

Black rings

Firenze 23-24 April 2007

Context: hi-d gravity & black holes

Motivation:

- String / M-Theory
- LXD-TeV gravity scenarios: BH's at colliders
 \downarrow
 $D > 4$: we know little of their physics!
- AdS/CFT: 4D QFT physics from hi-d gravity
- D as a tunable parameter to study gravity & bh's (analogous to "unphysical" lagrangians in QFT, or N in $SU(N)$)
what properties are deg/ind of D ?
- Natural Theoretical interest / Mathematical physics: math of Lorentzian manifolds

Plan

- Day 1:- vacuum (neutral):
- hi-d bh's
 - Ring coords
 - Neutral black ring

Day 2 - charges & dipoles

Day 3: susy & micro

References:

- review w/ H. Reall hep-th/0608012 (more details in hep-th/0608076 w/ Kraang + Virmani)
- for an alternative view on The susy case Benat Warner hep-th/0701216

- 4D black holes : virtually all is known (the simplest macroscopic objects in the universe)

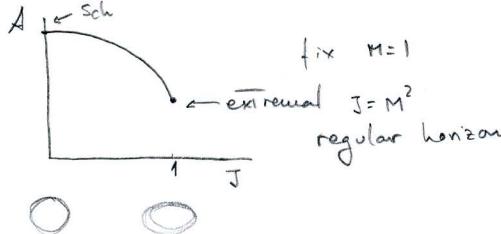
$$\text{Schwarzschild} \quad ds^2 = \left(1 - \frac{\mu}{r}\right) dt^2 + \frac{dr^2}{1 - \frac{\mu}{r}} + r^2 d\Omega_2^2 \quad \mu = 2GM$$

$$\text{Kerr} \quad (1963) \quad ds^2 = -dt^2 + \frac{\mu r}{\Sigma} (dt + a \sin^2\theta d\phi)^2 + \Sigma \left(\frac{dr^2}{\Delta} + d\theta^2 \right) + (r^2 + a^2) \sin^2\theta d\phi^2 \quad \mu = 2GM$$

$$J = Ma$$

$$\Sigma = r^2 + a^2 \cos^2\theta$$

$$\Delta = r^2 + a^2 - \mu$$



Horizon at $\Delta = 0$

$$r = r_+ = M + \sqrt{M^2 - a^2}$$

$$= M(1 + \sqrt{1 - \frac{J^2}{M^2}})$$

(G=1)

- $D > 4$

$$\text{Taub-NUT} \quad (1963) \quad ds^2 = \left(1 - \frac{\mu}{r^{D-3}}\right) dt^2 + \frac{dr^2}{1 - \frac{\mu}{r^{D-3}}} + r^2 d\Omega_{D-3}^2 \quad \mu \propto GM \quad \begin{array}{l} \text{Easy!} \\ \text{- Stable} \\ \text{- Unique} \end{array}$$

Myers+Perry : several rotation planes are possible: $t, x_1, x_2, x_3, x_4, x_5 \dots$
 $(\text{Cartan towers of } SO(D-1) \times U(1)^{\lfloor \frac{D-2}{2} \rfloor})$
 Found solutions in arbitrary D w/ rotation in all planes!

w/ only one rotation

$$ds^2 = -dt^2 + \frac{\mu r^{D-3}}{\Sigma} (dt + a \sin^2\theta d\phi)^2 + \Sigma \left(\frac{dr^2}{\Delta} + d\theta^2 \right) + (r^2 + a^2) \sin^2\theta d\phi^2 + r^2 \cos^2\theta d\Omega_{D-4}^2 \quad \mu \propto M \quad J \propto Ma$$

Apparently nothing new, but there

$$\Delta = r^2 + a^2 - \frac{\mu}{r^{D-3}}$$

are important differences: competition between
 grav. pot. and centrifugal

$$\frac{\Delta}{r^2} - 1 = -\frac{\mu}{r^{D-3}} + \frac{a^2}{r^2}$$

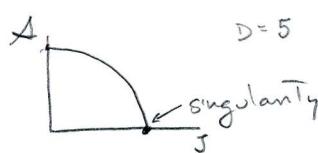
$\hookrightarrow D\text{-indep: confined to plane}$
 $\sim \text{grav. pot}$ $\sim D\text{-dep centrifugal pot}$

$$\text{Horizons: } \Delta = 0 \quad r_+^2 + a^2 - \frac{\mu}{r_+^{D-3}} = 0$$

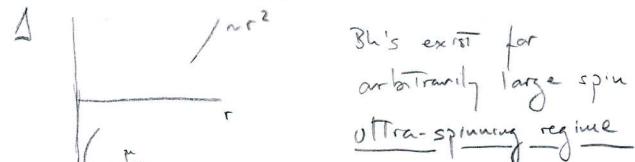
$D=4$: real solution for $a \leq M/2$

$D=5$: " " " " $a^2 \leq \mu$

$a = \mu$ $r_+ = \infty$ singular
 extremal limit



$D > 5$: real solution for all a !



