

What I will assume in these 4 sessions
is that everyone is familiar with the
basic ideas of AdS/CFT.

(Basically that some backgrounds of String theory
are dual to [capture the same information] as some
(gauge) field theories.)

The point of these meetings is to discuss
how one can learn about (particularly interesting)
QFT using String theory.

This topic is VERY extensive.

So, in these sessions I will focus on very
specific aspects of this duality
(but quite interesting ones!)

The problem of Physics behind these lectures is an old one:

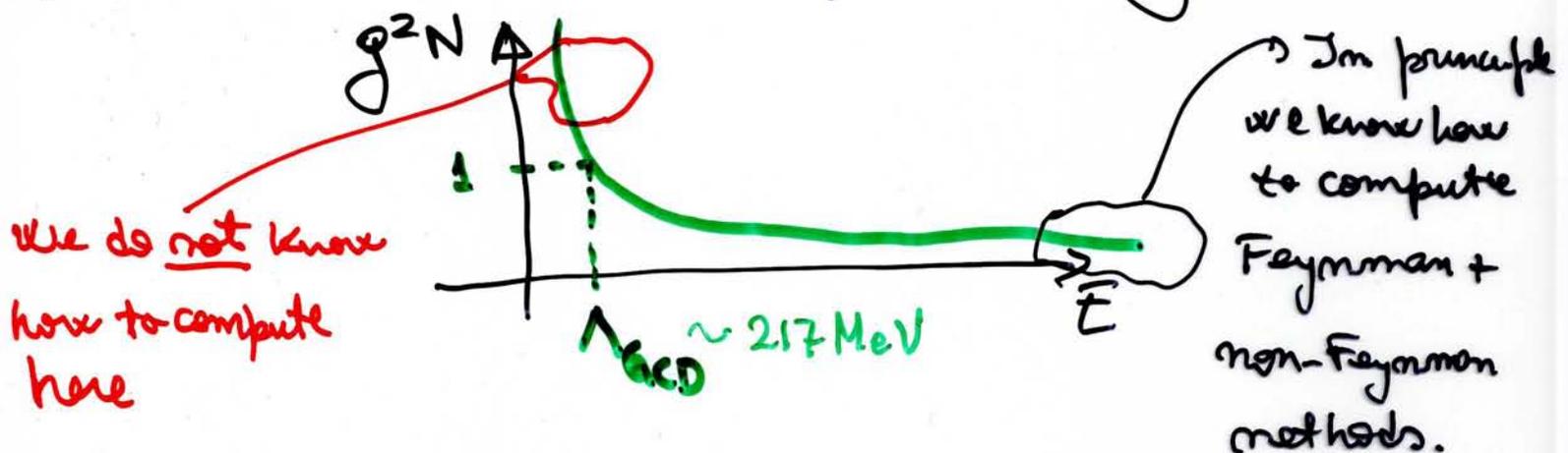
[How to compute non-perturbative aspects of
any given QFT (focus on gauge theories).]

The main Physics example to have in mind is "Nuclear Physics" (the "chemistry" of QCD)

$$\mathcal{L} = \bar{\psi} \left\{ -\frac{1}{4} F_{\mu\nu}^2 + i \not{D} \psi - m \psi \right\} \quad \left(\text{group is } \overset{SU(N)}{SU(3)} \right)$$

(and $k: 1, \dots, 6$)

We know that in this field theory



of course, this problem was not ignored and since 1973, people have developed methods to deal with the regime in red above

These methods range from very good ones to very "effective" ones

- Lattice : the "correct way". Limitations that will be overcome with more computing power and new algorithms (prevents on analytic understanding)
- GFT elaborations : very nice, not so successful

• Effective field theory : works by definition, but postpones the understanding of the connection between the effective Lagrangian and α_{QCD}

In these lectures I will use a method inspired on String theory, that is the so called Maldacena Conjecture (or AdS/CFT and its extensions)

I will focus on uses of this conjecture to learn about phenomenologically interesting field theories

This topic is so extensive that I will focus a little more. The basic problem I would like to address in these lectures is

"How to learn about non perturbative aspects of $\mathcal{N}=1$ SUSY gauge field theories in $3+1$ dimensions"

(the small # of SUSY \Rightarrow good caricature of QCD among other $3+1$ d interesting field theories).

Let me start by mentioning what I will not talk about. (even in this narrow topic there are different and very successful approaches.)

I will not discuss:

Models starting with $\mathcal{N}=4 \Rightarrow \mathcal{N}=1$

[ginsparg, Peticola, Pomati
Zaffaroni
Polchinski + Strassler
Freedman, Gubser, Penedes
Warner + ...]

Models based on cascading quivers

AdS/QCD [Erickson, Kotz, Son, Stephanou
Da Rold + Pomati
...]

[Klebanov + Tseytlin
Nekrasov + Strassler +
Witten + Dymarsky +
Seiberg + Murugan + Benini]

Models based on Black holes \Rightarrow Heavy Ion Physics [Policastro + Son
+ Strassler + ...]

But I will be quite pleased to discuss them outside the classes 14

So, what will I (try to) talk about?

• Models arising wrapped branes duals to
 $\mathcal{N}=1$ SYM and $\mathcal{N}=1$ SQCD

and discuss in detail how can we learn about
non perturbative aspects of these field
theories looking at (computing with) String theory.

The topic may look narrow, but is actually
quite extended, so, I will pick some particular
aspects.

Of course, outside the lectures, we can discuss
about everything I will not mention or other
things.

So, to summarize, the idea is to learn about
 $\mathcal{N}=1$ SYM and $\mathcal{N}=1$ SQCD from string theory
This is the topic of what follows

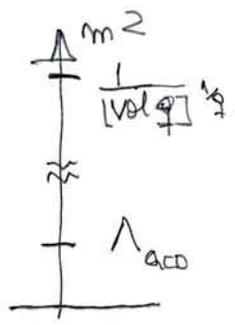
Wrapped branes

The idea is to construct a dual to a phenomenologically interesting field theory + some UV completion (that may have no relation with the theory itself)

One possibility is to take a D_p brane with $p > 3$ and "carefully" wrap it over a q manifold (if the q -manifold is "small" then we have an effective $p+1-q$ field theory. Two things should be clarified

- ① "carefully": so that this wrapping preserves some SUSY. (for example) BPS eqs
- ② "small": compared with the mass dynamical scale of the theory

the idea is:



\leadsto so that the dynamics are well understood (non perturbative field theory is well separated from the UV completion)

How to do this actually? let us see different examples

D_4 on S^1 \leadsto Witten 1998 \leadsto ~~the~~ wrapping is done with different

boundary conditions for fermions and bosons. \rightarrow a solution can be written that is non-SUSY.

