{m_a^2 + m_p^2}{2}}$ we have

$$\mathcal{L} = v^2 \left[\frac{m_a^2 + m_p^2}{2} - \mu^2 \right] \neq 0 !$$

- Need to stabilize the condensate with higher order terms!

LOOKS LIKE SMECTIC PHASE IN LIQUID CRYSTALS.

IV. Conclusions and Outlook

- We added a CS term to the hardwall model in order to reproduce the chiral anomaly
- At finite density holography tells us this affects the dispersion relations of transverse polarizations of (axial)vector mesons.
 - mixed polarization states
 - (axial)vector meson condensate

Lots left to be done:

- stabilize condensate with higher order terms, and examine its structure in detail
- interactions with pion ?
- add temperature

...

A quick note on 5d instantons vs. Skyrmions:

(nicely explained in e.g. Son & Stephanov Phys Rev D 69
arXiv: hep-ph/0304182)

(also in Hata, Sakai, Sugimoto, Yamamoto
arXiv: hep-th/0701028)

- 5d YM has a conserved topological current

$$\sqrt{g} j^M = \frac{1}{32\pi^2} \epsilon^{MNPR} F_{NP} F_{QR}$$

With topological charge

$$Q = \int d^2 z d^3 x \sqrt{g} j^0 = \frac{1}{32\pi^2} \int d^2 z d^3 x \underbrace{\epsilon^{0MNPQ} F_{MN} F_{PQ}}_{\text{usual instanton # (with } z \text{ instead of } t\text{)}}$$

- Recall that $d\omega_3 = \text{Tr}(F^2)$ or define another current

s.t. $K^0{}^M = 4 \epsilon^{0MNPQ} \text{Tr}(A_N \partial_P A_Q - \frac{2i}{3} A_N A_P A_Q)$

Say in this model we can define the pion in terms of

a Wilson line $\overbrace{z}^{\text{pion}} \int_0^z dz' A_z$

and A_H vanishes on the boundaries. Then, going to $A_z = 0$ gauge requires a gauge transformation by U , so $A_\mu \rightarrow U A_\mu U^{-1} + i U \partial_\mu U^{-1}$
(or just $i U \partial_\mu U^{-1}$ on the boundaries).

Now, using $K^0{}^M$ we can rewrite the top. charge Q :

$$Q = \int d^2 z d^3 x \frac{1}{32\pi^2} \partial_M K^0{}^M = \int d^3 x \frac{1}{32\pi^2} K^0 z / \Big|_{z=0}^{z_m} = \\ = \frac{1}{24\pi^2} \int d^3 x \epsilon^{ijk} (U^\dagger \partial_i U) (U^\dagger \partial_j U) (U^\dagger \partial_k U) \left. \right\}^{\text{top. charge}}_{\text{in Skyrme model}}$$