only in $D \geq 6$ There are black



$$g_{\phi\phi}(r=r_+) = \frac{(\mu r_+^{D-3})^2 \sin^2\theta}{\Sigma_+}$$

$$g_{\phi\phi}(r=r_+) = \Sigma_+$$

$$\sqrt{g_{tt} g_{rr}}|_{r=r_+} = \mu r_+^{D-3} \sin\theta$$

$$\sqrt{g_{\phi\phi}}|_{r=r_+} = \omega_{D-2} \mu r_+$$

$$k = \partial_t + \Omega_H \partial_\phi$$

$$\Omega_H = -\frac{\partial \phi}{\partial t}|_{r=r_+} = \frac{a}{r_+^{D-3}}$$

$$|k|^2 = 0 \text{ at } r=r_+$$

Also: $D > 5$ no inner horizon
 singularity spacelike
 $r=0$ is singular since $|\partial_\phi|^2 \rightarrow \infty$
 unlike 4D where $g_{\phi\phi} \equiv 0$ at $r=0, \theta=0$

- Are MP bls all there is? (in AF vacuum gravity)

No: 3 black rings in all $D \geq 5$

First example: $D=5$ black rings

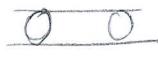
- 5D is actually fairly well-understood by now and we may be close to having a complete catalogue of bls.

Sector w/ $U(1)^3$ symmetry is completely integrable

time + 2 rotations

- $D > 5$ is less well known but progress is being made

- Heuristics of black rings:



black ring = $Schwar_{D-1} \times \mathbb{R}$



cut and glue



spin up: is it possible to balance forces?

Forces are: grav. D-dependent $\sim \frac{1}{R^{D-3}}$ (potential)

Tension } D-independent
centrifugal }

Grav is negligible for thin rings



$$R \gg r_+$$

Can spin balance tension? Yes, if $D \geq 5$

- $D=5$ black ring:

- Ring coordinates

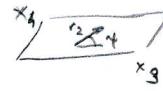
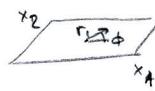
Flat space coords

$$x^1 = r_1 \cos \phi$$

$$x^3 = r_2 \cos \psi$$

$$x^2 = r_1 \sin \phi$$

$$x^4 = r_2 \sin \psi$$



There'll be J_+, J_ϕ . Set $J_\phi = 0$

$$ds^2 = dr_1^2 + r_1^2 d\phi^2 + dr_2^2 + r_2^2 d\psi^2$$

Find adapted coords for a ring-like source: equipotential surfaces

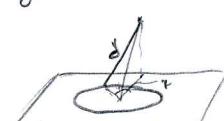
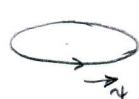
Not scalar source (possible but less convenient)

Int 2-form potential $B_{\mu\nu}$

source at $r_1=0, r_2=R, 0 \leq \psi \leq 2\pi$

$H = dB$
(eg test field of a fundamental string)

$$d^2 = r_1^2 + r_2^2 + R^2 - 2Rr_2 \cos \psi$$



Black ring: like previous, but w/ non-zero curvature encoded in functions

(Get this by double Wick from KK metric $\rightarrow Q$)
has been derived systematically using integrability, solitonic methods)

$$ds^2 = -\frac{F(\gamma)}{F(x)} \left(dt - CR \frac{1+\gamma}{F(\gamma)} dt \right)^2 + \frac{R^2}{(x-\gamma)^2} \frac{1}{F(x)} \left[-\frac{g(\gamma)}{F(\gamma)} d\gamma^2 - \frac{d\gamma^2}{g(\gamma)} + \frac{dx^2}{g(x)} + \frac{g(x)}{F(x)} d\phi^2 \right]$$

$$F(\gamma) = 1 + \lambda \gamma$$

$$g(\gamma) = (1 - \gamma^2)(1 + \nu \gamma)$$

$$C = \sqrt{\lambda(\lambda-\nu)\frac{1+\lambda}{1-\lambda}}$$

$$0 < \nu \leq \lambda < 1$$

$\lambda = \nu = 0$: flat

Horizon at $y = -1/\nu$ (where $G(y) = 0$)