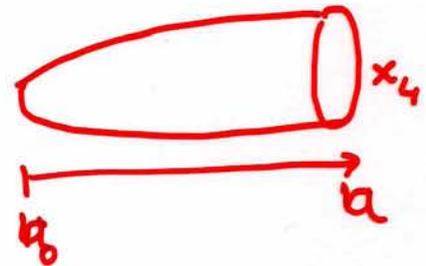
D4 brane wrapping S^1 with anti-periodic boundary conditions.
(Witten 1998).

$$ds^2 = \left(\frac{u}{R}\right)^{3/2} \left[-dt^2 + d\vec{x}_3^2 + f(u) dx_4^2 \right] + \left(\frac{R}{u}\right)^{3/2} \left[\frac{du^2}{f(u)} + u^2 d\Omega_4^2 \right]$$

$$F_4 = \frac{2\pi N_c}{\text{Vol } S^4} \cdot \omega_{S^4} \quad ; \quad e^{2\phi} = \int ds \left(\frac{u}{R}\right)^{3/2}$$

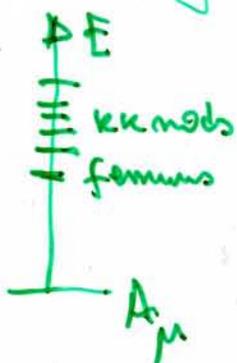
$$R = \pi \int ds N_c \alpha'^{3/2} \quad ; \quad f(u) = 1 - \left(\frac{u_0}{u}\right)^3$$

$\leadsto x_4$, Ad have a cigar topology

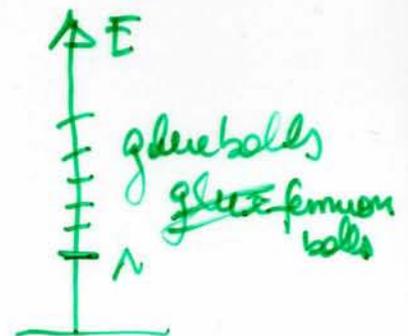


This background may be thought of as an analytic continuation from a D4 brane black hole

the weakly coupled spectrum



strongly coupled



In this case a particular ~~flat~~ background exists

$$dS^2 = \left(\frac{u}{R}\right)^{3/2} \left[-dt^2 + d\vec{x}_3^2 + f(u) dx_4^2 \right] + \left(\frac{R}{u}\right)^{3/2} \left[\frac{du^2}{f(u)} + u^2 d\Omega_4^2 \right]$$

$$F_4 = 2\pi N \text{ vol } \Omega_4$$

$$e^{2\phi} = g_s^2 \left(\frac{u}{R}\right)^{3/2}$$

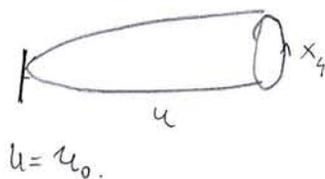
$$R = \pi g_s N_c (\alpha')^{3/2} \quad ; \quad f(u) = 1 - \left(\frac{u_0}{u}\right)^3$$

This can be thought of as taking the D4 brane hole and

continuum $x_4 \rightarrow it$
 $t \rightarrow i x_4$



The sub-manifold $x_4 - u \rightarrow \text{cylinder}$



$$g_s^2 = (2\pi)^2 g_s (\alpha')^{1/2}$$

$$g_4^2 = \frac{g_s^2}{2\pi R}$$

$$M_{\text{glueballs}} = \frac{1}{R}$$

$$T_{\text{QCD}} = \frac{1}{2\pi\alpha'} \left. \sqrt{\frac{g_{xx} g_{tt}}{g_{xx} g_{tt}}} \right|_{u=u_0} = \frac{1}{2\pi\alpha'} \left(\frac{u_0}{R}\right)^{3/2}$$

When wrapped D6 branes on a S^3 one obtains a metric that when lifted to 11 dimensional supergravity

reads:

$$dS^2 = dx_{1,3}^2 + \alpha^2 d\Omega^2 + \gamma^2 \tilde{\omega}_a^2 + \beta^2 (\omega^a - \frac{1}{2} \tilde{\omega}^a)^2$$

$$\alpha^2 = \frac{1}{1 - a^3/R^3} ; \quad \beta^2 = \frac{R^2}{9} \left(1 - \frac{a^3}{R^3}\right) ; \quad \gamma^2 = \frac{R^2}{12}$$

$$\tilde{\omega}_1 = \cos \tilde{\psi} d\tilde{\theta} + \sin \tilde{\psi} \sin \tilde{\theta} d\tilde{\varphi}$$

$$\omega_1 = \cos \psi d\theta + \sin \psi \sin \theta d\varphi$$

$$\tilde{\omega}_2 = -\sin \tilde{\psi} d\tilde{\theta} + \cos \tilde{\psi} \sin \tilde{\theta} d\tilde{\varphi}$$

$$\omega_2 = -\sin \psi d\theta + \cos \psi \sin \theta d\varphi$$

$$\tilde{\omega}_3 = d\tilde{\psi} + \cos \tilde{\theta} d\tilde{\varphi}$$

$$\omega_3 = d\psi + \cos \theta d\varphi$$

In IIA this looks like a background with

$$g_{\mu\nu}, \Phi, F_2$$

$$dS^2 = e^{4/3\Phi} (dx_{10} + A)^2 + e^{-2/3\Phi} dS_{10}^2$$

One can see that (a few) aspects of $N=1$ SYM are captured by the background above

Another possibility is to wrap a D6 brane on a 3 cycle

$\mathbb{S}^1, \phi, \mathbb{F}_2$. When lifted to 11 dimensions these solutions

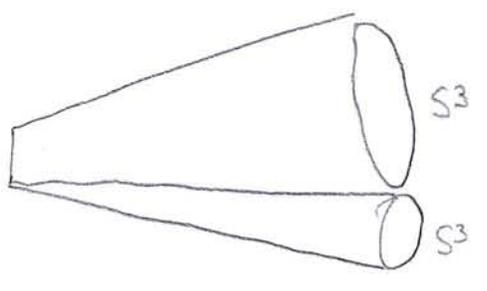
look like M theory on a G_2 manifold. The first solution I know of

was written by Bryant + Solomon and used by Atiyah, Mubarek, Vafa

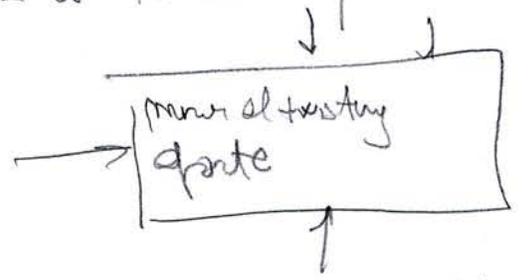
$$ds^2 = dx_{1,3}^2 + \alpha^2 dr^2 + \gamma^2 \tilde{\omega}_a^2 + \beta^2 (\omega_a - \frac{1}{2} \tilde{\omega}_a)^2$$

$$\alpha^2 = \frac{1}{1 - \frac{a^3}{r^3}}, \quad \beta^2 = \frac{r^2}{9} \left(1 - \frac{a^3}{r^3}\right), \quad \gamma^2 = \frac{r^2}{12}$$

when reduced to type IIA

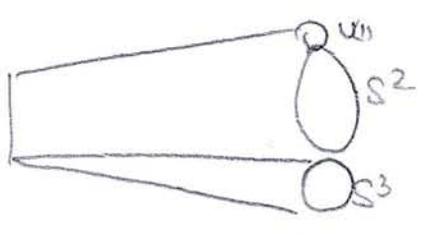


one must pick an invariant direction to do the reduction, the size of this direction is the value of the dilaton



$$ds_{II}^2 = e^{\frac{4}{3}\phi} (dx_{1,1} + A)^2 + e^{-\frac{2}{3}\phi} ds_{S^2}^2$$

So, a problem of interest in 2002 was to find G_2 manifolds with a stabilized U(1), that is:



this problem was solved by

BGGG

02, 03

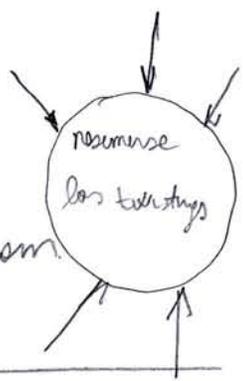
CGPL

Another possibility that was explored was to wrap a D5 brane on a small 2-cycle

There are two different ways of doing this wrapping.