$\Rightarrow y = \text{const}$ has Topology $S^1 \times S^2$: black ring

Examine S^2 , make $x = \cos\theta$

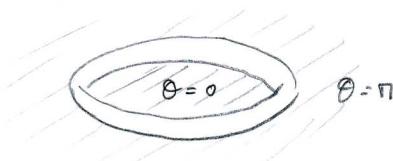
$$\frac{dx^2}{(1-x^2)(1+\nu x)} + \frac{(1+\nu x)}{1+\lambda x} (1-x^2) d\phi^2 = \frac{1}{1+\nu \cos\theta} \left(d\theta^2 + \frac{(1+\nu \cos\theta)^2}{1+\lambda \cos\theta} \sin^2\theta d\phi^2 \right)$$

Near $\theta \approx 0$ $d\theta^2 + \frac{(1+\nu)^2}{1+\lambda} \theta^2 d\phi^2 \Rightarrow \Delta\phi = 2\pi \frac{\sqrt{1+\lambda}}{1+\nu}$

Near $\theta = \pi$ $d\theta^2 + \frac{(1-\nu)^2}{1-\lambda} (\pi-\theta)^2 d\phi^2 \Rightarrow \Delta\phi = 2\pi \frac{\sqrt{1-\lambda}}{1-\nu}$

$$\lambda = \frac{2\nu}{1+\nu^2}$$

balance condition



If we impose regularity only at $\theta = \pi$
 (circle) \Rightarrow conical defect "membrane"
 extra force to balance ring
 No comes \Leftrightarrow equilibrium

Parameters: R, ν ($\lambda(\nu)$)
 \downarrow
 scale "Thickness" $\sim \frac{R_{S2}}{R}$
 (size of S^1)

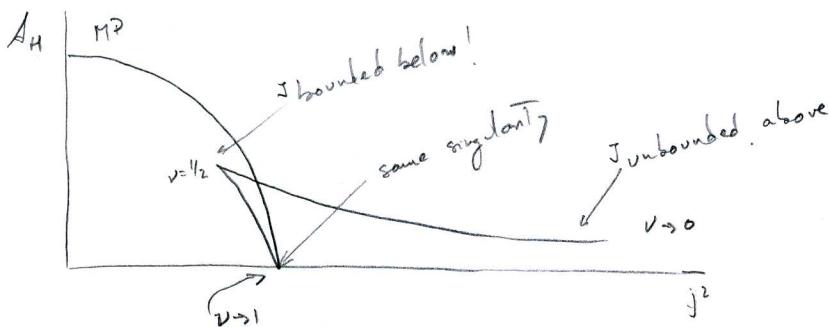
Physically: M, J, R balance fixes one of them in terms of the other two
 e.g. given M, J , Then R is fixed

One can measure for this solution M, J (going to asymptotic φ and comparing to linearized soln): $g_{tt} \approx -1 + \frac{GM}{r^2} + \dots$ $\partial_\phi \approx \frac{J}{r^3} + \dots$

A_H

Fix $M=1$, Then only variable is J . Can plot $A_H(J)$

$$j^2 = \frac{27\pi}{32G} \frac{J^2}{M^3} \quad (\text{or } M = \left(\frac{27\pi}{32G}\right)^{1/3})$$



$$A_H = \frac{3}{16} \sqrt{\frac{3}{\pi}} \frac{A_H}{(GM)^{3/2}}$$

$$A_H = 2\sqrt{2(1-j^2)}$$

$$\text{Ring } A_H = 2\sqrt{4(1-j^2)}$$

$$j^2 = \frac{(1+\nu)^2}{8\nu} \quad 0 \leq \nu < 1$$

$$\partial_{\mu} (\sqrt{g} H^{\mu\nu\rho}) = 0$$

$$B_{t+\phi} = \frac{R}{2\pi} \int_0^{2\pi} dt \frac{r_2 \cos \gamma}{r_1^2 + r_2^2 + R^2 - 2Rr_2 \cos \gamma}$$

$$= -\frac{1}{2} \left(1 - \frac{R^2 + r_1^2 + r_2^2}{\Sigma} \right)$$

$$\Sigma = \sqrt{(r_1^2 + r_2^2 + R^2)^2 - 4R^2 r_2^2}$$

Construct new gradient surfaces orthogonal to this: electric-magnetic dual potential
In 5D, A_μ is dual to $B_{\mu\nu}$ $B_{t+\phi} \xrightarrow{\text{dual}} A_\phi$: monopole field 

$$\partial_\mu (F_g F^{\mu\nu}) = 0$$

$$A_\phi = -\frac{1}{2} \left(1 + \frac{R^2 - r_1^2 - r_2^2}{\Sigma} \right)$$

Now set $B_{t+\phi} = -\frac{1}{2} (1 + \gamma)$ These define new coords (x, y) in terms of r_1, r_2

$$A_\phi = -\frac{1}{2} (1 + x)$$

$$r_1 = R \frac{\sqrt{1-x^2}}{x-\gamma} \quad r_2 = R \frac{\sqrt{\gamma^2-1}}{x-\gamma}$$