① Preserving $N=1$ in 3+1 \longrightarrow I will focus on this.

② Preserving $N=2$ in 3+1
 • Bigazzi, Cribane, Zaffaroni
 • Gauntlett, Kim, Moriella, Waldram.



We now focus on $N=1$ case. Let us first see the solution for a D5 on S^2

It is good to gain some intuition by starting with a flat D5 brane

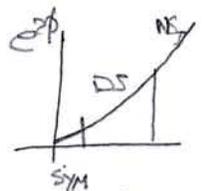
$$ds^2_{\text{string}} = \alpha' e^{\frac{\phi}{2}} \left[\frac{4dx^2_{1,5}}{\sqrt{g^2_{YM} N}} + \sqrt{g^2_{YM} N} \frac{du^2}{u} + \sqrt{g^2_{YM} N} u d\Omega_3 \right]$$

we took the decoupling limit

$$u = \frac{r}{\alpha'} = \text{fixed}$$

$$\alpha' \rightarrow 0$$

$$e^{\phi} = \frac{g^2_{YM} u}{(2\pi)^2 \sqrt{g^2_{YM} N}}$$



$$g^2_{YM} = g^2_{DS} \alpha' = \text{fixed}$$

$$\vec{F}_3 = -\frac{N}{4} \omega_1 \wedge \omega_2 \wedge \omega_3$$

$$\left. \begin{aligned} \omega_1 &= \cos \psi d\tilde{\theta} + \sin \psi \sin \tilde{\theta} d\tilde{\varphi} \\ \omega_2 &= -\sin \psi d\tilde{\theta} + \cos \psi \sin \tilde{\theta} d\tilde{\varphi} \\ \omega_3 &= d\psi + \cos \tilde{\theta} d\tilde{\varphi} \end{aligned} \right\}$$

$$d\Omega_3 = \frac{1}{4} (\omega_1^2 + \omega_2^2 + \omega_3^2)$$

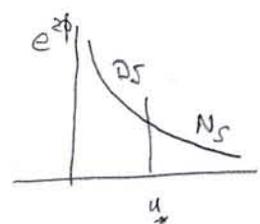
Nota

En realidad esta solución se debe pensar como el IR de una NS_5 brane

$$ds^2_{\text{string}} = dx^2_{1,5} + \tilde{\alpha}' (N \frac{du^2}{u^2} + N d\Omega_3)$$

$$e^{\phi} = \frac{N}{g^2_{YM}} \left(\frac{1}{u^2} \right)$$

$$\tilde{\alpha}' = g^2_{DS} \alpha' = \text{fixed}$$



$$\vec{H}_3 = -\frac{N}{4} \omega_1 \wedge \omega_2 \wedge \omega_3$$

one can check that the solution above preserves $U(1) \times SO(3)$

$$\delta\psi_\mu = 0$$

$$\delta\lambda = 0$$

$\leadsto \mathcal{E} =$

$$\prod_{x_0, x_1, x_2, x_3} \prod_{\partial\varphi} \eta = +\eta$$

$$\prod_{\partial\varphi} \eta = \prod_{\partial\tilde{\varphi}} \eta$$

$$\eta = i\eta^*$$

$$\mathcal{E} = e^{\frac{\phi}{8}} \sqrt{e^{\frac{\alpha}{2}} \prod_{\partial\tilde{\varphi}} \eta}$$

$$\cos \alpha = \coth 2r - \frac{2r}{\sinh^2(2r)}$$

$$\prod_{x_0, x_1, x_2, x_3} \left(\cos \alpha \prod_{12} + \sin \alpha \prod_{1\frac{1}{2}} \right) \mathcal{E} = \mathcal{E}$$

Let me comment something. if one considers the leading asymptotics

$$e^{2\phi} \sim e^{2\phi_0} \frac{e^{2r}}{4\sqrt{r}}$$

$$e^{2h} \sim r$$

$$a \sim 0$$

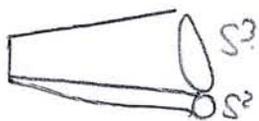
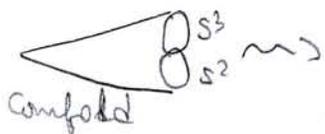
$$ds^2 = e^{\frac{\phi}{2}} \left[dx_{1,3}^2 + \alpha' g_{jk} N \left\{ dn^2 + r(d\theta^2 + \sin^2\theta d\varphi^2) + \frac{1}{4}(d\tilde{\theta}^2 + \sin^2\tilde{\theta} d\tilde{\varphi}^2) + \frac{1}{4}(d\psi + \cos\theta d\varphi + \cos\tilde{\theta} d\tilde{\varphi})^2 \right\} \right]$$

$$F_3 = -\frac{N}{4} \left[-\omega_1 \wedge \omega_2 \wedge \omega_3 + \sin\theta d\theta \wedge d\varphi \wedge \omega_3 \right]$$

this is also a solution. It is singular at $r \rightarrow 0$ $R \sim \frac{1}{r}$ (badly singular)

So, basically what the function $a(r)$ is doing is "resolving" the

singular behaviour by introducing the deformed conifold



So, the solution on which the DS wraps S^2 has to have a "topology" of the form.

$$ds^2 \sim e^{f_1} dx_{1,3}^2 + e^{f_2} d\Omega_2 + e^{f_3} d\Omega_3$$

but there will be some "complications." (to preserve SUSY)

In reality it is (coordinates $\{x^4, r, \theta, \varphi, \bar{\theta}, \bar{\varphi}, \psi\}$)

$$ds^2 = e^{\frac{f}{2}} \left[dx_{1,3}^2 + \alpha' \frac{1}{f_5} N \left\{ dr^2 + e^{2h} (d\theta^2 + \sin^2 \theta d\varphi^2) + \frac{1}{4} (\tilde{\omega}_1 + a d\theta)^2 + \frac{1}{4} (\tilde{\omega}_2 - a \sin \theta d\varphi)^2 + \frac{1}{4} (\tilde{\omega}_3 + \cos \theta d\varphi)^2 \right\} \right]$$

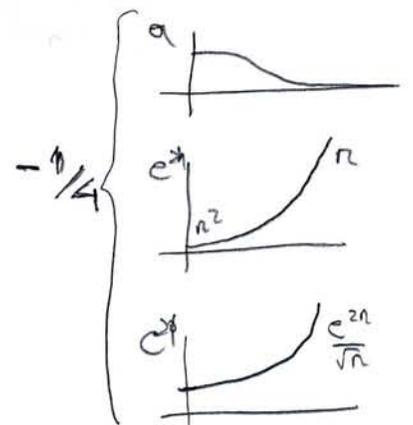
$$F_3 = + \frac{N_c}{4} \left[(\tilde{\omega}_1 + a d\theta) \wedge (\tilde{\omega}_2 - a \sin \theta d\varphi) \wedge (\tilde{\omega}_3 + \cos \theta d\varphi) + a' dr \wedge (\tilde{\omega}_1 \wedge d\theta + \sin \theta d\varphi \wedge \tilde{\omega}_2) + (1-a^2) \sin \theta d\theta \wedge d\varphi \wedge (\tilde{\omega}_3 + \cos \theta d\varphi) \right]$$