Note $\bullet -\infty \leq \gamma \leq -1 \quad -1 \leq x \leq 1$

$\bullet \gamma = -\infty$ is ring source $\begin{cases} r_1 = 0 \\ r_2 = R \end{cases}$

\bullet infinity $\begin{cases} r_1 \rightarrow \infty \\ r_2 \rightarrow \infty \end{cases} \quad x \rightarrow \gamma \rightarrow -1$

$$ds^2 = \frac{R^2}{(x-\gamma)^2} \left[(\gamma^2 - 1) dt^2 + \frac{dy^2}{\gamma^2 - 1} + \frac{dx^2}{1-x^2} + (1-x^2) d\phi^2 \right]$$

note $x \leftrightarrow \gamma$ symmetry \Rightarrow elec-mag duality

Can redefine (but still same foliation) $y = -\frac{R}{r} \quad x = \cos \theta$
 $0 \leq r \leq R \quad 0 \leq \theta \leq \pi$

$$ds^2 = \frac{1}{\left(1 + \frac{\cos \theta}{R}\right)^2} \left[\underbrace{(1 - \frac{r^2}{R^2}) R^2 dt^2}_{S^1} + \underbrace{\frac{dr^2}{1 - \frac{r^2}{R^2}}}_{S^2} + r^2 (d\theta^2 + \sin^2 \theta d\phi^2) \right]$$

$r = \text{const}$ ($\gamma = \text{const}$) are $S^1 \times S^2$: ring-like

: convenient for $r \ll R$

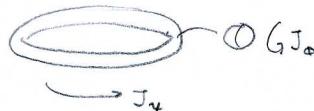
$$ds^2 = R^2 dt^2 + dr^2 + r^2 (d\theta^2 + \sin^2 \theta d\phi^2)$$

but not for $r \approx R$

Picture of x, y coords

Extensions:

- Two sprs J_+, J_-



Pomeransky + Senkov

- Ads not yet

- D>S : $S' \times S^1$
yes, in progress

↳ Extension definition of the isotropic plane

To big more difficult to solve the boundary value problem (inverse) Fourier transformation of the boundary to calculate the (inverse) Fourier transformation of the boundary to calculate the (inverse)

- charges, easy: tried early but failed

We know that is a first step to calculate the boundary value problem of the Dirichlet problem to get the boundary value problem of the Neumann problem. We have to solve the boundary value problem of the Neumann problem to get the boundary value problem of the Dirichlet problem. We have to solve the boundary value problem of the Neumann problem to get the boundary value problem of the Dirichlet problem.

$$\dots + \frac{((\lambda \eta \Omega)_n - \lambda)}{\lambda \eta \Omega - 1 - \lambda \eta \Omega} e^{i \lambda \eta \Omega t} \sin(\lambda \eta \Omega x) \text{ th } \sum_{n=0}^{\infty} n^2 \sum_{l=n}^{\infty} \frac{1}{\pi \Omega} = (u) \eta$$

For large values of n the expansion can be well approximated by the sum of the first few terms.

$$\dots + \frac{((\lambda \eta \Omega)_n \lambda \eta \Omega \sin(\lambda \eta \Omega t) - \lambda \eta \Omega \cos(\lambda \eta \Omega t)) \text{ th }}{\lambda \eta \Omega - 1 - \lambda \eta \Omega} \sum_{n=0}^{\infty} n^2 \sum_{l=n}^{\infty} \frac{1}{\pi \Omega} = (u) \eta$$

[6] similar solution for the boundary value problem of the Legendre test of T

$$\dots \left(\left(\overline{c(\pm \eta)(\sqrt{\eta \Omega} + 1)} \sqrt{\eta \Omega} + 1 \right) \pm \eta \right) \frac{\partial^n (\eta \Omega)}{\eta \Omega} - \frac{((\eta \Omega)_n \lambda \eta \Omega \sin(\lambda \eta \Omega t) - \lambda \eta \Omega \cos(\lambda \eta \Omega t)) \text{ th }}{\lambda \eta \Omega - 1 - \lambda \eta \Omega} \sum_{n=0}^{\infty} n^2 \sum_{l=n}^{\infty} \frac{1}{\pi \Omega} = (u) \eta$$

$\lambda \eta \Omega \mp \sqrt{\eta \Omega} = \pm \eta$ if $\eta \neq 0$

(u)η is the desired approximation for the desired value of u

$$\dots + \eta \lambda \eta \Omega \sum_{n=0}^{\infty} \frac{1}{\pi \Omega} = (u) \eta$$

approx

$$\dots + \left[\overline{c(\pm \eta)(\sqrt{\eta \Omega} + 1)} \sqrt{\eta \Omega} + 1 \right] \eta + \overline{c(\pm \eta)(\sqrt{\eta \Omega} + 1)} \eta + \frac{\partial^n (\eta \Omega)}{\eta \Omega} = \eta$$

$\lambda \eta \Omega \mp \sqrt{\eta \Omega} = \pm \eta$ if $\eta \neq 0$ and the boundary condition is known so it is possible to calculate the boundary condition for the boundary value problem of the Legendre test of T.

$$(D.S) \quad \dots + (u)^{\lambda \eta \Omega} \eta^2 + (u)^{\lambda \eta \Omega} \eta^2 = (u) \eta$$

To find the leading contribution of (u)η we have to develop the right side of the equation in powers of η . The result is not unique, it depends on whether it is zero or not. If it is zero we have to take the limit of the series for the value of u.

$$\left. \begin{aligned} & \text{if } |\eta| > |\eta|_0 = \text{min} \left(\text{Re } \left(\frac{\partial^n (\eta \Omega)}{\eta \Omega} \right), \text{Re } \left(\frac{\partial^n (\eta \Omega)}{\eta \Omega} \right) \right) \end{aligned} \right\} = u \eta \\ & \left. \begin{aligned} & \text{if } |\eta| < |\eta|_0 = \text{min} \left(\text{Re } \left(\frac{\partial^n (\eta \Omega)}{\eta \Omega} \right), \text{Re } \left(\frac{\partial^n (\eta \Omega)}{\eta \Omega} \right) \right) \end{aligned} \right\} = u \eta$$

Charged and easy rings & many rings

Charges and dipoles

$$\int R \cdot F^2$$

Ricci - Nordström has electric charge Q (electrovac)

Kerr - Newman has " " and rotation \rightarrow magnetic dipole, but not independent parameter

In 5D : Einstein - Maxwell + RN (easy)

"Kerr-Newman"-like : only recently numerically : very hard
Black ring w/ dipole (and no net charge!)

Einstein - Maxwell - Chern-Simons (w/ precise coupling) : 3MPV^{SUSY} and non-susy extension

bosonic sector of
minimal 5D sugra

Dipole ring (CS term irrelevant)
susy and non-susy rings

Independent dipoles imply continuous instead of discrete
non-uniqueness

$N=1, D=5$ sugra
8 supercharges

Spin(1,4) = Spin(1,1) = USp(2)
irreducible spinor becomes
a quaternion doublet
Multiply a solution to the
Killing spinor eq by a constant
quaternion \rightarrow general solution
has 4 real dof's

We'll be interested in the solutions to 5D sugras, and their 11D uplifts

so

$$I = \frac{1}{16\pi G} \int R \star 1 - \frac{1}{2} F \star *F - \frac{1}{6} G \star G \star A \quad 11D \text{ sugra}$$

Reduce on T^6 (could also do it in CY_3)

$$ds_5^2 = ds_3^2 + \mathbb{X}^1 (dz_1^2 + dz_2^2) + \mathbb{X}^2 (dx_3^2 + dx_4^2) + \mathbb{X}^3 (dx_5^2 + dx_6^2)$$

$$A = A^1 \wedge dz_1 \wedge dz_2 + A^2 \wedge dz_3 \wedge dz_4 + A^3 \wedge dz_5 \wedge dz_6$$

Assume $\mathbb{X}^1 \mathbb{X}^2 \mathbb{X}^3 = 1$ (T^6 has constant volume)
Volume is in a hypermultiplet
and decouples

$$I_5 = \frac{1}{16\pi G_5} \int R \star 1 - G_{ij} F^i \wedge \bar{F}^j - \frac{1}{6} C_{ijk} \bar{F}^i \wedge \bar{F}^j \wedge A^k$$

$$G_{ij} = \frac{1}{2} \begin{pmatrix} (\mathbb{X}^1)^{-2} & & \\ & (\mathbb{X}^2)^{-2} & \\ & & (\mathbb{X}^3)^{-2} \end{pmatrix} \quad C_{ijk} = 1 \text{ for permutations of } (123) \\ 0 \text{ otherwise}$$

\Rightarrow Minimal 5D sugra + 2 vector supermultiplets

If we set $A^1 = A^2 = A^3$ we get minimal 5D sugra
 $\mathbb{X}^1 = \mathbb{X}^2 = \mathbb{X}^3 = \text{const}$

$$\int R \star 1 - 2 F \star \bar{F} - \frac{8}{3\sqrt{3}} F \star \bar{F} \star A$$

$$A \equiv \frac{\sqrt{3}}{2} A^1$$

In CY_3 we get $N=1$ sugra + more vectors (+ hypers which decouple) - things generalize.