~~ϕ~~ $\phi(r)$

where the functions

$$a(r) = \frac{2r}{\sinh 2r} ; e^{2h} = r \coth 2r - \frac{r^2}{\sinh^2(2r)}$$

$$e^{2\phi} = e^{2\phi_0} \frac{\sinh 2r}{2e^{hr}}$$



The geometry is such that when $r \rightarrow 0$

$$ds^2 \sim dx_{1,3}^2 + \underbrace{dr^2 + r^2 d\Omega^2 + \frac{1}{4} (\tilde{\omega}_1 + d\theta)^2 + \frac{1}{4} (\tilde{\omega}_2 - \sin \theta d\varphi)^2 + \frac{1}{4} (\tilde{\omega}_3 + \cos \theta d\varphi)^2}_{\text{Deformed conifold}}$$

more geometry
Pop Tally

Deformed conifold

More general solutions

hep-th/0602027

Cosans, Núñez, Porets

one can propose a metric, F_3 ; ϕ

$$ds^2 = e^{2f(r)} \left\{ dx_{1,3}^2 + dr^2 + e^{2h} (d\sigma^2 + \sin^2\sigma d\varphi^2) + \right.$$

$$\left. \frac{e^{2g}}{4} \left[(\tilde{\omega}_1 + a d\sigma)^2 + (\tilde{\omega}_2 - a \sin\sigma d\varphi)^2 \right] + \frac{e^{2k}}{4} \left[\omega_3 + \cos\sigma d\varphi \right]^2 \right\};$$

$$F_3 = \frac{N_c}{4} \left[-(\tilde{\omega}_1 + b d\sigma) \wedge (\tilde{\omega}_2 - b \sin\sigma d\varphi) \wedge (\tilde{\omega}_3 + \cos\sigma d\varphi) + \right.$$

$$\left. b' d\sigma \wedge (\tilde{\omega}_1 \wedge d\sigma + \sin\sigma d\varphi \wedge \tilde{\omega}_2) + (1-b^2) \sin\sigma d\sigma \wedge d\varphi \wedge \tilde{\omega}_3 \right]$$

$\phi(r)$

write the BPS eqs (check that they solve the Einstein eqs) and find

$$b(r) = \frac{2r}{\sin 2\sigma}; \quad \phi(r) = f(r) \quad \text{and a one parameter}$$

family of solutions arise.

$$\begin{cases} a \sim 1 + \mu r^2 \\ e^{2k} \sim N_c \left(\frac{4}{6+3\mu} \right)^2 \sim N_c \frac{(20+36\mu+9\mu^2)}{15(2+\mu)} r^2 \end{cases}$$

$$\mu \in \left[-2, -\frac{2}{3} \right]$$

$$e^{2g} \sim N_c \frac{4}{6+3\mu}; \quad e^{2\phi} \sim e^{2f} (1 + \frac{4\mu}{6+3\mu})$$

$$e^{2h} \sim N_c \frac{4 r^2}{6+3\mu}$$

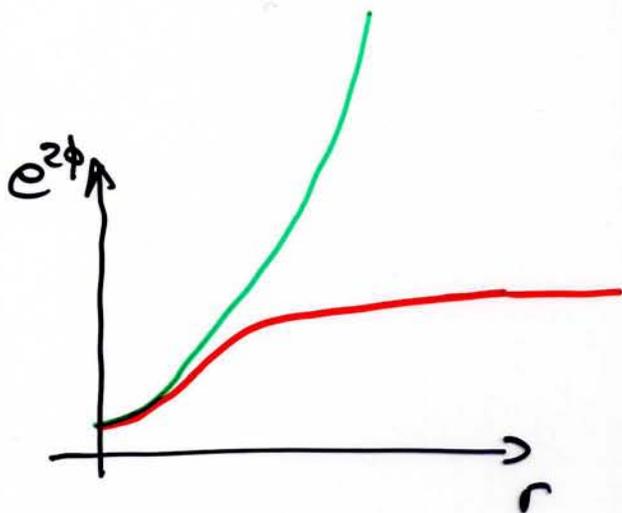
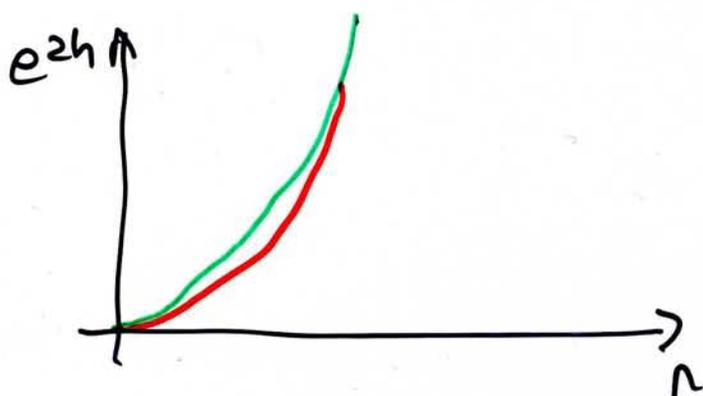
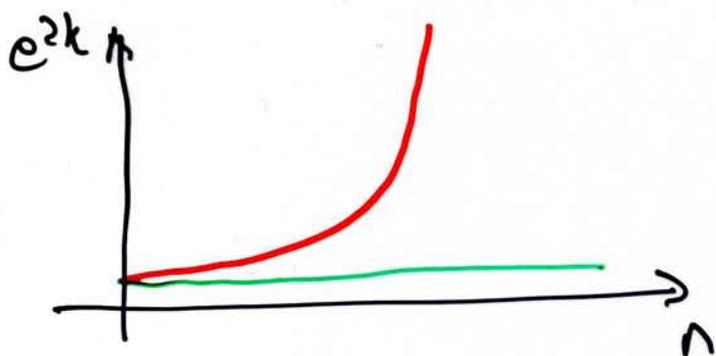
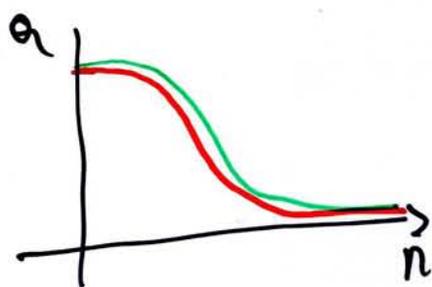
and plotting these functions

Plot of the deformed functions

$$\mu \in [-2, -2/3]$$

exact solution

$$\left[\begin{array}{l} \text{Mink}_3 \times \text{deformfold} \\ N_c = 0 \end{array} \right]$$



- exact solution $\mu = -2/3$
- new deformed solution

Notice the stabilization of the dilaton!

These new solutions present μ -dependent gauge theory dual quantities \rightsquigarrow quite interesting structure

See hep-th/0602027 (Section 3)

There are many checks that this solution captures very nicely non-perturbative effects of $N=1$ SYM:

Confinement of quarks
Screening of monopoles

Maldacena, Nunez

$U(1)_R$ Symmetry breaking / chiral symmetry breaking
Extension to S & CD

Casas, Nunez, Paredes

Hortnall
Portugues
Ginsparg

Instantons Maldacena Nunez, ...

Dipole deformation and dynamics of KK modes

Ginsparg Nunez

Gluon condensate Petruni, Zaffaroni; Ahnongy Leeyu Sonnenschein

Domain walls Comana, Marletti

Beta function di Vecchia, Lanza, Marletti; Bertolini-Marletti, ...