(Often take minimal sugra for simplicity of illustration)

Charges : $Q_i = \frac{1}{16\pi G_5} \int_{S^3} *F$ or $Q_i = \frac{1}{16\pi G_5} \int_{S^2} (\mathbb{X}^i)^{-2} *F^i$

This corresponds to having $A^i \sim \frac{1}{r^2}$ in 11D This is
a membrane charge eg Q_1 is membrane along 012

$$\text{BPS: } M \geq |Q_1| + |Q_2| + |Q_3|$$

Minimal: $Q \neq 0$ membranes $\begin{matrix} 012 \\ 0 \quad 34 \\ 0 \quad 56 \end{matrix}$ $\begin{matrix} Q_1 \\ Q_2 \\ Q_3 \end{matrix}$

But we can also have magnetic charges giving dipoles

Saw

A_ϕ monopole charge

$$q_i = \frac{1}{2\pi} \int_{S^2} F^i$$

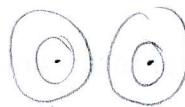
$$A_\phi = -\frac{1}{2}(1+x)$$

$x \in [-1, 1]$

It's a dipole since have opposite charges

Total magnetic charge is zero: an S^2 gives zero charge

Field is dipole-like



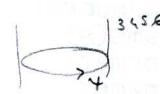
Analogue

Falls-off faster than monopole

In 11D These correspond To MS branes

eg $q^1 A_\phi \sim q^1 (1+x)$ is MS 0 3 4 5 6 7

q^2	0	1	2	5	6	7
q^3	0	1	2	3	4	7



In minimal sagra we'll have The Three of Them.

- Note: There's a black ring w/ only M5 dipoles, no M2 charges: Dipole black ring
 \Rightarrow continuous non-uniqueness: q is not conserved since t is contractible
 Conserved charges are M, J but need q to specify solution
 These are non-SUSY, since $M > \sum |Q_i| = 0$

- Will focus on SUSY rings. For simplicity will do minimal sagra

Solving Killing spinor eqs gives

$$ds^2 = -f^2 (dt + \omega)^2 + f^{-1} h_{mn} dx^m dx^n$$

h_{mn} : hyper-Kähler (will take flat)

f : scalar $f(x_m)$

ω : 1-form $\omega_m dx^m$

t-indep

$V = \frac{\partial}{\partial t}$ Timelike Killing constructed from Killing spinor

$$F = \frac{\sqrt{3}}{2} [d(f(dt + \omega)) - \Theta]$$

↳ electric ↳ magnetic

$$\Theta = -\frac{1}{3} f(d\omega + \star_4 d\omega)$$

These conditions are iff for SUSY

Then impose cons:

Bianchi: $d\Theta = 0$

Einstein then follows

Maxwell: $\nabla^2 f^{-1} = \Theta^2$

will take $\text{Lagrangian } dx^m dx^n = ds^2(E_4)$

Simple solutions: $\Theta=0 \quad \omega=0$: no dipole, no rotation: susy RN (Strominger-Vafa)

$$\nabla^2 f^{-1} = 0 : \quad f^{-1} = 1 + \frac{Q}{r^2} \quad (\text{harmonic function rules...})$$

$$A = \frac{\sqrt{3}}{2} f dt$$

- $\Theta=0, \omega \neq 0$: no dipole, rotating. BMFV

$$\text{again } \nabla^2 f^{-1} = 0 \quad f^{-1} = 1 + \frac{Q}{r^2}$$

$$\omega = \# \frac{J}{r^2} (\cos \theta dt + \sin \theta d\phi) \quad ds^2(E_4) = dr^2 + r^2(d\theta^2 + \sin^2 \theta d\phi^2)$$

$d\omega$ is anti-self-dual 2-form on \mathbb{R}^4

- $\Theta \neq 0, \omega \neq 0$: susy rotating black ring

more complicated since f^{-1} not harmonic & $d\omega$ not anti-self-dual

$$\text{Use ring coordinates for } E_4 : \quad ds^2 = \frac{R^2}{(x-\gamma)^2} \left(\frac{dt^2}{\gamma^2} + (y^2-1) dt^2 + (1-x^2) d\phi^2 + \frac{dx^2}{1-x^2} \right)$$

This will simplify enormously. The solution

is

$$\Theta = \frac{1}{3} q (dy \wedge dt + dx \wedge d\phi)$$

Recall Θ is related to dipole

we know that

This is a self-dual 2-form on the base, from our construction of the coordinates!