Correlators and holography Berg, Horack, Miick

Strings tensions Herzog Klebanov, Hortnall Portugues

Domain walls Maldacena, Nunez, Sonnenschein Leeyu

Finite temperature; Viscosity

Son, Starinets Koutan
Buchel Liu

Addition of flavors; S & CD. quenched, mesonic dynamics

Nunez Paredes Pando

glueballs; non susy deformations. Casas, Nunez; Pons Talavera

Baryonic vortex Sonnenschein, Leeyu; Pando, Aron Ahnongy Leeyu Sonnenschein

Veneziano-Yankelawicz - Superpotential Miick Evers Petruni Zaffaroni

Non-commutative version Mateos, Pons, Talavera

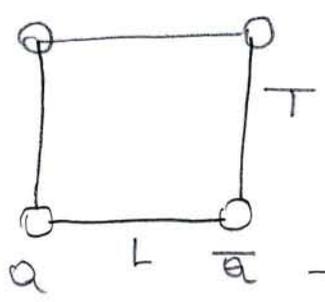
Bifurcation of flux tubes Sonnenschein, Leeyu

Tests of MN (D5 or S²) (All these tests are not influenced by the KK modes)

The first thing we should do is to see if the dual field theory confines. For this one would like to compute the

$\langle W, \text{Wilson loop} \rangle$

$$\langle W \rangle = \langle e^{i \oint A} \rangle$$



\leadsto action = in euclidean $S = E \cdot T$

\rightarrow ~~not~~ non dynamical quarks

if $E \sim \sigma L + \dots$
 \downarrow
 confinement ($\forall L$)

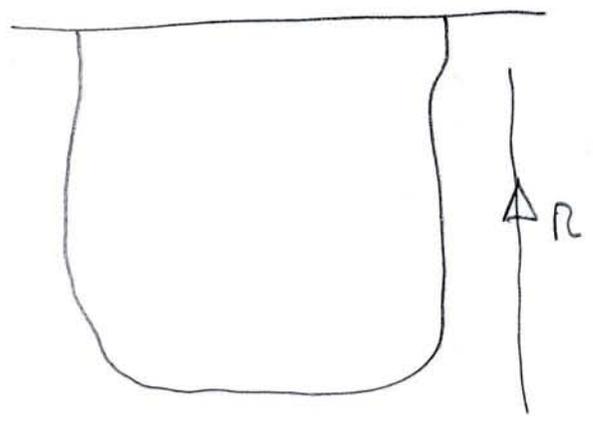
how to do this?

Maldacena
 Rey + Yee

\leadsto prescription

$$\langle W \rangle = e^{-S_{NG}}$$

for a string hanging from a D3 at ∞ (mass $\rightarrow \infty$ of the Q, \bar{Q})



$$\begin{aligned} t &= \tau & \tau \in [0, \infty] \\ x &= \sigma & \sigma \in [0, L] \\ R &= R(\sigma) \end{aligned}$$

background

Fortunately the Wilson loop was analyzed for a general background by Kuhn-Schreiber-Sonnenschein (1998) who proved a theorem

$$ds^2 = -g_{tt} dt^2 + g_{xx} dx^2 + g_{rr} dr^2 + \text{angular part}$$

$$ds_{\text{ind}}^2 = -g_{tt} d\tau^2 + (g_{xx} + g_{rr} r^2) d\sigma^2$$

$$S_{\text{NG}} = \frac{1}{2\pi\alpha'} \int d\sigma d\tau \sqrt{g_{tt} (g_{xx} + g_{rr} r^2)} = \frac{T}{2\pi\alpha'} \int d\sigma \sqrt{F^2 + G^2 r^2}$$

→ eq of motion $\frac{d}{d\sigma} \frac{\partial \mathcal{L}}{\partial \dot{r}} = \frac{\partial \mathcal{L}}{\partial r}$ → complicated but use Energy = $p \dot{q} - L$ is

conserved $p = \frac{G^2 r^2}{[F^2 + G^2 r^2]^{1/2}} \rightarrow \boxed{E = \frac{-F^2}{\sqrt{F^2 + G^2 r^2}} = -C}$ → \downarrow $F(r_0)$ → turning point

$$\frac{G}{F} \frac{dr}{\sqrt{F^2 - F_0^2}} = d\sigma$$

$$L = \int_0^L d\sigma = \int_{r_0}^{\infty} \frac{G}{F} \frac{dr}{\sqrt{F^2 - F_0^2}}$$

↖ insular class

$$E = F_0 L + 2 \int_{r_0}^{\infty} \frac{G}{F} (\sqrt{F^2 - F_0^2} - F) dr - 2 \int_0^{r_0} G dr$$

where the infinite mass of the quark was subtracted

$$m_q = \int_0^{\infty} G dr = 0$$

So, they proved that

• $F(r)$ is analytic near the IR

$$F = F(0) + a_k r^k$$

$k > 0$
 $a_k > 0$

• G is smooth $r \in [0, \infty)$

$$g \sim b_j r^j$$

$j > -1$

• $F, G > 0$ on $(0, \infty)$

• $F' > 0$ on $[0, \infty)$

$$\int \frac{G}{F^2} < \infty$$

$$\Rightarrow E = F(0) \cdot L + \begin{cases} \mathcal{O} \left[\log L e^{-\alpha L} \right] \\ \mathcal{O} \left[2\gamma - d L^{-n} \right] \end{cases}$$

ms if $F(0) \neq 0$ ms constant

if $F(0) = 0$ ms $E = -\frac{1}{L^\alpha}$

Kuro, Schreiber and Sonnenschein showed that

$F(x)$ is analytic

$$F = F(0) + a_k x^k$$

$$k > 0$$

$$a_k > 0$$

G is smooth

$$g \sim \frac{b}{x^{\alpha}}$$

$$\alpha > 1$$

$$F, G > 0 \text{ in } [0, \infty]$$

$$F' > 0 \text{ in } [0, \infty]$$

$$\int \frac{G}{F^2} < \infty$$

$$E = F(0) L + \begin{cases} \mathcal{O}\left[\frac{1}{g} L^{\alpha} e^{-\alpha L}\right] \\ \mathcal{O}\left[2\frac{1}{g} - \alpha L^{-\alpha}\right] \end{cases}$$

\leadsto if $F(0) \neq 0$ no component

$$\text{if } F(0) = 0 \leadsto E \sim \frac{1}{L^{\alpha}}$$

In our background

all the hypotheses are satisfied

$$F^2 = e^{\phi} = -g_{tt} + g_{xx}$$
$$G^2 = e^{-\phi} = -g_{tt} + g_{rr}$$

$$E_{\alpha\bar{\alpha}} = e^{2\phi_0} \cdot L + \text{corrections} \quad \leadsto \text{Wilson's confinement}$$

$$L = \int_0^{\infty} \frac{dr}{\sqrt{e^{\phi} - e^{\phi_0}}} \rightarrow \infty \quad \leadsto \text{the four is separated without problem}$$

usual ideas about confinement \leadsto quarks are ~~confined~~ ^{confined}
 \downarrow
monopoles are screened

how to compute screening of monopoles \leadsto compute the
t'Hooft loop \leadsto Wilson loop for magnetic charges.

In this case the "string" is simulated by a D3 brane
wrapping the cycle $\theta = \bar{\theta}$, $\varphi = 2\pi - \bar{\varphi}$, $\psi = \pi$.

~~Trapped~~

Screening of monopoles

The t' Hoft loop is computed using a D3 brane wrapped on a particular 2 cycle

The effective string computes the magnetic version of the Wilson loop.

$$\left. \begin{aligned} \theta = \bar{\theta} \quad , \quad \varphi = 2\pi - \bar{\varphi} \quad , \quad \psi = \pi \\ t = \tau \quad ; \quad x = \sigma \quad , \quad r = r(\sigma) \end{aligned} \right\} \text{ is the configuration for the D3 brane}$$

$$ds_{ind}^2 = e^\phi \left[-dt^2 + d\sigma^2 + r^2 d\tau^2 + \left(e^{2h} + \frac{(a-1)^2}{4} \right) \left(\frac{d\theta^2}{\sin^2 \theta} + dp^2 \right) \right]$$

$$\sqrt{\det g_{ind}} = e^{-\phi} \int dt d\sigma \cdot \left(\int d\theta dp \sin \theta \left[e^{2h} + \frac{(a-1)^2}{4} \right] \right) \sqrt{g_{tt} g_{\sigma\sigma} - g_{t\tau} g_{\tau\sigma} r^2}$$

\Rightarrow effective string

$$S = \underbrace{\left[e^{-\phi} e^{a\phi} \left(e^{2h} + \frac{(a-1)^2}{4} \right) 4\pi \right]}_{T_{eff} = 0} \cdot \int d\tau d\sigma \sqrt{g_{tt} g_{\sigma\sigma} - g_{t\tau} g_{\tau\sigma} r^2}$$

\Rightarrow so the monopoles are free \Rightarrow Screening of monopoles

The beta function (Di Vecchia-Louis-Morlotte, Bertolini-Marlotte)

one possible way of presenting this

The idea is to read the gauge coupling from a "brane probe"

For this, we wrap a D5 probe on a particular cycle defined by

$$\begin{aligned} \theta &= \tilde{\theta} & x^M & \text{fixed} \\ \varphi &= 2\pi - \tilde{\varphi} \\ \psi &= (2m+1)\pi \end{aligned}$$

How to define the gauge coupling?

we wrap a brane on ND5 sheet

$$g_{YM_6}^2 = g_s \alpha' N \cdot \left[\frac{g^2}{\text{Vol } \Sigma_6} \right] = L^2$$

$$g_{YM_4}^2 = \frac{g_{YM_6}^2}{\text{Vol } \Sigma_2} \quad ?$$

The idea is that the action on this probe

$$S = -T_5 \int d^5x e^{-\phi} \sqrt{\det(g + F^2)} + T_5 \int C_2 \wedge F_2 \wedge F_2$$

~~the idea is that expanding this around the background~~ the idea is that expanding this around the background

~~$$S = -T_5 \int d^5x e^{-\phi} \sqrt{\det(g + F^2)} + T_5 \int C_2 \wedge F_2 \wedge F_2$$~~

$$S \sim \underbrace{O \int d^4x F_{\mu\nu}^2}_{\sim \frac{1}{g^2}} + \underbrace{O \int d^4x F \wedge F}_{\sim \theta}$$

what each coefficient is?

Compute $\int \alpha_{ab} =$

$$\begin{aligned} \omega_1|_{\text{cycle}} &= \sin\psi d\theta + \sin\psi \sin\theta d\varphi \rightarrow (-1)^\theta d\theta \\ \omega_2|_{\text{cycle}} &= -\sin\psi d\theta + \cos\psi \sin\theta d\varphi \rightarrow +1 \sin\theta d\varphi \\ \omega_3|_{\text{cycle}} &= d\psi + \cos\theta d\varphi - \cos\theta d\varphi \rightarrow 0 \end{aligned}$$

$$ds_{\text{ind}}^2 = e^\phi \left\{ dx_{1/3}^2 + \alpha' g_s N \left[e^{2h} + \frac{(a-1)^2}{4} \right] (d\theta^2 + \sin^2\theta d\varphi^2) \right\}$$

	t	x ₁	x ₂	x ₃	θ	φ
t	-e ^ϕ	F _{tx₁}	F _{tx₂}	F _{tx₃}		
x ₁	-F _{tx₁}	e ^ϕ	F _{x₁x₂}	F _{x₁x₃}		
x ₂	-F _{tx₂}	F _{x₂x₁}	e ^ϕ	F _{x₂x₃}		
x ₃	-F _{tx₃}	F _{x₃x₁}	F _{x₃x₂}	e ^ϕ		
θ						
φ						

$$\Rightarrow \left[\det g_{ab} + 2\pi\alpha' F_{ab} \right] = \alpha' g_s N \left(e^{2h} + \frac{(a-1)^2}{4} \right) \sin\theta \cdot e^{3\phi} \left[\frac{1}{\alpha' g_s} \sqrt{-\det g_{\mu\nu}} e^{-2\phi} F_{\mu\nu}^2 \right]$$

So from here $e^{\int \sqrt{det(g_{\mu\nu})}}$

$$S_{D5} = - \left(\frac{1}{4 g_{YM}^2} \cdot \alpha'^3 N \left(e^{2h} + \frac{(a-1)^2}{4} \right) \int d\Omega_2 \sin^2 \theta \right) (2\pi \alpha')^2 \int d^4 x \frac{F^2}{\mu^2}$$

flat space

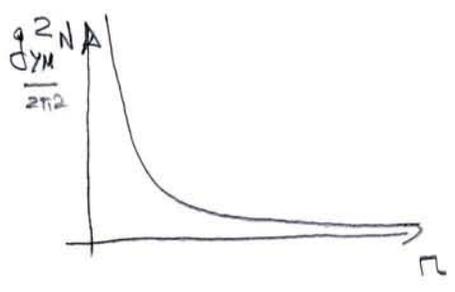
$$T_{D5} = \frac{1}{(2\pi)^5 g_{\alpha'^3}}$$

$$\frac{1}{(2\pi)^5 g_{\alpha'^3}} \cdot \alpha'^3 N \cdot 4\pi^2 \cdot 4\pi = \frac{16\pi^3 N}{(2\pi)^5}$$

$$\frac{1}{2\pi^2}$$

$$\frac{2 T^2}{g_{YM}^2 N} = \left[e^{2h} + \frac{(a-1)^2}{4} \right]$$

$$\left[e^{2h} + \frac{(a-1)^2}{4} \right] = r \tanh r$$



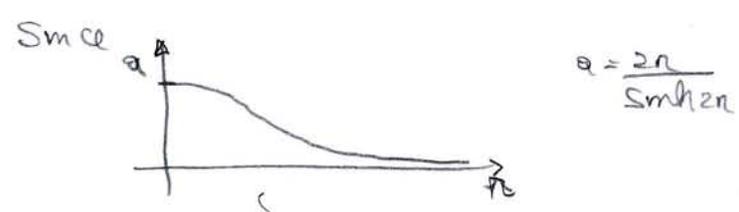
Now, can we compute the running of this coupling?
for this we need to propose some relation between

$$r \leftrightarrow \text{Energy.}$$

Cotrone, Bologno and Zaffaroni found that the gaugino condensate

$\langle \lambda \lambda \rangle$ was related with the function $a(r)$.