$$\text{so } d\Theta = 0$$

$$\text{Then } \nabla^2 f^{-1} = \Theta^2 = \frac{1}{2} q^2 \left(\frac{x-\gamma}{R} \right)^4$$

To solve this: first, The homogeneous eqn $\nabla^2 f^{-1} = 0$ is solved by $f^{-1} \propto x-y$ with δ -source at the ring

The complete eqn is solved in particular by

$$f^{-1} = -\frac{q^2}{4R^2} (x-\gamma)^2 \quad (\text{say, by trial and error})$$

$$f^{-1} = 1 + \frac{Q-q^2}{2R^2} (x-\gamma) - \frac{q^2}{4R^2} (x-\gamma)^2$$

The eq for ω is then also easily solved.

Requirements: absence of CTC's near source (positive semi-definite) " " Dirac-Misner singularities: ω globally defined

$$\omega_+ (x=\pm 1) = 0$$

$$\omega_+ (\gamma=-1) = 0$$

Horizon at $\gamma=-\infty$: $S^1 \times S^2$ (where $f \rightarrow 0$)

$x = \text{const}$

$$\text{Horizon geometry } ds_h^2 = L^2(Q, q, \gamma) dt^2 + \frac{q^2}{4} (d\theta^2 + \sin^2 \theta d\phi^2)$$

$$S^1 \times S^2$$

$$A_H = 2\pi L 4\pi \frac{q^2}{4} = 2\pi^2 L q^2$$

S^2 size controlled by dipoles

No CTC $\equiv L^2 \geq 0$

Near-horizon: $\text{AdS}_3 \times S^2$ locally

\downarrow radius $q/2$ \rightarrow controlled by dipoles, not by charges Q !

- Angular velocities vanish (as they must for a susy AF bh)

Remarks: - BMPV obtained as $R \rightarrow 0$. But area jumps discontinuously in limit.

- $J_4 > J_\phi$: always different parameters than BMPV: $J_4 = J_\phi$

- $R \rightarrow \infty$ limit: 5D busy black string with M_2, M_5, P charges: among the solutions obtained from the generating soln of Barcelini + Tringane

- Non-uniqueness: fix α_i, J_4, J_ϕ (M is fixed by α_i): 5 parameters

Solutions have 7 parameters: two dipoles remain unfixed

Double-continuous non-uniqueness

$$(01.c) \quad \text{(requiring maximal area fixes uniqueness in busy case, but not in neutral)}$$

$$+ (1/2 + v) \delta \tilde{\Sigma} + ((v + u)/\pi) \Psi = (P_0 + v)(\Psi - \tilde{\Psi}) = 0$$

- Multi-rings are possible

(Black rings in Taub-NUT space)

$$(11.c) \quad (v + u) \delta \tilde{\Sigma} + ((v + u)/\pi) \Psi = (P_0 + (v + u) \delta \frac{\pi}{2} + (v^2 + u^2)) \text{sol} \rightarrow \text{uniqueness in neutral case}$$

Note that we have considered a fixed value of $\tilde{\Sigma}$ - a priori it is not the intersection of a 5D black ring with the 4D spacetime. A generalization involving the relative motion of the two black holes is left for future work.

$$(21.c) \quad (v, u) \tilde{\Delta}_0 \tilde{\Lambda} \left((v^2 + u^2) \text{sol} \frac{1}{2} \right) (v, u)_1 \tilde{\Delta} \text{sol} \left[\dots \right] \frac{1}{\tilde{\Sigma}} = (v, u) \tilde{\Delta} \tilde{\Lambda}$$

$$(21.d) \quad (v, u) \tilde{\Delta}_0 \tilde{\Lambda} \left(\frac{1}{\tilde{\Sigma}} \text{sol} \frac{1}{2} \right) = (v, u)^+ \text{sol} \frac{1}{2} = (v, u)^+ x : |x| \leq 1$$

$$(21.e) \quad (v, u) \tilde{\Delta}_0 \tilde{\Lambda} \left((v^2 - 1)v + u \right) \frac{1}{\tilde{\Sigma}} = (v)^+ x : |x| \leq 1$$

$$\left(\frac{1}{\tilde{\Sigma}} \text{sol} \frac{1}{2} \right) = (v^2 - 1)v + u + \left(\frac{1}{\tilde{\Sigma}} \text{sol} \frac{1}{2} \right) = (v)^- x$$

Microscopics of black rings

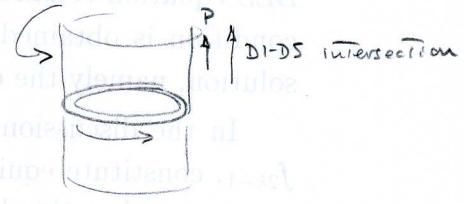
To identify micro, find a brane configuration w/ same charges.