This is VERY natural (though computationally involved)



$$a = \frac{2r}{\text{Smth} 2r}$$

$$a(r) \Leftrightarrow \langle \lambda \lambda \rangle = \left(\frac{\Lambda}{\mu} \right)^3$$

$$a = \left(\frac{\Lambda}{\mu} \right)^3$$

this is what lattice person want to see for LHS in a simulation

Now, the point is to compute

$$\frac{d g_{YM}}{d \log \mu / \Lambda} = -\frac{1}{3} \frac{d g_{YM}}{d \log a}$$

with

$$a = \frac{2\pi}{5\pi h^2 n}$$

$$\frac{2\pi^2}{g_{YM}^2 N} = n \log n$$

This gives a complicated expression.

$$-3 \frac{d \sqrt{\frac{2\pi^2}{N n \log n}}}{d \log \left(\frac{2\pi}{5\pi h^2 n} \right)} = f(n)$$

We gain some understanding by looking at this expression for $n \rightarrow \infty$

using that for large n $a \sim 4n e^{-2n} = \left(\frac{\Lambda}{\mu}\right)^3$

$$\beta \sim \frac{3}{2} (1 + 4e^{-2n}) =$$

$$\beta \frac{2\pi^2}{g_{YM}^2 N} = -\frac{3 g_{YM}^3 N}{4\pi} \left(\frac{1}{1 - \frac{g_{YM}^2 N}{2\pi^2}} \right)$$

} NSVZ
} Jans

Criticisms

- Computation done at $n \rightarrow \infty$ \rightarrow effect of the KK modes? \downarrow
- $n \rightarrow \infty$ is the only easy probe they decouple from the computation \rightarrow unchanged in the (Λ/μ)
- \downarrow
- may be the formula is more general
- Solutions with \rightarrow stabilized solution give the same result. in an approximate way.

The spectrum of this model

one starts by considering the theory on a D5 brane.

(5+1)d field theory with 16 SUSY's. The global symmetry is

$$SO(1,5) \times SO(4) \cong SU(4) \times SU(2)_A \times SU(2)_B$$

The fields on the D5 brane $A_M, 4 \times \phi; \lambda, \bar{\lambda}$ (weyl)

Transform as

	$SU(4)$	\times	$SU(2)_A$	\times	$SU(2)_B$
A_M	6		1		1
ϕ_i	1		2		2
λ	4		1		2
$\bar{\lambda}$	$\bar{4}$		2		1

Then, we wrap the D5 on $S^2 \rightsquigarrow SO(1,5) \rightarrow SO(1,3) \times \underbrace{SO(2)}_{(4,5)}$

$$\cong SO(4) \times SO(2) \cong SU(2)_L \times SU(2)_R \times U(1)_{45}$$

$(A_M \rightarrow A_\mu; m_2 = A_4 \pm iA_5); (\lambda \rightarrow \lambda_\alpha + \psi_{\dot{\alpha}}; \bar{\lambda} \rightarrow \bar{\lambda}_{\dot{\alpha}} + \bar{\psi}_\alpha)$

	$SU(2)_L$	\times	$SU(2)_R$	\times	$U(1)_{45}$	\times	$SU(2)_A$	\times	$SU(2)_B$
A_μ	2		2		0		1		1
m_i	1		1		± 2		1		1
ϕ_i	1		1		0		2		2
λ	2		1		+1		2		1
ψ	2		1		-1		1		2
$\bar{\lambda}$	1		2		$\bar{1}$		1		2
$\bar{\psi}$	1		2		+1		2		1

Now, the twisting cones. In order to satisfy the
 must mix $U(1)_{45}$ with $SU(2)_A$ or $SU(2)_B$ (a $U(1)$ inside them)

	$U(1)_A$	$U(1)_T$	
A_μ	0	0	
m_\pm	0	± 2	$U(1)_T = \text{diag}(U(1)_A, U(1)_B)$
f	± 1	± 1	$Q_T = Q_{45} + Q_A$
λ_α	± 1	$\begin{pmatrix} 2 \\ 0 \end{pmatrix}$	
λ_i	± 1	$\begin{pmatrix} 0 \\ -2 \end{pmatrix}$	
ψ_α	0	-1	
$\bar{\psi}_i$	0	+1	

So, now, we compactify with $U(1)_T$ (instead of $U(1)_{45}$)

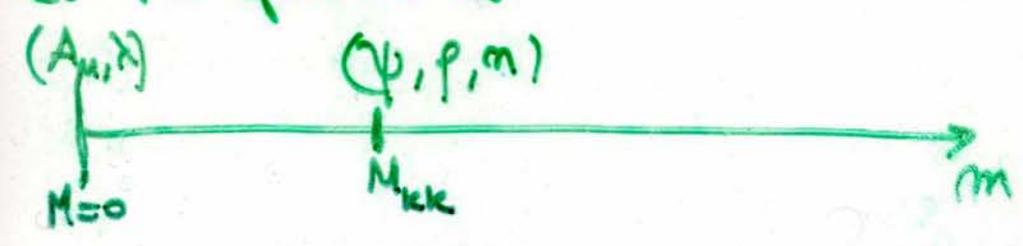
The T-Spin of each particle

$A_\mu, \lambda_\alpha, \bar{\lambda}_i$ T = Scalars $Q_T = 0$

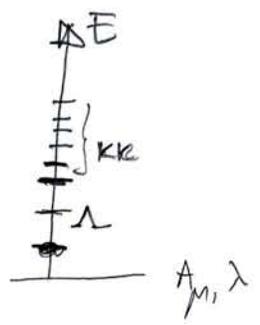
$\psi_\alpha, \bar{\psi}_i, f$ T = Spinors $Q_T = 1$

$m_\pm, \lambda_\alpha, \bar{\lambda}_i$ T = Vectors $Q_T = 2$

So the spectrum is



Now, one has a field theory whose spectrum of weak coupling is



one would like $\frac{\Lambda}{m_{KK}} \rightarrow 0$

but one actually gets $\frac{\Lambda}{m_{KK}} \sim m_{KK}$

This "degeneracy" between the strong coupling scale and the UV completion affects all supergravity duals. It is an effect of the supergravity approximation (note M theory for few N_{D0})

Now one may wonder, why so many tests are passed merely ~~while~~ ^{when} the effect of the KK modes could be noticeable?

~~Answer: ...~~

All the details to check when the KK modes do
(not) influence a computation one in hep-th/0505100
by Umur Gürsey and me.

We first write the full metric in the following appropriate form for the $SL(3, R)$ rotation in which the torus and the transverse parts have been separated as follows:

$$\begin{aligned}
 ds_{string}^2 = & e^\phi [dx_{1,3}^2 + \alpha' g_s N dr^2] + D_1 d\psi^2 + D_2 d\theta^2 + D_3 d\tilde{\theta}^2 + E_1 d\theta d\tilde{\theta} + \\
 & E_2 d\theta d\psi + E_3 d\tilde{\theta} d\psi + \frac{F}{\sqrt{\Delta}} [d\varphi + (\alpha_1 - C\beta_1)d\theta + (\alpha_2 - C\beta_2)d\tilde{\theta} + (\alpha_3 - C\beta_3)d\psi - Cd\tilde{\varphi}]^2 + \\
 & F\sqrt{\Delta}(d\tilde{\varphi} + \beta_1 d\theta + \beta_2 d\tilde{\theta} + \beta_3 d\psi)^2
 \end{aligned} \tag{1.1}$$

To simplify the notation let us define,

$$f = 4e^{2h} \sin^2 \theta + \cos^2 \theta + a^2 \sin^2 \tilde{\theta}, \quad g = a \sin \theta \sin \tilde{\theta} \cos \psi - \cos \theta \cos \tilde{\theta}. \tag{1.2}$$

Then, various functions in (1.1) are given as,

$$\begin{aligned}
 F &= \frac{\alpha' g_s N e^\phi}{4} \sqrt{f - g^2}, \quad \Delta = \frac{f - g^2}{f^2}, \quad C = \frac{g}{f}, \\
 \beta_1 &= \frac{f}{f - g^2} a \sin \psi \sin \tilde{\theta}, \quad \beta_2 = \frac{g}{f - g^2} a \sin \psi \sin \theta, \quad \beta_3 = \frac{f \cos \tilde{\theta} + g \cos \theta}{f - g^2} \\
 \alpha_1 &= \frac{a \sin \tilde{\theta} \sin \psi g}{f - g^2}, \quad \alpha_2 = \frac{a \sin \theta \sin \psi}{f - g^2}, \quad \alpha_3 = \frac{\cos \theta + g \cos \tilde{\theta}}{f - g^2} \\
 D_1 &= \frac{\alpha' g_s N e^\phi}{4(f - g^2)} (f \sin^2 \tilde{\theta} - g^2 - \cos^2 \theta - 2g \cos \theta \cos \tilde{\theta}), \\
 D_2 &= \frac{\alpha' g_s N e^\phi}{4} (a^2 + 4e^{2h} - \frac{f}{f - g^2} a^2 \sin^2 \psi \sin^2 \tilde{\theta}) \\
 D_3 &= \frac{\alpha' g_s N e^\phi}{4} (1 - \frac{a^2 \sin^2 \theta \sin^2 \psi}{f - g^2}), \quad E_1 = \frac{a}{2} \alpha' g_s N e^\phi (\cos \psi - \frac{g}{f - g^2} a \sin^2 \psi \sin \theta \sin \tilde{\theta}) \\
 E_2 &= -\frac{\alpha' g_s N e^\phi a \sin \psi \sin \tilde{\theta} (f \cos \tilde{\theta} + g \cos \theta)}{2(f - g^2)}, \quad E_3 = -\frac{\alpha' g_s N e^\phi a \sin \theta \sin \psi}{2(f - g^2)} (\cos \theta + g \cos \tilde{\theta}).
 \end{aligned} \tag{1.3}$$

Now, let us focus on the RR two form. It is useful to define four one forms as $\mathcal{A}^{(i)}, \mathcal{C}^{(i)}$, $i = 1, 2$

$$\mathcal{A}^{(1)} = \alpha_1 d\theta + \alpha_2 d\tilde{\theta} + \alpha_3 d\psi, \quad \mathcal{A}^{(2)} = \beta_1 d\theta + \beta_2 d\tilde{\theta} + \beta_3 d\psi, \tag{1.4}$$

and

$$\mathcal{C}^{(1)} = C_\theta^{(1)} d\theta + C_{\tilde{\theta}}^{(1)} d\tilde{\theta} + C_\psi^{(1)} d\psi, \quad \mathcal{C}^{(2)} = C_\theta^{(2)} d\theta + C_{\tilde{\theta}}^{(2)} d\tilde{\theta} + C_\psi^{(2)} d\psi, \tag{1.5}$$

and the two form $\tilde{c}_{\mu\nu} dx^\mu \wedge dx^\nu$, with components $\tilde{c}_{\theta\tilde{\theta}}, \tilde{c}_{\tilde{\theta}\psi}$, and all others being zero.

more details

In order to reproduce the RR two-form to the following form,

$$C^{(2)} = C_{12} D\varphi \wedge D\tilde{\varphi} - C^{(1)} \wedge D\varphi - C^{(2)} \wedge D\tilde{\varphi} - \frac{1}{2}(A^{(1)} \wedge C^{(1)} + A^{(2)} \wedge C^{(2)} - \tilde{c}); \quad (1.6)$$

the quantities defined above are determined as,

$$\begin{aligned} C_{12} &= \frac{1}{4}(a(r) \sin \theta \sin \tilde{\theta} \cos \psi - \cos \theta \cos \tilde{\theta}), \\ C^{(1)} &= \left(-\frac{\psi}{4} \sin \theta - C_{12}\beta_1\right)d\theta - \left(\frac{1}{4}a(r) \sin \theta \sin \psi + C_{12}\beta_2\right)d\tilde{\theta} - \beta_3 C_{12}d\psi, \\ C^{(2)} &= \left(\frac{1}{4}a(r) \sin \tilde{\theta} \sin \psi + C_{12}\alpha_1\right)d\theta + \left(\frac{\psi}{4} \sin \tilde{\theta} - C_{12}\alpha_2\right)d\tilde{\theta} + \alpha_3 C_{12}d\psi, \\ \tilde{c} &= \tilde{c}_{\theta\tilde{\theta}}d\theta \wedge d\tilde{\theta} + \tilde{c}_{\theta\psi}d\theta \wedge d\psi + \tilde{c}_{\tilde{\theta}\psi}d\tilde{\theta} \wedge d\psi \end{aligned} \quad (1.7)$$

with

$$\begin{aligned} \tilde{c}_{\theta\tilde{\theta}} &= \alpha_1\alpha_2 + \beta_1\beta_2 + C_{12}(\alpha_1\beta_2 - \alpha_1\beta_2) - \frac{1}{4}\psi \sin \tilde{\theta}\beta_1 + \frac{1}{4}a \sin \theta \sin \psi \alpha_1 \\ &\quad - \frac{1}{4}a \cos \psi + \frac{1}{4}a\beta_2 \sin \psi \sin \tilde{\theta}, \\ \tilde{c}_{\theta\psi} &= \alpha_1\alpha_3 + \beta_1\beta_3 - C_{12}(\alpha_1\beta_3 - \alpha_3\beta_1) - \frac{1}{4}\psi \sin \theta \alpha_3 + \frac{1}{4}a \sin \psi \sin \tilde{\theta}\beta_3, \\ \tilde{c}_{\tilde{\theta}\psi} &= \alpha_2\alpha_3 + \beta_2\beta_3 - C_{12}(\alpha_2\beta_3 - \alpha_3\beta_2) + \frac{1}{4}\psi \sin \tilde{\theta}\beta_3 - \frac{1}{4}a \sin \psi \sin \theta \alpha_3. \end{aligned} \quad (1.8)$$

So, after the $SL(3, R)$ rotation is done, we will end with a metric that reads

$$\begin{aligned} (ds_{string}^2)' &= \left(\frac{e^{2(\phi'-\phi)}F}{F'}\right)^{1/3} \left(e^\phi [dx_{1,3}^2 + \alpha' g_s N dr^2] + D_1 d\psi^2 + D_2 d\theta^2 + D_3 d\tilde{\theta}^2 + E_1 d\theta d\tilde{\theta} + \right. \\ &\quad \left. E_2 d\theta d\psi + E_3 d\tilde{\theta} d\psi \right) + \frac{F'}{\sqrt{\Delta}} [d\varphi + (\alpha_1 - C\beta_1)d\theta + (\alpha_2 - C\beta_2)d\tilde{\theta} + (\alpha_3 - C\beta_3)d\psi - Cd\tilde{\varphi}]^2 \\ &\quad + F'\sqrt{\Delta} (d\tilde{\varphi} + \beta_1 d\theta + \beta_2 d\tilde{\theta} + \beta_3 d\psi)^2 \end{aligned} \quad (1.9)$$

↳ using this metric and the RR fields above
one can disentangle when/how the KK
modes influence on computations

Many checks for this in hep-th/0505100.