Part focus on what charges: net charges q_i ?
dipole charges j_i ?

Both are possible, lead to alternative descriptions

Net charges q_1, q_2, q_3 : like BMPV
Can dualize $M2 \perp M2 \perp M2$ to $D1-D5-P$
Then q_i dualize to d_i , d_5 KKm

$$ds^2(DR) \times S^1 \times T^4$$



This was very successful for BMPV. Also, if only D1-D5 and KKm. This is not so straightforward for black ring.

Less successful for black ring.

The description is important since it contains both BMPV and ring as states of D1-D5-P system. This views black rings as superbranes.

If we focus on dipole charges. Then we can't include BMPV in the same description since it has none. But we'll be able to count entropy precisely.

↳ This means focusing on $M5 \perp M5 \perp M5 + P$ as momentum along intersection.

(2.8)
Recall $M5 + 1234 \dashv$ black rings as circular strings
 $M5 + 12 \quad 56 \dashv$
 $M5 + 3456 \dashv$
M2-charges correspond to fluxes on the worldvolume of the M5's

Use the same theory that describes $M5 \perp M5 \perp M5 + P$ for a straight intersection

$$M5 + 1234 \dashv \quad R \rightarrow \infty \text{ limit of black ring } (MSW)$$

$$\begin{array}{c} 1 \\ 2 \\ 3 \\ 4 \\ 5 \\ 6 \end{array}$$

MSW: smooth out intersection: single M5 wrapping a cycle outside CY_3

small $CY_3 \Rightarrow$ low energy dynamics is $(0,4)$ -easy $\begin{cases} \text{H1 model} \\ \text{intersection} \end{cases}$

(2.9) $c = n_1 n_2 n_3 (+\dots)$ \rightarrow numbers of M5:

$$V_6 \ll c : \text{sugra}$$

$$V_6 \gg c : \sigma\text{-model}$$

$$R \gg V_6 \text{ To get } \sigma\text{-model}$$

$$V_6 \gg 1 \text{ To neglect quantum corr.}$$

M2 charges: no fluxes

Turn them to zero for the moment

(2.10) $M5 + \text{momentum} \rightarrow J_+ = \hat{f}_0$ oscillator level

$$S_{\text{eff}} = m \sqrt{\frac{c}{6} \hat{f}_0} - m \sqrt{n_1 n_2 n_3 J_+} \quad \text{for straight string in 5D. (4D 4-ch bh)}$$

This is correct

Turn on M2 charges : They give rise to momentum zero modes. That contribute to total momentum \hat{q}_0 . So oscillator level (non-zero mode)

$$\hat{q}_0 = q_0 + \frac{1}{2} \left(\frac{N_1 N_2}{n_3} + \text{perm} \right) - \frac{1}{4 n_1 n_2 n_3} \sum_i^3 (N_i n_i)^2 + \frac{n_1 n_2 n_3}{4}$$

Setting $q_0 = -J_4$ in $S = 2\pi \sqrt{\frac{c}{6}} q_0$ gives perfect agreement with BH entropy

7-parameter match of the entropy !

(8) Shortcomings : doesn't explain why $Q_1 > q_2 q_3$
 R is fixed by equilibrium and not free.

~~Don't say anything about finite radius effects.~~

~~microscopic significance of non-uniqueness~~

(9)

\Rightarrow UV description : black rings are supertube-like, rotation has blown them into a tubular structure.

(10)

AT least shows that string Th has the required structure to accommodate different objects w/ the same conserved charges

(11)

Black rings seem to be a mixture between a BMPV bh and a supertube (but things are missing)

~~no no~~

- Non-vary versions exist, w/ 7 parameters, but would need 9-parameters to be able to go near-vary

- We may have all vary bhs of SD sigma. There are solutions w/ arbitrary profile

Have finite area but horizon is not regular
 $(\text{metric is continuous, curvatures are finite. But freely falling observers experience infinite tidal forces. Components of Riemann in parallel-prop basis diverge along geodesics})$

Thus says that horizon must be S^3 (or quotient) homogeneously squashed - like BMPV

$$S^1 \times S^2$$

T^3 (marginal, unlikely)

~~the only consistent manifold for which one can demand that two central charges vanish is T^3 (or integer side distinguishable at an equal [0] distance along each coordinate)~~

~~but it is naturally oriented, i.e. of bosonic origin, not supersymmetric shift~~

~~(0002 technicalities)~~